

# Random CMB Stuff

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**References:** arXiv:0802.3688

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The  $C_\ell$ 's can be defined as:

$$C_\ell = 4\pi \int d \ln k j_\ell^2(k D_*) \Delta_T^2(k). \quad (1)$$

The *dimensionless* power spectrum  $\Delta^2$  is defined from the power spectrum  $P(k)$  as:

$$\Delta_T^2(k) = \frac{k^3}{2\pi^2} P(k). \quad (2)$$

The power spectrum is *defined* from the two-point function as:

$$\langle \Theta(\mathbf{k})^* \Theta(\mathbf{k}') \rangle = (2\pi)^3 \delta(\mathbf{k} - \mathbf{k}') P(k). \quad (3)$$

Let's write

$$\Theta(\eta, \mathbf{k}) = \Theta(0, \mathbf{k}) \tilde{\Theta}(\eta, \mathbf{k}), \quad (4)$$

such that  $\tilde{\Theta}(0, \mathbf{k}) = 1$ . Then we have:

$$\langle \Theta(\mathbf{k})^* \Theta(\mathbf{k}') \rangle = \left\langle \Theta(0, \mathbf{k})^* \tilde{\Theta}(\eta, \mathbf{k})^* \Theta(0, \mathbf{k}') \tilde{\Theta}(\eta, \mathbf{k}') \right\rangle. \quad (5)$$

The  $\tilde{\Theta}$  are deterministic and independent of the initial conditions, so we can remove these from the correlator, which is an ensemble average over realisations of  $\Theta(\mathbf{k})$ .

$$\langle \Theta(\mathbf{k})^* \Theta(\mathbf{k}') \rangle = \langle \Theta(0, \mathbf{k})^* \Theta(0, \mathbf{k}') \rangle \tilde{\Theta}(\eta, \mathbf{k})^* \tilde{\Theta}(\eta, \mathbf{k}'). \quad (6)$$

The initial temperature fluctuations can be related to the initial curvature perturbations  $\zeta$  as

$$\Theta(0, \mathbf{k}) = -\frac{1}{5} \zeta(\mathbf{k}). \quad (7)$$

$$\langle \Theta(\mathbf{k})^* \Theta(\mathbf{k}') \rangle = \langle \Theta(0, \mathbf{k})^* \Theta(0, \mathbf{k}') \rangle \tilde{\Theta}(\eta, \mathbf{k})^* \tilde{\Theta}(\eta, \mathbf{k}') \quad (8)$$

$$= \frac{1}{25} \langle \zeta(\mathbf{k}) \zeta(\mathbf{k}')^* \rangle \tilde{\Theta}(\eta, \mathbf{k})^* \tilde{\Theta}(\eta, \mathbf{k}') \quad (9)$$

The power spectrum of initial curvature perturbations is:

$$\langle \zeta(\mathbf{k}) \zeta(\mathbf{k}')^* \rangle = (2\pi)^3 \frac{2\pi^2}{k^3} \Delta_\zeta^2(k) \delta(\mathbf{k} - \mathbf{k}'). \quad (10)$$

This gives:

$$\langle \Theta(\mathbf{k})^* \Theta(\mathbf{k}') \rangle = \frac{(2\pi)^3}{25} \frac{2\pi^2}{k^3} \Delta_\zeta^2(k) \delta(\mathbf{k} - \mathbf{k}') \tilde{\Theta}(\eta, \mathbf{k})^* \tilde{\Theta}(\eta, \mathbf{k}') . \quad (11)$$

Comparison with the definition of  $P(k)$  for the temperature fluctuations gives:

$$P(k) = \frac{\Delta_\zeta^2(k)}{25} \frac{(2\pi)^3}{4\pi k^3} \left| \tilde{\Theta}(\eta, \mathbf{k}) \right|^2 . \quad (12)$$

This gives a dimensionless power spectrum which is:

$$\Delta_T^2 = \frac{\Delta_\zeta^2(k)}{25} \left| \tilde{\Theta}(\eta, \mathbf{k}) \right|^2 . \quad (13)$$

We then just evaluate at recombination  $\eta = \eta_*$ .

Let's define the transfer function which takes us from initial temperature fluctuations to observed temperature fluctuations:

$$T(k) = \left| \tilde{\Theta}(\eta_*, \mathbf{k}) \right|^2 . \quad (14)$$

Thus, the CMB temperature anisotropy power spectrum is:

$$C_\ell = \frac{4\pi}{25} \int d \ln k j_\ell^2(k D_*) \Delta_\zeta^2(k) T(k) . \quad (15)$$

The final ingredient is the spectrum of initial curvature perturbations:

$$\Delta_\zeta^2(k) = A_s \left( \frac{k}{k_*} \right)^{n_s - 1} , \quad (16)$$

where  $A_s \approx 2.196 \times 10^{-9}$ ,  $k_* = 0.05 \text{ Mpc}^{-1}$  and  $n_s \approx 0.9603$ .