## On The Recovery of Planetary Energy Functions using Deep Neural Networks

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## 0.1 Problem Formulation

Two-body problem: in the previous formulation,

$$\min_{\phi} \sum_{t=1}^{T} \| \dot{\boldsymbol{r}}_{t} \cdot \nabla \phi \left( \boldsymbol{r}_{t} \right) \|_{2}^{2} \tag{1}$$

This converged in less than 3 epochs and did not produce any meaningful  $\phi()$ . As found last week the solution found would simply ignore any inputs and return 0

As discussed we experimented with adding a term to maximize the gradients of  $\phi()$  yielding the minimization:

$$\min_{\phi} \sum_{t=1}^{T} \left( \| \dot{\boldsymbol{r}}_{t} \cdot \nabla \phi \left( \boldsymbol{r}_{t} \right) \|_{2}^{2} + \frac{1}{\| \nabla \phi \left( \boldsymbol{r}_{t} \right) \|_{2}^{2}} \right)$$
 (2)

However as  $\nabla \phi$ () was highly variant solving this encountered issues with numeric stability and would underflow/overflow for various training hyper-parameters.

To combat the numerical stability we experimented with log of the norm of the gradient of phi:

$$\min_{\phi} \sum_{t=1}^{T} \left( \| \dot{\boldsymbol{r}}_{t} \cdot \nabla \phi \left( \boldsymbol{r}_{t} \right) \|_{2}^{2} - \log \left( \| \nabla \phi \left( \boldsymbol{r}_{t} \right) \|_{2}^{2} \right) \right)$$
(3)

This minimization again suffered from numerical stability.

Given our goal of minimizing the dot product  $\vec{r}_t \cdot \nabla \phi(r_t)$  we minimized the cosine distance between  $\vec{r}_t$  and  $\nabla \phi(r_t)$  and instead of maximizing the gradient we simple enforce that  $\|\nabla \phi(r_t)\| = 1$  along the trajectory:

$$\min_{\phi} \sum_{t=1}^{T} \left( \left| \frac{\dot{\boldsymbol{r}}_{t} \cdot \nabla \phi \left( \boldsymbol{r}_{t} \right)}{\|f_{t}\|_{2}^{2} * \| \nabla \phi \left( \boldsymbol{r}_{t} \right) \|_{2}^{2}} \right| - \left( 1 - \left\| \nabla \phi \left( \boldsymbol{r}_{t} \right) \right\|_{2}^{2} \right)^{2} \right)$$
(4)

This objective results in a well defined gradient for minimization. Producing a phi that is constant across  $r \in T$ . However this phi after convergence decreases consistently for each epoch across multiple planets. This may be a result of over-fitting however as it is only seen after 100k epochs for a small 50 unit fully connected neural network. I will post a histogram of phi as a function of training to validate these claims.

One of the potential concerns with the approach in (5) was unit-gradient magnitude was only enforced along the trajectory. We are currently testing the effect on regularizing the entire field to be unit length:

$$\min_{\phi} \sum_{t=1}^{T} \left| \frac{\dot{\mathbf{r}}_{t} \cdot \nabla \phi \left( \mathbf{r}_{t} \right)}{\|f_{t}\|_{2}^{2} * \|\nabla \phi(\mathbf{r}_{t})\|_{2}^{2}} \right| - \sum_{x \in R^{4}} \left( 1 - \|\nabla \phi(x)\|_{2}^{2} \right)^{2}$$
 (5)

## 0.2 Experiments