

PHY407: Lab 1

Date: September 17th. 2021

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Contributions:

- Q3. and the pseudocode for Q1b, Q2a completed by Brendan Halliday
- Q1, Q2 code, plots and explanatory notes completed by Summer Eldridge

Q1b.

Pseudocode:

```
# import important libraries
# store useful constants and initial values (i.e. G and initial positions and velocities)
# initialize an empty lists for x and y positions and velocities
# initialize an empty time list as well

# repeat this next section of code for as many increments of t as necessary

    # append initial time, x and y positions and velocities to their corresponding list
    # update time, x and y positions and velocities using the equations derived in part Q1 part a
    # i.e. use  $r = (x^2 + y^2)^{1/2}$ ,  $v_x = v_x - dt * G * x / r^3$ ,  $v_y = v_y - dt * G * y / r^3$ ,  $x = x + dt * v_x$ , and  $y = y + dt * v_y$ 
    # append updated values to their corresponding list

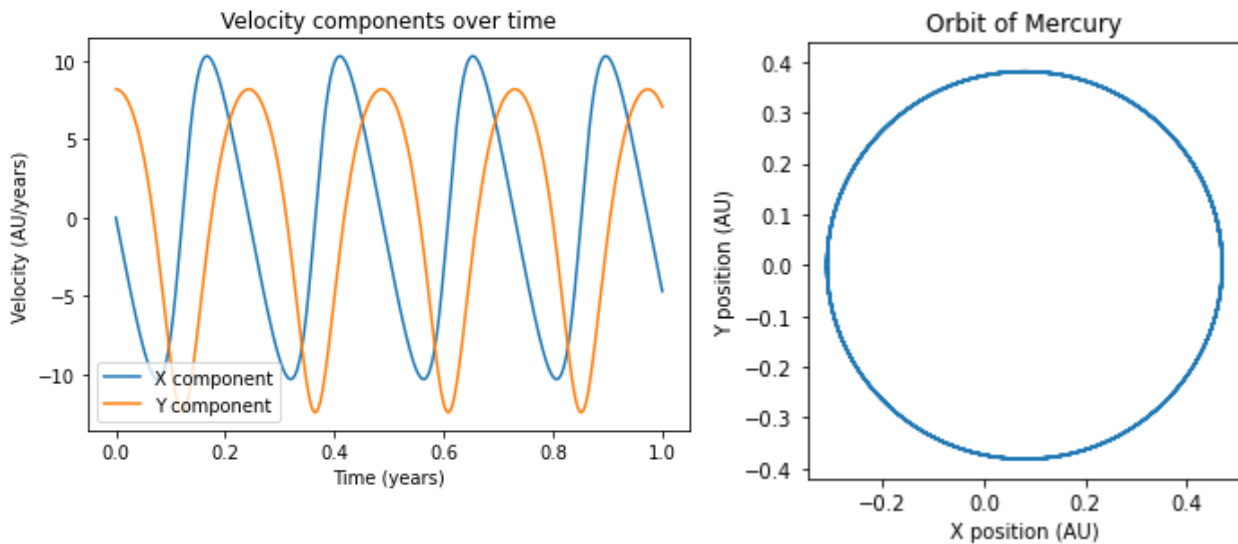
# plot orbits using list values previously calculated
# plot velocity components as a function of time
```

Q1c.

Code:

Code for this program is included as Lab01_Q1_c.py

Plots:

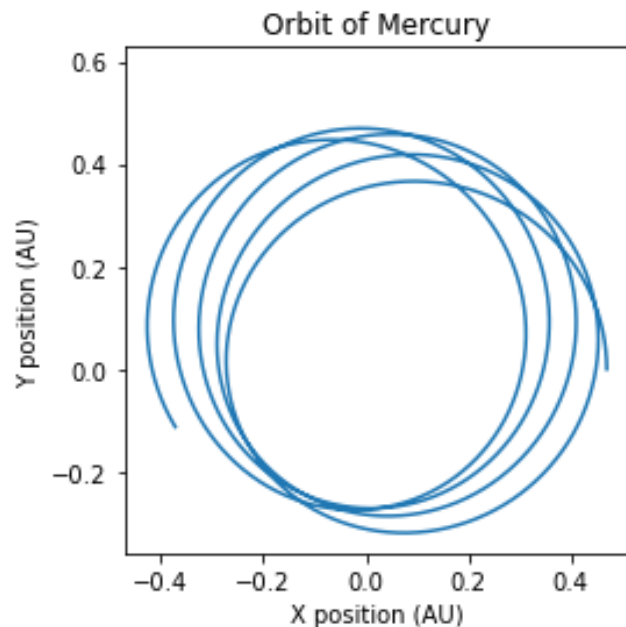


Q1d:

Code:

Code for this program is included as Lab01_Q1_d.py

Plots:



Q2a

Pseudocode:

```
# store useful constants such as G, Mass of Jupyter(Mj)
# and initial positions and velocities for both Earth and Jupyter
# where (ex, evx, jx, jvx) and (ey, evy, jy, jvy) are the positions and velocities of the x component and y
# component
# of Earth (e) and Jupyter (j) respectively.
# initialize T=10 and time increment dt=0.0001
# initialize empty lists for the x and y positions and velocities for both planets and time

#repeat next section for T/dt time increments

    #append updated positions and velocities of Jupyter and Earth to their corresponding lists

    #update the position and velocity of Jupyter using
    # jr = (jx ** 2 + jy ** 2) ** (1 / 2)
    # jvx = jvx - dt * G * jx * jr ** (-3)
    # jvy = jvy - dt * G * jy * jr ** (-3)
    # jx = jx + dt * jvx
    # jy = jy + dt * jvy

    # update the position and velocity of earth using:
    # the distance of Earth from the sun er = (ex ** 2 + ey ** 2) ** (1 / 2)
    # evx = evx - dt * (G * ex * er ** (-3) + G * Mj * (ex - jx) * ejr ** (-3))
    # evy = evy - dt * (G * ey * er ** (-3) + G * Mj * (ey - jy) * ejr ** (-3))
    # ex = ex + dt * evx
    # ey = ey + dt * evy
    # where we take into account the gravitational effect of Jupyter by using
```

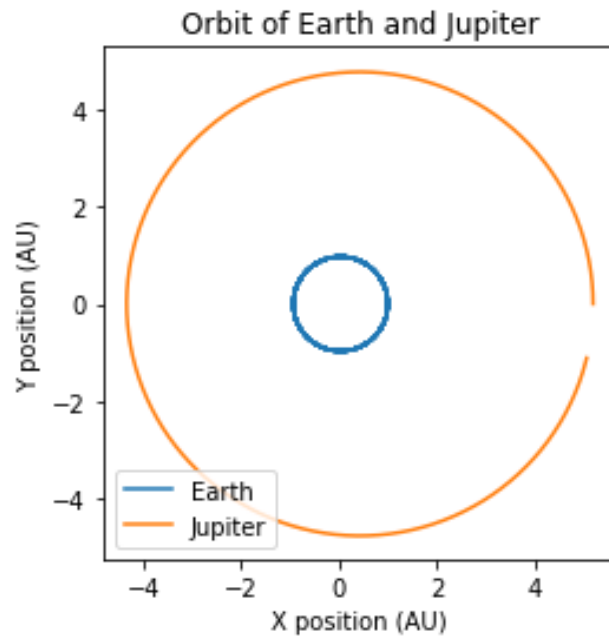
```
# the distance between the two planets ejr = ((ex - jx) ** 2 + (ey - jy) ** 2) ** (1 / 2)
```

```
#plot the orbits for both Earth and Jupyter
```

Code:

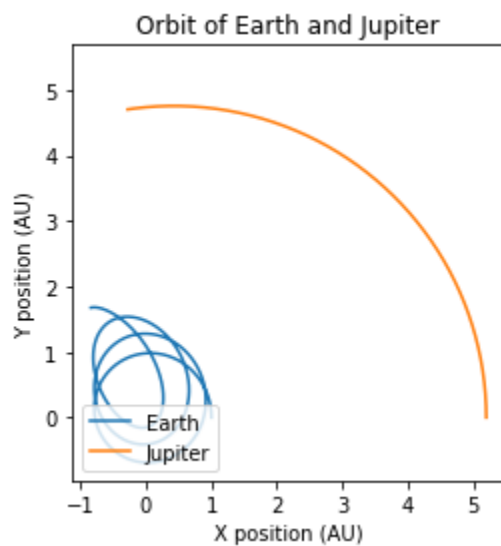
Code for this program is included as Lab01_Q2_a.py

Plots:



Q2b.

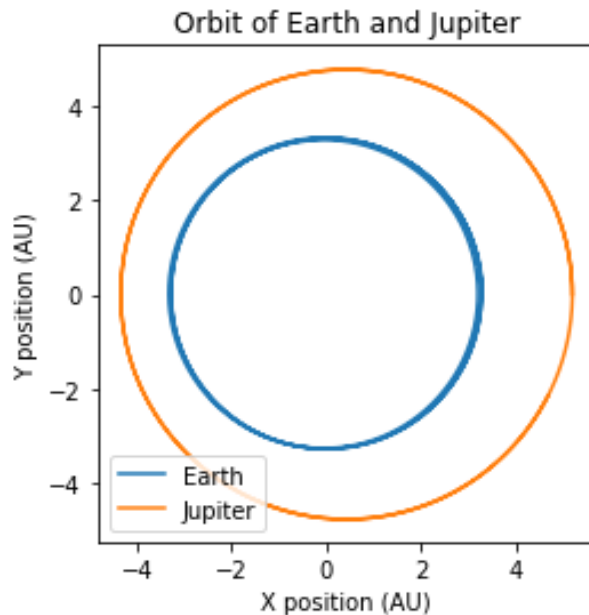
Plots:



If you go farther, earth swings too close to the sun and gets flung out of the solar system

Q2c.

Plots:



Q3a.

Pseudocode:

```
# import libraries and helper module Lab1_Q3
# store useful constants such as values for M and N

# use my_bins() (see appendix) function to generate 1000 bins between -5 to 5

# initialize empty list my_TIME for storing the time values
# for my_hist() (see appendix)

# This next block should be iterated for each value of N

    # generate R which is an array of N Gaussian distributed numbers
    # time my_hist() which is my histogram program from Lab1_Q3_helper_functions.py
    # append runtime to array my_TIME

#Appendix(These functions are defined at the beginning of Lab01_Q3_b.py)
```

```

#define function my_hist():
#This function takes in an array R of values and returns the histogram of those values
#initialize empty list
# repeat the code below for each bin
    # initialize accumulator to integer value zero for each bin
    # repeat the code below for each value in R
        # for each j in R, check if value is within
        # the bin under question
        # if so, add one to the accumulator
        # if not, do not change accumulator value
    # once each j has been checked, append accumulator value to empty list

#define my_bins()
# this function takes in an interval and number of bins and returns an array of equally spaced bins on
a given interval
# np.linspace is used

```

Timing Values for my_hist():

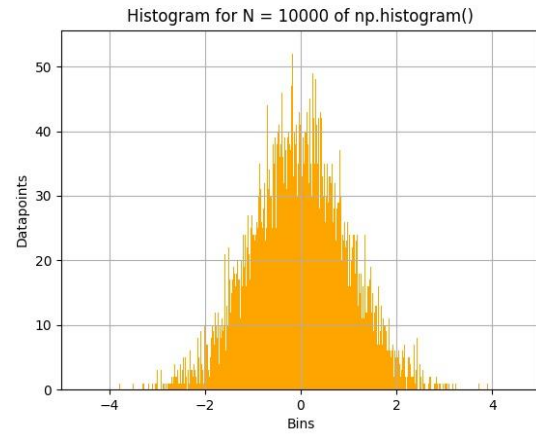
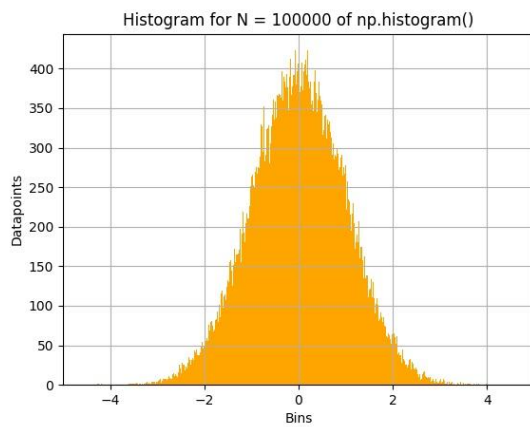
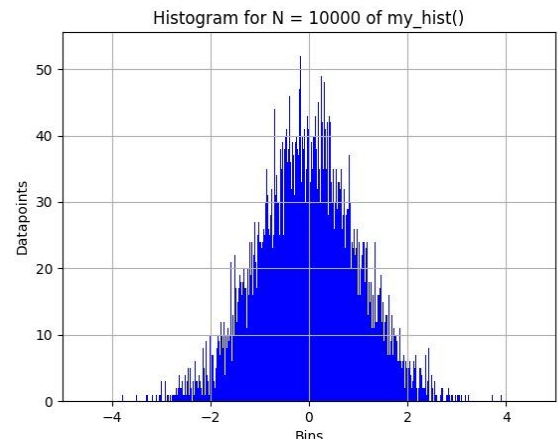
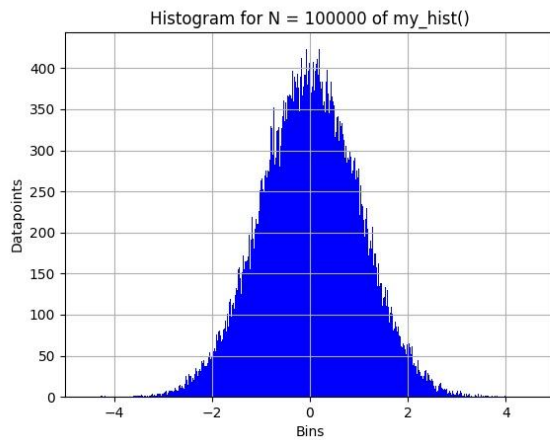
```

N=10, Time = 0.007s
N=100, Time = 0.07s
N=1000, Time= 0.6s
N=10000, Time= 6.7s
N=100000, Time= 63.1s
N=1000000, Time= 618.7s

```

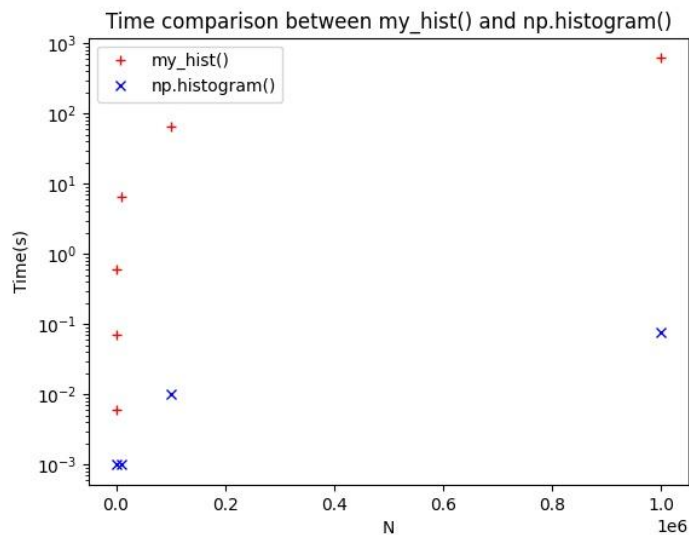
Q3b.

The following comparison shows that the two functions(my_hist() and np.histogram())produce identical results.



Timing graph:

The time graph below compares the speeds of the two functions where np.histogram is shown to be faster. This discrepancy is due to the fact that np.histogram uses a more efficient algorithm.



Code:

The code for this program is included in Lab01_Q3_b.py