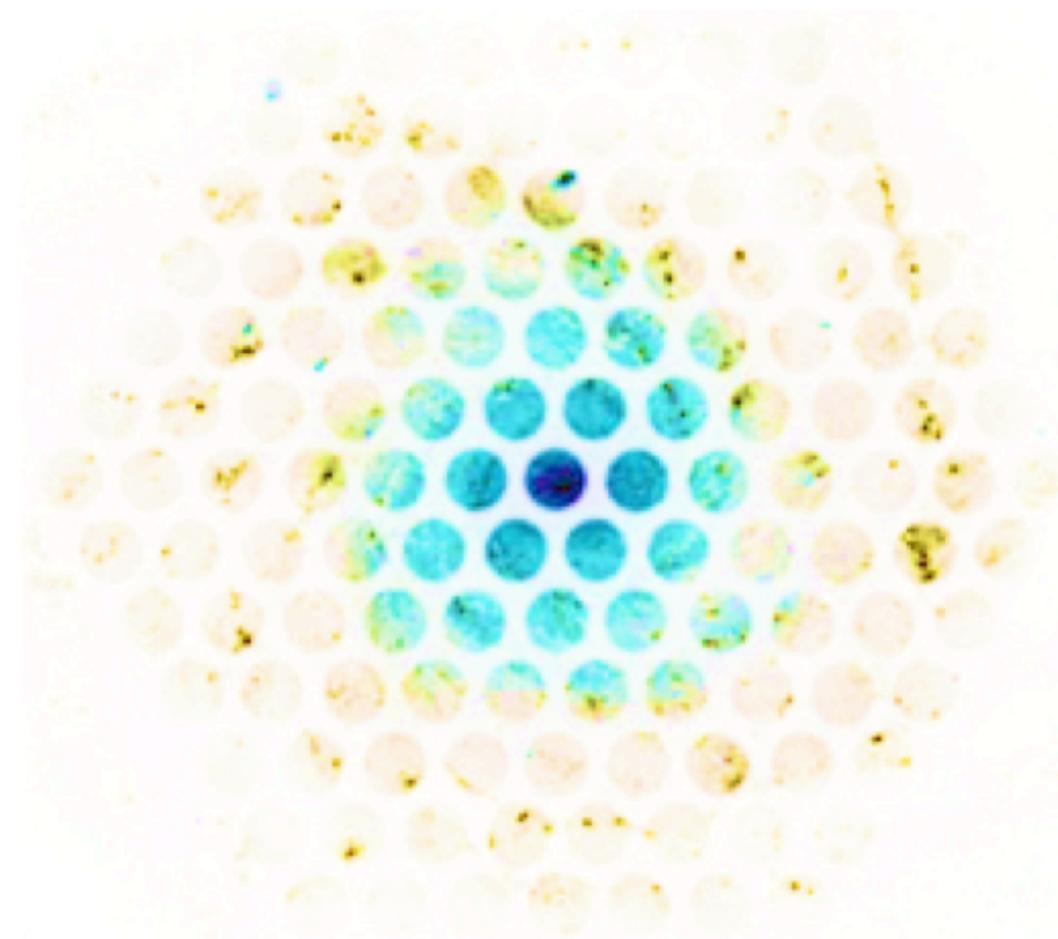


# Hidden in Plain Sight: a Deep Learning Approach to Finding Supernovae in Galaxy Maps



Brett Andrews

Physics & Astronomy

10.18.2018



University of Pittsburgh



# Type Ia Supernovae: Measuring the Accelerating Expansion of the Universe

- Accelerating expansion of the Universe
  - thought to be driven by dark energy
  - Nobel Prize in Physics 2011
- standardize brightness → measure distances
- Thermonuclear explosion of a white dwarf star.
  - White Dwarf: old, compact star as massive as Sun but size of Earth.
  - Normally stable but adding mass past a critical threshold makes them explode.

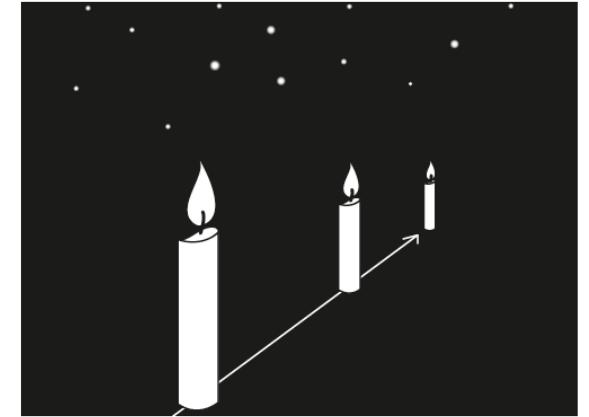
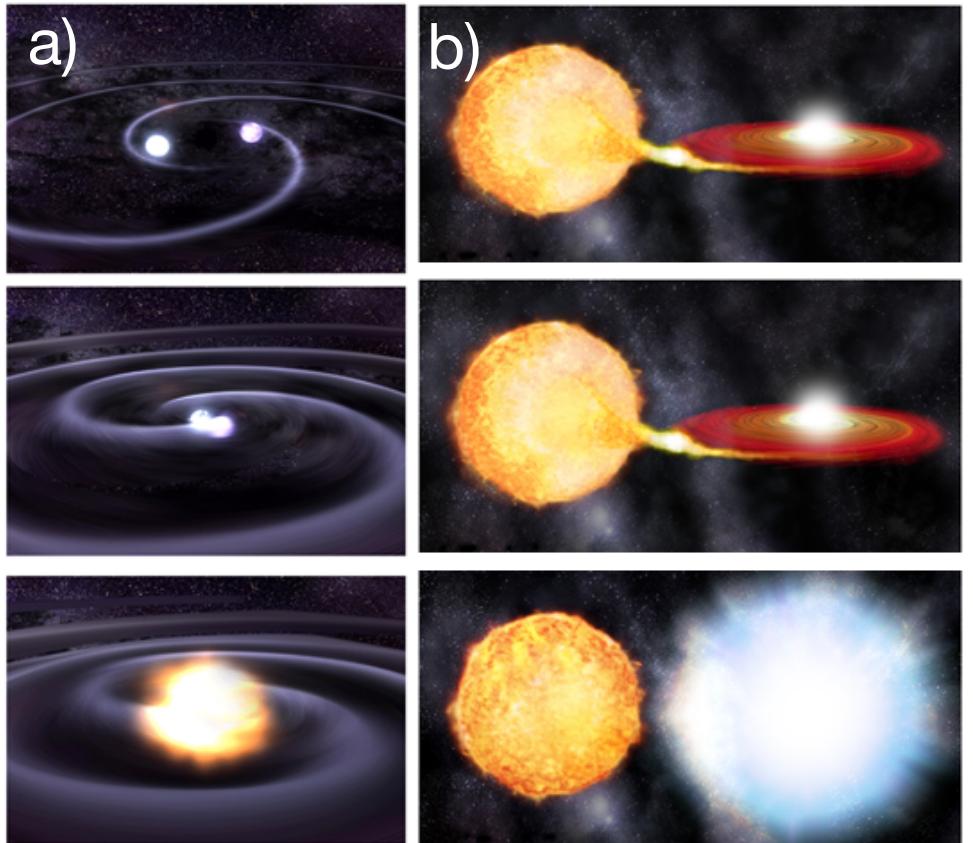


Image: nobelprize.org

# What is the companion star?

- a) another white dwarf
  - b) a normal star
- 
- Limits distance accuracy.
  - Might introduce systematic biases.
  - The **stellar population age** can help disentangle these two scenarios.



Images: Wikimedia Commons and Discover Magazine

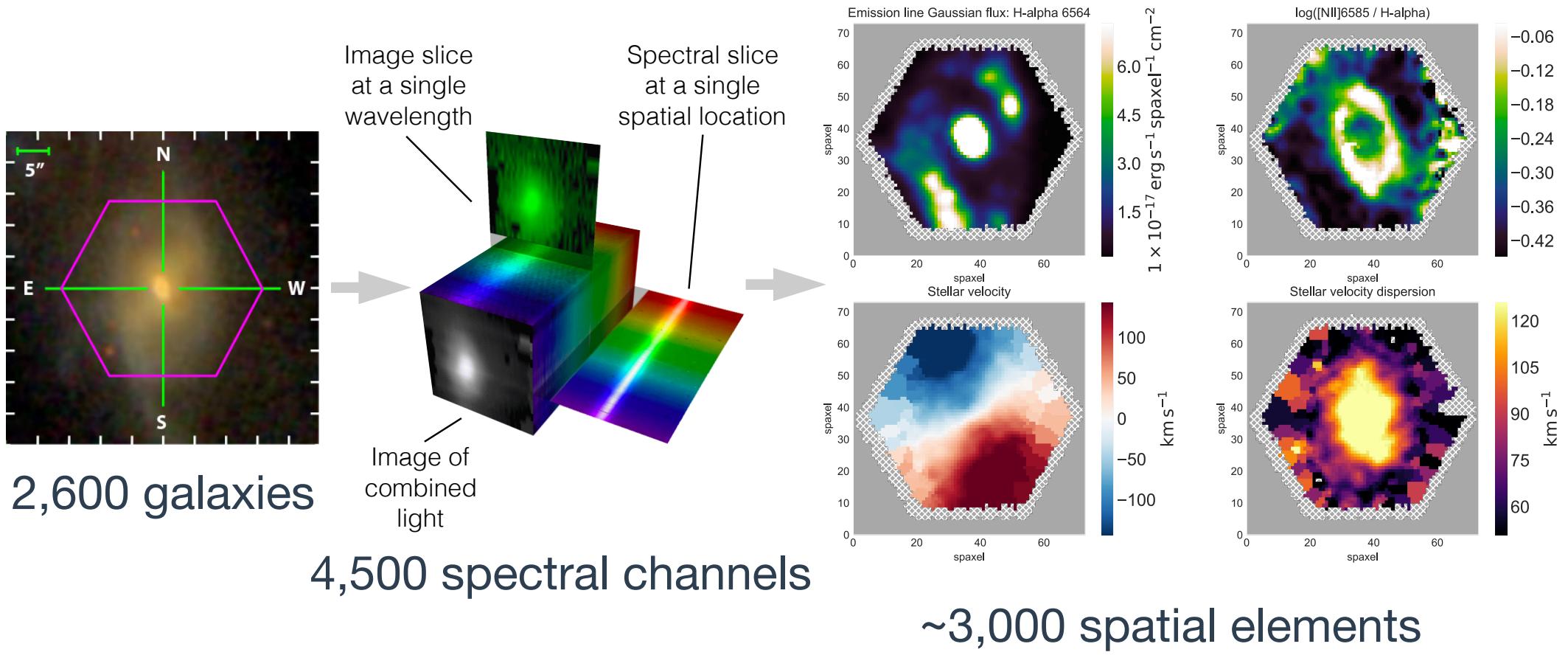
# Collaborators

- **Quanbin (Eric) Ma** (CMU Machine Learning; Facebook)
- Barnabas Poczos (CMU Machine Learning)
- **Qiong Zhang** (UBC Statistics)
- **Bo Chang** (UBC Statistics)
- Siamak Ravanbakhsh (UBC Computer Science)
- Ananth Tennenri (CMU Physics)

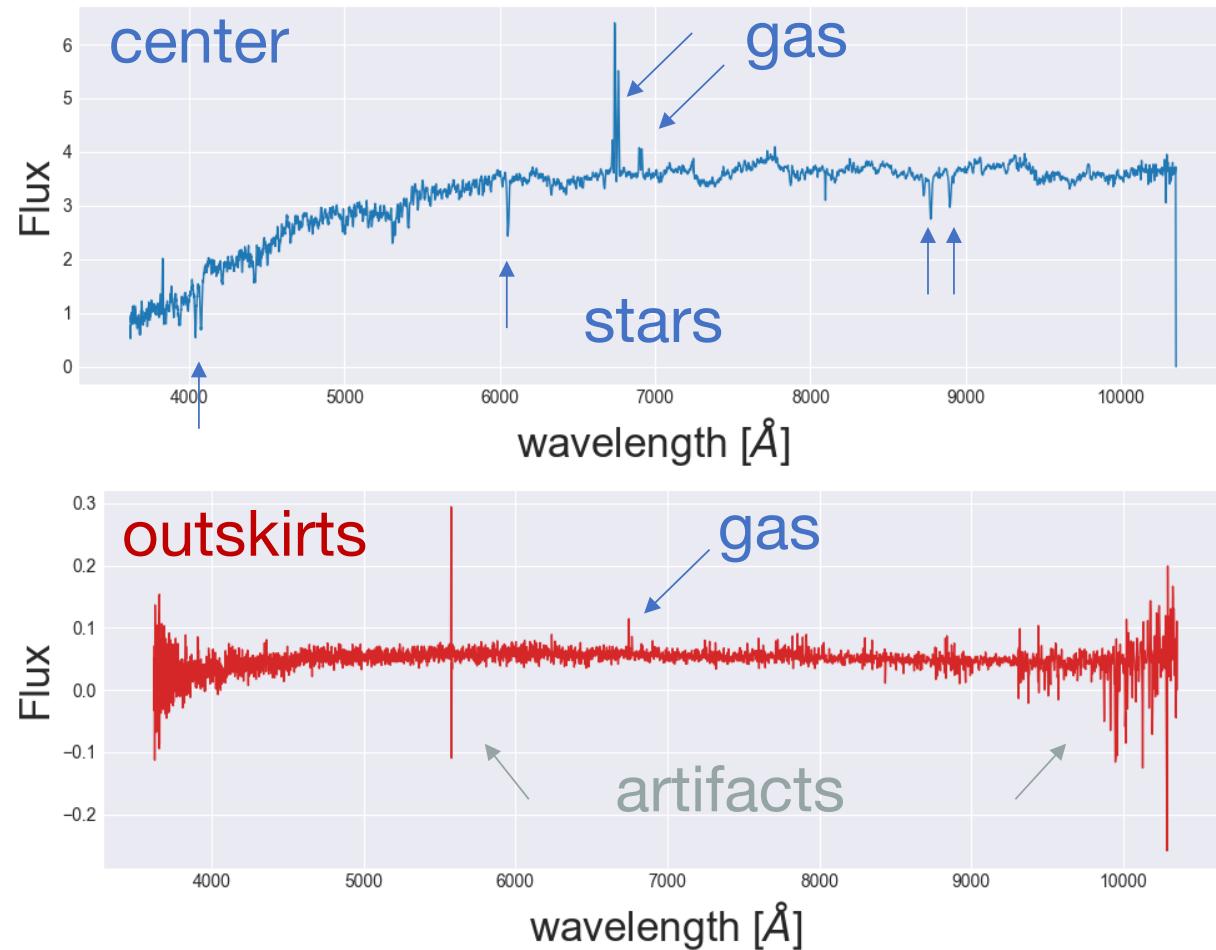
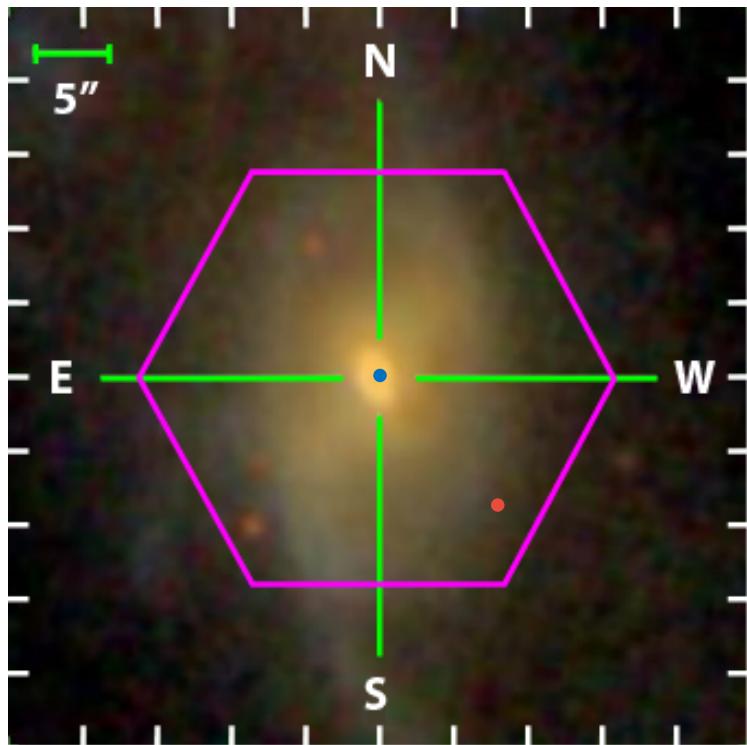
# Supernova Host Galaxy Spectroscopy

- Find supernovae in imaging and do resource intensive spatially-resolved spectroscopy of the host galaxy.
- Have a large data set of spatially-resolved spectroscopy of galaxies then look for supernovae (known or unknown).

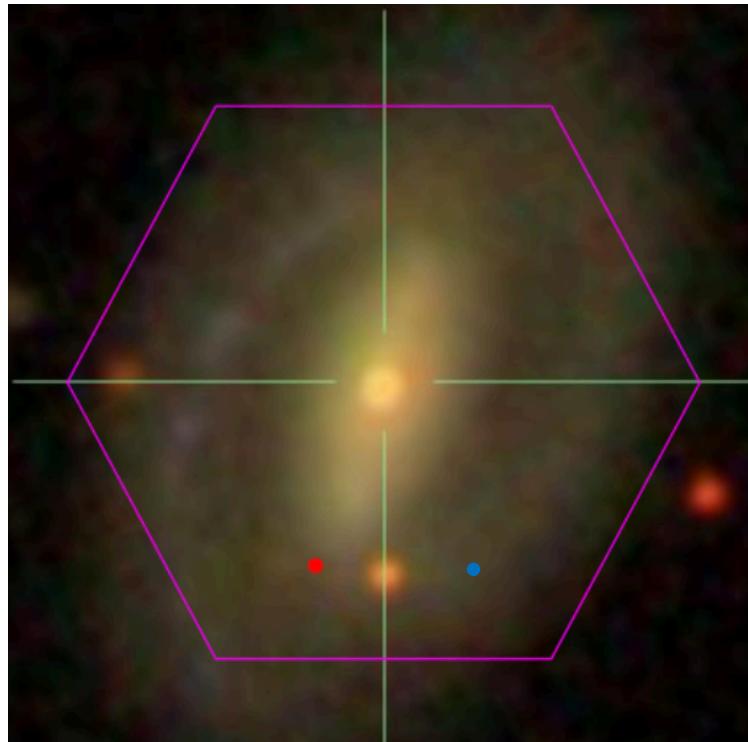
# SDSS-IV MaNGA Survey



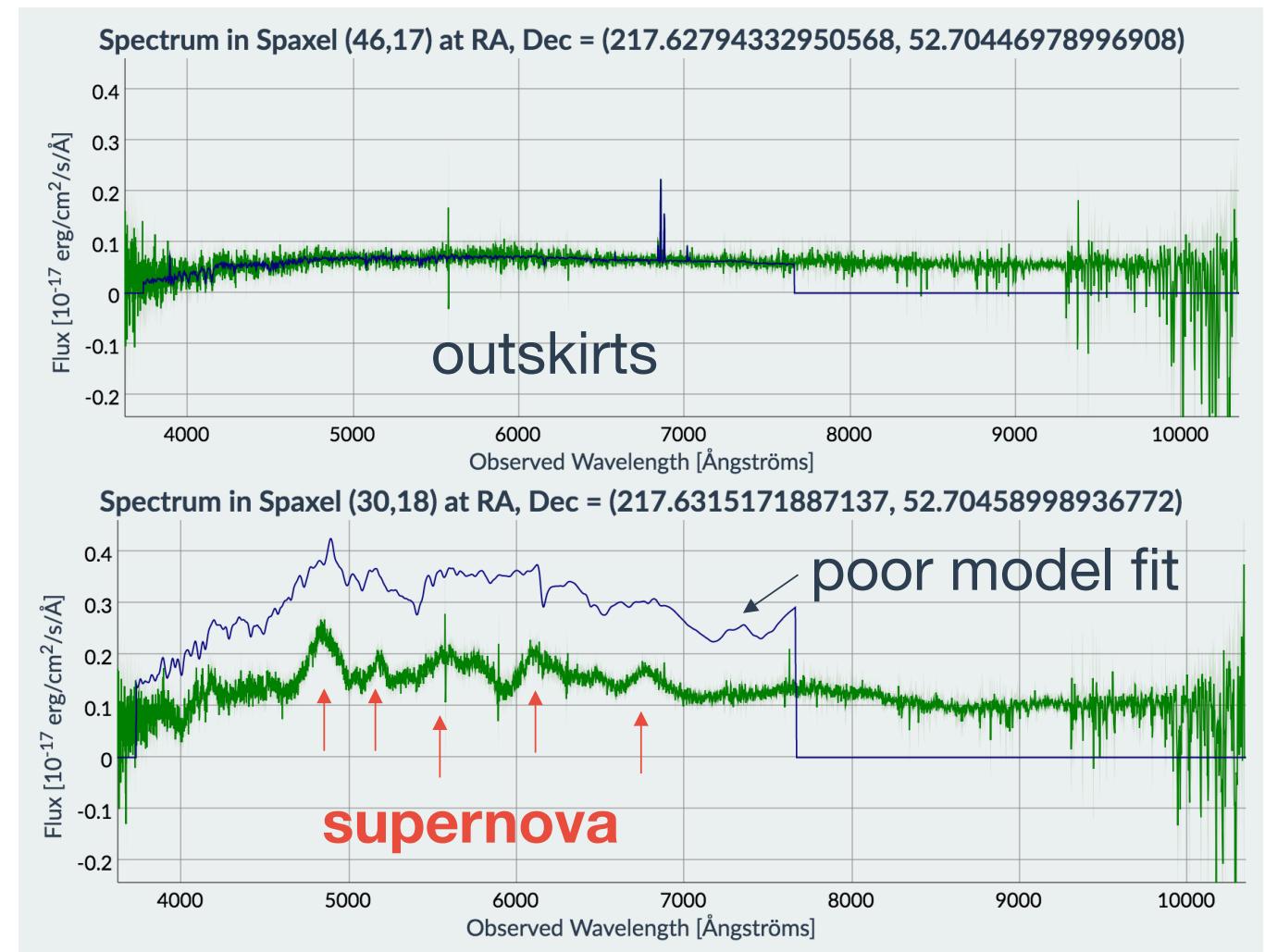
# Galaxy Spectra



# A Serendipitous Supernova



No supernova in image  
used to select galaxies.



# Finding Supernovae: Anomaly Detection

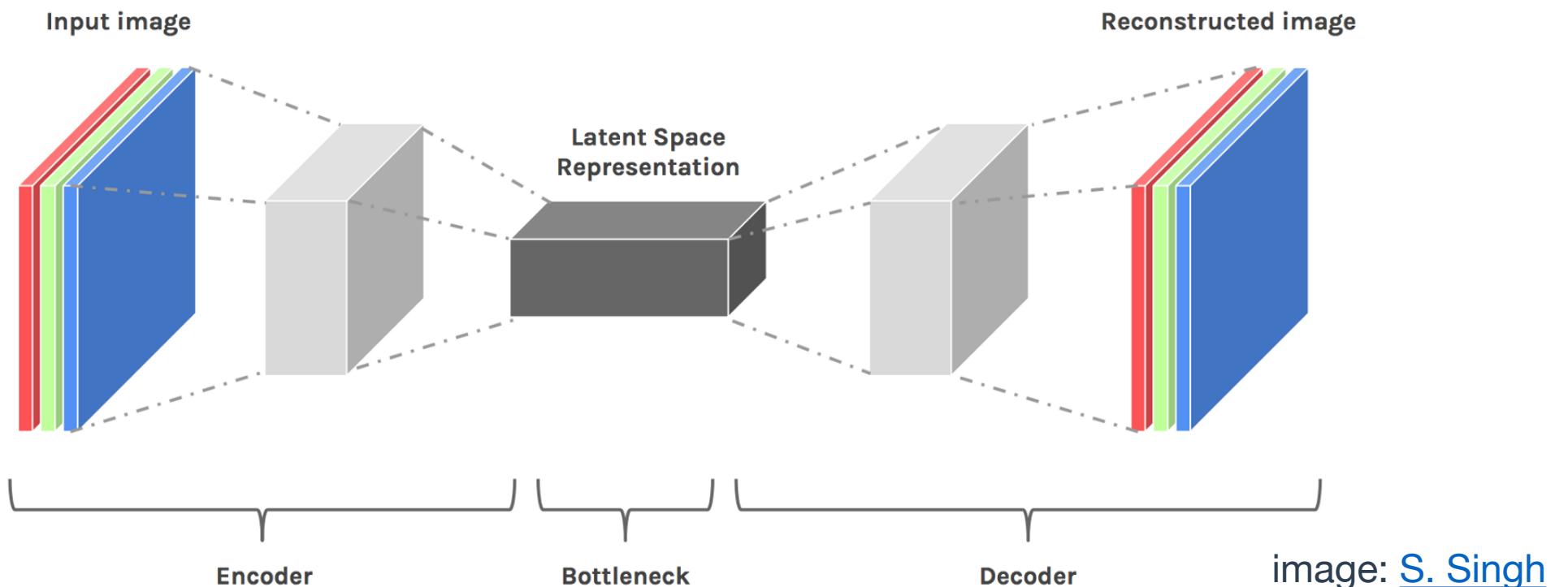
- Data set is too big to apply anomaly detection.

**2,600** galaxies \* **3,000** spatial elements \* **4,000** spectral channels  
~ **30,000,000,000** elements

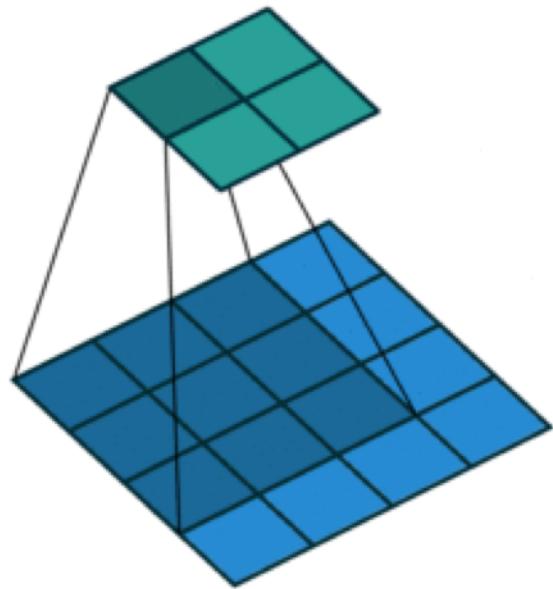
1. Reduce dimensionality with Convolutional Autoencoder.
2. Apply anomaly detection techniques.

# Convolutional Autoencoder (CAE)

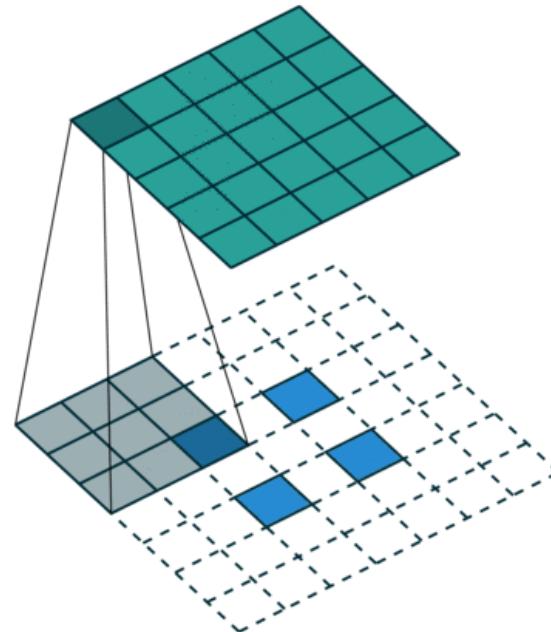
- Deep dimensionality reduction technique using a Convolutional Neural Network (unsupervised).



# What is convolution?



Convolution



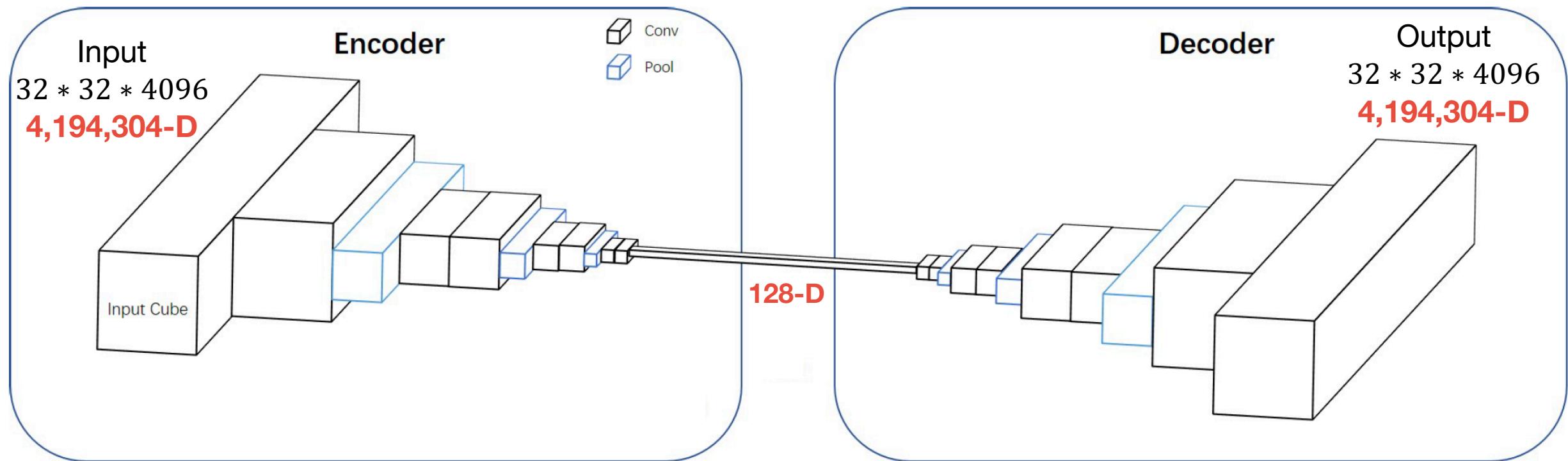
Transpose  
Convolution

Animations: [Dumoulin & Visin \(2016\)](#)

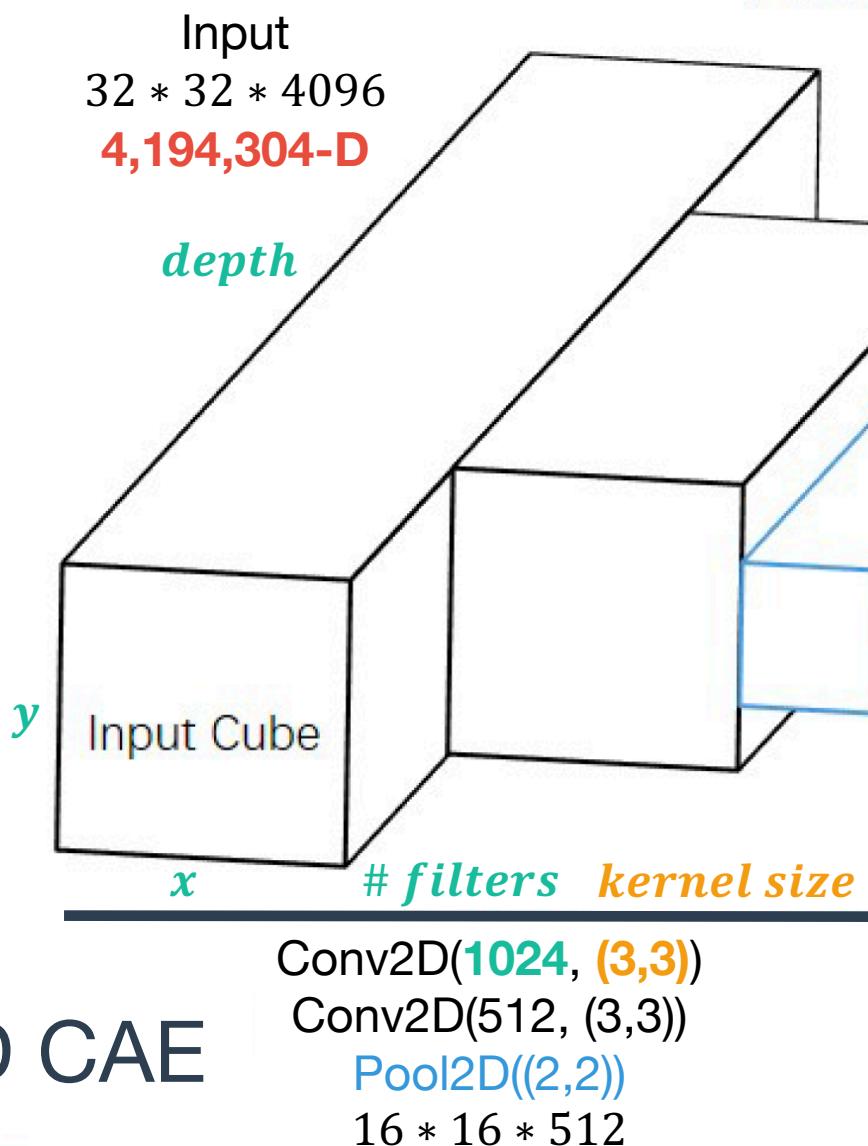
# CAE Architectures

- 2D CAE
  - standard for RGB images
- 3D CAE
  - kernel is 3x3x3 cube; take advantage of spectral correlations

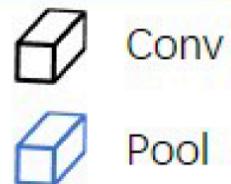
# Example Architecture (2D CAE)



## 2D CAE



## Encoder

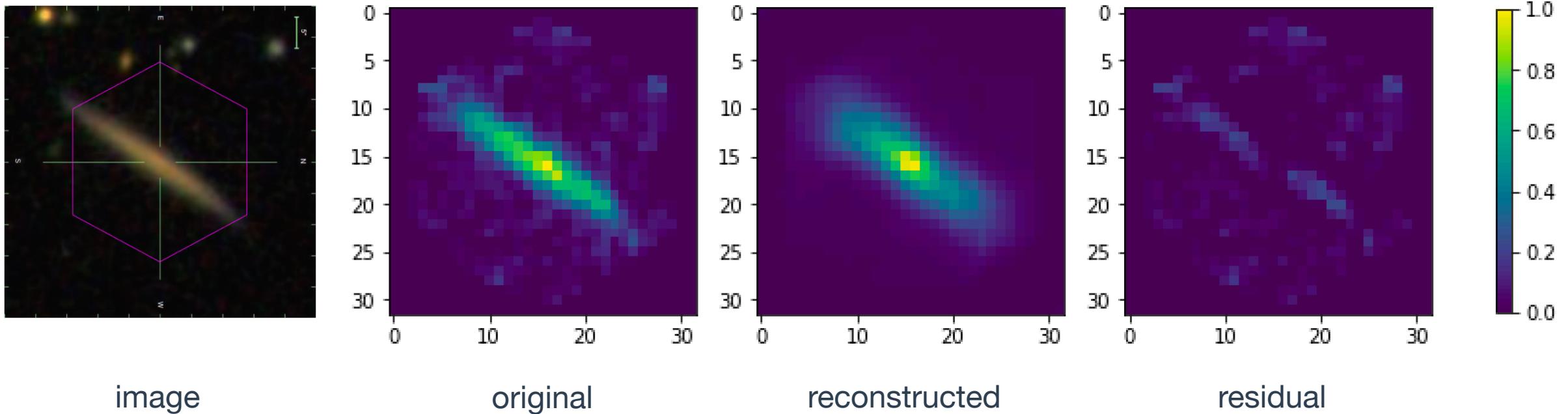


Conv2D(128, (3,3))  
Conv2D(128, (3,3))  
Pool2D((2,2))  
8 \* 8 \* 128

Conv2D(32, (3,3))  
Conv2D(32, (3,3))  
Pool2D((2,2))  
4 \* 4 \* 32

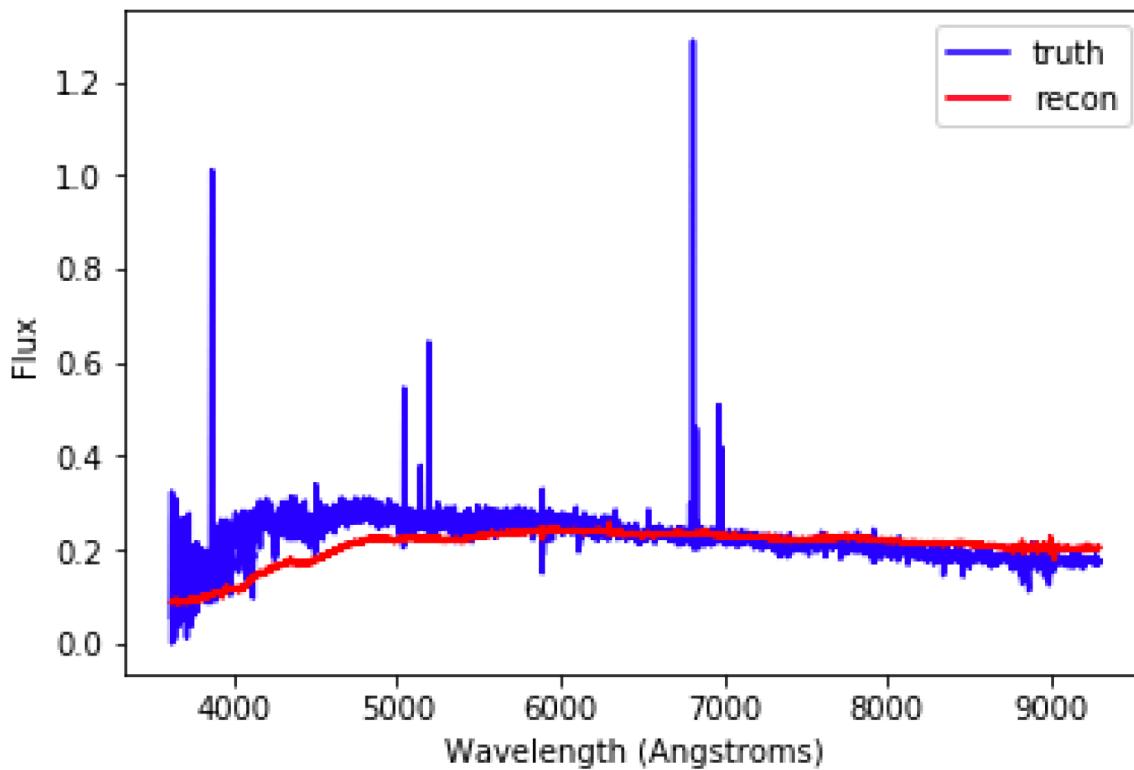
**128-D**  
Conv2D(8, (3,3))  
Conv2D(8, (3,3))  
4 \* 4 \* 8

# Reconstruction: Image Slice

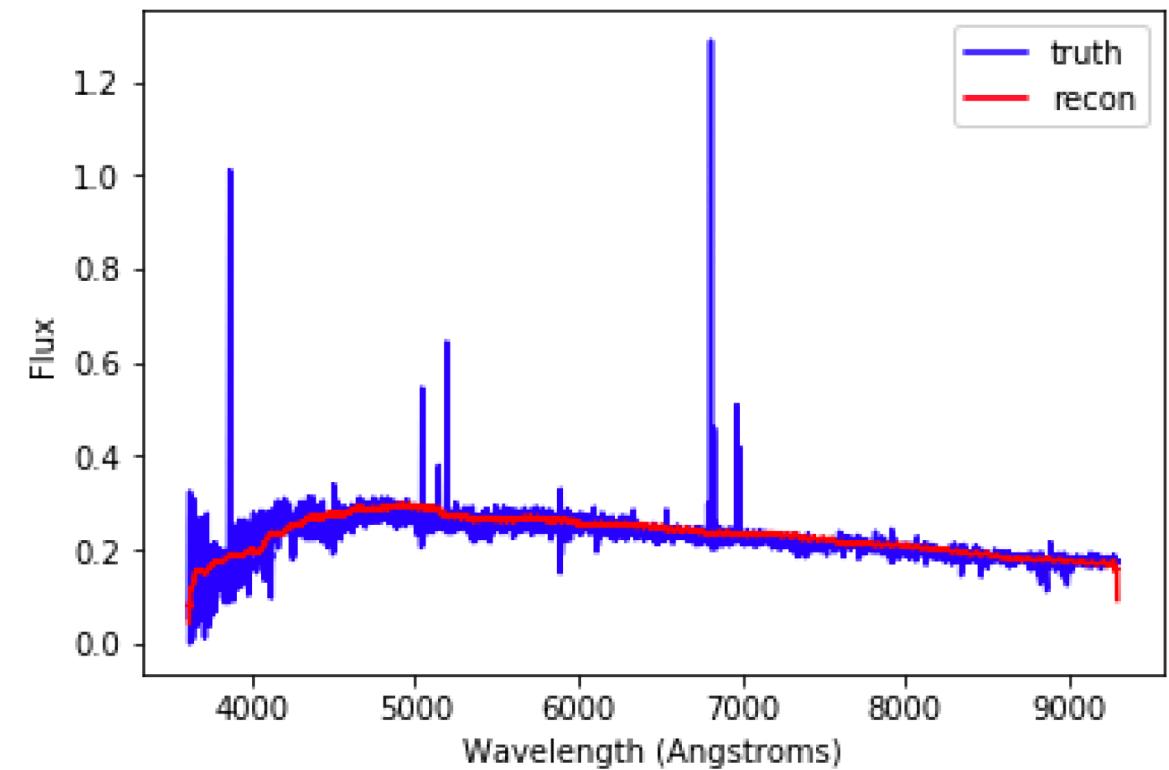


# Reconstruction: Spectral Slice

2D CAE

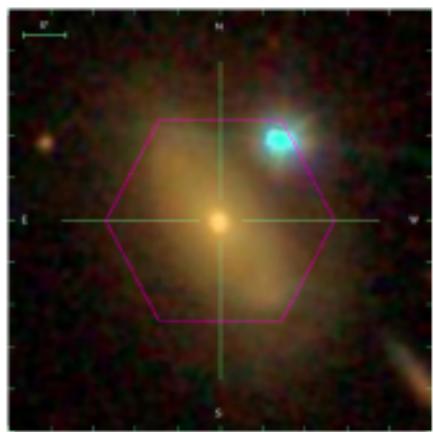


3D CAE

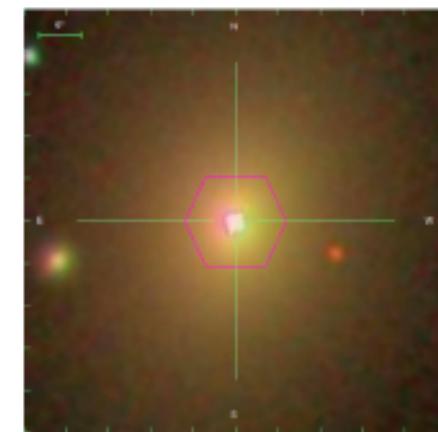
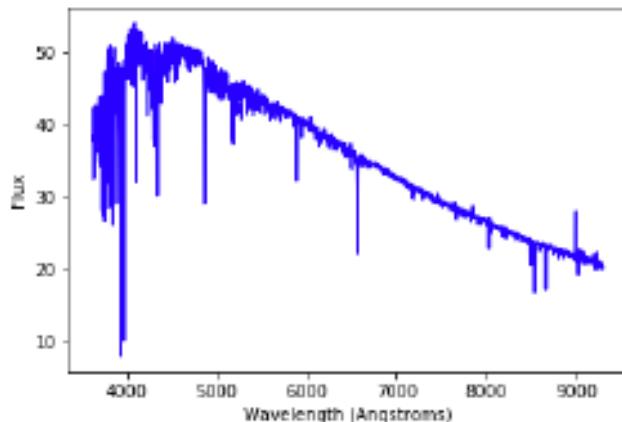


# Anomaly Detection on Low-D Encoding

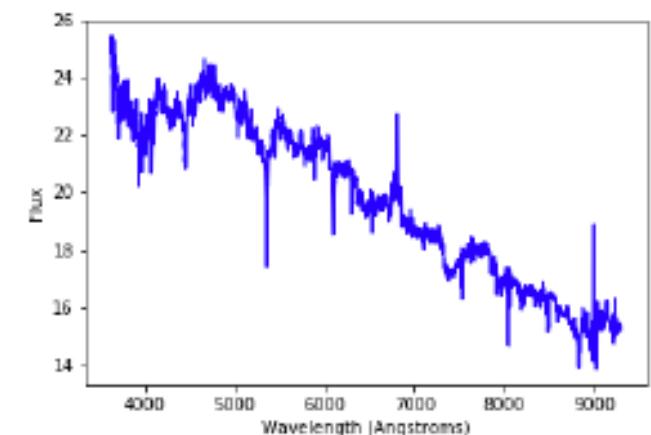
- Applied multiple anomaly detection methods, including isomap and k nearest neighbors.
- Could not recover our supernova example.
- But detected more extreme outliers:



foreground Milky Way star



blazar (beamed black hole accretion)



# Looking Ahead

- Supernova features are subtle.
- Need to improve reconstruction of sharp features.
  - incorporate signal-to-noise ratio
  - more galaxies (4x more will be available)
- Supernovae are only one science driver for anomaly detection.
- Other rare, interesting, and potentially unknown phenomenon could be discovered.