## **FLUKA: A Multi-Particle Transport Code**

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# Fluka:

## a multi-particle transport code

(Program version 2005)

Alfredo Ferrari, Paola R. Sala, Alberto Fassò, Johannes Ranft

#### Preface

The INFN-CERN Collaboration Agreement for the Maintenance and Development of the Fluka

### Acknowledgements

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The Fluka Collaboration:

Any portion of FLUKA so integrated, whether modified or unmodified shall continue to be subject to these license conditions.

7. Nothing in this chall be construed asant any rights any of th(c)-63sFLUKA versionsc

Johannes Ranft Johannes.Ranft@cern.ch Paola Sala Paola.Sala@mi.infn.it

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12 User routines 302

12.1 How to write, compile and link a user routine

## Part I

A summary description of FLUKA

## Chapter 1

Introduction

4 Introduction

#### 1.2.1 Physics

#### 1.2.1.1 Hadron inelastic nuclear interactions

The Fluka hadron-nucleon interaction models are based on resonance production and decay below a few GeV, and on the Dual Parton model above. Two models are used also in hadron-nucleus interactions. At momenta below 3–5 GeV/c the Peanut package includes a very detailed Generalised Intra-Nuclear Cascade (GINC) and a preequilibrium stage, while at high energies the Gribov-Glauber multiple collision mechanism is included in a less refined GINC. Both modules are followed by equilibrium processes: evaporation, fission, Fermi break-up, gamma deexcitation.

A quick look at Fl uka 5

#### 1.2.1.4 Transport of charged hadrons and muons

An original treatment of multiple Coulomb scattering and of ionisation fluctuations allows the code to handle accurately some challenging problems such as electron backscattering and energy deposition in thin layers even in the few keV energy range.

#### Energy loss:

- Bethe-Bloch theory [19-21, 29, 30]

\_

6 Introduction

#### 1.2.1.6 Electrons

#### 1.2.1.9 Neutrinos

\_

8 Introduction

#### Transport limits:

Secondary particles Primary particles

charged hadrons 1 keV-20 TeV ( ) 100 keV-20 Te81.1 ( ) neutrons the81.1 rm81.1al-2 Te81.1 ( ) Te81.1 rm81.20 TeV ( ) antineutrons 1 keV-20 TeV (

10 Introduction

transmutation, neutrino physics, shielding of free-electron lasers, calculation of tritium production at electron accelerators, energy amplifiers, maze design for medical accelerators, etc.

The recent addition of the simulation of heavy ion interactions allows also for applications to hadrotherapy.

## Chapter 2

## A FLUKA beginner's guide

2.1

12 Generalities

- 2.3 Building a FLUKA input
- 2.3.1 Generalities about FLUKA input

Beginner's guide

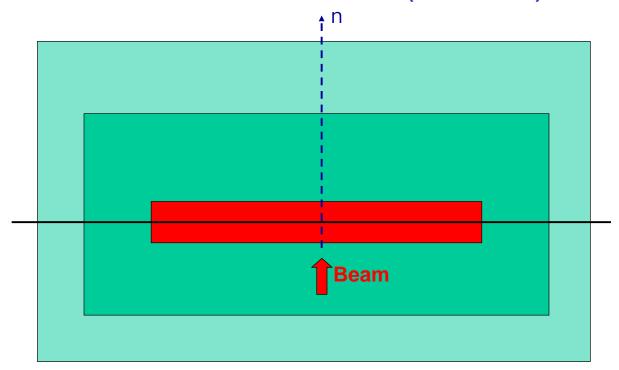
Beginner's guide 15

injection, the direction, etc., must be coherent with the geometrical description of the set-up, as discussed in the following section.

#### 2.6 The geometry

16 Materials

## **SKETCH OF THE GEOMETRY (not to scale!)**



Beginner's guide 17

18 Estimators and detectors

allowed, but is not necessary (except for region 2) because a region is blackhole by default unless another material has been associated to it. (Region 2, if not assigned a material, is COPPER by default).

The table entitled Regions: materials and fields, in the standard output, can be consulted to check that material assignment has been done as desired.

Beginner's guide

supplied by the user at the end of the geometry description (see 8.2.2). All other columns are normalised

giving in input the "volume" or "area" value multiplied or divided by those factors. Options USRTRACK, USRCOLL and USRBDX can also calculate energy fluence, if the "particle" type is set = 208.0 (energy) or 211.0 (electron and photon energy).

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independent runs (at least 4 or 5), each with a di erent independent initialisation, using the seeds written by the program at the end of each run. The rfl uka script provided with the code on UNIX and LINUX platforms takes care of this task, provided the following card is issued in the input file:

The seeds of the random number generator are printed on a special file in hexadecimal form at the end of each group of histories (the size of a group depends on the number of histories requested in the START card).

22 Sample input

GEOEND  *+1+2+3+4+5+6+7+8  MATERIAL 4.0 9.0122 1.848 5.0 BERYLLIU  *+1+2+3+4+5+6+7+8
* Be target, 1st and 2nd half
ASSI GNMAT 5.0 3.0 4.0
* External Black Hole
ASSI GNMAT 1.0 1.0
* Vacuum
ASSI GNMAT 2.0 2.0
*+1+2+3+4+5+6+7+8
* e+e- and gamma production threshold set at 10 MeV
EMFCUT -0.010 0.010 1.0 5.0 PROD-CUT
* score in each region energy deposition and stars produced by primaries
SCORE 208.0 210.0
* Boundary crossing fluence in the middle of the target (log intervals, one-way)
USRBDX 99.0 209.0 -47.0 3.0 4.0 400. pi Fl uenUD

24 Running Fl uka

At the end of each cycle the output files will be copied onto the running directory, the temporary

Beginner's quide 25

```
Removing links

Removing temporary files

Saving output and random number seed

Saving additional files generated

Moving fort. 47 to /home/fasso/Fluka/test/example005_fort. 47

Moving fort. 48 to /home/fasso/Fluka/test/example005_fort. 48

Moving fort. 49 to /home/fasso/Fluka/test/example005_fort. 49

Moving fort. 50 to /home/fasso/Fluka/test/example005_fort. 50

Moving fort. 51 to /home/fasso/Fluka/test/example005_fort. 51
```

End of FLUKA run

At this time, in the working directory, the following new O014les exist:

```
      exampl e001_fort. 47
      exampl e002_fort. 47
      exampl e005_fort. 47

      exampl e001_fort. 48
      exampl e002_fort. 48
      exampl e005_fort. 48

      exampl e001_fort. 49
      exampl e002_fort. 49
      exampl e005_fort. 49

      exampl e001_fort. 50
      exampl e002_fort. 50
      exampl e005_fort. 50

      exampl e001_fort. 51
      exampl e002_fort. 51
      exampl e005_fort. 51

      exampl e001. out
      exampl e002_out
      exampl e005_out

      exampl e005. err
      exampl e005. err
```

In Chapter 9 the user can O014nd a comprehensive description of the content of the content of the content of the purpose of this beginner's guide, it can just be pointed out that, according to the content of the USRBDX command, the O014les with extension t. 47

26 Results

For each estimator file the program will show the content of 37(the) Take card of 37(the) take input file,

Beginner's guide 27

```
Detector n: 1(1) piFluenUD

(Area: 400. cmq,
distr. scored: 209 ,
from reg. 3 to 4,
one way scoring,
fluence scoring)

Tot. resp. (Part/cmq/pr) 8.6904905E-04 +/- 0.6976866 %
(--> (Part/pr) 0.3476196 +/- 0.6976866 %)
```

28 Results

```
      1. 5672133E-04 +/-
      44.01294
      %
      2. 1093644E-04 +/-
      34. 72458
      %

      7. 4201569E-05 +/-
      33. 68359
      %
      7. 2452240E-05 +/-
      33. 54827
      %

      8. 6934262E-05 +/-
      62. 03180
      %
      1. 0245090E-04 +/-
      99. 00000
      %

      1. 6312006E-04 +/-
      82. 06016
      %
      1. 3002084E-04 +/-
      52. 15991
      %
```

Soon after, the cumulative fluence distribution as a function of energy is also given:

```
**** Cumulative Fluxes as a function of energy ****

**** (integrated over solid angle) ****
```

#### Energy boundaries (GeV):

49. 99992	40. 27077	32. 43475	26. 12349	21.04029
16. 94620	13.64875	10. 99293	8. 853892	7. 131072

Beginner's guide

\*\*\*\* Double diff. Fluxes as a function of energy \*\*\*\*

Solid angle minimum value (sr): 0.000000

Beginner's guide 31

#### 2.14.3 Binning estimator

To analyse the binary output from USRBIN

2.16 **Biasing**Althhagh able to perform fully analogue particle transport calculations (i.e., to reproduce faithfully actual particle histories), in many cases of very non-uniform radiation fields, such as those encountered in shielding design, only a very small fraction of all the histories contributes to the desired response (dose, fluence) in the regions of interest, for instaalst,

# Chapter 3

### Installation

The Fluka

34 Installation

If the source code is present, the INCLUDE files needed to compile the program may be grouped into three files emfadd. add, fl ukaadd. add and I owneuadd. add.

A Makefile and a number of auxiliary programs split these files into individual routines and INCLUDE files, which are placed in 30+1 separate directories and compiled. The object files are inserted in a FLUKA library I i bfl ukahp. a. A shell script I fl uka links all routines into an executable fl ukahp (the name is the leswhide29(i)14amwhir,a exet

Preconnected files 35

## Chapter 4

## FLUKA modules (Fortran files)

Since several years, the Fluka

Fluka modules 37

User oriented routines (see description in Chap. 12):

The "FLUKA User Routines" mentioned at point 3) in the FLUKA User License are those (and only those) contained in the directory usermvax, both in the source and binary versions of the code.

ABSCFF : absorption coe cient (for optical photons)

COMSCW

Particle codes 43

Table 5.2: Fluka generalised particles (to be used only for scoring)

Fluka	Fluka	Description
name	number	
	40	Low-energy neutrons (used only in some input options)
ALL-PART	201	All transportable particles
ALL-CHAR	202	All charged particles
ALL-NEUT	203	All neutral particles
ALL-NEGA	204	All negative particles
ALL-POSI	205	All positive particles
NUCLEONS	206	Protons and neutrons
NUC&PI+-	207	

44 Pre-defined materials

#### 5.2 Pre-defined materials

Materials can be easily defined by option MATERIAL (p. 142

### Chapter 6

### General features of FLUKA input

The input of Fluka consists of a text file containing a sequence of option lines (often called "cards") which

48 Fluka input

-100.0 0.0 -21200.0 100.0 0.0 -20800.0 STOP

#endif

### Chapter 7

### Description 9f FLUKA input options

There are more than 70 option keywords available for input in Fluka. A summary is given in Section 7.1,

50 Input Commands

EVENTBIN

Introduction 51

ROT-DEFIni defines rotations/translations to be applied to user-defined binnings

ROTPRBIN sets the storage precision (single or double) and assigns possible rotations/translations

for a given user-defined binning (USRBIN or EVENTBIN)

RQMD defines some I/O parameters relevant to the heavy ion event generator Rqmd

SCORE defines the (generalised) particles to be scored by region

SOURCE tells Fluka to call a user-written source routine

START defines the number of primary particles to follow, gets a primary particle from a beam

or from a source, starts the transport and repeats until the predetermined number of

primaries is reached

STEPSIZE sets the maximum step size in cm (by region) for transport of charged particles

STERNHEIme allows users to input their own values of the density e ect parameters

STOP stops input reading

TCQUENCH sets scoring time cut-o s and/or Birks quenching parameters

THRESHOLd defines the energy threshold for star density scoring, and sets thresholds for elastic and

inelastic hadron reactions

TIME-CUT sets transport time cut-o s
TITLE gives the title of the run

USERDUMP requests a collision file and defines the events to be written

52 Input Commands

For what concerns item 4, the user has a choice of several options to request the estimation of various radiometric quantities. Of course, there is no much point in running the program without requesting any result, but in a phase of input preparation it is quite common to have a few runs without any scoring commands. A typical minimum input containing only specifications for the above items 1, 2 and 3 will still produce some useful information. Looking at the standard Fluka output, the user can do several consistency checks, and can get some better insight into the problem from the final statistics and energy

Introduction 53

and needs to read its own input, as explained in 7.56.

## 7.1.3.2 General setting options

The di cult task of choosing the best settings for a calculation problem is made much easier by existence of

Introduction 55

In early Monte Carlo programs results could depend critically on the size of the step, mainly due to

#### 1.1.3.6 Time cut-o s

For time-dependent calculations, two time cut-o options are available: one for particle transport, TIME-CUT,

Introduction 57

particle type input via the BEAM command can only be a generic heavy ion.

counter of crossings, not weighted by inverse cosine: but despite a widespread credence, current is only

Introduction 59

## 7.1.6.1 Simple biasing options

Introduction 61

#### 7.1.8 Miscellaneous

Option COMMENT (7.8) is not often used. Its main function is to insert acertain number of comment lines

62 ASSIGNMAt

#### 7.2 ASSIGNMAt

#### Example:

*+1.	+ 2	. + 3	. + 4	. + 5	. + 6	. + 7 +
MATERI AL	13.0	27.0	2.7	10.0	0.0	O.O ALUMINUM
ASSI GNMA	10.0	1.0	15.0	0.0	1.0	0.0
ASSI GNMA	2.0	5.0	17.0	6.0	-1.0	0.0
ASSI GNMA	2.0	16.0	18.0	2.0	0.0	0.0

- The above definitions mean that all regions from 1 to 15 are aluminium with a magnetic field, except regions 5 and 11 which are vacuum without any magnetic field. Regions 16, 17 and 18 are also
- vacuum without field.
- Note that in the above example material 10 has been defined
- overriding the pre-defined FLUKA aluminium material, but keeping

66 BEAM

# Example:

*	The following B	EAM card r	efers to	a 100 keV	pencil-lik	e
*	electron beam:					
*+	1 + 2 +	3 +	4	+5	+6	+7+
BEAM	-1. E-4	0.0	0.0	0.0	0.0	1.0 ELECTRON
*	The next option	card desc	ribes a p	arallel pr	oton beam	with a
*	momentum of 10.	0 +/- 0.2	GeV/c, wi	th a gauss	ian profil	e in
*	the x-direction	and in th	ie y-direc	tion descr	ibed by st	andard
*	deviations sigm	$a_x = 1.$ c	m (FWHM =	2.36 cm)	and sigma_	y = 0.5
*	cm (FWHM = 1.18)	cm).				
* +	1 + 2 +	3 +	4	+5	+ 6	+7+
BEAM	10.0	0. 2	0.0	-2. 36	-1.18	1.0 PROTON
*	The next exampl	e concerns	a negati	ve muon be	am of 2 Ge	·V
*	kinetic energy,	with a di	vergence	of 3 mrad.		
* +	1 + 2 +	3 +	4	+5	+ 6	+7+
BEAM	-2.0	0.0	3.0	0.0	0.0	1.0 MUON-
*	The next BEAM c	ard descri	bes a 137	-Cs isotro	pic source	!
BEAM	-661. 7E-6	0.0	1. E4	0.0	0.0	1.0 PHOTON

7.4	BE.	Δ٨	Λ	X	F٩
1.4	$\mathbf{D}$	$\neg$	/   一	$\sim$	டல

68 BEAMAXES

7.5	B	F.	Α	M	lΡ	$\mathbf{O}$	S
1.5	$\mathbf{-}$	_	$\overline{}$	ıν		${}$	•

70 BEAMPOS

3. Beam divergence and transversal profile defined by option BEAM (p. 7.3), as well as polarisation defined by option POLARIZAti (p. 7.57), are meaningful only if the beam direction is along the positive z-axis, unless a command BEAMAXES is issued to establish a beam reference frame di erent from the geometry frame (see p. 7.4).

#### Examples:

```
* A beam parallel to the x-axis starting at a point of

* coordinates -0.1, 5.0, 5.0:

*...+...1....+...2...+...3...+...4...+...5...+...6...+...7...

BEAMPOS -0.1 5.0 5.0 1.0 0.0 0.0

* A beam perpendicular to the x-axis, with director cosines

* 0., 1/sqrt(2), -1/sqrt(2) with respect to x, y and z,

* starting at point 0.0, 0.0, 0.0:

*...+...1...+...2...+...3...+...4...+...5...+...6...+...7...+...

BEAMPOS 0.0 0.0 0.0 0.0 0.7071068 0.0 NEGATIVE
```

## 7.6 BIASING

72 BIASING

74 BIASING

## Example:

*+1	. + 2	+3	. + 4	. + 5	+6	. + 7 +
BIASING	2.0	0.0	10.0	7.0	11.0	2.0
BIASING	2.0	0.0	15.0	8.0	9.0	0.0
BIASING	-1.0	0.0	3.0	4.0	0.0	0.0
BIASING	1.0	0.7	0.4	3. 0	8.0	O. OPRINT

76 COMMENT

## Examples:

\*...+...1...+...2...+...3...+...4...+...5...+...6...+...7...+... COMMENT 3.0

78 COMPOUND

#### 7.9 COMPOUND

Defines a compound or mixture, made of several materials, or even a mixture of di erent isotopes

See also LOW-MAT, MATERIA	L, MAT-PROP

neutron data set is identified either by name (if equal to a Fluka

## 7.11 DCYSCORE

Associates selected scoring detectors of given estimator type with user-

82 DCYTIMES

#### 7.12 DCYTIMES

#### 7.13 DEFAULTS

Sets Fluka defaults suitable for a specified kind of problems. Starting from Fluka99.5 (June 2000) the standard defaults are those described under NEW-DEFAults below. That is, if no DEFAULTS card is issued the code behaves as if a card with NEW-DEFAults was given. be given.

See also GLOBAL

86 DEFAULTS

4.

90 DELTARAY

Notes

7 4 5		r	-	СТ
7.15	U	ᄆ	ᆮ	$\cup$ $\square$

92 DETECT

Note: if no trigger region is given (i.e., no region with negative sign), a simple event-by-, a sigNote7.2-11.9552870tak

94 DISCARD

# 7.16 DISCARD

# 7.17 DPMJET

Defines some I/O parameters relevant to the heavy ion event generator Dpmjet.

See also BME, EVENTYPE, MYRQMD, PHYSICS, RQMD

Option DPMJET

96 ELCFIELD

# 7.18 ELCFIELD

7.19	EMF				

98 EMF-BIAS

7.20	EMF-	-BIAS				

WHAT(5) = upper bound of the region indices where the selected leading particle biasing has to be played

102 EMFCUT

## 7.21 EMFCUT

Sets the energy thresholds for electron and photon production in di erent materials, and electron and photon transport cut-o s in selected regions.

Input Commands	103

106 EMFCUT

108 **EMFFIX** 

# Example:

*+1.	+ 2	+ 3.	+ 4 .	+ 5.	+ 6	+	7
MATERIAL	13.	26. 98	2.6989	3.	0.	0.	ALUMI NUM
MATERIAL	82.	207.20	11. 35	4.	0.	0.	LEAD
MATERIAL	29.	63.546	8. 96	12.	0.	0.	COPPER
MATERIAL	6.	12.000	2.00	26.	0.	0.	CARBON
MATERI AL	7.	14.000	0.0012	27.	0.	0.	NI TROGEN
MATERIAL	8.	16.000	0.0014	28.	0.	0.	OXYGEN
MATERIAL	1.	1.000	0.0001	29.	0.	1.	HYDROGEN
MATERIAL	0.	0.0	1.0000	30.	0.	0.	TISSUE
COMPOUND	5.57E-3	26.0	1. 118E-3	27.	2.868E-2	28.	TISSUE
COMPOUND	6. 082E-2	29.0	0.	0.	0.	0.	TISSUE
EMFFIX	3.	0. 15	4.	0. 15	12.	0.15	
EMFFIX	30.	0.05	0.	0.	0.	0.	PRINT

In this example, a maximum energy loss per step of 15% is requested for aluminium, copper and lead, while a more accurate 5% is requested

for tissue

# 7.23 EMFFLUO

Activates a detailed treatment of photoelectric interactions and of the fol-

110 EMFRAY

# 7.24 EMFRAY

Activates Rayleigh (coherent) scattering and Compton binding corrections

WHAT(1) 1.0:

7.25	EVEN	NTBIN				

112 **EVENTBIN** 

#### Example 1:

\*..+...1...+...2...+...3...+...4...+...5...+...6...+...7...+.... EVENTBIN 10.0208.0 25.0150.0200.0180. Firstscore EVENTBIN -150.0100.0-20.0 75.0 50.020.0 &

- \* In the above example, the user requests an event-by-event scoring of
- energy deposition (generalised75.566TfarticIthe 25(Ie,)-52iionaf

\*thests =heaf

- \* -1ts don-15(le,)-522he -1tsan01ts dof

# 7.26 EVENTDAT

For cal	lorimetry	onl	V.

116 EVENTYPE

7.27

118 **EXPTRANS** 

Example:

\*...+...1...+...2...+...3...+...4...+...5...+...6...+...7...+.... EXPTRANS 1. 0.8 10. 18. 8. 0. EXPTRANS -1. 7. 8. 0. 0. 0. 0.  $^{\circ}$  Exponential transformation is requested for photons ,050particle no. 7,051

## 7.29 FLUKAFIX

Sets the size of the step of muons and charged hadrons to a fixed fraction of the kinetic energy in di erent materials

See also EMFFIX, MULSOPT, STEPSIZE

WHAT(1) = fraction of the kinetic energy to be lost in a step (must not be > 0.2)

Default : if option DEFAg(thc)B(TS)JJ/F8.963T4.938Td(is)-334(u)1(se)-1(d)-333(ith)JJ/F508966T55.4590Td(SE

120 FREE

# 7.30 FREE

Activates free-format input

See also GLOBAL

# 7.31 GEOBEGIN

122 GEOBEGIN

and debugging messages. Minor tracking problems, however, are reported on the error message file (logical unit 15, see Chap. 3 and Sec. 9.4), unless reporting has been de-activated by setting

0 = 0 =	7.32	GEOEND
	7.32	GEOEND

124 GEOEND

4. It must be stressed too that the geometry debugger can be very time consuming, so don't ask for 100  $\mu m$ 

126 GLOBAL

WHAT(5): flag to request free format in the geometry input for bodies and regions. This format is described in 8.2.3.2 and 8.2.6.3, and requires the use of names (alphanumerical 8-character

128 IONFLUCT

### 7.35 IONFLUCT

Calculates ionisation energy losses of charged hadrons, muons, and electrons/positrons with ionisation fluctuations.

See also DELTARAY

### WHAT(1)

- 1.0: switches on restricted energy loss fluctuations for hadrons and muons
- -1. 0: switches o restricted energy loss fluctuations for hadrons and muons
- = 0.0: ignored

**Default**: restricted energy loss fluctuations for hadrons and muons are activated if option DEFAULTS is missing or if it is used with

# <u>Note</u>

1. The energy loss fluctuation algorithm is fully compatible with the DELTARAY option (p. 89

7.37	LAM-	-BIAS				

|**WHAT(2)**| 1.0:

134 LAM-BIAS

5. The biasing function for the decay direction is of the form  $e^{-\frac{1-cos(\cdot)}{2}}$  where  $\cdot$  is the polar angle between

	$\sim$	Λ.	, ,	<b>`</b>	^	
7.38	1 11	./\	/⊢	2 I	/\	•
/ .oo	 $\mathbf{C}$	v	/_E	<i>.</i>	$\boldsymbol{-}$	

136 LOW-BIAS

# <u>Notes</u>

1. The groups are numbered in

7 39	LOW	$I$ _ $D$	$\cap$	/NI
7.39	1 ()//	ノーレン	しりひい	/ I VI

138 LOW-MAT

# 7.40 LOW-MAT

140 LOW-NEUT

# 7.41 LOW-NEUT

#### **Notes**

1. In Fluka, transport of neutrons with energies lower than a certain threshold is performed by a multigroup algorithm. For the neutron cross-section library currently used by Fluka, this threshold is 0.0196 GeV. The multi-group transport algorithm is described in Chap. 10.

- 2. Evaporation option (see EVENTYPE, p. 116) is mandatory if LOW-NEUT is requested. If low-energy neutrons are not transported (because of the chosen DEFAULTS, or because a DEFAULTS card is absent), the energy of neutrons beler threshold (default or set by PART-THR, p. 172) is deposited on the spot. This is true also for evaporation neutrons. If there is no interest in transporting low-energy neutrons, but this feature is not desired, it is suggested trequest LOW-NEUT, and to use LOW-BIAS (p. 135) with a group cut-o WHAT(1) = 1.0.
- 3. Gamma data are used only for capture gamma generation and not for transport (transport is done via the 85cm0g0G1006.9620Td[lect)1(ro)]TF628.966Tf22.5510Td[M)]TF498.966Tf10.0570Td[agneti)1(c)]TF628.966Tf

142 MATERIAL

7 42	Λ	1	Δ٦	ΓF	D	IA	ı
14/	ı١	<i>,</i>	$\boldsymbol{-}$		$\mathbf{r}$	_	

144 MAT-PROP

7.43

### For SDUM = LOWNTEMP:

WHAT(1) > 0.0: Temperature ratio  $(T_{actual}/T_{xsec})$  with respect to the nominal one. The nominal temperature  $T_{xsec}$  (given by WHAT(3), see below) is the temperature for which the neutron cross-sections of the Fluka material(s) concerned have been prepared (see Tag(r) $\dagger\theta$ .3, p. 295). See Notes 4, 5

146 MAT-PROP

Notes	
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150 MCSTHRES

7.45

152 MGNFIELD

# Example:

\*...+....1....+....2....+....3....+....4....+....5....+....6....+....7....+...

### 7.46 MULSOPT

Sets the tracking conditions for multiple Coulomb scattering (MCS), for both hadrons/muons and  $e^+e^-$ . Can also be used to activate single scattering.

See also EMFFIX, FLUKAFIX, MCSTHRES, STEPSIZE

For SDUM = GLOBAL, GLOBEMF, GLOBHAD:

WHAT(1): controls the step optimisation for multiple Coulomb scattering (see Note 1) and the number

atomic electrons is switched o

: Fano correction for inelastic interactions of charged hadrons and muons on atomic electrons [48] is switched on

= 154 MULSOPT

WHAT(6)<sup>N</sup> = step length in assigning indices. ("...in steps of WHAT(6)") Default = 1.0

SDUM =

< 0.0: single scattering is not activated

= 0.0: ignored

**Default**: single scattering is not activated

WHAT(6) (meaningful only if single scattering is activated at boundaries and when step is too short: see WHAT(4) above)

> 0.0: number of single scatterings to be performed when crossing a boundary = 0.0: ignored < 0.0: resets the default

Default = 1.0

### <u>Notes</u>

1.

156 MUPHOTON

# 7.47 MUPHOTON

# 7.48 MYRQMD

Not yet implemented. Prepared for new QMD generator

158 OPEN

### 7.49 OPEN

Defines input/output files to be connected at run-time.

WHAT(1) > 0.0: logical unit number of a *formatted* file to be opened < 0.0: logical unit number of an

160 OPT-PROD

### 7.50 OPT-PROD

Requests and controls production of Cherenkov, Transition and Scintillation Radiation in specified materials.

See also OPT-PROP, Chap. 13, and axamples in SDUM = CERE-OFF: switches o Cherenkov production

For

WHAT(1) - WHAT(3): not used

WHAT(4) - WHAT(6): assignment to materials, see below

SDUM = CERE-OFF

SDUM = TRD-OFF

For

162 OPT-PROD

#### Notes

1. Optical photons such as those produced by Cherenkov e ect are distinguished b b Fluka name (OPTIPHOT) and by t6eir Fluka id-nber ( -1), as sho in 5.1.

2. To transport optical photons, it is necessary to define the optical properties of t28h6e relevan materials b of option OPT-PROP (p. 164). Users can also write t6eir own routines USRMED (p. 323), which is called at ever-64382(step64381(and)-382(at)-380(b)-29(oundar-64382(c)1(ros)-1(sings64381(whe)-1(n)-381(act)1(iv)57(ated64381(w)-382(at)-380(b)-382(at)-382(a

164 OPT-PROP

# 7.51 OPTgPROP

SDUM = RESET

For SDUM = METAL: flag the material as metal

WHAT(1) =  $1^{st}$  optical property (not used at the moment)

WHAT(2)  $0.02^{nd}$  optical property (not used at the moment)

2

Default =

166 OPT-PROP

### For all previous SDUMs:

WHAT(4) = lower bound of the indices of materials to which the indicated optical properties refer ("From material WHAT(4)...")

Default = 3.0

WHAT(5) = upper bound of the indices of materials to which the indicated optical properties refer ("...to material WHAT(5)...")

Default = WHAT(4)

WHAT(6) = step length in assigning indices

("...in steps of WHAT(6)")

3.0

#### Example 1:

\* Optical photon transport requested between 3.E15 and 7.E15 rad/s

\* (4.77E5 and 1.11E6 GHz, or 314 to 628 nm) for materials 6,9,12,15 and 18

\*...+....1....+...2...+...3...+...4...+...5...+...6...+...7...+...

OPT-PROP 3.E15 6.E15 7.E15 6.0 18.0 3. OM-LIMIT

\* User routine USRMED called when an optical photon is going to be transported

\* in materials 6, 12 and 18

MAT-PROP 1.0 0.0 0.0 6.0 18.0 6. USERDIRE

#### Example 2:

\* Material 11 has a reflectivity index = 0.32

\*...+....1....+....2....+....3....+....4....+....5....+....6....+....7....+.... OPT-PROP 0.0 0.0 0.32 11.0 0.0 0. METAL

#### Example 3:

 $^{\star}...+...1...+...2...+...3...+...4...+...5...+...6...+...7...+...$ 

 $<sup>^{\</sup>star}$  Optical photon transport requested between 300 and 600 nm for water.

<sup>\* (</sup>material 9). The optical properties are for the Na D line (589 nm)

170 PAIRBREM

# 7.52 PAIRBREM

Controls simulation of pair production and bremsstrahlung by high-energy

### <u>Notes</u>

1. Initialisation of bremsstrahlung and pair production by heavy charged particles is very demanding in com-

172 PART-THRes

# 7.53 PART-THRes

Sets di erent energy transport cut-o s for hadrons, muons and neutrinos.

See also EMFCUT, LOW-BIAS, THRESHOLd

The meaning of WHAT(1) depends also on the value of

#### **Notes**

1. When applied to neutrons, the cut-o energy defined by PART-THRes refers to the energy boundary between high-energy and low-energy neutrons, i.e., the upper limit of the first energy group in the multigroup transport

174 PHOTONUC

### 7.54 PHOTONUC

Activates gamma interactions with nuclei.

# Example 1:

- \*Gi antResonanceandQuasi deuteronphotonucl earinteracti onsarerequested \*inmateri al 18. Thephotonhadroni ci nteracti onl engthi sarti fi ci al l y
- \*shortenedbyafactor0.02inordertoimprovestatistics

176 PHYSICS

### 7.55 PHYSICS

Allows to override the standard Fluka

178 PHYSICS

For SDUM = EVAPORATion

7.56	DΙ	$\cap$	T	$\subseteq$	F	$\cap$	N	1
7.30	Гι	_~		J	ഥ`	$\smile$	ıv	ı

180 PLOTGEOM

than one of the points P2 on the current line satisfies the quoted condition for a given P1, then only the nearest one to that P1 is taken.

182 POLARIZAti

7.57 P	OLA	RIZAti					
WHAT(	1)	1.0:					

2. The program takes care of properly normalising the cosines unless they are badly unnormalised (in the latter case the code would reset to no polarisation). If WHAT(4) 1.0, the code makes sure that the two vectors are orthogonal within the minimum possible rounding errors.

3.

184 RADDECAY

### 7.58 RADDECAY

requests simulation of radioactive decays and sets the corresponding biasing and transport conditions

See also DCYTIMES, DCYSCORE, IRRPROFIle, RESNUCLEI

WHAT(1):

186 RANDOMIZE

# 7.59 RANDOMIZe

# 7.60 RESNUw0]187

188 RESNUCLEIRESNUCLEI

190 ROT-DEFIni

# 7.61 ROT-DEFIni

Defines rotations and translations to be applied to binnings.

See also EVENTBIN, ROTPRBIN, USRBIN

**WHAT(1)** : assigns a transformation index and the corresponding rotation axis 0.0

$$j = 3$$
:
$$\begin{array}{cccc} \vdots & X_{\text{new}} & \vdots & \vdots & \vdots \\ \vdots & Y_{\text{new}} & \vdots & \vdots & \vdots \\ \vdots & Z_{\text{new}} & \vdots & \vdots & \vdots \\ \end{array}$$

192 ROTPRBIN

7.62	ROT	PRBIN			

194 RQMD

7.64

196 SCORE

- 4. The SCORE card defines the following scoring:
  - (a) scoring by region the density of stars produced by the selected particles (if applicable, i.e., if the particles are hadrons but not low-energy neutrons photons or muons, or families of them). Stars produced by primary particles can be scored with id-number 210.0, all stars with 201.0.

Results will be in stars/cm $^3$  p-energy particle if region volumes have been input by setting T=3

### 7.65 SOURCE

Invokes the use of a user-defined source routine SOURCE to sample the primary particles.

See also BEAM, BEAMPOS, POLARIZAti, USRICALL us1knsd-3278koesened]TJ/F189.963Tf2161773ITd[(SOURCE)]TJ/F49.963Tf21.8380Td[(B.40278kT)83()40278kT)

198 SOURCE

5. In old versions of Fluka, the call to SOURCE was requested by means of a flag in card START

## 7.66 START

Defines the termination conditions, gets a primary from a beam or from a source and starts the transport.

See also

200 STEPSIZE

# 7.67 STEPSIZE

# Example:

*+1	. + 2	+ 3	. + 4	. + 5	+ 6	. + 7 +	⊦
ASSI GNMAT	2.0	15.0	30.0	5.0	1.0	0.0	
* A magnetic	$\   \text{field is} \\$	present in	vacuum r1g	ions 15,	20, 25 and	30.	
MGNFI ELD	20.0	0. 2	0. 10	0.0	0.0	0.0	
STEPSI ZE	-0.05	0.0	20.0	25.0	0.0	0.0	
STEPSI ZE	0.3	0.0	15.0	0.0	0.0	0.0	

202 STERNHEIme

# 7.68 STERNHEIme

Allows to input Sternheimer density e ect parameters

## 7.69 STOP

Stops the execution of the program

See also START

WHAT(1) - WHAT(6) and SDUM: not used

**Default** (option STOP not given): no e ect (the program stops at the end of the run when the conditions set in the START command (p. 199

204 TCQUENCH

## 7.70 TCQUENCH

Sets time cut-o s and/or quenching factors when scoring using the USRBIN or the EVENTBIN options.

See also TIME-CUT

WHAT(2) > 0.0: first Birks law coe cient in g/(MeV)

# 7.71 THRESHOLd

206 TIME-CUT

# 7.72 TIME-CUT

# 7.73 TITLE

210 USERWEIG

#### Notes

1. These weights are really extra, i.e., the results are multiplied by these weights at scoring time, but printed titles, headings and normalisations are not necessarily valid. It is the user's responsibility to interpret correctly the output. Actually, it is recommended to insert into standard output a user-written notice informing about the extra weighting

2. Setting the incident particle weight to a value di erent from

# 7.76 USRBDX

6. When scoring neutron fluence or current, and the requested energy bin structure overlaps with that of the low-energy neutron groups, bin boundaries are forced to coincide with group boundaries and no bin can be smaller than the corresponding group.

214 USRBDX

```
PARAMETER ( MXSCOR = MXENER*MXANGL )
PARAMETER ( NMXGRP = 100 )
                               ! # of low-energy neutron groups
 LOGICAL LFUSBX, LWUSBX, LLNUSX
CHARACTER RUNTIT*80, RUNTIM*32, TITUSX*10, FILNAM*80
DIMENSION EBXLOW(MXUSBX), EBXHGH(MXUSBX), ABXLOW(MXUSBX),
           ABXHGH(MXUSBX), DEBXBN(MXUSBX), DABXBN(MXUSBX),
           AUSBDX(MXUSBX), NEBXBN(MXUSBX), NABXBN(MXUSBX),
&
           NR1USX(MXUSBX), NR2USX(MXUSBX), ITUSBX(MXUSBX),
&
           IDUSBX(MXUSBX), IGMUSX(MXUSBX), LFUSBX(MXUSBX),
&
           LWUSBX(MXUSBX), LLNUSX(MXUSBX), TITUSX(MXUSBX),
&
&
           ENGMAX (NMXGRP+1), SCORED(MXSCOR), ELIMIT(MXUSBX),
           MX (MXUSBX), NHI GH (MXUSBX)
DOUBLE PRECISION CUMUL, ANGINT, EN1, EN2, ELIMIT, DIFF
\label{eq:write} \textit{WRITE(*,*)'} \  \  \, \textit{Type the name of the input file:'}
READ (*, '(A)') FILNAM
 LQ = INDEX(FILNAM, ' ') - 1
 OPEN (UNIT=1, FILE=FILNAM, STATUS='OLD', FORM='UNFORMATTED')
 OPEN (UNIT=2, FILE=FILNAM(1:LQ)//'.txt', STATUS='NEW')
```

END IF

= 8.0:

220 USRBIN

WHAT(5) =

222 USRBIN

7. When scoring energy deposition (generalised particles 208.0 and 211.0), it is recommended to set in the first USRBIN card WHAT(1) = 10.0, 11.0, ... 17.0 (rather than

226 USRBIN

# 7.78 USRCOLL

Defines a detector for a collision fluence estimator

See also USRBDX, USRBIN, USRTRACK

The full definition of the detector may require two successive cards. The second card, identified by the character " &

# 7.79 USRICALL

230 USROCALL

# 7.80 USROCALL

Calls user-dependent output.

See also

# 7.81 USRTRACK

Defines a detector for a track-length fluence estimator.

234 USRTRACK

236 USRYIELD

### 7.82 USRYIELD

238 USRYIELD

#### For SDUM = BEAMDEF:

WHAT(1) = (projectile particle index)

**Default** = IJBEAM (beam particle)

WHAT(2) = target particle index (used by the code to define the c.m.s. frame)

**Default** = 1.0 (proton)

WHAT(3) = projectile momentum

Default = PBEAM (beam momentum)

WHAT(4,5,6) = projectile direction cosines

**Default** = UBEAM, VBEAM, WBEAM (beam direction cosines)

**Default** (option USRYIELD

240 USRYIELD

```
* Lorentz transformations, Feynman X etc.

* RUNTIM = date and time of the run (as printed on standard output)

* RUNTIT = title of the run (as given by card TITLE)

* SCORED = result array

* SGUSYL = aUopted cross section (if any)
```

242 USRYIELD

## Example:

*+1.	+	. + 3	. + 4	+5	. + 6	+8
USRYI ELD	1399. 0	13.	21.0	3.0	2. 0	1. 0TotPi +(E)
USRYI ELD	50.0	0.001	100. 03. 14	1159265	0. 0	3.0 &

<sup>\*</sup> Score double differential yield of positive pions going from region 3 to \* region 2 with a polar angle between 0 and pi with respect to the beam

Default (if no WW-PROFIle

246 WW-FACTOr

5. It can also be said that WW-FACTOr and BIASING

248 WW-PROFIle

(option

7.85	WW-	-THRESh					

# Chapter 8

# Combinatorial Geometry

#### 8.1 Introduction

The Combinatorial Geometry (CG) used by FI uka is a modification of the package developed at ORNL for the neutron and gamma-ray transport program Morse [44] which was based on the original combinatorial

Body Data 253

the GLOBAL command, p. 125). The remaining 60 characters can be used for any alphanumeric string at the user's choice.

### 8.2.3 Body data

Box 255

8.2.4.3 Sphere. Code: SPH

A SPH (Fig. 8.3) is defined by 4 numbers:  $V_{x_i}$   $V_{y_i}$   $V_z$ 

- $^{\star}$  centred at point x=5, y=5, z=5, its axis making equal angles to  $^{\star}$  the three coordinate axes).

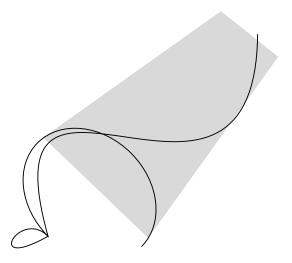


Fig. 8.6: Truncated Right Angle Cone (TRC)

*+1+.	2 + 3 + 4 .	+ 5 + 6	+7+.
ELL 003	-400.0	0.0	0.0
	400.0	0. 0	0.0
	1000.		

The same example in free format:

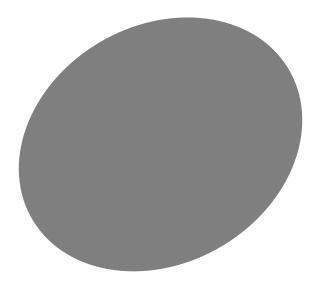


Fig. 8.7:

Infinite half-space 261

Infinite Cylinders 264

\*...+...1...+...2...+...3...+...4...+...5...+...6...+...7.. XCC 013 -480.0 25.0 300.0 
\* (an infinite cylinder of radius 300 cm, with axis defined by y=-480, 
\* z=25)

*.	+	1.	+ 2 + 3 .	+	4 +	5 +	. 6 + 7.	+.
	ZEC	101	15.	0		319.0		33.0
			80.	0				

The same body, described in free format:

ZEC ChimneyA 15.0 319.0 33.0 80.0

### NOTE

for this so-called "contiguity list", and it is not essential that it be exact (if left blank, it is set to 5).

Default = 1.0

WHAT(4) =

270 voxel geometry

voxel geometry 271

```
PARAMETER ( NX = 256 )
PARAMETER ( NY = 256 )
PARAMETER ( NZ = 220 )
DIMENSION GOLEM(NX, NY, NZ)
INTEGER*2 GOLEM
CHARACTER TITLE*80
DIMENSION IREG(1000), KREG(1000)
INTEGER*2 IREG, KREG
CALL CMSPPR
DO IC = 1, 1000
   KREG(IC) = 0
OPEN(UNIT=30, FILE='ascii_segm_golem', STATUS='OLD')
READ(30, *) GOLEM
NO=0
MO = 0
DO IZ=1, NZ
   DO IY=1, NY
       DO IX=1, NX1, NY
```

272 voxel geometry

- The usual region list of NR regions, with the space occupied by body NB+1 (the "voxel volume") subtracted. In other words, the NR regions listed must cover the whole available space, except the space corresponding to the "voxel volume". This is easily obtained by subtracting body NB+1 in the relevant region definitions, even though this body is not explicitly input at the end of the body list. The code will automatically generate and add several regions:

- a region NR+1

## Chapter 9

### Output

The output of Fluka consists of:

- a main (standard) output, written on logical output unit LUNOUT (predefined as 11 by default)
- a scratch file, of little interest to the user, written on output unit LUNGEO (16 by default). However, if the rfl uka

Main Output 275

to option LOW-NEUT, p. 141

Main Output 277

278 Main Output

bremsstrahlung by heavy particles (set with PAIRBREM, p. 170).

(k) Photonuclear reaction requests

280 Random Number Seeds

– Energy deposited by hadron and muon dE/dx

RANDOMIZe (p. 186, is written in hexadecimal format.

### 9.4 Error messages

282 EVENTBIN Output

1.

```
Cartesian binning n. 1 "Cufront ", generalised particle n. 208

X coordinate: from -2.1100E-01 to 5.5910E+00 cm, 58 bins (1.0003E-01 cm wide)

Y coordinate: from 0.0000E+00 to 5.4010E+00 cm, 53 bins (1.0191E-01 cm wide)

Z coordinate: from 0.0000E+00 to -1.0000E-03 cm, 1 bins (-1.0000E-03 cm wide)

Data follow in a matrix A(ix,iy,iz), format (1(5x,1p,10(1x,e11.4)))
```

288 USERDUMP Output

 $\ensuremath{^{\star\star\star\star\star}}$  Test track-length/coll. reading program for the manual

\*\*\*\*

DATE: 10/25/ 4, TIME: 10:32:59

### 9.7 RAY output

Tracking RAY pseudoparticles (Chap. 14, p. 336) produces only an unformatted file. No formatted output is available.

9.8-720.902cmBT/F830.580Td[(RA)1(YUs-1()]TJ/r-generted)-3-1()(out76tput)]TJ/F89.963Tf-53.5190Tc

# Chapter 10

# Low-energy neutrons in FLUKA

10.1

Neutron library 293

The standard cross-section set has 72 neutron energy groups and 22 gamma groups, a structure which has been chosen for practical considerations. Gamma energy groups are used only for (n, ) production, since transport of photons in Fluka is continuous in energy and angle and is performed through the EMF module.

Each material is identified by an alphanumeric name (a string not longer than 8 characters, all in upper case), and by three integer identifiers. Correspondence with Fluka

294 Neutron library

Table 10.1: Neutron energy group structure of the 72-group ENEA liroubrary

Neutron	Lower limit	Upper limit
group n.	(GeV)	(GeV)
1	$1.7500 \cdot 10^{-2}$	$1.9600 \cdot 10^{-2}$
2	$1.4918 \cdot 10^{-2}$	$1.7500 \cdot 10^{-2}$
3	$1.3499 \cdot 10^{-2}$	$1.4918 \cdot 10^{-2}$
4	$1.2214 \cdot 10^{-2}$	$1.3499 \cdot 10^{-2}$
5	$1.1052 \cdot 10^{-2}$	$1.2214 \cdot 10^{-2}$
6	$1.0000 \cdot 10^{-2}$	$1.1052 \cdot 10^{-2}$
7	$9.0484 \cdot 10^{-3}$	$1.0000 \cdot 10^{-2}$
8	$8.1873 \cdot 10^{-3}$	$9.0484 \cdot 10^{-3}$
9	$7.4082 \cdot 10^{-3}$	$8.1873 \cdot 10^{-3}$
10	$6.7032 \cdot 10^{-3}$	$7.4082 \cdot 10^{-3}$
11	$6.0653 \cdot 10^{-3}$	$6.7032 \cdot 10^{-3}$
12	$5.4881 \cdot 10^{-3}$	$6.0653 \cdot 10^{-3}$
13	$4.9659 \cdot 10^{-3}$	$5.4881 \cdot 10^{-3}$
14	$4.4933 \cdot 10^{-3}$	$4.9659 \cdot 10^{-3}$
15	$4.0657 \cdot 10^{-3}$	$4.4933 \cdot 10^{-3}$
16	$3.6788 \cdot 10^{-3}$	$4.0657 \cdot 10^{-3}$
17	$3.3287 \cdot 10^{-3}$	$3.6788 \cdot 10^{-3}$
18	$3.0119 \cdot 10^{-3}$	$3.3287 \cdot 10^{-3}$
19	$2.7253 \cdot 10^{-3}$	$3.0119 \cdot 10^{-3}$
20	$2.4660 \cdot 10^{-3}$	$2.7253 \cdot 10^{-3}$
21	$2.2313 \cdot 10^{-3}$	$2.4660 \cdot 10^{-3}$
22	$2.0190 \cdot 10^{-3}$	$2.2313 \cdot 10^{-3}$
23	$1.8268 \cdot 10^{-3}$	$2.0190 \cdot 10^{-3}$
24	$1.6530 \cdot 10^{-3}$	$1.8268 \cdot 10^{-3}$
25	$1.4957 \cdot 10^{-3}$	$1.6530 \cdot 10^{-3}$
26	$1.3534 \cdot 10^{-3}$	$1.4957 \cdot 10^{-3}$
27	$1.2246 \cdot 10^{-3}$	$1.3534 \cdot 10^{-3}$
28	$1.1080 \cdot 10^{-3}$	$1.2246 \cdot 10^{-3}$
29	$1.0026 \cdot 10^{-3}$	$1.1080 \cdot 10^{-3}$
30	$9.0718 \cdot 10^{-4}$	$1.0026 \cdot 10^{-3}$
31	$8.2085 \cdot 10^{-4}$	$9.0718 \cdot 10^{-4}$
32	$7.4274 \cdot 10^{-4}$	$8.2085 \cdot 10^{-4}$
33	$6.0810 \cdot 10^{-4}$	$7.4274 \cdot 10^{-4}$
34	$4.9787 \cdot 10^{-4}$	$6.0810 \cdot 10^{-4}$
35	$4.0762 \cdot 10^{-4}$	$4.9787 \cdot 10^{-4}$
36	$3.3373 \cdot 10^{-4}$	$4.0762 \cdot 10^{-4}$

4 4. . 4

Available materials 295

Table 10.2: Gamma energy group structure of the ENEA library

Gamma Lower limia Upper limia

296 Available materials

... Continues...

298 Available materials

. . . Continues. . .

## Chapter 11

## Collision tape

### 11.1 What is a collision tape and what is its purpose

A U134collision tape" is a U014le where quantities describing selected events are recorded in the coufse of the c

- 3) data concerning local (point) energy deposition
- 4) any combination of the three above

Data are written on the collision tape in single precision and unformatted, but it is also possible for the user to modify the MGDRAW

- = 10: elastic interaction recoil
- = 11: inelastic interaction recoil
- = 12: stopping particle
- = 14: escape
- ICODE = 2x: call from subroutine EMFSCO (electromagnetic part of Fluka);
  - = 20: local energy deposition (i.e. photoelectric)
  - = 21: below user-defined transport cut-o (but larger than EMF production cut-o )
  - = 22: below both user-defined transport cut-o and EMF production cut-o
  - = 23: escape
- ICODE = 3x: call from subroutine KASNEU (low-energy neutron part of Fluka)
  - = 30: target recoil
  - = 31: neutron below threshold
  - = 32: escape
- ICODE = 4x: call from subroutine KASHEA (heavy ion part of Fluka)
  - = 40: escape
- ICODE = 5x: call from subroutine KASOPH (optical photon part of Fluka)
  - = 50: optical photon absorption
  - = 51: escape

In Case 3, the following variables are written:

First record:

 $2270.\ 2673g0G1001-538.\ 6308d \hbox{\tt [(2-270.\ 2673g0179.\ 63Tf-52.\ 1308d \hbox{\tt [(2270.\ 2673051)]TJ/F179.\ 983Tf-22)}}\\$ 

# Chapter 12

User routines

User routines 303

#### 12.1.1 INCLUDE files

304 COMSCW

#### 12.2 Description of available user routines

#### 12.2.1 ABSCFF: user defined ABSorption Coefficient

Argument list (all variables are input only)

WVLNGT: photon wavelength (in cm)

OMGPHO: angular frequency ( = 2 ) of the photon (in  $s^{-1}$ )

MMAT : material index

Function ABSCFF

FORMFU 307

12.2.7

308 LATTIC

LUSRBL defines another discrete (integer) variable (by default: lattice number)

#### **Argument list** (all variables are input only)

IJ: particle type PCONTR: particle momentum

XFLK, YFLK, ZFLK: particle position

MREG : current region LATCLL : current lattice cell

I CALL: internal code calling flag (not for general use)

FUSRBV defines a continuous (double precision) variable (by default: pseudorapidioy with respect to the Z axis)

#### Argument list (all variables are input only)

IJ : particle type PCONTR : particle momentum

XFLK, YFLK, ZFLK: particle position

MREG : current region

ICALL: internal code calling flag (not for general use)

The 3 functions are called at track-length events. What is scored is the particle track-length multiplied by the particle's weight, possibly modified by a user-written FLUSCW (1(c)120.6), as a function of the 3 variables defined by MUSRBR, LUSRBL and FUSRBV.

#### 12.2.10 LATTIC: symmc transformation for lattice geometry

Subroutine LATTIC is activated by one or more LATTICE cards in the geometry input (1(c)812.9). It is expected to transform coordinates and direction cosines from any lattice cell (defined by card LATTICE) to

MAGFLD 309

LATTIC returns the tracking point coore-43cnateSB( card).

When 1(g)-430284(I)1(attic430283(option)-282(is30283(activ)56(ated,)-293(1(g)-430283(trac)

310 MGDRAW

The magnetic field spatial distribution is often read and interpolated from an external field map. Note that in any case the direction cosines must be properly normalised in double precision (e.g., BTX = SQRT(ONEONE - BTY\*\*2 - BTZ\*\*2)), even if B = 0.0.

Please read carefully the notes on option MGNFIELD (p. 151).

12.2.12

MGDRAW 311

MTRACK: number of energy deposition events along the track

JTRACK: type of particle

ETRACK: total energy of the particle

WTRACK: weight of the particle

NTRACK values of XTRACK, YTRACK, ZTRACK: end of each track segment MTRACK values of DTRACK: energy deposited at each deposition event

CTRACK:

312 MGDRAW

I CODE : see argument list JTRACK, ETRACK, WTRACK : see MGDRAW above

OPHBDX 313

Information about the secondary particles produced is available in COMMON GENSTK

#### 12.2.16 RFLCTV: user defined ReFLeCTiVity

**Argument list** (all variables are input o-ly)

WVLNGT: photo-wavelength (i-cm)

OMGPHO: angular frequency (=2) of the photo- (i- s<sup>-1</sup>)

MMAT : material i-dex

Functio- RFLCTV returns a user-defined of the current material for an optical photo- of the give-wavelength or frequency.

It is activated by command OPT-PROP with SDUM = METAL and WHAT(3) < -99. See Sec. 7.51 a-d Chap. 13 for more informatio-.

12.2.17

#### 12.2.19.4 Sampling from a generic distribution

One way to sample a value  $\boldsymbol{X}$ 

INCLUDE '(SOURCM)' INCLUDE '(CHEPSR)'

These values can be used as parameters or switches for a multi-source routine capable to handle several cases, or to identify an external file to be read, etc., without having to compile and link again the routine.

In the SOURCE routine there are a number of mandatory statements, (clearly marked as such in accompanying comments) which must not be removed or modified. The following IF blo s27(k)-253(init)1(ialise)-1(s)-254(th)b(e)s256(t)exist.

12.2.24)

At the time SOURCE is called, the particle bank FLKSTK

STUPRE, STUPRF 319

TZFLK (NPFLKA) = SQRT ( ONEONE - TXFLK(NPFLKA)\*\*2 - TYFLK(NPFLKA)\*\*2 )

**UBSSET** 321

#### )GNONA:

... Continuation of the UBSSET argument lict

### (WHAT(1) in card LOW-BIAS Argument lict

I GCUTO: Cut-o group index for low-energy neutrons in region I R (WHAT(1) in card LOW-BIAS, p. 135)
I GNONA: Non-anald@ue absorption group limit for low-energy neutrons in

322 USREOU

12.2.22 U ser defined D e C ay D I R ection biasing and L ambda (for only)

326 Examples

That option sets the quantum e ciency as a function of photon energy *overall* through the problem

The examples cresented here consider 0.5 GeV muons in a box of 4  $\times$  4  $\times$  4 m $^3$ . In order to avoid

328 Examples

TITLE								
Test of Cherenkov light production in Liquid Argon								
DEFAULTS						PRECISIO		
*+1	+ 2	. + 3	. + 4	+ 5	+ 6	+ 7 +		
BEAM	-10.000					MUON+		
BEAMPOS	0.0	0.0	190. 0			NEGATI VE		
DELTARAY	-1.0			18.0	18. 0			
PAIRBREM	-3.0			18.0	18. 0			
MUPHOTON	-1.0			18.0				
PHOTONUC	-1.0			3.0	100. 0			
EVENTYPE			6.0					
DISCARD	27. 0	28.0	43.0	44.0	5.0	6.0		
GEOBEGI N						COMBINAT		
	Test							
*+1+2+3+4+5+6+7+								
* A large box for the blackhole								
RPP 1 -9999999. +999999999999999999999999. +99999999.								
* A smaller				000 0	000 0	202 2		
RPP 2	-200.0	+200.0	-200.0	+200.0	-200. 0	+200. 0		
END *== Region D	ofinitions							
		=======	=======	=======	=======	=======		
•	ackhol e +1 -	2						
BL1		2						
	quid Argon							
LG3 END	+2							
GEOEND								
	foloctron	and photon	tranchor	+				
SWITCH OF	f electron	anu photon	i ti anspoi	ι				

EMF

\*...+...1....+...2....+...3...+...4...+...5...+...6...+...7....+
USRBDX 1.0 -55.0 2.0 1.0
USRBDX 0.0 120.0 &
USERDUMP 2.

330 Examples

```
INCLUDE '(CASLIM)'
     INCLUDE ' (COMPUT)'
INCLUDE ' (FHEAVY)'
     INCLUDE '(FLKSTK)'
INCLUDE '(GENSTK)'
      INCLUDE ' (MGDDCM)'
      INCLUDE ' (PAPROP)'
     INCLUDE ' (SOURCM)'
     INCLUDE '(STARS)'
      INCLUDE '(TRACKR)'
     CHARACTER*20 FILNAM
     LOGICAL LFCOPE
     SAVE LFCOPE
     DATA LFCOPE / . FALSE. /
*_____
     lcode = 1: call from Kaskad
     lcode = 2: call from Emfsco
     lcode = 3: call from Kasneu
     lcode = 4: call from Kashea
     lcode = 5: call from Kasoph
     IF ( .NOT. LFCOPE ) THEN
        LFCOPE = .TRUE.
        IF ( KOMPUT . EQ. 2 ) THEN
           FILNAM = '/'//CFDRAW(1:8)//' DUMP A'
        ELSE
           FILNAM = CFDRAW
        END IF
        WRITE(*,*) 'TRAJECTORY OPEN!'
        WRITE(*,'(A)') 'FILNAM = ', FILNAM
        OPEN ( UNIT = IODRAW, FILE = FILNAM, STATUS = 'NEW', FORM =
               'UNFORMATTED' )
    &
     END IF
С
С
     Write trajectories of optical photons
     IF(JTRACK . EQ. -1) THEN
        WRITE (IODRAW) NTRACK, MTRACK, JTRACK, SNGL (ETRACK),
             SNGL (WTRACK)
        WRITE (IODRAW) ( SNGL (XTRACK (I)), SNGL (YTRACK (I)),
             SNGL (ZTRACK (I)), I = 0, NTRACK ),
    &
             (SNGL (DTRACK (I)), I = 1, MTRACK),
    &
             SNGL (CTRACK)
        WRITE(IODRAW) SNGL(CXTRCK), SNGL(CYTRCK), SNGL(CZTRCK)
     ENDIF
     RETURN
 Boundary-(X)crossing DRAWing:
     lcode = 1x: call from Kaskad
            19: boundary crossing
      lcode = 2x: call from Emfsco
             29: boundary crossing
     lcode = 3x: call from Kasneu
```

\* 39: boundary crossing

\* Icode = 4x: call from Kashea

\* 49: boundary crossing

*	<u> </u>	
^	RETURN	. ^
*		
*=		: *
*		*
*	USer dependent DRAWing:	*
*		*
*	code = 10x: call from Kaskad	*
*	100: elastic interaction secondaries	*
*	101: inelastic interaction secondaries	*
*	102: particle decay secondaries	*
*	103: delta ray generation secondaries	*
*	104: pair production secondaries	*
*	105: bremsstrahlung secondaries	*
*	<pre>lcode = 20x: call from Emfsco</pre>	*
*	208: bremsstrahlung secondaries	*
*	210: Moller secondaries	*
*	212: Bhabha secondaries	*
*	214: in-flight annihilation secondaries	*
*	215: annihilation at rest secondaries	*
*	217: pair production secondaries	*
*	219: Compton scattering secondaries	*
*	221: photoel ectric secondaries	*

334 Examples

```
SUBROUTINE VXREAD (NCASE)
PARAMETER ( MXH = 2000 )
PARAMETER ( MXPR = 300 )
DIMENSION XH (MXH), YH (MXH), ZH (MXH), DH (MXH),

EPR (MXPR), WPR (MXPR), XPR (MXPR), YPR (MXPR),

ZPR (MXPR), TXP (MXPR), TYP (MXPR), TZP (MXPR),
           IPR (MXPR)
LUNSCR = 33
REWIND (LUNSCR)
NEVT=0
DO 4000 I=1, 200000000
   READ (LUNSCR, END=4100) NDUM, MDUM, JDUM, EDUM, WDUM
   IF(I.EQ. 1) WRITE(*, *) 'NDUM, MDUM, JDUM, EDUM, WDUM', NDUM, MDUM, JDUM
         , EDUM, WDUM
   \mathsf{NEVT} \ = \ \mathsf{NEVT} \ + \ 1
+----*
   Real tracking data:
   IF ( NDUM .GT. O ) THEN
      NTRACFT. I FMT. I FJT. I FET. I FWT.
       &K(UNSQR)T. WRITE(*, *71 'NDUM, MT. = NEVTREWIND(LUNSCR)
      NE(+=58(5)5WRXHE) 7-140 122-10. 958Td[(IFW050*, *71)-525('NDUMUNSC1,)]TOJ
```

# Chapter 14

## Use of RAY pseudo-particles

Pseudo-particles are called RAY

RAY pseudo-particles 337

RAY pseudo-particles

ALITOT =

# Chapter 15

## Examples on the material/compound definitions

15.1 Use of

340 Examples

\* Lead material n. 16, normal temperature, overrides pre-defined mat. 16 MATERIAL 82.0 207.19 11.35 16.0 LEAD

### Chapter 16

## History of FLUKA

#### 16.1 Introduction

The history of Fluka goes back to 1962-1967. During that period, Johannes Ranft was at CERN doing work

Protection Group, and of Jorma Routti, Chairman of the Department of Technical Physics at the Helsinki

Code structure 343

to maintain and update, the possibility to insert freely comments in input, and a special attention devoted to portability (Fluka87 could run only on IBM under VM-CMS).

The greatest importance was attached to numerical accuracy: the whole code was converted to double precision (but the new allocation scheme allowed for implementation also in single precision on 64-bit computers). As a result, energy conservation was ensured within  $10^{-10}$ .

A decision was also made to take systematically maximum advantage from the available machine precision, avoiding all unnecessary rounding and using consistently the latest recommended set of the physical

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in CG by Ferrari, and a powerful debugger facility was made availabla.

final revision and update of the algorithm were made in 1991. In 1995, the Fano correction for multiple

The main lines of the work developed mostly in Milan by Ferrari and Sala starting from 1990 can be summarised as follow [41,70]:

 further develop and improve the high energy DPM-based part of the models. This was performed in 4 main stages, which eventually led to an almost completely new code still based on the same physics foundations with soa

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In 1995, a newly developed Fermi Break-up model, with a maximum of 6 bodies in the exit channel,

Two biasing techniques were implemented by Fassò and Ferrari, which are applicable only to low-energy neutrons.

- Neutron Non Analogue Absorption (or survival biasiqu) was derived from Morse where it was systematically applied and out of user control. In Fl uka it was generalised to give full freedom to the user to fix the ratio between scattering and absorption probability in selected regions and within a chosen energy range. While it is mandatory in some problems in order to keep neutron slowing down under control, it is also possible to switch it o completely to get an analogue simulation.
- Neutron Biased Downscattering, also for low-energy neutrons, gives the possibility to accelerate or slow down the moderating process in selecte28(e)-37ratinw

Dpmjet interface 355

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