ZTF24aahgqwk in NGC 3443

Light Curve Notebook

This notebook begins with the 36 stacked images produced by our Calibration Notebook, and produces a light curve, consisting of 18 Sloan r' and 18 Sloan g' data points.

See a least squares notebook for background for the method used below.

```
In [1]:
        import os
        import numpy as np
        from scipy.optimize import least squares
        from astropy import units as u
        from astropy.nddata import CCDData
        from astropy.io import fits
        from ccdproc import ImageFileCollection
        import astroalign as aa
        import matplotlib.pyplot as plt
         %matplotlib inline
        from math import log10, floor, sqrt, log, exp, pi
        # THIS COMMENT IS THE LONGEST A LINE CAN BE AND STILL RENDER COMPLETELY WHEN PRINTING IN LANDSCAPE MODE.
        # THIS COMMENT IS 72 CHARACTERS WITHOUT COUNTING THE NEWLINE AT THE END.
        # This notebook needs to be able to find the stacked images.
        home directory = os.path.expanduser('~')
        supernova project directory = os.path.join(home directory, 'Projects', 'supernova-observation')
        stacked directory = os.path.join(supernova project directory, 'analyses', 'ZTF24aahgqwk', 'stacked')
        # The 36 images are in the stacked directory. There were 18 observation sessions with 2 filters each.
         # filters
```

```
filters = ['r', 'g']
filter full_names = ["Sloan r'", "Sloan g'"]
# observation dates (UTC)
observation dates = [
    '2024-03-20',
    '2024-03-21',
    '2024-03-23',
    '2024-03-27',
    '2024-04-02',
    '2024-04-03',
    '2024-04-04',
    '2024-04-06',
    '2024-04-10',
    '2024-04-11',
    '2024-04-13',
    '2024-04-17',
    '2024-04-21',
    '2024-04-22',
    '2024-04-23',
    '2024-04-29',
    '2024-04-30',
    '2024-05-02'
# We will need to specify rectangles surrounding the target and the reference stars.
# use named tuples to improve readability
from collections import namedtuple
Point = namedtuple('Point', 'x y')
Extent = namedtuple('Extent', 'width height')
Rectangle = namedtuple('Rectangle', 'center extent')
# Various utilities
```

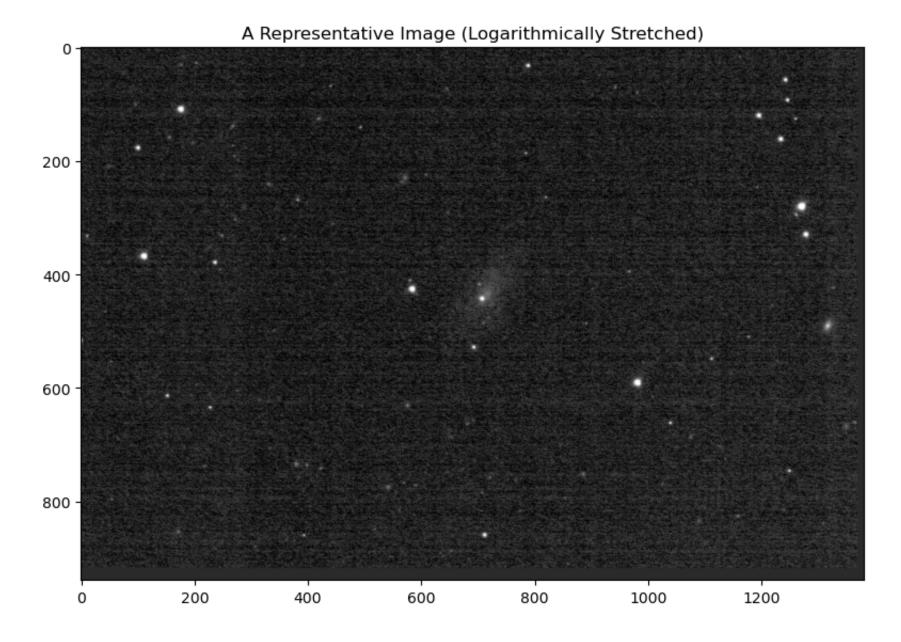
```
def file for date and filter(date, filter):
    return os.path.join(stacked directory, date + '-' + filter + '_stacked.fit')
def stacked_image_for_date_and_filter(date, filter):
    file = file for date and filter(date, filter)
    return CCDData.read(file, unit=u.adu)
# Log stretch utility
def log stretch transform(black point, saturation range):
    log saturation range = log10(saturation range)
    def fn(pixel value):
        pixel value -= black point
        if pixel value <= 1.0:</pre>
            return 0
        else:
            log_pixel_value = log10(pixel_value)
            if log_pixel_value >= log_saturation_range:
                return 255;
            else:
                return floor(256 * log_pixel_value / log_saturation_range)
    return fn
```

Specify the Regions of Interest for the Target and Reference Stars

```
In [2]: # Guarantee the extent widths and heights are odd so the loops do not have to handle even and odd cases.
EXTENT_HALF_WIDTH = 10
EXTENT_WIDTH = 2 * EXTENT_HALF_WIDTH + 1
EXTENT_HEIGHT = EXTENT_WIDTH
EXTENT = Extent(EXTENT_WIDTH, EXTENT_HEIGHT)
CENTERS = [
    Point(708, 443), # target
    Point(177, 109), # reference star at far upper left
    Point(112, 368), # reference star at far left
    Point(585, 426), # reference star just left of center
    Point(982, 591), # reference star right of center
    Point(1271, 280), # reference star at far right
]
CENTER_COUNT = len(CENTERS)
```

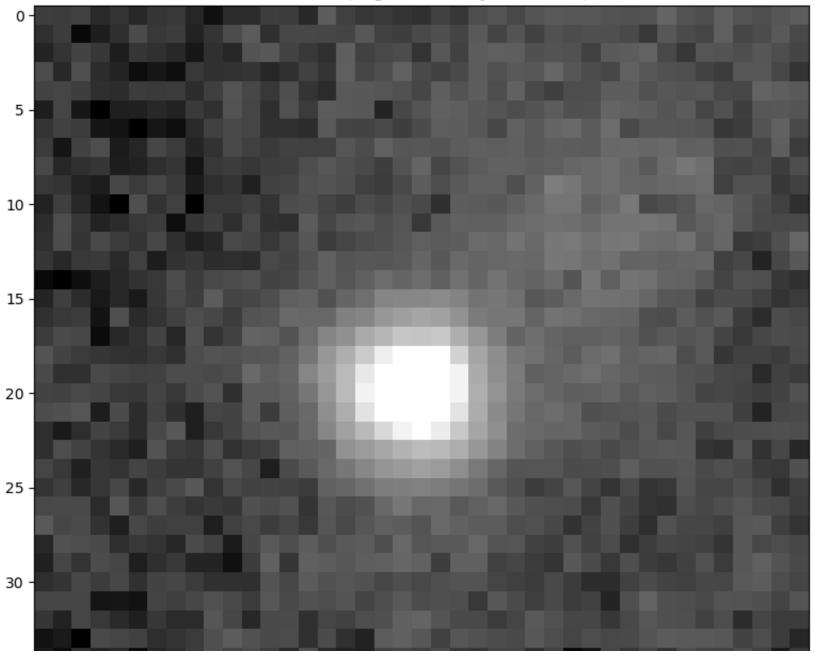
Display a Representative Image

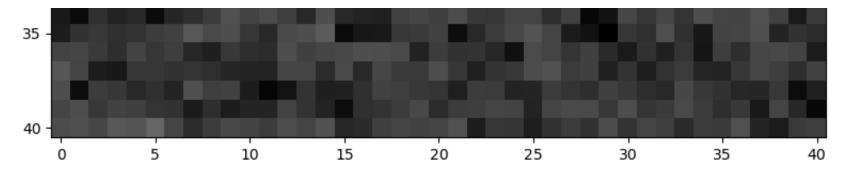
```
In [3]: first image = stacked image for date and filter('2024-04-03', 'r')
        # Log stretch
        stretch function = log stretch transform(8, 50)
        stretch transform = np.vectorize(stretch function)
        stretched_image = stretch_transform(first_image.data)
        # Display the image
        fig, axes = plt.subplots(1, 1, figsize=(8, 8))
        axes.imshow(stretched image, cmap='gray')
        axes.set title("A Representative Image (Logarithmically Stretched)")
        plt.tight layout()
        plt.show()
        WARNING: FITSFixedWarning: 'datfix' made the change 'Set MJD-OBS to 60403.212974 from DATE-OBS'. [astrop
        y.wcs.wcs]
        WARNING: FITSFixedWarning: 'obsfix' made the change 'Set OBSGEO-X to -2381449.053 from OBSGEO-[LBH].
        Set OBSGEO-Y to -4483194.922 from OBSGEO-[LBH].
        Set OBSGEO-Z to 3851220.317 from OBSGEO-[LBH]'. [astropy.wcs.wcs]
```



```
In [4]: DISPLAY_WHICH_CENTER = 0
        display extent half = 20
        display_extent_width = 2 * display_extent_half + 1
        display extent height = display extent width
        display extent = Extent(display extent width, display extent height)
        display center = CENTERS[DISPLAY WHICH CENTER]
        display left = display center.x - display extent half
        display right = display left + display extent width
        display top = display center.y - display extent half
        display bottom = display top + display extent height
        subframe = stretched image[display top:display bottom, display left:display right]
        # Display the representative subtracted dark
        fig, axes = plt.subplots(1, 1, figsize=(8, 8))
        axes.imshow(subframe, cmap='gray')
        axes.set_title("Subframe (Logarithmically Stretched)")
        plt.tight_layout()
        plt.show()
```

Subframe (Logarithmically Stretched)





```
In [5]: def sigma squared for fwhm(fwhm):
            return fwhm**2 / (8 * log(2))
        def make gaussian with flux(fwhm, flux, center x, center y):
            sigma squared = sigma squared for fwhm(fwhm)
            flux factor = flux / 2 / pi / sigma squared
            def gaussian(x, y):
                delta x = x - center x
                delta y = y - center y
                return flux factor * exp(-(delta x**2 + delta y**2) / 2 / sigma squared)
            return gaussian
        def model data for parameters(extent, total background, fwhm, flux, center x, center y):
            model data = np.zeros([extent.height, extent.width])
            model data.fill(total background)
            gaussian = make gaussian with flux(fwhm, flux, center x, center y)
            for j in range(EXTENT HEIGHT):
                for i in range(EXTENT WIDTH):
                    model data[j, i] += gaussian(i, j)
            return model data
        # IS THERE A WAY TO USE NP. VECTORIZE?
        PINDEX TARGET BACKGROUND = 0 # NB: THE TARGET BACKGROUND IS IN ADDITION TO THE GENERAL BACKGROUND
        PINDEX_GENERAL_BACKGROUND = 1
        PINDEX_FWHM = 2
        POFFSETINDEX FLUX = 0
```

```
POFFSETINDEX CENTER X = 1
POFFSETINDEX CENTER Y = 2
def roi residuals(image_data, target_index, center, extent, parameter_vector):
    left = center.x - extent.width // 2
    right = left + extent.width
    top = center.y - extent.height // 2
    bottom = top + extent.height
    roi data = image data[top:bottom, left:right]
    target background = parameter vector[PINDEX TARGET BACKGROUND]
    general background = parameter vector[PINDEX GENERAL BACKGROUND]
    total background = target background + general background if target index == 0 else general background
    fwhm = parameter vector[PINDEX FWHM]
    base index = 3 + 3 * target index
    flux = parameter vector[base index + POFFSETINDEX FLUX]
    center x = parameter vector[base index + POFFSETINDEX CENTER X]
    center_y = parameter_vector[base_index + POFFSETINDEX CENTER Y]
    model data = model data for parameters(extent, total background, fwhm, flux, center x, center y)
    return roi data - model data
def make residuals function(image data, centers, extent):
    def residuals function(parameter vector):
        all roi residuals = [
            roi residuals(image data, i, center, extent, parameter vector)
            for i, center in enumerate(centers)
        return np.concatenate(all roi residuals).ravel()
    return residuals function
initial guess for general background = 0.0
initial guess for target background = 0.0
initial guess for fwhm = 5.0
initial parameter_vector = [
    initial guess for target background,
    initial guess for general background,
    initial guess for fwhm
```