

HW 10: Physics 545

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1

The total entropy is the sum of the phonons + rotons. The distribution of the rotons with spin degeneracy g_r is $n_{\mathbf{p}} \approx g_r \xi e^{-\beta \frac{(p-p_0)^2}{2m}}$, with $\xi = e^{-\beta \Delta}$. We are assuming the temperature is small compared to the energy gap, $\beta \Delta \gg 1$, $\xi \ll 1$, and $\beta = 1/T$. The entropy of the rotons bosons is:

$$S_r = \frac{g_r}{(2\pi\hbar)^3} \int d^3\mathbf{p} \quad (1 + n_{\mathbf{k}}) \ln(1 + n_{\mathbf{k}}) - n_{\mathbf{k}} \ln(n_{\mathbf{k}}) \quad (1)$$

$$= \frac{4g_r\pi}{(2\pi\hbar)^3} \int dp \quad p^2 \left[(1 + \xi e^{-\beta \frac{(p-p_0)^2}{2m}}) \ln(1 + \xi e^{-\beta \frac{(p-p_0)^2}{2m}}) \right. \quad (2)$$

$$\left. - \xi e^{-\beta \frac{(p-p_0)^2}{2m}} \left(\ln(\xi) - \beta \frac{(p-p_0)^2}{2m} \right) \right] \quad (3)$$

$$\approx \frac{4g_r\pi\xi}{(2\pi\hbar)^3} \int dp \quad p^2 e^{-\beta \frac{(p-p_0)^2}{2m}} \left(\beta \Delta + \beta \frac{(p-p_0)^2}{2m} \right) \quad (4)$$

And the first term in the integral can be discarded in the low temperature expansion $1 + \xi e^{-\beta \frac{(p-p_0)^2}{2m}} \approx 1$ because the gaussian is always less than 1.

We make the substitution $x = (p - p_0)/c$, $c = \sqrt{2mT}$

$$\begin{aligned} S_r &\approx \frac{4g_r\pi\xi c}{(2\pi\hbar)^3 T} \int_{-p_0/c}^{\infty} dx \quad (cx + p_0)^2 e^{-x^2} \left(\Delta + x^2 T \right) \quad (5) \\ &\approx \frac{4g_r\pi\xi c}{(2\pi\hbar)^3 T} \int_{-p_0/c}^{\infty} dx \quad \left(c^2 T x^4 + 2TCp_0 x^3 (\Delta c^2 + Tp_0^2) x^2 + 2cp_0 \Delta x + p_0^2 \Delta \right) e^{-x^2} \end{aligned}$$

Now, we keep only the lowest order in T (noting that $c \propto \sqrt{T}$) and extending the lower bound $-\frac{p_0}{\sqrt{2mT}} \rightarrow -\infty$.

$$S_r \approx \frac{g_r p_0^2 \Delta}{\hbar^3} \sqrt{\frac{m}{2\pi^3 T}} e^{-\Delta/T} \quad (7)$$

The phonon contribution can be calculated similarly using the phonon distribution to be $n_{\mathbf{p}} = (e^{\beta u p} - 1)^{-1}$:

$$S_p = \frac{g_p}{(2\pi\hbar)^3} \int d^3\mathbf{p} \quad (1 + n_{\mathbf{k}}) \ln(1 + n_{\mathbf{k}}) - n_{\mathbf{k}} \ln(n_{\mathbf{k}}) \quad (8)$$

$$= \frac{4g_r\pi}{(2\pi\hbar)^3} \int dp \quad p^2 (e^{\beta u p} - 1)^{-1} \left[e^{\beta u p} \left(\beta u p - \ln(e^{\beta u p} + 1) \right) \right. \quad (9)$$

$$\left. + \ln(e^{\beta u p} - 1) \right] \quad (10)$$

$$= \frac{4g_r\pi}{(2\pi\hbar)^3} \int dp \quad \beta u p^3 e^{\beta u p} (e^{\beta u p} - 1)^{-1} - p^2 \ln(e^{\beta u p} - 1) \quad (11)$$

The change of variables is to $x = ap$, $a = \beta u = u/T$

$$S_p = \frac{4g_r\pi}{a(2\pi\hbar)^3} \int dx \quad x^3 e^x (e^x - 1)^{-1} - x^2 \ln(e^x - 1) \quad (12)$$

$$S_p = \frac{4g_r\pi}{a^3(2\pi\hbar)^3} \left[-\frac{x^3}{3} \ln(e^x - 1) \right]_0^\infty + \frac{4}{3} \int dx \quad x^3 e^x (e^x - 1)^{-1} \quad (13)$$

$$S_p = \frac{2g_r\pi^2 T^3}{45u^3 \hbar^3} \quad (14)$$

Mathematica gives result of the integral in brackets to be $4\pi^4/45$.

Because the entropy is an additive property, the total phonon-roton entropy is $S_{tot} = S_p + S_r$:

$$S_{tot} = \frac{2g_r\pi^2 T^3}{45u^3 \hbar^3} + \frac{g_r p_0^2 \Delta}{\hbar^3} \sqrt{\frac{m}{2\pi^3 T}} e^{-\Delta/T} \quad (15)$$

The heat capacity is $C_v = T \frac{\partial S}{\partial T}$

$$C_v = \frac{6g_r\pi^2 T^3}{45u^3 \hbar^3} + \frac{g_r p_0^2 \Delta}{\hbar^3} \sqrt{\frac{m}{2\pi^3}} \left[-\frac{1}{2} T^{-3/2} + \Delta T^{-5/2} \right] e^{-\Delta/T} \quad (16)$$

For the density of the normal component, we use the expression from class:

For the phonon distribution:

$$\rho_n^{phon} = \frac{4\pi}{3(2\pi\hbar)^3} \int dp \quad p^4 \left(-\frac{\partial n}{\partial \epsilon} \right) \quad (17)$$

$$\rho_n = -\frac{4\pi}{3u(2\pi\hbar)^3} \int dp \quad p^4 \left(\frac{\partial n_p}{\partial p} \right) \quad (18)$$

$$\rho_n = \frac{16\pi}{3u(2\pi\hbar)^3} \int dp \quad p^3 n_p \quad (19)$$

$$\rho_n = \frac{16\pi}{3ua^4(2\pi\hbar)^3} \int dp \quad \frac{x^3}{e^x - 1} \quad (20)$$

$$\rho_n = \frac{2\pi^2 T^4}{45u^5 \hbar^3} \quad (21)$$

$$(22)$$

And we have used integration by parts in the last step. The integral is the DEBYE INTEGRAL= $\pi^4/15$.

And for the roton distribution, using the same substitutions as for the roton entropy:

$$\rho_n = \frac{4\pi\beta\xi}{3(2\pi\hbar)^3} \int dp \quad p^4 e^{-\beta(p-p_0)^2/2m} \quad (23)$$

$$\rho_n = \frac{4\pi\beta\xi c}{3(2\pi\hbar)^3} \int_{-p_0/c}^{\infty} dx \quad (cx + p_0)^4 e^{-x^2} \quad (24)$$

$$\rho_n \approx \frac{4\pi\beta\xi c}{3(2\pi\hbar)^3} \int_{-\infty}^{\infty} dx \quad \left(c^4 x^4 + 4p_0 c^3 x^3 + 6p_0^2 c^2 x^2 + 4p_0^3 c x + p_0^4 \right) e^{-x^2} \quad (25)$$

Again, we use the low temperature limit to extend the lower limit $-p_0/c \rightarrow \infty$ and we keep only the lowest order in T which is the last term in the brackets:

$$\rho_n^{rot} \approx \frac{p_0^4}{6\hbar^3} \sqrt{\frac{2m}{\pi^3 T}} e^{-\beta\Delta} \quad (26)$$

2a

The energy of a single vortex is given by $E_v = \frac{1}{2} \int d^3\mathbf{r} \rho_s \mathbf{v}_s^2$. We can write this as the energy per length along the cylinder ϵ_v

$$\epsilon_v = \frac{\rho_s}{2} \int dr d\phi \quad r \Theta(r - r_c)^2 \Gamma_n^2 / r^2 \quad (27)$$

$$\epsilon_v = \rho_s \Gamma_n^2 \pi \int dr \quad \Theta(r - r_c)^2 / r \quad (28)$$

$$\epsilon_v = \rho_s \Gamma_n^2 \pi \left[\ln(r) \right]_{r_c}^R \quad (29)$$

$$\epsilon_v = \pi \rho_s \left(\frac{\hbar n}{m} \right)^2 \ln(R/r_c) \quad (30)$$

and we use $\Gamma_n = \frac{\hbar}{m} n$ and $\mathbf{v}_s = \frac{\Gamma_n}{r} \hat{r}$.

The superfluid momentum is $\mathbf{p}_s = m \mathbf{v}_s = \frac{\hbar n}{r} \hat{\phi}$, and angular momentum is $\mathbf{L} = \mathbf{r} \times \mathbf{p}$. The total angular momentum of the vortex is:

$$\mathbf{L}_s = \int d^3\mathbf{r} \quad \mathbf{r} \times \mathbf{p} \quad (31)$$

$$\mathbf{L}_s = \int dr d\theta dz \quad r((r\hbar n/r)\hat{z} - (z\hbar n/r)\hat{r}) \quad (32)$$

$$(33)$$

the \hat{r} term integrates to 0, and the \hat{z} part we write as angular momentum per length \mathbf{l}_s :

$$\mathbf{l}_s = \pi \hbar n (R^2 - r_c^2) \hat{z} \quad (34)$$

The critical frequency ω_c is when $\epsilon_s = |\mathbf{l}_s| \omega_c$:

$$\omega_c = \rho_s \left(\frac{\hbar n}{m^2} \right) \frac{\ln(R/r_c)}{R^2 - r_c^2} \quad (35)$$

2b

The circulation along a circle of radius r inside the star is:

$$2\pi \Gamma_r = \int_C \mathbf{v}_s \cdot d\mathbf{l} = \int \int \vec{\nabla} \times \mathbf{v}_s \cdot d\mathbf{s} \quad (36)$$

$$= 2\pi \Gamma_s N(r) \quad (37)$$

$\Gamma_s = \hbar/m$ is the single quantized circulation of one vortex and m is the mass of the condensate particles (probably neutrons i guess...)

$N(r) = \pi r^2 n_v$ is the number of vortices contained in the circle of radius r with n_v the area density of vortices which is constant.

The average velocity of this circulation is $\mathbf{v}_s = \frac{\Gamma_r}{r} \hat{\phi}$, and the angular momentum associated with this circulation of all the particles with total mass M is:

$$\mathbf{L} = M \Gamma_r \hat{z} = 2\pi^2 \hbar n_v r^2 \frac{M}{m} \hat{z} \quad (38)$$

Extending the radius of the circle to the edge of the sphere $r = R$, and assuming the mass $M = M_{star}$. Then we can equate this angular momentum with that of the spinning sphere of mass M and frequency Ω , $\mathbf{L} = \frac{2}{5} M R^2 \Omega \hat{z}$

$$n_v = \frac{m \Omega}{5 \pi^2 \hbar} \quad (39)$$