# HW 12: Physics 545

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### 1a

#### FIGURE ATTACHED

The density of states  $D(\epsilon)$  for this 2D dispersion of graphene is:

$$D(\epsilon) = \sum_{\mathbf{k}s} \delta(\epsilon - \epsilon_{\mathbf{k}s}) \tag{1}$$

$$= 2\pi V/(2\pi\hbar)^2 \int dk \quad k(\delta(\epsilon - v_f k) + \delta(\epsilon + v_f k))$$
 (2)

$$= D_0|\epsilon| \tag{3}$$

with  $D_0 = \frac{V}{2\pi\hbar^2 v_f^2}$ , and we note that it is independent of spin s

Also, it's kind of "gapped" in that there are no states at zero energy.

## 1b

At zero temperature, with  $\mu(0) = 0$ , the contribution to number of electrons is only from s = -1 because the Fermi distribution function  $f_{\mathbf{k}1} = 0$  and  $f_{\mathbf{k}-1} = 1$ 

$$N_0 = D_0 \int_0^\infty d\epsilon \quad \epsilon \tag{4}$$

The above is divergent, but we can still compare with the electron number for finite temperature:

$$N(T) = \sum_{\mathbf{k}s} (e^{\beta(\epsilon_{\mathbf{k}s} - \mu)} + 1)^{-1}$$
(5)

$$= D_0 \int_{-\infty}^{\infty} d\epsilon \ |\epsilon| (e^{\beta(\epsilon-\mu)} + 1)^{-1}$$
 (6)

$$= D_0 \int_0^\infty d\epsilon \ \epsilon \left[ 1 + (e^{\beta(\epsilon - \mu)} + 1)^{-1} - (e^{\beta(\epsilon + \mu)} + 1)^{-1} \right]$$
 (7)

$$= N_0 + D_0 \int_0^\infty d\epsilon \ \epsilon \left[ (e^{\beta(\epsilon - \mu)} + 1)^{-1} - (e^{\beta(\epsilon + \mu)} + 1)^{-1} \right]$$
 (8)

$$= N_0 + D_0 \left[ \int_{-\mu}^{\mu} dx \frac{x+\mu}{e^{\beta x} + 1} + 2\mu \int_{\mu}^{\infty} dx \frac{1}{e^{\beta x} + 1} \right]$$
 (9)

For conservation of particles we require  $N(T) = N_0$ . Therefore, the term in brackets (8) must be zero, so  $\mu = 0$ 

#### 1c

To calculate the specific heat we use the form  $C = \frac{dU}{dT}$ , U is the internal energy.

$$U = D_0 \int_{-\infty}^{\infty} d\epsilon \ |\epsilon| \epsilon f(\epsilon)$$
 (10)

$$= D_0 \int_0^\infty d\epsilon \ \epsilon^2 (f(\epsilon) - f(-\epsilon))$$
 (11)

$$= D_0 \int_0^\infty d\epsilon \ \epsilon^2 \left(2f(\epsilon) - 1\right) \tag{12}$$

$$= 2T^{3}D_{0}\int_{0}^{\infty} dx \frac{x^{2}}{e^{x}+1} - D_{0}\int_{0}^{\infty} d\epsilon \epsilon^{2}$$
 (13)

The second term is the ground state energy  $E_0$  and does not contribute to the specific heat because there is no temperature dependence there.

The integral in the first term can be found in tables and the result is

$$\delta U = 3T^3 D_0 \zeta(3) + E_0 \tag{14}$$

$$C = 9T^2 D_0 \zeta(3) \tag{15}$$

The equation for minimum of energy is:

$$\frac{\partial F}{\partial \boldsymbol{\eta}} = 0 = \sum_{i=1,2} \alpha \eta_i + \beta_1 (\boldsymbol{\eta} \cdot \boldsymbol{\eta}^*) \eta_i + \beta_2 (\boldsymbol{\eta} \cdot \boldsymbol{\eta}) \eta_i^* + \beta_3 |\eta_i|^2 \eta_i$$
 (16)

for the 4 orientation cases to consider we can write the above as a function of magnitude  $\eta_{01}$ ,  $\eta_{10}$ ,  $\eta_{11}$ ,  $\eta_{1i}$ , noting that  $\eta_{01}$  and  $\eta_{10}$  are the same.

$$0 = \alpha \eta_{10} + \beta_1 |\eta_{10}|^2 \eta_{10} + \beta_2 \eta_{10}^2 \eta_{10}^* + \beta_3 |\eta_{10}|^2 \eta_{10}$$
(17)

$$\Rightarrow |\eta_{10}|^2 = \frac{-\alpha}{\beta_1 + \beta_2 + \beta_3} \tag{18}$$

$$0 = \alpha \eta_{11} + 2\beta_1 |\eta_{11}|^2 \eta_{11} + 2\beta_2 \eta_{11}^2 \eta_{11}^* + \beta_3 |\eta_{10}|^2 \eta_{11}$$
(19)

$$\Rightarrow |\eta_{11}|^2 = \frac{-\alpha}{2\beta_1 + 2\beta_2 + \beta_3} \tag{20}$$

$$0 = \alpha \eta_{1i} + 2\beta_1 |\eta_{1i}|^2 \eta_{1i} + \beta_3 |\eta_{1i}|^2 \eta_{1i}$$
(21)

$$\Rightarrow |\eta_{1i}|^2 = \frac{-\alpha}{2\beta_1 + \beta_3} \tag{22}$$

The free energy of these states is:

$$F[\eta_{10}] = \frac{-\alpha^2/(2\beta_1)}{1 + \beta_2/\beta_1 + \beta_3/\beta_1}$$
 (23)

$$F[\eta_{11}] = \frac{-\alpha^2/(\beta_1)}{2 + 2\beta_2/\beta_1 + \beta_3/\beta_1}$$
 (24)

$$F[\eta_{1i}] = \frac{-\alpha^2/(\beta_1)}{2 + \beta_3/\beta_1} \tag{25}$$

For each possible value of  $x = \beta_2/\beta_1$  and  $y = \beta_3/\beta_1$  the system will choose the state with lowest free energy. The conditions on this choice are

choose 10 over 11: y < 0choose 10 over 1i: y < -2xchoose 11 over 1i: x < 0

The regions are mapped out in the figure

### 3a

The diagonalized Hamiltonian is:

$$\mathcal{H} = \sum_{\mathbf{k},s} E_{\mathbf{k}s} b_{\mathbf{k}s}^{\dagger} b_{-\mathbf{k}s} \tag{26}$$

and the b operators are defined through the Bogolioubov transformation:

$$a_{\mathbf{k}s} = u_{\mathbf{k}}b_{\mathbf{k}s} - (i\sigma_y)_{ss'}v_{\mathbf{k}}b_{-\mathbf{k}s'}$$
(27)

which results in the Bogolioubov-de Gennes equations

$$E_{\mathbf{k}s} \begin{pmatrix} u_{\mathbf{k}} \\ v_{\mathbf{k}} \end{pmatrix} = \begin{pmatrix} \xi_{\mathbf{k}} & \Delta^* \\ \Delta & \xi_{\mathbf{k}\bar{s}} \end{pmatrix} \begin{pmatrix} u_{vk} \\ v_{\mathbf{k}} \end{pmatrix}$$
 (28)

In the above definition for b operators I omitted spin indices on the amplitudes in anticipation of the final result which is

$$u_{\mathbf{k}} = \sqrt{\frac{\Delta}{E}} e^{\theta_E/2} \tag{29}$$

$$v_{\mathbf{k}} = \sqrt{\frac{\Delta}{E}} e^{-\theta_E/2} \tag{30}$$

The eigenvalues are

$$E_{\mathbf{k}s} = \sqrt{\xi_{\mathbf{k}}^2 + |\Delta|^2} - \mu_B H s \tag{31}$$

with  $\theta_E = \cosh^{-1}(E/\Delta)$ .

## 3b

After using the form of a operators in terms of the diagonal b's, the spin magnetisation reduces to

$$M = \mu_B \sum_{\mathbf{k}} f_{\mathbf{k}\uparrow} - f_{\mathbf{k}\downarrow} \tag{32}$$

where  $f_{\mathbf{k}s} = (e^{\beta E_{\mathbf{k}s}} + 1)^{-1}$  and  $\beta = 1/T$ . This is good! We can preform the sum over states by using the density of states for a superconductor  $N = N_0 |E|/\sqrt{E^2 - \Delta^2}$  and  $N_0$  is the normal DOS at the Fermi level.

$$M = \mu_B N_0 \int_{\Lambda}^{\infty} dE \frac{E}{\sqrt{E^2 - \Delta^2}} (f_{\mathbf{k}\uparrow} - f_{\mathbf{k}\downarrow})$$
 (33)

The linear response to a field will be  $M_{lr} = \chi(T)H$  and we can Taylor expand the Fermi functions in H  $f_{\mathbf{k}s} \approx f_{\mathbf{k}} + \frac{s\beta}{4cosh^2(\beta E/2)}H$  to get the desired result

$$\chi(T) = (1/2)\mu_B N_0 \beta \int_{\Delta}^{\infty} dE \frac{E}{\sqrt{E^2 - \Delta^2}} \frac{1}{\cosh^2(\beta E/2)}$$
(34)

$$= (1/2)\mu_B N_0 \int_{\beta\Delta}^{\infty} dx \frac{x}{\sqrt{x^2 - (\beta\Delta)^2}} \frac{1}{\cosh^2(x/2)}$$
 (35)

$$= (1/2)\mu_B N_0 \int_{\beta\Delta}^{\infty} dx \sqrt{x^2 - (\beta\Delta)^2} \frac{\tanh(x/2)}{\cosh^2(x/2)}$$
 (36)

and we integrated by parts in the last step

### 3c

The two limits of interest are  $\beta\Delta >> 1$ , and  $\Delta = \Delta_0 = constant$  for low temperature and  $\beta\Delta << 1$  for near  $T_c$  when self consistent solution  $\Delta(T)$  is first appearing as a second order phase transition. The limit near  $T_c$  would not apply if it was a first order transition.

#### $\beta \Delta_0 >> 1$ , Low Temperature

The low T limit allows write  $tanh(x/2) \approx 1$  and  $cosh(x/2) \approx \frac{1}{2}e^{x/2}$ 

$$\chi(T) = 2\mu_B N_0 \int_{\beta\Delta}^{\infty} dx \sqrt{x^2 - (\beta\Delta)^2} e^{-x}$$
 (37)

$$= 2\mu_B N_0 (\beta \Delta)^2 \int_1^\infty dy \sqrt{y^2 - 1} e^{-\beta \Delta y}$$
 (38)

The integral is equivalent to the n=1 modified Bessel function of the second kind.

$$K_1(z) = z \int_{1}^{\infty} dy \sqrt{y^2 - 1} e^{-zy}$$
 (39)

$$\chi(T) = 2\mu_B N_0 \frac{\Delta_0}{T} K_1(\Delta_0/T) \tag{40}$$

 $\beta\Delta << 1$ , Near  $T_c$  Here we employ Leibniz rule to Taylor expand the

integral

$$\int_{\beta\Delta}^{\infty} dx \sqrt{x^2 - (\beta\Delta)^2} \frac{\tanh(x/2)}{\cosh^2(x/2)} \approx \int_{0}^{\infty} dx \quad x \frac{\tanh(x/2)}{\cosh^2(x/2)}$$
(41)

$$- (\beta \Delta)^2 / 2 \int_0^\infty dx \frac{1}{x} \frac{\tanh(x/2)}{\cosh^2(x/2)}$$
 (42)

The first integral is easy and evaluates to 2.

The second integral is not easy, but mathematica can do it and the result is  $-28\zeta'(-2) \approx .852557$ . At this point we can also use the limiting expression for  $\Delta(T)$  near  $T_c$  which was written in class.

$$\Delta(T) = \sqrt{\frac{2}{7|\zeta'(-2)|}} T_c \sqrt{1 - T/T_c} \tag{43}$$

Now we can finally write the two limits using normal state susceptibility  $\chi_N = \mu_B N_0$ 

$$\frac{\chi^{T_c}(T)}{\chi_N} = 1 - 2\frac{1 - T/T_c}{(T/T_c)^2} \tag{44}$$

$$\frac{\chi^0(T)}{\chi_N} = 2\frac{\Delta_0}{T} K_1(\Delta_0/T) \tag{45}$$

The zero temperature expansion is a very slow growing function of T, and plots show that  $\frac{\chi^0(T)}{\chi_0} < 0.04$  for  $T < 0.2\Delta_0$ . This is expected from the form of the magnetisation in the beginning of part 3b combined with the gapped density of states.

Near  $T_c$  the normal state  $\chi_N$  is recovered and the temperature dependence of the decrease in  $\chi$  is due to the onset of the order parameter  $\Delta(T)$ . Looking at plots, one can see that the  $\chi$  is quickly reduced to near zero for  $T/T_c \approx 0.75$ 

FIGURES ATTACHED