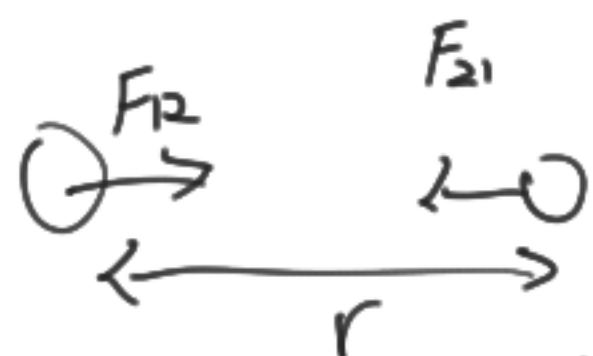


G field & E field

G field
due to
mass



$$|F_{12}| = |F_{21}| = \frac{GM_1m_2}{r^2}$$

Energy: $\frac{F \cdot r}{\frac{GM_1m_2}{r}}$

Definition of g-field:

$$g = \frac{F_g}{m} = \frac{GM}{r^2}$$

unit: Nkg^{-1}

Gravitational force exerted per
kilogram of the object's mass



Electrostatic force



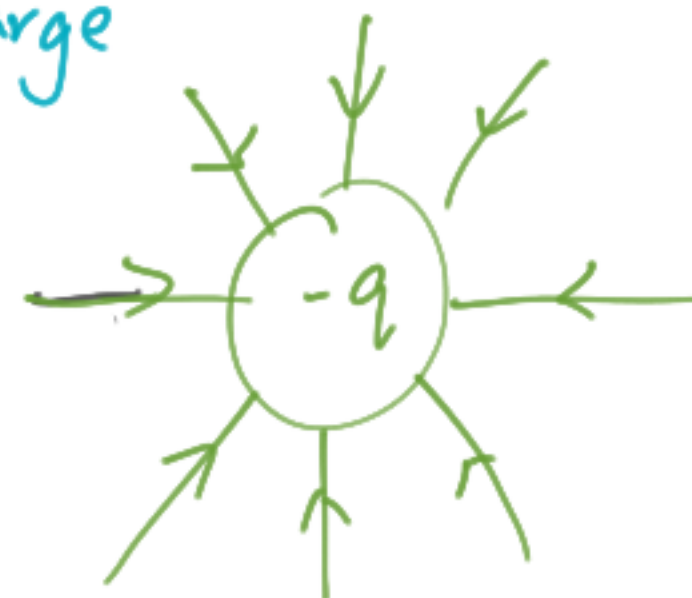
$$F = \frac{1}{4\pi\epsilon_0} \frac{q_1q_2}{r^2}$$

are the forces exerted on both
particles equal?

E field:

$$\frac{F}{q} = \frac{1}{4\pi\epsilon_0} \frac{q_1}{r^2}$$

Electric force exerted on a positive test
charge per unit charge



line density
= field strength

Gravitational potential energy

$$U = \left[-\frac{GmM}{r} \right] \leftarrow \int \frac{GmM}{r^2} dr$$

$r \rightarrow \infty \quad U \rightarrow 0 \quad (\text{Max})$

$\oplus \xrightarrow{1m} \ominus$

$$V_i = 0 \quad V_f = \frac{1}{4\pi\epsilon_0} \frac{-q_1q_2}{1}$$

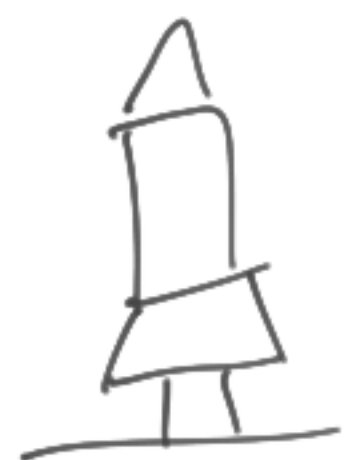
$$-\Delta U = \int F \cdot dr$$

$\frac{GmM}{r^2} = F$

$W = F \cdot s$

Useful in energy calculations!

Example: we are firing a rocket from Earth, find v_{min} required s.t. the rocket escapes the earth's gravitational field.



$GPE \uparrow \quad r \uparrow$

$r \uparrow \quad F \downarrow \quad F \propto \frac{1}{r^2}$

$$U = \left[-\frac{GmM}{r} \right]$$

$r \uparrow \quad U \uparrow \quad F \downarrow$

$$U = -\frac{GmM}{r}$$

$KE \rightarrow GPE$

$$\frac{1}{2}mv^2 = \left(0 - \left(-\frac{GmM}{R_e} \right) \right)$$

$$\frac{1}{2}mv^2 = \frac{GmM}{R_e} \quad v = \sqrt{\frac{2GmM}{R_e}}$$

larger

Electrical potential energy

$$U_E = \frac{1}{4\pi\epsilon_0} \frac{q_1q_2}{r} \leftarrow \int \frac{1}{4\pi\epsilon_0} \frac{q_1q_2}{r^2} dr$$

$\oplus \ominus \rightarrow (-)$ $\oplus \oplus \rightarrow (+)$ $\ominus \ominus \rightarrow (+)$

Lower U_E is always more stable.

$$\frac{-k}{4\pi\epsilon_0} \frac{1}{r} \rightarrow -\infty$$

$$\frac{k}{4\pi\epsilon_0} \frac{1}{r} \rightarrow 0$$

$$\frac{k}{4\pi\epsilon_0} \frac{1}{r} \rightarrow 0$$

Gravitational potential

$$\Phi = \frac{U}{m} \text{ (J kg}^{-1}\text{)}$$

Electrical potential

$$V = \frac{U}{q} \text{ (V / J C}^{-1}\text{)}$$

$500 = \frac{\Delta U}{q_1}$

$\Delta U = U_f - U_i \quad U = -\frac{GmM}{r}$

$\Delta \Phi = 3 \text{ J/kg}$



$F_e = mg$

$qE = mg \quad E = \frac{mg}{q} \quad v = \frac{mgd}{q}$

$m = 10 \text{ kg}$



Q9.

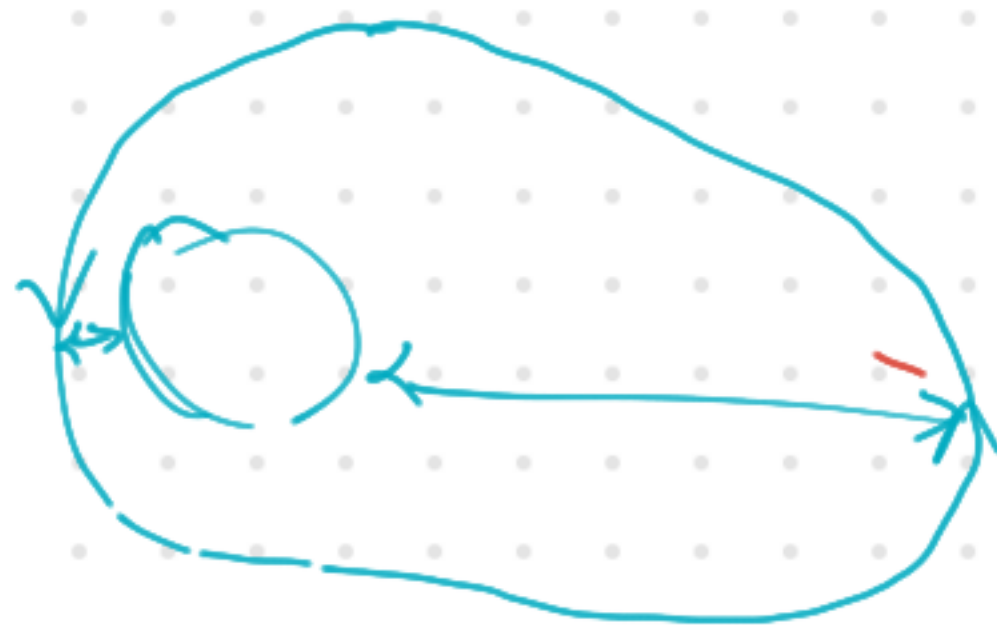
Astronomers observing stars at the centre of our galaxy have suggested that many of them are orbiting a supermassive black hole. The mass of this black hole is $9.2 \times 10^{36} \text{ kg}$.

The star S0-2 is in a highly elliptical orbit around the position of the black hole.

At its point of closest approach, S0-2 is at a distance of $1.8 \times 10^{13} \text{ m}$ from the centre of the black hole.

At the most distant point of its orbit, S0-2 is $2.7 \times 10^{14} \text{ m}$ from the black hole.

(i) Show that the change in gravitational potential between the closest and most distant points in this orbit is about $3 \times 10^{13} \text{ J kg}^{-1}$.



$$U = -\frac{GMm}{r}$$

$$\Phi = \frac{U}{m} \rightarrow$$

$$\Phi = -\frac{GM}{r}$$

$$\begin{aligned} \Delta\Phi &= \Phi_f - \Phi_i \\ &= -\frac{GM}{r_f} - \left(-\frac{GM}{r_i}\right) \\ &= -\frac{GM}{r_f} + \frac{GM}{r_i} \\ &\approx 3.18 \cdot 10^{13} \text{ J kg}^{-1} \end{aligned}$$

(ii) At its point of closest approach, the star is travelling at a speed of $8.1 \times 10^6 \text{ m s}^{-1}$. Calculate the speed of S0-2 at the furthest point in its orbit using the change in gravitational potential.
mass of S0-2 = $2.4 \times 10^{31} \text{ kg}$

(3)

Conservation of energy!

$$\begin{aligned} \Delta KE &= -\Delta U \\ \frac{1}{2} m (v_f^2 - v_o^2) &= -m \Delta\Phi \\ v_o^2 - v_f^2 &= 2\Delta\Phi \\ v_f &= 1.042 \cdot 10^6 \text{ ms}^{-1} \end{aligned}$$

Speed =

Q2.

// grav field strength

The acceleration of free fall at the surface of the Earth is 9.81 m s^{-2} .
The mass of the Earth is M and the diameter of the Earth is D .

Which of the following gives the acceleration of free fall, in m s^{-2} , at the surface of a planet with diameter $\frac{D}{2}$ and mass $\frac{M}{9}$?

☐ A $\frac{9.81 \times 2}{9}$

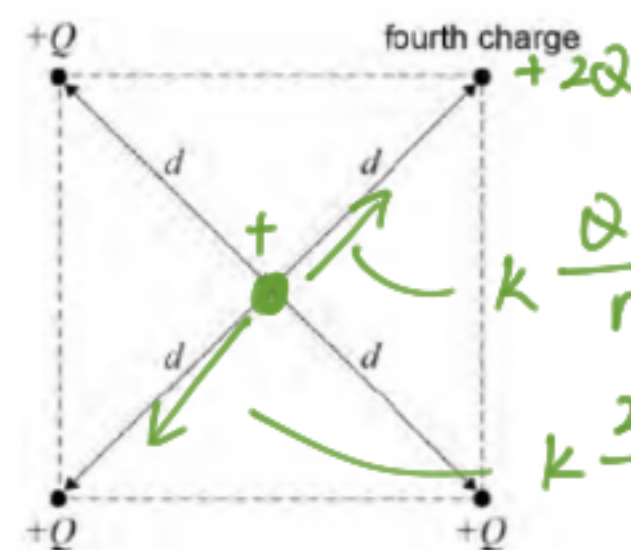
☒ B $\frac{9.81 \times 4}{9}$

☐ C $\frac{9.81 \times 2}{3}$

☐ D $\frac{9.81 \times 9}{4}$

$$a = \frac{GM}{r^2} = \frac{F_g}{m}$$

8. Four positive charges are fixed at the corners of a square as shown.



$$V = \frac{1}{4\pi\epsilon_0} \frac{q}{r} = \frac{V}{q/r}$$

$$= \frac{1}{4\pi\epsilon_0} \frac{q_1}{r} + \frac{1}{4\pi\epsilon_0} \frac{q_2}{r}$$

The total potential at the centre of the square, a distance d from each charge, is $\frac{5Q}{4\pi\epsilon_0 d}$

Three of the charges have a charge of $+Q$

What is the magnitude of the fourth charge?

- A $-\frac{7Q}{4}$ ☐
- B Q ☐
- C $\sqrt{2}Q$ ☐
- D $2Q$ ☒

(Total 1 mark)

$$V = \frac{1}{4\pi\epsilon_0} \sum \frac{q_i}{r_i}$$

$$V = \frac{1}{4\pi\epsilon_0} \left(\frac{Q}{d} + \frac{Q}{d} - \frac{Q}{d} + \frac{q'}{d} \right)$$

$$= \frac{Q}{4\pi\epsilon_0 d} + \frac{1}{4\pi\epsilon_0} \frac{q'}{d} = \frac{5Q}{4\pi\epsilon_0 d}$$

$$\frac{q'}{4\pi\epsilon_0 d} = \frac{4Q}{4\pi\epsilon_0 d}$$

$$\frac{1}{4\pi\epsilon_0} \left(\frac{2Q}{d^2} \right) - \frac{1}{4\pi\epsilon_0} \left(\frac{Q}{d^2} \right)$$

$$|\vec{E}| = \frac{Q}{4\pi\epsilon_0 d^2} \text{ N C}^{-1} / \text{ V m}^{-1} \quad \vec{E} = \frac{V}{d}$$

4.

The figure shows a moon of mass m in a circular orbit of radius r around a planet of mass M where $m \ll M$.

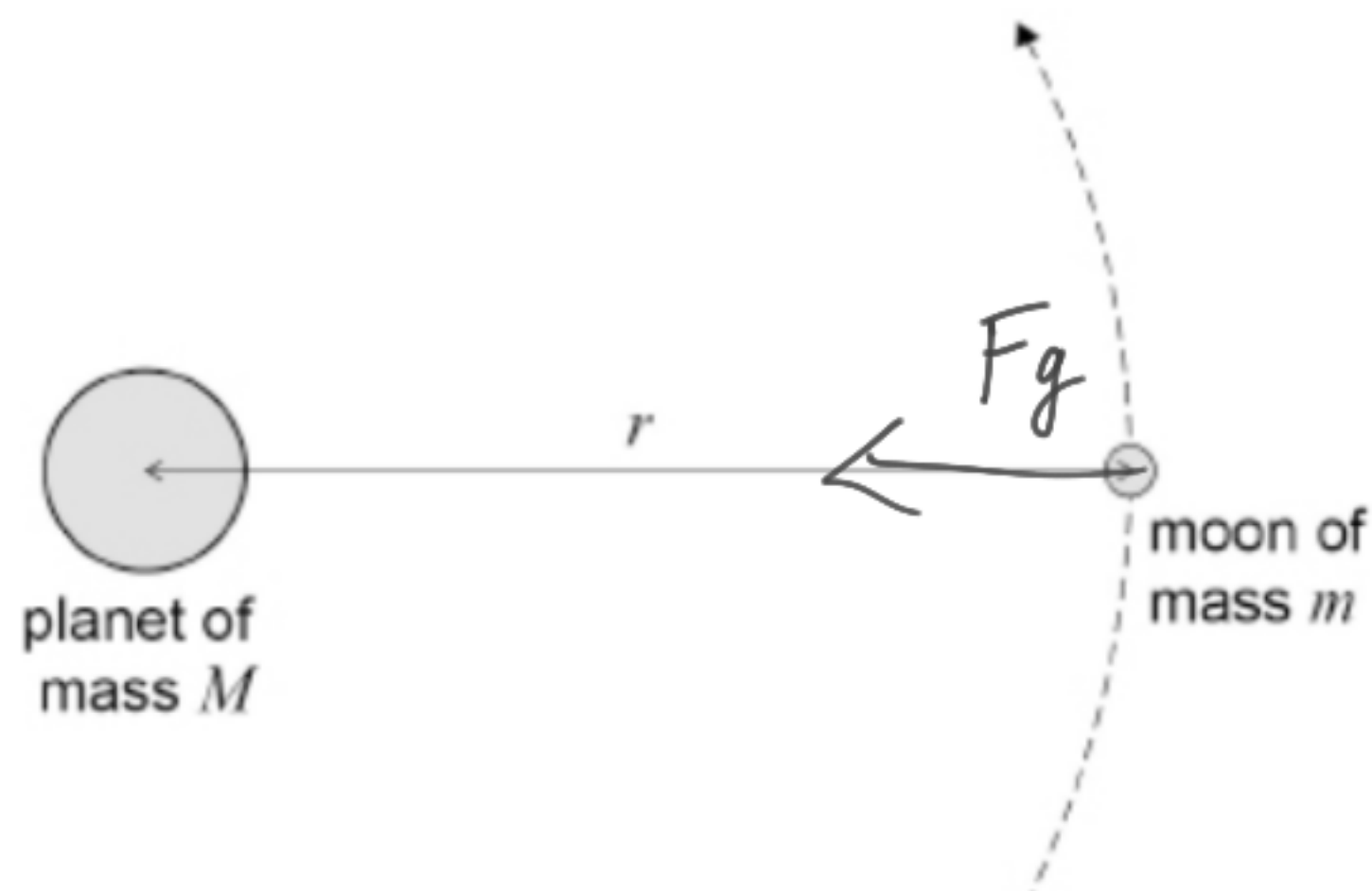


Table 1 gives data for two of the moons of the planet Uranus.

Table 1

Name	T / days	r / m
Miranda	1.41	1.29×10^8
Umbriel	4.14	X

The moon has an orbital period T .

T is related to r by

$$T^2 = kr^3$$

where k is a constant for this planet.

(a) Show that $k = \frac{4\pi^2}{GM}$

$$\frac{GM}{r^2} = r\omega^2$$

$$\frac{1}{\omega^2} = \frac{1}{GM} r^3 \rightarrow$$

Kepler's 3rd law

$$T^2 = \frac{4\pi^2}{GM} r^3 \quad T^2 \propto r^3 \rightarrow \left[\frac{T^2}{r^3} \right] = \frac{4\pi^2}{GM} = \text{constant}$$

(b) Calculate the orbital radius **X** of Umbriel.

$$\frac{1.41^2}{(1.29 \times 10^8)^3} = \frac{4.14^2}{X^3}$$

$$X^3 = 2.65 \cdot 10^8 \text{ m}$$

- (c) Calculate the mass of Uranus.

$$\frac{T^2}{r^3} = \frac{4\pi^2}{GM}$$

$$M = \frac{4\pi^2 r^3}{GT^2}$$

mass = $8.60 \cdot 10^{25}$ kg

Table 2 gives data for three more moons of Uranus.

Table 2

Name	Mass / kg	Diameter / m
Ariel	1.27×10^{21}	1.16×10^6
Oberon	3.03×10^{21}	1.52×10^6
Titania	3.49×10^{21}	1.58×10^6

- (d) Deduce which moon in **Table 2** has the greatest escape velocity for an object on its surface.

Assume the effect of Uranus is negligible.

Planet escape velocity, given r & M

$$\frac{1}{2}mv_{\min}^2 = G\frac{mM}{r}$$

$$v_{\min} = \sqrt{\frac{2GM}{r}}$$

Titania has largest v_{escape}

(3)

- (e) A spring mechanism can project an object vertically to a maximum height of 1.0 m from the surface of the Earth.

$$E_s = mgh$$

Determine whether the same mechanism could project the same object vertically to a maximum height greater than 100 m when placed on the surface of Ariel.

$$\Delta h \ll r_a$$

$$g_a = \frac{GM}{r^2}$$

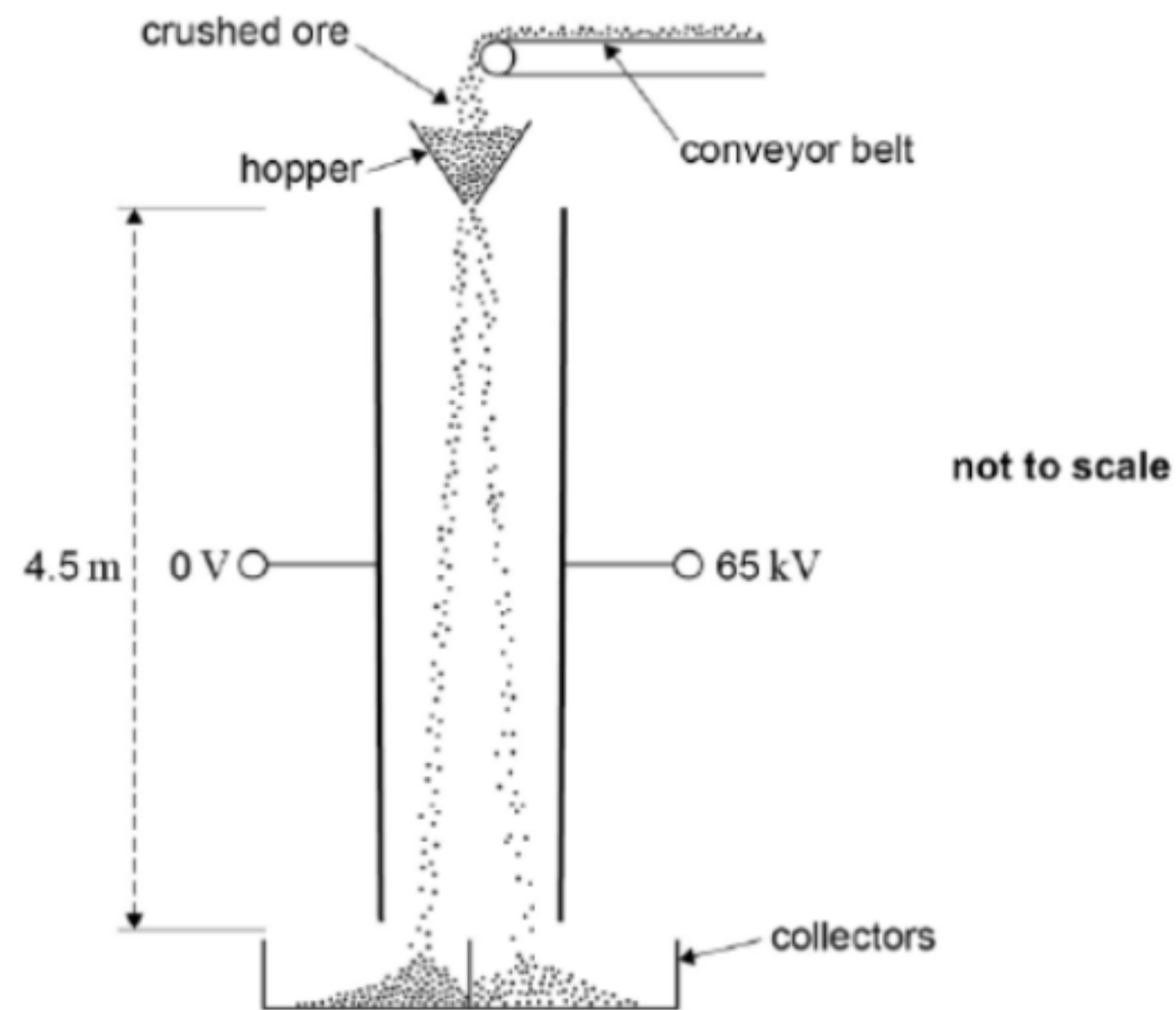
$$g_a = 0.25 \text{ ms}^{-2}$$

$$m g_a h_a = m g_e h_e$$

$$h_a = 40 \text{ m} < 100 \text{ m} \therefore \text{No!}$$

10.

The figure below shows a system that separates two minerals from the ore containing them using an electric field.



The crushed particles of the two different minerals gain opposite charges due to friction as they travel along the conveyor belt and through the hopper. When they leave the hopper they fall 4.5 metres between two parallel plates that are separated by 0.35 m.

- (a) Assume that a particle has zero velocity when it leaves the hopper and enters the region between the plates.

Calculate the time taken for this particle to fall between the plates.

- (b) A potential difference (pd) of 65 kV is applied between the plates.

Show that when a particle of specific charge $1.2 \times 10^{-6} \text{ C kg}^{-1}$ is between the plates its horizontal acceleration is about 0.2 m s^{-2} .

- (c) Calculate the total horizontal deflection of the particle that occurs when falling between the plates.

horizontal deflection = _____m

- (d) Explain why the time to fall vertically between the plates is independent of the mass of a particle.

(2)

- (e) State and explain **two** reasons, why the horizontal acceleration of a particle is different for each particle.

(4)

(Total 12 marks)

5. A rocket carrying an artificial satellite is launched vertically from the Earth. When the rocket is at a certain height from the Earth's surface, it expels 2.60×10^3 kg of gas per second with a certain speed v towards the Earth's centre. As a result, an average thrust of 5.20×10^6 N is produced. Neglect air resistance.

(a) (i) Assuming that the speed of the rocket is negligible, estimate v . (2 marks)

(ii) At a certain instant, the total mass of the rocket and the artificial satellite is $3.60 \times 10^5 \text{ kg}$ while the acceleration due to gravity at the rocket's position is 8.56 m s^{-2} . Estimate the acceleration a of the rocket at this position. (2 marks)

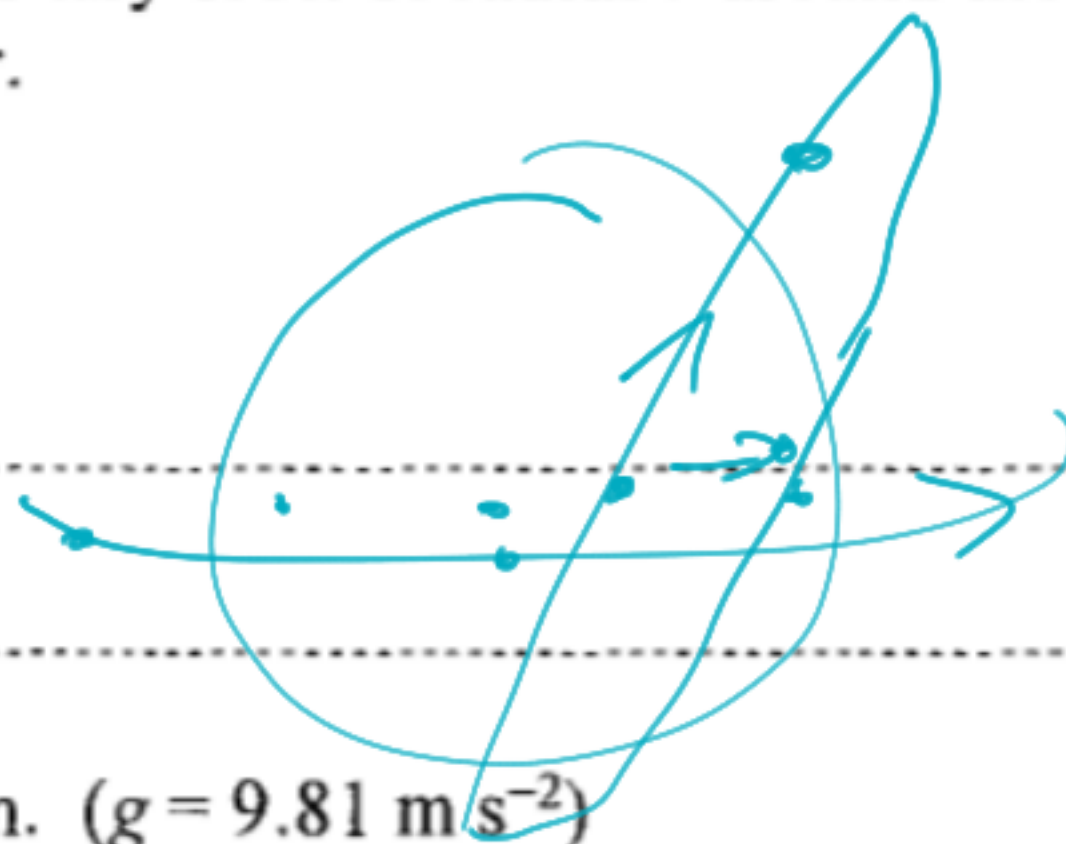
(iii) Suppose the rocket keeps expelling gas at the same rate for a few seconds. Would the rocket's acceleration increase, decrease or remain unchanged in that duration? Explain. (2 marks)

*(b) The artificial satellite is put in the geostationary orbit of radius r around the Earth. It appears to be always stationary above an observer at the equator.

(i) State the period of this satellite.

(1 mark)

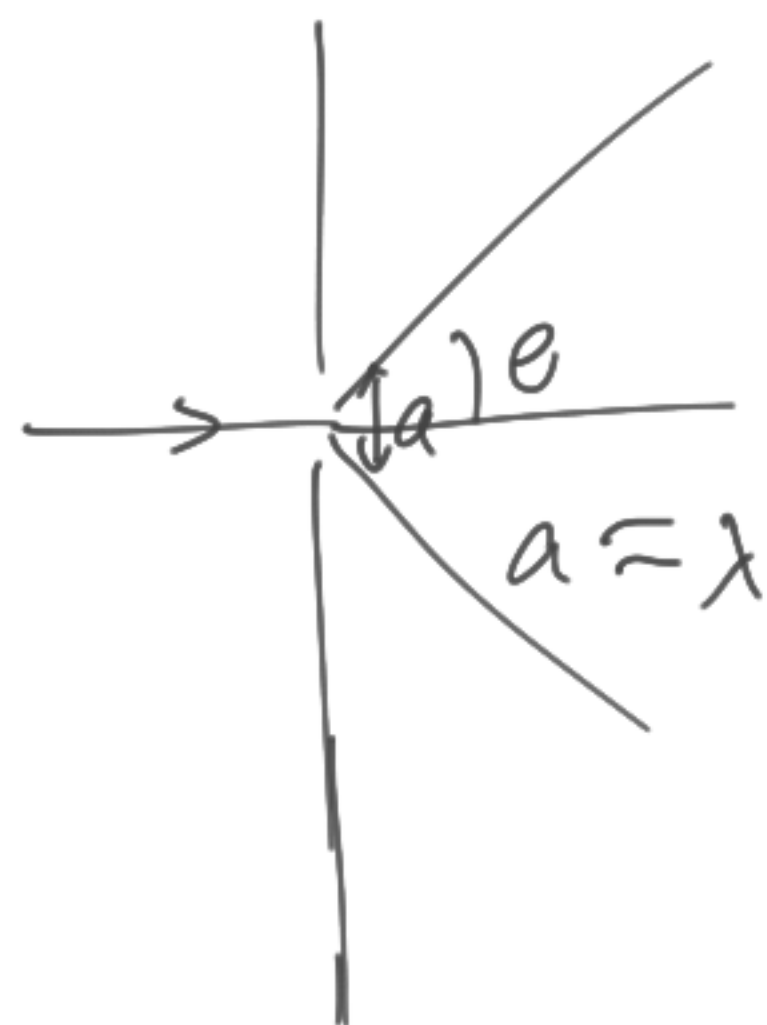
1 day



(ii) Show that r is approximately 42000 km. ($g = 9.81 \text{ m s}^{-2}$)

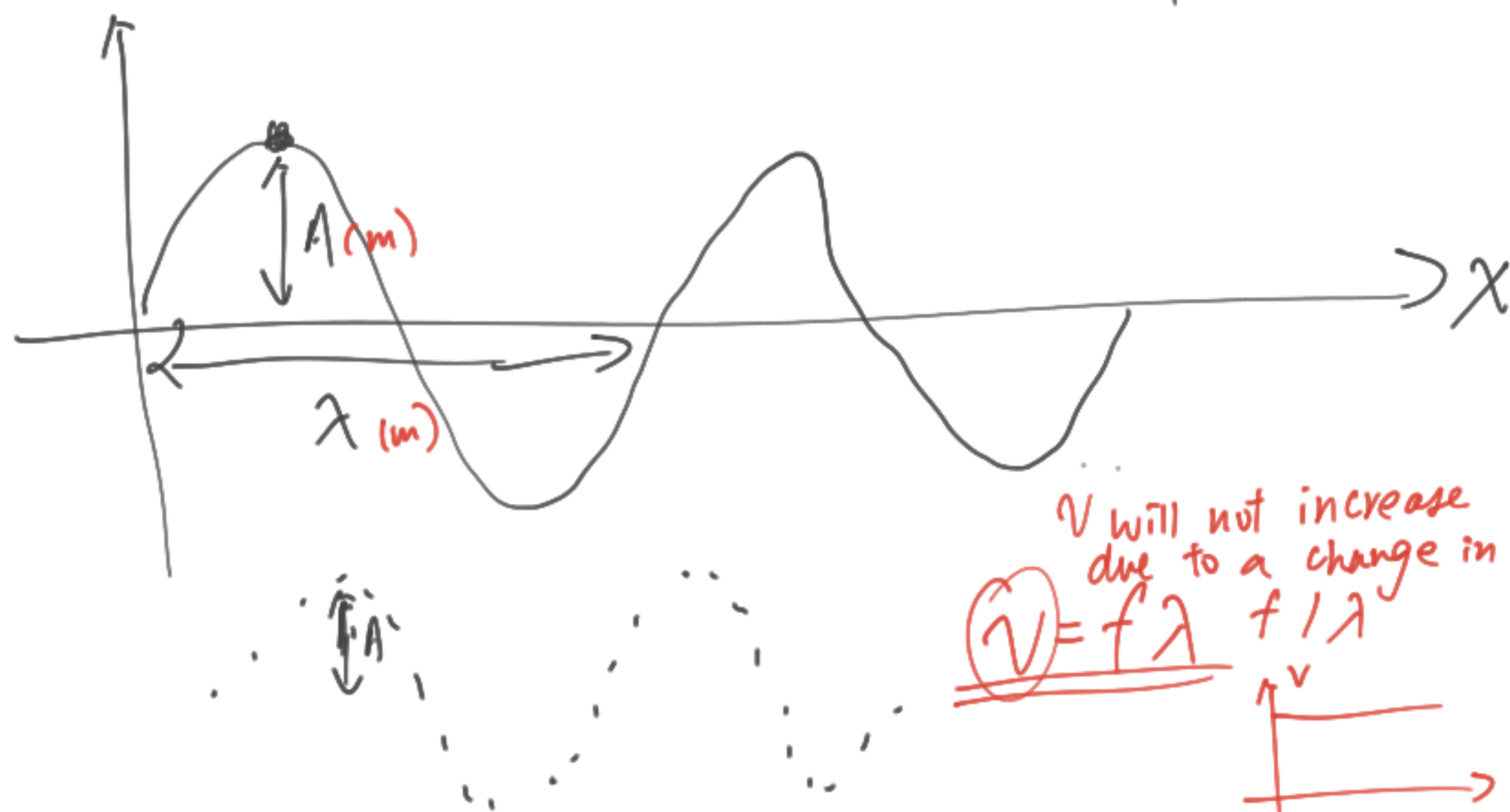
(2 marks)

Given: radius of the Earth = $6.37 \times 10^6 \text{ m}$



$$a \gg \lambda$$

Diffraction
is significant



v will not increase
due to a change in

$v = f\lambda$

$f \propto 1/\lambda$

v

λ



$$f = \frac{1}{T} \text{ (Hz) (s}^{-1}\text{)}$$

\hookrightarrow how many complete wave form
per second