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Program for finding Schrödinger's Egtn Wexternal forces.

The Schrödinger Equation with External Forces

We shall derive Schrödinger's Wave Egth for the wavefon 4(15,t) of a QM particle of mass m moving in the presence of external forces. Program is:

- Start from a generalized version of the free particle wave egts for 4;
- use packet solutions for the wave for 4;
- · define particle (packet) motion via expectation values;
- impose condition that the particle's motion obeys the Correspondence Principle, in the sense that -- on average -- m moves per Newton's Laws.
- 1) We work in 1D for simplicity (generalization to 3D is Nobrious). Write a generalization of Schrödinger's free particle extr...

$$\rightarrow i \frac{\partial}{\partial t} \psi(x,t) = \mathcal{H} \psi(x,t), \quad \forall y = \frac{1}{2m} p^2 + Q,$$

where: p=-ikopx, per Eq. (36) above;

and : Q = Q(x, p; t), due to external forces, is to be found.

For free particles, Q=0, and : $i\hbar \frac{\partial}{\partial t} \psi = -\frac{\hbar^2}{2m} (\partial^2/\partial x^2) \psi$. Ultimately, we will show that Q= potential energy V associated with the external forces.

Now, w.n.t. packet-type solutions to Eq.(38), <u>define</u> m's momentum by $\rightarrow \langle p \rangle = m \frac{d}{dt} \langle x \rangle, \langle \rangle \Rightarrow \text{expectation value w.n.t.} \Psi. \tag{39}$

This is a Correspondence Principle-type statement... on average, the particle's momentum will correspond to the Newtonian form: $p = m \frac{d}{dt} \times$. The definition of m's momentum per Eq. (39) will impose a restriction on our to-be-found Q... viz-Q is <u>real</u>: $Q^* = Q$. Now we show that.

In Eq. (39), colculate (all
$$\int = \int_{-\infty}^{\infty}$$
)...

$$\frac{d}{dt}\langle x \rangle = \frac{d}{dt} \int \psi^* \{x\} \psi \, dx = \int \left[\left(\frac{\partial \psi^*}{\partial t} \right) x \psi + \psi^* x \left(\frac{\partial \psi}{\partial t} \right) \right] dx; \qquad (40)$$

but $\Omega : \frac{\partial \psi}{\partial t} = -\frac{i}{\hbar} \left[-\frac{\hbar^2}{2m} \frac{\partial^2}{\partial x^2} + Q \right] \psi, \quad \text{and} \quad (2) : \frac{\partial \psi^*}{\partial t} = +\frac{i}{\hbar} \left[-\frac{\hbar^2}{2m} \frac{\partial^2}{\partial x^2} + Q^* \right] \psi^*,$

Solve $\left[\frac{d}{dt} \langle x \rangle = \frac{i \hbar}{2m} \int \left[\psi^* x \left(\frac{\partial^2 \psi}{\partial x^2} \right) - \left(\frac{\partial^2 \psi^*}{\partial x^2} \right) x \psi \right] dx + V,$

where $: V = \frac{i}{\hbar} \int (Q^* - Q) \psi^* x \psi \, dx.$

This improbable-looking mess must be made to fit the definition in Eq. (39), viz: $\frac{d}{dt}\langle x \rangle = \frac{1}{m}\langle p \rangle = \frac{1}{m}\int \psi^*\{-i\hbar\partial/\partial x\}\,\Psi\,dx$. Evidently, we must simplify (41). We do that by twice partial-integrating the second term RHS in (41), imposing that 24 & $34/\partial x$ transh at ∞ ...

$$= -\left[qh_{\star} \frac{\partial x}{\partial x}(xh) qx = -h_{\star} \frac{\partial x}{\partial x}(xh)\Big|_{t\infty}^{-\omega} + \left[h_{\star} \frac{\partial x}{\partial x}(xh)qx\right] + \left[h_{\star} \frac{\partial x}{\partial x}(xh)q$$

$$= + \int \psi^* \frac{\partial}{\partial x} \left(\psi + x \frac{\partial \psi}{\partial x} \right) dx = 2 \int \psi^* \frac{\partial \psi}{\partial x} dx + \int \psi^* x \frac{\partial^2 \psi}{\partial x^2} dx.$$

From the last expression in (42), we can form the integral in (41), viz ...

$$\rightarrow \int \left[\psi *_{\chi} \left(\frac{\partial^{2} \psi}{\partial x^{2}} \right) - \left(\frac{\partial^{2} \psi *}{\partial x^{2}} \right) \chi \psi \right] dx = -2 \int \psi * \left(\frac{\partial \psi}{\partial x} \right) dx. \tag{43}$$

Use of this result in Eq. (41) allows us to write ...

$$\frac{d}{dt}\langle x \rangle = \frac{1}{m} \int \psi^* \{-i\kappa \frac{\partial}{\partial x}\} \psi dx + \mathcal{V} = \frac{1}{m} \langle p \rangle + \mathcal{V};$$

$$\frac{\partial}{\partial x} \text{ if } \frac{\langle p \rangle = m \frac{d}{dt} \langle x \rangle}{dt}, \text{ then } \mathcal{V} = \frac{i}{k} \int (Q^* - Q) \psi^* x \psi dx = 0,$$

$$\text{for all } \psi \dots \text{ possible only if } Q^* = Q, \text{ i.e. } \underline{Q} \text{ is real. } \mathbf{QED}. \tag{44}$$

2) Now we impose Newton II on the motion, i.e. ma = F, or $\frac{d}{dt} p = F$. Since we are working with expectation values, we need to calculate...

by partial integrations $\psi * \frac{\partial}{\partial x} \left(\frac{\partial^2 \psi}{\partial x^2} \right) = -\left(\frac{\partial \psi^*}{\partial x} \right) \frac{\partial}{\partial x} \left(\frac{\partial \psi}{\partial x} \right) = +\left(\frac{\partial^2 \psi^*}{\partial x^2} \right) \frac{\partial \psi}{\partial x} \Rightarrow \text{ in Eq.(9) is } \equiv 0.$

 $\frac{d}{dt}\langle p \rangle = \int \psi^* \{-\frac{\partial Q}{\partial x}\} \psi dx = \langle -\partial Q/\partial x \rangle. \tag{46}$

This result provides substantial information on what the "unknown" Q can be. Newton II can be written: $\frac{d}{dt} p = F = -\frac{\partial V}{\partial x}$, where V is the potential for [V = V(x,t)] in general [V = V(x,t)] from which the force [V = V(x,t)] in general [V = V(x,t)] from which the force [V = V(x,t)] in an expectation Comparison with (46) shows that: [V = V(x,t)] from which the force [V = V(x,t)] in an expectation Value sense, so that: [V = V(x,t)] space-independent cost. We take the cost [V = V(x,t)] would be uniform everywhere at a given time, and would just set a relative zero of energy. Thus we claim...

[4/ d (p) = (F), then/ Q = V = system potential energy. (47)

This is enough to establish the 1D Schrödinger Eqta for a particle on in an external field as...

$$i\hbar \frac{\partial}{\partial t} \psi = 46 \psi$$
, $^{1/4} 46 = \frac{b^2}{2m} + V$ $V = P.E. of field$. (48)

Remarks on Schrödinger's external field extr. H= system Flamiltonian. Sch. (20

REMARKS On Schrödinger's external field wave est, Eq. (48).

1. Generalization from 1D to 3D is simple ...

$$46(1D) = \frac{1}{2m} p^{2} + V(x_{i}t) \rightarrow 46(3D) = \frac{1}{2m} p^{2} + V(r_{i}t), \quad p = -i\hbar \nabla;$$

$$54y \quad i\hbar \frac{\partial}{\partial t} \psi(r_{i}t) = \left[-\frac{\hbar^{2}}{2m} \nabla^{2} + V(r_{i}t)\right] \psi(r_{i}t) \quad SCHRÖDINGER'S WAVE EQTN \\ \underline{(\dot{m} 3D, particle m in field)}. \quad (49)$$

2: A continuity extra can be written for Eq. (49), just as for a free particle ...

if
$$\frac{\partial \psi}{\partial t} = -\frac{k^2}{2m} \nabla^2 \psi + V \psi \leftarrow \text{multiply on left by } \psi^*$$
and
$$-i \frac{\partial \psi^*}{\partial t} = -\frac{k^2}{2m} \nabla^2 \psi^* + V \psi^* \leftarrow \text{multiply on left by } \psi$$
then subtract

$$\implies i \, h \left[\psi^* \left(\frac{\partial \psi}{\partial t} \right) + \left(\frac{\partial \psi^*}{\partial t} \right) \psi \right] = -\frac{h^2}{2m} \left[\psi^* (\nabla^2 \psi) - (\nabla^2 \psi^*) \psi \right];$$

3. The addition of the P.E. term in V in Eq. (49) has not affected the form of the containinty extra; p & I in (50) we the same form as in Eq. (19) for the free particle (only 4 will change its content when V + 0). As before, we get conservation of probability from the continuity extra...

$$\rightarrow \frac{\partial}{\partial t} \int_{\Omega} \rho \, d^3r = \frac{\partial}{\partial t} \int_{\Omega} |\Psi|^2 d^3r = -\int_{\Omega} \nabla \cdot \mathbf{J} \, d^3r = -\oint_{\Omega} \mathbf{J} \cdot d\mathbf{S} \rightarrow 0, \qquad (51)$$

for 4 & V4 that vanish at 00. Then: [14/2 d3r = time indept. crist = 1.

4. We note that Ho in Eq. (49) is the total system <u>Hamiltonian</u> (operator), since:

(46) = ⟨p²/2m⟩ + ⟨V⟩ = K.E. + P.E. = ⟨E⟩, total system energy (in an expectation value sense). Then, in the same expectation value sense...

[464 = it 24 => ⟨46⟩ = ⟨E⟩ = ⟨it 0/0t⟩, for exp. values w.a.t. 4;

[50] m¹s total energy E → Eq. = it 0/0t, as suggested in Eq.(4), p. Sch. 2...

3) For QM systems where the potential energy V is independent of time, a simplification of the Schrödinger Eqt., Eq. (49) is possible. As follows...

Suppose V = V(r) is time-indept. Let: $\Psi(r,t) = u(r)f(t)$.

Then: $i \frac{\partial \Psi}{\partial t} = \left(-\frac{h^2}{2m} \nabla^2 + V\right) \Psi$, becomes a <u>separable PDE</u>, as: $\frac{i \hbar}{f} \cdot \frac{d f}{d t} = \frac{1}{u} \left(-\frac{h^2}{2m} \nabla^2 + V\right) u = W$, a const indept. of $r \neq t$.

for of t only

for of r only

arbitrary enst

... so: $(i \frac{h}{f}) \frac{d f}{d t} = W$ has solution: $f(t) = Ae^{-(i/h)}Wt$...

and $W(r,t) = Au(r) \exp[-ii/h)Wt$, is a solution for wavefor Ψ ,

With $W(r,t) = Au(r) \exp[-ii/h)Wt$, is a solution const.

We can choose the multiplicative cost A in f(t) to be A=1; then 4 & 21 have be same normalization, as: $\int_{\infty} |\psi|^2 d^3r = \int_{\infty} |u|^2 d^3r = 1$.

Son W= E = total system energy, const in time (i.e. $\frac{dW}{dt} = 0 \Rightarrow E = const$). (54)

Summarizing these results, we can state ...

If the P.E. V = V(r) is time-independent, then a particular solution to Schrödinger's Eqtra [Eq. (49)] is: \frac{\psi(r,t) = u(r) exp(-\frac{i}{h}Et)}{h}, \frac{m}{h}Eis} \text{ the (constant) total system energy (E=\langle K.E.+P.E.\rangle), and u(r) obeys \frac{[-(\frac{r}{2}/2m)\sigma^2 + V(r)]u(r) = Eu(r)}{(system energy E = constant).} \frac{\sigma \text{System energy} E = constant)}{(\sigma \text{U}(r)) \text{ is called a "Stationary state Solution."}}