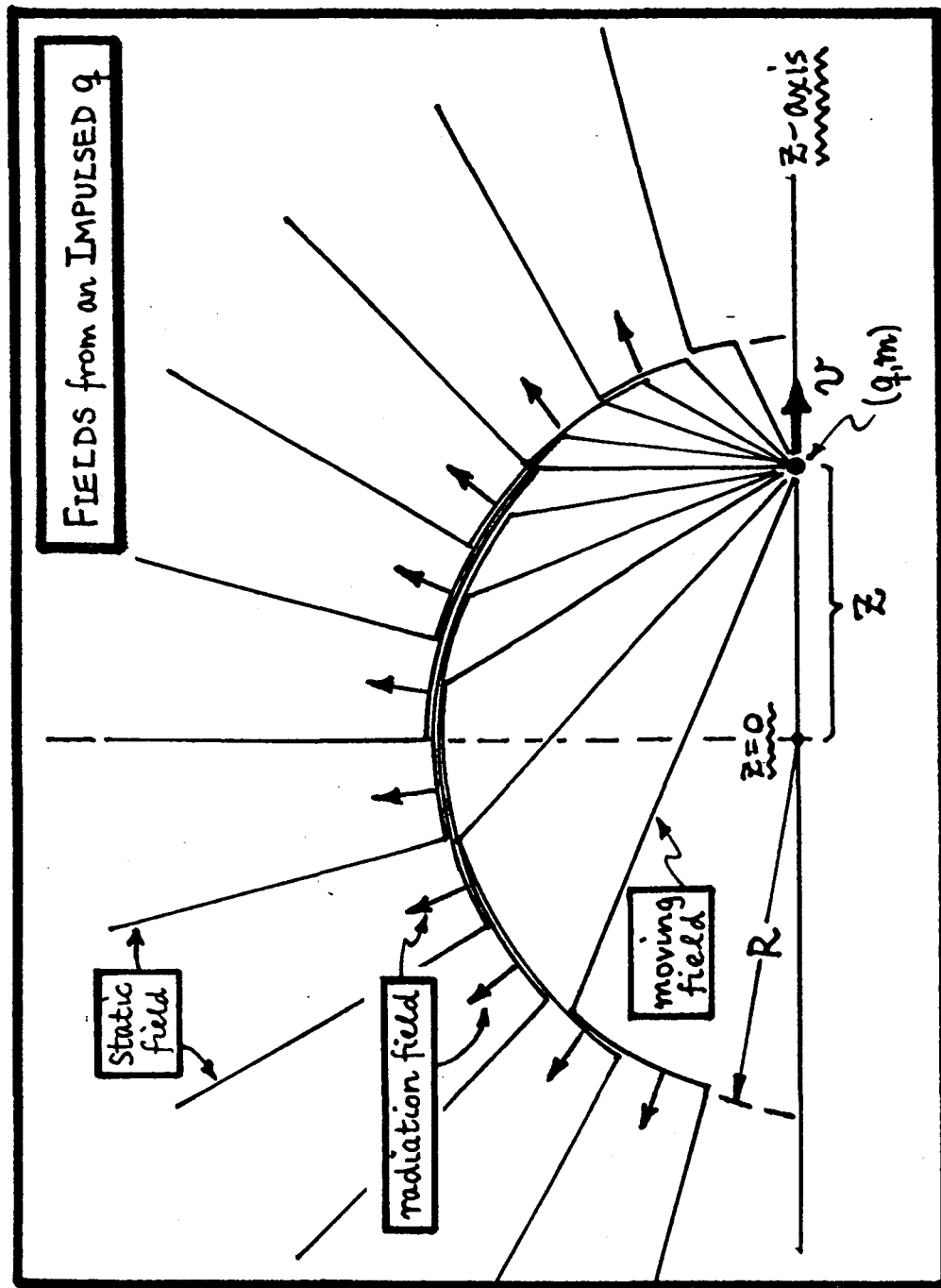


RADIATION REACTION

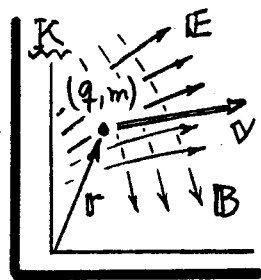


Incompleteness of the Lorentz Law as an Equation-of-Motion.

RR(1)

The Equation-of-Motion for a Single Charge q . Radiation Reaction.

- 1) For a classical (spinless) particle of charge q & mass m moving in lab (reference) frame \underline{K} at velocity $\underline{v} = d\underline{r}/dt$ (t is \underline{K} 's time), through external electric & magnetic fields \underline{E} & \underline{B} , an eqn-of-motion is provided by the Lorentz Force Law [Jk² Eq. (11.124), or Eq. (1) of $\phi 520$ notes p. COV1]:



$$\underline{\underline{F = dp/dt = q(\underline{E} + \frac{\underline{v}}{c} \times \underline{B})}}. \quad (1)$$

Here $\underline{p} = \gamma m \underline{v}$ is the particle's relativistic 3-momentum ($\gamma = \frac{1}{\sqrt{1-\beta^2}}$, $\beta = v/c$), and \underline{E} & \underline{B} are specified at \underline{K} cds \underline{r} & t . In addition to (1), we have...

$$\underline{\underline{\frac{dK}{dt} = q \underline{E} \cdot \underline{v}}} \quad \sqrt{\text{w}} \quad K = (\gamma - 1)mc^2 = \text{particle's relativistic K.E.} \quad (2)$$

Eq. (2) is the work-energy theorem for (q, m) ; it tells how fast the particle is gaining the mechanical energy K by action of the \underline{E} -field. Eqs. (1) & (2) can be combined into a single "manifestly covariant" force law [Jk² Eq. (11.144), or Eqs. (7) & (26) of $\phi 520$ notes pp. COV 2 & 9]. For $F^{\alpha\beta}$ the field tensor...

$$\boxed{m \dot{u}^\alpha = (q/c) F^{\alpha\beta} u_\beta}, \quad u^\alpha = 4\text{-velocity} = \gamma(c, \underline{v}). \quad (3)$$

The "dot" here means $d/d\tau$, $\sqrt{\text{w}} \quad \tau = \text{particle's proper time}$ ($d\tau = \frac{1}{\gamma} dt$).

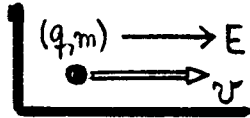
A point to be made at the outset is that all the fields that appear in these eqns-of-motion are external fields, created in \underline{K} by sources outside (q, m) ... nowhere in Eqs. (1)-(3) is there mention of (q, m) 's own (self) fields. Should there be? (q, m) 's self \underline{E} -field certainly doesn't accelerate the particle.

BUT, these eqns allow (q, m) to accelerate, and -- in accelerating -- (q, m) must radiate away EM energy, which is carried away and lost to the system [system = (q, m) + the external \underline{E} & \underline{B}] by (q, m) 's own fields. A radiative loss term appears nowhere in Eqs. (1)-(3), but it should because such an energy

An Example of a Radiation Reaction Force f_{RR} .

RR12

loss must change (inhibit) the motion of (q, m) . So, Eqs (1)-(3) cannot be a complete description of (q, m) 's motion -- there must be missing terms that account for (q, m) 's radiation... terms implicitly dependent on (q, m) 's self-fields.

- 2) What we are talking about can be seen clearly in the case of a 1D  nonrelativistic motion of (q, m) accelerated in an external electric field E . According to Eq. (2), the field does work on the particle at rate qEv . This work will appear as a K.E. increase at rate $\frac{d}{dt}(\frac{1}{2}mv^2)$, and must also supply the radiation energy loss at rate P_{rad} . So Eq. (2) should read...
- $$\rightarrow \frac{d}{dt}(\frac{1}{2}mv^2) + P_{rad} = qEv. \quad (4)$$

For the radiation loss term we can try the Larmor rate: $P_{rad} = (2q^2/3c^3)\dot{v}^2$ [Jk² Eq. (14.22)], so that (4) yields...

$$\frac{d}{dt}(\frac{1}{2}mv^2) + \frac{2q^2}{3c^3}\dot{v}^2 = qEv \quad \left\{ \begin{array}{l} \text{do the } \frac{d}{dt}, \text{ divide by } v, \text{ and} \\ \text{define: } \tau_0 = 2q^2/3mc^3 \end{array} \right. \left\{ \begin{array}{l} \text{time} \\ \text{scale} \end{array} \right.$$

$$\text{ny} \quad \underline{m\dot{v} = qE - f_{RR}} \quad \text{w//} \quad \underline{f_{RR} = m\tau_0(\dot{v}^2/v)}. \quad (5)$$

Eq. (5) is the force law in this case, now modified by a "radiation reaction" force f_{RR} that opposes (q, m) 's motion. f_{RR} is present because during the motion that produces an acceleration \dot{v} , the particle must radiate at rate $f_{RR}v$.

f_{RR} in the form given in (5) is neither complete nor correct (see ASIDE), even non-relativistically. But the manner in which it enters the energy Eq. (4) and motion Eq. (5) is characteristic: f_{RR} represents a loss term that opposes the motion. It would enter the covariant Lorentz law of Eq. (3) as...

$$\boxed{m\dot{u}^\alpha = (q/c)F_{ext}^{\alpha\beta}u_\beta - f_{RR}^\alpha}. \quad (6)$$

Solution for the Larmor correction in Eq. (5).

RR(3)

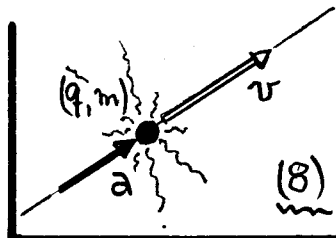
ASIDE Solutions to Eq. (5): 1D motion with Larmor loss term.

1. Write (5) as : $\dot{v} = a - (\tau_0/v) \dot{v}^2$ \int $a = qE/m$ is external acceleration, (7)
 $\tau_0 = 2q^2/3mc^3$ is a scale time.

2. If $a=0$, have : $\dot{v} = -v/\tau_0 \Rightarrow v(t) = v(0) \exp[-(t/\tau_0)]$. Peculiar... any moving charge spontaneously comes to a stop w/o being decelerated. What?

3. If $a \neq 0$, (7) is a quadratic eqn for \dot{v} , w/ solution:

$$\rightarrow \dot{v} = \frac{1}{2\tau_0} [\sqrt{v^2 + 4a\tau_0 v} - v] \quad \int \text{root sign chosen so that: } \dot{v} \rightarrow 0 \text{ when } a \rightarrow 0.$$



This eqn can be integrated to find $t = t(v)$. For $v = v_0$ @ $t = 0$ & $a = \text{const}...$

$$\boxed{at = U(v) - U(v_0)} \quad \int \text{w/ } U(v) = u(v) + a\tau_0 \ln[1 + \frac{1}{2a\tau_0} u(v)], \quad (9)$$

$$\text{and: } u(v) = \frac{1}{2} (v + \sqrt{v^2 + 4a\tau_0 v}).$$

4. Analyse solution of Eq. (9) for both { acceleration : $a > 0$
 deceleration : $a < 0$. Two cases of interest...

(A) Acceleration of (q, m) from rest : $v_0 = 0$, and $a = |a| > 0$.

$$\left. \begin{array}{l} \text{at long times:} \\ t \gg \tau_0, v \gg a\tau_0 \end{array} \right\} \underline{v(t) \approx at - a\tau_0 \ln(t/\tau_0)}, \quad \underline{\dot{v}(t) \approx a[1 - (a\tau_0/v)]}; \quad (10A)$$

$$\left. \begin{array}{l} \text{at short times:} \\ t \ll \tau_0, v \ll a\tau_0 \end{array} \right\} \underline{v(t) \approx \frac{1}{4} at^2/\tau_0}, \quad \underline{\dot{v}(t) \approx \frac{1}{2} at/\tau_0} \quad \int \text{NOTE: } \dot{v}^2/v = \frac{a}{\tau_0} = \text{const.} \quad (10B)$$

The long-time motion is (nearly) Newtonian. Short-time motion is radically new.

NOTE : at short times : $f_{RR} = m\tau_0 \frac{\dot{v}^2}{v} \approx ma$, which cancels $F_{ext} = ma$. Strange!

(B) Deceleration of a moving (q, m) : $v_0 > 0$, and $a = -|a| < 0$.

Per Eq. (9) : $|a|t = U(v_0) - U(v)$, but now with : $u(v) = \frac{1}{2} (v + \sqrt{v^2 - 4|a|\tau_0 v})$.

At $t > 0$, $v(t)$ declines toward 0. BUT, when $v < 4|a|\tau_0$, the $\sqrt{\quad}$ becomes imaginary; solution of Eq. (9) does not apply at the end of the motion, i.e. for velocities in the range : $4|a|\tau_0 > v > 0$. We cannot stop (q, m) in a sensible way.

ASIDE (cont'd) 1D motion with Larmor loss term.

5. The glitch just noted -- our inability to account for the motion (or even whereabouts) of a charged particle which is being decelerated and which has slowed down to a velocity $v < 4/27 c_0$ -- certainly suggests that the Larmor form of a radiation reaction force, $f_{RR} = m\tau_0 \dot{v}^2/v$, as written in Eq. (5), is not yet complete. Some small part of (q,m) 's inertial reaction to an external force is missing. Otherwise the motion analysed in Eqs. (7)-(10) seems plausible.

3) What is missing in $f_{RR}(\text{Larmor}) = m\tau_0 \dot{v}^2/v$ is very likely the effects of (q,m) 's fields "in close", i.e. in the static & induction zones of its radiation field. The Larmor radiation rate $f_{RR}v$ as written only accounts for the radiation zone fields, i.e. those parts of $E(\text{self})$ & $B(\text{self})$ which $\propto 1/R$, where R = distance between (q,m) and the observer. But (q,m) 's Liénard Wiechert fields [Jk² Eqs (14.13) & (14.14)] contain parts $\propto 1/R^2$ as well as $1/R$, so a full expansion of (q,m) 's Poynting vector (from whence the radiation loss comes) is:

$$\rightarrow E(\text{self}) \times B(\text{self}) \propto \frac{1}{R^2} \left(\begin{smallmatrix} \text{radiation} \\ \text{zone} \end{smallmatrix} \right) + \frac{1}{R^3} \left(\begin{smallmatrix} \text{induction} \\ \text{zone} \end{smallmatrix} \right) + \frac{1}{R^4} \left(\begin{smallmatrix} \text{static} \\ \text{zone} \end{smallmatrix} \right). \quad (11)$$

We've used only the first term RHS: this surely constitutes radiation (loss) and surely does hinder (q,m) 's motion via the above f_{RR} . The 2nd & 3rd terms cannot be radiation (since $\oint (1/R^{3,4}) dS \rightarrow 0$ for a large sphere at ∞)... but it would seem that they could contribute to (q,m) 's inertial reaction in an equation of motion. After all, the induction & static zone fields -- with their nontrivial energy densities -- must continually adjust their configurations during an externally applied acceleration.

We shall not pursue these notions (they quickly lead to questions re what happens as $R \rightarrow 0$... what is (q,m) 's structure). Instead, we review some orthodoxy.

4) It is possible to argue that searching for correct & complete radiation reaction corrections in classical electrodynamics is largely irrelevant -- because these corrections are negligible in most cases of interest, or because quantum-mechanical limitations become dominant long before we reach the distance & time scales at which RR corrections become important. The arguments go as follows.

① Consider nonrelativistic 1D motion of (q, m) at const acceleration a .

Starting from rest, in time $0 \rightarrow T$, the particle (q, m) exhibits:

$$\left[\begin{array}{l} \text{kinetic energy: } K = \frac{1}{2} m (aT)^2, \text{ radiated energy: } E_{\text{radn}} = \left(\frac{2q^2}{3c^3} a^2 \right) T; \\ \text{ } E_{\text{radn}} \sim K \text{ only if } T \sim 2\tau_0, \text{ } \tau_0 = \frac{2}{3} \frac{q^2}{mc^3}. \end{array} \right. \quad (12)$$

τ_0 is the scale time defined in Eq. (5)... it is extremely small for an elementary particle (e.g. $\tau_0 = \frac{2}{3} \frac{e^2}{mc^3} \approx 6.26 \times 10^{-24}$ sec, for an e).

The distance traveled during the time when E_{radn} is significant is also very small, certainly \ll classical charge radius $\tau_0 = \frac{q^2}{mc^2} = \frac{3}{2} c\tau_0$ (for an e : $\tau_0 = \frac{e^2}{mc^2} = 2.82 \times 10^{-13}$ cm). Altogether, E_{radn} has a crucial role in the motion only over negligible times τ_0 & distances τ_0 .

② The time $\tau_0 = 2q^2/3mc^3$ and distance $\tau_0 = q^2/mc^2$ scales are QMly inaccessible for measurements on elementary particles. If we try measuring times down to $\Delta t \sim \tau_0$, then we will impart to (q, m) random energies of a size:

$$\rightarrow \Delta E \sim \hbar / \tau_0 = mc^2 / \left(\frac{2}{3} \frac{q^2}{\hbar c} \right), \quad (13)$$

According to the Uncertainty Principle. If $q = \pm e$, then $\Delta E \sim 205 mc^2$, so q goes ballistic. Similarly, localization to $\Delta x \sim \tau_0 \Rightarrow$ uncontrollable momentum up to $\Delta p \sim \hbar / \tau_0 = mc / (q^2 / \hbar c) \sim 137 mc$, when $q = \pm e$. Conclusion: measurements at time & distance scales τ_0 & τ_0 are impossible.