5) Now we turn to Topic II on p. QF1 -- i.e. quantization of An. We do this by analogy with the well-known QM of the simple harmonic oscillator (SHO). The exercise provides a natural context for the "photon" concept. The reason for invoking the SHO is that the "normal modes" of the Em radiation field can be thought of -- by means of Frurier's Theorem -- as a continuous distribution of elementary oscillators.

Single mode, standing-wave realization of the EM field. [Davydov, 980]. QF6

Define terms ...

(1) k = wave vector, $k = |k| = \text{wave number} \left\{ \begin{array}{l} \text{wave length} : \lambda = 2\pi/k, \\ \text{Alar freq.} : \omega = kc. \end{array} \right.$ (15a)

(2) $\hat{\epsilon}$ = unit polarization vector $\begin{cases} \hat{\epsilon} = \text{unit vector in direction of } E_n \cdot \underline{\text{NOTE}} : \text{there } \underline{\text{115b}} \end{cases}$ are two indept polarizations for each transverse wave.

(3) Drop Subscript "n" for the radiation fields. All fields calculated henceforth are exclusively free radiation fields (no sources present, etc.)*

To begin, consider vector potential for a <u>standing</u> EM wave at wave vector k and poisrization ê (a single mode); it is...

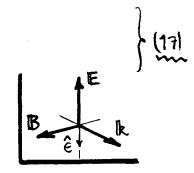
→ A = ê 5 cos(k·r), obeying: (\(\nabla^2 + k^2\) A = 0, Standing wave legtin, (16)

The amplitude & of A, which is arbitrary in time, turns out to be the generalized "position" coordinate for this elementary EM field oscillator.

For a free radiation field, we can choose $\phi = 0 \notin \nabla \cdot A = 0$ (Contomb gange). Then the fields accompanying A of Eq. (16) are:

[electric field: $E = -\frac{1}{c} \partial A/\partial t = -\hat{\epsilon}(\xi/c) \cos(k \cdot r);$ [magnetic field: $B = \nabla \times A = -(k \times \hat{\epsilon}) \zeta \sin(k \cdot r).$

We have a transverse EM wave, since E·B=0, and both fields are I lk. The directions k, E& B are mutually I.



The energy density of the standing wave in Eq. (17) is, locally ...

Over a macroscopie volume V (linear dimension >> $\lambda = \frac{2\pi}{k}$), total field energy is:

By this time, we are thinking of the problem as that of an atom enclosed in a box with perfectly reflecting walls. Photons (radiation) in the box bonnce around randomly (& freely), bothing the atom in Arad, etc.



Analogy: rad field energy + SHO. How SHO's work.

$$\mathcal{E} = \int U d^3x \rightarrow \overline{U} V = \frac{V}{8\pi} \left[(\ddot{\zeta}/c)^2 \langle \cos^2(lk \cdot r) \rangle + k^2 \zeta^2 \langle \sin^2(k \cdot r) \rangle \right]$$

$$(avg. value in V) \frac{1}{2}, over many $\lambda'^5$$$

i.e.
$$\frac{\dot{E}}{2} = \frac{1}{2} \left[(V/8\pi c^2) \dot{\zeta}^2 + (Vk^2/8\pi) \dot{\zeta}^2 \right]$$
Lurge bex of volume V. (19)

This expression is the direct analogue of the total energy of a SHO, as

... for SHO of mass m, natural frequency w, and generalized position q...

$$E_{SH0} = \frac{1}{2} \left[m \dot{q}^2 + m \omega^2 q^2 \right].$$

(here q = coordinate,) (20)

Evidently the two expressions are identical, if we define:

(21)

We shall take this analogy serionsly, because the <u>structure</u> of a quantized field theory must deal with the SHO decomposition of those fields, and thus must incorporate SHO quantization to some degree. Eqs. (19)-(21) are a beginning.

ASIDE Recollections on how QM SHO's work. [Ref. Sakurai, Sec. 2.3].

1. Let Q= generalized position of the SHO motion, and P= m Q = canonical momentum.

-Auentize SHO by imposing
$$[Q,P]=i\hbar$$
 on $H=\frac{1}{2m}P^2+\frac{1}{2}m\omega^2Q^2$. (22)

Get eigenfons M). In Q-space, these are Hermite polynomials of degree N=0,1,2,3...

Get eigenenergies: En=(N|HIN)=(N+2) ħω. (23)

N=3 N=2 N=0 Q

NOTE: for N=0, get "zero-point energy": Eo = ½ tow. This ______ Q energy is really present, and is necessary for compatibility Uncertainty Principle.

2. In the SHO state IN), we can say the energy is: EN = E0+ energy of Nquanta, we can say the energy is W. These quanta well later be

identified with photons, so the SHO state IN) signifies there are N photons present, each with energy trw. It is fair to mention photons here, because—ly Einstein's work on the photoelectric effect (ca. 1905)—the photon had already been identified as a working quantum, carrying energy trw. NOTE: even when no photons are present (i.e. N=0 and "bacuum state" 10>), the mode we still has the "zero-point" energy Eo=\frac{1}{2}thw, like half a photon there, for free. These energies are called the "zero-point oscillations" of the field, and cannot be turnin out % violating the Uncertainty Principle. This zero-point behavior causes Big Trouble... e.g. the overall vacuum 10> has & energy.

3. Some SHO matrix elements will be useful to us. They are...

END of ASIDE

* E.g. the average energy of an assembly of QM SHO'S at frequency ω , in thermal equilibrium at temperature T is (1) k=Boltzmann cost of EN=(N+\frac{1}{2})\tau\): $\overline{E}(\omega) = (\sum_{N=0}^{\infty} E_N e^{-E_N/kT})/(\sum_{N=0}^{\infty} e^{-E_N/kT}) = \hbar\omega/(e^{\kappa\omega/kT}-1) + \frac{1}{2}\kappa\omega$ $\frac{E(\omega) = (\sum_{N=0}^{\infty} E_N e^{-E_N/kT})/(\sum_{N=0}^{\infty} e^{-E_N/kT})}{e^{-E_N/kT}} = \frac{\hbar\omega}{e^{-E_N/kT}}$

The total field energy: J. E(w) × [*modes] × dw, then picks up a contribution from the zero-point energies which goes as Jow dw → ∞. Called an "ultraviolet catastrophe".

) Now, to quantize the EM field in the mode (Ik; ê), we just take over the SHO quantization verbation. That is, where & is the amplitude of the field vector potential [i.e. A=ê5 cos(k.r) of Eq. (16)], we define -- as in Eq. (21):

[S = q, generalized position; $p = m\dot{q} = (V/8\pi c^2)\dot{S}$, canonical mom²; $kc = \omega$, natural frequency; $V/8\pi c^2 = m$, field mass (Sic).

Then, imposing: [4, p] = it, quantizes the field energy $\overline{\epsilon}$ in Eq. (19), and we get:

field SHO { eigenstates | N > ← a state (at freq. ω) with N photons present; (27) eigenenergies: EN=(N+½) πω ← energy of N photons + 3ero-pt.

It's that simple. The field can now supply or absorb photons via N -> N71.

The idea of emission & absorption of photons by the field his a change N > N = 1 in the photon occupation number now finds a beautifully clear representation in the "ladder operators" a & at of the SHO. We define:

The SHO quantization condition: [q.p]=it, is now replaced by: [a,at]=1 (Sakurai, Egs. (2.3.2) & (2.3.3), etc). Only nonvanishing matrix elements are...

 $(N-1|a|N) = \sqrt{N}$, i.e., $a|N\rangle = \sqrt{N}|N-1\rangle$ field loses one photon; $(N+1|a^{\dagger}|N) = \sqrt{N+1}$, i.e., $a^{\dagger}|N\rangle = \sqrt{N+1}|N+1\rangle$ field gains one photon.

Here the notion of the field emitting (i.e. supplying) or absorbing pho- in tons is born... these are just processes where a or at operate on the field SHO states (N) of Eq. (27), supplying or absorbing a quantum of energy to.

With this quantization scheme, we can now say the radiation field is quantized in the mode (k; ê) because all its transactions with other (quantum) systems will occur in terms of discrete quanta (photons) of energy tow. The field quantities themselves become operators, as follows. The inverse of Eqs. (28) give the "position" 5 and momentum" p = m 5 of A as...

$$\int \zeta = (4\pi \hbar c / V k)^{1/2} (a^{\dagger} + a), \quad p = i (V \hbar k / 16\pi c)^{1/2} (a^{\dagger} - a);$$

$$\int_{V} A = \hat{\epsilon} \zeta \cos(k \cdot r) = (\frac{4\pi \hbar c}{V k})^{\frac{1}{2}} \hat{\epsilon} (a^{\dagger} + a) \cos(k \cdot r).$$
(30)

The at & a do not operate on the (atom's) space coordinate it; they just operate on the photon states IN) in the field. Then it is clear that A is an operator, in that A|N) will generate states |N±1> in the field.

Now, when matrix elements of Ybint ~ A (Tadn). p (Charge) [See Eq. (11b)] are taken w.n.t. direct product states | n(aton) (N(photons)) > [See Eq. (12)], as indicated in the theory sketch in Eq. (14), we have a <u>mechanism</u> for ensuring that $\langle m(M)| \text{Home} | n(N) \rangle \sim \langle (M)| \text{Al}(N) \rangle \cdot \langle m| \text{pln} \rangle$ will properly describe $n \rightarrow m$ for the atom, with $N \rightarrow M = N \pm 1$ for the field interpreted as the emission or absorption of a photon during the transition.

We have work to do to generalize these notions. At this point, we have only quantized the standing wave $(k;\hat{\epsilon})$ vector potential A of Eq.(30). We need to put in time dependence, to get traveling waves, and we should generalize to a k-spectrum for arbitrary photon fields. This is \sim straight forward.

In passing, we note the field Hamiltonian for freq. w can be written This comes in handy, later, when we track the field energy.