

5) As a first application of Eq. (11), we find the $\Theta(V)$ amplitudes. For $\mu=0$:

$$\rightarrow i\hbar a_k^{(1)}(t) = \sum_n a_n^{(0)} \int_{t_0}^t V_{kn}(\tau) e^{i\omega_{kn}\tau} d\tau. \quad (12)$$

First order transition amplitude for $m \rightarrow k$.

tD15

Choose $a_n^{(0)} = \delta_{nm}$, i.e. system initially in eigenstate m (for $t < t_0$). Then...

$$\boxed{a_k^{(1)}(t) = -(i/\hbar) \int_{t_0}^t V_{km}(\tau) e^{i\omega_{km}\tau} d\tau}, \text{ from initial state } m. \quad (13)$$

The system wavefun is by now...

$$\begin{aligned} \psi(x,t) &= \sum_n [a_n^{(0)} + a_n^{(1)}(t) + \dots] \phi_n(x) e^{-i\omega_n t} \\ \text{or} \rightarrow \psi(x,t) &\simeq \underbrace{\phi_m(x) e^{-i\omega_m t}}_{\text{initial state}} + \sum_k \underbrace{a_k^{(1)}(t) \phi_k(x) e^{-i\omega_k t}}_{\text{states mixed into } m \text{ by } V(x,t)}. \end{aligned} \quad (14)$$

The overlap of ψ in state k , i.e. the amplitude for state k to appear here is:

$$\langle \phi_k | \psi \rangle \simeq \delta_{km} e^{-i\omega_m t} + a_k^{(1)}(t) e^{-i\omega_k t}. \quad (15)$$

Suppose $k \neq m$. Then the 1st term RHS here vanishes, and...

$$\boxed{|\langle \phi_k | \psi \rangle|^2 \simeq |a_k^{(1)}(t)|^2 = (1/\hbar^2) \left| \int_{t_0}^t V_{km}(\tau) e^{i\omega_{km}\tau} d\tau \right|^2}. \quad (16)$$

INTERPRETATION. This is the probability (to lowest order in V) of a transition $m \rightarrow k$ induced by V (which supplies or absorbs the the excitation energy $(\omega_m - \omega_k)$). Hence $a_k^{(1)}$ is called the "1st order transition amplitude" for $m \rightarrow k$.

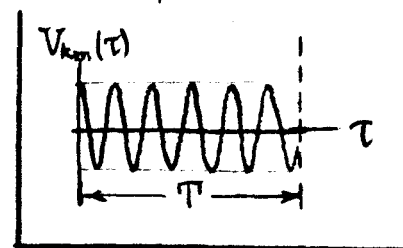
6) With Eq. (16) in hand, we do an example of how a time-dependent coupling $V(x,t)$ actually drives transitions out of state m into states $k \neq m$.

EX. PULSED HARMONIC PERTURBATION

$$V(x,t) = \hbar \Omega(x) [e^{+i\omega t} + e^{-i\omega t}] = 2\hbar \Omega(x) \cos \omega t \quad \begin{cases} \text{for } 0 \leq t \leq T, \text{ and} \\ 0, \text{ otherwise} \end{cases}$$

$$\text{or} \rightarrow V_{km}(\tau) = \hbar \Omega_{km} [e^{i\omega\tau} + e^{-i\omega\tau}], \text{ over } 0 \leq \tau \leq T$$

$$\text{where; } \hbar \Omega_{km} = \langle \phi_k(x) | V(x) | \phi_m(x) \rangle. \quad (17)$$

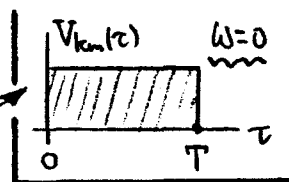


Example of Pulsed Harmonic Perturbation

td(6)

Such a coupling $V(x,t)$ could represent the effect of a laser pulse (monochromatic, at freq. ω) shining on an atom [initially in state m] for some finite time T .

Notice that when $\omega \rightarrow 0$, then $V_{km}(t) = 2\hbar\Omega_{km} = \text{const}$, $0 \leq t \leq T$, and zero, otherwise... so the "atom" is exposed to a const field pulse.



Put $V_{km}(t)$ of Eq. (17) into the first-order amplitude of Eq. (13)...

$$a_k^{(1)}(t \gg T) = -\frac{i}{\hbar} \hbar \Omega_{km} \int_0^T (e^{i\omega\tau} + e^{-i\omega\tau}) e^{i\omega_{km}\tau} d\tau,$$

$$\xrightarrow{\text{So}} a_k^{(1)}(t \gg T) = \Omega_{km} \left[\underbrace{\frac{1 - e^{i(\omega_{km} + \omega)T}}{\omega_{km} + \omega}}_{\textcircled{1}} + \underbrace{\frac{1 - e^{i(\omega_{km} - \omega)T}}{\omega_{km} - \omega}}_{\textcircled{2}} \right] \quad \begin{matrix} \checkmark t \gg T \text{ means} \\ \text{pulse } V \text{ is} \\ \text{finished.} \end{matrix} \quad (18)$$

REMARKS on Eq. (18).

1. V "small" here $\Rightarrow |\Omega_{km}| \ll |\omega_{km}|$. Then $a_k^{(1)}(t \gg T)$ is appreciable only at the "resonances": $\omega_{km} \approx \mp \omega$, when: $a_k^{(1)}(t \gg T) \approx -i\Omega_{km}T$. This can become large as $T \rightarrow \text{large}$ (and requires higher order pertⁿ theory to handle completely).

2. The resonant condition $\omega_{km} \approx \mp \omega$ can be written in terms of energies as...

(19)

$E_k^{(0)}[\text{final}] = E_m^{(0)}[\text{initial}] \mp \hbar\omega$

EMISSION { term ① in (18),
(-) sign at left.

ABSORPTION { term ② in (18),
(+) sign at left.

Interpretation: the resonant behavior of $a_k^{(1)}(t \gg T)$ in Eq. (18) is associated with either the emission

(term ①) or absorption (term ②) of a "photon" of energy $\hbar\omega$, $\omega \approx |\omega_{km}|$, from V -field.

3. If we are interested in absorptive processes only (which pump energy into the "atom"), then in Eq. (18) only term ② is important, and we can write...

for $E_m^{(0)} \rightarrow E_k^{(0)} > E_m^{(0)}$, absorption, the $m \rightarrow k$ transition probability is:

$$|a_k^{(1)}|^2 = 4|\Omega_{km}|^2 \sin^2 \frac{1}{2}(\omega_{km} - \omega)T / (\omega_{km} - \omega)^2 \quad \begin{matrix} \text{for irradiation by pulse} \\ \text{of length } T \text{ @ frequency } \omega. \end{matrix} \quad (20)$$

This is the lowest order result. The emission probability is gotten via $\omega \rightarrow (-)\omega$ here.

Further general remarks on 1st order amplitude

td(7)

7) Two other general remarks can be made about the first-order amplitude in Eq.(13):

4. Even though emissive or absorptive transitions $m \rightarrow k$ are induced by V , the population of the initial state m is barely depleted (in this order of pertⁿ theory).

The argument goes as follows. By time t after turn-on, V has induced the state

$$\rightarrow \psi(x, t) = \underbrace{\phi_m(x)}_{\text{initial state}} e^{-i\omega_m t} + \sum_n \underbrace{a_n^{(1)}(t)}_{\text{admixture due to } V} \phi_n(x) e^{-i\omega_n t}, \text{ to } O(V). \quad (21)$$

The amplitude of ϕ_m still present in ψ is ...

$$\rightarrow \langle \phi_m | \psi \rangle = e^{-i\omega_m t} [1 + a_m^{(1)}(t)] = e^{-i\omega_m t} \underbrace{\left[1 - \frac{i}{\hbar} \int_{t_0}^t V_{mm}(\tau) d\tau \right]}$$

$$\text{So } \langle \phi_m | \psi \rangle = e^{-i\omega_m t} e^{-\frac{i}{\hbar} \int_{t_0}^t V_{mm}(\tau) d\tau}, \quad \underbrace{\hspace{10em}}_{= e^{-\frac{i}{\hbar} \int V_{mm} d\tau}, \text{ to } O(V);}$$

$$\left. \begin{array}{l} \text{only population of initial} \\ \text{state } m \text{ at time } t \end{array} \right\} \underline{\underline{|\langle \phi_m | \psi \rangle|^2 = \left| e^{-\frac{i}{\hbar} \int V_{mm} d\tau} \right|^2 \rightarrow 1.}} \quad (22)$$

All that is needed here is to assume V is real (necessary to ensure $\mathcal{H}_0 + V$ is Hermitian).

The depletion of state m by transitions $m \rightarrow k$ is at most an $O(V^2)$ effect.

5. Consider $m \rightarrow k$ couplings via $V_{km}(\tau)$ which last a "long" time [e.g. could be $V_{km}(\tau) \propto e^{-\alpha\tau^2}$, a Gaussian; all we really need is duration $\Delta\tau \gg 1/|\omega_{km}|$].

Formally, the $m \rightarrow k$ transitions are not finished till $t \rightarrow \infty$, so we need:

$$\rightarrow a_k^{(1)}(\infty) = -\frac{i}{\hbar} \int_{-\infty}^{\infty} V_{km}(\tau) e^{i\omega_{km}\tau} d\tau, \quad (23)$$

to find out the final $m \rightarrow k$ transfer. Except for the factor $(i\hbar)^{-1}$ out in front, the required amplitude is just the Fourier Transform of $V_{km}(\tau)$.

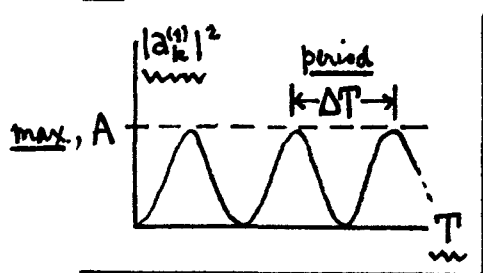
So the way a quantum system reacts to an impressed time-dependent perturbation $V(\tau)$ is to Fourier analyse $V(\tau)$ w.r.t. its own natural freqs. ω_{km} .

8) We backtrack to the Pulsed Harmonic Perturbation of Eqs. (17)-(20) above, in order to learn more about the $m \rightarrow k$ transition dynamics.

Consider absorptive probability for $m \rightarrow k$ per Eq. (20), viz.

$$|a_k^{(1)}|^2 = \frac{4|\Omega_{km}|^2}{(\omega_{km}-\omega)^2} \sin^2 \frac{1}{2}(\omega_{km}-\omega)T \quad \text{for } m \rightarrow k \text{ via pulse of duration } T \text{ @ freq. } \omega. \quad (24)$$

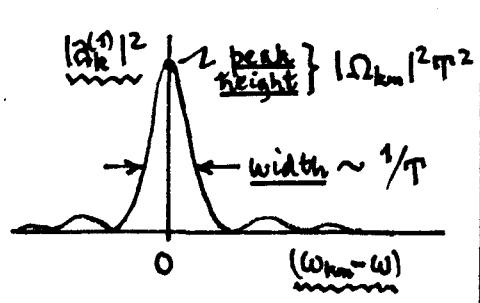
A. Plot $|a_k^{(1)}|^2$ vs. pulse duration T ; all other parameters fixed.



max.: $A = 4|\Omega_{km}|^2/(\omega_{km}-\omega)^2$; period: $\Delta T = 2\pi/|\omega_{km}-\omega|$.

System exhibits a "quantum oscillation" $m \rightleftharpoons k$ out of and back into initial state m @ freq. $|\omega_{km}-\omega|$.

B. Plot $|a_k^{(1)}|^2$ vs. off-resonance freq. $\omega_{km}-\omega$; all other parameters fixed

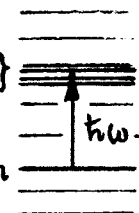


Area under curve \sim peak height \times width $\sim T$. As $T \rightarrow$ large, transition $m \rightarrow k$ with $E_k^{(0)} = E_m^{(0)} + \hbar\omega$ becomes more certain. Curve width $\sim 1/T$ is consistent with energy uncertainty $\Delta E \sim \hbar/T$ for $m \rightarrow k$.

NOTE: $(|a_k^{(1)}|^2/T)$ is a nascent delta-function.

C. Suppose the absorption is $m \rightarrow \{k\}$, a set of final states k . Then consider:

$$\left\{ \begin{array}{l} \text{total transition} \\ \text{probability: } m \rightarrow \{k\} \end{array} \right\} P_m = \sum_k |a_k^{(1)}|^2, \quad \text{w/ } E_k^{(0)} \approx E_m^{(0)} + \hbar\omega. \quad (25)$$



[Keep in mind we are working to lowest order (leading order) in V].

Now suppose the final states $\{k\}$ lie close together in energy, and define

$\rightarrow p(k) dE_k =$ number of final states k with energy in range $E_k^{(0)}$ to $E_k^{(0)} + dE_k$. (26)

$p(k)$ is called the "density of states" function. Now in these terms, write P_m

Derivation of Fermi's Golden Rule

tD(9)

$$\rightarrow P_m = \sum_k |a_k^{(1)}|^2 \rightarrow \int_{\{k\}} |a_k^{(1)}|^2 \rho(k) dE_k = \int_{\{k\}} |\Omega_{km}|^2 \frac{\sin^2 \frac{1}{2}(\omega_{km} - \omega)T}{[\frac{1}{2}(\omega_{km} - \omega)]^2} \rho(k) \hbar d\omega_k$$

... let: $x = \frac{1}{2}(\omega_{km} - \omega)T$, so: $dx = \frac{1}{2}T d\omega_k$. Then...

$$\left[P_m = 2\hbar T \int_{\{k\}} \rho(k) |\Omega_{km}|^2 \left(\frac{\sin^2 x}{x^2} \right) dx. \right. \quad (27)$$

The integrand in (27) has a strong peak @ $x=0$, which corresponds to strict energy conservation: $E_k^{(1)} \equiv E_m^{(0)} + \hbar\omega$, for $m \rightarrow k$ via absorption of photon $\hbar\omega$. But the integrand also has a finite width $\Delta x \sim T \Delta\omega_k$, which is demanded by the final state energy uncertainty $\Delta E_k = \hbar \Delta\omega_k \sim \hbar/T$ for $m \rightarrow k$ in a finite time T . We integrate over all such "uncertain" transitions.

Assume $\rho(k)$ & Ω_{km} are "slowly varying" functions of k (or ω_k) near the $x=0$ peak in the integrand of (27). Take them out of the integral, so as to write...

$$\left[\begin{array}{l} \text{Davydov Eq. (93.4)} \\ \text{Sakurai Eq. (5.6.34)} \end{array} \right] \rightarrow P_m \approx 2\hbar T [\rho(k) |\Omega_{km}|^2]_{\text{AVG.}} \underbrace{\int_{-\infty}^{\infty} \left(\frac{\sin^2 x}{x^2} \right) dx}_{\pi} \quad (28)$$

The "AVG." means an average (typical) value of $\rho(k) |\Omega_{km}|^2$ near $E_k^{(1)} = E_m^{(0)} + \hbar\omega$.

The integral = π , and we put in $\Omega_{km} = \frac{1}{\hbar} \langle k | V | m \rangle$ [see Eq. (17)]. Then:

$$\boxed{W(m \rightarrow \{k\}) = P_m/T = \frac{2\pi}{\hbar} \left(|\langle k | V | m \rangle|^2 \rho(k) \right)_{\text{AVG.}}} \quad \text{FERMI'S GOLDEN RULE} \quad (29)$$

W is the transition probability per unit time for $m \rightarrow \{k\}$, induced by the coupling V . $\rho(k)$ is the density of final states k , and (on average) energy is conserved: $E_k^{(1)} = E_m^{(0)} + \hbar\omega$, with the "photon" $\hbar\omega$ supplied by the V -field.

NOTE that $W(m \rightarrow \{k\})$ is independent of the time T over which V acts... that is the surprising feature of this calculation. The result for W can be derived on very general grounds (S -matrix theory), and evidently is a controlling fact for all transition rate calculations. To leading order, anyway.