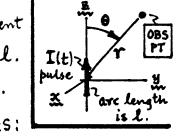
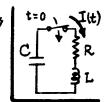


- (79)[20 pts]. In Sec. 12.9, Jackson quotes the Proca Lagrangian Lp [Eq. 112,91)], and de- (25) rives Proca's wave egts for a massive photon field [Eq. (12.93)]. Along the way, he claims the "Lorentz gauge is now required for current conservation."
- (A) Show why Jackson's claim about the Lorentz gauge is justified.
- (B) Assume the initial choice for Lp is the Coventz gange: Da Aa=0, so the current Ja is conserved:  $\partial^{\alpha} J_{\alpha} = 0$ . Now consider a gauge transform:  $A_{\alpha} \rightarrow A_{\alpha}' = A_{\alpha} + \partial_{\alpha} G$ . What condition on the gange for G is needed to maintain current conservation? The theory is nonsense 1/0 conserved currents!). What gauge are you in now?
- (C) With the current-conserving gauge transforms allowed in part (B), show that the transformed Lagrangian is:  $L_p = L_p - \partial^{\alpha}U_{\alpha}$ , where  $U_{\alpha}$  is a vector field that depends on Ja & Aa. Find Va explicitly. Are Lp & Lp "gange equivalent" in the sense of problem (3)?
- (D) Make a statement regarding the gange freedom of Lp (as a massive vector field L).
- (80 [20 pts]. Consider a 1D are discharge along the z-axis: a current pulse I(t) begins at time t=0 and flows along a path of length l. An observer at position  $(r, \theta)$ ,  ${}^{W}r>> l$ , detects the arc's radiation.



- (A) Start from the arc's Poynting vector S derived in class (NOTES: p. Rad 7, Eq (18)]. The energy/unit time radiated into solid \$ dΩ is T2S dΩ, so: dE/dΩ = 5- r2S dt, is the total energy radiated per unit solid 4. Convert this to to a frequency integral:  $d\mathcal{E}/d\Omega = \int_{-\infty}^{\infty} \sigma(\omega) d\omega$ [Parseval's Theorem]. Show the spectrum for is  $\sigma(\omega) = \frac{\sin^2\theta}{8\pi^2c^3} \ell^2\omega^2 \int_{-\infty}^{\infty} I(t)e^{-i\omega t} dt$
- (B) A model for I(t) is the discharge of a capacitor C (switched on at t=0, w/ c to the control of the control overdamped case): I(t) = (Vo/LP) e-yt sinh rt, my y = R , r= Jy2-(1/LC).



Sketch this Ilt) pulse vs. t, and roughly estimate the pulse risetime & duration.

- (C) Calculate the arc frequency spectrum O(W) for the I(t) model of part (B). Sketch O(W) VS. W. Over what frequency range is the arc radiation appreciable?
- (D) Find the total energy radiated by the arc. Compare it with SI2Rdt = discharge. HINT: See R. Robiscoe & Z. Sui, J. Appl. Phys. <u>64</u>, 4364 (Nov. 1988).

## \$520 Solutions



(19) [20 pts.]. Analyse gange constraints on Proca Lagrangian.

1. The Proca Lagrangian is 
$$Jk^{\text{P}}$$
 Eq. (12.91):  

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The egtis-of-motion (i.e. 0 [0 Lp/0(0 Aa)] = 0 Lp/0 Aa) are Eq. (12.92):

$$\rightarrow \partial^{\beta} F_{\beta\alpha} + \mu^{2} A_{\alpha} = \frac{4\pi}{c} J_{\alpha} , \quad \text{W} F_{\beta\alpha} = \partial_{\beta} A_{\alpha} - \partial_{\alpha} A_{\beta} . \quad \text{(2)}$$

To see how current Ix is conserved, operate through Eq. (2) by 20, so that

$$\rightarrow \partial^{\alpha} \partial^{\beta} F_{\beta\alpha} + \mu^{2} \partial^{\alpha} A_{\alpha} = \frac{4\pi}{c} \partial^{\alpha} J_{\alpha}. \tag{3}$$

It is easy to show that with Fox the antisymmetric field tensor defined in (2), the first term IHS in (3) varishes: Da OB Fpa = 0. Then (3) yields

$$\rightarrow \frac{\partial^{\alpha} J_{\alpha} = (\mu^{2} c/4\pi) \partial^{\alpha} A_{\alpha}}{\partial^{\alpha} J_{\alpha} = 0}, \frac{\partial^{\alpha} J_{\alpha} = 0}{\partial^{\alpha} A_{\alpha} = 0} \leftrightarrow \frac{\partial^{\alpha} J_{\alpha} = 0}{\partial^{\alpha} A_{\alpha} = 0} \leftrightarrow \frac{\partial^{\alpha} J_{\alpha} = 0}{\partial^{\alpha} A_{\alpha} = 0} \leftrightarrow \frac{\partial^{\alpha} J_{\alpha} = 0}{\partial^{\alpha} J_{\alpha} = 0}.$$
(4)

This justifies Jk s claim that the Lorentz gange is required for current constn.

(B) 2: Under a gange transform: Aa -> Aa = Aa + daG, the field tensor Fpa remains unchanged. The field extrs (2) become:  $\partial F F \beta \alpha + M^2 [A \alpha + \partial \alpha G] = \frac{4\pi}{C} J \alpha$ , and operation through this extr by 2d produces the counterpart of Eq. (4):

$$\rightarrow \partial^{\alpha} J_{\alpha} = (\mu^{2} c/4\pi) [\partial^{\alpha} A_{\alpha} + (\partial^{\alpha} \partial_{\alpha})G]. \tag{5}$$

If the original gauge was Lorentz, then 2x Aa=0. According to (5), current is conserved in the new gauge (Aa + DaG) only if the gauge function is restricted:

$$\rightarrow (\partial^{\alpha} \partial_{\alpha}) G = \left(\frac{1}{c^2} \frac{\partial^2}{\partial t^2} - \nabla^2\right) G = 0 \implies G = \text{ free field scalar.}$$
 (6)

But then we are still in the Lorentz gauge (Jk" p. 221), since da A' =0 also. In any case, current can be conserved for Lp, for gauge fons G obeying Eq. (6).

3. When  $A_{\alpha} + A_{\alpha}' = A_{\alpha} + \partial_{\alpha}G$ , the Proca Lagrangian in Eq. (1) becomes...

$$(C) \rightarrow \mathcal{L}'_{P} = -\frac{1}{16\pi} F_{\alpha\beta} F^{\alpha\beta} - \frac{1}{C} J_{\alpha} (A^{\alpha} + \partial^{\alpha} G) + \frac{\mu^{2}}{8\pi} (A_{\alpha} + \partial_{\alpha} G) (A^{\alpha} + \partial^{\alpha} G)$$

$$= \mathcal{L}_{P} - \frac{1}{C} J_{\alpha} \partial^{\alpha} G + \frac{\mu^{2}}{8\pi} [A_{\alpha} \partial^{\alpha} G + A^{\alpha} \partial_{\alpha} G + (\partial_{\alpha} \partial^{\alpha}) G]. \qquad (?)$$

We've used the fact that For is invariant under the gauge transform, and have gathered together the terms that form  $L_p$ . Term 0 = 0 for current conservation [Eq. 6) above ], terms  $0 \neq 3$  corn line to give  $2A^{\alpha} \partial_{\alpha}G$ , and term  $0 \neq 0$  can be written as:  $J_{\alpha} \partial^{\alpha}G = \partial^{\alpha}(J_{\alpha}G) - (\partial^{\alpha}J_{\alpha})G$  (current conservation again). Then...

$$\rightarrow \mathcal{L}_{p}' = \mathcal{L}_{p} - \frac{1}{c} \partial^{\alpha} (J_{\alpha}G) + \frac{\mu^{2}}{4\pi} A_{\alpha} \partial^{\alpha}G. \qquad (8)$$

For the third term on the RHS here, we can write

$$A_{\alpha} \partial^{\alpha} G = \partial^{\alpha} (A_{\alpha} G) - (\partial^{\alpha} A_{\alpha}) G$$
, by Toventz gauge. (9)

Soly 
$$\mathcal{L}_{p}' = \mathcal{L}_{p} - \partial^{\alpha} \left[ \left( \frac{1}{c} J_{\alpha} - \frac{\mu^{2}}{4\pi} A_{\alpha} \right) G \right] ; \alpha = 0, 1, 2, 3.$$

4: Let  $U_{\alpha} = (\frac{1}{c} J_{\alpha} - \frac{M^2}{4\Pi} A_{\alpha}) G$ , and integrate  $L_p'$  of Eq. (10) over a hypervolume with invariant volume element  $d^4x = dx^{\alpha} dx^{1} dx^{2} dx^{3}$ , with the  $x^{\alpha} = ct cd$ .

Tanging from time  $t_1$  to  $t_2$  (fixed endpoints of the motion). Then the <u>action</u> is

$$\rightarrow A_{P}' = \int_{1}^{2} \mathcal{L}_{P}' d^{4}x = A_{P} - \int_{1}^{2} (\partial^{\alpha} U_{\alpha}) d^{4}x. \qquad (11)$$

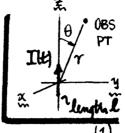
The  $\alpha=0$  term in the integral gives just:  $\int d^3x \, U_0|_{t_1}^{t_2}$ , fixed at the endpts; it contributes nothing to the variation  $\delta A_p^2$ . The  $\alpha=1,2,3$  terms can--by Gauss' Theorem-- be transformed to integrals over hypersurfaces at  $\infty$ , where they vanish. Thus we get  $\delta A_p^2=\delta A_p=0$  together, and  $\mathcal{L}_p^2 \notin \mathcal{L}_p$  of Eq.(10) are Jange equivalent... they will give the same extra-of-motion. We can state:

For current conservation (and gauge equivalence), Lp is totally restricted to the Loventz gauge (3ª Aa = 0). Any further gauge freedom requires additional terms for Lp.

(P)

## 3 [20 pts]. Analyse frez. spectrum for a 1D arc discharge.

(A) Poynting vector as derived in class [class notes, p. Rad 7, Eg. (18)]:



$$\longrightarrow S(r,t) = \left(\frac{\sin^2\theta}{4\pi r^2}\right) \frac{1}{c^3} \left[\tilde{I}(t') l\right]^2 \leftarrow \text{energy/unit time & avea at observer.}$$

The energy/unit time & solid & is 72 S, and if we integrate this over all t, we get

$$\frac{d\mathcal{E}}{d\Omega} = \int_{-\infty}^{\infty} r^2 S(r,t) dt = \left(\frac{\sin^2 \theta}{4\pi c^3}\right) \ell^2 \int_{-\infty}^{\infty} \left[\dot{\mathbf{I}}(t')\right]^2 dt' \leftarrow \frac{radiated energy}{\text{per unit solid } \dot{\mathbf{X}}}. \quad (2)$$

The integral can be converted to an integration over a freq. variable w by means of Parseval's Theorem for Fourier Integrals, which states...

$$\longrightarrow \int_{-\infty}^{\infty} |F(t)|^2 dt = \int_{-\infty}^{\infty} |f(\omega)|^2 d\omega, \quad f(\omega) = (1/\sqrt{2\pi}) \int_{-\infty}^{\infty} F(t) e^{-i\omega t} dt. \quad (3)$$

We identify Flt) in Eq. (3) with I(t) in Eq. (2), so we can write ...

$$\frac{d\varepsilon}{d\Omega} = \int_{-\infty}^{\infty} \sigma(\omega) d\omega, \quad \sigma(\omega) = \left(\frac{\sin^2\theta}{4\pi c^3}\right) \frac{\ell^2}{2\pi} \left| \int_{-\infty}^{\infty} \dot{\mathbf{I}}(t) e^{-i\omega t} dt \right|^2. \quad (4)$$

Juckson calls the frequency-angle spectrum. There we have, as desired...

$$\frac{d^2I}{d\omega d\Omega} = \sigma(\omega) = \left(\frac{\sin^2\theta}{8\pi^2c^3}\right) \ell^2\omega^2 \left| \int_0^{\infty} I(t)e^{-i\omega t} dt \right|^2.$$
 (5)

Two details in going Eq. (4) > (5): the {F.T. of I(t)} \rightarrow iw {F.T. of I(t)}, by purtial integration (for any I(t) which vanishes as t > ± \pi); and the lower limit on the integral is put to zero [for I(t)'s which vanish @ t < 0].

(B) For the possive CRL act described, the act extres are: I=-Cv, C + FRI+ Li. It is easily verified that the solution for the current

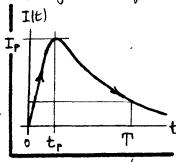
† G. Arfken "Math Methode for Physicists" (Academic Press, 3rded, 1985), Eg. 115.55). \* Since we integrate over all times, the distinction between t & t'=t-\(\frac{T}{c}\) is vanimportant. NOTE: most of the material in parts (A)-(C) is worked out in R. Robicoc & Z. Sui, J. Appl. Phys. <u>64</u>, 4364 (1988).

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I which results from the initial conditions: V(0) = Vo, I(0) = 0, is just the given:



Ilt) shows no oscillations so long as P is real, i.e. so long as CR2/L >4; this is the condition for overdamping. The small-t behavior is: IIt) ~ (Vo/2I)t, while as t > large we have:



Ilt) ~ (Vo/2IP) e-(Y-P)t. Roughly speaking: Ilt) goes torough a peak (Ip~ VolR) at time tp~ I/R (which is the "visetime"), then falls off exponentially in a cheracteristic time (i.e. duratim") [T~ RC]. The overall behavior is sketched.

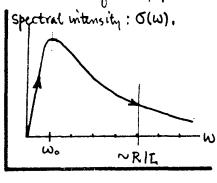
(C) For Ilt) = (Vo/LT) e-8t sinh Tt, the spectrum of Eq. (5) requires the F.J.:

$$\longrightarrow \int_{0}^{\infty} I(t) e^{-i\omega t} d\omega = \frac{V_{o}}{L\Gamma} \int_{0}^{\infty} e^{-\gamma t} \left( \frac{e^{\Gamma t} - e^{-\Gamma t}}{2} \right) e^{-i\omega t} dt = \frac{V_{o}/L}{(\omega_{o}^{2} - \omega^{2}) + 2i\gamma \omega},$$
 (7)

Where Wo= 1/TEC is the cet natural frequency. Taking the absolute square, find ...

$$O(\omega) = \left(\frac{\sin^2\theta}{8\pi^2c^3}\right) \frac{L^2 V_o^2}{R^2} \left[\frac{4\gamma^2 \omega^2}{(\omega_o^2 - \omega^2)^2 + 4\gamma^2 \omega^2}\right]. \quad (8)$$

As a for of w, the spectrum peaks @ w=wo, then falls off slowly [as ~(wo/w)2]. Beyond w= wo, the spectrum does not fall to In of its peak value until w~



2/n (R/L); if n = 10 line, if d2I/dwdn is detectable out to 10% of its peak value), then spectrum freq. range is OSW & 7(R/I).

(D) From Eq. (2), with dΩ = 2π sin θ dθ, the total arc radiation energy is ...

$$\Rightarrow \varepsilon_{rrs} = \int_{4\pi}^{9} d\Omega \int_{-\infty}^{9} r^{2} dt = \frac{\ell^{2}}{2c^{3}} \int_{0}^{\pi} \sin^{3}\theta d\theta \int_{-\infty}^{9} [\dot{I}(t)]^{2} dt = \frac{\ell^{2}}{2c^{3}} \cdot \frac{4}{3} \cdot \int_{0}^{9} [\dot{I}(t)]^{2} dt.$$

To get MKS mits, the RHS must be multiplied by (1/47160). Then, for the m pulse: I(t)=(Vo/L)e-rt sinh Pt, calculate: 50 I2dt = 1/48 (Vo/L)2, so that

$$\stackrel{\mathcal{E}}{\approx} \longrightarrow \underbrace{\mathcal{E}_{\text{red}} = \left(\frac{1}{4\pi\epsilon_{\bullet}}\right) \cdot \frac{2\ell^{2}}{3c^{3}} \cdot \frac{1}{4\gamma} \left(\nabla_{o}/\Gamma\right)^{2} = \dots, \quad \stackrel{\text{def}}{\approx} \left(\frac{2}{4\pi\epsilon_{\bullet}}\right)^{2} \cdot \frac{2\ell^{2}}{3c^{3}} \cdot \frac{1}{4\gamma} \left(\nabla_{o}/\Gamma\right)^{2} = \dots, \quad \stackrel{\text{def}}{\approx} \left(\frac{2}{4\pi\epsilon_{\bullet}}\right)^{2} \cdot \frac{2\ell^{2}}{3c^{3}} \cdot \frac{1}{4\gamma} \left(\nabla_{o}/\Gamma\right)^{2} = \dots, \quad \stackrel{\text{def}}{\approx} \left(\frac{2}{4\pi\epsilon_{\bullet}}\right)^{2} \cdot \frac{2\ell^{2}}{3c^{3}} \cdot \frac{1}{4\gamma} \left(\nabla_{o}/\Gamma\right)^{2} = \dots, \quad \stackrel{\text{def}}{\approx} \left(\frac{2}{4\pi\epsilon_{\bullet}}\right)^{2} \cdot \frac{2\ell^{2}}{3c^{3}} \cdot \frac{1}{4\gamma} \left(\nabla_{o}/\Gamma\right)^{2} = \dots, \quad \stackrel{\text{def}}{\approx} \left(\frac{2}{4\pi\epsilon_{\bullet}}\right)^{2} \cdot \frac{2\ell^{2}}{3c^{3}} \cdot \frac{1}{4\gamma} \left(\nabla_{o}/\Gamma\right)^{2} = \dots, \quad \stackrel{\text{def}}{\approx} \left(\frac{2}{4\pi\epsilon_{\bullet}}\right)^{2} \cdot \frac{2\ell^{2}}{3c^{3}} \cdot \frac{1}{4\gamma} \left(\nabla_{o}/\Gamma\right)^{2} = \dots, \quad \stackrel{\text{def}}{\approx} \left(\frac{2}{4\pi\epsilon_{\bullet}}\right)^{2} \cdot \frac{2\ell^{2}}{3c^{3}} \cdot \frac{1}{4\gamma} \left(\nabla_{o}/\Gamma\right)^{2} = \dots, \quad \stackrel{\text{def}}{\approx} \left(\frac{2}{4\pi\epsilon_{\bullet}}\right)^{2} \cdot \frac{2\ell^{2}}{3c^{3}} \cdot \frac{1}{4\gamma} \left(\nabla_{o}/\Gamma\right)^{2} = \dots, \quad \stackrel{\text{def}}{\approx} \left(\frac{2}{4\pi\epsilon_{\bullet}}\right)^{2} \cdot \frac{2\ell^{2}}{3c^{3}} \cdot \frac{1}{4\gamma} \left(\nabla_{o}/\Gamma\right)^{2} = \dots, \quad \stackrel{\text{def}}{\approx} \left(\frac{2}{4\pi\epsilon_{\bullet}}\right)^{2} \cdot \frac{2\ell^{2}}{3c^{3}} \cdot \frac{1}{4\gamma} \left(\nabla_{o}/\Gamma\right)^{2} = \dots, \quad \stackrel{\text{def}}{\approx} \left(\frac{2}{4\pi\epsilon_{\bullet}}\right)^{2} \cdot \frac{2\ell^{2}}{3c^{3}} \cdot \frac{1}{4\gamma} \left(\nabla_{o}/\Gamma\right)^{2} = \dots, \quad \stackrel{\text{def}}{\approx} \left(\frac{2}{4\pi\epsilon_{\bullet}}\right)^{2} \cdot \frac{2\ell^{2}}{3c^{3}} \cdot \frac{1}{4\gamma} \left(\nabla_{o}/\Gamma\right)^{2} = \dots, \quad \stackrel{\text{def}}{\approx} \left(\frac{2}{4\pi\epsilon_{\bullet}}\right)^{2} \cdot \frac{2\ell^{2}}{3c^{3}} \cdot \frac{1}{4\gamma} \left(\nabla_{o}/\Gamma\right)^{2} = \dots, \quad \stackrel{\text{def}}{\approx} \left(\frac{2}{4\pi\epsilon_{\bullet}}\right)^{2} \cdot \frac{2\ell^{2}}{3c^{3}} \cdot \frac{2\ell^{2$$

Here Zo= Tholeo = 377 12, and wo= 1/VIC. The total discharge energy Edis = SoRI2dt = 1 CV2 (clearly), so: Erad/Edic = (Zo/6TR)(Wol/c)2. This ratio < 10-6, typically.