

Phys. 506 Final Exam

1/28/71

1/28/71 ① "In a certain QM system, it is found that the eigenfcn  $\psi(x)$  of energy  $E$  is translationally invariant, i.e. if  $\psi(x)$  is a solution to  $H\psi = E\psi$ , then so is  $\psi(x+\Delta x)$ , where  $\Delta x$  is an arbitrary translation of the coordinate origin. Show, as a result of this, that the system momentum (operator)  $p$  must commute with the Hamiltonian  $H$ , i.e.  $[H, p] = 0$ , so that  $p$  is a constant of the motion, which in turn means  $\psi$  describes a free particle."

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Have:  $H\psi(x) = E\psi(x)$  and  $H\psi(x+\Delta x) = E\psi(x+\Delta x)$ .

Suppose  $\Delta x$  is small. Then by Taylor series expansion...

$$\psi(x+\Delta x) \simeq \psi(x) + \Delta x \left( \frac{\partial \psi}{\partial x} \right)_{\Delta x=0} = \psi(x) + \frac{i}{\hbar} \Delta x p \psi(x), \quad p = \frac{\hbar}{i} \frac{\partial}{\partial x}$$

$$\therefore 2^{nd} \text{ eq.} \Rightarrow H\psi(x) + \frac{i}{\hbar} \Delta x H p \psi(x) = E\psi(x) + \frac{i}{\hbar} \Delta x p E \psi(x)$$

$\nwarrow$  cancel  $\nearrow$

$$= H\psi(x)$$

$\therefore H p \psi(x) = p H \psi(x)$ , i.e.  $[H, p] \psi(x) = 0$

Since this is true for general  $\psi$ , must have  $[H, p] = 0$ . QED

2/3/71 ② "The wavefcn describing the motion of a free particle starting out at  $(x, t)$  and moving to  $(x', t')$  obviously depends only on the differences between initial and final coordinates. Consequently, the free particle propagator  $G_0$  is at most a fcn of  $x'-x$  and  $t'-t$ . A full Fourier integral representation of  $G_0$  must then be of the form (in 1D)

$$G_0(x'-x, t'-t) = \frac{1}{(2\pi)^2} \int_{-\infty}^{+\infty} dk \int_{-\infty}^{+\infty} d\omega g(k, \omega) e^{ik(x'-x)} e^{-i\omega(t'-t)}$$

$G(k, \omega)$  is known as the free particle propagator in momentum space. Derive an expression for  $G$ , by taking into account the fact that  $G_0$  obeys the point-source Schrödinger eqn

$$\left( i\hbar \frac{\partial}{\partial t'} - \frac{p'^2}{2m} \right) G_0 = \hbar \delta(x'-x) \delta(t'-t).$$

Operating through the Fourier integral by  $\left( i\hbar \frac{\partial}{\partial t'} - \frac{p'^2}{2m} \right)$ , we have

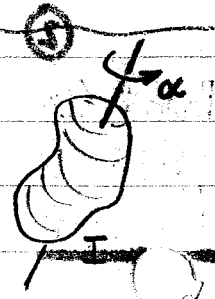
$$\begin{aligned} \frac{1}{(2\pi)^2} \int_{-\infty}^{+\infty} dk \int_{-\infty}^{+\infty} d\omega \left[ (\hbar\omega - \frac{\hbar^2 k^2}{2m}) G(k, \omega) \right] e^{ik(x'-x)} e^{-i\omega(t'-t)} &= \\ = \hbar \delta(x'-x) \delta(t'-t) &= \frac{\hbar}{(2\pi)^2} \int_{-\infty}^{+\infty} dk \int_{-\infty}^{+\infty} d\omega e^{ik(x'-x)} e^{-i\omega(t'-t)} \end{aligned}$$

The latter integral is the std rep<sup>n</sup> of  $\delta(x'-x) \delta(t'-t)$ . The only way the integrals can be identical is if  $[ ] = \hbar$ . So

$$G(k, \omega) = 1 / \left( \omega - \frac{\hbar k^2}{2m} \right) \quad \underline{\underline{QED}}$$

N.B. In order to complete the defn of  $G(k, \omega)$ , we have to decide how to go around the pole at  $\omega = \hbar k^2 / 2m$ . This is done by letting  $\omega \rightarrow \omega + i\epsilon$ .  $\epsilon \rightarrow 0_+$  gives  $\theta(t'-t)$ , and the standard (retarded) propagator, while  $\epsilon \rightarrow 0_-$  generates the advanced  $G_0$ .

- ③ a) A "plane rotator" is a rigid body constrained to rotate (with arbitrary angular momentum) about a fixed axis in space. The rotation can be specified by choosing a point on the body and giving its azimuthal  $\phi$  w.r.t. the rotation axis. Suppose the body has moment of inertia  $I$  about the rotation axis. For the QM plane rotator, solve the time-independent



Schrödinger eqn for the allowed energies  $E_m$  and <sup>normalized</sup> eigenfns  $\Psi_m(\alpha)$  of the rotation. What is the degeneracy of the state with energy  $E_m$ ?

b) Suppose, at time  $t=0$ , the rotator is in a state specified by a wavefn  $\Psi(\alpha, 0) = C \sin^2 \alpha$ , where  $C$  is a normalizing constant. What is the wavefn  $\Psi(\alpha, t)$  for  $t > 0$ ?

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a) Pick rotation axis to be  $z$ -axis of rect. co. system. Rotational energy is

$$H = L_z^2 / 2I, \quad L_z = z\text{-comp. of } \mathbf{L} \text{ momentum}$$

$$\text{QM by: } L_z = -i\hbar \frac{\partial}{\partial \alpha}, \quad \text{so } H = -\frac{\hbar^2}{2I} \frac{\partial^2}{\partial \alpha^2}$$

$$\text{S. eqn: } H\Psi = E\Psi \Rightarrow -\frac{\hbar^2}{2I} \frac{\partial^2}{\partial \alpha^2} \Psi(\alpha) = E\Psi(\alpha)$$

$$\text{or } \left( \frac{\partial^2}{\partial \alpha^2} + m^2 \right) \Psi(\alpha) = 0, \quad m = \sqrt{2IE/\hbar^2}$$

$$\text{Soln is: } \Psi(\alpha) = C e^{-im\alpha}$$

For  $\Psi(\alpha)$  single-valued over  $0 \leq \alpha \leq 2\pi$ , need  $m=0, \pm 1, \pm 2, \dots$

$\therefore$  allowed energies are:  $E_m = (\hbar^2/2I)m^2, \quad m=0, \pm 1, \pm 2, \dots$

Degeneracy of level  $E_m$  is 2 for any  $m \neq 0$ .

Normalized eigenfns are  $\Psi_m(\alpha) = \frac{1}{\sqrt{2\pi}} e^{-im\alpha}$

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b) Can now expand any  $\Psi(\alpha, t)$  in terms of  $\Psi_m(\alpha)$  by

$$\Psi(\alpha, t) = \sum_{m=-\infty}^{+\infty} c_m \Psi_m(\alpha) e^{-\frac{i}{\hbar} E_m t}$$

By orthogonality of  $\psi_m(\alpha)$ , i.e.  $\int_0^{2\pi} \psi_m^*(\alpha) \psi_n(\alpha) d\alpha = \delta_{mn}$ , have

$$\Psi(\alpha, 0) = \sum_m C_m \psi_m(\alpha) \Rightarrow C_m = \int \psi_m^*(\alpha) \Psi(\alpha, 0)$$

Now if  $\Psi(\alpha, 0) = C \sin^2 \alpha$ , we have

$$C_m = \frac{C}{\sqrt{2\pi}} \int_0^{2\pi} e^{+im\alpha} \sin^2 \alpha d\alpha$$

$$\text{Write: } \sin^2 \alpha = \left( \frac{1}{2i} (e^{+i\alpha} - e^{-i\alpha}) \right)^2 = \frac{1}{2} - \frac{1}{4} (e^{+2i\alpha} - e^{-2i\alpha})$$

$$\therefore C_m = \frac{C}{\sqrt{2\pi}} \int_0^{2\pi} \left[ \frac{1}{2} e^{+im\alpha} - \frac{1}{4} (e^{i(m+2)\alpha} + e^{i(m-2)\alpha}) \right] d\alpha$$

$$\text{But } \int_0^{2\pi} e^{ik\alpha} d\alpha = \frac{1}{ik} (e^{2\pi ik} - 1) = \begin{cases} \text{with } k \text{ an integer...} \\ 0 \text{ for } k \neq 0 \\ 2\pi \text{ for } k = 0 \end{cases}$$

$\Rightarrow$  we have  $C_m$  values only for  $m=0$  &  $m=\pm 2$ , i.e.

$$C_0 = \frac{1}{2} C \sqrt{2\pi}, \quad C_{+2} = -\frac{1}{4} C \sqrt{2\pi}, \quad C_{-2} = -\frac{1}{4} C \sqrt{2\pi}$$

So the desired  $\Psi(\alpha, t)$  is

$$\begin{aligned} \Psi(\alpha, t) &= C \sqrt{2\pi} \left\{ \frac{1}{2} \frac{1}{\sqrt{2\pi}} - \frac{1}{4} \frac{1}{\sqrt{2\pi}} \left[ e^{2i\alpha} e^{-\frac{i}{\hbar} E_2 t} + e^{-2i\alpha} e^{-\frac{i}{\hbar} E_2 t} \right] \right\} \\ &= \frac{C}{2} \left\{ 1 - \cos 2\alpha \times e^{-i(2\hbar/I)t} \right\} \leftarrow \begin{array}{l} E_2 = 2\hbar^2/I \\ \text{Agrees with Ter Haar} \\ \text{p. 103, prob. (1.34)} \end{array} \end{aligned}$$

④ The expectation value of  $1/r^2$  in the state  $|n, l, m\rangle$  of a hydrogen-like atom (potential:  $V(r) = -Ze^2/r$ ) is calculated to be

Prob. #④  
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(May 1992).

$$\langle 1/r^2 \rangle = \langle n, l, m | \frac{1}{r^2} | n, l, m \rangle = \left( \frac{Z}{a_0} \right)^2 / n^3 (l + \frac{1}{2}), \quad a_0 = \hbar^2 / m e^2.$$

Use this to show that  $\langle 1/r^3 \rangle$  in the same state is given by

$$\langle 1/r^3 \rangle = \langle nlm | \frac{1}{r^3} | nlm \rangle = \left( \frac{Z}{a_0} \right)^3 / n^3 l(l+\frac{1}{2})(l+1).$$

Hint: do not use explicit wavefunctions. Instead, look at the equation of motion for an electron in orbit.

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Eqn of motion is :  $mv^2/r = Ze^2/r^2 \Rightarrow L^2/r^3 = Zme^2/r^2$

Where  $L = mvr$  is the orbital momentum. QMly, the eqn of motion must be the same, in an expectation value sense, so we can write...

$$\langle L^2/r^3 \rangle = Zme^2 \langle 1/r^2 \rangle$$

But  $L^2 = l(l+1)\hbar^2$  w.r.t. state  $|nlm\rangle$ , So we have...

$$l(l+1)\hbar^2 \langle 1/r^3 \rangle = Zme^2 \langle 1/r^2 \rangle$$

$$\text{or } \langle 1/r^3 \rangle = \frac{(Z/a_0)}{l(l+1)} \langle 1/r^2 \rangle, \quad a_0 = \hbar^2/me^2$$

$$= \left( \frac{Z}{a_0} \right)^3 / n^3 l(l+1)(l+\frac{1}{2}) \quad \text{QED}$$

- ⑤ "Let  $\vec{A} = (A_x, A_y, A_z)$  be a QM vector operator, and consider the quantity  $\vec{I} = \psi^\dagger \vec{A} \psi$ , which is the integrand of the expectation value of  $\vec{A}$  in state  $\psi$ . Under a rotation of the coordinate system by an infinitesimal angle about any one of the coordinate axes, there are two equivalent ways to describe how  $\vec{I}$ , and hence  $\vec{A}$ , transforms. One either transforms  $\psi$ , leaving  $\vec{A}$

unchanged (i.e.  $\psi \rightarrow \psi'$ , so that  $\vec{I} \rightarrow \vec{I}' = \psi'^{\dagger} \vec{A} \psi'$ ), or one transforms  $\vec{A}$ , leaving  $\psi$  unchanged (i.e.  $\vec{A} \rightarrow \vec{A}'$ , so that  $\vec{I} \rightarrow \vec{I}' = \psi^{\dagger} \vec{A}' \psi$ ). By equating the two equivalent forms for the transformed  $\vec{I}$ , derive a commutation for  $\vec{A}$  and  $\vec{J}$ , where  $\vec{J}$  is the total system angular momentum.

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It is sufficient to consider rotation by  $\delta\phi$  about z-axis. The results for rotations about the x & y axes will follow by the usual cyclic permutation of x, y & z. The transf<sup>n</sup> of  $\psi$  is

$$\psi \rightarrow \psi' = (1 - i\delta\phi J_z) \psi, \quad J_z \text{ in units of } \hbar$$

Only  $A_x$  &  $A_y$  are affected by a z-axis rotation. Since we ultimately will end up relating  $A_x$  &  $A_y$  to  $J_z$ , it is sufficient to consider one at a time, i.e. we shall look at  $I_x$  only.  $A_x$  transforms as

$$A_x \rightarrow A'_x = A_x - \delta\phi A_y$$

Setting the two alternate expressions for  $I'_x$  equal, we find

$$\psi'^{\dagger} A_x \psi' = \psi^{\dagger} A'_x \psi$$

the two [ ] must be identical

$$\psi^{\dagger} [(1 + i\delta\phi J_z) A_x (1 - i\delta\phi J_z)] \psi = \psi^{\dagger} [A_x - \delta\phi A_y] \psi$$

$$\text{i.e. } (A_x + i\delta\phi J_z A_x) (1 - i\delta\phi J_z) = A_x - \delta\phi A_y, \text{ neglect } \mathcal{O}(\delta\phi)^2 \text{ on LHS}$$

$$-i\delta\phi (A_x J_z - J_z A_x) = -\delta\phi A_y, \text{ cancel } \delta\phi \text{ and mult. by } -i$$

$$\boxed{\Rightarrow} [J_z, A_x] = i A_y \Rightarrow \text{generalization } [J_x, A_y] = i A_z, \alpha\beta\gamma = \vec{123}$$

Had we used  $I_y$ , we would have gotten same thing. Generalization is "obvious".

- ⑥ "A hydrogen-like atom (potential:  $V(r) = -Ze^2/r$ ) is in its ground state  $|100\rangle$ , with total energy  $E_1$  given by the usual Bohr formula. Calculate the probability that the electron will be found at a distance from the nucleus greater than its energy would permit from a classical standpoint."

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Bohr ground state energy is:  $E_1 = -Z^2 e^2 / 2a_0$

Classically:  $\frac{mv^2}{r} = \frac{Ze^2}{r^2}$ , and total energy  $= \frac{1}{2}mv^2 - \frac{Ze^2}{r} = -\frac{Ze^2}{2r}$

Classical total energy  $= E_1 \Rightarrow r = a_0/Z = r_0$

Normalized ground state wavefunction is:  $\psi_{100}(r) = \left(\frac{Z^3}{\pi a_0^3}\right)^{1/2} e^{-Zr/a_0}$

Probability that e found at  $r \geq r_0$  is

$$P = \int_{r_0}^{\infty} |\psi_{100}(r)|^2 \times 4\pi r^2 dr = \frac{1}{2} \left(\frac{2Z}{a_0}\right)^3 \int_{r_0}^{\infty} r^2 e^{-\left(\frac{2Z}{a_0}\right)r} dr$$

$$\text{or } P(x_0) = \frac{1}{2} \int_{x_0}^{\infty} x^2 e^{-x} dx, \text{ with } x_0 = \frac{2Z}{a_0} r_0 = 2$$

The integral can be done by partial integrations...

$$\begin{aligned} \int_{x_0}^{\infty} x^2 e^{-x} dx &= -x^2 e^{-x} \Big|_{x_0}^{\infty} + 2 \int_{x_0}^{\infty} x e^{-x} dx = x_0^2 e^{-x_0} - 2 \left[ x e^{-x} \Big|_{x_0}^{\infty} - \int_{x_0}^{\infty} e^{-x} dx \right] \\ &= (x_0^2 + 2x_0 + 2) e^{-x_0} \leftarrow \text{Checks with Dwight, p. 127} \end{aligned}$$

$$\therefore P(x_0) = \frac{1}{2} (x_0^2 + 2x_0 + 2) e^{-x_0} \leftarrow \text{Note } P(0) = 1, \text{ as expected}$$

$$\text{Desired: } P(2) = 5e^{-2} = 0.6767$$

⑦ "An operator  $F$  depends on the position vector  $\vec{x}$  and particle momentum  $\vec{p}$  only through the combinations  $\vec{x}^2$ ,  $\vec{p}^2$  and  $\vec{x} \cdot \vec{p}$ ; that is, considered as a function of  $\vec{x}$  and  $\vec{p}$ ,  $F = F(\vec{x}^2, \vec{p}^2, \vec{x} \cdot \vec{p})$  only. Let  $\vec{L}$  be the system orbital angular momentum operator, and denote the eigenstates of  $\vec{L}^2$  and  $L_z$  by  $Y_{lm}(\vartheta, \varphi) = |lm\rangle$ .

a) Calculate the commutator bracket  $[\vec{L}, F]$ .

b) State all that can be said about the matrix elements  $\langle l'm' | F | lm \rangle$ .

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$[\vec{L}, \vec{x}^2] = 0$  since  $\vec{L}$  commutes with all sph. symmetric fns

$[\vec{L}, \vec{p}^2] = 0$  as has been shown (e.g. Merzbacher, p. 176).

So we must worry about  $[\vec{L}, \vec{x} \cdot \vec{p}]$ . Taking 1<sup>st</sup> comp...

$$[\vec{L}_1, \vec{x} \cdot \vec{p}] = \sum_{\alpha} [L_1, x_{\alpha} p_{\alpha}] = \sum_{\alpha} \{ x_{\alpha} [L_1, p_{\alpha}] + [L_1, x_{\alpha}] p_{\alpha} \}$$

Recall  $[L_i, p_j] = i\hbar p_k$ ,  $ijk = \vec{123}$ , and similarly for comps of  $\vec{x}$

(This was prob. ⑮ in PHYS. 505). Then we have...

$$\begin{aligned} [L_1, \vec{x} \cdot \vec{p}] &= \{ x_2 \underbrace{[L_1, p_2]}_{+i\hbar p_3} + \underbrace{[L_1, x_2]}_{+i\hbar x_3} p_2 \} + \{ x_3 \underbrace{[L_1, p_3]}_{-i\hbar p_2} + \underbrace{[L_1, x_3]}_{-i\hbar x_2} p_3 \} \\ &= i\hbar \{ x_2 p_3 + x_3 p_2 \} - i\hbar \{ x_3 p_2 + x_2 p_3 \} = 0 \end{aligned}$$

So  $\vec{L}$  commutes with  $F$ , i.e.  $[\vec{L}, F] = 0$ .

This means  $\vec{L} \nmid F$  have the eigenfns  $|lm\rangle$  in common. So

$$\langle l'm' | F | lm \rangle = F_{lm} \delta_{l'm'}$$

where  $F_{lm}$  is the average value of  $F$  in the state  $lm$ .