### **DEPARTMENT OF PHYSICS**

# M.S. COMPREHENSIVE / PH. D. QUALIFYING EXAMINATION NOVEMBER 26, 1984

#### DEPARTMENT OF PHYSICS

#### M.S. COMPREHENSIVE and PH.D. QUALIFYING EXAM

MONDAY, 26 NOVEMBER 1984, 8 AM-12 NOON

Answer each of the following eight (8) questions.

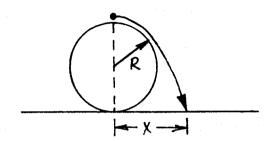
All questions are of equal weight.

Begin your answer to each question on a <u>new</u> sheet of paper. Solutions to different questions must <u>not</u> appear on the same sheet of paper.

Label each page of your answer sheets as follows:

- A. Your name in upper left-hand corner.
- B. Problem number, and page number for that problem, in upper right-hand corner.

- 1. A particle starts from rest at the top of a frictionless sphere of radius R and slides off the sphere under the force of gravity.
  - (a) How far below its starting point does it get before flying off the sphere?
  - (b) How far horizontally does it land from the sphere (x)?

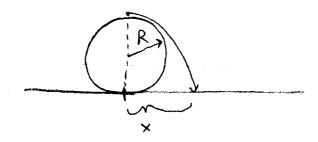


Question 1-8 sequentially one after the other with space for figures allocated when fitting several questions per page.

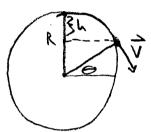
1. It A particle starts from rest at the top of a frictionless sphere of radius R and stides of the sphere under the force of gravity.

(b) How far below its starting point does it get before flying of the sphere?

(b) How far horizontally does it land from the sphere (x)?



(a) the particle can stay on the surface of the sphere only as long as the normal component of the force is greater than or equal to the man of the particle times its contripctal acceleration



we will acceleration needed for particle to contine in circular path

now take the normal component of the gravitational force?



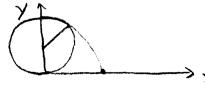
FN = mg rine

The particle will fly iff the ophere when

$$sin \theta = 2/3$$
, or  $h = R/3$ 

how For below starting point it Ayr of

(b) first lets not up a coordinate system:



The constraint of staying on the sphere's justace vanisher at the point

X = R cos0 = (5/3) R

y = 2R-h = (5/3) R V = 129R/3 => Vx = V sin G = (2 129R/3)/3

Vy = -V cos6 = - (109R/3)/3

we now we there values of x,y, vx, vy as initial conditions for the remaining unconstrained free fall of the particle:

$$may = 0$$

$$may = -g$$

$$x = x_0 + V_x^{\circ} t$$

$$y = y_0 + V_y^{\circ} t - \frac{1}{2}gt^2$$

$$y = \frac{5}{3}R - \frac{\log R}{3}(\frac{1}{3})t - \frac{1}{2}gt^2$$

ret y=0, rolve for t, relatitute into xitil to find x.

$$g t^{2} + \frac{2}{3} \left[ \frac{109R}{3} t - \frac{19}{3} k = 0 \right]$$

$$t = \frac{1}{29} \left\{ -\frac{2}{3} \left[ \frac{4 \cdot 109R}{3} + \frac{40}{3} R_{9} \right]^{2} \right\}$$

$$\left( + \frac{4 \cdot 109R}{500} + \frac{40}{3} R_{9} \right)^{2} \left\{ + \frac{4 \cdot 109R}{500} + \frac{40}{3} R_{9} \right]^{2} \right\}$$

$$t_{ground} = \frac{1}{3} \left[ \frac{10}{13} - \left[ \frac{10}{3} \right] \right] \frac{R}{9}$$

50

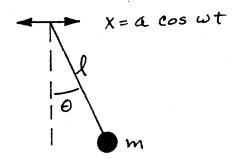
$$X = (5/3) R + \frac{2}{3} \frac{2\pi R}{3} \left[ \frac{10}{3} - \frac{10}{3} \right] \frac{R}{3}$$

$$X = (2012 + 515) R \approx R$$

2. A pendulum bob of mass m is suspended by a string of length 1 from a moving point of support. The point of support moves to and fro along a horizontal x-axis according to the equation x = acosωt.

Assume that the pendulum swings only in the vertical plane containing the x-axis. Let the position of the pendulum be described by the angle  $\theta$  which the string makes with a line vertically downward.

- (a) Set up the equation of motion for arbitrary amplitude.
- (b) Find the steady-state amplitude for small oscillations as a function of m,l,a, and  $\boldsymbol{\omega}$



2. A pendulum beb of mass m is suspended by a string of length I from a point of support moves to and for along a horizontal x-axis according to the equation  $X = a \cos \omega t$ .

N=accsut m

Assume that the pendulum swings only in the vertical plane containing the x-axis. Let the pesition of the pendulum be described by the angle of which the string makes with a line vertically downward

a) Set up the Lagrangian function for arbitrary amplitude.

6) Find the amplitude of steady-state oscillations; as a function of m, l, a, and w.

 $X = a\cos\omega t$   $X = a\cos\omega t + l\sin\omega$   $Y = l\cos\omega$   $M = a\cos\omega t + l\sin\omega$ 

 $T = \frac{1}{2}m(\hat{x}^2 + \hat{y}^2) = \frac{1}{2}m(-a\omega \sin \omega t + l\cos \theta \hat{\theta})^2$ + 12 sin 2002 = \frac{1}{2}m \frac{1}{2}awsmint #-2awleast & sin wt + 82 62 }

V = - maleose

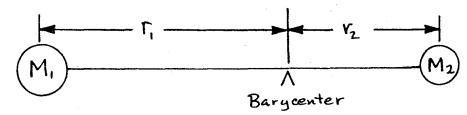
L= T-V= \frac{1}{2}m\\ a^2w^2 sin^2w^2 + \land 6 - 2awlees & sin w + \text{0} + mgs cas A

<u> Db</u> = m s² \(\hat{\theta}\) - mawl coso sin wt

36 = amulsinosinut o - malsino

ml'é + maul sui e sui vt é - mauil cord court - mand sind sind & + mgl suid = 0 5 ml² 6 + mgl suit = maw²l cose cossit b) Small Oscillations, set sin 0 = 0 Dinde by ml  $\stackrel{\circ}{\ominus} + (9/2) = a \cos \omega + (\frac{\omega^2}{2})$ ege Driven Oscillator. Duding to = awig with Dord ( phase of driver) B= 17/2 ( 8 = 0 no domping) aw2 (w2-w2) (eos 18) + aw<sup>2</sup> coswt Try state, selve for state, selve for A gives

3. If two massive bodies revolve about each other under their mutual gravitational force, they each travel along an elliptical path about their common center of mass, the so-called barycenter.



- (a) Derive Kepler's third law from Newton's laws for the special case when each orbit is a circle. Do <u>not</u> assume that the mass of one body can be neglected relative to that of the other.
- (b) Use Kepler's third law to calculate the mass of the moon given the following data:

$$G = 6.672 \times 10^{11} \text{ N-m}^2/\text{kg}^2$$

$$M(earth) = 5.977 \times 10^{24} \text{ kg}$$

$$R(earth-moon) = 384404 km$$

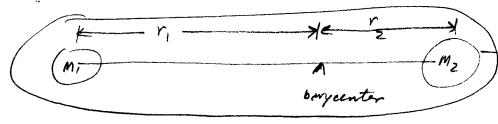
(c) How far from the center of the earth is the barycenter?

#### M.S. Comprehensive/Ph.D. Qualifying Exam

Larry D. Kirkpatrick

If two massive bodies revolve about each other under their mutual gravitational force, they each travel along an elliptical path about their common center of mass. the so-called barvcenter.

a. Derive Kepler's third law from Mewton's laws for the special case when each orbit is a circle. Do not assume that the mass of one body can be neglected relative to the other.



b. Use Kepler's third law to calculate the mass of the moon given the following data:

-11 2 2  $6 = 6.672 \times 10$  N-m /kg 24M(earth) = 5.977 × 10 kg

R(earth-moon) = 384404 km

T(siderea) of moon) = 27.322 days

c. How far from the center of the earth is the barycenter?

Equations the gravitational force to the contripetal

$$\frac{G M_1 M_2}{(r_1 + r_2)^2} = \frac{M_1 V_1^2}{r_1}$$

$$= \frac{M_1 V_1^2 r_1}{T^2}$$

$$= \frac{M_2}{(r_1 + r_2)^2} = \frac{4\pi^2 r_1}{T_2^2}$$

 $\frac{Gm_1}{(r_1+r_2)^2} = \frac{4\pi^2 r_2}{T_{\mathbf{E}}^2}$ 

Adding 
$$G(m, +m_2) = 4\pi^2(r, +r_2)$$
  
 $(r, +r_2)^2 = 7^2$ 

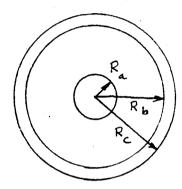
or 
$$T^2 = \frac{4\pi^2}{6(m_1 + m_2)} (r_1 + r_2)^3$$

b) 
$$m_1 + m_2 = \frac{4\pi^2}{6\tau^2} (r_1 + r_2)^3 = 6.031 \times 10^{24} \text{ kg}$$
 $m_e = 5.977 \times 10^{24} \text{ kg} \Rightarrow m_m = 0.054 \times 10^{24} \text{ kg}$ 

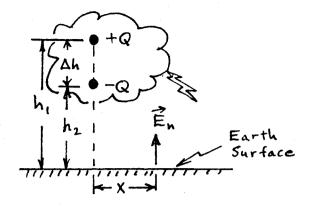
(should be 0.074 × 10<sup>24</sup> kg)

=) 
$$r_e = \frac{m_m}{m_{e+m_m}} r = 3442 \text{ km}$$
 (should be 4672 km)

- 4. An infinitely long insulating rod of radius  $R_a$  has a uniform volume charge density  $\rho > 0$ . The rod has  $\epsilon = \epsilon_0$ . The rod is inside of, and on the axis of, an infinitely long concentric conductor of inner radius  $R_b > R_a$  and outer radius  $R_c$ . The conductor has a charge density per unit length of  $\lambda > 0$ .
  - (a) Find the electric field vector for all r, where r is measured from the axis of symmetry.
  - (b) Find the values of the surface charge densities on the inner and outer surfaces of the conductor in terms of  $\rho$  and  $\lambda$ .



5. Inside a thunderstorm, a separation of charge is generated, with charge + Q at altitude  $h_1$ , and (-) Q at  $h_2 < h_1$ . Assume the sizes of these charge concentrations are small relative to:  $h_1$ ,  $h_2$  and  $\Delta h = h_1 - h_2$ . Also, assume the earth's surface can be approximated by a conducting plane. Measure distance x along the surface, with x = 0 directly below the  $\pm$  Q charges.



- (a) Find the electric field  $\boldsymbol{E}_n$  normal to the earth's surface at distance x from the storm.
- (b) Calculate the charge density induced on the surface directly below the storm (x = 0).
- (c) Let  $h_{1,2} = h \pm 1/2 \Delta h$ , with  $\Delta h < h$ . Find the x-value for which  $E_n$  vanishes. Neglect terms of order  $(\Delta h/h)^2$ .

# E&M: Electrostatics

6 Nov. 84

h + 2-a > 3

h2 En earth Surface

Inside a thunderstorm, a separation of charge is generated, with charge + Q at altitude  $h_1$ , and (-) Q at  $h_2 < h_1$ . Assume the size of these charge concentrations are small relative to:  $h_1$ ,  $h_2$  and  $\Delta h = h_1 - h_2$ . Also, assume the lattice surface can be approximated by a conducting plane.

Measure distance x along the surface, with x=0 directly below the ± Q charges.

A. Find the electric field En normal to the earth's surface at distance & from the storm.

B. Calculate the Charge density induced on the surface directly below the storm (x=0).

C. Let  $h_{1,2} = h \pm \frac{1}{2} \Delta h$ , with  $\Delta h \ll h$ . Find the x-value for which En vanishes. Neglect terms of order  $(\Delta h/h)^2$ .

Solution: Use method of images. +Q has image -Q at distance hy below plane; -Q has image +Q at hz below plane; these pairs ensure plane = equipotential. +Q  $\frac{1}{2}$   $\frac{1}{2$ 

high Similarly /  $E_n(y-Q) = + \frac{2kQh_2}{\gamma_2^3}$ ,  $\gamma_2^2 = h_2^2 + \chi^2$ 

A. Net normal field is: En = Enlby+Q)+ Enlby-Q)...

 $E_{n} = 2kQ\left(\frac{h_{2}}{r_{z}^{3}} - \frac{h_{1}}{r_{1}^{3}}\right) = 2kQ\left[\frac{h_{2}}{(h_{2}^{2} + \chi^{2})^{3/2}} - \frac{h_{1}}{(h_{1}^{2} + \chi^{2})^{3/2}}\right]$ 

B. Field directly below storm (x=0) is ...

 $E_{no} = 2kQ\left(\frac{1}{h_2^2} - \frac{1}{h_1^2}\right) = \frac{2kQ}{h_1^2h_2^2}\left(h_1^2 - h_2^2\right) = 4kQ\frac{h\Delta h}{h_1^2h_2^2} > 0.$ 

The charge density is:  $\sigma_0 = \frac{Q}{E_{no}/4\pi k} = \frac{Q}{2TT} \left( \frac{1}{h_2^2} - \frac{1}{h_1^2} \right) > 0$ . Both  $\sigma_2 \in E_n$ 

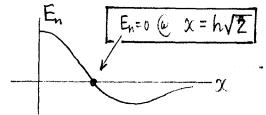
C. The denominators appearing in part A are ...

Sup  $\left(\frac{2}{h_{1,2}+\chi^2}\right)^{-\frac{3}{2}} \simeq \left(\frac{1}{h_{+}^2+\chi^2}\right)^{-\frac{3}{2}} \left[1 + \frac{3}{2} \frac{h\Delta h}{h_{+}^2+\chi^2}\right]$ , by Binomial Expansion

Use this in En of part A to find ...

$$E_{n} \simeq \frac{2 k Q}{(h^{2} + \chi^{2})^{3/2}} \left[ (h - \frac{1}{2} \Delta h) \left( 1 + \frac{3}{2} \frac{h \Delta h}{h^{2} + \chi^{2}} \right) - \left( h + \frac{1}{2} \Delta h \right) \left( 1 - \frac{3}{2} \frac{h \Delta h}{h^{2} + \chi^{2}} \right) \right]_{-}^{-}$$

 $E_n \simeq \frac{2kQ\Delta h}{(h^2+\chi^2)^{3/2}} \cdot \left(\frac{2h^2-\chi^2}{h^2+\chi^2}\right), \text{ to } \Theta\left(\frac{\Delta h}{h}\right).$ 



For large x, this  $\sim$  dipole field, as it must. En >0 for  $0 \le x < h\sqrt{2}$ , En <0 for  $h\sqrt{2} < x \rightarrow \infty$ , and the cross-over is  $@x = h\sqrt{2}$ , as shown.

6. Find the atmospheric pressure p as a function of altitude z above the earth's surface on the assumption that the temperature, T, decreases with altitude, according to  $T = T_0$  (1-az). At the surface, z = 0,  $p = p_0$ , and  $T = T_0$ .

Maybe too easy for Comp ? Fall 84

6.

Find they atmospheric pressure, as a function of altitude, above the earth's surface on the assumption that the temperature, T, decreases with altitude, according to T= To (1- xz). At the surface, z=0 P=Po and T=To.

Take T=To (1-02) + treat air as ideal gos

PV = RT or  $g = \frac{MP}{V} = \frac{MP}{RT}$ 

Static equil. => F= \( \bar{f} = \bar{g} \)  $\frac{dp}{dz} = -gg = -\frac{Mg}{gz} P$ 

 $\frac{dP}{P} = -\frac{Mq}{RT} \frac{dz}{(1-xz)}$ 

Integrale from Po -> P

het u= 1- x = en (P/P0) = - Mg (1-02) du = - xd= = Mg lu (1- az)

=> \ P = Po (1- xz) Mg/2RTo

7. Evaluate the integral  $I = \int_{0}^{\pi} \frac{\cos\theta \ d\theta}{5 - 4\cos\theta}$  by the contour method.

7. Evaluate the integral  $I = \int_{0}^{T} \frac{\cos \theta}{5 - 4\cos \theta}$ by contain method

Soln: Note  $\cos\theta$  is even  $f(n, of \theta)$ , so  $I = \frac{1}{2} \int_{0}^{2\pi} \frac{\cos\theta}{5 - 4\cos\theta} d\theta$ 

Then let  $z=e^{i\theta}$ , so I becomes a contour integral about

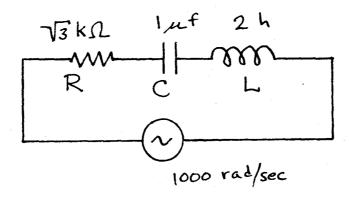
the unit circle:  $dz = ie^{i\theta}d\theta = izd\theta, \text{ or } d\theta = -i\frac{dz}{3}$   $\cos\theta = \frac{i}{2}\left(z + \frac{i}{3}\right)$ 

 $I = \frac{1}{2i} \oint \frac{d3}{3} \frac{\frac{1}{2}(3^{\frac{1}{2}})}{5 - 2(3^{\frac{1}{2}})} = -\frac{1}{4i} \int \frac{d3}{3} \frac{(3^{2}+i)}{3(23^{2}-53+2)}$ 

 $= -\frac{1}{8i} \oint \frac{dz}{3} \frac{(3^{2}+1)}{3(3-2)(3-\frac{1}{2})} = -\frac{2\pi i}{8i} \underbrace{\begin{cases} \frac{+1}{(-2)(-\frac{1}{2})} + \frac{(\frac{(5)^{2}+1)}{2}}{(\frac{1}{2})(-\frac{3}{2})} \end{cases}}_{(\frac{1}{2})(\frac{-3}{2})}$ 

 $= -\frac{\pi}{4} \left\{ 1 - \frac{5}{3} \right\} = +\frac{\pi}{6}$ 

- 8. You are given an L,C,R, series circuit connected to a generator operating at an angular frequency of 1000 radians per second. L is two Henries, C is one microFarad and R is  $\sqrt{3}$  kiloOhm.
  - (a) Find the impedance seen by the generator.
  - (b) If the generator current is given by  $I_0 \cos \omega t$ , what is the phase angle  $\phi$  of the voltage across the generator?
  - (c) If the generator frequency is allowed to vary, the circuit exhibits a resonance behavior. Discuss the properties of resonance in this circuit, e.g. frequency, voltage, current, phase, width, etc.



8. Your form on L, C, R, series circuit connected to generation operating at an angular frequency of 1000 radius per second, I in two Henries, C your microfassel and Ris 73 Addison. a) Fine the impedance seen by the generator b) the that the generator current is give by I could what is the place angle of the vollage across the generales?

(2) If the generales presency is allowed to very the circuit I shibits a resonance behavior. Dus cur the properties of resources, ago francis, V3KR IMF 2h m-11-3300-0/000 rad/ sie

a) 
$$d = 10^{3} \cdot 10^{6} = 1 \cdot 10^{3} \cdot$$

#### DEPARTMENT OF PHYSICS

#### M.S. COMPREHENSIVE and PH.D. QUALIFYING EXAM

#### MONDAY, 26 NOVEMBER 1984, 1 PM-5PM

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- A. Your name in upper left-hand corner.
- B. Problem number, and page number for that problem, in upper right-hand corner.

- 9. (a) Consider a system of spin 1/2. What are the eigenvalues and normalized eigenvectors of the operator  $AS_y + BS_z$ , where  $S_y$ ,  $S_z$  are the angular momentum operators, and A and B are real constants?
  - (b) Assume that the system is in a state corresponding to the upper eigenvalue. What is the probability that a measurement of  $S_y$  will yield the value + 15/2?

The Pauli matrices are

$$\sigma_{x} = \begin{pmatrix} 0 & 1 \\ 1 & 0 \end{pmatrix} \qquad \sigma_{y} = \begin{pmatrix} 0 & -i \\ i & 0 \end{pmatrix} \qquad \sigma_{z} = \begin{pmatrix} 1 & 0 \\ 0 & -1 \end{pmatrix}$$

Secretary (a) 
$$AS_{g} + BS_{g} = \frac{1}{2} \begin{pmatrix} B & -iB \\ iB & -B \end{pmatrix}$$
 Find a vial after  $S_{g} = \frac{1}{2} \begin{pmatrix} B & -iB \\ iB & -B \end{pmatrix}$  Find a vial after  $S_{g} = \frac{1}{2} \begin{pmatrix} B & -iB \\ iB & -B \end{pmatrix}$   $S_{g} = \frac{1}{2} \begin{pmatrix} A_{g} \\ A_{g} \end{pmatrix}$ . Solve reg (1 (1 coeffs =  $a$ ).

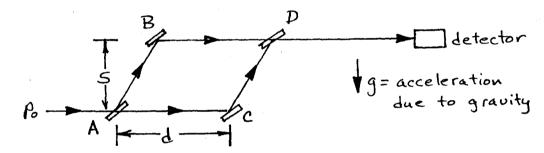
By  $-iB_{g} = \lambda d_{g}$   $d_{g} = \frac{1}{2} \begin{pmatrix} A_{g} \\ A_{g} \end{pmatrix}$ . Solve reg (1 (1 coeffs =  $a$ ).

By  $-iB_{g} = \lambda d_{g}$   $d_{g} = \frac{1}{2} \begin{pmatrix} A_{g} \\ B_{g} - \lambda^{2} + B^{2} = 0 \end{pmatrix}$   $A_{g} = \frac{1}{2} \begin{pmatrix} B_{g} + B_{g}^{2} \end{pmatrix}^{1/2}$ 
 $A_{g} = \begin{pmatrix} A_{g} \\ A_{g} \end{pmatrix} d_{g}$  Normalize;  $A_{g} = \frac{1}{2} \begin{pmatrix} A_{g} \\ B_{g} - A_{g} \end{pmatrix} d_{g} = \frac{1}{2} \begin{pmatrix} A_{g} + A_{g} + A_{g} \\ B_{g} - A_{g} + A_{g} \end{pmatrix} d_{g} = \frac{1}{2} \begin{pmatrix} A_{g} + A_{g} + A_{g} \\ A_{g} + A_{g} + A_{g} \end{pmatrix} d_{g} = \frac{1}{2} \begin{pmatrix} A_{g} + A_{g} + A_{g} \\ A_{g} + A_{g} + A_{g} \end{pmatrix} d_{g} = \frac{1}{2} \begin{pmatrix} A_{g} + A_{g} + A_{g} \\ A_{g} + A_{g} + A_{g} \end{pmatrix} d_{g} = \frac{1}{2} \begin{pmatrix} A_{g} + A_{g} + A_{g} \\ A_{g} + A_{g} + A_{g} \end{pmatrix} d_{g} = \frac{1}{2} \begin{pmatrix} A_{g} + A_{g} + A_{g} \\ A_{g} + A_{g} + A_{g} \end{pmatrix} d_{g} + \frac{1}{2} \begin{pmatrix} A_{g} + A_{g} + A_{g} \\ A_{g} + A_{g} + A_{g} \end{pmatrix} d_{g} + \frac{1}{2} \begin{pmatrix} A_{g} + A_{g} + A_{g} + A_{g} \\ A_{g} + A_{g} + A_{g} \end{pmatrix} d_{g} + \frac{1}{2} \begin{pmatrix} A_{g} + A_{g} + A_{g} + A_{g} + A_{g} \end{pmatrix} d_{g} + \frac{1}{2} \begin{pmatrix} A_{g} + A_{g} + A_{g} + A_{g} + A_{g} \end{pmatrix} d_{g} + \frac{1}{2} \begin{pmatrix} A_{g} + A_{g} + A_{g} + A_{g} \end{pmatrix} d_{g} + \frac{1}{2} \begin{pmatrix} A_{g} + A_{g} + A_{g} + A_{g} \end{pmatrix} d_{g} + \frac{1}{2} \begin{pmatrix} A_{g} + A_{g} + A_{g} + A_{g} \end{pmatrix} d_{g} + \frac{1}{2} \begin{pmatrix} A_{g} + A_{g} + A_{g} + A_{g} \end{pmatrix} d_{g} + \frac{1}{2} \begin{pmatrix} A_{g} + A_{g} + A_{g} + A_{g} \end{pmatrix} d_{g} + \frac{1}{2} \begin{pmatrix} A_{g} + A_{g} + A_{g} + A_{g} \end{pmatrix} d_{g} + \frac{1}{2} \begin{pmatrix} A_{g} + A_{g} + A_{g} + A_{g} \end{pmatrix} d_{g} + \frac{1}{2} \begin{pmatrix} A_{g} + A_{g} + A_{g} + A_{g} \end{pmatrix} d_{g} + \frac{1}{2} \begin{pmatrix} A_{g} + A_{g} + A_{g} + A_{g} \end{pmatrix} d_{g} + \frac{1}{2} \begin{pmatrix} A_{g} + A_{g} + A_{g} + A_{g} \end{pmatrix} d_{g} + \frac{1}{2} \begin{pmatrix} A_{g} + A_{g} + A_{g} + A_{g} \end{pmatrix} d_{g} + \frac{1}{2} \begin{pmatrix} A_{g} + A_{g} + A_{g} + A_{g} \end{pmatrix} d_{g} + \frac{1}{2} \begin{pmatrix} A_{g} + A_{g} + A_{g} \end{pmatrix} d_{g} + \frac{1}{2} \begin{pmatrix} A_{g} + A_{g} + A_{g} + A_{g} \end{pmatrix} d_{g} + \frac{1}{2} \begin{pmatrix} A_{g} + A_{g} + A_{g} + A_{g} \end{pmatrix} d_{g} + \frac{1}{2} \begin{pmatrix} A_{g} + A_{g} + A_{g} + A_{g} \end{pmatrix} d_{g} + \frac{1}{2} \begin{pmatrix} A_{g} + A_{g} + A_{g} + A_{g$ 

 The Colella-Overhauser-Werner experiment (Phys. Rev. Lett. 34, 1472(1975)) measures the effects of gravitation on quantum interference.

A beam of nonrelativistic neutrons with momentum po and mass M is split and Bragg reflected through a carefully prepared silicon crystal. The neutrons in the upper beam, ABD, will have a different wavelength over the path BD than the neutrons in the lower beam AC. The difference is due to the difference in gravitational potential between the paths BD and AC. The crystal has been constructed so that the path length ABD is equal to the path length ACD. (The paths ABD and ACD are within the Si crystal.)

During the experiment, interference fringes are observed with the crystal held at various angles to the vertical. When B and C are at the same height (all paths horizontal) the paths ABD and ACD are indistinguishable and there is no phase shift. When the crystal is rotated  $90^{\circ}$  about the incident beam axis so that ABCD lies in the vertical plane, the phase shift between the two beams is maximized.



- (a) Calculate the phase shift,  $\Delta \phi$ , between the two beams when the crystal is oriented vertically (as shown above), as measured when the beams recombine at D. You may assume that  $p_0^2/2M \gg Mgs$ .
- (b) If  $M = 1.67 \times 10^{-27} \text{kg}$ ;  $g = 9.8 \text{ m/sec}^2$ ;  $M = 1.05 \times 10^{-34} \text{ J. sec}$ ; d = 3 cm, s = 2 cm, and  $p_0 = 4 \times 10^{-24} \text{ kg m/sec}$  (i.e.,  $v_0 \sim 10^3 \text{ m/sec}$ ), then how big is  $\Delta \phi$ ? Is this an observable shift in an interference pattern?

Note: ignore the bending of the neutron beams caused by ''g''
(This is a higher order effect).

### ucurelativistic

13) The Colella - Overhauser - Werner experiment (Phys. New Lett. 34, 1472 (1975))
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about the incident  $\sqrt{g} = acceleration due to gravity$ 

(a) Calculate the phase shift,  $\Delta \phi$ , between the two beaut when the crystal is oriental vertically (as shown above), as measured when the beaut recombine at D. You may assure that  $\rho_0^2/2M >> Mgs$ 

(b) If M=1.67×10-27kg; g=9.8 m/rec2; t=1.05×10-34 J. sec; d=3cm S=2cm, and po=4×10-24kg m/rec (ie., Vo~103m/rec), then how big is Dp? Is this an observable rhift in an interference pattern?

Note: ignore the bending of the nextra beaut caused by "g" (This is a higher order effect).

(a) nectrons injected with p=po 
$$\lambda = \frac{h}{po}$$

along upper path, the nextrus kinetic energy is reduced

$$\Rightarrow b_{5}^{-1} = b_{5}^{0} - 5 L_{5}^{2} ds \qquad bab = b \cdot \left[1 - \frac{5 L_{5}^{2} s}{5 L_{5}^{2} s}\right]$$

5. 
$$\lambda_{up} = \frac{h}{\rho_0 \sqrt{1 - \frac{2M_2^2 f}{\rho_0^2}}} = \frac{\lambda_0}{\sqrt{1 - \frac{2M_2^2 f}{\rho_0^2}}}$$

The total phase shift will be ap = (lup-lo) . 2 . 2 ...

$$\Delta \phi = \left[ \frac{1}{1 - \frac{2m^2 g \Gamma}{\rho_0^2}} - 1 \right] \cdot \frac{\rho_0 d}{t_0}$$

or, to first order in Mgs = legit rina for >> Mgs

$$\Delta \phi = \frac{m_g^2}{t_{po}} \ d.s$$

$$\frac{16}{5} \frac{m^2 g}{k_{10}} \lambda.s = ad$$

$$\frac{(1.67\times10^{-27})^2(9.8)(.03)(.02)}{(1.05\times10^{-34})(4\times10^{-24})} = 39 \text{ radians}$$

or aN = 
$$\frac{\Delta \Phi}{2\pi}$$
 ~ 6.2 fringer

- 11. A 50 Gev electron hits a proton at rest. Assume a head-on collision.
  - (a) What is the total energy in the center of mass frame?
  - (b) Assume further that the collision is perfectly elastic. What are the final energies of the electron and the proton in the lab frame?

11. Q. A 50 Go Delectron hits a proton exclision. Assume a head on

a) What is the total energy in the C.M. frame?

Center of mass

b) Assume further that the collision is perfectly clastic. What are the final energies of the electron and the proton in the lab frame.

A. Betore the colling. the momentum

A vectors are:

PR = (0, Eo/c)

Pe = (pmeV, pmec) = (E/c, E/c) E = 50 GeV

Parts of Pr. Pa die oguel and opposites

Po = Tole (-shx, chx)

Pe= E/c (-shx, chx) Pe= E/c (cx, e-x)

 $Ee^{x} = Eoshx \approx Eo(\pm e^{x})$   $e^{x} = \sqrt{2E} \approx 10$   $D' = (50)(\pm e) = 500 \text{ We for each}$ 

Pe' = (50)(to) = sur V/cfor conte particle.
Total Memory = (10 GeV)

1) After the collision, the space marte et Pripe change sign: 中 = Eo/c (+ shx, shx)

Now transform back, using yard:

Pc"= E/c(-e-2x, e-2x)

To"= Eole (Zohxchx, ch'x+sh'x)

The energies are:

Ee" = (so)(to)2 = 0.5 GeV (now morning)

Ep" = (1) (+) (100) = 50 GeV (actually 49.5)

Note at these ultra relativistic energies, Missime no longer matters! The particles a billiard we get a billiard balleffect in which the particles balleffect in which the particles trade energy.

12. Useful integral 
$$\int_{0}^{\infty} e^{-x^{2}} dx = \pi/2$$

A box has a very dilute gas of  $U^{235}$  and  $U^{238}$  atoms. Particle densities are  $n_{235}$  and  $n_{238}$  respectively. Mass densities are  $\rho_{235}$  and  $\rho_{238}$  respectively. The gas is at temperature T. A small hole of area S opens into a vacuum. The total rates of escape  $N_{235}$  and  $N_{238}$  of  $U^{235}$  and  $U^{238}$  atoms are given by

- (a) Use dimensional analysis to find the dimensions of the f factor.
- (b) What is the ratio of  $f_{235}/f_{238}$ ?
- (c) What is the ratio of  $N_{235}/N_{238}$ ?
- (d) Evaluate f using a kinetic theory model in which the probability of an atom having speed v is the Boltzmann distribution

$$P(V) = N \exp \left(-\frac{Mv^2}{2kT}\right) v^2$$

N is a normalization constant. M is the mass of an atom, k is Boltzmann's constant.

#12 Useful Integral fe-xdx = JTT/2 A box has a very dilute gas of U<sup>235</sup> and U<sup>238</sup> atoms. Particle densities are N235 and N238 respectively. Man densities are firs and first respectively of the gas is at temperature T. A small hole of area . S opens into a vacuum. The rate of lescope of 4235 and 4238 atoms is given by N<sub>235</sub> = S N<sub>235</sub> f<sub>235</sub> atoms/sec N<sub>238</sub> = S N<sub>238</sub> f<sub>238</sub> atms/nec 1) Use demensional analysis the dimensions of the factor. (2) What is the ratio of fr35/fr38? 3) What is the rater of N235/N238?

4 Evaluate + using a Kingles theory model in which the probability Jan atom having speed V is. the Boltzman distribution P(v) = N(e-Mv) V2 N is a riomaly atem constant. M is the man of an atom, k is Boltzmans constant.

$$N = \frac{1}{T} \quad A = L^2 \quad N = \frac{1}{L^3}$$

$$f = L/T \quad a \quad \text{velocit}. \quad \text{If}$$

$$\text{out he projected to toyed scale of}$$

$$\text{thank speed ungar} \quad f \quad \sqrt{\frac{kT}{M}}$$

$$\text{so is } \quad N \quad N/M \quad M = \frac{P}{N}$$

$$N \sim \frac{N^{3/2}}{P^{1/2}} = \frac{N_{235}}{N_{235}} = \frac{N_{235}}{N_{235}} = \frac{N^{3/2}}{N_{235}} =$$

$$= \frac{An\pi}{8} \nabla \qquad \qquad \int \frac{\pi kT}{8} \nabla = \frac{\pi kT}{8M}$$

$$= -\frac{d}{d\alpha} \int v dv e^{-\alpha v^2} dv \qquad \qquad \int \frac{e^{-\alpha v^2} v^3 dv}{e^{-\alpha v^2} v^2 dv} = \frac{8kT}{\pi M}$$

$$= -\frac{d}{d\alpha} \int v dv e^{-\alpha v^2} dv \qquad \qquad \int \frac{e^{-\alpha v^2} v^2 dv}{e^{-\alpha v^2} v^2 dv}$$

13. A one-dimensional harmonic oscillator of mass m and natural frequency  $\boldsymbol{\omega}$  is described at t=0 by the wave function

$$\psi(x,0) = (u_0+u_1)/\sqrt{2}$$

where

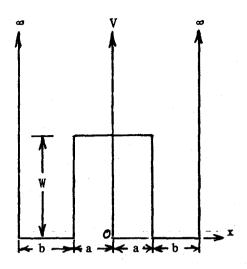
$$u_0 = -\frac{1}{\pi^{1/4} \Lambda} e^{-x^2/2\Lambda^2}$$

$$u_1 = \sqrt{2} (\frac{x}{\Lambda}) u_0$$

$$\Delta = \sqrt{\hbar/m\omega}$$

Both  $u_0$  and  $u_1$  are normalized. For t>0, compute  $\langle E \rangle$  and  $\langle x \rangle$ .

J. Hermanson A me-dimensional harmonic oscillator \( \) is described out t=0 by the wavefunction  $\Psi(x,0) = (u_0 + u_1)/\sqrt{2}$ where  $u_0 = \frac{1}{\pi^{14}\sqrt{\Delta}} e^{-\chi^2/2\Delta^2}$  $u_1 = \sqrt{2} \left( \frac{x}{4} \right) u_0$ 1 = Vt/mw Both u, and u, are normalized. For t>0, compute (E) and (x). Soln: 4(x,t)= (u0 e-iwt/2 + u, e-3 iwt)/1/2 = auo+bu, (E)= 1a12 km + 16/2 3 km =  $\left(4+\frac{3}{4}\right)\frac{\hbar\omega}{2} = \hbar\omega$  [constant!] (x)= (a\*b+b\*a) (u, x)u, dx [Since  $\int u_0^2 x dx = \int u_1^2 x dx = 0$ ] = = = (e-inteint) \ \int \ u, 2 dx  $\left[\int u_{i}^{2}dx=1\right]$ 2 coswt Foscillates at classical Seq. ]

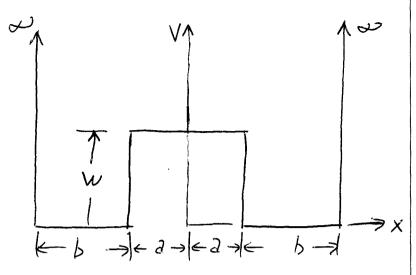


A particle of mass m is in this 1-d well.

- a) Find the barrier height W such that the ground state energy  $E_{_{\mbox{\scriptsize O}}}=$  W.
- b) Find and describe the normalized ground state wave function for all x.

#14

Schmidt: Quantum Machanies



A particle of mass m is in this 1-d well.

a) Find the barrier height w such that the ground state energy Eo = W.

b) Describe the ground state wave function for all X.

## Telmidt: Quantum Madanics Seletion

a) Ground state wave function must have gero corvative, be egumetric, and have no zero chossings:

-th dry + V4 = E4 for V=0, 4 is sinusuidal:

4 = A sin & X' (X'= X+2+6)

 $t\frac{t^2 d^2}{2m} = E_0 = W$ 

4 (x)

Also, to be flat,  $db = \frac{\pi}{2}$ ,  $d = \frac{\pi}{2b}$ 

 $\frac{\hbar^2 H^2}{8mb^2} = E_0 = W$ 

b)  $\int 4^{4}4 dx = 1$   $A^{2}b + 2A^{2}d = 1$   $A = \sqrt{\frac{1}{b+2a}}$ 

$$-2-b \le x \le -2$$

$$-2 \le x \le 3 + 6$$

$$\forall = A \sin \left[ \chi \left( x + 3 + 6 \right) \right]$$

$$-2 \le x \le 3 + 6$$

$$\forall = A \sin \left[ \chi \left( 2 + 6 - x \right) \right]$$

#### 15. Sodium is an alkali metal with z=11.

- a) What is the electron configuration of the ground state of an isolated sodium atom?
- b) What are the L,S, and J quantum numbers of the state? How are they expressed in spectroscopic notation?
- c) What is the most probable first excited state? What J values are possible for it?

#15.

Sodium i, an alkali metal with Z=11. Samisolated a) What is the electron configuration of the ground state nation?

b) What are the L, 5, and J quantum numbers of the

state? How are they expressed in spectroscopic notation?

c) What is the most probable first excited state? What I values are possible for it

A. (a) (15)2 (25)2 (2p) (35)

(b) n=3  $^{2}Sy_{2}$   $S=\frac{1}{2}$ , L=0,  $J=\frac{1}{2}$ 

2 Pi 2 P3/2 (c)  $(3p) = 1 = 1 \Rightarrow J = \frac{1}{2}$ 

- 16. Consider an optical fiber with refractive indices  $n_{core}=1.552$  and  $n_{cladding}=1.550$ .
  - a) Calculate the critical angle for propagation and illustrate your results with a carefully labeled sketch.
  - b) For 10 km of this fiber, calculate the difference in propagation time between a ray propagating along the optic axis and one taking the longest path inside the fiber.
  - c) Comment on the significance of your result in part b for optical fiber communications at high pulse rates.
  - d) What other effect could limit the performance of an optical fiber system using short pulses?

#!6.	Consider an opteral fiber with refractive indices
	More = 1,552 and notadding = 1,550.
	a) Calculate the critical angle for propagation and illustrate b) For 10km of this fiber, calculate the difference camp in propagation time between a ray propagating whete
	b) For 10 km of this fiver, calculate the diffice co couplings
	along the optic axis and one taking the longest
	path inside the fiber.
	c) Comment on the significance of your result in
	part b for optical filer communications at
	high pulse rates.
	d) What other effect could limit the performance
	of an optical fiber system using short pulses?
A	v v
	Oc = 87°
	Oc-01 1/////// clad oc=87° core
	9c= 81
	5) l'err 3°=10 km
	*
Ĺ	$2l=l'-l=13  m \qquad n \stackrel{\text{d}}{=} = 70 \text{ nsec}$
	c) Pulses must be at least this, for apart or the
	Real systems would be seriously limited by this design,
	d) Chromatic dispersion arise from da

Since short pulses contain a spread of is, this also leads to pulse stretching.