DEPARTMENT OF PHYSICS PH. D. COMPREHENSIVE EXAMINATION SEPTEMBER 24-25, 1984

DEPARTMENT of PHYSICS

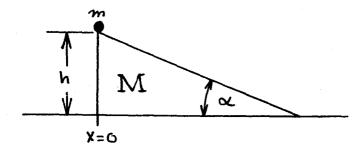
PH.D. COMPREHENSIVE EXAMINATION

MONDAY, SEPT. 24, 1984, 9-12 AM

Answer each of the following questions. Each question carries equal weight. Begin your answer to each question on a <u>new</u> sheet of paper; solutions to different questions must <u>not</u> appear on the same sheet of paper. Label <u>each page</u> of your exam as follows:

- A. Your name in upper left-hand corner.
- B. Problem number, and page number for that problem, in upper right hand corner.

1. A point particle of negligible size, and mass m, slides without friction down the inclined plane shown below. The incline also slides without friction on the table and has mass M. The incline angle is α and the side height is h.



If the particle starts at the top of the incline at t=0, and the left edge of the incline is initially at X=0, as shown, at what point on the table does the particle actually hit the table?

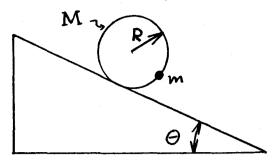
1. Particle ou sliding, inclined plane Where does particle actually hit the table? Describe position of incline as X, and position of mass as: X = X + scosx where s is position of m' along incline (note 2 degrees of freedom) X = X + S CESX hagrangian b= T-V - ssina = 1/2 (M+m) x2 + 1/2 m s2 + m x s cosx V = mg (h-ssina) L= M-V= 1/2 (M+m) X2+1/2ms2+mXscosx-mg(h-ssinx) X is cyclic Kignorable), so Px = constant = (M+m) X + mscosa

is cyclic (ignorable), so

Px = constant = (M+m) \(\bar{X} \) + m \(\scale \) \(\scale \) \(\bar{X} = -\frac{m\cos \alpha}{(M+m)} \) \(\bar{S} \)

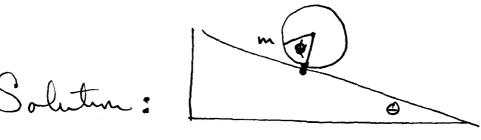
Now find s(t): dt (DE) - DE = 0

2. A hollow cylinder of radius R and mass M is loaded at one point by a mass m. (That is, m is firmly attached to the cylinder.) The cylinder is on an inclined plane of angle θ .



Beyond what critical angle of the inclined plane θ_c will the cylinder roll down the hill regardless of its initial orientation? You may assume the cylinder rolls without slipping. For $\theta < \theta_c$, find the frequency of small rolling oscillations of the cylinder.

* * * * * * * * * * *



& is angle of roll relative to me beig in contact with plane

$$V(\phi) = -Mg sm \theta R \phi$$

$$+ mg [sm \theta R \phi + R[cn(\theta + \phi) - cn \theta]]$$

$$\frac{\partial V(\phi)}{\partial \phi} = 0 = (M+m)gRsmo + mgRsm(0+\phi)$$

$$Sm(\theta+\phi) = \frac{M+m}{m} sm\theta \leq 1$$

$$T = \frac{1}{2} M R^{2} \dot{\phi}^{2} + \frac{1}{2} M R^{2} \dot{\phi}^{2} + T_{m}$$

$$T_{m} = \frac{1}{2} m R^{2} \dot{\phi}^{2} \left[2 - 2 \cos \phi \right]$$

$$T = \left[M + m(1 - \cos \phi) \right] R^{2} \dot{\phi}^{2}$$

$$V'' = mgR cn(\theta+\phi_0) \qquad sm(\theta+\phi_0) = |\underline{M+m}|sm\theta_0$$

$$TM = 2 [\underline{M+m(1-m\phi_0)}] R^2$$

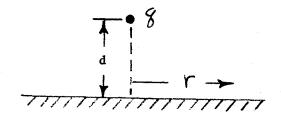
$$W^2 = \frac{g}{R} \frac{m}{2(\underline{M+m(1-m\phi_0)})} \frac{cn(\theta+\phi_0)}{cn(\theta+\phi_0)}$$

$$Who sm(\theta+\phi_0) = (\underline{M+m}|sm\theta_0) = \frac{sm\theta_0}{sm\theta_0}$$

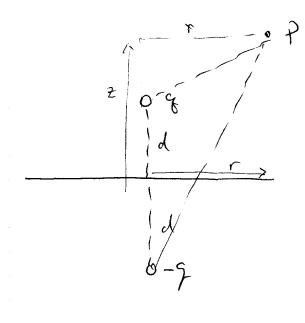
$$= \frac{g}{R} \frac{m}{2(\underline{M+m(1-cn\phi_0)})} \frac{\sqrt{sm^2\theta_0 - sm^2\theta_0}}{sm\theta_0}$$

$$= \frac{g}{R} \frac{M+m}{2(\underline{M+m(1-m\phi_0)})} \frac{\sqrt{sm^2\theta_0 - sm^2\theta_0}}{sm\theta_0}$$

3. A point charge q is placed a distance d from an infinite conducting plate. Find the charge density at the surface of the plate as a function of the distance r from the point on the plate directly under the charged point.



(Hint: use the image charge method.)



Using an image charge
The single charge and plate
can be replaced by a positive
and a negative charge in
fel space. The potential at
day point can then be colculated $V_{p} = \frac{1}{418} \sqrt{(z-c)^{2}+r^{2}} + \frac{-9}{(z+c)^{2}+r^{2}}$

From Jauss's law the electric field can be related to the charge dusity

$$\oint \vec{E} \cdot d\vec{S} = \frac{8}{6}$$

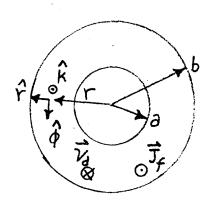
$$\vec{E} A = \frac{rA}{6}$$

$$\vec{E} = \vec{F}$$

E is related to the potential as $E = -\frac{\partial V}{\partial E}\Big|_{z=0}$

$$= \frac{-9d}{2\pi} \frac{1}{(d^2+r^2)^{3/2}}$$

4. The hollow wire shown in cross section at right has a given uniform current density \overline{J}_f and electron drift velocity \overline{v}_d . Find the magnitudes and directions of \overline{B} and of the Hall field \overline{E}_h , for a $\langle r \langle b \rangle$. The unit vectors at an arbitrary point in the wire are labeled \hat{r} , $\hat{\phi}$ and \hat{k} .



From Ampères (20), $2\Pi \Gamma B = M_0 J_{+}\Pi(\Gamma^2 - \lambda^2) \quad \vec{B} = M_0 J_{+}(\Gamma^2 - \lambda^2) \quad \vec{\Phi}$ $= -e(\vec{E}_{h} + \vec{V}_{x} \vec{R}) = 0 \quad (\text{no radial dright in equilibrium})$ $\vec{E}_{h} = -\vec{V}_{x} \vec{K} = + V_{x} \vec{K} \times M_0 J_{x} (r^2 - \lambda^2) \vec{\Phi} = -M_0 J_{x} J_{x} (r^2 - \lambda^2)^{\frac{1}{2}}$ $= -v_{x} \vec{K} = + v_{x} \vec{K} \times M_0 J_{x} (r^2 - \lambda^2) \vec{\Phi} = -M_0 J_{x} J_{x} (r^2 - \lambda^2)^{\frac{1}{2}}$ $= -v_{x} \vec{K} = -v_{x} \vec{K} \times M_0 J_{x} (r^2 - \lambda^2) \vec{\Phi} = -M_0 J_{x} J_{x} (r^2 - \lambda^2)^{\frac{1}{2}}$

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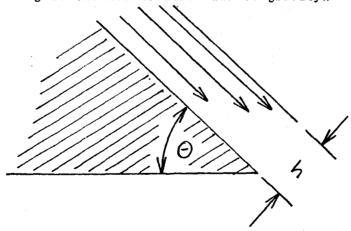
MONDAY, SEPT. 24, 1984, 2-5 PM

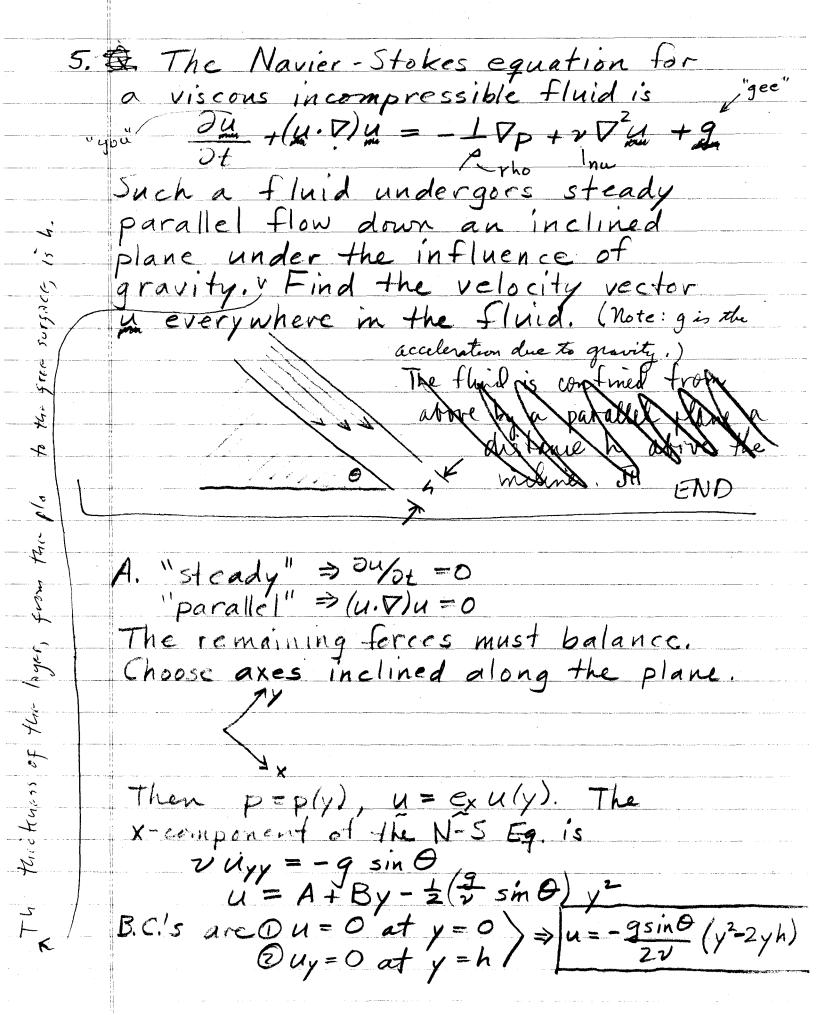
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- 5. The Navier-Stokes equation for a viscous incompressible fluid is

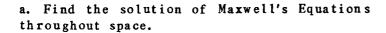
$$\frac{\partial \vec{u}}{\partial t} + (\vec{u} \cdot \vec{\nabla}) \vec{u} = -\frac{1}{\rho} \vec{\nabla} p + \nabla^2 \vec{u} + \vec{g}$$

Such a fluid undergoes steady parallel flow down an inclined plane under the influence of gravity. The thickness of the layer, from the plane to the free surface, is h. Find the velocity vector \vec{u} everywhere in the fluid. (Note: \vec{g} is the acceleration due to gravity.)

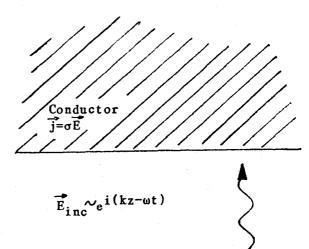




6. A plane electromagnetic wave is normally incident on a conductor of finite conductance σ .



- b. Calculate the fraction of power absorbed in the conductor as a function of frequency $\boldsymbol{\omega}_{\star}$
- c. Calculate the penetration range of radiation into the conductor as a function of



For part c, assume σ >>\\wedge \epsilon\$, where \varepsilon\$ is the dielectric permittivity.

Salitan:
$$\nabla x E = -\overline{B}$$
 $\nabla x B = \overline{E} + 4\pi r \overline{J}$
 $= \overline{E} + 4\pi r \overline{G} \overline{E}$

So in conduction $\nabla^2 E = \overline{E} + 4\pi r \overline{G} \overline{E}$

In five spine $\overline{E} = \widehat{E} \left(e^{i} \left(k x - w^4 \right) + R e^{i - k x - w^4 J} \right)$

An conduction $\overline{E} = \widehat{E} \cdot T = e^{i} \left(k^2 - w^4 \right)$
 $E \cdot \overline{B} = Continuous at boundary$
 $1 + R = T$
 $1 + R = T$

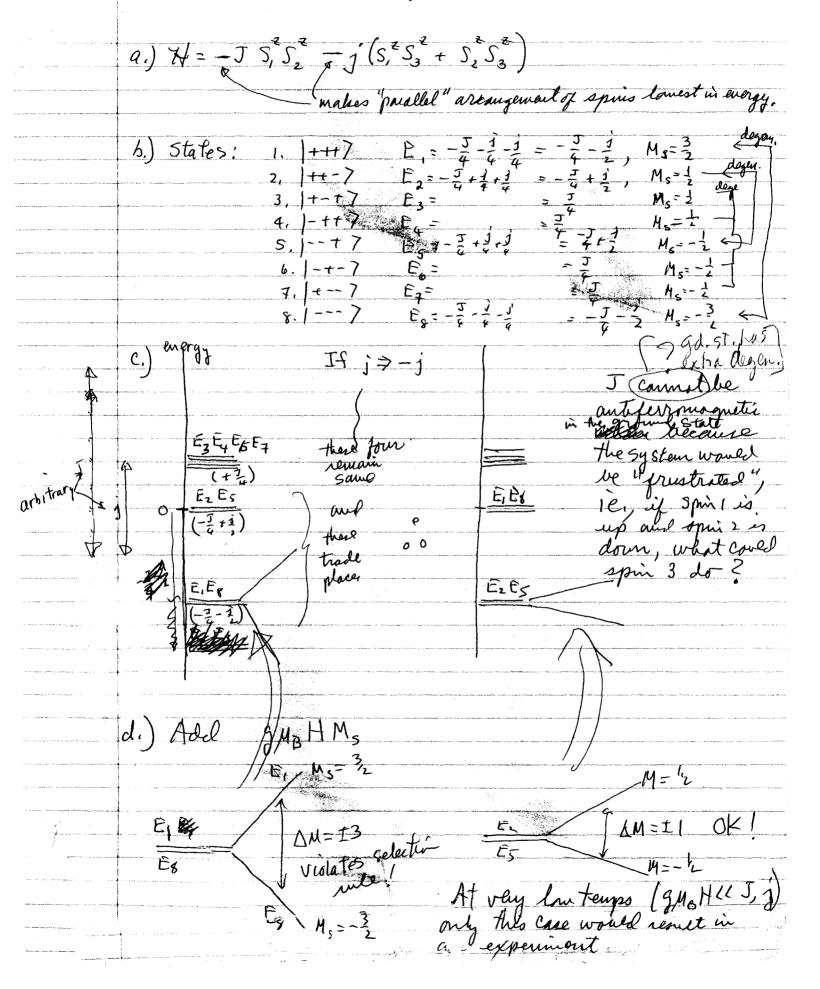
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7. A system of three spin-1/2 particles is pairwise exchange-coupled according to the Ising-model Hamiltonian.

$$H = -JS_1^z S_2^z - j(S_1^z S_3^z + S_2^z S_3^z); |J| >> |j|.$$

- a) What must be the signs of J and j if the ground state is ferromagnetic?
- b) Determine the eigenstates and their energies for arbitrary J and j.
- c) Sketch the energy-level diagram for J=2j>0. Also for J=-2j>0. What is peculiar about the case J<0?
- d) Sketch the Zeeman splitting of the ground state in the two cases where J>0. If the selection rules for an experiment are $\Delta M=\pm 1$, which case would have detectable transitions within the ground state at very low temperatures?

Note - the problem was revised. This solution was written for the original version.



8. An electron in the spinor state at t=0,

$$[\chi_0] = \begin{pmatrix} \cos\frac{\theta}{2} \\ \frac{\theta}{\sin\frac{\theta}{2}} \end{pmatrix} \qquad e^{i\vec{k}\cdot\vec{r}} \qquad [\theta, \vec{k} = constant]$$

is subjected to a uniform magnetic field B in the z-direction, i.e. the quantization axis.

- a) Interpret the state $[\chi_o]$.
- b) Determine $[\chi]$ for t>0 and interpret it.

8. QUI Hermanson "ky" ove" An electron in the Equipor) state at t=0, chi [] = (=) e'k'r [Q, k = constant] is subjected to a unitorn magnetie field B in the 7-direction, i.e. He quantitation axis. a) Interpret the state [X]. END 6) Defermine [X] for to and interpretit : a) Conjute Spin projections $\langle S_{x} \rangle = (\cos \frac{1}{2}, \sin \frac{1}{2}) \pm (0) (\sin \frac{1}{2})$ $= \pm (\cos \frac{1}{2} \sin \frac{1}{2} + \sin \frac{1}{2} \cos \frac{1}{2})$ $=\frac{t}{2}\sin\theta$ $(S_y) = \frac{t}{2} \left(\cos \frac{1}{2}, \sin \frac{1}{2} \right) \left(\frac{0-i}{i} \right) \left(\frac{\cos \frac{1}{2}}{\sin \frac{1}{2}} \right)$ $\langle S_{+} \rangle = \frac{1}{2} \left(\cos \frac{\pi}{2}, \sin \frac{\pi}{2} \right) \left(\frac{10}{0-1} \right) \left(\frac{\cos \frac{\pi}{2}}{\sin \frac{\pi}{2}} \right)$ $=\frac{1}{2}\cos\theta$ The spin is aligned along an axis is in the xz plane and moves along k. 2 1950.

6) When
$$B \neq 0$$
, $H = \frac{b^2}{2m} + H$

$$= -g(\frac{MB}{t})S_z B ; MB = \frac{gt}{2me} < 0 \text{ for } e^{-\frac{gt}{2me}}$$

$$= -(\frac{gt}{2}) \text{ fine } S_z$$

$$= \omega_0 S_z ; \omega_0 = -(\frac{gt}{2}) \text{ fine } > 0$$

Now $[\chi] = e^{-iHt/t} (\cos \frac{gt}{2}) e^{ik\cdot t}$

$$= e^{ik\cdot t} - \frac{Et/t}{k} e^{-i\omega_0} (\frac{S_t}{t}) t (\cos \frac{gt}{2}) e^{-ik\cdot t}$$

$$= e^{ik\cdot t} - \frac{Et/t}{k} e^{-i\omega_0} (\frac{S_t}{t}) t (\sin \frac{gt}{2}) e^{-ik\cdot t}$$

$$= e^{i} \phi (\cos \frac{gt}{2} e^{-i\omega_0} t/2) e^{-ik\cdot t}$$

$$= e^{i} \phi (\sin \frac{gt}{2} e^{-i\omega_0} t/2) e^{-ik\cdot t}$$

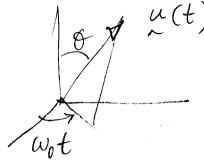
$$= e^{i} \phi (\cos \frac{gt}{2} e^{-i\omega_0} t/2) e^{-ik\cdot t}$$

$$= e^{i} \phi (\cos \frac{gt}{2} e^{-i\omega_0} t/2) e^{-i\omega_0}$$
And $(S_x) = \frac{t}{2} \sin \theta \cos \omega_0 t$

And
$$\langle 5_x \rangle = \frac{t}{z} \sin \theta \cos \omega_0 t$$

 $\langle 5_y \rangle = \frac{t}{z} \sin \theta \sin \omega_0 t$
 $\langle 5_z \rangle = \frac{t}{z} \cos \theta$

The spin precesses about B with frequency wo:



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- 9. In a crystal lattice of N sites, m Schottky defects are formed, each defect corresponding to one of the N original atoms being removed from the lattice and leaving a vacancy behind. An energy ε (ε >0) is required to form one such defect.
- a) Find the average number of Schottky defects when the crystal is at temperature T.
- b) Calculate the contribution to the specific heat associated with defect creation.
- c) Suppose a volume decrease δ is associated with the formation of each defect. Find the equilibrium volume at temperature T and pressure p, assuming the volume of the perfect crystal is V_{0} .

* * * * * * * * * * *

5. In a crystal lattice of W sites, m Schottky defects are formed, each defect corresponding to one of the Noriginal atoms being removed from the lattice and leaving a vacancy behind. An energy & (6>0) is required to form one such detect. epsilon & on spinwiter

- a) Find the average number of Schotthy defects when the crystal is at temperature (T.) uc
- b) Calculate the contribution to the specific heat associated with detect creation.
- c) Suppose a volume decrease of were associated with the formation of each defect. Find the equilibrium returne at temperature (T) and pressure Q, assuming the volume of the persect crystal is Va, le END

Som o Calculate partition fon, realizing that there are (M)

ways of picking the n atoms to be removed:
$$\frac{Z}{Z} = \sum_{m=0}^{N} {N \choose m} e^{-mE/kT} = (1+e^{-E/kT})^{N}$$

$$\langle m \rangle = \frac{\partial}{\partial (-\epsilon_{kT})} \ln \hat{z} = \frac{N}{1 + e^{-\epsilon_{kT}}} = \frac{N}{1 + e^{\epsilon_{kT}}}$$

or calculate F= U-T5= MEO - TkB ln(N) $\langle m \rangle$ found from $\frac{\partial F}{\partial m} = 0$, so $0 = \epsilon_0 - k_B T \ln \frac{N - k_B}{\langle m \rangle}$

Usma Stirlings

b) either write
$$F = -k_0 T \ln 2 = -Nk_0 T \ln (1 + e^{-\epsilon/k_T})$$

4 use $C = T \frac{\partial S}{\partial T} = -T \frac{\partial^2 F}{\partial T^2}$

(or use
$$C = \frac{\partial U}{\partial T} = \frac{\partial}{\partial T} \langle n \rangle \in$$

$$= \frac{\partial}{\partial T} \left(\frac{N \epsilon}{1 + e^{\epsilon kT}} \right) = \frac{N k_0 \langle \epsilon \rangle^2}{(1 + e^{\epsilon kT})^2} \frac{e^{\epsilon kT}}{(1 + e^{\epsilon kT})^2}$$

c) Now
$$-\frac{pV_0}{kT} = e^{-\frac{pV_0}{kT}} (1+e^{-\frac{pV_0}{kT}})^N$$

(where some the free energy is explicitly a for of p, we denote it of for Gibbs)

10. N atoms of a monatomic gas in a box of volume V have a Maxwell Boltzmann velocity distribution

$$n_o(v) = \frac{N}{V} \left(\frac{m}{2\pi k_B T}\right)^{3/2} 4\pi v^2 e^{-mv^2/2k_B T}$$

where $n_{O}(v)$ dv is the number density of atoms with speeds in the interval dv at v, T=absolute temperature, M=mass of each atom, and k_{B} =Boltzmann's constant. A small hole is made in the box, so that atoms can leak out.

- a) Find an expression for the velocity distribution n'(v) of escaping atoms i.e. the number (per unit time and unit surface area of the hole) escaping with speeds in the interval dv at v. Explain qualitatively why n' differs in functional form from n_0 .
- b) Find the rms velocity of escaping atoms, and compare it with the rms velocity of atoms inside the container. Based on your result, explain whether the remaining gas will become hotter or colder.

each atom, and & = Boltzmann's constant. A small hole is made in the box, so that atoms can leak out.

an expression for the relocity distribution n'(v) of escaping atoms - ie, the member (per unit time and surface area of the hole) with speeds in the interval du at v.
in functional form tuste: Explain qualitatively why m' differs from ma tailferent would still

b) Find the rms velocity of escaping atoms, and compare it with The rms velocity of atoms inside the container, Based on your result, with the remaining gas become hotter or colder.

END

Soln!

In time dt, all atoms with velocity 2 will escape through the hole provided they lie in the slanked cylinder shown at left, with volume AV dt. Thus we need volume Avidt. to multiply no by Avidt to get the number escaping in time at with vector velocity to. We then integrate over all angles of F, subject to the restriction Ut >0. In palar roords, this amounts to 05 q < ZTT, 050 < T/Z.

 $n'(v) dt A = \int_{a}^{b} d\rho \left(\frac{\partial n}{\partial u} \partial v \right) v \cos\theta \quad m_{o}(v) \quad A ot$ where the 1st factor of cos & comes from the transcription to

polar coordinates, and the 2nd from vz. Thus.

$$m'(v) = \underline{M_0(v)} v$$

This differs from no for the physical reason that faster-moving atoms strike the walls more often, and therefore are more likely to & cocape.

 $(v) = \int_{-\infty}^{\infty} \frac{1}{\sqrt{v}} dv = \int_{-\infty}^{\infty} \frac{1}{\sqrt{v}} \frac{1}{\sqrt{v}} dv = \int_{-\infty}^{\infty} \frac{1}{\sqrt{v}} \frac{1}{\sqrt{v}} \frac{1}{\sqrt{v}} dv$

$$= \left(\frac{2kT}{m}\right) \int_{0}^{\infty} \frac{5e^{-x^{2}}}{\sqrt{m}} dx = \left(\frac{2kT}{m}\right) \cdot 2 = \frac{4kT}{m}$$

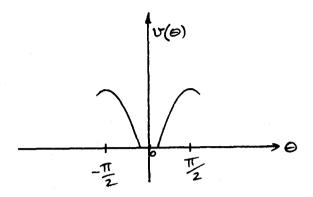
so for escaping atoms $v_{rms} = \sqrt{\frac{4kT'}{m}}$ $= \int_{0}^{\infty} \frac{1}{m_0(v)} \frac{1}{v^2} dv - \int_{0}^{\infty} \frac{1}{v^2} e^{-mv^2/2kT} dv = \int_{0}^{\infty} \frac{1}{v^2} e^{-mv^2/2kT} dv$ $= \int_{0}^{\infty} \frac{1}{m_0(v)} dv - \int_{0}^{\infty} \frac{1}{v^2} e^{-mv^2/2kT} dv = \int_{0}^{\infty} \frac{1}{v^2} e^{-mv^2/2kT} dv$

 $=\left(\frac{ZkT}{m}\right)^{3}$

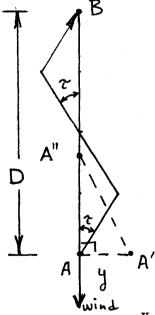
so for atoms in the container Trms = \frac{3kt}{m}

The slower atoms are left behind, and the container roots.

11. A sailboat sails at speed $v(\theta) = -\alpha + \beta \sin |\theta|$ when heading at an angle θ to the wind. $\alpha > 0$, $\beta > 0$



Find the optimum tacking angles τ in order to sail straight upwind at the fastest rate from A to B.



Starting from A' instead of A $(\frac{Y}{D} < \tan \tau)$, show that the fastest way to get to B is still to tack at angles τ . (Hint: calculate time to sail from A' to A' on line AB plus time to sail from A'' to B by optimum tacking, and minimize.)

dt = 0 $\frac{V'(0)}{V(0)} = \frac{sm0}{en0} = \frac{scn0}{-\alpha + ssm0}$ smol-x+(35md)= B(1-sm²0) 2 pgm20 - dsm0 - B=0 sm20 - 2/5 m0 - 1 20 SMT = 45 + (2+ (4))2

$$t = \frac{x}{v(0) \cos \theta} + \frac{D - x}{v(t) \cot \theta} + \frac{y}{x} = t \cos \theta$$

$$\frac{d+}{dx} = \left(\frac{1}{v(0) \cos \theta} - \frac{1}{v(t) \cot \theta}\right) - \frac{x}{v(0) \cos \theta} \frac{d}{d\theta} \left(v \cos \theta\right)$$

$$= 0$$

$$d = 0$$

$$d =$$

12. ³H nuclei collide with ⁴He nuclei to produce ⁶Li nuclei plus neutrons n. Find the kinetic energy threshold for this reaction in the lab frame where the helium nuclei are the targets. Mass defects of the nuclei are:

$$\Delta M(^3H) = 15.84 \text{ MeV/c}^2$$

$$\Delta M(^{4}He) = 3.61 \text{ Mev/c}^{2}$$

$$\Delta M(^{1}n) = 8.37 \text{ Mev/c}^{2}$$

$$\Delta M(^{6}Li) = 15.86 \text{ Mev/c}^{2}$$

(The mass defect of a nucleus is:

$$\Delta M = M(A,Z) - A\mu$$

where M(A,Z) is the actual mass of the nucleus, A is the number of baryons, and μ is the nuclear mass unit.)

Solution: Coms. of energy + numeritum in lab
$$\frac{P^2}{2M_1} + M_1c^2 + M_2c^2 = \frac{P^2}{2(M_3+M_4)} + (M_3+M_4)c^2$$

$$\frac{P^{1}}{2M_{1}}\left\{1-\frac{M_{1}}{M_{3}+M_{4}}\right\} = \left(M_{3}+M_{4}-M_{1}-M_{2}\right)c^{2}$$

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- 13. A rocket has engines which give it a constant acceleration of one g relative to its instantaneous inertial frame as measured by an accelerometer attached to the rocket. The rocket starts from rest near the earth. Ignore all gravitational effects.

Compute the proper time (τ) for the occupants of the rocket ship to travel the 30,000 light years from the earth to the center of the galaxy, assuming that they accelerate at one g for half the trip and decelerate at one g for the remaining half.

Suggestions:

Use the velocity and acceleration four-vectors. Note that $u^{\alpha} = dx^{\alpha}/d\tau$, $a^{\alpha} = du^{\alpha}/d\tau$, and that $g \approx 1$ year⁻¹ in units where c=1. Also note that the four-velocity and four acceleration are perpendicular.

* * * * * * * * * * *

Problem (1) solution

(a) Take the rockets motion to be along the x-axis. Let t be Earth time and T proper "rhip" time. The initial condition of rest year the Earth is then that

We have the following equations for the four-velocity, it and four-acceleration a:

$$\vec{u} \cdot \vec{u} = -1 = -(u^{\pm})^2 + (u^{\pm})^2 \qquad (normalization of four velocity)$$
 (1)

$$\vec{u} \cdot \vec{a} = 0 = -a^{\dagger}u^{\dagger} + a^{\chi}u^{\chi} \quad (\vec{a} \text{ orthogonal to } \vec{u})$$
 (2)

$$\vec{a} \cdot \vec{a} = g^2 = -(at)^2 + (at)^2$$
 (proper acceleration is g) (3)

$$(2) \Rightarrow at = a \times \left(\frac{u^{x}}{ut}\right) \quad \text{substituting this into (3), we get}$$

$$y^{2} = (a^{x})^{2} \left[1 - \left(\frac{u^{x}}{ut}\right)^{2}\right] = -\frac{(a^{x})^{2}}{(u^{t})^{2}} \left[-(u^{t})^{2} + (u^{x})^{2}\right] \quad \text{now one Eq. (1)}$$

$$y^{2} = (a^{x})^{2} / (u^{t})^{2} \quad \text{or} \quad \overline{a^{x} = q u^{t}}$$

$$(4)$$

Now differentiate Eq. (4) with repeat to proper time to get a differential equation for UN:

$$\frac{da^{x}}{d\tau} = \frac{\int_{0}^{2} u^{x}}{d\tau^{2}} = g \frac{du^{t}}{d\tau} = g a^{t} = g^{2} u^{x}$$

$$\int_{0}^{by} \frac{de^{t} n}{d\tau} = g \frac{du^{t}}{d\tau} = g a^{t} = g^{2} u^{x}$$

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$$\int_{0}^{by} \frac{de^{t} n}{d\tau} = g \frac{du^{t}}{d\tau} = g \frac{du^$$

The solutions to this equ,
$$\frac{d^2u^x}{d\tau^2} = g^2u^x$$
, are, obviously, just
$$u^x = A \exp \left[g \tau\right] + B \exp \left(-g \tau\right)$$
(7)

Since the initial condition is rest at T=0, and ax = dux = g at T=0,

we must have A = -B = 1, so that

$$u^{\times} = \sinh(9\tau) \tag{8}$$

and, by Eq. (1)

$$u^{t} = \cosh(g\tau) \tag{9}$$

To final X(T), we integrate Eq. (8), religent to the initial condition that X=0 at T=0:

$$\times (z) = g^{-1} \left[\cosh \left(g \tau \right) - 1 \right] . \tag{10}$$

In units with c=1 = 3x1010 cm

$$g = 980 \frac{cm}{s^{4}c^{2}} \cdot \frac{1 sec}{3 \times 10^{10} cm} = 3.27 \times 10^{-8} sec^{-1} = \frac{1}{3.06 \times 10^{7} sec} \approx \frac{1}{yr}$$

To get halfway to the galactic center requires x=15,000 light years, so

٥٥

For rich a large value, cost is well approximated by Zexp (9T), so

gz = In (30,000)

The deceleration half of the trip is identical, so the total

14. Find the quantum-mechanical eigenfunction $\Psi_n(k)$ and energy bands $E_n(k)$ of a one-dimensional empty lattice [V(x)=0] with lattice constant a; n and k are the band index and wave-vector. Illustrate your results with a sketch of the energy bands. Hint: Use Bloch's theorem to represent $\Psi_n(k)$ in terms of its periodic part $u_n(k)$.

14. Find the eigenfunction Post 1.c. kay

(A. Find the eigenfunction PMK) (mband index)

and draw the energy bands of a one dimensione

empty latter with latter constant a.

(Hint less the boundary condition for MA(X),

the guriodic part of the Block function). "you"

Find the quartum-mechanical eigenfurctions $Y_n(k)$ and energy tands $E_n(k)$ of a one-demensional empty letter [V(x)=0] with latter constant a; n and k are the band index and ware-rector. Heit: Use Block's theorem to represent $Y_n(k)$ in terms of its periodic part $u_n(k)$.

$$\nabla_{x}^{(k)} = 2^{2kx} \mathcal{U}_{k}^{(m)}(x)$$

$$\nabla_{x}^{(k)} \mathcal{V}_{m}^{(k)} = \frac{2m(v-F)}{h^{2}} \mathcal{V}_{m}(h)$$

$$= \frac{2kx}{h^{2}} (\nabla_{x}^{2} + z_{L} h \nabla_{x} - h^{2}) \mathcal{U}_{m}^{(m)}$$

$$(\nabla_{x}^{2} + z_{L} h \nabla_{x}) \mathcal{U}_{h}^{(m)} = (h^{2} - \frac{2mE}{h^{2}}) \mathcal{U}_{h}^{(m)}$$

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$$(\nabla_{x}^{2} + z_{L} h \nabla_{x}) \mathcal{U}_{h}^{(m)}$$

$$(\nabla_{x}^{2} + z_{L} h \nabla_{x}) \mathcal{U}_{h}^{(m)}$$

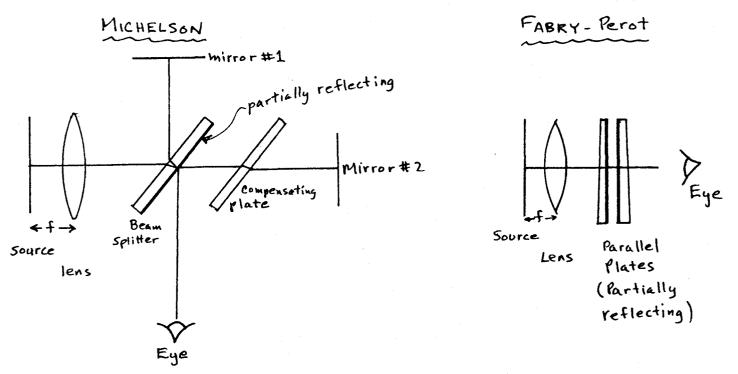
$$(\nabla_{x}^{2}$$

15. The mercury atom has the following energy levels expressed in terms of energy units $1/\lambda$.

- a) Explain the meaning of the spectroscopic notation above.
- b) What transitions will occur between these energy levels in a gas discharge? Explain in moderate detail.
- c) Briefly outline an experimental method for verifying the total angular momenta J assigned to the levels above.

n $l=0,1,2,3,\dots$ called $s,p,d,f\dots$
$^{3}P_{o}$: $^{2S+1}L_{J}$ gives notine of multiclection wavefu coupled to give $\vec{L}=\vec{l}+\vec{l}_{2}$ $\vec{S}=\vec{D}_{1}+\vec{D}_{2}$ and $\vec{J}=\vec{J}_{1}+\vec{J}_{2}$
b) $\Delta S=0$ $ \Delta L \leq 1$ $ \Delta J \leq 1$ $ \Delta m_{\overline{J}} \leq 1$ $ \Delta m_{\overline{J}} \leq 1$
derived (operator F does not act on S space of -> operator F is a vector (renk!) operator in L'space => max change ±1.000
Fluorescence will occur, since the closes in the discharge are excited by electron collisions
(onclude transition, are allowed between the triplets only (by electric dipole). Furtherwore must be 5 -> P to have parily change
Hence 35, 3P, only 6. 3 lines c) Zeemen effect: measure \ # sublevels 95 factors

- 16. Consider the circular pattern of fringes resulting from a Michelson interferometer which is illuminated by an extended monochromatic source and which is viewed by eye. A schematic drawing is given below.
- a) If the difference in distance between the beam splitter and the two mirrors is d=2 mm, find the order m of the central fringe for λ =500 nm and discuss whether it is bright or dark. (You may take these numbers to be exact.)
- b) Find the angular radius of the 3rd dark fringe seen off-axis.
- c) Describe the difference in the fringe pattern for a Fabry-Perot inteferometer (consisting of two parallel partially reflecting mirrors) vs. the Michelson interferometer and comment on their relative usefulness.



#16

a) path difference is $\Lambda = 2 d \cos \theta$ (See opties texts)

Depending on whether or not there is a phose shift upon reflection, a given Λ value can give either constructive or destructive interference. For example, an uncoated glass beamsplitter would introduce a net phose difference of π . If we ignore this effect, $\Lambda = m \lambda$ is the condition for constructive interference: $\Lambda = 2d = m \lambda$ $M = 2d = m \lambda$ $M = \frac{4 \times 10^{-3} \text{m}}{500 \times 10^{-9} \text{m}} = 8000$

b) ducreasing 0 => 1 = 2dcos0 Decreases

i all off-oxis parts of the pattern correspond to smaller module.

 13^{\pm} clarking $M = 7999 \frac{1}{2}$ 2^{nq} ... $M = 7998 \frac{1}{2}$ 3^{nq} ... $M = 7997 \frac{1}{2}$

 $cos\theta = \frac{7997.5}{8000}$ and $\theta = 1.43°$

c) Fabry - Perotoses multiple reflections, Multiple beam interference gives much sharper interference than the two-beam case of the Michelson, Hence, Fabry-Perot is more useful for spectroscopy. The Michelson, with Two I arms, played a key role in confirming special relativity theory. Loser versions have now proven that space is isotropic to better than 2.5 parts in 10'5