EMISSION

4 out Wif

E_f ¥

photon OUT@ energy

Er = kwif

ABSORPTION

F_f ω_{fi}

photon IN@ energy

Ez=kwfi

Interaction of a Quantum System my a Quantized EM Field

In our discussion of time-dependent perturbation theory, it was natural to speak of transitions 2 -> f in a quantum system as being either "emission" or absorption processes, wherein a "photon" of energy ~ to | was either given-up-to or gained-from some external field. The external field is of course Maxwell's EM field, and now we shall study why the above language is sensible.

Instead of separating the quantum system (Say an atom) from the external field, we shall consider the whole works, atom <u>plus</u> field, to be our quantum system... We do this in part because we want to be able to accurately track energy transfers between atom & field.

The EM field, now as part of the overall "quantum system", itself becomes quantized -- in just such a way that it can absorb the quantized energy towif when the atom emits it, or emit these when the atom absorbs it. These quantized energies to [Wifl are "photons", and the freed needs some quantum structure to be able to trade its photons with the (quentized) atom.

The most interesting feature of the theory is the notion of a quantized EM field. teld quantization is achieved by viewing the field as a (continuous) assembly of SHO's (via Fourier: $\mathbb{E}(\mathbf{r},t) = \int dt \int d^3x \, \mathbb{E}(\mathbf{k},\omega) e^{i(\mathbf{k}\cdot\mathbf{r}-\omega t)}$), then quantizing each mode (W, k) via the "well-known" rules for SHO quartization. We will use the annihilation (a) - creation (at) operator formalism for the field SHO's L Sakurai: 2.3].

tur discussion is in 3 steps, viz.

I. Interaction of a changed particle with an EM field.

II. Quantization of the radiation field & role of "photons".

The Calculation of transition amplitudes for the atom-field system.

Nonrelativistic (q, m) in EM potentials (o, A).

) From classical electrodynamics, the electric & magnetic fields [E & 1B are derivable from Scalar & vector potentials \$4 A as [Jackson, Eq. (6.31)]*

$$\rightarrow E = -\nabla \phi - \frac{1}{c} (\partial A/\partial t) , B = \nabla \times A .$$

Also the (nonrelativistic) Hamiltonian for a particle (q,m) in (\$\phi\$, \$A\$) is:

$$\rightarrow \mathcal{H}_{EM} = \frac{1}{2m} \left(p - \frac{q}{c} A \right)^2 + q \phi . \tag{2}$$

REMARKS on YEEn.

- 1: p is not the usual particle momentum. In fact: p=mv+(q/c)A, is the sum of the usual particle mechanical momentum and a correction due to the field.
- 2. But p is the momentum which is canonical for position or, in that Hamilton's letters of motion lead to the correct Torentz force Law, i.e.

$$\frac{\partial}{\partial \mathbf{r}} \mathcal{H}_{En} = (-) \frac{d}{dt} \mathbf{P} \Rightarrow \frac{d}{dt} (m \mathbf{v}) = q \left(\mathbf{E} + \frac{\mathbf{v}}{c} \times \mathbf{B} \right). \tag{3}$$

- 3. It is the canonical momentum p (and not mv) which will be replaced by the operator (-) it V when we do the QM on (9,m).
- 2) We assume the EM field in Hom can be split into two parts, which we shall discuss and treat separately. Namely...

$$\rightarrow (\phi, A) = (\phi, A)_{ext.}^{2} + (\phi, A)_{rod}^{2}.$$

(1) => external static fields (to be treated classically), e.g. pext = Conlomb potential,

Aux = \frac{1}{2} Bexe x & (Zeeman field). We choose Aext = 0, and set \frac{9}{4} \overline{\text{ent}} = \frac{V}{2}.

(2) => radiation fields (to be quantized). These obey the free-field waveletter, viz:

 $(\nabla^2 - \frac{1}{c^2} \frac{\partial^2}{\partial t^2}) A_{red} = 0$, $\phi_{red} = 0$. We choose the Coulomb gauge: $\nabla \cdot A_{red} = 0$, which describes the propagation of transverse EM waves [Jackson, Sec. 6.5].

I Jackson Eq. 112.14). Her of Eq. (2) is correct through terms of O(1/c2).

^{*} References are to J.D. Jackson" Classical Electrodynamics" (Wiley, 2nd Ed., 1975).

Reduction of Yben to building + coupling terms.

With these choices, Hem of Eq. (2) becomes ...

$$\rightarrow \text{Hoem} = \frac{1}{2m} \left(p - \frac{q}{c} A_n \right)^2 + V \quad \leftarrow A_n \text{ is } A_{rad} \text{ for the radiation field} \\
= \left(\frac{1}{2m} p^2 + V \right) - \frac{q}{2mc} \left(p \cdot A_n + A_n \cdot p \right) + \frac{q^2}{2mc^2} A_n^2. \tag{5}$$

①: With $p \rightarrow (-)$ it ∇ , this is the usual Schrödinger Hamiltonian; denote it by: $\frac{y_{6s} = \frac{1}{2m} p^2 + V}{Assume solutions} : y_{6s} \phi_n(r) = E_n \phi_n(r), \text{ known.}$ (6)

2: This is the compling between the bound particle (q,m) and the radiation field.
Ultimately An will be quentized. For now, note that with \$>1->1->1 to \$\mathbb{P}...

$$(\mathbf{p} \cdot \mathbf{A}_{n} + \mathbf{A}_{n} \cdot \mathbf{p}) \psi = -i t_{n} [\nabla \cdot (\mathbf{A}_{n} \psi) + \mathbf{A}_{n} \cdot (\nabla \psi)]$$

$$= -i t_{n} [\nabla \cdot (\mathbf{A}_{n} \psi) + 2 \mathbf{A}_{n} \cdot \mathbf{p} \psi]$$

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p. An + An p = 2 An p, w.r.t. wavefons \(\psi \).

3: We shall ignore this term as "small". We note: term 3/term 2 ~ 9An/cp. If

An is generated by motion of 9 in an atom of size as, then An ~ 9v/cas, and

the ratio: 3/2 ~ (9²/as)/mc² ~ (atom binding energy)/(rest energy) << 1. Term

3 is important only when (9,m) is subjected to intense external (laser) fields.

At this point, we have reduced Hem to just two terms, viz

This is not the entire Hamiltonian for the system: atom+field, but it does represent the particle binding, through Hbs, and the compling of the (spinless) particle's motion to the radiation field, through An. p.

Total System Hamiltonian: atom + field. Anticipation of the theory.

3) Two add-ons to the total system (atom + field) Hamiltonian... Bohr magnetion

A. Suppose q has spin $\frac{1}{2}$, and thus a magnetic moment $\mu = (\frac{9 \text{ th}}{2 \text{mc}})^{\prime} \sigma = (\frac{9}{\text{mc}}) \text{ S}$. μ couples to the radiation magnetic field Bn as...

B. The energy in the rediction field itself should be nicheded. It is ...

$$\rightarrow 46_n = \int d^3x \frac{1}{8\pi} (\mathbf{E}_n^2 + \mathbf{B}_n^2), \text{ classically [Jackson, Eq. (6.112)]}.$$
 (10)

Total, anly, this will be a sum 2(No+2) how over photomenergies two.

With these add-ons, we take as the total system (atom + field) Hamiltonian ...

This formulation completes Topic I as listed on p. QF1 above.

4) Before treating Topic II, we <u>anticipate</u> how the QM treatment of the overall system (atom+field) in Eqs. (11) can be developed. Quantization of the atom is "obvious"; we take it on faith that the field will be quantized in terms of SHO's at frequencies ω... where the energy in mode ω is (Nω+½) thwo, and Nω=0,1,2,... can be interpreted as the # "photons" in that mode. Then...

1. Let $\{\phi_n(\mathbf{r})\}\$ be the eigenfons of \mathcal{H}_S , i.e.: \mathcal{H}_S $\phi_n(\mathbf{r})=E_n\phi_n(\mathbf{r})$. $\leftarrow \underline{ATOM}$. Let $|(N_\omega)\rangle=|...N_\omega...\rangle$ n n \mathcal{H}_S , i.e.: \mathcal{H}_S $|(N_\omega)\rangle=E_{(N)}|(N_\omega)\rangle$. $\leftarrow \underline{FIELD}$.

Then eigenfens of Ho = Hos + Hon we the direct product states:

Interaction representation for atom+ field system.

Here: E(N) = 2 (Nw+2) tw, is the total energy (by photon counting) in the rad fed.

2. Use the direct product states of Eq. (12) as a basis when the coupling Home is turned on. The most general state of (atom+field) is the superposition:

 $\rightarrow \psi(t) = \sum_{n(N)} C_{n(N)}(t) |n(N)\rangle e^{-(i/\hbar)E_{n(N)}t}$, $N = \# \text{ photons at freq. } \omega$. (13)

(we've dropped the subscript "w" on N). With this 4, the Schrödinger wave extr.: it 34/2t = (46.+ 46int) 4, then gives -- exactly

iticm(m)(t) = 2 Cn(N)(t) (m (m) | 46 int | n(N) > e + (Em(M) - En(N))t,

(4)

En(N) = En + E(N) is total system energy, with atom at energy En, and radiation field with a distribution of (N) photons present [i.e. N1 photons at w2, N2 photons at w2, etc.]. An extr of true form of (14) was the starting point for our treatment of time-dependent perturbation theory (see class notes, p. tD3); there, as here, it was exact. The coefficients | Cn(N)|² are interpreted as the probabilities of actually finding the state | m(N) > at energy En(N).

3. A perturbation theory on Eq. (14) will yield "transition amplitudes", for changes in the overall system, when \(\(\mathbf{I}\)) → \(f(F)\), with the atom making a transition from an initial state \(\mathbf{I}\) to a final state \(f\), while the state of the radiation field changes \((\mathbf{I}\)) → (F). Overall, system energy must be conserved: if \(\mathbf{I}\) → forms (loses) energy, then \((\mathbf{I}\)) → (F) must lose (gain) energy.