

## Q → e Collisional Energy Transfer

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### Charged Particle Collisions: Energy Loss [Jk<sup>n</sup> Ch. 13].

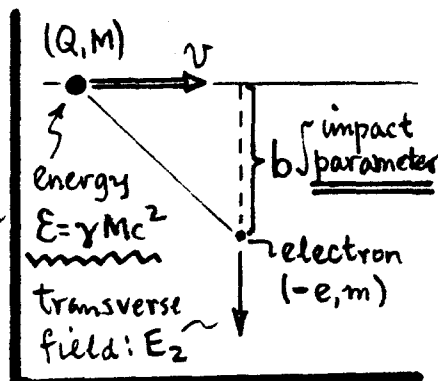
1) We consider the rate of loss of mechanical energy by a heavy, charged particle moving at high velocities through matter -- where it undergoes numerous "small" collisions with the ambient electrons and nuclei. The subject is important to:

- A. Cosmic-ray  $\phi$ : can identify high-E particles by tracks left in emulsions;
- B. High-energy  $\phi$ : specify collision products by bubble-chamber tracks & Cerenkov rad<sup>n</sup>;
- C. Solid-state  $\phi$  / Bio  $\phi$ : analyse radiation damage in materials & organisms.

2) The calculation begins with the simple picture at right...

The following assumptions are made:

- a) Major energy transfer from Q is to the electrons; the Q → e collisions cause negligible deviations in Q's path.
- b) Collisions of Q with nuclei ⇒ path deviations, but -- on average -- are rare, and cause ~ negligible energy loss.
- c) Q's energy loss is mechanical (transfers K.E. to electrons); radiation is ignored.
- d) Q's velocity  $v \gg$  orbital velocity of the e's; the e's end up nonrelativistic energy.



Now calculate the transverse momentum (impulse) delivered to e by passing Q...

→  $\Delta p = \int_{-\infty}^{\infty} e E_2(t) dt$ , <sup>2nd</sup>  $E_2(t) = \gamma Q b / (b^2 + \gamma^2 v^2 t^2)^{3/2}$  <sup>Jk<sup>n</sup> Eq. (11.152);</sup>  
t = time in e's frame

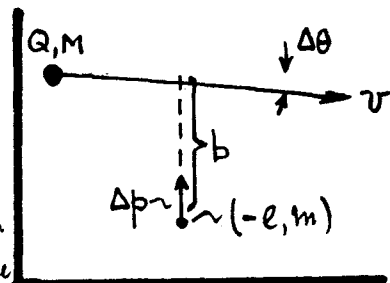
<sup>2nd</sup>  $\Delta p = \frac{e Q b}{v^3} \int_{-\infty}^{\infty} d\tau / [(b/v)^2 + \tau^2]^{3/2} = 2 Q e / v b$ . <sup>v (of Q) is assumed unchanged by collision. (1)</sup>

<sup>3rd</sup> Energy transferred to electron:  $\Delta E(b) = \frac{(\Delta p)^2}{2m} = \left( \frac{2 Q^2 e^2}{m v^2} \right) \frac{1}{b^2}$  <sup>Jk<sup>n</sup> Eq. (13.2)</sup> (2)

**REMARKS** on Q → e collision.

1. Deflection of M:  $\Delta \theta \approx \frac{\Delta p}{p} \approx \frac{2 Q e}{p v b}$ , <sup>2nd</sup>  $p = \gamma M v$ ;

<sup>2nd</sup>  $\left[ \Delta \theta \approx \frac{2}{\beta^2} \left( \frac{Q e / b}{E} \right) \sim 2 \left( \frac{\text{P.E. of Q at } b}{\text{total energy of M}} \right) \ll 1 \right]$  (3) <sup>deflection negligible</sup>



## Remarks on the $Q \rightarrow e$ collision. Lower & upper limits on $b$ .

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### REMARKS on $Q \rightarrow e$ collision (cont'd)

2.  $\frac{Q \rightarrow e \text{ energy}}{\text{transfer}} \} \underline{\underline{\Delta E(b) = \left( \frac{2Q^2 e^2}{mv^2} \right) \frac{1}{b^2}}}$   $\int$  depends on  $Q$  &  $v$ , but not on  $M$ ;  
impact parameter  $b \rightarrow \begin{pmatrix} \text{small} \\ \text{large} \end{pmatrix} \Rightarrow \Delta E(b) \rightarrow \begin{pmatrix} \text{large} \\ \text{small} \end{pmatrix}$ .

3.  $b \rightarrow 0 \Rightarrow \Delta E(b) \rightarrow \infty$ , nonsense. For head-on collisions ( $b \rightarrow 0$ ), must fix up  $b \dots$

LOWER LIMIT Maxm energy transfer for  $Q-e$  collision:  $\Delta E_{\max} \approx 2\gamma^2 mv^2$  (for  $M \gg m$ ),

$\Rightarrow \Delta E(b) \leq \Delta E_{\max} \Rightarrow \underline{\underline{b \geq b_0 = Qe/\gamma mv^2}}$ .  $\star$   $\nwarrow$  Jk<sup>n</sup> Eq. (13.5). (4)

For  $b < b_0$ , we must modify  $\Delta E(b)$  of Eq. (2).

4. A modification to  $\Delta E(b)$ , which respects  $\Delta E(b) \leq \Delta E_{\max}$  as  $b \rightarrow 0$  is:

$\Delta E(b) = \frac{2Q^2 e^2}{mv^2} / (b^2 + b_0^2)$ ,  $b_0 = Qe/\gamma mv^2$ .  $\int$  Jk<sup>n</sup> Eq. (13.7). (5)

5. What about an upper limit on  $b$ ? There is one, by following argument.

U. 2 LIMIT The assumption that  $v(\text{of } Q) \gg v(\text{of } e's)$  essentially treats the  $e's$  as being free.

But in fact they are bound. Now when  $b \rightarrow$  large, the  $Q-e$  collision time

$\Delta t \sim b/\gamma v$  can exceed the electron orbital period, so the  $e's$  will orbit their nuclei as though bound (certainly not free). For an "effective" collision, impose:

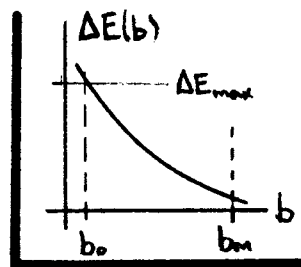
$\rightarrow \Delta t(\text{collision time}) \sim b/\gamma v < \frac{1}{\omega}$  ( $e$ -orbital period)  $\Rightarrow \underline{\underline{b < b_m \approx \frac{\gamma v}{\omega}}}$ . (6)

6. By comparing Eqs. (4) for  $b_0$  (lower limit) and (6) for  $b_m$  (upper limit),

we see easily that  $b_0 \ll b_m$  (this is true so long as  $\beta(\text{of } Q) >$

$\alpha \approx 1/137$ ). And, with these limits in mind, we expect that the loss

$\rightarrow \Delta E(b) \approx (2Q^2 e^2 / mv^2) \frac{1}{b^2}$ , holds over  $b_0 \leq b \leq b_m$ . (7)



$\star$  During collision, avg. electron velocity is  $v_e \sim \Delta p / 2m$ , and  $e$  moves  $d = v_e \Delta t$ ,  $\nwarrow$  collision time in  $e$ -frame  $\Delta t \sim b/\gamma v$ . Then:  $d \sim (\Delta p / 2m) \Delta t \sim Qe/\gamma mv^2 = b_0$ .

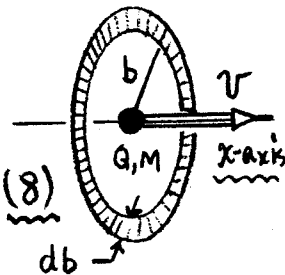
## Bohr Stopping Power Formula. Check upper limit $b_m$ .

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3) Now  $\Delta E(b)$  of Eq. (7) is good (?) for the collision of  $Q$  with a single electron. Look for an integrated result for  $Q$  colliding <sup>w/</sup> many electrons...

Solid:  $N$  atoms/vol.,  $Z$  electrons/atom

$$\Rightarrow \left. \begin{array}{l} \# \text{ electrons in vol. of length } \delta x, \\ \text{w/ impact parameters in } b \text{ to } b+db \end{array} \right\} \delta n = NZ \cdot 2\pi b db \delta x. \quad (8)$$



So  $Q$ 's energy loss to the  $\delta n$   $e$ 's:  $\delta E = [\Delta E(b)] \delta n$ , or...

$$\rightarrow \frac{\delta E}{\delta x} = 2\pi NZ [\Delta E(b)] b db = 2\pi NZ \left[ \left( \frac{2Q^2 e^2}{mv^2} \right) \frac{1}{b^2} \right] b db \quad \text{valid over } b_0 \leq b \leq b_m. \quad (9)$$

Integrate  $(\delta E/\delta x db)$  over  $b$  to get  $Q$ 's total energy loss/unit length...

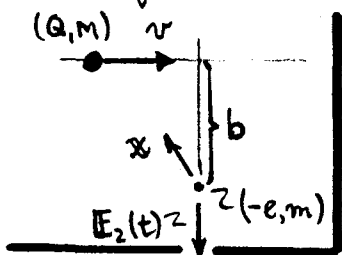
Jk<sup>2</sup> Eq. (13.13)  $\rightarrow$

$$\frac{dE}{dx} = \int_{b=b_0}^{b=b_m} (\delta E/\delta x db) db = 4\pi NZ \left( \frac{Q^2 e^2}{mv^2} \right) \ln B, \quad B = \frac{b_m}{b_0} \approx \frac{\gamma^2 m v^3}{\omega Q e} \quad (10)$$

This result for  $dE/dx$ , which measures the rate at which the material can stop  $Q$  by absorbing its K.E., is known as Bohr's Stopping Power Formula.

Much of the rest of Jackson's Chap. 13 concerns various ~small corrections to this formula--particularly adjustments to the impact parameter limits  $b_0$  &  $b_m$ .

4) The first adjustment to  $(\frac{dE}{dx})_{\text{BOHR}}$  concerns the accuracy of the upper limit  $b_m$   $\sim \gamma v/\omega$ . In Sec. 13.2, he does the "distant passage" problem (as suggested in Eq. (6) above) more carefully... with about the same result for  $b_m$ . To wit:



$(Q, M)$  passes by at  $v$  and generates (mainly) a transverse field  $E_2(t)$ ... but now electron is explicitly bound in a damped SHO...

$$\rightarrow \ddot{x} + \Gamma \dot{x} + \omega_0^2 x = -\frac{e}{m} E(t). \quad (11)$$

$E(t)$  can now include both  $E_2$  (transverse) &  $E_1$  (longitudinal). Details appear in Jk<sup>2</sup> Eqs. (13.15)-(13.33). The principal results are...

\* Worth noting:  $m$  = electron mass,  $\omega$  = electron orbit freq.,  $\gamma$  &  $v \leftrightarrow M$ 's motion.

Verify that upper limit  $b_m \approx \gamma v / \omega_0$  is reasonable.

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**RESULTS** "Distant passage": (Q, M) collision w/ bound (-e, m).

1. For an e bound in a (nonrelativistic) SHO as in Eq. (11), the energy transfer to the electron from any electric field impulse  $E(t)$  is...

Jk<sup>2</sup> Eq. (13.26)  $\Delta E = (e^2/2m) \left| \int_{-\infty}^{\infty} E(t) e^{i\omega_0 t} dt \right|^2$ ,  $\omega_0 = e$ 's SHO frequency. (12)

This is a general result, and holds even when  $E(t)$  is a (photon) radiation field.

2. In (12), we separately evaluate  $\Delta E$  for each component of  $E$  (passage), viz.†

$\left\{ \begin{array}{l} \text{only } E(t) = E_2(t), \text{ transverse, w/ } E_2(t) = \gamma Q b / (b^2 + \gamma^2 v^2 t^2)^{3/2}; \\ E(t) = E_1(t), \text{ longitudinal, w/ } E_1(t) = -\gamma Q v t / (b^2 + \gamma^2 v^2 t^2)^{3/2}. \end{array} \right\} \quad (13)$

Since  $E_1 \perp E_2$ , the integrals in (12) add as the sum of squares, with result:

Jk<sup>2</sup> Eq. (13.31)  $\Delta E(b) = \left[ \left( \frac{2Q^2 e^2}{m v^2} \right) \frac{1}{b^2} \right] \left\{ \xi^2 K_1^2(\xi) + \frac{1}{\gamma^2} \xi^2 K_0^2(\xi) \right\}$ ,  $\xi = \left( \frac{b}{\gamma v} \right) \omega_0$ ; (14)

previous result [Eq. (2)]      from transverse field  $E_2$       from longitudinal field  $E_1$       collision time  $\Delta t$  of Eq. (6)

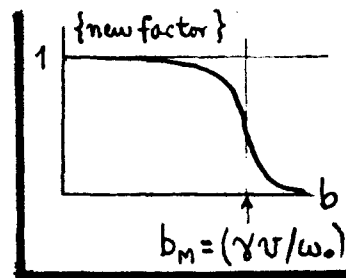
w/ modified Bessel fn<sup>‡</sup>:  $K_\nu(z) \approx \sqrt{\frac{\pi}{2z}} e^{-z} \left[ 1 + (4\nu^2 - 1) \frac{1}{2z} + \dots \right]$ , as  $z \rightarrow \infty$ ; (15)

‡  $K_\nu(z) \approx \frac{1}{2} \Gamma(\nu) \left( \frac{2}{z} \right)^\nu$ , as  $z \rightarrow 0$  ( $\nu \neq 0$ );  $K_0(z) \approx \ln \left( \frac{2}{z} \right) - 0.5772\dots$ ,  $z \rightarrow 0$ .

3. Eq. (14)  $\Rightarrow \Delta E(b) = [\text{previous result}] \times \{\text{new factor}\}$ , where the  $\{\}$  accurately for the electron binding. The correction factor behaves asymptotically as...

$\left\{ \text{new factor} \right\} \sim \begin{cases} 1 - O(\xi^2), & \text{as } \xi = \omega_0 \Delta t \rightarrow 0; \\ (1 + \gamma^2)^{-1/2} \xi e^{-2\xi}, & \text{as } \xi \rightarrow \infty. \end{cases} \quad (16) \Rightarrow$

The previous estimate of a cutoff in  $\Delta E(b)$  @  $b \approx b_m = \frac{\gamma v}{\omega_0}$  (i.e.  $\xi \approx 1$ ) in Eq. (6) was pretty good; there is little more at  $b > b_m$ .



‡ See, e.g., NBS Handbook (Abramowitz & Stegun), Chap. 9, Sec. 9.6.

† These fields were calculated in Jk<sup>2</sup> Sec. (11.10). See Jk<sup>2</sup> Eq. (11.152), p. 554.

## Modification of Bohr's Formula due to electron binding.

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5) We can use the exact result in Eq. (14) [for the energy transfer  $\Delta E(b)$  from  $(Q, M)$  to a bound electron  $(-e, m)$ ] over the impact parameter range...

$$\rightarrow \underline{b_0 = Qe/\gamma m v^2} \leq b \rightarrow \infty, \quad \text{or} \quad \underline{\xi_0 = \frac{Qe\omega_0}{\gamma^2 m v^3}} \leq \xi \rightarrow \infty. \quad (17)$$

The lower limit  $b_0$  is chosen to be consistent with the max. transfer  $\Delta E_{\max}$  in Eq. (4). We will adjust  $b_0$  below, but for the moment we have the transfer:

$$\rightarrow \Delta E(\xi) = \frac{2Q^2 e^2}{m v^2} \left( \frac{\omega_0}{\gamma v} \right)^2 \left\{ K_1^2(\xi) + \frac{1}{\gamma^2} K_0^2(\xi) \right\}, \quad \text{or} \quad \xi = (\omega_0/\gamma v) b. \quad (18)$$

Now we can consider  $Q$ 's energy loss/unit distance when colliding with a collection of electrons at different bound frequencies  $\omega_0 \rightarrow$  set of  $\omega_k$ 's. As follows...

[N atoms/vol., Z electrons per atom @ bound frequencies  $\{\omega_k\}$ ;  
the  $e$ 's have "oscillator strengths"  $f_k$   $\checkmark$   $f_k$  measures relative contribution from  $k^{\text{th}}$  electron, and  $\sum_{k=1}^Z f_k = Z$ , for norm.] (19)

In analogy to Eq. (10), form...

$$\begin{aligned} \rightarrow \frac{dE}{dx} &= N \sum_k f_k \int_{b_0}^{\infty} \Delta E(\xi) \cdot 2\pi b db \\ &= 4\pi N \left( \frac{Q^2 e^2}{m v^2} \right) \sum_k \int_{\xi_k}^{\infty} \left\{ K_1^2(\xi) + \frac{1}{\gamma^2} K_0^2(\xi) \right\} \xi d\xi, \quad \xi_k = \frac{Qe\omega_k}{\gamma^2 m v^3}. \end{aligned} \quad (20)$$

The integration can be done, with the result a modified Bohr formula...

Jk<sup>n</sup> Eq. (13.36)  $\rightarrow \boxed{\frac{dE}{dx} = 4\pi N Z \left( \frac{Q^2 e^2}{m v^2} \right) \left[ \ln B_c - \frac{1}{2} (v^2/c^2) \right]}$  (21)

$$\text{or} \quad \underline{B_c = 1.123 \gamma^2 m v^3 / \langle \omega \rangle Qe}, \quad \& \quad \ln \langle \omega \rangle = \frac{1}{Z} \sum_k f_k \ln \omega_k.$$

Compare with Eq. (10). This result is similar to the previous Bohr formula... but now we have an explicit  $\mathcal{O}(v/c)^2$  correction, and also a modification to the log argument:  $B = \gamma^2 m v^3 / \omega Qe \rightarrow B_c = 1.123 (\omega/\langle \omega \rangle) B$ . This is Bohr's work.

o) The weak point remaining in  $\left( \frac{dE}{dx} \right)$  of Eq. (21) is the question of close encounters:  $b \rightarrow b_0 = Qe/\gamma m v^2$ , classically. Here, there can be QM corrections, when the