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Solar Flares & CMEs

Problem Set 3

```
In[1]:= Clear["Global`*"]
```

Part a.

Define array to store numerical values of constants

```
In[2]:= vals = {};
```

The problem statement provides the following expression for the radiative loss function

```
In[3]:=  $\Lambda = \Lambda_0 (10^{-6} T)^{\alpha};$   
$Assumptions =  $T > 0 \ \&\& \ \Lambda_0 > 0 \ \&\& \ \alpha \in \text{Reals};$   
AppendTo[vals,  
   $\Lambda_0 \rightarrow \text{Quantity}[1.2 \times 10^{-22}, \text{"Ergs"} * \text{"Centimeters"}^3 * \text{"Seconds"}^{-1} * \text{"Kelvins"}^{1/2}];$ 
```

In Lecture 9, the radiative cooling time is defined as

```
In[6]:=  $\left( \tau_{\text{rad}} = \frac{3 k T}{n \Lambda} \right) // \text{Framed}$   
$Assumptions = $Assumptions &&  $k > 0 \ \&\& \ n > 0;$ 
```

```
Out[6]:= 
$$\frac{3 \times 1000000^{\alpha} k T^{1-\alpha}}{n \Lambda_0}$$

```

Part b.

Also in Lecture 9, we are given an expression for the conductive cooling time as

```
In[8]:=  $\tau_{\text{cond}} = \frac{21 k n L^2}{8 \kappa_0 T^{5/2}};$ 
$Assumptions = $Assumptions && L > 0 &&  $\kappa_0$  > 0;
AppendTo[vals,
   $\kappa_0 \rightarrow \text{Quantity}[10^{-6}, \text{"Ergs"} / (\text{"Seconds"} * \text{"Centimeters"} * \text{"Kelvins"}^{(7/2)})]$ ];
AppendTo[vals, k  $\rightarrow 1.38 \times 10^{-16}$  ergs/K];
```

Set the two cooling times equal and solve for the electron number density

```
In[12]:= (nb = Part[Solve[ $\tau_{\text{rad}} == \tau_{\text{cond}}$ , n] // FullSimplify, 2, 1, 2]) // Framed
```

Out[12]=
$$\frac{2^{\frac{3}{2}+3\alpha} \times 125^\alpha T^{\frac{7}{4}-\frac{\alpha}{2}} \sqrt{\frac{\kappa_0}{\Delta\theta}}}{\sqrt{7} L}$$

Solve for where the derivative with respect to temperature of the above expression is positive as a function of α .

```
In[13]:= ( $\alpha b = \text{Reduce}[D[nb, T] > 0, \alpha]$  // Simplify) // Framed // Quiet
$Assumptions = $Assumptions &&  $\alpha b$ ;
```

Out[13]= $2\alpha < 7$

Part c.

Solve Equation 2 in the problem statement for number density. This is the peak number density

```
In[15]:=  $nc = \frac{\epsilon}{3 L k T};$ 
$Assumptions = $Assumptions &&  $\epsilon > 0$ ;
```

Again, solve the condition where the two cooling times are equal for the peak temperature, with the number density replaced with the above expression

```
In[17]:= eqc =  $\tau_{\text{rad}}$  ==  $\tau_{\text{cond}}$  /. n → nc // FullSimplify;
(Tstar = Part[Solve[eqc, T] // FullSimplify, 1, 1, 2] // Quiet) // Framed
```

$$\text{Out[18]} = \left(\frac{9}{7} \right)^{-\frac{2}{-11+2\alpha}} \left(\frac{8^{1+2\alpha} \times 15625^\alpha k^2 K0}{\varepsilon^2 \Delta 0} \right)^{-\frac{2}{-11+2\alpha}}$$

So then the peak number density is found by plugging the peak temperature back into the expression for peak number density

```
In[19]:= (nstar = nc /. T → Tstar // FullSimplify) // Framed
```

$$\text{Out[19]} = \frac{3^{-1-\frac{4}{-11+2\alpha}} \times 7^{-\frac{2}{-11+2\alpha}} \varepsilon \left(\frac{8^{1+2\alpha} \times 15625^\alpha k^2 K0}{\varepsilon^2 \Delta 0} \right)^{-\frac{2}{-11+2\alpha}}}{k L}$$

Part d.

Equation 3 in the problem statement is a differential equation in pressure and time

```
In[20]:= p = 2 n k T /. n → nb /. T → T[t];
rhs = -5 n^2 Δ /. n → nb /. T → T[t];
oded =  $\frac{3}{2}$  D[p, t] == rhs
```

$$\text{Out[22]} = \frac{3 \times 2^{\frac{3}{2}+3\alpha} \times 125^\alpha k \left(\frac{11}{4} - \frac{\alpha}{2} \right) \sqrt{\frac{K0}{\Delta 0}} T[t]^{\frac{7}{4}-\frac{\alpha}{2}} T'[t]}{\sqrt{7} L} == -\frac{40 K0 T[t]^{7/2}}{7 L^2}$$

Solve the differential equation and write in the correct form

```
In[23]:= $Assumptions = $Assumptions && Ts > 0 && t ∈ Reals;
Td1 = Part[DSolve[{oded, T[0] == Ts}, T[t], t] // Quiet, 1, 1, 2];
Td2 = Td1;
Td2[[2, 1, 2, 2]] = Td1[[2, 1, 2, 2]] (Ts/Tstar)^(1/Td1[[2,2]]);
Td3 = Td2 // PowerExpand // Simplify // PowerExpand // Simplify;
Td3[[4, 1]] = Td3[[4, 1]] // Apart;
Td4 = Td3;
Td4[[4, 1]] = Td3[[4, 1]]/Td3[[4, 1, 1]] // Apart;
Td5 = Td4 Td3[[4, 1, 1]]^Td3[[4,2]];
posCond = Reduce[-Td5[[5, 2]] > 0];
$Assumptions = $Assumptions && posCond;
(Td6 = (Td5/Td5[[-1]] // Simplify) Td5[[-1]]) // Framed
```

$$\text{Out[34]} = \frac{T_s \left(1 - \left(2^{-2 + \frac{34}{11-2\alpha} - \frac{52\alpha}{11-2\alpha}} \times 3^{-1 + \frac{3}{-11+2\alpha} + \frac{-2\alpha}{-11+2\alpha}} \times 5^{\frac{11}{11-2\alpha} - \frac{44\alpha}{11-2\alpha}} \times 7^{\frac{4}{-11+2\alpha} - \frac{2\alpha}{-11+2\alpha}} k^{\frac{14}{-11+2\alpha}} t (3 + 2\alpha) \varepsilon^{\frac{3}{11-2\alpha} + \frac{2\alpha}{11-2\alpha}} \kappa \theta^{-\frac{4}{-11+2\alpha} + \frac{2\alpha}{-11+2\alpha}} \right)^{\frac{7}{11-2\alpha}} \right)}{\left(L(-11 + 2\alpha) \right)^{-\frac{4}{3+2\alpha}}}$$

Write in terms of ω and μ

```
In[35]:= wd = Td6[[2, 1, 2]] / t;
μd = -Td6[[2, 2]];
$Assumptions = $Assumptions && ω > 0 && μ > 0;
Td = Td6 /. wd → ω /. -μd → -μ
```

$$\text{Out[38]} = T_s (1 + t \omega)^{-\mu}$$

The solution will tend to zero if μ is positive

```
In[39]:= posCond // Framed
```

$$\text{Out[39]} = \boxed{\alpha > -\frac{3}{2}}$$

Part e.

The contribution function is approximated as

```
In[40]:= Gλ = Gλ0 Exp[-Log[T / Tλ]^2 / σλ^2];
$Assumptions = $Assumptions && Tλ > 0 && σλ > 0 && Gλ0 > 0;
```

The emissivity is related to the contribution function through

In[42]:= $\epsilon \lambda = n^2 G \lambda / . n \rightarrow nb$

Out[42]=
$$\frac{2^{3+6\alpha} \times 125^{2\alpha} e^{-\frac{\log\left(\frac{T_\lambda}{T_s}\right)^2}{\sigma \lambda^2}} G \lambda \theta T^{\frac{7}{2}-\alpha} K \theta}{7 L^2 \Delta \theta}$$

Take the derivative with respect to temperature to find the maximum emissivity

In[43]:= $(T\lambda pk = \text{Part}[\text{Solve}[D[\epsilon \lambda, T] == 0, T], 1, 1, 2] // \text{FullSimplify}) // \text{Framed}$

Out[43]=
$$e^{\frac{1}{4}(7-2\alpha)\sigma\lambda^2} T \lambda$$

Find whether β is greater or less than unity

In[44]:= $re = T\lambda pk[[1]] > 1$

Out[44]= $e^{\frac{1}{4}(7-2\alpha)\sigma\lambda^2} > 1$

In[45]:= $\text{Reduce}[re] // \text{FullSimplify} // \text{Framed}$

Out[45]= True

Therefore β is greater than unity.

Part f.

Find the time at which maximum emissivity is reached

In[46]:= $T\lambda pk == Td$

Out[46]= $e^{\frac{1}{4}(7-2\alpha)\sigma\lambda^2} T \lambda == T_s (1 + t \omega)^{-\mu}$

In[47]:= $(t\lambda = \text{Part}[\text{Solve}[T\lambda pk == Td, t] // \text{FullSimplify} // \text{Quiet}, 1, 1, 2]) // \text{Framed}$

Out[47]=
$$\frac{-1 + \left(\frac{e^{\frac{1}{4}(7-2\alpha)\sigma\lambda^2} T \lambda}{T_s} \right)^{-1/\mu}}{\omega}$$

The lifetime of the emission of the spectral line is defined through the relation

```
In[48]:= ελf = ελ /. T → Td;
fr =  $\frac{1}{ελf} D[D[ελf, t], t] == -\frac{2}{Δτλ^2} /. t → tλ // FullSimplify$ 
Out[49]=  $\left( \frac{e^{\frac{1}{4}(7-2\alpha)\sigma\lambda^2} T\lambda}{Ts} \right)^{2/\mu} \Delta\tau\lambda \mu^2 \omega^2 == \frac{\sigma\lambda^2}{\Delta\tau\lambda}$ 
```

Solve for $\Delta\tau\lambda$

```
In[50]:= (Δτλf = Part[Solve[fr, Δτλ], 2, 1, 2]) // Framed
Out[50]= 
$$\frac{\left( \frac{e^{\frac{1}{4}(7-2\alpha)\sigma\lambda^2} T\lambda}{Ts} \right)^{-1/\mu} \sigma\lambda}{\mu \omega}$$

```

Part g.

Plug in the values given in Table I in the problem statement

```
In[51]:= FeXXI = {Gλ0 → 1.64 × 10-25 ergs * cm3/s, Tλ → Quantity[11.51 × 106, "Kelvins"], σλ → 0.3};
FeXVIII = {Gλ0 → 1.43 × 10-25 ergs * cm3/s, Tλ → Quantity[6.91 × 106, "Kelvins"], σλ → 0.44};
FeXIV = {Gλ0 → 6.13 × 10-25 ergs * cm3/s, Tλ → Quantity[1.99 × 106, "Kelvins"], σλ → 0.25};
FeIX = {Gλ0 → 37.84 × 10-25 ergs * cm3/s, Tλ → Quantity[0.82 × 106, "Kelvins"], σλ → 0.42};
```

Save the other given numerical values to memory

```
In[55]:= AppendTo[vals, L → Quantity[5 × 109, "Centimeters"]];
AppendTo[vals, ε → Quantity[2 × 1012, "Ergs" * "Centimeters"-2]];
AppendTo[vals, α → -1/2];
```

Evaluate the time of peak emission

```
In[58]:= tλ /. Ts → Tstar /. μ → μd /. ω → ωd /. vals /. FeXXI // N // Framed
tλ /. Ts → Tstar /. μ → μd /. ω → ωd /. vals /. FeXVIII // N // Framed
tλ /. Ts → Tstar /. μ → μd /. ω → ωd /. vals /. FeXIV // N // Framed
tλ /. Ts → Tstar /. μ → μd /. ω → ωd /. vals /. FeIX // N // Framed
```

Out[58]= - 161.537 s

Out[59]= 134.086 s

Out[60]= 2498.32 s

Out[61]= 4240.92 s

Evaluate the duration of the emission

```
In[62]:= Δτλf /. Ts → Tstar /. μ → μd /. ω → ωd /. vals /. FeXXI // N // Framed
Δτλf /. Ts → Tstar /. μ → μd /. ω → ωd /. vals /. FeXVIII // N // Framed
Δτλf /. Ts → Tstar /. μ → μd /. ω → ωd /. vals /. FeXIV // N // Framed
Δτλf /. Ts → Tstar /. μ → μd /. ω → ωd /. vals /. FeIX // N // Framed
```

Out[62]= 271.039 s

Out[63]= 462.561 s

Out[64]= 558.348 s

Out[65]= 1303.97 s

```
In[66]:= ωd /. vals
```

Out[66]= 0.00050801 per second