

RadLib: a radiative heat transfer model library for CFD

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Abstract

Ca. 100 words

Keywords: radiative heat transfer, reacting flows, CFD

Required Metadata

Current code version

Ancillary data table required for subversion of the codebase. Kindly replace examples in right column with the correct information about your current code, and leave the left column as it is.

Nr.	Code metadata description	Please fill in this column
C1	Current code version	TODO 2.1
C2	Permanent link to code/repository used for this code version	<i>github.com/BYUignite/RadLib</i>
C3	Code Ocean compute capsule	N/A
C4	Legal Code License	MIT license (MIT)
C5	Code versioning system used	Git
C6	Software code languages, tools, and services used	C++, Python 3
C7	Compilation requirements, operating environments & dependencies	TODO
C8	If available Link to developer documentation/manual	TODO
C9	Support email for questions	davidlignell@byu.edu

Table 1: Code metadata (mandatory)

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1. Motivation and significance

Why did we make this? Radiation is hard to deal with in CFD. It's complicated. You can't really get analytic solutions to practical problems. In lots of cases, that's fine because radiation can be safely neglected; it's not usually the dominant mode of heat transfer. However, we do combustion CFD. Radiation isn't always important to combustion problems, but when it is, simulations can get grossly inaccurate. So we need a nice way to account for radiation in our CFD simulations.

When is radiation important to combustion simulations? In high-temperature systems, radiation can often be neglected in the beginning and middle of the simulation (assuming, say, a jet flame or a counterflow flame or something like that). At those times (which I'm going to call "early flame"), convective heat transfer dominates pretty heavily. However, radiation becomes more important for late-stage flame phenomena, either after some time has gone by (radiation time scales?), or physically high up in a flame where convection doesn't dominate as much as below, or just in areas that are relatively far away from the flame sheet itself where the reactions are happening. There's also soot, which involves radiation heavily, but soot particles are so big and slow that their time and length scales (and their radiative time scales) mean that radiation doesn't become important until late in the flame evolution.

So what's the problem? Why hasn't this been done already? Radiation is complicated. So is combustion. Combustion in particular can be so complex that direct simulations are too computationally expensive for us to simulate configurations that are practical for engineering systems. Direct simulations are usually used as a research tool. Other modeling approaches exist (LES, ODT, etc.) that lower the computational cost and allow us to simulate practical things. So far so good.

Radiation is a little like combustion in that its core mechanisms are complex, physically and mathematically. In addition, it's directional AND depends on the wavelength of the energy involved (other heat transfer doesn't have the wavelength dependence); its governing equations integrate over direction and wavelength, which makes things extra complicated. RADIATION GOVERNING EQUATIONS HERE. Convective heat transfer follows easy rules, but radiation doesn't. There are simple systems in which we can boil things down to analytic solutions, but most of the time, the simple equations don't apply, and that's often true in combustion systems where there are so many different length and time scales involved. The fundamental equations of radiative heat transfer are big and mathematically complex; they don't have analytic solutions except in the simplest geometries. So what do we do? One of two things. One, we do ray tracing (sometimes referred

41 to as Monte Carlo simulations. These can be extremely accurate, but super
42 computationally expensive. Essentially the equivalent of a direct numerical
43 solution. Two, we simplify the the equations. We make assumptions. And
44 so on. This is potentially less accurate, but more practical. By simplifying
45 the fundamental governing equations, we create models that apply to various
46 situations, systems, and geometries (i.e. black body assumptions, directional
47 assumptions, etc.). There are models developed for CFD and specific com-
48 bustion systems. These are the models that we've put into practice here in
49 such a way that they're easy to apply to various systems. Modular organi-
50 zation for this purpose.

51 So far there isn't an easy way to access and use radiation models. There
52 are simple ones in Cantera (check this, there might not be any), but they
53 make too many assumptions or don't work well for combustion or what have
54 you. RadLib is a radiation library of models that you can apply to any
55 simulation type, and we'd like to add it to Cantera, too. Basically, there
56 isn't a reliable way to do radiation calculations in CFD simulations without
57 coding the models yourself, which is extra hard because they're complicated
58 and hard. So we've done it for you and put them in a library that easy for
59 anyone to use. [1]

60 Models we're using:

- 61 • planck mean (optically thin?)
- 62 • two flux model (not sure if we even have this)
- 63 • weighted sum of grey gases (CITE Bordbar and others)
- 64 • SLW model (CITE Solovjov and others)

65 Discuss pros and cons of each model for combustion simulations? You still
66 have to pick the right one for your simulation and situation. That discussion
67 might fit better in software description section. See example section for
68 comparisons of models.

69 **2. Software description**

70 *2.1. Model descriptions*

71 You need two separate (but very related) models for any given case: one to
72 get radiation absorption coefficients and one for the solution to the governing
73 equations. The first can happen independently of anything else, but the
74 second requires the first. You can sort of mix and match models, but usually
75 we don't, so it's not presented that way in the software.

76 *2.1.1. Planck Mean*

77 *2.1.2. WSGG*

78 *2.1.3. RCSLW*

79 *2.2. Software Architecture*

80 Object-oriented C++ class library. Examples folder contains sample
81 driver scripts for using the library.

82 Python version somewhere? Corresponding examples?

83 **3. Illustrative Examples**

84 Simple comparisons of each of these models as implemented to some ray-
85 tracing results (equivalent to a direct numerical solution) to establish validity.
86 Illustrates when to use various models and why you might choose one over
87 another. Plots and discussion go here.

88 **4. Impact**

89 As far as I'm aware, there isn't an easy way to incorporate radiative heat
90 transfer into a CFD simulation. Most cases use optically thin assumption
91 (which works in some cases, but not in others) or neglect radiation entirely
92 (also applicable sometimes, but not always). Unfortunately, this means that
93 when you simulation some configuration in which radiation might be impor-
94 tant to the overall heat transfer, you won't get accurate simulation results.
95 Furthermore, if your study is about something else entirely, your radiation
96 model (or lack thereof) becomes a source of error that may be very difficult
97 to separate from other sources of error in your simulation study (i.e. soot
98 modeling studies). As of right now, if you want any detailed radiation treat-
99 ment, you have to code it yourself, which is difficult and requires external
100 validation.

101 RadLib can make researchers' lives easier by providing a library of preval-
102 idated radiation models that can be switched out with no difficulty. Addi-
103 tional models can be added easily using the provided modular framework.
104 Researchers can even use Radlib as a tool for comparing models. No need
105 to code multiple different complex radiation models yourself just to decide
106 which one works best for your simulation. And no need to puzzle through
107 the literature to figure out which model(s) might be best for your simula-
108 tion, either; instead, you can test them yourself. It saves tons of time and
109 effort that researchers can now put toward results rather than code or model
110 development.

111 By putting all these models side by side in a modular framework, RadLib
112 also provides a structure on which new or altered radiation models can be

113 tested against existing ones. Maybe more comparative studies can be done,
114 especially for more complex simulation cases. We plan to add radiation
115 models as appropriate to RadLib, too.

116 RadLib opens up new horizons for CFD simulations, especially in combus-
117 tion cases. It is designed such that it can be easily incorporated into existing
118 research codes. Our group plans to use it with ODT [CITE OTHER SOFT-
119 WAREX PAPER HERE?] alongside soot model library (in development) to
120 study late-flame phenomena such as soot oxidation, flame extinction, and
121 soot-flame breakthrough. These topics in particular are difficult because
122 they require accurate simulation data over a relatively long computational
123 time in addition to good models for both soot chemistry (which is an active
124 research area) and radiation heat transfer (which is the difficult part that
125 RadLib can solve). Using RadLib for such studies allows us to separate and
126 quantify sources of error that may occur due to various models, which is
127 super difficult if you only have one model to work with or it hasn't been
128 validated well.

129 RadLib can also be used outside of combustion CFD research for anything
130 involving radiative heat transfer. Potential research areas that could benefit
131 include atmospheric and climate sciences, interstellar phenomena, improving
132 efficiency of energy-producing processes, safety in chemical plant design, etc.
133 We aren't experts in these areas, but radiative heat transfer is a universal
134 phenomena that applies to any system or process that involves heat transfer.

135 This software has not been used outside of the current research group,
136 and it is not used in any commercial settings at this time.

137 5. Conclusions

138 Set out the conclusion of this original software publication.

139 6. Conflict of Interest

140 We wish to confirm that there are no known conflicts of interest associated
141 with this publication and there has been no significant financial support for
142 this work that could have influenced its outcome.

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References

- [1] D. O. Lignell, V. B. Lansinger, J. Medina, M. Klein, A. R. Kerstein, H. Schmidt, M. Fistler, M. Oevermann, One-dimensional turbulence modeling for cylindrical and spherical flows: model formulation and application, *Theoretical and Computational Fluid Dynamics* 32 (4) (2018) 495–520. doi:10.1007/s00162-018-0465-1.

Current executable software version

Ancillary data table required for sub version of the executable software: (x.1, x.2 etc.) kindly replace examples in right column with the correct information about your executables, and leave the left column as it is.

Nr.	(Executable) software meta-data description	Please fill in this column
S1	Current software version	TODO 2.1
S2	Permanent link to executables of this version	TODO For example: https://github.com/combogenomics/DuctApe/releases/tag/DuctApe-0.16.4
S3	Legal Software License	MIT license (MIT)
S4	Computing platforms/Operating Systems	Linux, OS X, Microsoft Windows
S5	Installation requirements & dependencies	TODO
S6	If available, link to user manual - if formally published include a reference to the publication in the reference list	TODO For example: http://mozart.github.io/documentation/
S7	Support email for questions	davidlignell@byu.edu

Table 2: Software metadata (optional)