Galactic Industries Corp. Universal Data Format Specification

Document History:

	_	
10-21-94	SCS	Original version (GRAMSUI.TXT).
05-18-95	JHD/BMG	Add NMR log text extensions.
07-13-95	JHD	Add AIA log text extensions.
09-18-96	JHD/PEH	Update NMR log text extensions, reformat to split DRIVER commands from converter log text entries, add sections for other techniques (NIR & UV/VIS), add new DRIVER commands for UV/VIS instruments
10-09-96	JHD/DEA/PEH	Incorporated changes from DEA & PEH to NIR section and DRIVER commands sections respectively.
11-08-96	JHD	Split DRIVER commands into separate document: "Galactic Industries Universal Driver Specification". Current document now only covers the Galactic SPC file format Log information.
09-03-97	JHD	Added definition of fexper main header parameter for experimental technique. Extend unit type tables and add new descriptions of SPC header parameters and 4D data sets. Also added more examples and detailed descriptions of header parameters and file structures. Added all header files as Appendices.

This document contains confidential information and is the exclusive property of Galactic Industries Corporation. Distributing or reproducing this document or any part thereof without the express written permission of Galactic Industries Corporation is expressly prohibited.

Spectra $Calc_{\circledR}$, Lab $Calc_{\circledR}$, $GRAMS/386_{\circledR}$, $GRAMS/3D_{\circledR}$ and $GRAMS/32_{\circledR}$ are registered trademarks of Galactic Industries Corp. Array Basic is a trademark of Galactic Industries Corporation. MS-DOS, Windows, Windows 95 and Windows NT are trademarks of Microsoft Corp. All other trademarks are the property of their respective owners.

Table of Contents

Purpose	1
Other Relevant Documents	3
Galactic SPC Binary File Format (SPC.H)	3
Galactic File Converter Specification (GCDLL.H)	3
GRAMS DDE Interface (GRAMSDDE.H)	3
Galactic Industries Universal Driver Specification	3
Converting Data Files To and From Galactic SPC Format	4
Conversion from a Foreign Format to Galactic SPC	4
Conversion from Galactic SPC to a Foreign Format	4
The Galactic SPC File Format	
The SPCHDR Structure (Main File Header)	7
The SUBHDR Structure (Subfile Header)	
The SSFTC Structure (Subfile Offset Pointers Directory)	
The LOGSTC Structure (Log Data Header)	
Data File Extensions	
Universal Log Text Parameters	
Log Text Information Format	
General Data File Description Parameters	
FT-IR Data Files	
NIR (Quant) Data Files	
UV/VIS (NIR) Data Files	
NMR Data Files	
Mass Spectroscopy Data Files	43
Chromatogram Data Files	
Diode Array Chromatography Detector Data Files	
Diode Array / CCD Spectrometer Data Files	
Xray Diffraction Spectrometer Data Files	
Fluorescence Spectrometer Data Files	
Data Files Created from AIA Data Interchange (ANDI/netCDF) Files	
Appendix A - SPC.H Header File	
Appendix B - GCDLL.H Header File	
Appendix C - GRAMSDDE.H Header File	61

Purpose

This document provides a guide for programmers to implement data reading or data writing routines for files in Galactic Industries' SPC data file format. This format provides for storage of a variety of different types of data taken from laboratory analytical instrumentation. As part of its products, Galactic Industries provides a series of data file conversion utilities that can translate data files taken from industry standard formats (i.e. ASCII, JCAMP, AIA/ANDI/netCDF), or from vendor specific binary formats into the Galactic SPC format. All of Galactic's software products use the SPC format to store and retrieve spectral and chromatographic data.

The Galactic SPC format consists of a main header describing the contents of the file followed by a binary storage area for the instrumentation data and an optional "Log block" for storing other relevant information about the data which cannot be stored in the main header block. In general, this block is used to store ASCII strings of important instrument or sample information that may be relevant to visualizing or processing the data when it is loaded into a Galactic software program. It is also used by Galactic software to store information about modifications made to the attached data. However, the Log block may also be used to store binary data as well. The text area of this block is known as the Log Text block, while the binary portion is called the Log Binary block. In addition to the main header and data areas (which are always present), data files in Galactic SPC format can contain none, either or both Log Text and Log Binary data in the Log block at the end of the file.

The Log Text block is a "free form" area where ASCII text information is stored. Although there is no defined format for the information in Log Text block (it can be used to store any type ASCII text information), Galactic has defined a system of text "keys" that are used to store important information when files are translated into SPC format from another source file. In addition, data acquisition software written by Galactic Industries that interfaces directly to laboratory instrumentation use this system of "keys" to store relevant sample and instrumental parameter information along with the raw data.

The Log Text block is primarily for use with Galactic-written data file converters, but information may be placed in this block by software from other companies who write/convert data files to the Galactic SPC file format or by Galactic-written data processing programs. Note that most of the described "key" parameters are optional, and Galactic file converters need not fill in all possible parameters in the Log Text block unless otherwise noted by the technique (i.e. FT-IR, NIR, UV/VIS, NMR, etc.) specification in this document. This set of key names is designed to take care of the following situations in an attempt to maintain data file "completeness" when it is brought into the Galactic software environment:

- 1. Data acquisition parameters used on the instrument to collect the data using Galactic software or similar parameters from a foreign file imported with a Galactic-written file converter.
- Data processing parameters required to properly manipulate, process and/or display the data.
- 3. Other relevant data that may exist in a foreign file for which there is no corresponding structure in the Galactic SPC file format which is required for describing the sample or data.
- 4. Other relevant data that may exist in a foreign file which may be required if the file is to be exported back to the exact same foreign format.
- 5. Processing histories and electronic signatures.

The main goal behind implementation of these standard "keys" is to allow data to be transferred between Galactic SPC files and foreign file formats with minimal loss of important information. While in most cases complete interchangability cannot achieved, the goal is to provide a mechanism to transfer as much relevant information as possible when data is brought into the Galactic SPC file format.

In addition to textual information, some file formats have large portions of extra information that must be maintained as a complete binary block within the SPC file, but this information is not required by Galactic software to display or process the spectrum or chromatogram. The Galactic SPC format provides a Log Binary block for storage of this information. Possible uses of this block include:

- 1. Storing a copy of a large block of binary information that is required for calculations, but should not be displayed when the file is loaded into Galactic software (i.e. imaginary data in an NMR spectrum).
- 2. Storing a copy of the file header for a complex file format so that the SPC file may be exported back to the foreign format completely.
- 3. Storing a complete copy of the foreign binary file such that data integrity and GLP (Good Laboratory Practices) can be maintained when the file is brought back to the native system or the data is archived for future use.

In order to simplify and standardize the process of developing conversion routines, this document was written to describe the set of standard ASCII Log Text "keys" and other storage mechanisms used to store information in a Galactic SPC file. Thus, the hope is that programmers both within Galactic Industries and at other organizations can build utilities to translate data to and from Galactic's SPC file format and maintain the maximum data completeness possible.

Other Relevant Documents

The following documents are also relevant for programmers implementing conversion routines to/from Galactic's SPC file format. In addition, these documents should be referred to by Galactic programmers implementing data acquisition routines.

The following documents, source code files, some sample program source code and a copy of this specification are all available from Galactic Industries Corp. in the "Galactic SPC File Software Developer's Kit". This kit is available from Galactic Industries Corp. as a Pkzip file called SPC SDK.ZIP.

Galactic SPC Binary File Format (SPC.H)

The source code header file SPC.H describes the complete SPC data file structure and all relevant header parameters. This file is absolutely required to properly understand the format and construct properly formatted SPC data files. See Appendix A.

Galactic File Converter Specification (GCDLL.H)

This document describes how to write a Windows DLL-based file converter that will plug into the Galactic Convert engine (GRAMSCNV.EXE). DLL's written to this specification may also be called directly by other Windows programs for conversion services. It describes the function calls and argument lists that must be supported in the DLL and provides for both 16-bit and 32-bit interfaces. There are sample code libraries available as well. This specification can only be used for writing converter DLLs that are compatible with Galactic Convert V2.00 or higher. Prior versions of the convert engine will not work with converters written to this specification.

This source code header file is provided to Galactic programmers for implementation of conversion routines within Galactic products, but can also be provided to other programmers who wish to use the Galactic Convert engine to translate files to/from the SPC format as well. Refer to this file for more information. See Appendix B.

GRAMS DDE Interface (GRAMSDDE.H)

This is a source code file that describes the complete Windows Dynamic Data Exchange (DDE) Client and Server interface in all GRAMS Level II compliant products. This includes information on how to develop a DDE server to interface with GRAMS as the client, and how to send data to GRAMS when it is the server. There are specific items in the SPC data file structure which must be filled in when sending data via DDE that are not required when simply converting data to disk-based SPC files for loading and manipulation in the GRAMS products. This file, other documentation and sample source code are included in a complete developer's kit for the GRAMS DDE interface available from Galactic Industries upon request. See Appendix C.

Galactic Industries Universal Driver Specification

The Log Text entries used for the data file Log Text parameters are, in almost all cases, the same arguments that are used by the DRIVER command interface to Galactic-written ABC data acquisition drivers. This document describes the interfaces that are allowed and those that <u>must</u> be supported when writing a Galactic ABC data acquisition driver. This document is not required to develop data file conversion routines for SPC files, and is primarily only of interest to Galactic programmers when developing data acquisition drivers.

Converting Data Files To and From Galactic SPC Format

When converting data files between different formats, it is desirable to maintain as much information as possible from the original data. This is certainly required by GLP, but is also a good idea to maintain the "completeness" of the data as usually every piece of information stored in a file is relevant to the data contained therein.

Internally, Galactic Industries has set two policies for all the conversion routines it creates:

- All attempts should be made to maintain data integrity when importing data from the native
 format into Galactic SPC. However, unless the native file format is simple enough to easily
 reconstruct converters will not support export from Galactic SPC files back to the native
 format. Galactic will continue to support and enhance the export capabilities of the
 converters for "industry standard" formats such as JCAMP and ASCII.
- As of December 1996 all Galactic-written converters are designated to follow the Galactic File Converter (GCDLL) Specification (see information in source code file GCDLL.H) to provide as much portability and flexibility as possible in terms of both the conversion software and the data files.

Conversion from a Foreign Format to Galactic SPC

When data is converted from foreign file formats into Galactic SPC format, every attempt should be made to include all relevant information that is available in the foreign file. This includes filling in the standard SPC main file header items (i.e. X units, Y units, Date, Time, Resolution Text, Memo Text) as well as all instrument parameters possible in the Log Text block.

Galactic has defined sets of standand Log Text keys for commonly used parameters for a variety of instrumental techniques. When at all possible, these keys must be used when converting data to the SPC format to insure the completeness of the data file. While the use of the Log Text parameters is not required to create a valid Galactic SPC file, it allows a greater degree of interchangeability of data to other foreign formats. In general, the Galactic-defined Log Text keys are specific only to each different instrumental technique. In addition, at this time, complete sets of keys are not defined for some techniques. As new conversion routines are developed either by Galactic or third parties that define parameters for these techniques, they will be adopted into the Galactic recommend set and published in future updates of this document. For complete information on the currently defined Log Text keys for instrumental parameters, see the "Universal Log Text Parameters" chapter later in this document

Conversion from Galactic SPC to a Foreign Format

When data is exported from a Galactic SPC file to a foreign file format, the converting program should look for any and all relevant items in the Log Text block to determine common parameters that can be exported along with the raw data. Note that some foreign formats do not support fields for all the data that may exist in the Log Text block, thus those parameters must be ignored on export. In addition, some foreign file formats will support many more parameter fields than are available in the parameters in the Log Text block. In this case, the export must fill in these fields with reasonable defaults that will allow the foreign software to load the exported data file without failure. Finally, when exporting data from a Galactic SPC file with no Log Text block, or no relevant entries in the Log Text block that match required parameters in the foreign file format, the converter should fill in these fields with reasonable defaults. These defaults should be set such that the foreign software can load the exported data file without failure.

In some cases, it will not be possible to properly export data from a Galactic SPC file to a foreign file format. There are many possible reasons for this:

- 1. Critical information required for constructing the foreign file is not available in the Galactic SPC file and there are no reasonable default values that can be universally used for all possible data files in the foreign format.
- The documentation on the foreign file format is not complete enough to determine reasonable default values for required parameters.
- 3. The data structures of the foreign format are too complex to reconstruct, or reconstruction would result in loss of accuracy or precision in the data. Examples of this are file formats that do not use the standard Intel word size and/or ordering, or use byte or word compression methods on the raw data.
- 4. The information in the foreign file format is not stored in a single data file, but uses complex structures to split the information into many different files and/or directories.

The Galactic SPC File Format

NOTE: The following information requires a basic understanding of the Galactic SPC binary format structures. Please refer to the SPC.H header file for reference to the variables and structures discussed here.

There are actually 2 different forms of the Galactic SPC file format: a simple array oriented record format used by the DOS-based products Spectra $Calc_{\odot}$ and Lab $Calc_{\odot}$, and a redesigned version used by the newer Windows-based products such as $GRAMS/386_{\odot}$, $GRAMS/32_{\odot}$ and other future products. The differences between these formats are documented in the SPC.H source code. The former format only allows for evenly spaced data arrays, while the latter allows for non-evenly spaced data, and dynamic data array sizes in multi record files to provide complete support for all possible different types of analytical data.

All conversion routines written by Galactic Industries after July 1996 only create SPC data files in the latter format. While future Galactic software products will continue to read all the different formats, they will only write the newer redesigned file format.

All the information in this document only refers to files stored in the "newer" GRAMS version of the SPC file format. The main reason being, that the older format does not allow for Log data storage. If back compatibility with the DOS products is required for the SPC files, then the old format must be used. However note that none of the information in this document regarding the Log Text and Log Binary information is relevant. It is highly recommended that programmers implementing data file translation routines to store data in Galactic's SPC file format only support the newer file format. However, programmers wishing to write routines to translate SPC data files into other formats must be prepared to read all the different SPC data formats.

As noted previously, the SPC format was designed to be flexible to allow for storage of many different data types. The final construction of an SPC file depends very heavily on the type of data it will hold. The SPC format also allows for different representations of the binary "Y" data. The ftflgs, fversn and fexp parameters in the SPCHDR structure (see SPC.H) determine both mechanism of the Y data storage (single or multi-record, evenly spaced X axis or data pairs, enumerated label types or custom strings, constant or dynamic record length, etc.), the format version type of the file (i.e. old or new) and the binary type of the Y values.

The remainder of this section is devoted to describing the important structures and parameters in Galactic SPC files and differentiating the data storage mechanisms offered by the format.

The SPCHDR Structure (Main File Header)

This section describes all the parameters in the SPCHDR structure and how they are used to describe the information stored in the SPC data file. One piece of terminology to note is that in SPC files, the X axis is always the horizontal axis (i.e. wavelength/frequency for spectra, time for chromatograms), the Y axis is always the vertical axis (response), and the Z axis is always the planar axis (event, time or secondary wavelength/frequency).

Ftflgs: File Data Type Flags

This determines structural organization of the SPC data file. As noted above, the structural layout of an SPC file will depend on the type of data it is being use to store. This value is set by OR'ing together the data file type bit flags:

Defined Flag	Ftflgs Value	Description			
TSPREC	0x01h	Y data blocks are 16 bit integer (only if fexp is NOT 0x80h)			
TCGRAM	0x02h	Enables fexper in older software (not used)			
TMULTI	0x04h	Multifile data format (more than one subfile)			
TRANDM	0x08h	If TMULTI and TRANDM then Z values in SUBHDR structures			
		are in random order (not used)			
TORDRD	0x10h	If TMULTI and TORDRD then Z values are in ascending or			
		descending ordered but not evenly spaced. Z values read			
		from individual SUBHDR structures.			
TALABS	0x20h	Axis label text stored in fcatxt separated by nulls. Ignore			
		fxtype, fytype, fztype corresponding to non-null text in			
		fcatxt.			
TXYXYS	0x40h	Each subfile has unique X array; can only be used if TXVALS			
		is also used. Used exclusively to flag as MS data for			
		drawing as "sticks" rather than connected lines.			
TXVALS	0x80h	Non-evenly spaced X data. File has X value array preceding			
		Y data block(s).			

In all cases of SPC files the ftflgs value in SPCHDR can be OR'ed with TSPREC to flag the Y data arrays being stored in 16 bit (rather than 32 bit) values. However, this only applies for block scaled integer format data values. If fexp is set to 0x80h to flag floating point data values, DO NOT set TSPREC! See the description of the SPCHDR parameter fexp for more details.

TALABS is used to flag axis unit label strings being stored in fcatxt rather than using the fxtype, fytype, fztype enumerated label types. See the description of the SPCHDR parameter fcatxt for more details.

The following diagrams show the layout of different types of SPC files. The parameters in the square brackets [] indicate the other possible ftflgs bit values that can be OR'ed with the main values to produce the same type of file, with other attributes. Note in all cases, the LOGSTC and Log Data blocks are optional with some exceptions. See the description of the SPCHDR parameter flogoff for more details.

SPCHDR			
SUBHDR			
Data			
LOGSTC (optional)			
Log Data (optional)			

Single File, Evenly Spaced X Values

Ftflgs = Null

This is the most common form of SPC files. It is typically used to store data for a single spectrum or chromatogram where the X axis data point spacing is evenly throughout the data array. Galactic software calculates the X value of a given point in the Data array from the data point index from the beginning of the array using fnpts, ffirst and flast in SPCHDR:

 $X \text{ value} = \text{index} \times (\text{flast} - \text{ffirst})/(\text{fnpts} - 1).$

SPCHDR			
SUBHDR			
Subfile Data			
SUBHDR			
Subfile Data			
LOGSTC (optional)			
Log Data (optional)			

Multifile, Evenly Spaced X Values

Ftflgs = TMULTI [| TORDRD]

This form is generally used for multidimensional data sets. Examples include hyphenated chromatography-spectroscopy techniques (except GC-MS) and full spectrum kinetics experiments. It is also used for response map data from imaging experiments, as well as multidimensional spectroscopies such as 2D-IR, and 2D-NMR. As with the previous form, the X values are assumed to be evenly spaced.

The fnsubs parameter in SPCHDR determines the number of repeats of the SUBHDR and Subfile Data structures in the file.

The Z axis values can be stored in a variety of ways, depending on how they are spaced. This is determined by the settings of the TORDRD bits in the ftflgs parameter in SPCHDR. Files with TORDRD not set have evenly spaced Z values which are determined by fzinc in SPCHDR, or by difference from the subtime and subnext values in the first SUBHDR if fzinc is null. Files with non-evenly spaced Z values will have TORDRD, and the Z value for each subfile is stored in the leading SUBHDR for every subfile. (Although allowed, TRANDM should not be used)

SPCHDR			
X Data			
SUBHDR			
Subfile Y Data			
LOGSTC (optional)			
Log Data (optional)			

Single File, Unevenly Spaced X values

Ftflgs = TXVALS

This form is used for single record data sets where the spacing of the X values is not uniform. This is typically used for data such as Raman spectra from CCD/diode cameras, and TOF MS spectra. Some chromatography systems change the data collection rate during the run, which would also require this format.

In this format, the entire array of X values is stored in the file in IEEE 32-bit floating point format. There exactly the same number of data points (specified by fnpts) in both the X and Y value data blocks. Although the X values are fully specified by the X array, the ffirst and flast parameters must also be set to the first and last values in the X array, respectively. The X values must be in either increasing or decreasing order; not random order.

SPCHDR			
X Data			
SUBHDR			
Subfile Y Data			
SUBHDR			
Subfile Y Data			
LOGSTC (optional)			
Log Data (optional)			

Multifile, Unevenly Spaced X Values, Common X Array

Ftflgs = TMULTI | TXVALS [| TORDRD]

This is effectively the mutifile form of the previous format. This form is generally used for multi-spectral kinetics experiments using Raman/CCD/Diode detectors. It can also be used for multi-chromatogram Diode array runs if the time point spacing is not constant (in this case the wavelength values for each chromatogram would be the Z values).

Note that in this form, the same set of X values is used for all the subfiles in the multifile. Thus every subfile must have the same number of data points as the X array which is specified by fnpts.

TORDRD can be applied to specify the spacing of the Z values as with multifiles with evenly spaced X values (above).

SPCHDR			
SUBHDR			
Subfile X Data			
Subfile Y Data			
SUBHDR			
X Data			
Subfile Y Data			
Directory (optional)			
LOGSTC (optional)			
Log Data (optional)			

Multifile, Unevenly Spaced X Values, Unique X Arrays

Ftflgs = TMULTI | TXYXYS | TXVALS [| TORDRD]

This format is generally only used for one type of data: GC-MS spectra. This format allows every subfile to have a unique X array, thus a different number of data points. In this case, the fnpts parameter in SPCHDR should normally be set to null since the number of data points in each subfile is specified by the subnpts parameter in each SUBHDR structure. However, if the optional Directory of subfile pointers (SSFSTC) is included at the end of the file, then fnpts must be non-null and hold an offset pointer to the beginning of the directory.

When <code>TXYXYS</code> is included in <code>ftflgs</code>, all Galactic software will draw the data points a "sticks" originating from Y=0 (the traditional representation for MS spectra) rather than connecting the data points.

Note that a variant of this format can also be used for a single thresholded or centroided MS spectra (i.e. non-multifile). In this case, do not include TMULTI with the other settings for ftflgs. The Directory is not necessary for a single spectrum.

Fversn: SPC File Format Version

This is the file format version; normally this should always be set to 0x4Bh for "new format" Galactic SPC files. Although SPC.H makes mention of a 0x4C setting for new format SPC files that allows for a different word order, most Galactic software products do not support it.

Files in the older Spectra/Lab Calc format use 0x4D for this parameter. However, while the offset and format of this value are the same between the old and new format SPC files, the rest of the headers themselves are substantially different. This document only describes the newer format information. For more documentation the old SPC file format refer to the OSPCHDR structure in SPC.H.

Fexper: Instrumental Experiment Technique

This is the "experiment type" flag. In most Galactic SPC files in circulation today, the fexper parameter has been left set at null. However, Galactic plans to use this value as an internal flag to identify the instrumental technique that created the original data. This will allow future products be "smart" and tailor the interface to perform operations that are typically done on that type of data, while possibly disallowing other operations that don't make any sense. All future file conversion routines should take advantage of these flags when creating SPC files.

There are a series of predefined flags in SPC.H for setting this parameter documented in the following table. Note that these flags are mutually exclusive and <u>cannot</u> be OR'ed together.

Defined Flag	Fexper Value	Instrumental Technique			
SPCGEN	0x00h	General (could be anything)			
SPCGC	0x01h	Gas Chromatogram			
SPCCGM	0x02h	General Chromatogram (same as SPCGEN and TCGRAM			
		in ftflgs).			
SPCHPLC	0x03h	HPLC Chromatogram			
SPCFTIR	0x04h	FT-IR, FT-NIR, FT-Raman Spectrum (Can also be used for			
		scanning IR.)			
SPCNIR	0x05h	NIR Spectrum (Usually multi-spectral data sets for			
		calibration.)			
SPCUV	0x06h	UV-VIS Spectrum (Can be used for single scanning UV-			
		VIS-NIR.)			
	0x07h	Not Defined – Do not use.			
SPCXRY	0x08h	X-ray Diffraction Spectrum			
SPCMS	0x09h	Mass Spectrum (Can be GC-MS, Continuum, Centroid or TOF.)			
SPCNMR	0x0Ah	NMR Spectrum			
SPCRMN	0x0Bh	Raman Spectrum (Usually Diode Array, CCD, etc. not for			
		FT-Raman.)			
SPCFLR	0x0Ch	Fluorescence Spectrum			
SPCATM	0x0Dh	Atomic Spectrum			
SPCDAD	0x0Eh	Chromatography Diode Array Spectra			

Fexp: Data Block Scaling Exponent or Floating Point Data Flag

This is the scaling exponent for the Y data values in the Data Block. In most SPC files created by Galactic software, the Y values in the Data Block(s) are represented as fixed-point signed fractions scaled by a single fexp exponent value. These are similar to integers except that the binary point is above the most significant bit rather than below the least significant bit. This format was designed to allow for storing data in higher precision than is offered by IEEE 32-bit floating point representation, but uses the same amount of disk space.

To convert the Galactic fixed format Y values to floating point representation:

FloatY = $(2^Exponent)^IntegerY/(2^32)$

or: $FloatY = (2^Exponent)^IntegerY/(2^16)$ (only if TSPREC in ftflgs)

For example, a Y value of 0x40000000h represents 0.25 and a value of 0xC0000000h represents -0.25. If the fexp exponent is 2 then they represent 1 and -1 respectively.

Note that if the file is a multifile (see ftflgs) then there are separate exponents for each subfile stored in subexp in the preceding SUBHDR. This same conversion applies to the Y data for that particular Data Block immediately following the SUBHDR.

SPC files can also be formatted to store the Y data values as IEEE 32-bit floats. In this case the fexp value (or subexp values in the SUBHDR structures for multifiles) must be set to 0x80h. Currently, all Galactic-developed software products can read any and all SPC file formats, but new converters save SPC files using the floating-point format. However older Galactic converters always saved SPC files in the fixed integer format.

When developing new data file conversion routines, the Y data should always be stored in 32-bit IEEE floating point representation. The only exception would be if the original format of the data required higher precision than is offered by this format (i.e. if the original data was 32-bit integer). As of December 1996 all new Galactic-written data file conversion routines will output SPC files using the IEEE 32-bit floating point format unless more precision is required due to the original format of the data.

Fnpts: Number of Data Points or TXYXYS Directory Offset

This is the number of data points in the X and Y data array(s) in the file. If the TXYXYS bit is turned on in the ftflgs parameter, then fnpts will normally be null since the number of data points is different for every subfile and is specified by the subnpts parameter in the SUBHDR structures.

However, the SPC file format allows for a separate Directory of subfile pointers to be stored immediately following the last subfile in the multifile. A non-null value for fnpts flags the existence of the Directory, and the fnpts parameter must be the offset to the beginning of the Directory. The Directory is a series of fnsubs repeats of the SSFSTC structure which gives offset pointers to the beginning of the SUBHDR for each subfile. This allows software to quickly "look up" an individual subfile without having to scan through all the subfiles. It also allows software to modify a single subfile, change its size, and place it at the end of the subfile list, as long as the corresponding pointer in the directory is also changed. This also allows for the subfiles to be stored in the multifile in a random Z order, although this is generally not recommended.

Ffirst: X Value of First Data Point

X value corresponding to the first data point in the Y data array. In the case of XY type files (ftflgs includes TXVALS) this should be the same as the first value in the X data array. Ignored for files where ftflgs includes TXYXYS.

Flast: X Value of Last Data Point

X value corresponding to the last data point in the Y data array. In the case of XY type files (ftflgs includes TXVALS) this should be the same as the last value in the X data array. Ignored for files where ftflgs includes TXYXYS.

Fnsub: Number of Subfiles

Number of repeats of SUBHDR and data blocks in the file. In cases where ftflgs does not include TMULTI, this will normally be 1 or null.

Fxtype, Fytype, and Fztype: Enumerated X,Y and Z Axis Unit Label Types

Galactic has defined a series of enumerated axis unit "types" that are allowed for SPC files. This allows for software to know the type of data stored in the file, and whether certain operations are allowed (i.e. convert to Absorbance from Transmission units). The SPC.H file defines the enumerated X, Y and Z axis unit type numbers. For convenience, they are duplicated here, however the ONLY source of the currently defined unit strings is maintained in the SPC.H source code. The fxtype and fztype unit parameters use the same enumerated list, while the fytype has its own list.

Note that not all the unit type values listed are supported by all Galactic products. The lists have been extended as the new products and routines for new data types were developed. The values listed here are in decimal, NOT hexidecimal.

It is not necessary to use one of the enumerated type values listed. If the unit type string required is not in the lists below, then SPC file main header block provides for "custom text" unit labels to be stored in the header instead. See the fcatxt parameter for more information.

Defined Flag	Fxtype or	X or Z Unit Label		
	Fztype Value			
XARB	0	Arbitrary		
XWAVEN	1	Wavenumber (cm-1)		
XUMETR	2	Micrometers		
XNMETR	3	Nanometers		
XSECS	4	Seconds		
XMINUTS	5	Minutes		
XHERTZ	6	Hertz		
XKHERTZ	7	Kilohertz		
XMHERTZ	8	Megahertz		
XMUNITS	9	Mass (M/z)		
XPPM	10	Parts per million		
XDAYS	11	Days		
XYEARS	12	Years		
XRAMANS	13	Raman Shift (cm-1)		
XEV	14	Electron Volts (eV)		
ZTEXTL	15	X,Y,Z text labels in fcatxt (old 4Dh version only)		
XDIODE	16	Diode Number		
XCHANL	17	Channel		
XDEGRS	18	Degrees		
XDEGRF	19	Temperature (F)		
XDEGRC	20	Temperature (C)		
XDEGRK	21	Temperature (K)		
XPOINT	22	Data Points		
XMSEC	23	Milliseconds (mSec)		
XUSEC	24	Microseconds (uSec)		
XNSEC	25	Nanoseconds (nSec)		
XGHERTZ	26	Gigahertz (GHz)		
XCM	27	Centimeters (cm)		
XMETERS	28	Meters (m)		
XMMETR	29	Millimeters (mm)		
XHOURS	30	Hours		
XDBLIGM	255	Double interferogram (no display labels)		

Defined Flag	Fytype Value	Y Unit Label		
YARB	0	Arbitrary Intensity		
YIGRAM	1	Interferogram		
YABSRB	2	Absorbance		
YKMONK	3	Kubelka-Munk		
YCOUNT	4	Counts		
YVOLTS	5	Volts		
YDEGRS	6	Degrees		
YAMPS	7	Milliamps		
YMETERS	8	Millimeters		
YMVOLTS	9	Millivolts		
YLOGDR	10	Log (1/R)		
YPERCNT	11	Percent		
YINTENS	12	Intensity		
YRELINT	13	Relative Intensity		
YENERGY	14	Energy		
	15	***** NOT USED *****		
YDECBL	16	Decibel		
	17 – 18	***** NOT USED ******		
YDEGRF	19	Temperature (F)		
YDEGRC	20	Temperature (C)		
YDEGRK	21	Temperature (K)		
YINDRF	22	Index of Refraction [N]		
YEXTCF	23	Extinction Coeff. [K]		
YREAL	24	Real		
YIMAG	25	Imaginary		
YCMPLX	26	Complex		
YTRANS	128	Transmission (ALL TYPES >=128 ARE ASSUMED TO HAVE INVERTED PEAKS!)		
YREFLEC	129	Reflectance		
YVALLEY	130	Arbitrary or Single Beam with Valley Peaks		
YEMISN	131	Emission		

Fpost: DDE Posting Disposition Code

Posting disposition code for SPC data sent to GRAMS/386 or GRAMS/32 via DDE. See GRAMSDDE.H for more information. Do not use when creating data file conversion routines; should normally be set to null.

Fdate: Data File Collection Date/Time

This is the date and time the data in the file was collected/stored. This is <u>not</u> the same as the date and time stamp put on the SPC file by the operating system when it is stored. The information is encoded as unsigned integers into a 32-bit value for compactness as follows (*most significant bit is on the left*):

Year (12 bits) Month (4 bits)	Day (5 bits)	Hour (5 bits)	Minute (6bits)
-------------------------------	--------------	---------------	----------------

Fres: Resolution Description Text

This is a short string (9 characters, but must be null terminated) describing the resolution of the data in the file. This is not necessarily the same as the data point spacing. For example, in and FT-IR spectrum, a commonly used resolution for data collection is 4 cm⁻¹, however the data point spacing for these spectra is ~1.93 cm⁻¹. In this case the text might say "4 cm-1". This string is not required, but if it is not used, it must be set to null.

Fsource: Source Instrument Description Text

This is a short string (9 characters, but must be null terminated) identifying the source instrument name or model. This string is not required, but if it is not used, it must be set to null.

Fpeakpt: Interferogram Peak Point Number

If the data in the file is an interferogram, (normally flagged by fytype=2), then this is the data point index in the Y array of the ZPD point, or point of maximum absolute response. Note that the data point index starts at zero (0) at the first value in the Y data array. If this is not known, or the data is not an interferogram set this value to null.

Fspare[8]: Array Basic Value Storage

These are 8 IEEE 32-bit floating point values used by Galactic data processing routines (primarily written into the SPC file by an Array Basic application) to store private values inside an SPC file. These were widely used in products prior to GRAMS/386 V3.0, however have been supplanted by the use of textual parameters in the Log Text block. In general, it is a good idea to leave these values set to null.

Fcmnt: Memo/Comment/Description Text

This is a 130 character text string memo or comment that describes the data in the file. This string must be null terminated.

Fcatxt: Custom Axis Unit Label Text

Note that when custom unit labels are used, ftflgs must include TALABS. When using "custom unit" labels, some data processing tasks in Galactic software require data having specific enumerated unit types in order to work correctly. In addition that conversions from these SPC files out to other foreign formats may not be reliable.

Flogoff: File Offset to LOGSTC Structure

This is an offset pointer to the beginning of the file Log Data. Normally this points to the beginning of the LOGSTC structure at the end of the file. If there is no Log Data in the SPC file, then flogoff must be set to null.

Fmods: File Modification Flags

These are a series of bit-encoded flags that identify certain classes of operations that have been applied to the data since it was originally collected. The bit codes correspond to letters that are representative of the operation performed. These were widely used in GRAMS products prior to GRAMS/386 V3.0, however have been supplanted by the use of textual parameters in the Log Text block. Refer to SPC.H for more detailed information.

Fprocs: DDE Processing Code

Post processing code for SPC data sent to GRAMS/386 or GRAMS/32 via DDE. See GRAMSDDE.H for more information. Do not use when creating data file conversion routines; should normally be set to null.

Flevel: Chromatogram Calibration Level

For Galactic internal use only; do not use. Must be set to null.

Fsampin: Chromatogram Sample Injection Number

For Galactic internal use only; do not use. Must be set to null.

Ffactor: Chromatogram Concentration Multiplier

For Galactic internal use only; do not use. Must be set to null.

Fmethod: Chromatogram Method/Program/Data Filenames List

For Galactic internal use only; do not use. Must be set to null.

Fzinc: Multifile Z Value Subfile Increment

For multifiles with evenly spaced Z axes (ftflgs includes TMULTI, but <u>not TORDRD</u>), this is the spacing value between the individual subfiles. The initial value is taken from the subfirst value in the first subfile's SUBHDR, and then the Z value for subsequent files is given by:

$$Z \text{ value}(n) = \text{subfirst}(1) + (\text{fzinc} * n)$$

where "n" is the subfile number. However, the Z value spacing can also be calculated from the first subfile in a different way. If fzinc is null, then the subfile spacing is calculated by:

$$fzinc = subnext(1) - subfirst(1)$$

Note that this only applies to multifiles that use even Z spacing. All other multifiles must have explicit values set in subfirst in every single SUBHDR. In general, it is a good idea to put values in for subfirst in every SUBHDR, even if the file will have even Z spacing. This allows future products to randomly seek through the file and retrieve the complete data for a single subfile without having to rely on information in SPCHDR other than fnpts, fnsubs and ftflgs.

Fwplanes: Multifile 4D Data Number W Planes

The fwplanes allows a using a multifile to store 4D data from experiments like spectral surface mapping. This is accomplished by identifying a series of subfiles to be interpreted as a volume of data with ordinate Y values along three dimensions (X,Z,W).

When fwplanes is non-null, then groups of subfiles are interpreted as separate planes along a W axis. The fwplanes value gives the total number of planes (groups of subfiles) in the SPC file and thus it must divide evenly into the total number of subfiles (fnsub).

Fwinc: Multifile 4D W Plane Increment Value

For 4D files (fwplanes is non-null) fwinc is the constant W spacing value between the planes of data. If fwinc is non-zero, then the W axis values are evenly spaced beginning with the subwlevel values in the SUBHDR structure for the first subfile and incremented by fwinc after each group of fwplanes subfiles. Thus the W value for subsequent planes is given by:

where "n" is the W plane number.

If fwinc is null, then the planes are allowed to have non-evenly-spaced W axis values as given by the subwlevel in the SUBHDR structure for the first subfile in each plane's group. However, the W axis values must be ordered so that the plane values always increase or decrease. Also all subfiles in the plane should have the same subwlevel value.

In general, equally-spaced W planes are recommended since some software may not handle fwinc=0.

Fwtype: W Axis Unit Label Type

The fwtype gives the W axis unit type value. This uses the same enumerated list of enumerated axis unit types as fxtype and fztype. Note that unlike the other 3 axis unit labeling methods, there is no allocated storage for a W axis custom unit string in fcatxt. Thus the W axis can only be labeled using a standard Galactic unit type.

Freserv: Reserved Storage for Future Expansion

Do not write any data in this area. In addition, make sure that this area of the SPCHDR structure is initialized to nulls to avoid problems with any future definitions of parameters in this area.

The SUBHDR Structure (Subfile Header)

In the new Galactic SPC file format, every individual data block is preceded by an SUBHDR structure. This gives information on the data in the following data block.

Subflgs: Subfile Data Flags

This value should always be set to null for files that are created by conversion routines. However if a subfile in an SPC file has been manipulated by Galactic software and saved back to disk, it is possible that some bit flags will be set in this value. See SPC.H for more information.

Subexp: Subfile Data Block Scaling Exponent or Floating Point Data Flag

Similar to fexp in SPCHDR, this is the scaling exponent for the data in the following Y data block for a single subfile in a multifile. See the previous description of fexp for information on

the scaling calculations. If the value of subexp is 0x80h, then the Y data block for the subfile is stored as IEEE32-bit float.

Note that the data storage format of the Y data blocks should be homogeneous throughout all subfiles in the multifile. This means that the subfiles should either all be stored as floating point format (subexp=0x80h for all subfiles) or they should all be stored as block scaled integers. If the latter format is used, each subfile can have a different exponent, but the file cannot also have individual subfiles where the Y data stored in floating point format mixed in.

Subindx: Subfile Index

Integer index of the subfile number in the whole multifile. Note that these should increase monotonically starting at zero (0) for the first subfile, going up to fnsubs - 1 for the last subfile.

Subtime: Subfile Z Axis Value

This is the Z value for the subfile. This is typically the time the trace data in the following data block was collected during the experiment (i.e. spectra in hyphenated chromatography experiments), but can actually be any value representing a 3rd dimension for the data collected.

This value can be null in the subfile SUBHDR structures if ftflgs includes only TMULTI (i.e. does <u>not</u> include TORDRD) and the fzinc value is non-null. However, remember the Z value of the first subfile must be set to the starting Z value for the series, otherwise zero is assumed.

If ftflgs does include TORDRD, then every subfile SUBHDR must have a value for subtime.

Subnext: Subfile Ending Z Axis Value

This is the "end" Z value for the subfile. In some experiments there is a time span over which the data for the subfile was collected, but this may not necessarily be the same as the spacing between the SUBHDR subtime values. For example, FT-IR kinetics runs can have multiple subfiles each of which was calculated from a number of co-added scans. If the time to collect the co-added scans for a single subfile was only 10 seconds, but the subfile spacing was set to be 30 seconds during the run, then the subnext parameter should be used to indicate the time actually required to collect the subfile. Thus the subtime value is the time point in the experiment that the collection of the subfile data began, while the subnext is the time that the subfile data was completed, and the next subfile would have a subtime different from both of these values.

Unless this capability is required, the value of the subnext parameter should be set to null, or set to the same value as the subtime. However, if the multifile is set up to have an evenly spaced Z axis and the fzinc parameter is null, then the subtime and subnext values in the very first subfile must be set to different values to indicate the Z value spacing.

Subnois: Subfile Peak Picking Noise Level

This is used by Galactic software to store the calculated noise level in the Y data block for Galactic's peak picking algorithm. This value should always be set to null for files that are created by conversion routines. However if a subfile in an SPC file has been manipulated by Galactic software and saved back to disk, it is possible that a value will be stored for subnois.

Subnpts: Subfile Number of Data Points

Indicates the number of data points in the following subfile X and Y data arrays for files with ftflgs set to include TXYXYS. For all other file types, it must be set to null.

Subscan: Subfile Number of Scans

The number of co-added scans collected to produce the subfile data. Generally only set by Galactic-written data acquisition routines. In general, this value should be set to null for files that are created by conversion routines.

Subwlevel: Floating W axis value (if fwplanes non-zero)

This is the subfile W axis values for 4D data sets. This value will only be used when the fwplanes value in the SPCHDR structure is non-null. Refer to fwplanes and fwinc in the SPCHDR documentation for more information.

Subresv[4]: Subfile Reserved Storage for Future Expansion

Do not write any data in this area. In addition, make sure that this area of the SUBHDR structure is initialized to nulls to avoid problems with any future definitions of parameters in this area.

The SSFTC Structure (Subfile Offset Pointers Directory)

This structure will generally only be found in files with the TXVALS and TXYXYS flags set in the ftflags value in the SPCHDR structure. This structure is effectively a lookup pointer to an individual subfile record in the multifile. In this type of multifile the subfile records are variable length, and therefore software cannot perform calculated jumps to specific subfiles. A directory of fnsubs repeats of these subfile pointer structures allows software to quickly jump to any subfile

The existence of the directory is flagged by a non-null value for the fnpts parameter in the SPCHDR structure. In this case, the fnpts parameter is interpreted as an offset pointer in the file to the beginning of the directory structure.

Ssfposn: Byte Offset to beginning of subfile SUBHDR(n)

Byte offset from the beginning of the file to the beginning of the SUBHDR structure for the subfile.

Ssfsize: Byte size of complete subfile record

Byte size in the complete subfile record (SUBHDR + X data block + Y data block).

Ssftime: Z value of subfile record

This is effectively a duplicate of the subfile SUBHDR subtime value.

The LOGSTC Structure (Log Data Header)

The Galactic SPC file Log Data is a free-form dynamic data block at the end of the file. The existence of log data is flagged by a non-null value of flogoff in the SPCHDR structure which is the byte offset to the beginning of the LOGSTC structure.

The LOGSTC structure normally immediately follows the main binary data block (or subfile directory in the case of TXYXYS files), and immediately precedes the Log Text block.

Logsizd: Byte size of log disk block

The logsizd describes the size (in bytes) of the complete log block (LOGSTC structure and Log data block).

Logsizm: Byte size of log memory block required

The logsizm should be the logsizd size of the block rounded to the next larger multiple of 4096.

Logtxto: Byte offset to Log Text data

The logtxto is the byte offset FROM THE BEGINNING OF LOGSTC to the beginning of the Log Text data block.

Logbins: Byte size of log binary block

This is the size in bytes of the private binary data block which is assumed to immediately follow the LOGSTC structure, and immediately precede the Log Text data block. This binary block area can be used by other programs to store their own private data inside an SPC file. In general, Galactic software programs do not use the data in this area, however it is carried along intact every time the file is modified and then saved back to disk. The one notable exception is for NMR data files, which use this area to store the imaginary array from quadrature detection NMR experiments.

Note that this binary data block is read into memory when the file is loaded and thus can cause memory usage to go up if this private data block is very large.

Logsizd: Byte size of log disk block

This is the size in bytes of a private binary block which is assumed to immediately follow the logbins data block. The purpose of this block is the same as for logbins, however this data is not loaded into memory along with the file when using Galactic software products. Thus when the file is modified and saved back to disk, this data is not carried along.

Logspar: Reserved storage for future expansion

Do not write any data in this area. In addition, make sure that this area of the LOGSTC structure is initialized to nulls to avoid problems with any future definitions of parameters in this area.

Data File Extensions

Galactic'c DOS based software packages supported only a few file extensions for data files. For Spectra Calc, only files with ".SPC" extensions were supported. For Lab Calc, data files with either ".SPC" or ".CGM" extensions were recognized. Again, as mentioned above, these files used a different file format with no Log block for storing extra data.

However, Galactic's GRAMS/386 and GRAMS/32 software packages allow for a variety of file extensions on the data files according to the following table. These were designed to allow users to more easily organize and segregate their data by analytic techniques when many files were stored in the same directory. Creating files in GRAMS/386 or GRAMS/32 using these extensions (either via the use of Array Basic programs, or the GRAMS Convert Application file converters), also sets the fexper parameter in SPCHDR to match the experiment extension (although newer designed file converters force the fexper value to an internally set value rather than letting the user choose the type).

In general these experiment-oriented file extensions are not used by many users. Thus the standard convention for file naming is to use ".SPC" file extensions for files that contain spectral data, and ".CGM" extensions for chromatographic data.

Some experiments create both sets of spectral data and chromatographic data (i.e. GC-IR, GC-MS, HPLS-DAS, etc.). The SPC file format does not provide a way to store both pieces of information in the same file. However, all Galactic software products will automatically "link" the

data together if the spectral data is stored in a Galactic SPC file with a ".SPC" file extension, and the chromatogram data is stored in the same directory, with the same base name, but using a ".CGM" file extension (i.e. DATA.SPC and DATA.CGM). Therefore, when presented with data file from hyphenated techniques, many Galactic-written file conversion routines will store the spectral data in a multi-record ".SPC" file, and the chromatogram(s) in a separate ".CGM" file.

Extension	Data File Type	
SPC	Spectrum (General case)	
CGM	Chromatogram (General case)	
GC	Gas Chromatogram	
LC	Liquid Chromatogram	
HC	HPLC Chromatogram	
FIR	Infrared (FT-IR) Spectrum	
IR	Near Infrared (NIR) Spectrum	
VIS	Visible Spectrum	
UV	Ultraviolet (Visible) Spectrum	
XRY	X-ray Spectrum	
MS	Mass Spectrum	
NMR	Nuclear Magnetic Resonance (NMR) Spectrum	
RMN	Raman Spectrum	
EFL	Fluorescence Spectrum	
AAS	Atomic Spectrum	
DAS	Diode Array Spectrum	
OTR	Other Data Type	

Universal Log Text Parameters

Log Text Information Format

All information in the Log Text block must be in ASCII characters. The parameter keys are stored in the Log Text block of the SPC file as a string characters, followed by an equals sign ("=", ASCII 61), and another string of characters. Each entry should terminate with a carriage return (CR, ASCII 13) and a line feed (LF, ASCII 10). The lines can be in one of the following formats:

```
KeyName = ########

or

KeyName = Text
```

where "######" represents a numerical value as ASCII text, and "Text" represents a text string of characters. Each particular parameter always uses either a numeric value or a quoted text value (although numbers may optionally be followed by text after a space). Most parameters use numbers. The following sections denote what parameters are followed by ASCII values, and which are followed by text strings. Some parameters that are followed by text strings require that the text match a certain string in a list of allowed strings. All routines that read information from the Log Text block must be written such that the key names are not context sensitive (i.e. "Scans = " can be interpreted as easily as "SCANS ="), but by convention the key names are capitalized when writing information into the Log Text.

Note that it is not required to use every parameter for a given file type. This is especially true if the parameter does not exist in the file being translated or is not used by the piece of instrumentation/software writing the SPC file. However, all attempts should be made to translate as many or all parameters as possible to maintain data portability between different file formats.

For example, a Galactic SPC file containing an FT-IR spectrum may contain the following information in the Log Text block:

```
SCANS = 10

SPEED = 15

LWN = 15799.7

APOD = Triangular

PHASE = Self
```

Note also that some numeric parameter values have inherent units. However, the units are not always expressed. In the Log Text, the units are optional but <u>must</u> be separated from the number by at least one space character (ASCII 32) if used:

```
BEGX = 4000 cm-1
```

General Data File Description Parameters

Before considering storing information in the Log Text block, all reasonable attempts must be made to fill in the information in the main header block of the SPC file. The main header block stores most of the general information required to properly display data in Galactic software. This includes the X limits, number of points, number of sub-files, the axis types, resolution text, memo text, and instrument source text. Therefore, such items such must always be stored in the main header block in the areas provided (see SPC.H). However, these same items can optionally be duplicated in the Log Text block

The following table lists the Log Text keys for the common parameters that can possibly be found within every Galactic SPC file. Note that these are duplicates of parameters that are already part of the main SPC file header.

Log Text Key	Туре	Description
BEGX	Value	X position of first trace point
ENDX	Value	X position of last trace point
NPTS	Value	Number of trace points (0=XY pairs)
BEGZ	Value	Z position of first subfile
ENDZ	Value	Z position of last subfile
NSUBS	Value	Number of subfiles
XTYPE	Value	X axis type number (as defined in SPC.H)
YTYPE	Value	Y axis type number (as defined in SPC.H)
ZTYPE	Value	Z axis type number (as defined in SPC.H)
RES	Text string	Collection X axis resolution (not necessarily the same
		as the data point spacing!)
		(Set to -1 if automatic or unknown)
NAME	Text string	Data file name (optionally including path and
		extension)
MEMO	Text string	Primary comment/description text (up to 129 bytes)
COMMENT	Text string	Secondary comment/description text (use MEMO
		first)
USER	Text string	User's or Analyst's name
MODEL	Text string	The unique instrument name without spaces. This is
		used in log text and with "IGET" to determine the type
		of instrument or data acquisition or the file converter
		name. By convention "MODEL=name" is placed at
		the beginning of log text for a file.

Galactic Enumerated Unit Labels

The SPC.H file defines the enumerated X, Y and Z axis unit type numbers. For convenience, they are also duplicated below. Be aware that this list is NOT used in compiling code and the ONLY source of the currently defined unit strings is maintained in these source code header files. In addition, it is not necessary to use one of the enumerated type values listed. The SPC file main header block provides for "custom text" unit labels to be stored in the header instead. Note however, that when using "custom unit" labels, some data processing tasks in Galactic software require that the data use a certain enumerated unit type, and that conversions from these SPC files out to other foreign formats may not be reliable. Also note that not all unit type values are supported by all Galactic products as the lists have been extended as the new products and routines for new data types were developed.

Possible settings for XTYPE and ZTYPE and the corresponding unit label:

Value	X or Z Unit Label			
0	Arbitrary			
1	Wavenumber (cm-1)			
2	Micrometers			
3	Nanometers			
4	Seconds			
5	Minutes			
6	Hertz			
7	Kilohertz			
8	Megahertz			
9	Mass (M/z)			
10	Parts per million			
11	Days			
12	Years			
13	Raman Shift (cm-1)			
14	Electron Volts (eV)			
15	X,Y,Z text labels in fcatxt (old 4Dh version only)			
16	Diode Number			
17	Channel			
18	Degrees			
19	Temperature (F)			
20	Temperature (C)			
21	Temperature (K)			
22	Data Points			
23	Milliseconds (mSec)			
24	Microseconds (uSec)			
25	Nanoseconds (nSec)			
26	Gigahertz (GHz)			
27	Centimeters (cm)			
28	Meters (m)			
29	Millimeters (mm)			
30	Hours			
255	Double interferogram (no display labels)			

Possible settings for YTYPE and the corresponding unit label:

Value	Y Unit Label
-1	Reference Arbitrary Energy
0	Arbitrary Intensity
1	Interferogram
2	Absorbance
3	Kubelka-Munk
4	Counts
5	Volts
6	Degrees
7	Milliamps
8	Millimeters
9	Millivolts
10	Log (1/R)
11	Percent
12	Intensity
13	Relative Intensity
14	Energy
15	***** NOT USED *****
16	Decibel
17 – 18	***** NOT USED ******
19	Temperature (F)
20	Temperature (C)
21	Temperature (K)
22	Index of Refraction [N]
23	Extinction Coeff. [K]
24	Real
25	Imaginary
26	Complex
128	Transmission (ALL TYPES >=128 ARE ASSUMED TO HAVE INVERTED PEAKS!)
129	Reflectance
130	Arbitrary or Single Beam with Valley Peaks
131	Emission

FT-IR Data Files

The following Log Text keys can be used in Galactic SPC files containing FT-IR spectra.

SCANS SCANSBG Value Number of co-added scans Number of co-added scans in background file NAMEBG Text string Background spectrum filename GAIN Value Detector Gain Factor SPEED Text string Scan Speed or Velocity text (may be text like "Slow") VELOCITY Value Scan Speed or Velocity number (-1 = automatic or unknown) LWN Value Laser Wavenumber (usually ~15799.7 cm-1) RAMANFREQ RAMANPWR Value Raman Laser Frequency in Wavenumbers RAMANPWR Value B-Stop aperture width BSTOP Value B-Stop aperture width DET Text string Detector type name SRC Text string Source type name BSPLITTER Text string Delay before scanning ZFF Value Zero Filling Factor (0=none, 1=2x, 2=4x, etc.) IRMODE APOD Text item Apodization function type name: "None" (boxcar apodization) "Triangular" "Weak" Norton-Beer "Medium" Norton-Beer "Happ-Genzel" "Bessel" "Cosine" "Filler" "Kaiser" Bessel "Gaussian" "Tradezoidal" PHASE Text item Phase correction type name: "Self" "Magnitude" "Background" "Stored" Phase "None" Number of interferogram points used for phase correction "Correction "Correct	Log Text Key	Туре	Description
SCANSBG Value Number of co-added scans in background file	SCANS	Value	Number of co-added scans
GAIN Value Detector Gain Factor SPEED Text string Scan Speed or Velocity text (may be text like "Slow") VELOCITY Value Scan Speed or Velocity number (-1 = automatic or unknown) LWN Value Laser Wavenumber (usually ~15799.7 cm-1) RAMANFREQ Value Raman Laser Frequency in Wavenumbers RAMANPWR Value Raman Laser Power JSTOP Value B-Stop aperture width BSTOP Value B-Stop aperture width DET Text string Detector type name SRC Text string Source type name SRC Text string Beam Splitter name PURGE Value Purge Delay before scanning ZFF Value Zero Filling Factor (0=none, 1=2x, 2=4x, etc.) IRMODE Value FITR Laser Sampling Frequency and Raman Mode 2 = Mid-IR mode 1 = Near-IR mode 2 = Raman-shift (negative => Raman) APOD Text item Apodization function type name: "None" (boxcar apodization) "Triangular" "Weak" Norton-Beer "Medium" Norton-Beer "Medium" Norton-Beer "Medium" Norton-Beer "Kaiser" Bessel "Cosine" "Filler" "Kaiser" Bessel "Gaussiann" "Tradezoidal" PHASE Text item Phase correction type name: "Self" "Magnitude" "Background" "Stored" Phase "None" Wumber of interferogram points used for phase	SCANSBG	Value	
GAIN Value Detector Gain Factor SPEED Text string Scan Speed or Velocity text (may be text like "Slow") VELOCITY Value Scan Speed or Velocity number (-1 = automatic or unknown) LWN Value Laser Wavenumber (usually ~15799.7 cm-1) RAMANFREQ Value Raman Laser Frequency in Wavenumbers RAMANPWR Value Raman Laser Power JSTOP Value B-Stop aperture width BSTOP Value B-Stop aperture width DET Text string Detector type name SRC Text string Source type name SRC Text string Beam Splitter name PURGE Value Purge Delay before scanning ZFF Value Zero Filling Factor (0=none, 1=2x, 2=4x, etc.) IRMODE Value FITR Laser Sampling Frequency and Raman Mode 2 = Mid-IR mode 1 = Near-IR mode 2 = Raman-shift (negative => Raman) APOD Text item Apodization function type name: "None" (boxcar apodization) "Triangular" "Weak" Norton-Beer "Medium" Norton-Beer "Medium" Norton-Beer "Medium" Norton-Beer "Kaiser" Bessel "Cosine" "Filler" "Kaiser" Bessel "Gaussiann" "Tradezoidal" PHASE Text item Phase correction type name: "Self" "Magnitude" "Background" "Stored" Phase "None" Wumber of interferogram points used for phase	NAMEBG	Text string	Background spectrum filename
Value	GAIN		
Value Scan Speed or Velocity number (-1 = automatic or unknown)	SPEED	Text string	Scan Speed or Velocity text (may be text like "Slow")
LWN	VELOCITY	Value	Scan Speed or Velocity number (-1 = automatic or
RAMANFREQ Value Raman Laser Frequency in Wavenumbers RAMANPWR Value Raman Laser Power JSTOP Value J-Stop aperture width BSTOP Value B-Stop aperture width DET Text string Detector type name SRC Text string Source type name SPLITTER Text string Beam Splitter name PURGE Value Zero Filling Factor (0=none, 1=2x, 2=4x, etc.) IRMODE Value Zero Filling Factor (0=none, 1=2x, 2=4x, etc.) FTIR Laser Sampling Frequency and Raman Mode 2 = Mid-IR mode 1 = Near-IR mode -2 = Raman-shift (negative => Raman) APOD Text item Apodization function type name: "None" (boxcar apodization) "Triangular" "Weak" Norton-Beer "Medium" Norton-Beer "Medium" Norton-Beer "Filler" "Kaiser" Bessel "Gaussian" "Tradezoidal" PHASE Text item Phase correction type name: "Self" "Magnitude" "Background" "Stored" Phase "None" PHASEPTS Value Number of interferogram points used for phase	LWN	Value	/
RAMANPWR Value Raman Laser Power JSTOP Value J-Stop aperture width BSTOP Value B-Stop aperture width DET Text string Detector type name SRC Text string Source type name BSPLITTER Text string Beam Splitter name PURGE Value Purge Delay before scanning ZFF Value Zero Filling Factor (0=none, 1=2x, 2=4x, etc.) IRMODE Value FTIR Laser Sampling Frequency and Raman Mode 2 = Mid-IR mode -2 = Raman-shift (negative => Raman) APOD Text item Apodization function type name: "None" (boxcar apodization) "Triangular" "Weak" Norton-Beer "Medium" Norton-Beer "Happ-Genzel" "Bessel" "Cosine" "Filler" "Kaiser" Bessel "Gaussian" "Tradezoidal" PHASE Text item Phase correction type name: "Self" "Magnitude" "Background" "Stored" Phase "None" PHASEPTS Value Number of interferogram points used for phase			
JSTOP			
BSTOP Value B-Stop aperture width DET Text string Detector type name SRC Text string Source type name BSPLITTER Text string Beam Splitter name PURGE Value Purge Delay before scanning ZFF Value Zero Filling Factor (0=none, 1=2x, 2=4x, etc.) IRMODE Value FTIR Laser Sampling Frequency and Raman Mode 2 = Mid-IR mode 1 = Near-IR mode 2 = Raman-shift (negative => Raman) APOD Text item Apodization function type name: "None" (boxcar apodization) "Triangular" "Weak" Norton-Beer "Happ-Genzel" "Bessel" "Cosine" "Filler" "Kaiser" Bessel "Gaussian" "Tradezoidal" PHASE Text item Phase correction type name: "Self" "Magnitude" "Background" "Stored" Phase "None" PHASEPTS Value Number of interferogram points used for phase			
DET Text string Detector type name SRC Text string Source type name BSPLITTER Text string Beam Splitter name PURGE Value Purge Delay before scanning ZFF Value Zero Filling Factor (0=none, 1=2x, 2=4x, etc.) IRMODE Value FTIR Laser Sampling Frequency and Raman Mode 2 = Mid-IR mode 1 = Near-IR mode 2 = Raman-shift (negative => Raman) APOD Text item Apodization function type name: "None" (boxcar apodization) "Triangular" "Weak" Norton-Beer "Medium" Norton-Beer "Strong" Norton-Beer "Happ-Genzel" "Bessel" "Cosine" "Filler" "Kaiser" Bessel "Gaussian" "Tradezoidal" PHASE Text item Phase correction type name: "Self" "Magnitude" "Background" "Stored" Phase "None" PHASEPTS Value Number of interferogram points used for phase			
SRC Text string Source type name BSPLITTER Text string Beam Splitter name PURGE Value Purge Delay before scanning ZFF Value Zero Filling Factor (0=none, 1=2x, 2=4x, etc.) IRMODE Value FTIR Laser Sampling Frequency and Raman Mode 2 = Mid-IR mode 1 = Near-IR mode -2 = Raman-shift (negative => Raman) APOD Text item Apodization function type name: "None" (boxcar apodization) "Triangular" "Weak" Norton-Beer "Medium" Norton-Beer "Strong" Norton-Beer "Happ-Genzel" "Bessel" "Cosine" "Filler" "Kaiser" Bessel "Gaussian" "Tradezoidal" PHASE Text item Phase correction type name: "Self" "Magnitude" "Background" "Stored" Phase "None" PHASEPTS Value Number of interferogram points used for phase			
BSPLITTER Text string Beam Splitter name PURGE Value Purge Delay before scanning ZFF Value Zero Filling Factor (0=none, 1=2x, 2=4x, etc.) IRMODE FIIR Laser Sampling Frequency and Raman Mode 2 = Mid-IR mode 1 = Near-IR mode -2 = Raman-shift (negative => Raman) APOD Text item Apodization function type name: "None" (boxcar apodization) "Triangular" "Weak" Norton-Beer "Medium" Norton-Beer "Happ-Genzel" "Bessel" "Cosine" "Filler" "Kaiser" Bessel "Gaussian" "Tradezoidal" PHASE Text item Phase correction type name: "Self" "Magnitude" "Background" "Background" "Stored" Phase "None" PHASEPTS Value Number of interferogram points used for phase			, ,
PURGE ZFF Value Zero Filling Factor (0=none, 1=2x, 2=4x, etc.) IRMODE Value FTIR Laser Sampling Frequency and Raman Mode 2 = Mid-IR mode 1 = Near-IR mode -2 = Raman-shift (negative => Raman) APOD Text item Apodization function type name: "None" (boxcar apodization) "Triangular" "Weak" Norton-Beer "Medium" Norton-Beer "Happ-Genzel" "Bessel" "Cosine" "Filler" "Kaiser" Bessel "Gaussian" "Tradezoidal" PHASE Text item Phase correction type name: "Self" "Magnitude" "Background" "Stored" Phase "None" PHASEPTS Value Number of interferogram points used for phase			
Zero Filling Factor (0=none, 1=2x, 2=4x, etc.) IRMODE			,
IRMODE Value FTIR Laser Sampling Frequency and Raman Mode 2 = Mid-IR mode 1 = Near-IR mode -2 = Raman-shift (negative => Raman) APOD Text item Apodization function type name: "None" (boxcar apodization) "Triangular" "Weak" Norton-Beer "Medium" Norton-Beer "Medium" Norton-Beer "Happ-Genzel" "Bessel" "Cosine" "Filler" "Kaiser" Bessel "Gaussian' "Tradezoidal" PHASE Text item Phase correction type name: "Self" "Magnitude" "Background" "Stored" Phase "None" PHASEPTS Value Number of interferogram points used for phase			
2 = Mid-IR mode 1 = Near-IR mode -2 = Raman-shift (negative => Raman) APOD Text item Apodization function type name: "None" (boxcar apodization) "Triangular" "Weak" Norton-Beer "Medium" Norton-Beer "Strong" Norton-Beer "Happ-Genzel" "Bessel" "Cosine" "Filler" "Kaiser" Bessel "Gaussian" "Tradezoidal" PHASE Text item Phase correction type name: "Self" "Magnitude" "Background" "Stored" Phase "None" PHASEPTS Value Number of interferogram points used for phase			
APOD Text item Apodization function type name: "None" (boxcar apodization) "Triangular" "Weak" Norton-Beer "Medium" Norton-Beer "Strong" Norton-Beer "Happ-Genzel" "Bessel" "Cosine" "Filler" "Kaiser" Bessel "Gaussian" "Tradezoidal" PHASE Text item Phase correction type name: "Self" "Magnitude" "Background" "Stored" Phase "None" PHASE PHASE Value Number of interferogram points used for phase	IRMODE	value	
APOD Text item Apodization function type name: "None" (boxcar apodization) "Triangular" "Weak" Norton-Beer "Medium" Norton-Beer "Strong" Norton-Beer "Happ-Genzel" "Bessel" "Cosine" "Filler" "Kaiser" Bessel "Gaussian" "Tradezoidal" PHASE Text item Phase correction type name: "Self" "Magnitude" "Background" "Stored" Phase "None" PHASEPTS Value Number of interferogram points used for phase			
APOD Text item Apodization function type name: "None" (boxcar apodization) "Triangular" "Weak" Norton-Beer "Medium" Norton-Beer "Strong" Norton-Beer "Happ-Genzel" "Bessel" "Cosine" "Filler" "Kaiser" Bessel "Gaussian" "Tradezoidal" PHASE Text item Phase correction type name: "Self" "Magnitude" "Background" "Stored" Phase "None" PHASEPTS Value Number of interferogram points used for phase			
"None" (boxcar apodization) "Triangular" "Weak" Norton-Beer "Medium" Norton-Beer "Strong" Norton-Beer "Happ-Genzel" "Bessel" "Cosine" "Filler" "Kaiser" Bessel "Gaussian" "Tradezoidal" PHASE Text item Phase correction type name: "Self" "Magnitude" "Background" "Stored" Phase "None" PHASEPTS Value Number of interferogram points used for phase	ADOD	T ('t	
"Triangular" "Weak" Norton-Beer "Medium" Norton-Beer "Strong" Norton-Beer "Happ-Genzel" "Bessel" "Cosine" "Filler" "Kaiser" Bessel "Gaussian" "Tradezoidal" PHASE Text item Phase correction type name: "Self" "Magnitude" "Background" "Stored" Phase "None" PHASEPTS Value Number of interferogram points used for phase	APOD	I ext Item	
"Weak" Norton-Beer "Medium" Norton-Beer "Strong" Norton-Beer "Happ-Genzel" "Bessel" "Cosine" "Filler" "Kaiser" Bessel "Gaussian" "Tradezoidal" PHASE Text item Phase correction type name: "Self" "Magnitude" "Background" "Stored" Phase "None" PHASEPTS Value Number of interferogram points used for phase			
"Medium" Norton-Beer "Strong" Norton-Beer "Happ-Genzel" "Bessel" "Cosine" "Filler" "Kaiser" Bessel "Gaussian" "Tradezoidal" PHASE Text item Phase correction type name: "Self" "Magnitude" "Background" "Stored" Phase "None" PHASEPTS Value Number of interferogram points used for phase			
"Strong" Norton-Beer "Happ-Genzel" "Bessel" "Cosine" "Filler" "Kaiser" Bessel "Gaussian" "Tradezoidal" PHASE Text item Phase correction type name: "Self" "Magnitude" "Background" "Stored" Phase "None" PHASEPTS Value Valu			
"Happ-Genzel" "Bessel" "Cosine" "Filler" "Kaiser" Bessel "Gaussian" "Tradezoidal" PHASE Text item Phase correction type name: "Self" "Magnitude" "Background" "Stored" Phase "None" PHASEPTS Value Valu			
"Bessel" "Cosine" "Filler" "Kaiser" Bessel "Gaussian" "Tradezoidal" PHASE Text item Phase correction type name: "Self" "Magnitude" "Background" "Stored" Phase "None" PHASEPTS Value Value Value Value Value Value Value			
"Cosine" "Filler" "Kaiser" Bessel "Gaussian" "Tradezoidal" PHASE Text item Phase correction type name: "Self" "Magnitude" "Background" "Stored" Phase "None" PHASEPTS Value			
"Filler" "Kaiser" Bessel "Gaussian" "Tradezoidal" PHASE Text item Phase correction type name: "Self" "Magnitude" "Background" "Stored" Phase "None" PHASEPTS Value Value Value Value Value Value Value Value "Filler" "Massel "Tradezoidal" "Austria "Tradezoidal" "Tradezoidal" "Tradezoidal" "Tradezoidal" "Tradezoidal" "Tradezoidal" "Name: "Self" "Magnitude" "Background" "Stored" Phase "None"			
"Kaiser" Bessel "Gaussian" "Tradezoidal" PHASE Text item Phase correction type name: "Self" "Magnitude" "Background" "Stored" Phase "None" PHASEPTS Value			
"Gaussian" "Tradezoidal" PHASE Text item Phase correction type name: "Self" "Magnitude" "Background" "Stored" Phase "None" PHASEPTS Value			
PHASE Text item Phase correction type name: "Self" "Magnitude" "Background" "Stored" Phase "None" PHASEPTS Value Value Value Value Value Tradezoidal" Phase correction type name: "Self" "Magnitude" "Background" "Stored" Phase "None"			
PHASE Text item Phase correction type name: "Self" "Magnitude" "Background" "Stored" Phase "None" PHASEPTS Value Value Value Value Value Value Phase correction type name: "Self" "Magnitude" "Background" "Stored" Phase "None" Phase "None"			
"Self" "Magnitude" "Background" "Stored" Phase "None" PHASEPTS Value Number of interferogram points used for phase	PHASE	Text item	
"Background" "Stored" Phase "None" PHASEPTS Value Number of interferogram points used for phase			· ·
"Background" "Stored" Phase "None" PHASEPTS Value Number of interferogram points used for phase			"Magnitude"
"Stored" Phase "None" PHASEPTS Value Number of interferogram points used for phase			
"None" PHASEPTS Value Number of interferogram points used for phase			"Stored" Phase
PHASEPTS Value Number of interferogram points used for phase			
	PHASEPTS	Value	
OUTOUIOT			correction

Log Text Key	Туре	Description
IGRAMSIDE	Text item	Interferogram mode name:
		"Double" sided
		"Single" sided
		"Left" single sided
		"Right" single sided
IGRAMDIR	Text item	Interferogram collection direction name:
		"Bi" interferometer scans in both direction
		"Uni" interferometer scans in one direction
		"Forward" uni-directional
		"Reverse" uni-directional
POLARIZER	Value	Polarizer Angle
LOWPASS	Value	Low Pass Filter Cutoff
HIGHPASS	Value	High Pass Filter Cutoff
FILTER	Text string	Optical Filter Name
SMOOTH	Value	Number of averaged readings per data point.
BEAM	Text item	Single beam mode name:
		"Sample" (Does not replace existing background).
		"Background" for future ratioing with sample.
AVGTIME	Value	Time for A/D averaging of each point reading (msec).
		If the A/D has a fixed time and point readings are
		averaged, then this time should be divided by the
		approximate A/D time to get the number of readings.
BDELAY	Value	Begin delay in seconds
SDELAY	Value	Scan delay in seconds
CHANNEL	Value	Detector or beam path channel used for acquisition.

NIR (Quant) Data Files

Many NIR instrument software packages use multi-record file formats where each record contains a spectrum and additional information and/or data related to the individual sample. Since the Galactic SPC multifile format does not allow for easy storage of this information within the subfile records, the extra information should be imported according to the following specifications.

In addition, these files can have "constituent" information in them which is used for quantitative calibration purposes. This information is not imported into the Galactic SPC file, but is stored in a separate ASCII file known as a Galactic Calibration File List (CFL). The format of the CFL files is described here as well as the Log Text keys.

The following Log Text keys can be used in Galactic SPC files containing NIR spectra.

Log Text Key	Туре	Description
NODE	Value	Identifies the instrument connection node number
PPROCXX	Text item	Text description of additional processing applied to entire data set. There can be multiple entries allowed and they are NOT required to be in numerical order, and are NOT required to have a complete sequence (i.e. file could have only PPROC2 and PPROC8). Some possible values are, (although others are allowed): "Constant Multiplication" "File Multiplication" "Constant Division" "File Division" "Constant Addition" "File Multiplication" "Constant Subtraction" "File Subtraction" "Normalize" "Mean Center" "Variance Scaled" "Smooth" "SG Smooth" "SG Smooth" "Second Derivative" "Log Y" "MSC" (Multiplicative Scatter Correction) "SG 1st Derivative" "SG 2nd Derivative" "SNV" (Standard Normal Variate) "SNV + Detrend" "Divide by Wavelength (xxx)" (xxx is wavelength value) "FFT"
CHANNEL	Value	Detector or beam path channel used for acquisition.

Log Text Key	Туре	Description
SENSORCNT	Value	If the "SENSORCNT=" value is non-zero, each "SUBFILEx=" line is followed by a "SENSORVALx=" line where "x" is the subfile number beginning with "1" for the first subfile.
SENSORNAMEX	Text string	Sensor names. If the sensor count is non-zero, then the "SENSORCNT=" label is followed by "SENSORNAME1=" "SENSORNAME2=", etc. up to the number of sensors identified by SENSORCNT.
SUBFILEX	Formatted text string	Extra text information for subfile number "x". This is a set of comma delimited strings in the following format: MM/DD/YY, HH:MM:SS, "Memo text" Example: SUBFILE43= 04/03/1996, 22:31:45, "MeOH"
SENSORVALx	Comma delimited values	Sensor values for subfile number "x" as comma delimited floating point values. Must be SENSORCNT number of values.

There can be (optionally) at least one subfile label section for each subfile in the file. If used, the subfile section starts with the label "SUBFILEx=" (where "x" is the subfile number beginning with "1") followed by the subfile date and time. The date and time are fixed width fields of the form MM/DD/YYYY HH:MM:SS separated by commas. The time should be in 24 hour format. The last item on the "SUBFILEx=" label line is the subfile description in double quotes. If no description is available, "" is used.

If the "SENSORCNT=" value is non-zero, each "SUBFILEx=" line is followed by a "SENSORVALx=" line where "x" is the again subfile number. Following the label are the sensor values as comma delimited text of floating point values.

Constituent Concentration Data - Calibration File List (CFL File)

Many NIR instrument file formats store concentration values and constituent name information used to create quantitative calibrations from the spectral data. The Galactic SPC file has no room for this information as the number of constituents and thus the size of the extra data is completely dynamic. While it is possible that the Log Text key name specifications could be extended in the future to include this information as part of the Galactic SPC file, the current mechanism is to write the information into a Galactic ASCII text Calibration File List (CFL) file. There are currently two formats allowed for the CFL file, however all future converters should only import to the "New" format. When importing the concentration data, the base name of the CFL file should match the base name of the output Galactic SPC data file (i.e. importing the foreign file "SAMPLE.DAT" would generate a Galactic multifile of the spectral data called "SAMPLE.SPC", and a Galactic CFL file called "SAMPLE.CFL" of the concentration data).

When exporting data to a foreign NIR data file format, the Galactic-written converters should check for the existence of a matching CFL file. If there is a CFL file that matches the base name of the Galactic SPC file being exported, the conversion routine MUST export both the spectral data and concentration data. Also note that export converters will need to support BOTH CFL file formats as there is no guarantee that the CFL file is always in the new format and the user is not expected to manually translate it.

The old CFL format is from previous versions of Galactic's quantitative applications and is still supported for import. However, this format is no longer used and new foreign file converters and Array Basic applications are only expected to save concentration data into the new CFL format. However, file converters must support the both the old and new CFL file formats on export.

Old CFL Format Example

	JMBER			NTS					
	JMBER								
	e Calibra	ation file	e for PLS	Splus tu	torial - S	ynthetic	test mix	tures	
Α									
В									
С									
1	2	3	4	5	6	7	8	9	10
GOOD	#1								
.1	.35	.55							
GOOD									
.15	.4	.45							
GOOD									
.2	.45	.35							
GOOD		.00							
.25	.5	.25							
GOOD		.23							
		•							
.3	.1	.6							
GOOD		_							
.35	.15	.5							
GOOD									
.4	.2	.4							
GOOD	#8								
.45	.25	.3							

In the old CFL file format it is very important that there are NO BLANK LINES, especially at the top of the file.

Line 1	Contains the number of components (M). A comment may follow, however there must be AT LEAST ONE SPACE between the comment and the number.
Line 2	Contains the number of samples (N). A comment may follow, however there must be AT LEAST ONE SPACE between the comment and the number.
Line 3	Contains a comment line.
Line 4	Contains the name of the first component. Continue typing the name of each component on a separate line up to the number of components (M) listed in Line 1. Component names may be up to 15 characters long.
Line 4+M	The next line simply labels 10 columns for visual examination.
Line 5+M	The rest of the lines contain the sample file name (with NO FILE EXTENSION!), and on the following line(s), the corresponding concentration data. Follow the example as shown. The values may be delimited by a space, comma or tab. If you have more than 10 components in the unknown, the concentration data may continue on the next line but the name of the next file must start on the following line. There must be N (from Line 2) sample name, concentration line pairs in the file.

The new Galactic CFL format was designed specifically for users that create calibration lists using text editor or spreadsheet programs. It is currently only supported by the PLSplus/IQ quantitative analysis product, however all future products will use it as well. It uses a comma delimited format to store the spectral file names, and concentration values (if applicable). Therefore, it is not necessary to have the number of components and number of samples in the file header.

New CFL Format

```
Sample Calibration file for PLSplus/IQ tutorial - Synthetic test mixtures
File/Comp,Constit A,Constit B,Constit C
                      0.35.
GOOD.SPC #1,0.10,
                             0.55
GOOD.SPC #2,0.15,
                      0.40,
                             0.45
GOOD.SPC #3,0.20,
                      0.45,
                             0.35
GOOD.SPC #4,0.25,
                      0.50,
                             0.25
GOOD.SPC #5,0.30,
                      0.10,
                             0.60
GOOD.SPC #6,0.35,
                      0.15,
                             0.50
                             0.40
GOOD.SPC #7,0.40,
                      0.20,
GOOD.SPC #8,0.45,
                      0.25,
                             0.30
```

In the new CFL file format, it is very import that there are NO BLANK LINES!. The number of samples (N) is determined by counting the number of non-blank lines after the component list line (Line 2). The count is stopped when the first blank line, or a line that has fewer concentration values than the first sample is encountered.

- Line 1 File comment (MUST NOT BE BLANK!).
- Line 2 Must start with "File/Comp", then list constituent names using commas as separators. All characters between the commas are considered significant, thus ", Stuff ," will result in a constituent name " Stuff ". The number of constituents (M) is determined by the number of comma separated constituent names found on this line. If there are no names following the "File/Comp" leader, then the rest of the file is assumed to have NO CONCENTRATION DATA and there must be no concentration values following the comma after the file name.
- Line 3 File name of 1st sample spectrum, followed by the comma separated concentration values of the constituents listed in line 2. Files can optionally have the extension, but if it is not present, then ".SPC" is assumed. When reading concentration values any non-numeric characters (i.e. spaces) between the commas are ignored.
- Line N+2 File name and concentrations for last sample spectrum.

UV/VIS (NIR) Data Files

The following Log Text keys can be used in Galactic SPC files containing UV/VIS/NIR spectra. Note that these are different from the parameters used by Galactic NIR (Quant) Data Files. This set of parameters is primarily used by Galactic ABC data acquisition drivers to store relevant data collection parameters. However, some foreign file formats may included this information as well, so this list is provided to help achieve data completeness for those files that do.

Log Text Key	Туре	Description
NAMESTD	Text string	Standard reference correction filename
NAMEZR	Text string	Zero reference (dark reference) correction filename
NAMEBG	Text string	Background or baseline spectrum filename
GAIN	Value	Detector Gain Factor
SLIT1	Value	Slit aperture width (-1 = automatic or unknown)
SLIT2	Value	Slit aperture width (-1 = automatic or unknown)
DET	Text string	Detector type name
SRC	Text string	Source type name
BEAM	Text item	Beam mode name: "Double" "Single Front" "Single Rear" "Double Reverse" "Dual Single"
BDELAY	Value	Begin Delay in seconds
SDELAY	Value	Scan delay in seconds
PMT	Value	Photo multiplier Tube voltage (-1 = automatic or unknown)
DETCHG	Value	Detector changeover (NIR to UV/VIS) wavelength in nm
DETCOR	Value	Detector correction 0 = off 1 = on
SRCCHG	Value	Source changeover wavelength in nm
GRTCHG	Value	Grating changeover wavelength in nm
INDEPENDENT	Value	Spectral range independent collection: 0 = off, use same instrument parameters for whole spectrum 1 = allow different parameters for UV/VIS and NIR range
SCANMODE	Text item	Scan type: "Spectrum" (0) "Time scan" (1) "Multi-wl time" scan (2)
SIGNOISE	Value	Signal to noise processing 0 = no S/N processing 1 = use S/N processing
SNLEVEL	Value	Acceptable S/N ratio
SNTIMEOUT	Value	Time out (in sec) for S/N processing in scan
NIR_RES	Value	NIR data interval (INDEPENDENT = 1 only)

Log Text Key	Туре	Description
NIR_AVERAGING	Value	NIR averaging time (INDEPENDENT = 1 only)
NIR_SBW	Value	NIR slit width (INDEPENDENT = 1 only)
NIR_ENERGY	Value	NIR energy level (INDEPENDENT = 1 only)
NIR_SLITHT	Value	NIR slit height (INDEPENDENT = 1 only):
		0 = Full
		1 = 1/3 height
CORRECTION	Value	Sample baseline correction:
		0 = off
		1 = on
CORR_TYPE	Text item	Type of correction name:
		"Normal" (0)
		"Zero line" (1)
		"Zero SRA" (2)
		"Zero*Std.Ref." (3) (uses NAMESTD file)
SCANTYPE	Text item	Scan type name:
		"Sample" (0)
		"Baseline" (1)
		"Zero line" (2)

NMR Data Files

The following information describes how the Galactic SPC format was extended to provide support for 1D NMR data. This makes use of the both the Log Text block (for storing acquisition parameters) and the Log Binary block (for storing the imaginary data in complex number data sets). Note that some Log text parameters are REQUIRED, and must be present in the Log Text block in order for Galactic software to properly process the data. Also, some parameters may only be present in processed NMR spectra, and not FID data.

NOTE: These parameters have been defined for 1D NMR data only. This specification will be extended if there is a decision made to accommodate 2D NMR data in the future.

NOTE: All the following parameters <u>MUST</u> appear as the first set of information in the Galactic file text header record. In addition, the last item in the NMR part of the Log Text parameters must be "NMREND=NMREND". This is used by the Galactic NMR processing routines to recognize if additional processing has been applied to the data file. It is also used to flag that there is no more NMR-related information in the Log Text after that point.

Complex (Imaginary) Data Storage - Binary Log Block

NOTE: The following information requires an understanding of the Galactic SPC binary format. Refer to the SPC.H header file for a complete description of the Log data area.

Most NMR instruments generate complex data (data that has arrays of both "real" and "imaginary" numbers). In order to meet the dual requirements that the data must display the "real" part of the spectrum properly for plotting, general manipulation and publication, but retain the "imaginary' part for future NMR-related manipulations (such as re-phasing), Galactic has designed a split file format.

In all cases, whether the data is an FID or Spectrum, only the "real" portion is stored in the main data block of the SPC file. The imaginary data is stored in an additional binary block as part of the Log Text block at the end of the file (see SPC.H for more information). Although the main data block in an SPC file can be stored as either 32 bit-block scaled integers or 32-bit IEEE floating point values, the "imaginary" binary data block can ONLY be stored as 32-bit IEEE floats. When loading an SPC file for display of the Y values, Galactic software only reads the data from the main data block, and ignores any information in the log text block.

When creating a Galactic NMR file for SPC data, the flogoff value in the SPCHDR structure MUST be set to the byte offset to the beginning of the LOGSTC structure. This normally immediately follows the main binary data block:

The logsizd describes the size (in bytes) of the complete log block (LOGSTC, binary block and text data). The logsizm should be the logsizd size of the block rounded to the next larger power of 4096. The logtxto is the byte offset FROM THE BEGINNING OF LOGSTC to the beginning of the log text data. In other words, the binary log data should immediately follow the LOGSTC structure in the file. The logbins specifies the size of the binary block (in bytes) that follows LOGSTC.

Note that the size of the main data block and the binary log block MUST be exactly the same. This is required since the Galactic NMR processing routines require that the "real" and "imaginary" array have exactly the same number of data points.

Informational Parameters

These parameters are used to include extra information about the sample and source of the data in the file. These are all optional parameters and are not required for proper processing.

Log Text Key	Туре	Description
INSTRUM	Text string	Text description of the instrument/software
SOLVENT	Text string	Text name of the solvent used (i.e. "CDCl3", "D20", "DMSO", "Acetone", "CD30D", "MeOH", etc.)
NUCLEUS	Text string	Nucleus name for tuning frequency "1H" for Proton "13C" for Carbon 13 "19F" for Fluorine 19 General format is "AtomicnumberElement" for all others.

Acquisition Parameters

These parameters are generally related to the modes used to collect the data from the NMR instrument. Many of these are REQUIRED since Galactic uses them during data processing.

Log Text Key	Туре	Description
DETMODE	Text item	Detection Quadrature mode name (REQUIRED): "SIM" for simultaneous (use Complex FFT) "SEQ" for sequential (use Redfield FFT) "REAL" for real-only (use Real FFT)
NUCFREQ	Value	Nucleus tuning frequency in MHz (REQUIRED) (aka Carrier Freq., Observed Freq., Center Freq.)
SW_HZ	Value	Spectral band width measured in Hertz (REQUIRED if SW_ppm not used) (aka Sweep width, Spectral width)
DWELL	Value	Dwell time in Seconds (Should equal 1/SW_Hz)
DELAY	Value	Acquisition delay in Seconds
ACQTIME	Value	Acquisition time in Seconds (REQUIRED) (Should be equal to NPTS*DWELL, or NPTS/SW_HZ.)
SCANS	Value	Actual number of scans performed (aka Count of actual responses recorded (CT))
REQSCAN	Value	Number of scans requested during acquisition (aka Number of transients (NT), Number of Acquisitions (NA))
FLTFREQ	Value	Filter Frequency in Hertz
PULSWD	Value	Pulse Width in Microseconds
SW_PPM	Value	Spectral band width measured in PPM (REQUIRED if SW_HZ not used) (aka Sweep width, Spectral width)
AB_APPS	Text boolean	File subsequently processed by AB program "TRUE" or "FALSE"
FIDFILE	Text string	Text filename (no path!) of Galactic FID file (*.FIG) corresponding to spectrum This tag can only exist if the spectrum or FID was either processed with Galactic's NMR routines OR was read into the NMR routine. In general, only files processed using Galactic's NMR routines will have this field set.
FID	Text boolean	Current data type indicator (REQUIRED) "TRUE" for raw FID data "FALSE" for FFT'd data (REQUIRES Processing Parameters below)
SPC_REAL	Text boolean	REAL data only, no IMAG record. (REQUIRED) "TRUE" or "FALSE"

Processing Parameters

These parameters are generally added to the log text block by the Galactic NMR processing applications. In general, file converters will not be able to get this information from the foreign file format or the values will be different than what is expected by the Galactic processing routines.

Log Text Key	Туре	Description
DCOFF	Value	Direct Current Offset correction. This is the data point number threshold to use when calculating the DC offset. This value indicates the number of data points, starting from the last value, and ending at the first data point, that is to be used in the DC offset calculation.
ZFF	Text item	Zero Filling Factor Text field that is used to indicate then 2 ⁿ K value by which to increase the data size by. "NONE" "2xNPTS" (i.e. 2 times the number of data points in the file) "4xNPTS" "8xNPTS" "16xNPTS" "32xNPTS"
APOD	Text item	Apodization filter function type name. (Note this is the same key for Universal Parameters, but the list of allowed strings is different.) "NONE" "MATCH_EXP" (Matched Exponential) "MATCH_GAUSS" (Matched Gaussian) "BOXCAR" (Boxcar) "COS_SQ" (Cosine) "SINE_BELL" (Sine) "TRAPEZOID" (Trapezoid) "LORENTZ-GAUSS" (Lorentz-Gauss) "OPT_RES" (Optimal Resolution)
APODP(0) APODP(1)	Value	Some apodization filters require one or more fit parameter (REQUIRED if APOD present)
PHMOD	Text item	Phase mode used for spectrum (NOT SUPPORTED YET) "NORMAL" (Normal spectrum phasing using PH0,PH1,PV0&PV1) "MAGNITUDE" (Magnitude spectrum = sqrt(Real^2+Imag^2)) "POWER" (Power spectrum = (Real^2+Imag^2)) (Note if PHMOD<>"Normal" then user cannot rephase.)
PH0	Value	Zero order phase angle (degrees)
PH1	Value	First order phase angle (degrees)
PV0	Value	Zero order pivot (data point #, 0=first)
PV1	Value	First order pivot (data point #, 0=first)
AUTOTYP	Text item	Text name of autophasing method used (NOT SUPPORTED)

Log Text Key	Туре	Description
SPC_REV	Text boolean	Reverse spectrum flag
		"TRUE" the file has been flipped
		"FALSE" file read in does not change
DELTA_PPM	Value	Shift in PPM is the value that is subtracted from the
		full range beginning value (SW_PPM) and ending
		value (0) resulting in the new range in PPM.

Integration Parameters

The following are used to display the integrals for the file when it is loaded into the Galactic NMR processing routines. These are entirely optional, but if one is present, they all must be.

Log Text Key	Туре	Description
INTBAS	Text boolean	All integral regions were baselined "TRUE" or "FALSE"
INTOFF	Text boolean	Integral regions were offset "TRUE" or "FALSE"
BASE	Text boolean	Interactive baseline indicator "TRUE" or "FALSE"
NORM	Value	Region normalization value
MASK_ON	Text boolean	Use MASK regions to first excluding spectral information for calculating integrals and for display purposes "TRUE" or "FALSE"
IREGNx	Space delimited values	Integral parameters for a single integral region. There can be any number of IREGNx log text entries (where "x" is the region number), and they must appear in sequence and in increasing order (i.e. IREGN1, IREGN2, IREGN3, etc.) The data is stored as a list of text values that are space (ASCII 32) delimited. The format is as follows: Left_Edge Right_Edge Area Normalized_Area Multiplier Offset Tilt The edge values must be specified in terms of PPM The Offset and Tilt values are the baseline correction parameters for the integral region. They are specified in terms of a fraction (0-100%) of 1% of the largest Y value in the masked real data array.
MREGNx	Space delimited values	Integral masking parameters for a single mask region. There can be any number of MREGNx log text entries (where "x" is the region number), and they must appear in sequence and in increasing order (i.e. MREGN1, MREGN2, MREGN3, etc.)The data is stored as a list of text values that are space (ASCII 32) delimited. The format is as follows: Left_Edge Right_Edge The edge values must be specified in terms of PPM.

Peak Picking Parameters

These are used to display the peak positions for the file when it is loaded into the Galactic NMR processing routines. These are entirely optional, but if the peaktable is present, they all must be with the exception of the PEAK_MODE.

Log Text Key	Туре	Description
PEAK_TBL	Text string	Text filename of Galactic peak table for spectrum.
		Normally this will be <datafile>.PKN.</datafile>
PEAK_MODE	Text item	Peak picking mode used:
		"MASK_PKP" presets a masked region where peaks
		are not to be picked, then auto picks the rest of the
		spectrum using the threshold and option values.
		If the user adds values to the table manually, both modes become obsolete.
THRESH	Value	Peak picking threshold value (corresponds to Galactic
INKESH	value	Threshold Y Level)
SENS	Value	Peak picking sensitivity value (corresponds to Galactic
		Sensitivity)
SMOOTH	Value	Peak picking smoothing value (corresponds to Galactic
		Smoothing)
		Valid only if PICK_TYPE is set to EXTREMA (1)
PICK_TYPE	Value	Peak picking type:
		0 = Maximum
		1 = Extrema (also REQUIRES SMOOTH value)
WINDOW	Value	Size of peak window in data points when Adding a new
		peak. (Window used will actually be +/- WINDOW
PK MARK	Text item	from the current position). Peak mark type:
FK_WAKK	1 ext item	"None"
		"Arrow"
		"V"
		"Numbers"
PK LABEL	Value	Peak labeling type number. Corresponds to the
_		GRAMS internal Array Basic code number for peak
		labeling [i.e. 7 for Smart Labels].
LBL_FORM	Value	Peak label digits. Integer number of digits after
		decimal point in peak labels. Allowed values are 1 to
		4.
TR_OBJ	Value	Trace box object number. Allowed value is 1 only
PEAK_MASK	Text boolean	Peak marking masking used:
		"TRUE" the mask is used to block out portions of the
		spectrum. This is always TRUE if PEAK_MODE =
		MASK_PKP.
		"FALSE" for no masking

Log Text Key	Туре	Description
PREGNx	Space delimited values	Peak picking masking parameters for a single region. There can be any number of PREGNx log text entries (where "x" is the region number), and they must appear in sequence and in increasing order (i.e. PREGN1, PREGN2, PREGN3, etc.) The data is stored as a list of text values that are space (ASCII 32) delimited. The format is as follows: Left_Edge Right_Edge The edge values must be specified in terms of PPM.

Other Considerations to Note for NMR Data

Array Basic NMR processing programs must add the appropriate processing log text entries when manipulating data. This includes FFTing FIDs to spectra, phasing, and referencing. If the log text processing information is already present after import, but the data is FID type, then Galactic written processing retains the parameters imported from the instrument in memory and upon FFTing the FID will automatically apply those values. These values can be re-processed in the software at any point.

NMR FID data must always have the X axis in Seconds, with FFP=0 and FLP=ACQTIME [or NPTS/SW_Hz]. The maximum intensity of the data must be at the FFP end with the decay of the signal going towards the FLP end of the trace.

NMR spectra that have been FFT'd from FIDs must have X axis units in either Hertz (spanning - BANDWD/2 to +BANDWD/2) or PPM. Default is for the axis to be in PPM.

Mass Spectroscopy Data Files

Currently there are no established standards for Log Text parameters for Mass Spectroscopy data files. However, any standard that is established will have to differentiate between the different types of spectrometers (i.e. quadrapole, sector, TOF, etc.) and deal with specific parameters related to each. It is likely that this category can cover all single scan as well as hyphenated (i.e. GC-MS) Mass Spectroscopy data.

Chromatogram Data Files

Currently there are no established standards for Log Text parameters for Chromatography data files. However, any standard that is established will have to differentiate between the different types of chromatographs (i.e. HPLC, GC, SEC, GPC, etc.) and deal with specific parameters related to each. As chromatography data has a large variety of information required to maintain complete GLP compliance, it is likely that only a subset of all possible information available will be able to be carried along in a Galactic SPC file.

Diode Array Chromatography Detector Data Files

Currently there are no established standards for Log Text parameters for Diode Array Chromatography Detector data files.

Diode Array / CCD Spectrometer Data Files

Currently there are no established standards for Log Text parameters for Diode Array / CCD Spectrometer data files. Note this category is NOT the same as Diode Array Chromatography Detectors, but instead for CCD spectrometers used for applications like Raman spectroscopy, or spectroscopic imaging.

Xray Diffraction Spectrometer Data Files

Currently there are no established standards for Log Text parameters for Xray Diffraction Spectrometer data files.

Fluorescence Spectrometer Data Files

Currently there are no established standards for Log Text parameters for Fluorescence Spectrometer data files.

Data Files Created from AIA Data Interchange (ANDI/netCDF) Files

There are many other data parameters which may be needed from time to time. When a setting is not available in the above list, the AIA name from the AIA FTIR or Chromatography Data Interchange specifications can be used. Alternatively, a new Galactic parameter name can be devised. However, to avoid duplications, it is important that all names used, either AIA or Galactic, be recorded in a single place (which is this document for now). A Galactic parameter should be used if possible. A new Galactic name should be created only if the equivalent AIA name has not already been used.

In order to assure proper transfer of key AIA sample information from file to file, any AIA names used as keys in the log text should appear between "AIA_info_begin" and "AIA_info_end" entries. The key names should be the AIA names as they appear in the template specification, complete with underbars as needed.

For example (taken from an AIA Chromatography V1.0 file):

```
AIA info begin
dataset_completeness = C1+C2
aia_template_revision = 1.0
netcdf_revision = 2.0
languages = English
dataset origin = PE Nelson Turbochrom
operator_name = JDX
company_method_name = gasoline
sample id = 1
sample name = Gasoline
sample injection volume = 1.00
sample amount = 1.00
detection_method_name = gasoline
detector name =
detector_unit = uV
retention_unit = time in seconds
uniform_sampling_flag = Y
autosampler_position =
AIA_info_end
```

Currently, in Galactic's file converters for the AIA/ANDI/netCDF file formats, there is no attempt made to map the instrument or other parameters to the standard or terchnique-specific Log Text entries.

Appendix A - SPC.H Header File

```
******************
   FILENAME: spc.h
   AUTHOR:
              Steven Simonoff 11-25-90 (from spc.inc)
              5-6-97 (New X,Y types, fexper types, float ssftime, 4D data)
       Contains the Spectrum Information Header structure definitions.
       Based on zspchdr.inc from Lab Calc (tm)
       Must #include <windows.h> before this file.
   Copyright (C) 1986-1997 by Galactic Industries Corp.
   All Rights Reserved.
*************
{}^{\star} The following defines a trace header and file format.
* All floating values in SPC files are in the IEEE standard formats.
* There are two basic formats, new and old. FVERSN flags the format and
* also serves as a check byte.
                                   The new format has header values in the file
* exactly as they appear below. The old format has slightly different file
* header formatting which is translated as it is read into memory. New features
^{\star} like XY data and audit log information are not supported for the old format.
* The new format allows X,Y pairs data to be stored when the TXVALS flag is set.
* The main header is immediately followed by any array of fnpts 32-bit floating
* numbers giving the X values for all points in the file or subfiles. Note
* that for multi files, there is normally a single X values array which applies
* to all subfiles. The X values are followed by a subfile header and fixed
* point Y values array or, for multi files, by the subfiles which each consist
 of a subfile header followed by a fixed-point Y values array. Note that
* the software may be somewhat slower when using X-values type files.
* Another X,Y mode allows for separate X arrays and differing numbers of
* points for each subfile. This mode is normally used for Mass Spec Data.
* If the TXYXYS flag is set along with TXVALS, then each subfile has a
^{\star} separate X array which follows the subfile header and preceeds the Y array.
* An additional subnpts subfile header entry gives the number of X,Y values
* for the subfile (rather than the fnpts entry in the main header). Under
* this mode, there may be a directory subfile pointers whose offset is
\ensuremath{^{\star}} stored in the fnpts main header entry. This directory consists of an
* array of ssfstc structures, one for each of the subfiles. Each ssfstc
* gives the byte file offset of the begining of the subfile (that is, of
* its subfile header) and also gives the Z value (subtime) for the subfile
^{\star} and is byte size. This directory is normally saved at the end of the
* file after the last subfile.
                                    If the fnpts entry is zero, then no directory
* is present and GRAMS/32 automatically creates one (by scanning through the
* subfiles) when the file is opened. Otherwise, fnpts should be the byte
* offset into the file to the first ssfstc for the first subfile. Note
* that when the directory is present, the subfiles may not be sequentially
* stored in the file. This allows GRAMS/32 to add points to subfiles by
* moving them to the end of the file.
* Y values are represented as fixed-point signed fractions (which are similar
^{\star} to integers except that the binary point is above the most significant bit
* rather than below the least significant) scaled by a single exponent value.
* For example, 0x40000000 represents 0.25 and 0xC0000000 represents -0.25 and
* if the exponent is 2 then they represent 1 and -1 respectively. Note that
^{\star} in the old 0x4D format, the two words in a 4-byte DP Y value are reversed.
* To convert the fixed Y values to floating point:
       FloatY = (2^Exponent)*FractionY
* or: FloatY = (2^Exponent)*IntegerY/(2^32)
                                                     -if 32-bit values
* or: FloatY = (2^Exponent)*IntegerY/(2^16)
                                                     -if 16-bit values
* Optionally, the Y values on the disk may be 32-bit IEEE floating numbers.
* In this case the fexp value (or subexp value for multifile subfiles)
* must be set to 0x80 (-128 decimal). Floating Y values are automatically
* converted to the fixed format when read into memory and are somewhat slower.
* GRAMS/32 never saves traces with floating Y values but can read them.
```

```
* Thus an SPC trace file normally has these components in the following order:
                      Main header (512 bytes in new format, 224 or 256 in old)
        SPCHDR
        [X Values]
                          Optional FNPTS 32-bit floating X values if TXVALS flag
       SUBHDR Subfile Header for 1st subfile (32 bytes)
Y Values FNPTS 32 or 16 bit fixed Y fractions scaled by exponent
[SUBHDR] Optional Subfile Header for 2nd subfile if TMULTI flag
[Y Values] Optional FNPTS Y values for 2nd subfile if TMULTI flag
... Additional subfiles if TMULTI flag (up to FNSUB total)
[Log Info] Optional LOGSTC and log data if flogoff is non-zero
                          FNPTS 32 or 16 bit fixed Y fractions scaled by exponent
* However, files with the TXYXYS ftflgs flag set have these components:
        SPCHDR Main header (512 bytes in new format)
        SUBHDR
                          Subfile Header for 1st subfile (32 bytes)
       X Values
Y Values
Y Values
Y Values
FNPTS 32-bit floating X values
Y Values
FNPTS 32 or 16 bit fixed Y fractions scaled by exponent
SUBHDR |
Subfile Header for 2nd subfile
Y Values |
FNPTS 32-bit floating X values for 2nd subfile
Y Values |
FNPTS Y values for 2nd subfile
                          Additional subfiles (up to FNSUB total)
        [Directory] Optional FNSUB SSFSTC entries pointed to by FNPTS
                        Optional LOGSTC and log data if flogoff is non-zero
        [Log Info]
* Note that the fxtype, fytype, and fztype default axis labels can be
* overridden with null-terminated strings at the end of fcmnt.
* TALABS bit is set in ftflgs (or Z=ZTEXTL in old format), then the labels
* come from the fcatxt offset of the header. The X, Y, and Z labels \,
* must each be zero-byte terminated and must occure in the stated (X,Y,Z)
* order. If a label is only the terminating zero byte then the fxtype,
\mbox{*} fytype, or fztype (or Arbitrary Z) type label is used instead. The
* labels may not exceed 20 characters each and all three must fit in 30 bytes.
\star The fpost, fprocs, flevel, fsampin, ffactor, and fmethod offsets specify \star the desired post collect processing for the data. Zero values are used
* for unspecified values causing default settings to be used. See GRAMSDDE.INC
* Normally fpeakpt is zero to allow the centerburst to be automatically located.
* If flogoff is non-zero, then it is the byte offset in the SPC file to a
* block of memory reserved for logging changes and comments. The beginning
* of this block holds a logstc structure which gives the size of the
* block and the offset to the log text. The log text must be at the block's * end. The log text consists of lines, each ending with a carriage return
* and line feed. After the final line's CR and LF must come a zero character
* (which must be the first in the text). Log text requires V1.10 or later.
^{\star} The log is normally after the last subfile (or after the TXYXYS directory).
* The fwplanes allows a series of subfiles to be interpreted as a volume of
* data with ordinate Y values along three dimensions (\bar{X}, Z, W). Volume data is
* also known as 4D data since plots can have X, Z, W, and Y axes. When
* fwplanes is non-zero, then groups of subfiles are interpreted as planes
* along a W axis. The fwplanes value gives the number of planes (groups of
* subfiles) and must divide evenly into the total number of subfiles (fnsub).
'If the fwinc is non-zero, then the W axis values are evenly spaced beginning
* with subwlevel for the first subfile and incremented by fwinc after each
* group of fwplanes subfiles. If fwinc is zero, then the planes may have
* non-evenly-spaced W axis values as given by the subwlevel for the first
* subfile in the plane's group. However, the W axis values must be ordered so
* that the plane values always increase or decrease. Also all subfiles in the
plane should have the same subwlevel. Equally-spaced W planes are recommended
* and some software may not handle fwinc=0. The fwtype gives the W axis type.
```

Galactic Universal Data Format Specification 9/4/97

```
/************************
* In the old 0x4D format, fnpts is floating point rather than a DP integer,
* ffirst and flast are 32-bit floating point rather than 64-bit, and fnsub
* fmethod, and fextra do not exist. (Note that in the new formats, the
* fcmnt text may extend into the fcatxt and fextra areas if the TALABS flag
* is not set. However, any text beyond the first 130 bytes may be
* ignored in future versions if fextra is used for other purposes.)
* Also, in the old format, the date and time are stored differently.
* Note that the new format header has 512 bytes while old format headers
^{\star} have 256 bytes and in memory all headers use 288 bytes. Also, the
* new header does not include the first subfile header but the old does.
* The following constants define the offsets in the old format header:
* Finally, the old format 32-bit Y values have the two words reversed from
* the Intel least-significant-word-first order. Within each word, the
* least significant byte comes first, but the most significant word is first.
typedef struct
   {
  BYTE oftflgs;
   WORD opeakpt;
   WORD onscans;
   float ospare[7];
   char ocmnt[130];
   char ocatxt[30];
   char osubh1[32]; /* Header for first (or main) subfile included in main header */
   } OSPCHDR;
* This structure defines the subfile headers that preceed each trace in a
* multi-type file. Note that for evenly-spaced files, subtime and subnext are
* optional (and ignored) for all but the first subfile. The (subnext-subtime) * for the first subfile determines the Z spacing for all evenly-spaced subfiles.
* For ordered and random multi files, subnext is normally set to match subtime.
* However, for all types, the subindx must be correct for all subfiles.
* This header must must always be present even if there is only one subfile.
* However, if TMULTI is not set, then the subexp is ignored in favor of fexp.
* Normally, subflgs and subnois are set to zero and are used internally.
                    /* Subflgs bit if subfile changed */
/* Subflgs bit if peak table file should not be used */
#define SUBCHGD 1
#define SUBNOPT 8
\#define SUBMODF 128 /* Subflgs bit if subfile modified by arithmetic */
typedef struct
  BYTE subflgs; /* Flags as defined above */
char subexp; /* Exponent for sub-file's Y values (80h=>float) */
WORD subindx; /* Integer index number of trace subfile (0=first);
                      /* Integer index number of trace subfile (0=first) */
   float subnext;
float subnois;
                     /* Floating time for next trace (May be same as beg) */
/* Floating peak pick noise level if high byte nonzero */
                     /* Integer number of subfile points for TXYXYS type */
   DWORD subnpts;
  DWORD subscan; /* Integer number of co-added scans or 0 (for collect) */
float subwlevel; /* Floating W axis value (if fwplanes non-zero) */
   char subresv[4]; /* Reserved area (must be set to zero) */
   } SUBHDR;
#define FSNOIS fsubhl+subnois+3 /* Byte which is non-zero if subnois valid */
```

```
/* This structure defines the entries in the XY subfile directory. */
/* Its size is guaranteed to be 12 bytes long. */
typedef struct
    /^{\star} This structure defines the header at the beginning of a flogoff block. ^{\star}/
^{\prime \star} The logsizd should be large enough to hold the text and its ending zero. ^{\star \prime}
/* The logsizm is normally set to be a multiple of 4096 and must be */
^{\prime \star} greater than logsizd. It is normally set to the next larger multiple. ^{\star \prime}
/* The logdsks section is a binary block which is not read into memory. */
typedef struct
                              /* log block header format */
                             /* byte size of disk block */
    DWORD logsizd;
    DWORD logsizm;
                              /* byte size of memory block */
                             /* byte offset to text */
    DWORD logtxto;
                              /* byte size of binary area (immediately after logstc) */
    DWORD logbins;
    DWORD logdsks; /* byte size of disk area (immediately after logbins) */
char logspar[44]; /* reserved (must be zero) */
    } LOGSTC;
/* Possible settings for fxtype, fztype, fwtype. */
^{\prime\star} XEV and XDIODE - XMETERS are new and not supported by all software. ^{\star\prime}
/* Arbitrary */
#define XARB 0
#define XEV 14  /* eV */
#define ZTEXTL 15  /* XYZ text labels in fcatxt (old 0x4D version only) */
#define XDIODE 16  /* Diode Number */
#define XCHANL 17  /* Channel */
#define XDEGRS 18  /* Degrees */
#define XDEGRS 19  /* Temperature (F) */
#define XDEGRC 20  /* Temperature (C) */
#define XDEGRC 21  /* Temperature (K) */
#define XPOINT 22  /* Data Points */
#define XMSEC 23  /* Milliseconds (mSec) */
#define XUSEC 24  /* Microseconds (uSec) */
#define XNSEC 25  /* Nanoseconds (nSec) */
#define XGHERTZ 26  /* Gigahertz (GHz) */
#define XGHERTZ 26 /* Gigahertz (GHz) */
#define XCM 27 /* Centimeters (cm) */
#define XMETERS 28 /* Meters (m) */
#define XMOURS 30 /* Hours */
#define XDBLIGM 255 /* Double interferogram (no display labels) */
```

```
/* Possible settings for fytype. (The first 127 have positive peaks.) */
/* YINTENS - YDEGRK and YEMISN are new and not supported by all software. */
/* Kubelka-Monk */
#define YKMONK 3
#define YCOUNT 4 /* Counts */
#define YVOLTS 5 /* Volts */
#define YDEGRS 6 /* Degrees */
#define YAMPS 7 /* Milliamps */
#define YMVOLTS 9 /* Millivolts */
#define YLOGDR 10 /* Log(1/R) */
#define YLOGDR 11 /* Percent */
#define YPERCNT 11
                        /* Percent */
                         /* Intensity */
#define YINTENS 12
                        /* Relative Intensity */
#define YRELINT 13
                        /* Energy */
#define YENERGY 14
                        /* Decibel */
#define YDECBL 16
#define YDEGRF 19 /* Temperature (F) */
#define YDEGRC 20 /* Temperature (C) */
#define YDEGRK 21 /* Temperature (K) */
#define YINDRF 22 /* Index of Refraction [N] */
#define YEXTCF 23 /* Extinction Coeff. [K] */
#define YEXTCF 23 /* Page 1 */
#define YEAL 24 /* Real */
#define YIMAG 25 /* Imaginary */
#define YCMPLX 26 /* Complex */
                        /* Transmission (ALL HIGHER MUST HAVE VALLEYS!) */
#define YTRANS 128
#define YREFLEC 129 /* Reflectance */
#define YVALLEY 130 /* Arbitrary or Single Beam with Valley Peaks */
                         /* Emission */
#define YEMISN 131
/* Possible bit FTFLGS flag byte settings. */
/* Note that TRANDM and TORDRD are mutually exclusive. */
/* Code depends on TXVALS being the sign bit. TXYXYS must be 0 if TXVALS=0. */
/* In old software without the fexper code, TCGRAM specifies a chromatogram. */
/* FMODS spectral modifications flag setting conventions: */
/* "A" (2^01) = Averaging (from multiple source traces) */
    "B" (2^02) = Baseline correction or offset functions */
    "C" (2^03) = Interferogram to spectrum Computation */
/* "D" (2^04) = Derivative (or integrate) functions */
    "E" (2^06) = Resolution Enhancement functions (such as deconvolution) */
/* "I" (2^09) = Interpolation functions */
    "N" (2^14) = Noise reduction smoothing */
    "0" (2^15) = Other functions (add, subtract, noise, etc.) */
/* "S" (2^19) = Spectral Subtraction */
    "T" (2^20) = Truncation (only a portion of original X axis remains) */
/* "W" (2^23) = When collected (date and time information) has been modified */
/* "X" (2^24) = X units conversions or X shifting */
    "Y" (2^25) = Y units conversions (transmission->absorbance, etc.) */
/* "Z" (2^26) = Zap functions (features removed or modified) */
```

Galactic Universal Data Format Specification 9/4/97

Appendix B - GCDLL.H Header File

MRU: 10-20-96 dea add GCI_SHWLST & GCC_SHWLST flags

DESC: Defines for Galactic file converters written in 'C'

Compile with FLAT32=1 for 32-bit version This file should be included after Windows typedefs in windows.h and windef.h as it uses standard Windows data types.

IMPORTANT! Changes to this header file must be reflected in the assembly include file "GCDLL16.INC"

OVERVIEW:

Galactic DLL File Converters are standard 16-bit Windows DLL's renamed with a *.CN4 extension (32-bit converters are renamed to a *.CN5 extension). Currently only 16-bit converters are supported by GRAMSCNV.EXE but all converters should be written so that they can be compiled to a 32-bit DLL for future Galactic products. 16-bit converters should be given an 8.3 DOS filename where the base filename (8 characters) describes the instrument manufacturer and the model (as in "NIC_OMNC"), the extension will always be ".CN4" 32-bit converters should use Win32 long filenames of up to 32 characters containing the same information. The base filename is used as the converter name in the converter selection listbox.

Each file converter must support the following EXPORT'ed calls:

```
BOOL EXPORT GCGetInfo (lpGCGETINFOS); // ordinal @2
BOOL EXPORT GCCheckData (lpGCCHECKDATAS); // ordinal @3
int EXPORT GCConvert (lpGCCONVERTS); // ordinal @4
```

All API calls take a single pointer to a structure (defined below) to pass information between the caller and the converter. The first call (GCGetInfo) returns information about the converter including its version, extension(s) for source files, prompt stings and functions supported to the caller. The GCCheckData() call allows the caller to see if the converter can handle a particular piece of data. This is useful when the caller does not know the format of the data and needs to find a converter which understands it. The final call imports or exports foreign data to or from the Galactic data format. Conversions are done from disk file to disk file or the Galactic data file may be supplied (on export) or returned (on import) in a block of global memory. A converter may also optionally support importing a foreign file from a memory file and/or exporting a foreign file to a memory file. The converter is responsible for all file I/O including opening/creating the source/destination files and cleaning up if an error occurs. However, in general a file converter should not have to do any user I/O, all user input is normally handled by the caller. The format of the Galactic data file is defined in SPC.H (only the new Galactic format is supported).

Note that a caller to the converter does not have to call these API's in a particular order so a converter should not, for example, require GCGetInfo() to be called before GCConvert() in order for the converter to function.

```
with a math coprocessor available.
               All strings are NULL terminated byte character text strings
               unless otherwise noted.
               Note: 16-bit DLL converters written to this specification
               (version 1) do not have to support long filenames. However
               32-bit converter DLL's must support long filenames.
   Copyright (C) 1996 Galactic Industries Corp. All Rights Reserved.
#ifndef FLAT32
                                      // 16-bit version
#define EXPORT
#define DllExport
#define SPECLEVEL 1 // level 1 of the DLLL converter specification
                              // 32-bit version
#else
#define EXPORT
#define DllExport __declspec( dllexport )
#define SPECLEVEL 1001 // +1000 flags 32 bit converter
#endif
#define MAXSTRING 80 // max string length for strings in GCGETINFOS #define MAXVERTXT 16 // max string length for version text
// ***********************************
// Structure passed by GCGetInfo() call to converter DLL. The first
// item (size) is filled in by caller, all others filled in by converter.
// This function provides information to the caller about the converter
// type and version. The return value is normally TRUE. It should
// return FALSE it there is an error initializing the converter and it
// is not available for use (insufficient memory for example).
\ensuremath{//} The caller may not modify any of returned strings (so they may be declared
// LPCSTR in the converter). All returned strings are currently 8-bit
// character type (UNICODE is not supported).
//
      GCInfoFlags defines
               GCI_EXPORTOK 0x0001L // set if export supported by converter
#define
#define
              GCI_IFILESEL 0x0002L // set if converter handles
import fsel
             GCI_MEMIMPOK 0x0004L // memory source import supported
GCI_MEMEXPOK 0x0008L // memory dest export supported
GCI_SRCDIR 0x0010L // foreign filename is a directory
GCI_SHWLST 0x0020L // converter supports GCC_SHWLST flag
#define GCI_MEMIMPOK 0x0004L
#define
#define
#define
typedef struct
               DWORD GCInfoSSize; // length, in bytes, of the structure
               DWORD GCSpecLevel; // must return 1 (16bit) or 1001 (32bit)
               DWORD GCInfoFlags; // returns converter function flags LPSTR GCVersion; // retrns ptr to cnv version string
               LPSTR GCDescription; // retrns ptr to cnv description string
               LPSTR GCImpExt; // retrns ptr to extension(s) for import LPSTR GCImpPrompt; // retrns ptr to title for imp file sel
               LPSTR GCExpPrompt; // retrns ptr to title for exp file sel LPSTR GCRegFormat; // retrns ptr to clipboard/DDE fmt strng
               } GCGETINFOS, FAR *lpGCGETINFOS;
```

Converters can expect to run on an Intel 80386 or better

```
// All LPSTR variables point to strings no longer than MAXSTRING (including
// terminating NULL) except GCVersion which is no longer than MAXVERTXT in
// length.
// GCInfoSSize is the number of bytes in the structure. It is filled in
// by the caller before calling this function and may be used by the converter
// to check the specification level supported by the caller.
// GCSpecLevel is the DLL converter specification supported. Currently, only
// level 1 of the specification exists (this spec) and this value should
// always return 1 (this spec) for 16-bit converters or 1 + 1000 for
// 32-bit converters.
// GCInfoFlags returns a set of OR'ed flags indicating the options
// supported by the converter. Currently only the GCI_EXPORTOK,
// GCI_IFILESEL, GCI_IMPMEMOK, GCI_EXPMEMOK and GCI_SRCDIR flags
// are defined (all other bits must be set to 0).
// The GCI_EXPORTOK flag should be set by the converter to indicate
// that it supports export from Galactic format to the native format. If
// this flag is not set GRAMSCNV will disable its "Export" button.
// If the GCI_IFILESEL flag is set, the calling program will not display
// a file selector when the user tries to import a file. Instead, the
// calling program will call the converter exported function GCFileSel()
// defined below. The converter is then responsible for selecting the
\ensuremath{//} foreign file to import. MOST CONVERTERS SHOULD NOT SET THIS FLAG!
// The only converters which may need to set this flag are those in which
// the foreign file is encapsulated in a larger database and there is no
// way for a file based selector to choose a file. Bypassing the
// caller's import file selector may make a converter inoperable with
// some Galactic programs and may cause the converter to fail when
// called directly via the command line or DDE. In addition, the normal // GRAMSCNV file selector supports multiple file import while the
// GCFileSel() call only supports a single file at a time.
// All converters must support import to a Galactic data file in memory. If
// the converter also supports import from a foreign file in memory, it
// should return the GCE_MEMIMPOK flag set. All converters should
// try to support import from memory if the source format can be // represented with a single file. The sample converter includes routines
// which make the import file form transparent to the caller.
// All converters must support export from a Galactic data file in memory. If
\ensuremath{//} the converter also supports export to a foreign file in memory, it should
// return the GCE_MEMEXPOK flag set. All converters should try to support // export to a memory file if the destination format can be represented
// with a single file. The sample converter includes routines which
// make the export file form transparent to the caller. If a converter
// that normally creates additional files on import (e.g. \star.CFL, \star.CGM) is
// called to convert a file with the destination Galactic file in memory
// none of the additional files are created. Similarly, if a converter
// normally exports files in addition to the main foreign data file,
// none of the additional files are created if the destination file is
// a memory file. If the converter returns the GCI_SRCDIR flag set it
// indicates that the converter expects a directory name as the source
// filename on import and as the destination filename on export. Note that
// the destination directory may not exist on file export, the converter is
// responsible for creating the directory if it does not already exist.
// Multiple file conversion when the GCI_SRCDIR is not supported by
// the current GRAMSCNV.EXE
// The GCI_SHWLST flag is set by converters with file formats which have
// multiple data files in a single disk file but normally do not require an
// internal file selector (i.e. do not have the GCI_IFILESEL flag set).
// These formats convert a default internal data file, but if the GCC_SHWLST
// flag is set when GCConvert is called on import or export the converter
// presents a list of data files available in the input file (or destination // file) and allows the user to choose which file to convert.
```

```
// GCVersion points to a text string containing the converter version
// This string must end in a "D4" to indicate a 16-bit DLL converter
// (i.e. *.CN4) or end in "D5" for a 32-bit converter (i.e. *.CN5)
// For example "1.03D4\0". The maximum version string length is MAXVERTXT.
// The converter version must be incremented any time the converter source
// code is modified resulting in a new executable (*.CN4 or *.CN5) file.
// In addition the version information in the converter's version resource
// should also track this version number. The sample converters place the
// version information at the top of the <cvnname>.H file to set the converter
// version in both places making it easier to update the file. We recommend
// incrementing the version number before editing any source files in a
// converter to be modified.
// GCDescription points to a text string which describes the converter, i.e. // "ASCII File with Header" \, This string is displayed next to the
// base file converter name in the initial GRAMSCNV dialog. The string
// normally should not exceed 60 characters (less if mostly uppercase letters)
// in order to fit the dialog box and is limited to MAXSTRING in all cases.
// GCImpExt points to a NULL terminated text string used by the file import
// file selector for the source file extension(s). Multiple extensions may be
// specified. This string has the '|' characters (ASCII 0x7c) replaced with
// NULL characters and the result is passed directly to the Windows Common
\ensuremath{//} Dialog file open dialog routine. This string contains one or more pairs of
// '|' separated strings specifying filters. The first string in each pair // describes a filter (for example, "ASCII Files (*.asp)"); the second // specifies the filter pattern (for example, "*.asp"). Multiple filters can
// be specified for a single item; in this case, the semicolon (;) is used to
// separate filter pattern strings - for example, "*.txt;*.doc;*.bak". The
// string in the buffer must be terminated by a null character. GRAMSCNV adds // "All Files (*.*) | *.*" to the string returned by GCGetInfo(). If the last
// character in the GCImpExt[] string is a ('|'), then the "All Files
// (*.*)" entry will not be added to the file type dropdown combobox.
// A converter should not supress the "All Files" addition without a very good
// reason. The filter strings must be in the proper order, the system does
// not change the order. Extensions should be in lowercase. For example:
//
       "Data Files (*.dat)|*.dat"
//
       "Text Files | *.txt; *.prn; *.cap"
//
       "Data Files (*.dat) | *.dat | Capture Files (*.cap) | *.cap"
//
//
       "Data Files Only | *.dat | "
//
// The GCImpExt[] string is limited to a maximum of MAXSTRING characters. The
// first extension is always used as the default on import if no extension is
// typed and is the default for export if no extension is typed. Note that
// the '|' character is NOT allowed in the string except as the separator.
// Using the '|' in place of NULL separated strings allows the extension
// specifying string to be placed as a single entry in the converter's
// stringtable resource. If the GCI_SRCDIR flag is set, the extension string
// is ignored and the return pointer should be set to NULL.
// GCImpPrompt points to a text string used as the prompt for the source file
// selector on import. This string normally should not exceed 60 characters
// (less if mostly uppercase letters) in order to fit the dialog box and is
// limited to MAXSTRING in all cases. It should of the form "Convert from
// ASCII File with Header to Galactic File"
// GCExpPrompt points to a text string used as the prompt for the source file
// selector on export. It should normally be no longer than 60 characters
// (and is always limited to MAXSTRING characters) and usually
// is of the form "Convert from Galactic File to ASCII File with Header"
// GCRegFormat points to a string specifying the registered clipboard
format(s)
// supported by the converter. Multiple clipboard formats may be specified
// separated by semicolons ';'. For example: "CF_OLDINST;CF_NEWINST"
```

```
// If no clipboard or DDE registered formats are supported this pointer
// should be set to NULL.
// Structure passed by GCCheckData() call to converter DLL. All items in
// this structure are filled in by the caller.
// This function queries the converter to see if it can import a particular
// data file. The converter returns TRUE if it handles this format or // FALSE if it does not. The handle of the open file, the name of the
// open file and a pointer to the first 256 bytes in the file are normally
// passed to this function. If the caller has the source file only in memory,
// then the file handle (GChFile) is NULL and GCChkBytes is set to the total
// bytes in the memory file pointed to by GCFileData (which may be much
// larger than 256 bytes - converters should always do a DWORD compare
// against the number of pointed to by GCFileData). If the data is
// passed solely in memory the filename pointer (GCFileName) will normally
// be NULL. This function should return as quickly as possible if it
// does not handle the data. If there are less than 256 bytes in the
// data file - unused bytes should be set to 0.
typedef struct
            DWORD GCChkDSize; // length, in bytes, of the structure
            DWORD GCChkBytes; // actual number of bytes in GCFileData
            LPSTR GCFileName; // ptr to name of open file GChFile
            LPVOID GCFileData; // ptr to first GCChkBytes bytes in file
#ifndef FLAT32
            HFILE GChFile;
                               // handle of file to check or NULL
#else
            HANDLE GChFile; // handle of file to check or NULL
#endif
            } GCCHECKDATAS, FAR *lpGCCHECKDATAS;
// GCChkDSize is the number of bytes in the structure. It is filled in
// by the caller before calling this function.
// GCChkBytes is the actual number of bytes in the buffer pointed to by
// GCFileData. This a number between 0 and 256 if GChFile is not NULL,
// otherwise it is the total number of bytes in the memory file. Some
// converters may use this value to quickly qualify or reject a file.
// GCFileName points to the name of the file being tested. This may be NULL.
// The file name is provided for those cases where the converter needs to
// open or check for the presence of an auxiliary file to test the format.
// GCFileData points to a buffer holding the first GCChkBytes bytes in
// the data file. If GChFile is not NULL this will not exceed 256 bytes.
// GChFile is the handle of an open data file (HFILE type opened with the
// OpenFile() function call if 16-bit or HANDLE type opened with CreateFile()
// function call if 32-bit) or NULL if the entire data file is pointed to by
// GCFileData.
// The converter MUST NOT CLOSE the data file GChFile, but it does not have
// to reset the file pointer to the beginning of the file if it reads from
// the file or moves the file pointer. The file pointer is always reset to
// the beginning of the file before this function is called.
// ********************
// Structure passed by GCConvert() call to converter DLL.
// All items are filled in by caller (GCErrorMsg and GCDestMemFile are
// set to NULL by the caller and GCDestMemFSize is set to OL.
// GCSrcMemFile and GCSrcMemFSize are set to NULL and 0L respectively
// by the caller if not used). The return value is 0 if no error occurs
// or an error code (see GCDLLERR.H) if the conversion fails.
      GCCvtFlags defines
```

```
// set if export operation
// set to supress user interaction
// show multiple file choice if any
                             0 \times 0001 L
#define
             GCC_EXPORT
                             0 \times 0002 L
#define
              GCC_NOUSER
              GCC_SHWLST
#define
                             0 \times 0004 L
// If the GCC EXPORT flag is set on entry, the file pointed to GCSrcFileNm
// (or in the memory file GCSrcMemFile) is a Galactic data file which
// should be exported to the converter's native format. Note that the
// converter should handle this case (with an error message) even if the
// converter did not return the GCI_EXPORTOK flag on the call to GCGetInfo()
// since some callers may call this API without first calling GCGetInfo().
// If the GCC_NOUSER flag is set by the caller then the converter should
// convert the file with no user interaction (however it should still return // and error code and message if applicable). This flag is normally ignored
// since most file conversions do not require any interaction in order to
// convert a file. However some converters give the user conversion options
// or warning messages and these should not be displayed if the GCC_NOUSER
// flag is set. If the converter cannot continue without user interaction it
// should return the GCE_NOUSER error code. Note that a converter can still
// call GCCallback with messages even if this flag is set as it is up to the
// calling program whether to display them or not.
\ensuremath{//} The GCC_SHWLST flag is used only by file formats which have multiple
// data files in a single disk file but normally do not require an internal
// file selector (i.e. do not have the GCI_IFILESEL flag set). These formats
// convert a default internal data file, but if the GCC_SHWLST flag is set
// on import or export the converter presents a list of data files available
// in the input file (or destination file) and allows the user to choose
// which file to convert. This flag is ignored if the file format does not
// contain more than one data file or if the GCE_NOUSER flag is set.
typedef struct
       DWORD GCCvnSSize;
                                  // length, in bytes, of the structure
      DWORD GCCvtFlags;
LPSTR GCSrcFileNm;
                                 // GCC_EXPORT and/or GCC_NOUSER
      LPSTR GCSrcFileNm; // ptr to source filename
HGLOBALGCSrcMemFile; // handle for source memory file
DWORD GCSrcMemFSize; // bytes in source memory file
       LPSTR GCDestFileNm;
                                 // ptr to destination filename
                                  // handle for destination memory file
       HGLOBAL GCDestMemFile;
       DWORD GCDestMemFSize;
                                  // bytes in destination memory file
      FARPROC GCCallback; // caller timesharing callback or NULL LPSTR GCErrorMsg; // returns optional ptr to err msg text HWND GChParent; // handle of parent window
       FARPROC GCCallback;
} GCCONVERTS, FAR *lpGCCONVERTS;
// GCCnvSSize is the number of bytes in the structure. It is filled in
// by the caller before calling this function.
// GCCvtFlags contains the action flags defined above, all other bits
// will be set to 0.
// GCSrcFileNm is a pointer to the name of the file to open (or NULL if
// the source is a memory file). It is a maximum of MAX_PATH bytes long.
// The converter cannot modify this buffer, it must copy the filename to
// a local buffer if it needs to modify the string. The file is not open
// when the call to GCImport() or GCExport() is made. The converter is
// responsible for opening (in read only mode) and closing the file.
// Normally, when called by GRAMSCNV.EXE this string will contain a filename
// and extension but no path. GRAMSCNV.EXE changes the current directory
// to the directory containing the source file before calling the GCConvert()
// function. However, other users of the converter may choose to pass a full
// pathname instead. The converter must handle both cases. The converter
// does not have to restore the original directory if it changes the current
// directory. If this pointer is NULL then the source file is avaiable in
// the block of global memory whose handle is passed in GCSrcMemFile.
```

```
// GCSrcMemFile is the handle of a block of global memory containing
// the source data file if GCSrcFileNm is NULL. The converter must not
// free this memory block but it must unlock it before returning to the
// caller. If the converter does not support import from a memory file
// it should return GCE_BADOP.
// If a source memory file is supplied, GCSrcMemFSize is set to the total
// number of bytes in the memory file (not the amount of memory allocated
// which may be greater). This allows the user of the memory file to ensure // that they do not try to access memory beyond the end of the file if the
// file format contains errors.
// GCDestFileNm is a pointer to the name of the file to create (or NULL if
// the destination is a memory file). It is a maximum of MAX_PATH bytes
// long. The converter cannot modify this buffer, it must copy the filename
// to a local buffer if it needs to modify the string. The file is not open
// when the call to GCImport() or GCExport() is made. The converter is
// responsible for creating this file. The converter does NOT need to check
// to see if the file already exists, the caller is responsible for this.
// The converter should delete any partial destination file(s) if an error
// occurs. This string may contain a path or may contain just a filename and
// extension. The converter must handle both cases. In particular, make sure
// the converter creates the destination file before changing the current
// directory if the destination file does not include a path. If GCDestFileNm
^{\prime}/ is NULL the converter should create the destination data file as a memory
// file.
// If GCDestFileNm is NULL the converter should return the destination
// file in a block of global memory allocated with the GHND flag. On return,
// the memory handle should be stored in GCDestMemFile and GCE_OK returned.
// The memory handle returned must be unlocked by the converter before
// returning. The caller is responsible for FREE'ing the handle returned.
// If an error occurs (and the converter returns any value other than GCE_OK)
// the converter is responsible for freeing any memory allocated and should
// return a NULL value in GCDestMemFile. If insufficient memory exists to
// return the data file in memory, the converter should return GCE_MEMORY.
// The caller can then choose to call the function again with a temporary
// disk file as the destination name. If the converter does not support
// export to a memory file it should return GCE_BADOP.
// If a destination memory file is returned, GCDestMemFSize is set to the
// total number of bytes in the memory file (not the amount of memory
// allocated which may be greater). This allows the user of the memory
// file to ensure that they do not try to access memory beyond the end of
// the file if the file format contains errors.
// If a converter that normally creates additional files on import
// (e.g. *.CFL, *.CGM) is called to import a file to a Galactic memory
// file none of the additional files are created. Similarly, if a converter
// normally exports suplemental files in addition to the main foreign data
// file, none of the additional files are created if the destination file is
// a memory file. The destination file must be a named disk file
// GCCallback is a pointer to a function which allows Windows non-
// preemptive timesharing in while the converter is in a tight conversion
// loop and allows the converter to set status messages in the caller's
// dialog window. It also checks for abort by the user. The converter
// must call this function often (no less than twice per second) while
// converting to allow Windows multitasking to occur and to check for
// user abort. The return value is a BOOL. TRUE to abort the conversion,
// FALSE to continue. If this function returns TRUE, the converter should
// end its conversion and return (GCE_CANCEL) to the caller as soon as
possible.
// Note that as long as the converter calls this function often, it does not
// have to call PeekMessage() to allow Windows multitasking.
// The callback function is declared as follows:
```

```
//
//
      BOOL CALLBACK GRCheckAbort (LPSTR szStatus)
//
// The LPSTR parameter should be set to NULL if no status message needs
// to be displayed. IMPORTANT! If the caller does not support the
// callback function it will be set to NULL. The converter should not
// call a NULL callback function!
// GCErrorMsg is set to NULL by the caller. If GCConvert() returns GCE_OK // or GCE_CANCEL, this value is ignored and the caller does not put up any
// error dialog. If any other error code is returned, the caller displays
// the error in an error dialog. If GCErrorMsg is returned NULL, then the
// caller uses its own message string (see GRDLLERR.H for standard error
// strings) based on the error code returned. If GCErrorMsg is not NULL,
// the string it points to is used for the error message. Note that you
// can OR GCE_WARN or GCE_INFO to the error code to display a Warning (!)
// or Information (I) dialog respectively rather than an Error (Stop Sign)
// dialog. Also, you can add the source or destination filename to the end
// of the message string (either the msg returned by the converter or the
// internal error messsage) by OR'ing the GCE_SRC or GCE_DEST flags to the
// returned error code, however the default error dialog already displays
// the source file name in the dialog title. If the GCE_WARN or the GCE_INFO
// flags are not set, the error dialog is preceded by a Galactic error beep.
// If a converter returns a custom error message in GCErrorMsg for a condition
// which is not defined in GCDLLERR.H, the function should return GCE_UNKNOWN.
// GChParent is the handle of the parent window. This window may be hidden
// if the conversion is taking place in the background or it may be iconic.
// The converter should not popup a dialog or window if the parent window
// is not the active application. It should wait until the parent window
// becomes active to put up an error. Normally, converters do not need // to do any user interaction and may ignore this value
// ********************
          The following call is optional and should
//
            not be supported by most file converters
//
// Structure passed by optional GCFileSel() call to converter DLL.
// All items except GCFSelBuffer are filled in by caller. The return
// value is count of filenames in buffer, 0 if import should be canceled,
// or -1 if an error occurs.
//
      GCFileSel defines
          GCF_LONGFN 0x0001L // set if long filenames allowed GCF_MULTIOK 0x0002L // set if multiple null sep names ok
#define
#define
typedef struct
      DWORD GCFSelSSize; // length, in bytes, of the structure
      DWORD GCFSelflags; // GCF_LONGFN and/or GCF_MULTIOK
LPSTR GCFSelBuffer; // ptr to returned filename(s)
HWND GCFhParent; // handle of parent window

GCFILESELS, FAR *lpGCFSELS;
// GCFSelSSize is the number of bytes in the structure. It is filled in
// by the caller before calling this function.
// GCFSelFlags are set by the caller to indicate what options are
// supported by the converter. If the GCF_LONGFN flag is not set, the
// filename returned in GCFSelBuffer must conform to the DOS 8.3 filenaming
// conventions. The filenames may include a full path or a filename only.
// If no path is returned, the current directory is assumed. The GCFileSel()
// call may change the current directory.
// If the GCF_MULTIOK flag is set, the converter may return multiple NULL
// separated filenames in the GCFSelBuffer. The value returned by the
```

```
// function is the actual number of filenames in the buffer. The maximum
// number of filenames returned must be less than 32768 (in the 16-bit
// version of the converter) since the return value must be positive.
// Currently, the 2.00 version of GRAMSCNV does not set either flag (so the
// return value must be -1, 0 or 1). All undefined bits in GCFSelFlags
// are set to 0 by the caller.
// The converter must point GCFSelBuffer to a buffer containing the file(s)
// selected separated by NULL characters. Each filename returned is copied
// to the GCSrcFileNm buffer in the GCCONVERTS structure before GCConvert()
// is called. The caller does no parsing or checking of the filename(s)
// returned. GRAMSCNV does NOT check for the validity of the filename
// returned but merely passes it back to GCConvert(). The maximum length of
// a single filename is MAX_PATH. Note that if the converter allocates // memory for the filename buffer, it is responsible for freeing it when it // unloads.
// GCFhParent is the window handle of the parent window which may be needed
// if the converter needs to put up a dialog window or file selector.
// ************
// Required functions exported by converter
// *************
DllExport BOOL WINAPI EXPORT GCGetInfo (lpGCGETINFOS);
                                                       // ordinal @2
DllExport BOOL WINAPI EXPORT GCCheckData (lpGCCHECKDATAS); // ordinal @3
DllExport UINT WINAPI EXPORT GCConvert (lpGCCONVERTS); // ordinal @4
// ************
// The following function is only reguired if
// the GCI_IFILESEL flag is set on the call to
// GCGetInfo
// *************
DllExport int WINAPI EXPORT GCFileSel (lpGCFSELS); // ordinal @5
```

Appendix C - GRAMSDDE.H Header File

```
/**********************************
                  gramsdde.h
   FILENAME:
   AUTHOR: David Abrams 11-22-91 (from gramsdde.inc)
            10-25-92 Add infostc, ABControl posting choices, xfrflag=XFABC.
            Collect DDE Transfer Protocol Definitions.
            Must #include <windows.h> before this file.
   Copyright (C) 1991-1992 Galactic Industries Corp. All Rights Reserved.
*********************
* This structure comes at the beginning of a GRAMS-file-format DDE data
block.
  The image of a GRAMS file (such as an .SPC format file) immediately
follows.
* If the first xfrchk four bytes of a DDE transfer block are not: 254,24,6,18
  (=0xFE,^X,^F,^R or 0x120618FE) then the data block is not a GRAMS format
* file. This means that CF_TEXT data can double as a GRAMS file image so
that
* servers which do not support this GRAMS format automatically send text
data.
  If the check words match, then the XFRBEG area is 512 bytes long and is
  immediately followed by an exact image of a GRAMS disk file. Note that
  it is important that the xfrname contain the filename extension (such as
  .SPC, .PKS, .GDB, etc) so that the type of GRAMS file is known.
  extension is not case sensitive. The xfrmtd list is a series of file path
  names separated by spaces, tabs, or commas and ending with a final zero.
  The filenames may optionally include disk and directory and should always
  include the extension.
  The xfrchan and xfrexpr settings are needed only for unsolicited (POKE)
data
 and allow GRAMS to determine the channel to be used to process the data.
  If xfrchan is non-zero then it determines the channel. Otherwise, the
  channel with a matching experiment code is used. Otherwise, the experiment
  implied by the xfrname extension is used to determine the channel or
channel
 1 is used.
* Note that many settings be "unspecified" (0) so that GRAMS' default
settings
* are used instead of the supplied settings. Also some settings are
duplicated
* in the SPC header and can be supplied there rather than in this block.
^{\star} Most 0 values for parameters are "unspecified" so GRAMS uses its own
  defaults. The xfrflag byte cannot be defaulted an must have a valid
* (usually 0).
* Note that xfrpost is equivalent to fpost in the SPC header.
equivalent:
  xfrproc=fprocs, xfrlevl=flevel, xfrsamp=fsampin xfrfac=ffactor,
  xfrmtd=fmethod   The method list can have comma-separated filenames with
* extensions.
                 However, only the *.MTD method name is placed back in the
* fmethod for processing.
*******************
/* Four check bytes at beginning of data block to flag disk image. */
#define VXFRCHK0 254
#define VXFRCHK1 24
#define VXFRCHK2 6
```

#define VXFRCHK3 18

```
typedef struct
      BYTE xfrchk[4]; /* Must be 120618FE Hex (0xFE, Ctrl-X, Ctrl-F, Ctrl-R) */
      DWORD xfrsize; /* Data size including 512 for header (0=use actual
size)*/
      BYTE xfrflag; /* Flags (must always be specified, see below) */
      BYTE xfrchan; /* Channel to use (1-4 or \overset{-}{0} if unspecified) */
      WORD xfrfree; /* Unused (may be used by collect server) */
      WORD xfrform; /* Clipboard format (not currently used) */
BYTE xfrexpr; /* Experiment code to use plus one (0 if unspecified) */
      BYTE xfrexpr; /* Experiment code to aborate the strength of th
      BYTE xfrproc; /* Processing type (see below) */
      BYTE xfrlevl; /* Calibration level plus one (1 = not calibration data)
      WORD xfrsamp;
                                 /* Sub-method injection sample number (1=first) */
      float xfrfac;
                                 /* Floating-point concentration factor (IEEE 32-bit) */
      char xfrmtd[48*3]; /* Method/program/reference/library/archive fnm list
                               /* Mode argument when xfrflag=XFABC */
      BYTE xfrabcm;
      BYTE xfrabcn; /* Database number for channel when xfrflag=XFABC */
WORD xfrabck; /* Database type key when xfrflag=XFABC */
WORD xfrabcf; /* Database dbflags (LSB) & dbftflg (MSB) creation
                                         /* Reserved (must be zero) */
      BYTE xfrresv[82];
      char xfrname[256]; /* File path name (with optional disk and directory)
      } XFRBEG;
/******************************
* These are the xfrflag flag definitions.
* When XFMORE is set, then GRAMS assumes that the run will continue to produce
* more files in the future and requests another DDE file from the server.
* When XFDISK is set, then GRAMS reads the file from disk to do any
* When XFSUBF is set, GRAMS keeps multifile open and adds subfiles to end.
* The XFHOTLK allows server to switch to hot link protocol at run time.
* Whether or not XFNOERROR is set, GRAMS returns any processing error code
* in the DDE_FAPPSTATUS bits of wStatus (lpdwResult from DdeClientTranaction).
* A zero code indicates no error. See ERROR.INC for a list of code numbers.
* The XFHOMED flag allows Array Basic programs and other processing files
* normally located in GRAMS' "home" directory to be located instead in the
* data file's directory without having to specify complete path names for
*************************
run */
#define XFDISK
                             2 /* Data file is on the disk (and nothing follows
xfrbeg) */
#define XFSUBF
                             4 /* Data is a subfile in a series from a multifile */
#define XFNSHOW 8 /* Do not show collect data in GRAMS even if foreground */
#define XFHOTLK 16 /* GRAMS should use the hot-link protocol for rest of run
#define XFNOERR 32 /* GRAMS should skip showing error messages for data */
#define XFHOMED 64 /* Sets "home" directory to match data's during
processing*/
#define XFABC
                               128 /* Alternate ABC-style processing and posting via
@colPost*/
/***************************
* These are the xfrproc post processing type definitions.
* Beware that PPDEFLT (0) is reserved for "unspecified". Otherwise, PPKNOWN=1
* MUST be added into the sum of the bit settings. For example, to specify
* compute (PPCOMP) and search (PPSRCH) be used, the xfrproc values must be
* set to: PPCOMP+PPSRCH+PPKNOWN = 66. For no processing: PPNONE+PPKNOWN = 1.
```

```
* Most of the bit settings may be added together to specify multiple actions
* as long as 1 is added to the result. However, PPTRANS, PPABSRB, and PPKMUNK
* are mutually exclusive. Only one of the three may be added into the sum.
* Also, "?" within filename denotes the experiment's beginning letter:
* "S"=SPC (General), "C"=CGM, "G"=GC, "L"=LC, "H"=HC (HPLC), "F"=FIR (FTIR),
* "I"=IR (NIR), "V"=VIS, "U"=UV (UV/VIS), "X"=XRY (X-Ray), "M"=MS (Mass
  "N"=NMR, "R"=RMN (Raman), "E"=EFL (Excitation Fluorescence), "A"=AAS (Atomic), "D"=DAS (Diode Array), "O"=OTR (Other).
* Although the experiment letters are limited to "ACDEFGHILMNORSXUV",
* they need not be used for the types of experiments implied above.
* The code number for the first "S" experiment is 0 and for the "O" last is
* Array Basic .ABP or .DLL program extensions are used (rather than .AB or
* so that the special post processing programs do not show in the DoProg menu.
* During any post processing, the file is temporarily the current #S database.
* Processing is done in the order listed below. In some systems, the PPCOMP
* progam may also do the background ratio and output transmission, absorbance,
* or k-munk spectra. If the xfrmtd list has a SPC-type spectrum filename,
* then it its assumed to the desired background spectrum. (Otherwise, the
* spectrum with the same experiment type in GRAMS' background list is used.)
* If PPPEAK is specified, the a *.MTD method filename from the xfrmtd list
* is used as the peak picking method. (If none is present, then the default
* method chosen in GRAMS's Peak menu is used.) If PPSRCH is specified than
* a LIB-type library database filename in the xfrmtd list is searched. (If
* no library is listed, then the library or libraries from GRAMS's library
; with matching experiments are used.) Finally, any *.AB* or *.DLL Array Basic
; program filenames listed in xfrmtd are executed in order.
#define PPDEFLT 0 /* Use default settings (xfrproc is "unspecified") */
#define PPKNOWN 1 /* MUST BE ADDED TO THE SUM OF SETTINGS INCLUDED BELOW!
#define PPNONE 0 /* No post processing */
#define PPCOMP 1 /* Compute (run PPCOMP?.ABP) */
#define PPDLLC 2 /* Compute with DLL (run PPCOMP?.DLL) */
#define PPTRANS 4 /* Transmission (run PPTRANS?.ABP) */
#define PPABSRB 8 /* Absorbance (run PPABSRB?.ABP) */
#define PPKMUNK 4+8 /* Kuebelka-Munk (run PPKMUNK?.ABP) */
#define PPRAK 22 /* CPAMS byilt in rock midting and report
#define PPPEAK 32 /* GRAMS built-in peak picking and reporting */
#define PPSRCH 64 /* Library Search associated w/experiment LIB driver
                    128 /* General user-written post prgm. (run PPUSER?.ABP)
#define PPUSER
/****************************
^{\star} These are the xfrpost posting file dispositon settings. For PSTAPPD, the
* is appended to a .GDB general archive database (or .DBS custom database) if
* one is listed in xfrmtd. Otherwise it is appended to the current #R
#define PSTDEFT 0
                      /* Use default setting (xfrpost is "unspecified") */
                    /* Save file to disk (but remove from memory) */
/* Append to end of #R database or xfrmtd .GDB database
#define PSTSAVE 1
#define PSTAPPD 2
* /
#define PSTMERG 3
                    /* Merge into current #R database row */
                      /* Save as new background for experiment */
#define PSTBACK 4
#define PSTNONE 5
                      /* Do not save after processing (if any) */
                      /* Do not save and keep in memory as #S */
#define PSTKEEP 6
#define PSTBOTH 7
                      /* Both disk save & keep in #S memory (ABC Driver Only!)
```

Galactic Universal Data Format Specification 9/4/97

```
*/
#define PSTASK 8 /* Query user: save, keep, or both (ABC Driver Only!) */
```