## Recap from last time

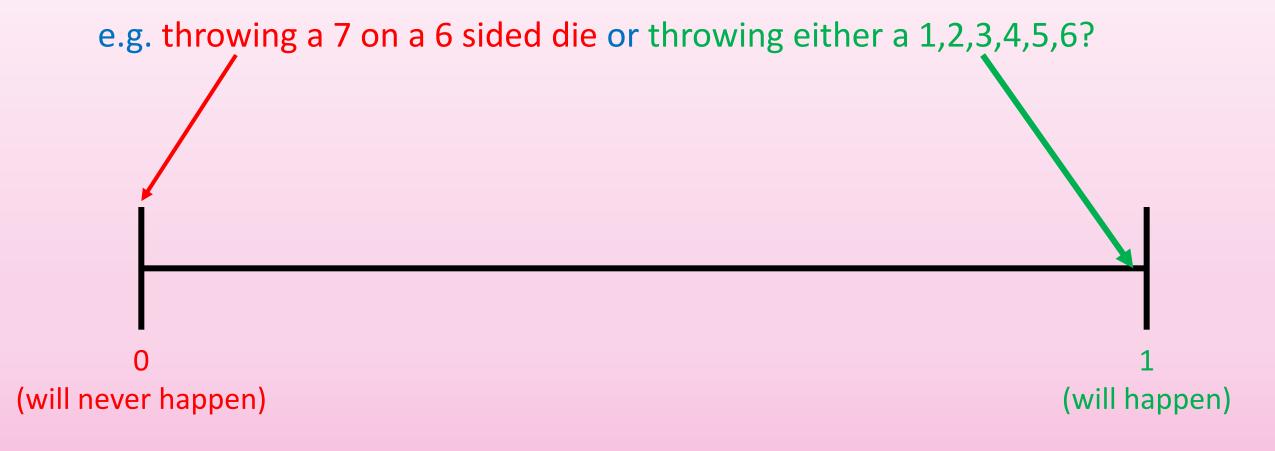
#### **Assumptions:**

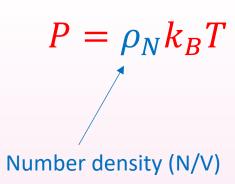
- Atmosphere is isothermal (in reality, dT/dh ~ 10 K km<sup>-1</sup>) [5.8]
- Treat the atmosphere as consisting of an ideal gas (good assumption for majority of component gases at low pressure – not water vapour)
- 3) The Earth is flat (~15 km << 6400 km)
- 4) The atmosphere is stationary and thus in mechanical equilibrium



#### Recap from last time

Probability exists on a number line between 0 and 1



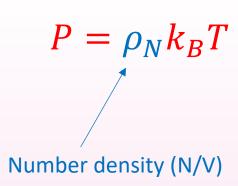


Forces: forces due to pressure and gravity (in equilibrium)

$$0 = P(h)A - P(h + dh)A - m(\rho_N dh)Ag$$

$$0 = dPA - m(\rho_N dh)Ag$$

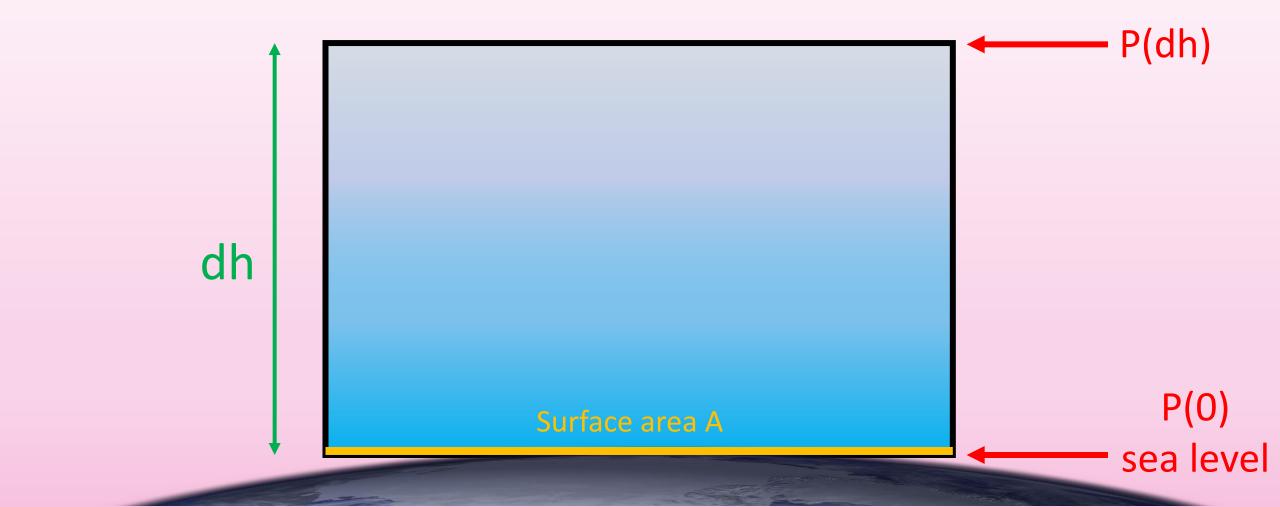




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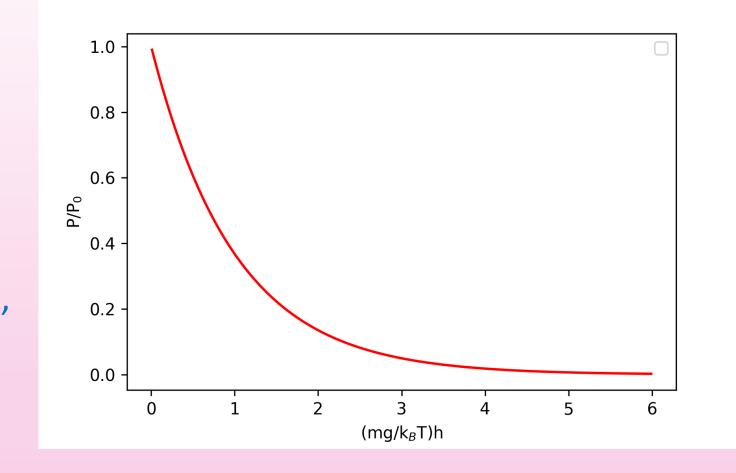


#### Isothermal model of the atmosphere

$$\rho_N(h) = \rho_N(0) e^{-\left(\frac{mgh}{k_B T}\right)}$$

$$P(h) = P(0)e^{-\left(\frac{mgh}{k_BT}\right)}$$

 $\frac{mgh}{k_BT}$  must be dimensionless, and so  $\frac{k_BT}{mg}$  must have dimensions of h -> scale height,  $h_0$  (at which  $P(h_0) = P(0)e^{-1}$ )



# Scale heights

Particle	Mass	Scale height
O <sub>2</sub> gas	32 amu = 5.3137e-26 kg	~10 km
N <sub>2</sub> gas	28 amu = 4.6495e-26 kg	~10 km
Covid virus	~10 <sup>-18</sup> kg	~0.5 mm
Smallest known virus	~10 <sup>-21</sup> kg	~50 cm
Bacteria	~10 <sup>-15</sup> kg	~5 μm
Dr Stuart Pirrie	~10 <sup>2</sup> kg	~50 ym (10 <sup>-24</sup> m is 1 ym)

#### Isothermal model meets probability

We have already established that the number density (number of fluid molecules per unit volume),  $\rho_N(0)$ , can be related to atmospheric pressure,  $P_{at}(=P(0))$ , by

$$\rho_N(0) = \frac{P_{at}}{k_B T}$$

If we wanted to determine the total number of gas molecules, N, in our slab of atmosphere (with surface area A), we can integrate across all possible heights,  $h = 0 \rightarrow h \equiv \infty$ ,

$$A\int\limits_{0}^{\infty}\rho_{N}(h)\,\mathrm{d}h=N$$

### Isothermal model meets probability

We can then, through solving the integral, relate the total number, N, to the number density  $\rho_N(0)$ :

$$\rho_N(0) = \frac{N}{Ah_0} \qquad \text{With } h_0 = \frac{k_B T}{mg}$$

Here, we can see that  $\rho_N(0)$  is in fact the average number density of the slab of atmosphere between h=0 and  $h=h_0$ 

We can then define a new quantity,  $Pr(h) = \frac{A\rho_N(h)}{N}$ , which is therefore the contribution from one molecule

#### Isothermal model meets probability

As  $A \int_0^\infty \rho_N(h) dh = N$ , we can show that

$$\int_0^\infty Pr(h) dh = \frac{A}{N} \int_0^\infty \rho_N(h) dh = \frac{N}{N} = 1$$

Thus, it is clear that the quantity Pr(h) is in some way a probability (normalised to be equal to 1 between 0 and infinity) – and the quantity Pr(h) dh gives the probability of finding a given molecule between h and h + dh

#### Probability density functions

We can also work out the function form of the probability by just looking at the expression

$$Pr(h) = \mathsf{C}\rho_N(0)\mathrm{e}^{-\left(\frac{mgh}{k_BT}\right)}$$

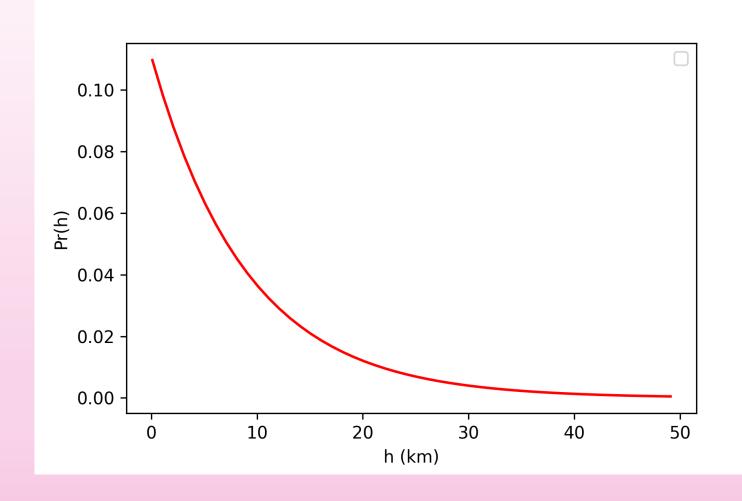
If we include some constant (C), we can make sure the integral of the function is equal to 1 and hence represents a probability to make a probability density function – in our case, this takes the form

$$Pr(h) = \frac{mg}{k_B T} e^{-\left(\frac{mgh}{k_B T}\right)} = \frac{1}{9020} e^{-\left(\frac{h}{9020}\right)} = \frac{1}{1} e^{-\left(\frac$$

### Probability density functions

$$Pr(h) = \frac{1}{9020} e^{-\left(\frac{h}{9020}\right)}$$

Without any microscopic information regarding the motion of these molecules, we can gather the information we're interested in just from this distribution



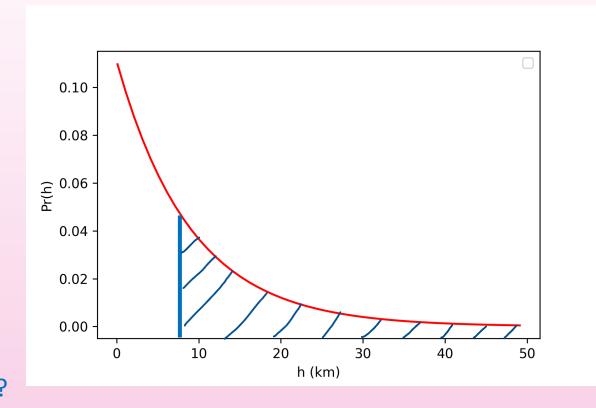
### Probability density functions

$$Pr(h) = \frac{1}{9020} e^{-\left(\frac{h}{9020}\right)}$$

Q1: For a random molecule in the atmosphere, what is the probability that it can be found above 8 km?

Q2: At any given time, what proportion of molecules in the atmosphere have a height greater than 8 km?

Q3: Averaged over a long timescale, what fraction of the time does a particular molecule spend at an altitude > 8 km?



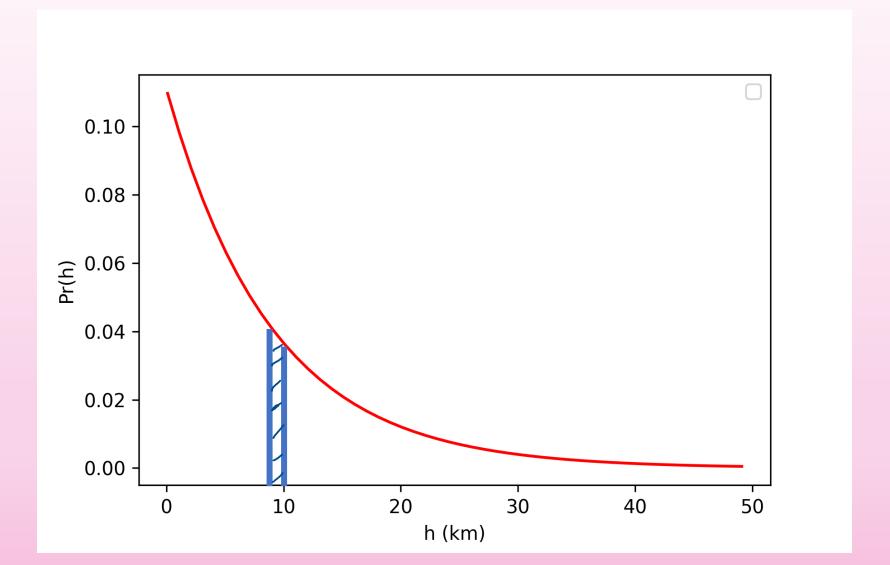
#### Identical questions!

A: 
$$Pr(h > 8 \ km) = \int_{8 \ km}^{\infty} Pr(h) \ dh = 0.41$$

#### More PDF examples

Roughly, what is the probability of finding a particle between 9 km and 10 km?

A: ~ 0.04



#### Boltzmann factors

Remember that the quantity  $\frac{mgh}{k_BT}$  must be dimensionless... what physically does it mean?

mgh = (gravitational) potential energy

 $k_BT$  = thermal energy

$$Pr(h) = \frac{A\rho_N(h)}{N} = \frac{mg}{k_B T} e^{-\left(\frac{mgh}{k_B T}\right)}$$

$$Pr(h) \propto e^{-\frac{E}{k_B T}}$$

This is the Boltzmann factor – gives the probability of measuring a certain energy state at a given temperature