

# Numerical Simulation of acoustic and electromagnetic scattering problems

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# Chapter 1

## The acoustic scattering Problem in full space

### 1.1 Introduction

We study the wave equation in full space  $\mathbb{R}^d$ ,  $d \in \{1, 2, 3\}$ .

$$\frac{\partial^2}{\partial t^2}p + \gamma \frac{\partial}{\partial t}p = c^2 \Delta p$$

Where

$c = c(x)$  speed of sound

$\gamma = \gamma(x)$  damping coefficient

Assume *time-periodic behaviour*

$$p(x, t) = \operatorname{Re}[u(x)e^{-i\omega t}]$$

with frequency  $\omega$  and a real-valued function  $u$ .

Since

$$\begin{aligned}\frac{\partial^2 p(x, t)}{\partial t^2} &= \operatorname{Re}[-\omega^2 u(x)e^{-i\omega t}] \\ \frac{\partial p(x, t)}{\partial t} &= \operatorname{Re}[-i\omega u(x)e^{-i\omega t}] \\ \Delta p(x, t) &= \operatorname{Re}[\Delta u(x)e^{-i\omega t}]\end{aligned}$$

for all times  $t > 0$ , we infer

$$-\omega^2 u - \gamma i\omega u = c^2 \Delta u, \quad \text{in } \mathbb{R}^d.$$

This is equivalent to

$$\Delta u + \frac{\omega^2}{c^2} \left( 1 + \frac{i\gamma}{\omega} \right) u = 0.$$

We assume, that  $c = c_0$  is constant in free space (reference value). We can now define a *wave number*

$$\kappa = \underbrace{\frac{\omega}{c_0}}_{>0} > 0$$

and the *index of refraction*

$$n(x) = \frac{c_0^2}{c(x)^2} \left( 1 + i \frac{\gamma}{\omega} \right).$$

This results in the *Helmholtz-Equation*

$$\Delta u + \kappa^2 n u = 0, \quad \text{in } \mathbb{R}^d.$$

We assume that there exists  $\alpha > 0$  such that

$$c(x) = c_0 \text{ and } \gamma(x) = 0, \quad \forall |x| > \alpha,$$

that means inhomogeneities of the medium lie inside some bounded region (inside a ball). Furthermore we assume, that outside this ball of radius  $\alpha$ ,  $\overline{B_\alpha(0)}$ , there are sources that generate plane waves, that is functions of the type

$$u^{in}(x) = e^{ikx \cdot \hat{\theta}},$$

where  $|\hat{\theta}| = 1$  for  $\hat{\theta} \in \mathbb{R}^d$  and  $\cdot$  denotes the inner product in  $\mathbb{R}^d$ . Then,  $u^{in}$  satisfies

$$\Delta u^{in} + k^2 u^{in} = 0.$$

(just calculation). This  $u^{in}$  generates a *scattered field*  $u^s$ . The *total field*

$$u = u^{in} + u^s$$

satisfies

$$\Delta u + \kappa^2 n u = 0.$$

Furthermore we assume *Sommerfeld's radiation condition* ( $r = |x|$ ):

$$\frac{\partial u^s}{\partial r} - i\kappa u^s(x) = o\left(r^{\frac{1-d}{2}}\right), \quad \text{as } r = |x| \rightarrow \infty$$

uniformly in  $\frac{x}{|x|}$ .

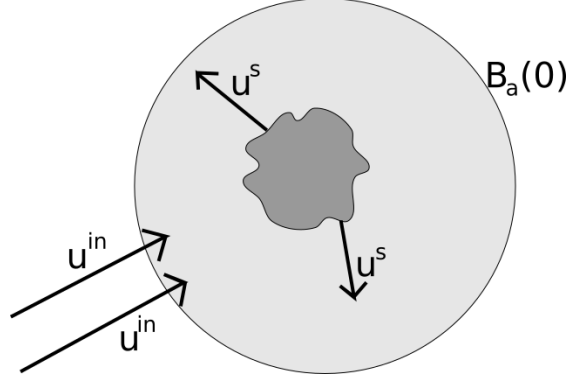


Figure 1.1: Visualization of the problem

## 1.2 Theory for the direct Scattering Problem in $\mathbb{R}^3$

Throughout this section  $d = 3$  holds! Furthermore

1.  $\hat{\theta} \in \mathbb{R}^3$  with  $|\hat{\theta}| = 1$  defines the *indicent field*

$$u^{in}(x) = e^{-i\kappa\hat{\theta}\cdot x}, \quad x \in \mathbb{R}^d$$

2.  $\kappa \in \mathbb{R}, \kappa > 0$
3.  $n \in L^\infty(\mathbb{R}^3), \operatorname{Re} n \geq 0, \operatorname{Im} n \geq 0$  and  $n(x) = 1$  for all  $x \in \mathbb{R}^3 \setminus B_\alpha(0)$  and some  $\alpha > 0$

We recall

$$H_{loc}^1(\mathbb{R}^3) := \{u : \mathbb{R}^3 \rightarrow \mathbb{C} : u|_K \in H^1(K), \quad \text{for every } K = B_R(0) \text{ and any } R > 0\}$$

The *Scattering Problem* reads as follows

**Definition 1.1 (Scattering Problem (S)).** Given  $\hat{\theta}, \kappa, n$  as above. Seek  $u \in H_{loc}^1(\mathbb{R}^3)$  such that

$$\Delta u + \kappa^2 n u = 0$$

in  $\mathbb{R}^3$  in the weak sense, that is

$$\int_{\mathbb{R}^3} (\nabla u \cdot \nabla \bar{\Psi} - \kappa n u \bar{\Psi}) \, dx = 0$$

for any  $\Psi \in H^1(\mathbb{R}^3)$  with compact support, and s.t.

$$u^s = u - u^{in}$$

satisfies Sommerfeld's radiation condition ( $d = 3$ , sharper version)

$$\frac{\partial u^s}{\partial r} - i\kappa u^s = O\left(\frac{1}{r^2}\right), \quad \text{uniformly in } \hat{x} = \frac{x}{|x|} \text{ as } r = |x| \rightarrow \infty$$

*Remark 1.2.* Regularity-theory tells us, that the radiation condition is well-defined, given the assumptions from [Definition 1.1](#).

**Theorem 1.3 (Rellich's Lemma)**

Let  $u$  satisfy  $\Delta u + \kappa^2 u = 0$  for every  $|x| > \alpha$ . The following property

$$\lim_{R \rightarrow \infty} \int_{|x|=R} |u(x)|^2 \, dS = 0$$

implies, that  $u(x) = 0$  for all  $|x| > \alpha$ .

*Remark 1.4.* Vice versa, if the property in [Theorem 1.3](#) does not hold, then  $u$  cannot vanish for all  $|x| > \alpha$ .

PROOF. Employs spherical Bessel functions, see e.g. ■

For the proof of uniqueness, we require results on periodic differential equations. We recall *Fourier representations of periodic functions*:

Define

$$Q := (-\pi, \pi)^3 \subseteq \mathbb{R}^3.$$

The functions

$$\{\varphi_j : j \in \mathbb{Z}^3\}$$

with

$$\varphi_j(x) = \frac{1}{(2\pi)^3} e^{ij \cdot x}, \quad \forall x \in Q, j \in \mathbb{Z}^3$$

define a complete orthonormal system (ONS) of  $L^2(Q)$  and every  $g \in L^2(Q)$  has the expansion

$$g = \sum_{j \in \mathbb{Z}^3} g_j \varphi_j, \quad \text{with } g_j = \int_Q g \varphi_j \, dx, \quad \text{for } j \in \mathbb{Z}^3.$$

Panceval's inequality shows

$$\|g\|_{L^2(Q)}^2 = \sum_{j \in \mathbb{Z}^3} |g_j|^2.$$

Furthermore, if  $g \in L^2(Q)$  and  $\sum_{j \in \mathbb{Z}^3} |j|^2 |g_j|^2 < \infty$ , where  $|j| = |j_1| + |j_2| + |j_3|$ , then  $g \in H^1(Q)$ .

Define

$$H_{per}^1(Q) := \left\{ g \in L^2(Q) : \sum_{j \in \mathbb{Z}^3} (1 + |j|)^2 |g_j|^2 < \infty \right\}$$

and identify  $L^2(Q)$  and  $H_{per}^1(Q)$  with the corresponding periodic functions in  $\mathbb{R}^3$  by

$$g(2\pi j + x) = g(x), \quad \text{for } x \in Q, j \in \mathbb{Z}^3.$$

**Lemma 1.5**

Let  $p \in \mathbb{R}^3, \alpha \in \mathbb{R}$  and  $\hat{e} = \begin{bmatrix} 1 \\ i \\ 0 \end{bmatrix}$ . Then, for every  $t > 0$  and every  $g \in L^2(Q)$  there exists a unique solution  $w = w_t(g) \in H_{per}^1(Q)$  to the differential equation

$$\Delta w + \lambda_t \cdot \nabla w + \mu_t w = g \quad (*)$$

understood weakly, that is

$$\forall \Psi \in C_c^\infty(Q) : \quad \int \left( -\nabla w \cdot \nabla \bar{\Psi} + (\lambda_t \cdot \nabla w + \mu_t w) \bar{\Psi} \right) dx = \int_Q g \bar{\Psi} dx$$

for  $\lambda_t := 2t\hat{e} - ip$ ,  $\mu_t = -(it + \alpha)$ . It holds that

$$\|w\|_{L^2(Q)} \leq \frac{1}{t} \|g\|_{L^2(Q)},$$

this means that

$$L_t : L^2(Q) \rightarrow L^2(Q) \text{ defined by } g \mapsto w_t(g)$$

defines a bounded operator

$$\|L_t\| \leq \frac{1}{t}, \quad \forall t > 0.$$

PROOF. With the Fourier expansions

$$g = \sum_{j \in \mathbb{Z}^3} g_j \varphi_j \quad \text{and} \quad w = \sum_{j \in \mathbb{Z}^3} w_j \varphi_j,$$

Eq. (\*) transforms to

$$\forall j \in \mathbb{Z}^3 : \quad c_j w_j = g_j, \quad \text{with } c_j = -|j|^2 + ij \cdot \lambda_t + \mu_t.$$

We have

$$|c_j| \geq |\operatorname{Im} c_j| \stackrel{\text{insert}}{=} |2tj \cdot \hat{e} - t| = t \underbrace{|2\hat{e} \cdot j - 1|}_{\geq 1} \geq t.$$

Thus, the operator

$$(L_t g) := \sum_{j \in \mathbb{Z}^3} \frac{g_j}{c_j} \varphi_j, \quad \text{with } g \in L^2(Q),$$

is well-defined and satisfies

$$\|L_t\| \leq \frac{1}{t}, \quad \forall t > 0.$$

It remains to show, that  $w$  is really a solution to the differential equation. In order to do so, let  $\Psi \in C_c^\infty(Q)$ . As  $\Psi$  has compact support, function-value and all derivatives are zero on the boundary of  $Q$ , and therefore we have also that  $\Psi$  is periodic. We can represent

$$\Psi = \sum_{j \in \mathbb{Z}^3} \Psi_j \varphi_j. \quad \blacksquare$$

Then

$$\begin{aligned} \int_Q \left( -\nabla w \cdot \nabla \bar{\Psi} + (\lambda_t \nabla w + \mu_t w) \bar{\Psi} \right) dx &= \sum_{j \in \mathbb{Z}^3} c_j w_j \cdot \bar{\Psi}_j \\ &= \sum_{j \in \mathbb{Z}^3} g_j \cdot \bar{\Psi}_j \\ &= \int_Q g \bar{\Psi} dx. \end{aligned}$$

**Theorem 1.6 (Unique continuation principle)**

Let  $u \in H_{loc}^1(\mathbb{R}^3)$  solve  $\Delta u + \kappa^2 n u = 0$ , where  $n \in L^\infty(\mathbb{R}^3)$  with  $n(x) = 1$  for  $|x| > \alpha$ , and let  $b \geq \alpha$ , such that  $u(x) = 0$  for all  $|x| \geq b$ . Then we have  $u = 0$  in  $\mathbb{R}^3$ .

*Remark 1.7.* Theorem 1.6 holds in a much more general version than we stated here.

PROOF. We introduce the scaling parameter  $\varrho = \frac{2b}{\pi}$  and the following function

$$w(x) := e^{\frac{i}{2}x_1 - t\hat{e} \cdot x} u(\varrho x), \quad x \in Q = (-\pi, \pi)^3$$

for some parameter  $t > 0$  and now  $\hat{e} = \begin{bmatrix} 1 \\ i \\ 0 \end{bmatrix} \in \mathbb{C}^3$ . Since  $w(x) = 0$  for all  $|x| \geq \frac{\pi i}{2}$  it can be extended to a  $2\pi$ -periodic function by

$$w(2\pi j + x) = w(x), \quad \forall j \in \mathbb{Z}^3, x \in Q,$$

with  $w \in H_{per}^1(Q)$ . It is readily seen that  $w$  satisfies

$$\Delta w + \lambda_t \cdot \nabla w + \mu_t w = -\varrho^2 \kappa^2 \tilde{n} w,$$



for

$$\lambda_t := 2t\hat{e} - \begin{bmatrix} i \\ 0 \\ 0 \end{bmatrix}$$

$$\mu_t := -\left(it + \frac{1}{4}\right)$$

and  $\tilde{n}$  is the periodic function

$$\tilde{n}(x + 2\pi j) = n(\varrho x), \quad x \in \bar{Q}, j \in \mathbb{Z}^3$$

(The proof of this is by straightforward differentiation). [Lemma 1.5](#) applies to this situation with  $g := -\varrho^2 \kappa^2 \tilde{n}w$  and yields

$$w = L_t g = -\varrho^2 \kappa^2 L_t(\tilde{n}w)$$

with  $\|L_t\| \leq \frac{1}{t}$ . This means that

$$\begin{aligned} \|w\|_{L^2(Q)} &\leq \frac{1}{t} \varrho^2 \kappa^2 \|\tilde{n}w\|_{L^2(Q)} \\ &\leq \frac{\varrho^2 \kappa^2 \|n\|_\infty}{t} \|w\|_{L^2(Q)}. \end{aligned}$$

This holds for all  $t > 0$ . Taking  $t \gg 1$  results in  $\|w\|_{L^2(Q)} = 0$  where  $w = 0$ . Thus  $u \equiv 0$ . ■

**Theorem 1.8 (Uniqueness)**

[Definition 1.1](#) admits at most one solution. That is,  $u^{in} \equiv 0$  implies  $u \equiv 0$ .

PROOF. Let  $u^{in} = 0$ . Recall the formula

$$|a - i\beta|^2 = |a|^2 + |\beta|^2 + 2 \operatorname{Im}(\beta \bar{a})$$

(elementary proof). We have

$$\begin{aligned} O\left(\frac{1}{R^2}\right) &= \int_{|x|=R} \left| \frac{\partial u}{\partial r} - i\kappa u \right|^2 dS \\ &= \int_{|x|=R} \left( \left| \frac{\partial u}{\partial r} \right|^2 + \kappa^2 |u|^2 \right) dS + 2\kappa \operatorname{Im} \int_{|x|=R} u \frac{\partial \bar{u}}{\partial r} dS \end{aligned}$$

The divergence-theorem implies

$$\begin{aligned} 2\kappa \operatorname{Im} \int_{|x|=R} u \underbrace{\frac{\partial \bar{u}}{\partial r}}_{=\nabla \bar{u} \cdot \underbrace{\nu}_{\text{normal}}} dS &= 2\kappa \operatorname{Im} \left[ \int_{|x|<R} \left( u \Delta \bar{u} + \underbrace{|\nabla u|^2}_{\in \mathbb{R}_0^+} \right) dx \right] \\ &\stackrel{\text{Helmholtz-eq}}{=} 2\kappa \operatorname{Im} \left[ \int_{|x|<R} \left( -\kappa^2 \bar{n} |u|^2 \right) dx \right] \\ &\geq 0, \end{aligned}$$

Since  $\text{Im } n \geq 0$ . Thus we have for large  $R \rightarrow \infty$ :

$$0 \leq \limsup_{R \rightarrow \infty} \int_{|x|=R} \left( \left| \frac{\partial u}{\partial r} \right|^2 + \kappa |u|^2 \right) dS \leq 0,$$

whence

$$\lim_{R \rightarrow \infty} \int_{|x|=R} |u|^2 dS = 0.$$

**Theorem 1.3** (Rellich) implies that  $u(x) = 0$  for all  $|x| > a$ . By the unique continuation principle we have  $u \equiv 0$  in  $\mathbb{R}^3$ . ■

We have shown uniqueness. For the existence proof we will construct solutions.

**Definition 1.9.** The function

$$\Phi(x, y) := \frac{1}{4\pi|x-y|} e^{i\kappa|x-y|}, \quad x, y \in \mathbb{R}^3, x \neq y$$

is called *fundamental solution* or *free space Green's function*.

$\Phi$  has the following properties

**Proposition 1.10 (Properties of the fundamental solution)**

1.  $\Phi(\cdot, y)$  satisfies the Helmholtz-equation  $\Delta u + \kappa^2 u = 0$  in  $\mathbb{R}^3 \setminus \{y\}$
2.  $\Phi$  satisfies the *empradiation condition*

$$\frac{x}{|x|} \cdot \nabla_x \Phi(x, y) - i\kappa \Phi(x, y) = O\left(\frac{1}{|x|^2}\right), \quad \text{as } |x| \rightarrow \infty$$

*uniformly in  $\hat{x} = \frac{x}{|x|} \in \mathcal{S}^2$  (Sphere)*

3.  $\Phi$  has the *asymptotic behaviour*

$$\Phi(x, y) = \frac{e^{i\kappa|x|}}{4\pi|x|} e^{-i\kappa\hat{x} \cdot y} + O\left(\frac{1}{|x|^2}\right), \quad \text{as } |x| \rightarrow \infty$$

*uniformly in  $\hat{x} = \frac{x}{|x|} \in \mathcal{S}^2$  (Sphere)*

PROOF. The first 2 properties are easily verified. We prove the third:

We calculate the the binomial formula

$$\begin{aligned}
 |x - y| - (|x| - \hat{x} \cdot y) &= \frac{|x - y|^2 - (|x| - \hat{x} \cdot y)^2}{|x - y| + (|x| - \hat{x} \cdot y)} \\
 &= \frac{|y|^2 - 2\hat{x} \cdot y + |x|^2 - |x|^2 + 2\hat{x} \cdot y - \overbrace{(\hat{x} \cdot y)^2}^{\leq |y|^2}}{|x| \left( 1 + \underbrace{\left| \hat{x} - \frac{y}{|x|} \right|}_{1 - \left| \frac{y}{|x|} \right|} - \frac{\hat{x} \cdot y}{|x|} \right)} \\
 &\leq \frac{|y|^2}{2|x| \left( 1 - \frac{|y|}{|x|} \right)}.
 \end{aligned}$$

Hence  $|x - y| = |x| - \hat{x} \cdot y + O\left(\frac{1}{|x|}\right)$ . Take the exponential of this to get

$$e^{i\kappa|x-y|} = e^{i\kappa|x|} e^{-i\kappa\hat{x} \cdot y} \left( 1 + O\left(\frac{1}{|x|}\right) \right).$$

By the above formula we also have

$$\frac{1}{|x - y|} = \frac{1}{|x|} + \left[ \frac{|x| - |x - y|}{|x - y||x|} \right] = \frac{1}{|x|} + O\left(\frac{1}{|x|^2}\right).$$

So

$$\begin{aligned}
 \frac{e^{i\kappa|x-y|}}{|x - y|} &= \frac{e^{i\kappa|x-y|}}{|x|} + O\left(\frac{1}{|x|^2}\right) \\
 &= \frac{1}{|x|} e^{i\kappa|x|} e^{-i\kappa\hat{x} \cdot y} + O\left(\frac{1}{|x|^2}\right),
 \end{aligned}$$

as  $|x| \rightarrow \infty$ . ■

*Remark 1.11.* The previous result states in particular that  $\Phi$  is determined by the function

$$e^{-i\kappa\hat{x} \cdot y}$$

on  $\mathcal{S}^2$  (up to pertubation). We will see that this property holds in a more general context.

With the help of the fundamental solution  $\Phi$  we can create

**Theorem 1.12**

Let  $\Omega \subseteq \mathbb{R}^3$  be bounded. For every  $\phi \in L^2(\Omega)$ , the function

$$v(x) := \int_{\Omega} \phi(y) \Phi(x, y) dy, \quad x \in \mathbb{R}^3$$

belongs to  $H_{loc}^1(\mathbb{R}^3)$  and satisfies the Sommerfeld radiation condition. Moreover,  $v$  is the only radiating solution (it is the only function in the set of solution that satisfies the Sommerfeld radiation condition) to

$$\Delta v + \kappa^2 v = -\phi$$

in the variational sense

$$\int_{\mathbb{R}^3} (\nabla v \cdot \nabla \bar{\Psi} - \kappa^2 v \bar{\Psi}) dx = \int_{\mathbb{R}^3} \phi \bar{\Psi} dx \quad (*)$$

for all  $\Psi \in H^1(\mathbb{R}^3)$  with compact support. For any  $R > 0$  with  $\Omega \subseteq B_R(0) =: K$  we have with  $c = c(R, \kappa, \Omega)$  that

$$\|v\|_{H^1(K)} \leq c \|\phi\|_{L^2(\Omega)}.$$

In other words, the mapping  $\phi \mapsto v$  is a bounded (continuous linear) operator from  $L^2(\Omega)$  to  $H^1(K)$ .

PROOF. (1) Since  $\frac{1}{r^2}$  is locally integrable in  $\mathbb{R}^3$ , the expression on the right-hand side in Eq. (\*) is well-defined. Provided  $\phi \in C^1(\bar{\Omega})$ , then we can interchange integration and differentiation (possible as we have a majorant, cf. PDE theory) and obtain

$$\Delta v + \kappa^2 v = \begin{cases} -\phi & \text{in } \Omega \\ 0 & \text{in } \mathbb{R}^3 \setminus \bar{\Omega} \end{cases}$$

This means that  $v$  solves Eq. (\*) for any  $\kappa \in \mathbb{C}$ .

(2) We cannot directly evaluate  $L^2$ -integrals of the gradient  $\nabla_x \Phi$ . To prove stability, we first consider the special case  $\kappa = i$  and  $\phi \in C^1(\bar{\Omega})$ . Then, the fundamental solution

$$\Phi_i(x, y) = \frac{1}{4\pi|x-y|} e^{-|x-y|}$$

decays exponentially for  $|x| \rightarrow \infty$ , and by approximation arguments Eq. (\*) holds for any  $\Psi \in H^1(\mathbb{R}^3)$ .

Taking  $\Psi = v$  we obtain

$$\|v\|_{H^1(\mathbb{R}^3)}^2 = \int_{\Omega} \phi \bar{v} \leq \|\phi\|_{L^2(\Omega)} \|v\|_{H^1(\mathbb{R}^3)}$$

(note that Eq. (\*) with  $\kappa = i$  becomes the  $H^1$ -Norm). By density, this defines a continuous operator

$$\phi \mapsto v \quad \text{from } L^2(\Omega) \text{ to } H^1(K).$$

(3) Let  $k > 0$ . We define

$$\Psi(x, y) := \Phi_k(x, y) - \Phi_i(x, y) = \frac{1}{4\pi|x-y|} [e^{i\kappa|x-y|} - e^{-|x-y|}]$$

It is easy to prove that  $\Psi$  and  $\nabla_x \Psi$  belong to  $L^2(K \times \Omega)$ . We sketch the crucial part. We calculate

$$4\pi|\nabla_x \Psi(x, y)| = \left| \underbrace{\frac{i\kappa e^{i\kappa|x-y|} - e^{-|x-y|}}{|x-y|}}_{O(\frac{1}{|x-y|})} + \frac{e^{i\kappa|x-y|} - e^{-|x-y|}}{|x-y|^3} \right|$$

$$\stackrel{*}{\leq} O\left(\frac{1}{|x-y|}\right) \quad \text{for small } |x-y|$$

\*: The denominator of the right fraction is in  $O(|x-y|)$

Thus,

$$\|\nabla_x \Psi\|_{L^2(K \times \Omega)} \leq C \int_K \int_\Omega \frac{1}{|x-y|^2} dy dx < \infty.$$

(4) With (3) we see that the mapping

$$\varphi \mapsto \int_\Omega \varphi(y) \Psi(\cdot, y) dy$$

is bounded from  $L^2(\Omega)$  to  $H^1(K)$ , because

$$\begin{aligned} \int_K \left| \nabla_x \int_\Omega \varphi(y) \Psi(x, y) dy \right|^2 dx &= \int_K \left[ \int_\Omega |\varphi(y) \nabla_x \Psi(x, y)| dy \right]^2 dx \\ &\leq \int_K \|\varphi\|_{L^2(\Omega)}^2 \|\nabla_x \Psi(x, \cdot)\|_{L^2(\Omega)}^2 dx \\ &\leq \|\varphi\|_{L^2(\Omega)}^2 \underbrace{\|\nabla_x \Psi\|_{L^2(K \times \Omega)}^2}_{\stackrel{(3)}{< \infty}}. \end{aligned}$$

This and (2) show that  $\varphi \mapsto v$  is also bounded from  $L^2(\Omega)$  to  $H^1(K)$  for  $\kappa > 0$ .

(5) The radiation condition follows from the radiation condition of  $\Phi$ :

$$\begin{aligned} \frac{\partial v}{\partial \nu} - i\kappa v &= \int_\Omega \varphi(y) \left( \frac{\partial}{\partial \nu} - i\kappa \right) \Phi(x, y) dy \\ &\leq \|\varphi\|_{L^2(\Omega)} O\left(\frac{1}{r}\right)^2. \end{aligned}$$

Uniqueness follows from (2). ■

*Remark 1.13. Proof of Theorem 1.12 (1).*

To prove that  $\Delta v(x) + \kappa^2 v(x) = -\varphi(x)$ , for  $x \in \Omega$ , it satisfies to verify

$$(\Delta + \kappa^2) \int_{B_\varepsilon(x)} \varphi(y) \Phi(x, y) dy = -\varphi(x),$$

for small  $\varepsilon > 0$ .

1) We readily see that

$$\kappa^2 \left| \int_{B_\varepsilon(x)} \varphi(y) \Phi(x, y) dy \right| \leq \frac{\kappa^2}{4\pi} \left| \int_{B_\varepsilon(x)} \frac{e^{i\kappa|y-x|}}{|x-y|} dy \right|$$

and the expression on the right tends to 0 for  $\varepsilon \rightarrow 0$ .

2) Since  $\Delta_x \Phi(x, y) = \Delta_y \phi(x, y)$ , we see

$$\begin{aligned} \Delta_y \int_{B_\varepsilon(x)} \varphi(y) \phi(x, y) dy &= \int_{B_\varepsilon(x)} \phi(y) \Delta_y \Phi(x, y) dy \\ &= \underbrace{- \int_{B_\varepsilon(x)} \nabla \varphi(y) \cdot \nabla \Phi(x, y)}_{=:A} + \underbrace{\int_{|y-x|=\varepsilon} \varphi(y) \frac{\partial \Phi(x, y)}{\partial \nu} dS(y)}_{=:B}. \end{aligned}$$

We have that

$$\nabla_y \Phi(x, y) = \frac{1}{4\pi} e^{i\kappa|x-y|} \left( \frac{1}{|x-y|^2} - \frac{1}{|x-y|^3} \right) (x-y),$$

and thus

$$A \leq \|\nabla \varphi\|_{L^\infty} \int_{B_\varepsilon(x)} |\nabla_y \Phi(x, y)| dy \rightarrow 0, \text{ as } \varepsilon \rightarrow 0.$$

We stay with  $B$ :

$$\begin{aligned} B &= \frac{1}{4\pi} \int_{\partial B_\varepsilon(x)} e^{i\kappa\varepsilon} \left( \frac{1}{\varepsilon} - \frac{1}{\varepsilon^2} \right) \varphi(y) dS(y) \\ &= \underbrace{\frac{1}{4\pi} e^{i\kappa\varepsilon} \left( \frac{1}{\varepsilon} - \frac{1}{\varepsilon^2} \right) 4\pi\varepsilon^2}_{=: e^{i\kappa\varepsilon}(\varepsilon + \varepsilon^{-1}) \rightarrow -1, \text{ as } \varepsilon \rightarrow 0} \underbrace{\oint_{\partial B_\varepsilon(x)} \varphi(y) dS(y)}_{\rightarrow \varphi(x), \text{ as } \varepsilon \rightarrow 0} \end{aligned}$$

**Theorem 1.14 (Lippmann Schwinger integral equation)**

Let  $\alpha > 0$  and  $K = B_\alpha(0)$ . If  $u \in H^1$  solves the [Definition 1.1](#), then  $u|_K \in L^2(K)$  satisfies the Lippmann Schwinger equation

$$u(x) = u^i n(x) - \kappa^2 \int_{|y| < \alpha} (1 - n(y)) \Phi(x, y) u(y) dy \quad (\text{LS})$$

for almost all  $x \in K$ .

Conversely, if  $u \in L^2(K)$  satisfies [Eq. \(LS\)](#), then it can be extended by the right-hand side of [Eq. \(LS\)](#) to the solution  $u \in H_{loc}^1(\mathbb{R}^3)$  of [Definition 1.1](#).

PROOF. Let  $u$  satisfy Definition 1.1 and define

$$v = \int_K \varphi(y) \Phi(\cdot, y) dy$$

for  $\varphi = \kappa^2(1-n)u \in L^2(K)$ . Theorem 1.12 states that  $v \in H_{loc}^1(\mathbb{R}^3)$  satisfies  $\Delta v + \kappa^2 v = -\varphi$ . Since

$$\Delta u + \kappa^2 u = \kappa^2(1-n)u$$

and

$$\Delta u^{in} + \kappa u^{in} = 0,$$

we have (addition of previous equations)

$$\Delta(v + u^s) + \kappa^2(v + u^s) = 0.$$

The uniqueness from Theorem 1.8 shows that  $v + u^s = 0$ . Therefore

$$u = u^{in} + u^s = u^{in} - v.$$

For the converse direction let  $u \in L^2(K)$  satisfy Eq. (LS). Define  $v$  as above, such that  $u = u^{in} - v$  in  $K$  (in the  $L^2$ -sense). By Theorem 1.12 we know that  $v \in H_{loc}^1(\mathbb{R}^3)$  by extending it to  $\mathbb{R}^3$ , and that  $\Delta v + \kappa v = -\varphi$ . This implies  $u \in H_{loc}^1(\mathbb{R}^3)$ . Furthermore

$$\Delta u + \kappa^2 u = \varphi = \kappa^2(1-n)u,$$

that is

$$\Delta u + \kappa^2 n u = 0.$$

This implies that  $u^s = -v$ . The radiation condition follows again from Theorem 1.12. ■

We now can derive existence for Definition 1.1.

**Theorem 1.15 (Existence of solutions for Definition 1.1)**

*Let  $\kappa, n, \theta$  be as assumed previously. Then, the Lippmann Schwinger equation (LS), and thus Definition 1.1, is uniquely solvable.*

PROOF. Define the following operator  $T : L^2(B_\alpha(0)) \rightarrow L^2(B_\alpha(0))$  by

$$(Tu)(x) = \kappa^2 \int_{|y| < \alpha} (1 - n(y)) \Phi(x, y) u(y) dy, \text{ for } |x| < \alpha.$$

Theorem 1.12 shows that  $T$  is bounded from  $L^2(B_\alpha(0))$  to  $H^1(B_\alpha(0))$ , and by the compact embedding  $H^1(B_\alpha(0)) \hookrightarrow L^2(B_\alpha(0))$ , the operator  $T$  is compact from  $L^2(B_\alpha(0))$  into itself. The Lippmann Schwinger equation reads

$$u + Tu = u^{in}$$

and by the Riesz theory (Fredholm alternative) for compact operators, the uniqueness of  $u$  (Theorems 1.8 and 1.14) implies its existence. ■

**Theorem 1.16 (Far-field pattern)**

Let  $u$  solve [Definition 1.1](#). Then we have

$$u(x) = u^{in}(x) + \frac{e^{i\kappa|x|}}{|x|} u_{\infty}(\hat{x}) + O\left(\frac{1}{|x|^2}\right), \text{ as } |x| \rightarrow \infty,$$

uniformly in  $\hat{x} = \frac{x}{|x|}$  (direction of  $x$ ). The function  $u_{\infty}(\hat{x})$  is given by

$$u_{\infty}(\hat{x}) = \frac{\kappa^2}{4\pi} \int_{|y| < \alpha} (n(y) - 1) e^{-i\kappa\hat{x} \cdot y} u(y) dy$$

is called far-field pattern or scattering amplitude

$$u_{\infty} : \underbrace{\mathcal{S}^2}_{\text{Sphere}} \rightarrow \mathbb{C}$$

is analytic on  $\mathcal{S}^2$  and determines  $u^s$  outside  $B_{\alpha}(0)$  uniquely ( $u_{\infty} = 0 \Leftrightarrow u^s(x) = 0$  for  $|x| > \alpha$ ).

The formula for  $u(x)$  follows from [Proposition 1.10](#) (iii):

$$\Phi(x, y) = \frac{e^{i\kappa|x|}}{|x|} e^{-i\kappa\hat{x} \cdot y} + O\left(\frac{1}{|x|^2}\right).$$

By [Theorem 1.3](#) (Rellich) we have

$$\lim_{R \rightarrow 0} \int_{|x|=R} |u(x)|^2 dS(x) = 0, \text{ for } |x| > \alpha$$

and this implies uniqueness. The analyticity follows from the formula (calculate derivative w.r.t  $\hat{x}$ ).

*Remark 1.17.* 1. The concept of far-field pattern is fundamental to inverse scattering theory.

2. The analyticity of  $u_{\infty}$  shows that the inverse scattering problem is ill-posed (little perturbations of  $u_{\infty}$  can destroy analyticity of the function!).