

PHYS 506 Assignment 5

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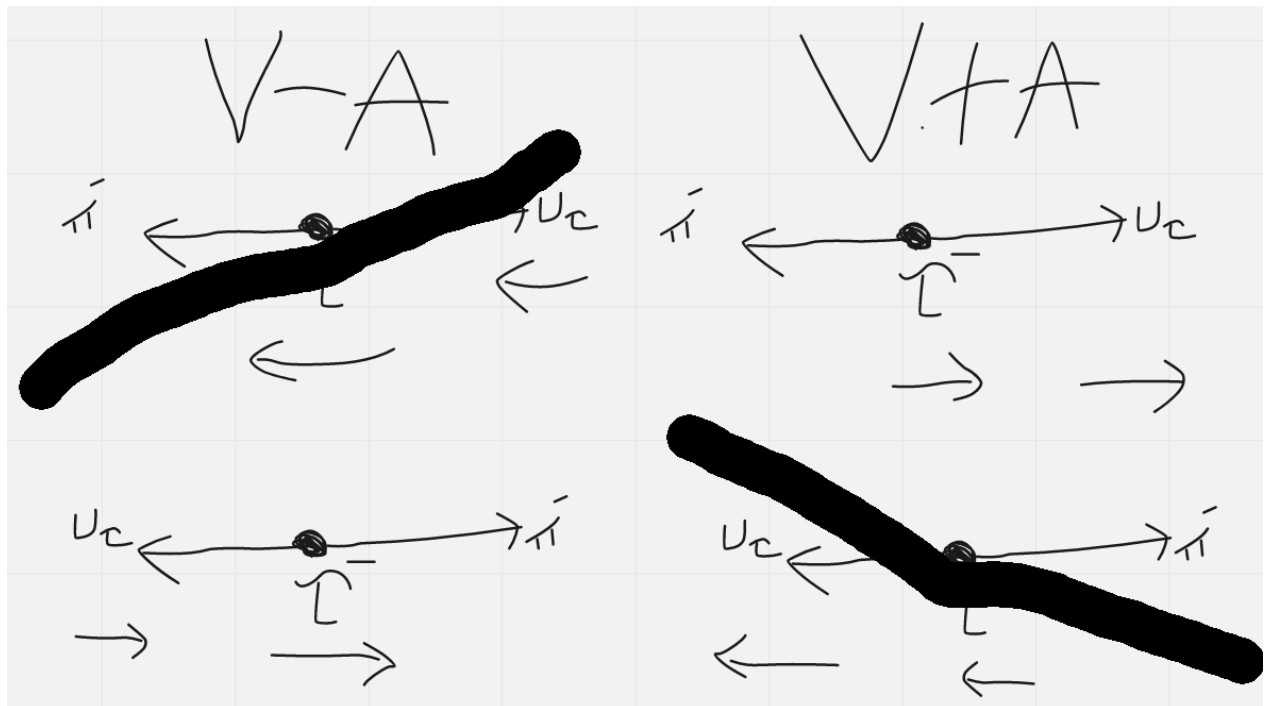
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- 1 Consider the decay at rest $\tau^- \rightarrow \pi^- \nu_\tau$, where the spin of the tau is in the positive z -direction and the ν_τ and π^- travel in the z -directions. Sketch the allowed spin configurations assuming that the form of the weak charged-current interaction is i) V-A and ii) V+A.

i) In this case, V-A \Rightarrow only LH chiral particle states / RH chiral anti-particle states have non-zero projections (the RH particle / LH anti-particle states go to zero). So in our interaction $\tau^- \rightarrow \pi^- \nu_\tau$, the neutrino is produced in a LH chiral state. Using the fact that neutrinos are nearly massless, we can say LH chiral \approx LH helicity for the neutrino, which gives us our spin anti-parallel to momentum, as shown on the left.

ii) In this case, we have the opposite thing where only RH particle / LH anti-particle states have non-zero projections \Rightarrow RH chiral neutrino \approx RH helicity neutrino \Rightarrow spin parallel to momentum.

The question says to make the spin of τ positive, so that's why 2/4 are crossed out.

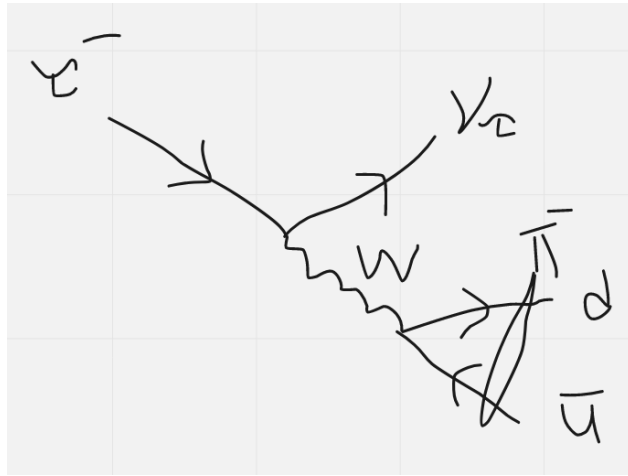


2 Calculate the partial decay width for the decay $\tau^- \rightarrow \pi^- \nu_\tau$ in the following steps.

a): Draw the Feynman diagram and show that the corresponding matrix element is

$$\mathcal{M} = \sqrt{2}G_F f_\pi \bar{u}(p_\nu) \gamma^\mu \frac{1}{2}(1 - \gamma^5) u(p_\tau) g_{\mu\nu} p_\pi^\nu$$

The Feynman diagram:



The parts:

- Incoming τ^- : $u(p_\tau)$
- Outgoing ν_τ : $\bar{u}(p_\nu)$
- W^- propagator: $\frac{-i[g_{\mu\nu} - q_\mu q_\nu / m_W^2]}{q^2 - m_W^2}$ (or appx $\frac{-ig_{\mu\nu}}{m_W^2}$)
- $\tau^- \rightarrow W^- \nu_\tau$ vertex: $\frac{-ig_W}{\sqrt{2}} \frac{1}{2} \gamma^\mu (1 - \gamma^5)$
- Outgoing pion with the vertex: $\frac{-ig_W}{\sqrt{2}} \frac{1}{2} f_\pi p_\pi^\nu$

Put those all together:

$$\begin{aligned}
-iM_{fi} &= \bar{u}(p_\nu) \frac{-ig_W}{\sqrt{2}} \frac{1}{2} \gamma^\mu (1 - \gamma^5) u(p_\tau) \frac{-ig_{\mu\nu}}{m_W^2} \frac{g_W}{\sqrt{2}} \frac{1}{2} f_\pi p_\pi^\nu \\
M_{fi} &= \frac{1}{m_W^2} \frac{g_W}{\sqrt{2}} \frac{1}{2} f_\pi \frac{g_W}{\sqrt{2}} \bar{u}(p_\nu) \gamma^\mu \frac{1}{2} (1 - \gamma^5) u(p_\tau) g_{\mu\nu} p_\pi^\nu \\
&= \frac{g_W^2}{4m_W^2} f_\pi \bar{u}(p_\nu) \gamma^\mu \frac{1}{2} (1 - \gamma^5) u(p_\tau) g_{\mu\nu} p_\pi^\nu \\
&= \sqrt{2} G_F f_\pi \bar{u}(p_\nu) \gamma^\mu \frac{1}{2} (1 - \gamma^5) u(p_\tau) g_{\mu\nu} p_\pi^\nu
\end{aligned}$$

That last line came from $\frac{G_F}{\sqrt{2}} = \frac{g_W^2}{8m_W^2}$ (for low q^2).

b): Taking the τ^- spin in the z -direction and the four-momentum of the neturino to be $p_\nu = p^*(1, \sin(\theta), 0, \cos(\theta))$, show that the leptonic current is $j^\mu = \sqrt{2}m_\tau p^*(-s, -c, -ic, s)$, where $s = \sin(\theta/2), c = \cos(\theta/2)$. For this configuration, the spinor for the τ^- can be taken to be u_1 for a particle at rest.

The leptonic current: $\bar{u}(p_\nu) \gamma^\mu \frac{1}{2} (1 - \gamma^5) u(p_\tau)$. Make the approximation that chirality = helicity for high-energy particles like this.

We have $p_\nu = p^*(1, \sin(\theta), 0, \cos(\theta)), p_\tau = (m_\tau, 0, 0, 0)$, i.e. $s_\tau = 0, c_\tau = 1$

$$\begin{aligned}
\bar{u}(p_\nu) \gamma^\mu \frac{1}{2} (1 - \gamma^5) u(p_\tau) &= \bar{u}(p_\nu) \gamma^\mu \frac{1}{2} (1 - \gamma^5) u_1(p_\tau) \\
&= \bar{u}(p_\nu) \gamma^\mu \frac{1}{2} (1 - \gamma^5) \sqrt{E_\tau + m_\tau} \begin{pmatrix} 1 \\ 0 \\ 0 \\ 0 \end{pmatrix} \\
&= \bar{u}(p_\nu) \gamma^\mu \frac{1}{2} \begin{pmatrix} 1 & 0 & -1 & 0 \\ 0 & 1 & 0 & -1 \\ -1 & 0 & 1 & 0 \\ 0 & -1 & 0 & 1 \end{pmatrix} \sqrt{2m_\tau} \begin{pmatrix} 1 \\ 0 \\ 0 \\ 0 \end{pmatrix} \\
&= \sqrt{2m_\tau} \frac{1}{2} \bar{u}(p_\nu) \gamma^\mu \begin{pmatrix} 1 \\ 0 \\ -1 \\ 0 \end{pmatrix}
\end{aligned}$$

For $u(p_\nu)$ we can use u_\downarrow since we had P_L , and $\phi = 0$ wlog since it's along the z axis:

$$\begin{aligned}
\bar{u}(p_\nu) &= u_\downarrow(p_\nu)^\dagger \gamma^0 \\
&= \sqrt{E_\nu} \begin{pmatrix} -s \\ c \\ s \\ -c \end{pmatrix}^\dagger \gamma^0 \\
&= \sqrt{p^*} \begin{pmatrix} -s & c & -s & c \end{pmatrix}
\end{aligned}$$

Calculate some matrix products:

$$\begin{aligned}
\gamma^0 u_1(p_\tau) &= \begin{pmatrix} 1 & 0 & 0 & 0 \\ 0 & 1 & 0 & 0 \\ 0 & 0 & -1 & 0 \\ 0 & 0 & 0 & -1 \end{pmatrix} \begin{pmatrix} 1 \\ 0 \\ -1 \\ 0 \end{pmatrix} \\
&= \begin{pmatrix} 1 \\ 0 \\ 1 \\ 0 \end{pmatrix} \\
\begin{pmatrix} -s & c & -s & c \end{pmatrix} \gamma^0 u_\downarrow(p_\tau) &= \begin{pmatrix} -s & c & -s & c \end{pmatrix} \begin{pmatrix} 1 \\ 0 \\ 1 \\ 0 \end{pmatrix} \\
&= -2s
\end{aligned}$$

$$\begin{aligned}
\gamma^1 u_1(p_\tau) &= \begin{pmatrix} 0 & 0 & 0 & 1 \\ 0 & 0 & 1 & 0 \\ 0 & -1 & 0 & 0 \\ -1 & 0 & 0 & 0 \end{pmatrix} \begin{pmatrix} 1 \\ 0 \\ -1 \\ 0 \end{pmatrix} \\
&= \begin{pmatrix} 0 \\ -1 \\ 0 \\ -1 \end{pmatrix} \\
(-s \quad c \quad -s \quad c) \gamma^1 u_\downarrow(p_\tau) &= (-s \quad c \quad -s \quad c) \begin{pmatrix} -1 \\ 0 \\ -1 \\ 0 \end{pmatrix} \\
&= -2c \\
\gamma^2 u_\downarrow(p_\tau) &= \begin{pmatrix} 0 & 0 & 0 & -i \\ 0 & 0 & i & 0 \\ 0 & i & 0 & 0 \\ -i & 0 & 0 & 0 \end{pmatrix} \begin{pmatrix} 1 \\ 0 \\ -1 \\ 0 \end{pmatrix} \\
&= \begin{pmatrix} 0 \\ -i \\ 0 \\ -i \end{pmatrix} \\
(-s \quad c \quad -s \quad c) \gamma^2 u_\downarrow(p_\tau) &= (-s \quad c \quad -s \quad c) \begin{pmatrix} 0 \\ -i \\ 0 \\ -i \end{pmatrix} \\
&= -2ic
\end{aligned}$$

$$\begin{aligned}
\gamma^3 u_\downarrow(p_\tau) &= \begin{pmatrix} 0 & 0 & 1 & 0 \\ 0 & 0 & 0 & -1 \\ -1 & 0 & 0 & 0 \\ 0 & 1 & 0 & 0 \end{pmatrix} \begin{pmatrix} 1 \\ 0 \\ -1 \\ 0 \end{pmatrix} \\
&= \begin{pmatrix} -1 \\ 0 \\ -1 \\ 0 \end{pmatrix} \\
(-s \quad c \quad -s \quad c) \gamma^2 u_\downarrow(p_\tau) &= (-s \quad c \quad -s \quad c) \begin{pmatrix} -1 \\ 0 \\ -1 \\ 0 \end{pmatrix} \\
&= 2s
\end{aligned}$$

Now put it all together:

$$\begin{aligned}
\bar{u}(p_\nu) \gamma^\mu \frac{1}{2} (1 - \gamma^5) u(p_\tau) &= \sqrt{2m_\tau} \frac{1}{2} \bar{u}(p_\nu) \gamma^\mu \begin{pmatrix} 1 \\ 0 \\ -1 \\ 0 \end{pmatrix} \\
&= \sqrt{2m_\tau} \frac{1}{2} \sqrt{p^*} (-2s \quad -2c \quad -2ic \quad 2s) \\
&= \sqrt{2m_\tau p^*} (-s, -c, -ic, s)
\end{aligned}$$

c): Write down the 4-momentum of the π^- and show that $|\mathcal{M}|^2 = 4G_F^2 f_\pi^2 m_\tau^3 p^* \sin^2(\theta/2)$.

We have $p_\nu = p^*(1, \sin(\theta), 0, \cos(\theta))$, $p_\tau = (m_\tau, 0, 0, 0)$. Conservation of momentum:

$$\begin{aligned}
p_\tau &= p_\nu + p_\pi \\
(m_\tau, 0, 0, 0) &= p^*(1, \sin(\theta), 0, \cos(\theta)) + p_\pi \\
(m_\tau - p^*, -p^* \sin(\theta), 0, -p^* \cos(\theta)) &= p_\pi
\end{aligned}$$

So then we have

$$\begin{aligned}
M_{fi} &= \sqrt{2}G_F f_\pi \bar{u}(p_\nu) \gamma^\mu \frac{1}{2}(1 - \gamma^5) u(p_\tau) g_{\mu\nu} p_\pi^\mu \\
&= \sqrt{2}G_F f_\pi \sqrt{2m_\tau p^*} (-s, -c, -ic, s) g_{\mu\nu} (m_\tau - p^*, -p^* \sin(\theta), 0, -p^* \cos(\theta)) \\
&= \sqrt{2}G_F f_\pi \sqrt{2m_\tau p^*} (-s(m_\tau - p^*) - c(p^* \sin(\theta)) + 0 + sp^* \cos(\theta)) \\
&= \sqrt{2}G_F f_\pi \sqrt{2m_\tau p^*} (-sm_\tau + p^*(s - c \sin(\theta) + s \cos(\theta))) \\
&= 2G_F f_\pi \sqrt{m_\tau p^*} (-sm_\tau + p^*(0)) \\
&= -2G_F f_\pi \sqrt{p^*} m_\tau^{\frac{3}{2}} \sin\left(\frac{\theta}{2}\right) \\
|M_{fi}|^2 &= 4G_F^2 f_\pi^2 m_\tau^3 p^* \sin^2\left(\frac{\theta}{2}\right)
\end{aligned}$$

d): Hence show that $\Gamma(\tau^- \rightarrow \pi^- \nu_\tau) = \frac{G_F^2 f_\pi^2}{16\pi} m_\tau^3 \left(\frac{m_\tau^2 - m_\pi^2}{m_\tau^2}\right)^2$

We know that for $a \rightarrow 1 + 2$,

$$\Gamma = \frac{p^*}{32\pi^2 m_a^2} \int |M_{fi}|^2 d\Omega$$

Where

$$p^* = \frac{1}{2m_a} \sqrt{[m_a^2 - (m_1 + m_2)^2][m_a^2 - (m_1 - m_2)^2]}$$

So in our case,

$$\begin{aligned}
p^* &= \frac{1}{2m_\tau} \sqrt{[m_\tau^2 - (m_\nu + m_\pi)^2][m_\tau^2 - (m_\nu - m_\pi)^2]} \\
&\approx \frac{1}{2m_\tau} \sqrt{[m_\tau^2 - m_\pi^2][m_\tau^2 - m_\tau^2]} \\
&= \frac{m_\tau^2 - m_\pi^2}{2m_\tau}
\end{aligned}$$

and

$$\begin{aligned}
\Gamma(\tau^- \rightarrow \pi^- \nu_\tau) &= \frac{p^*}{32\pi^2 m_\tau^2} \int |M_{fi}|^2 d\Omega \\
&= \frac{p^*}{32\pi^2 m_\tau^2} \int 4G_F^2 f_\pi^2 m_\tau^3 p^* \sin^2\left(\frac{\theta}{2}\right) d\Omega \\
&= \frac{G_F^2 f_\pi^2}{8\pi^2 m_\tau^2} (p^*)^2 m_\tau^3 \int \sin^2\left(\frac{\theta}{2}\right) d\Omega \\
&= \frac{G_F^2 f_\pi^2}{8\pi^2 m_\tau^2} \left(\frac{m_\tau^2 - m_\pi^2}{2m_\tau}\right)^2 m_\tau^3 \int \sin^2\left(\frac{\theta}{2}\right) d\Omega \\
&= \frac{G_F^2 f_\pi^2}{32\pi^2} m_\tau^3 \left(\frac{m_\tau^2 - m_\pi^2}{m_\tau^2}\right)^2 \int \sin^2\left(\frac{\theta}{2}\right) d\Omega \\
&= \frac{G_F^2 f_\pi^2}{32\pi^2} m_\tau^3 \left(\frac{m_\tau^2 - m_\pi^2}{m_\tau^2}\right)^2 (2\pi) \\
&= \frac{G_F^2 f_\pi^2}{16\pi} m_\tau^3 \left(\frac{m_\tau^2 - m_\pi^2}{m_\tau^2}\right)^2
\end{aligned}$$

e): Using the value of f_π obtained in the previous problem, find a numerical value for $\Gamma(\tau^- \rightarrow \pi^- \nu_\tau)$.

Well first let's do the previous problem: from the prediction of 11.25 and the measured value of the charged pion lifetime $\tau_\pi = 2.6033 \times 10^{-8} \text{ s}$ find f_π . Note charged pions decay almost always to muons + muon neutrinos.

$$11.25: \Gamma_\pi = \frac{1}{\tau_\pi} = \frac{G_F^2}{8\pi m_\pi^3} f_\pi^2 [m_\ell(m_\pi^2 - m_\ell^2)]^2$$

$$\begin{aligned}
\frac{1}{\tau_\pi} &= \frac{G_F^2}{8\pi m_\pi^3} f_\pi^2 [m_\mu(m_\pi^2 - m_\mu^2)]^2 \\
\sqrt{\frac{8\pi m_\pi^3}{\tau_\pi [m_\mu(m_\pi^2 - m_\mu^2)]^2 G_F^2}} &= f_\pi
\end{aligned}$$

using these:

- $m_\pi = 139.6 \text{ MeV}$
- $\tau_\pi = 2.6033 \times 10^{-8} \text{ s} = 3.9551808 \times 10^{16} \text{ GeV}^{-1}$
- $m_\mu = 105.7 \text{ MeV}$

- $G_F = 1.166\,38 \times 10^{-5} \text{ GeV}^{-2}$

we get $f_\pi = 0.128 \text{ GeV} \approx m_\pi$.

Then using $f_\pi = m_\pi$ and $m_\tau = 1.777 \text{ GeV}$, we find

$$\begin{aligned}\Gamma(\tau^- \rightarrow \pi^- \nu_\tau) &= \frac{G_F^2 m_\pi^2}{16\pi} m_\tau^3 \left(\frac{m_\tau^2 - m_\pi^2}{m_\tau^2} \right)^2 \\ &\approx 2.922 \times 10^{-13} \text{ GeV}\end{aligned}$$

f) Given that the lifetime of the τ^- lepton is measured to be $\tau_\tau = 2.906 \times 10^{-13} \text{ s}$, find an approximate value for the $\tau^- \rightarrow \pi^- \nu_\tau$ branching ratio.

Convert that time to a decay rate in GeV: $\tau_\tau = 2.906 \times 10^{-13} \text{ s} = 4.415 \times 10^{11} \text{ GeV}^{-1}$
 $\Rightarrow \Gamma_\tau = \frac{1}{\tau_\tau} = 2.265 \times 10^{-12} \text{ GeV}$

$$\begin{aligned}BR &= \frac{\Gamma(\tau^- \rightarrow \pi^- \nu_\tau)}{\Gamma_\tau} \\ &= \frac{2.922 \times 10^{-13} \text{ GeV}}{2.265 \times 10^{-12} \text{ GeV}} \\ &= 0.129\end{aligned}$$

3 Predict the ratio of $K^- \rightarrow e^- \bar{\nu}_e$ and $K^- \rightarrow \mu^- \bar{\nu}_\mu$ weak interaction decay rates and compare your answer to the measured value of $\frac{\Gamma(K^- \rightarrow e^- \bar{\nu}_e)}{\Gamma(K^- \rightarrow \mu^- \bar{\nu}_\mu)} = (2.488 \pm 0.012) \times 10^{-5}$.

Recall Thomson 11.25:

$$\Gamma(\pi^- \rightarrow \ell^- \bar{\nu}_\ell) = \frac{G_F^2}{8\pi m_\pi^3} f_\pi^2 [m_\ell(m_\pi^2 - m_\ell^2)]^2$$

This is for pions, but let's see if we can derive something similar for kaons.

Start the same way, in the CoM frame:

$$p_K = (m_K, 0, 0, 0), \quad p_\ell = p_3 = (E_\ell, 0, 0, p), \quad p_{\bar{\nu}} = p_4 = (p, 0, 0, -p)$$

(where p is the magnitude of the momentum for both the charge lepton and the antineutrino)

The weak leptonic current associated with the $\ell^- \bar{\nu}_\ell$ vertex is $j_\ell^\nu = \frac{g_W}{\sqrt{2}} \bar{u}(p_3) \frac{1}{2} \gamma^\nu (1 - \gamma^5) \nu(p_4)$

And again since the kaon is a bound $q\bar{q}$ state, the corresponding hadronic current cannot be expressed in terms of free particle Dirac spinors. However, the kaon current has to be a four-vector such that the four-vector scalar product with the leptonic current gives a Lorentz-invariant expression for the matrix element.

There might be a subtlety here since kaons don't have zero spin, but I'm pretty sure to some approximation we can use the same expression for pions, making a constant replacement and ending up with a similar expression as for pions but with the kaon mass:

$$\frac{\Gamma(K^- \rightarrow e^- \bar{\nu}_e)}{\Gamma(K^- \rightarrow \mu^- \bar{\nu}_\mu)} = \left[\frac{m_e(m_K^2 - m_e^2)}{m_\mu(m_K^2 - m_\mu^2)} \right]^2$$

Using $m_e = 0.511 \text{ MeV}$, $m_K = 494 \text{ MeV}$, $m_\mu = 106 \text{ MeV}$, we get

$$\frac{\Gamma(K^- \rightarrow e^- \bar{\nu}_e)}{\Gamma(K^- \rightarrow \mu^- \bar{\nu}_\mu)} \approx 2.55 \times 10^{-5}$$

The measured value:

$$\frac{\Gamma(K^- \rightarrow e^- \bar{\nu}_e)}{\Gamma(K^- \rightarrow \mu^- \bar{\nu}_\mu)} = (2.488 \pm 0.012) \times 10^{-5}$$

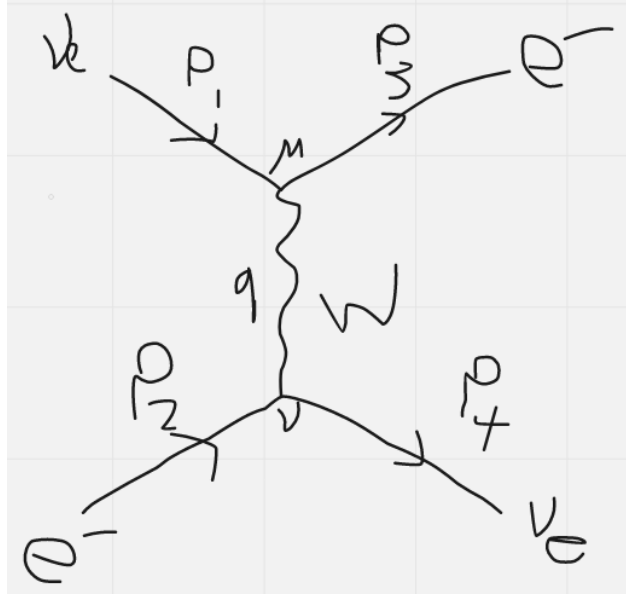
These are relatively close to each other – our calculation isn't within the uncertainty of the measured value, but it's the same order of magnitude at least, and we've made some approximations, especially with the spin thing.

4 Electron neutrino scattering cross section.

- a. Assuming that the process $\nu_e e^- \rightarrow e^- \nu_e$ only occurs by the weak charged-current interaction (ignore Z), show that $\sigma_{CC}^{\nu_e e^-} \approx \frac{2m_e E_\nu G_F^2}{\pi}$, where E_ν is in the lab frame where the struck e^- is at rest.

Remember in chapter 12.2 when we derived the charged-current cross section of neutrino scattering off a nucleus? Let's try to do something similar here.

We have a feynman diagram like the middle bit of Thomson figure 12.4, just with $d = \mu^-$, $u = \nu_\mu$ and $\mu = e$.



For a neutrino interacting with an electron at rest, the CoM energy squared is $s = (p_1 + p_2)^2 = (E_\nu + m_e)^2 - E_\nu^2 = 2m_e E_\nu + m_e^2$. For high-energy neutrinos we can ignore the m_e^2 term and just say $s = 2m_e E_\nu$.

Calculate the matrix element here, ignoring q^2 dependence as in 12.2.1:

- Incoming ν_e : $u(p_1)$
- Outgoing e^- : $\bar{u}(p_3)$
- W^- propagator: $\text{appx } \frac{-ig_{\mu\nu}}{m_W^2}$
- $e^- \rightarrow W^- \nu_e$ vertex 1: $\frac{-ig_w}{\sqrt{2}} \frac{1}{2} \gamma^\mu (1 - \gamma^5)$
- $e^- \rightarrow W^- \nu_e$ vertex 2: $\frac{-ig_w}{\sqrt{2}} \frac{1}{2} \gamma^\nu (1 - \gamma^5)$
- Incoming e^- : $u(p_2)$

- Outgoing ν_e : $\bar{u}(p_4)$

$$\begin{aligned}
 -iM_{fi} &= \left[\bar{u}(p_3) \frac{-ig_w}{\sqrt{2}} \frac{1}{2} \gamma^\mu (1 - \gamma^5) u(p_1) \right] \frac{-ig_{\mu\nu}}{m_W^2} \left[\bar{u}(p_4) \frac{-ig_w}{\sqrt{2}} \frac{1}{2} \gamma^\nu (1 - \gamma^5) u(p_2) \right] \\
 M_{fi} &= \frac{g_w^2}{2m_W^2} g_{\mu\nu} \left[\bar{u}(p_3) \gamma^\mu \frac{1}{2} (1 - \gamma^5) u(p_1) \right] \left[\bar{u}(p_4) \gamma^\nu \frac{1}{2} (1 - \gamma^5) u(p_2) \right]
 \end{aligned}$$

For high-energy neutrino scattering, let's approximate chiral states = helicity states, and from here the calculation of σ is exactly the same as in 12.2.1, yielding

$$\sigma_{CC}(\nu_e e^- \rightarrow \nu_e e^-) = \frac{G_F^2 s}{\pi}.$$

Using our $s = 2m_e E_\nu$, we find what the question wanted us to show:

$$\sigma_{CC}^{\nu_e e^-} \approx \frac{2m_e E_\nu G_F^2}{\pi}$$

- b. Using the above result, estimate the probability that a 10 MeV solar electron neutrino will undergo a charged-current weak interaction with an electron in the earth if it travels along a trajectory passing through the centre of the earth. Take the earth to be a sphere of radius $R = 6400$ km and uniform density $\rho = 5520$ kgm⁻³.

To get the probability of an interaction (which we know will be relatively small), we can use $\sigma = \frac{P}{n_e}$, where n_e is the surface area number density of electrons in our path, i.e. $n_e = N_e/A$.

Consider the cross section first:

$$\sigma_{CC}^{\nu_e e^-} = \frac{2m_e E_\nu G_F^2}{\pi}$$

$$G_F = 1.16638 \times 10^{-5} \text{ GeV}^{-2}$$

Mass volume density is ρ , get mass are density:

$$\begin{aligned}
 M/A &= \rho d \\
 &= \rho 2R
 \end{aligned}$$

Some fraction of the mass of the earth is due to electrons, say $\frac{1}{2000}$ since the earth is mostly composed of atoms with equal amounts of neutrons and protons, and $m_e \approx \frac{1}{1000}m_p$.

$$M_e/A = \frac{1}{1000}\rho R$$

The electron mass here can be written as $M_e = N_e m_e$,

$$\begin{aligned} N_e m_e / A &= \frac{1}{1000} \rho R \\ n_e m_e &= \frac{1}{1000} \rho R \\ n_e &= \frac{1}{m_e} \frac{1}{1000} \rho R \\ &= 3.878\,198\,7 \times 10^{37} \text{ m}^{-2} \end{aligned}$$

Use this to get the probability we're after:

$$\begin{aligned} P &= \sigma n_e \\ &= \frac{2m_e E_\nu G_F^2}{\pi} n_e \\ &= \frac{2E_\nu G_F^2}{\pi} \frac{1}{1000} \rho R \\ &= \frac{2(10 \text{ MeV})(1.166\,38 \times 10^{-5} \text{ GeV}^{-2})^2}{\pi} \frac{1}{1000} 5520 \text{ kg m}^{-3} 6400 \text{ km} \\ &= 6.68 \times 10^{-10} \end{aligned}$$

(There were some extra factors in there from implicit \hbar s and c s. In the spirit of natural units, I've left those out.)

5 T2K

The T2K experiment uses an off-axis ν_μ beam produced from $\pi^+ \rightarrow \mu^+ \nu_\mu$ decays. Consider the case where the pion has velocity β along the z -direction in the laboratory frame and a neutrino with energy E^* is produced at an angle θ^* with respect to the z' -axis in the π^+ rest frame.

- a. Show that the neutrino energy in the pion rest frame is $p^* = (m_\pi^2 - m_\mu^2)/2m_\pi$

We know that for any decay of the form $a \rightarrow 1 + 2$, we have

$$p^* = \frac{1}{2m_a} \sqrt{[m_a^2 - (m_1 + m_2)^2] [m_a^2 - (m_1 - m_2)^2]}$$

In our case, $\pi^+ \rightarrow \mu^+ \nu_\mu$

$$\begin{aligned} p^* &= \frac{1}{2m_\pi} \sqrt{[m_\pi^2 - (m_\mu + m_\nu)^2] [m_\pi^2 - (m_\nu - m_\mu)^2]} \\ &\approx \frac{1}{2m_\pi} \sqrt{[m_\pi^2 - (0 + m_\nu)^2] [m_\pi^2 - (0 - m_\mu)^2]} \\ &= \frac{m_\pi^2 - m_\mu^2}{2m_\pi} \end{aligned}$$

- b. Using a Lorentz transformation, show that the energy E and angle of production θ of the neutrino in the lab frame are $E = \gamma E^*(1 + \beta \cos(\theta^*))$, $E \cos(\theta) = \gamma E^*(\cos(\theta^*) + \beta)$, where $\gamma = E_\pi/m_\pi$.

In the rest frame, we have

$$\begin{aligned} p_\pi^* &= (m_\pi, 0, 0, 0) \\ p_\nu^* &= p^*(1, \sin(\theta^*), 0, \cos(\theta^*)) \\ p_\mu^* &= (m_\pi - p^*, -p^* \sin(\theta^*), 0, -p^* \cos(\theta^*)) \end{aligned}$$

And in the lab frame we'll have

$$\begin{aligned} p_\pi &= (E_\pi, 0, 0, m_\pi \beta) \\ p_\nu &= p(1, \sin(\theta), 0, \cos(\theta)) \\ p_\mu &= (E_\pi - p, -p \sin(\theta), 0, -p \cos(\theta)) \end{aligned}$$

First show $\gamma = \frac{E_\pi}{m_\pi}$ using the Lorentz transformations for momentum/energy (boost with $-\beta$ to get into the lab frame):

$$\begin{aligned} E_\pi &= \gamma(E_\pi^* + \beta p_{z,\pi}^*) \\ E_\pi &= \gamma((m_\pi) + \beta(0)) \\ E_\pi &= \gamma m_\pi \\ \gamma &= \frac{E_\pi}{m_\pi} \end{aligned}$$

Then on use the same transformations on the neutrino, and also use $p^* = E^*$ for the neutrino since it's essentially massless:

$$\begin{aligned} E &= \gamma(E^* + \beta p_{z,\nu}^*) \\ &= \gamma(E^* + \beta p^* \cos(\theta^*)) \\ &= \gamma(E^* + \beta E^* \cos(\theta^*)) \\ &= \gamma E^* (1 + \beta \cos(\theta^*)) \\ p_{z,\nu} &= \gamma(p_{z,\nu}^* + \beta E^*) \\ E \cos(\theta) &= \gamma(p^* \cos(\theta^*) + \beta E^*) \\ &= \gamma(E^* \cos(\theta^*) + \beta E^*) \\ &= \gamma E^* (\cos(\theta^*) + \beta) \end{aligned}$$

- c. Using the expressions for E^* and θ^* in terms of E and θ , show that $\gamma^2(1 - \beta \cos(\theta))(1 + \beta \cos(\theta^*)) = 1$

Derive a similar expression to the ones above for the unstarred E, θ , flipping the sign on β for the inverse transform:

$$\begin{aligned} E^* &= \gamma(E - \beta p_{z,\nu}) \\ &= \gamma(E - \beta p \cos(\theta)) \\ &= \gamma(E - \beta E \cos(\theta)) \\ &= \gamma E (1 - \beta \cos(\theta)) \end{aligned}$$

Use this:

$$\begin{aligned}
\gamma^2(1 - \beta \cos(\theta))(1 + \beta \cos(\theta^*)) &= \frac{1}{E} \gamma E (1 - \beta \cos(\theta)) \frac{1}{E^*} \gamma E^* (1 + \beta \cos(\theta^*)) \\
&= \frac{1}{E} E^* \frac{1}{E^*} E \\
&= \frac{EE^*}{EE^*} \\
&= 1
\end{aligned}$$

- d. Assuming that the pions have a flat energy spectrum in the range $1 - 5 \text{ GeV}$, sketch the form of the resulting neutrino energy spectrum at the T2K far detector (SuperK), which is off-axis at $\theta = 2.5^\circ$. Given that SuperK is 295km from the beam, explain why this angle was chosen.

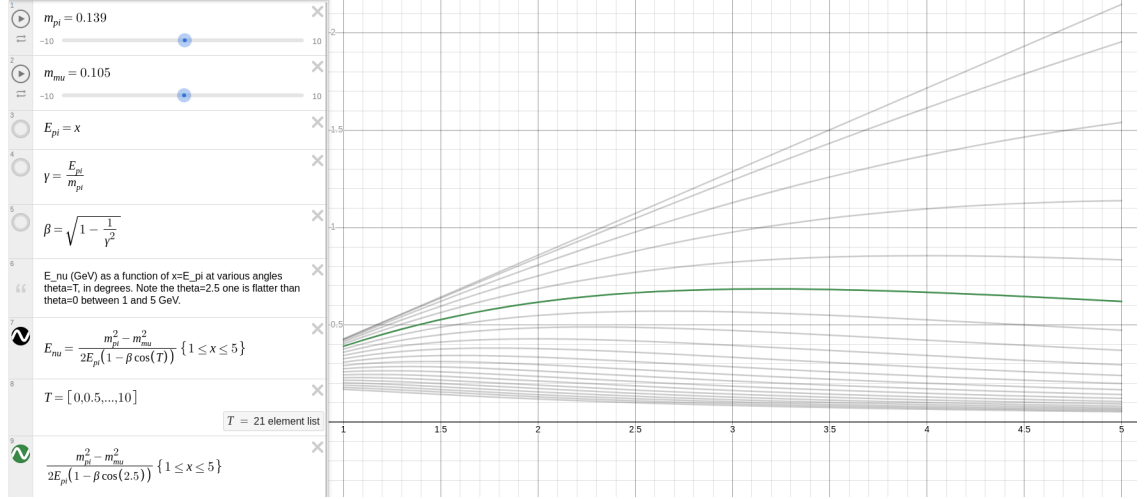
First, get an expression for β :

$$\gamma = \frac{1}{\sqrt{1 - \beta^2}} \implies \beta = \sqrt{1 - \frac{1}{\gamma^2}}$$

And use our other knowledge from above to get an expression for E_ν in terms of other variables we have:

$$\begin{aligned}
E_\nu^* &= \gamma E_\nu (1 - \beta \cos(\theta)) \\
p^* &= \gamma E_\nu (1 - \beta \cos(\theta)) \\
\frac{m_\pi^2 - m_\mu^2}{2m_\pi} &= \frac{E_\pi}{m_\pi} E_\nu (1 - \beta \cos(\theta)) \\
\frac{m_\pi^2 - m_\mu^2}{2E_\pi (1 - \beta \cos(\theta))} &= E_\nu \\
\frac{m_\pi^2 - m_\mu^2}{2E_\pi \left(1 - \sqrt{1 - \frac{m_\pi^2}{E_\pi^2}} \cos(\theta)\right)} &= E_\nu
\end{aligned}$$

This is something we can plot for different values of θ :



E_{ν} (GeV) as a function of E_{π} at various angles θ , in degrees.

Note the $\theta = 2.5$ line is flatter than $\theta = 0$ between 1 and 5 GeV.

I suppose at a large distance like 295km, it's a good idea to have a more energy-uniform beam of neutrinos, so you can tell which ones came from the experiment vs external sources and also so you only need to design detectors to measure within a smaller range of energies, hence the choice of 2.5° .