PHYS 509C Assignment 1

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Code for this assignment is here:

https://github.com/callum-mccracken/PHYS-509C-A1

It's in a bit of a strange format since I make it write the LaTeX file that I use for making the document you're reading, but here are the highlights:

- Open the file with numpy.loadtxt()
- Get the mean with numpy.mean()
- Get the standard deviation with numpy.std()
- Get the correlation coefficient with numpy.corrcoef()
- Get the skew with scipy.stats.skew()
- Use scipy.stats.chi2.pdf() for the chi-squared PDF
- Integrate using scipy.integrate.quad()

- 1 fakedata.out contains 200 observations of three random variables: X, Y, and Z (each variable in its own column, listed in that order). Calculate the following for this data:
- **A**. The mean values of X, Y, and Z.

$$\overline{X} = 49.85, \overline{Y} = -1.56, \overline{Z} = -19.38.$$

B. The standard deviations for all thre variables.

$$\sigma_X = 12.75, \sigma_Y = 13.63, \sigma_Z = 11.06.$$

C. The three correlation coefficients between the three variables.

$$C_{X,Y} = 0.30, C_{X,Z} = 0.72, C_{Y,Z} = -0.30.$$

D. The skew for X, Y, and Z.

$$Skew(X) = -0.10, Skew(Y) = 0.05, Skew(Z) = -0.31.$$

- Numerically calculate the probability that a number drawn from a χ^2 distribution with n=5 degrees of freedom will be larger than $\chi^2=5$. Do the same for n=10. Do not use a lookup table or a pre-existing function to evaluate the answer, but calculate it for yourself as if you had just discovered the χ^2 distribution for the first time.
 - $P(\chi_5^2 < 5) = 0.42$.
 - $P(\chi_{10}^2 < 5) = 0.89.$

3 Three independent random numbers X_1, X_2, X_3 are drawn from uniform distributions with means of 0 and variances of 1/3. Let Z be the sum of these three numbers. Derive the normalized probability distribution for Z.

A uniform distribution's PDF is

$$P(x) = \begin{cases} \frac{1}{b-a}, & x \in [a,b] \\ 0, & x \notin [a,b] \end{cases}$$

with mean $\frac{a+b}{2}$ and variance $\frac{(b-a)^2}{12}$.

Here we have a mean of $0 \implies b = -a \implies \frac{(b-a)^2}{12} = \frac{a^2}{3}$. And a variance of $\frac{1}{3} \implies a = -1, b = 1$.

$$P(x) = \begin{cases} \frac{1}{2}, & x \in [-1, 1] \\ 0, & x \notin [-1, 1] \end{cases}$$

Or in terms of heaviside step functions:

$$P(x) = \frac{1}{2}(H(x+1) - H(x-1))$$

If we have two independent variables of this type, we have:

$$Y = X_1 + X_2$$

$$P(y) = \int_{x_1, x_2 \mid x_1 + x_2 = y} P(x_1, x_2)$$

It's not super obvious to me what to do here, after some googling it seems this can be done with CDFs:

The CDF of Y is found using all possibile combinations of $x_1 + x_2 < y$. At

this point note $y \in [-2, 2]$.

$$F(y) = \iint_{x_1 + x_2 < y} P(x_1, x_2) dx_1 dx_2$$
$$= \int_{x_1 = -\infty}^{\infty} \int_{x_2 = -\infty}^{y - x_1} P(x_1, x_2) dx_1 dx_2$$

And if we take the derivative we'll get P(y):

$$P(y) = \frac{d}{dy} \int_{x_1 = -\infty}^{\infty} \int_{x_2 = -\infty}^{y - x_1} P(x_1, x_2) dx_1 dx_2$$

$$= \int_{x_1 = -\infty}^{\infty} \frac{d}{dy} \int_{x_2 = -\infty}^{y - x_1} P(x_1, x_2) dx_1 dx_2$$

$$= \int_{x_1 = -\infty}^{\infty} P(x_1, y - x_1) dx_1$$

Since we had two independent variables,

$$P(y) = \int_{x_1 = -\infty}^{\infty} P(x_1)P(y - x_1)dx_1$$

$$= \int_{x_1 = -\infty}^{\infty} \frac{1}{2} (H(x_1 + 1) - H(x_1 - 1)) \frac{1}{2} (H(y - x_1 + 1) - H(y - x_1 - 1)) dx_1$$

$$= \frac{1}{4} \int_{x_1 = -\infty}^{\infty} H(x_1 + 1)H(y - x_1 + 1) - H(x_1 + 1)H(y - x_1 - 1)$$

$$- H(x_1 - 1)H(y - x_1 + 1) + H(x_1 - 1)H(y - x_1 - 1) dx_1$$

Consider the products of steps we have:

- $H(x_1+1)H(y-x_1+1)$ To be non-zero: $x_1+1>0$ and $y-x_1+1>0$.
- $H(x_1+1)H(y-x_1-1)$ To be non-zero: $x_1+1>0$ and $y-x_1-1>0$.

- $H(x_1 1)H(y x_1 + 1)$ To be non-zero: $x_1 - 1 > 0$ and $y - x_1 + 1 > 0$.
- $H(x_1 1)H(y x_1 1)$ To be non-zero: $x_1 - 1 > 0$ and $y - x_1 - 1 > 0$.

So we have points of interest at $x_1 = -1, y - 1, y + 1, 1$. How these relate to each other depends on y.

Consider if the conditions above can be met simultaneously for $y \in [-2, 2]$, i.e. whether the products will be zero.

- $H(x_1+1)H(y-x_1+1)$ can be non-zero for $y \in [-2,2]$
- $H(x_1+1)H(y-x_1-1)$ can be non-zero for $y \in [0,2]$
- $H(x_1-1)H(y-x_1+1)$ can be non-zero for $y \in [0,2]$
- $H(x_1-1)H(y-x_1-1)$ is always zero for $y \in [-2,2]$

So for $y \in [-2, 0]$:

$$P(y) = \frac{1}{4} \int_{x_1 = -1}^{y+1} 1 - 0 - 0 + 0 dx_1$$
$$= \frac{1}{4} (y+2)$$

And for $y \in [0, 2]$:

$$P(y) = \frac{1}{4} \int_{x_1 = y - 1}^{1} 1 - 1 - 1 + 0 dx_1$$
$$= \frac{1}{4} (-y)$$

All together,

$$P(y) = \frac{1}{4}(y+2)H(y+2) - \frac{1}{2}yH(y) + (\frac{1}{4}y - \frac{1}{2})H(y-2)$$

Then take another convolution to get P(z) for $Z = Y + X_3$

$$\begin{split} P(z) &= \int_{y=-\infty}^{\infty} P_y(y) P_{x_3}(z-y) dy \\ &= \int_{y=-\infty}^{\infty} \left(\frac{1}{4} (y+2) H(y+2) - \frac{1}{2} y H(y) + (\frac{1}{4} y - \frac{1}{2}) H(y-2) \right) \\ &\qquad \times \left(\frac{1}{2} (H(z-y+1) - H(z-y-1)) \right) dy \\ &= \int_{y=-\infty}^{\infty} \left(\frac{1}{4} (y+2) H(y+2) - \frac{1}{2} y H(y) + (\frac{1}{4} y - \frac{1}{2}) H(y-2) \right) \frac{1}{2} H(z-y+1) \\ &\qquad - \left(\frac{1}{4} (y+2) H(y+2) - \frac{1}{2} y H(y) + (\frac{1}{4} y - \frac{1}{2}) H(y-2) \right) \frac{1}{2} H(z-y-1) dy \\ &= \int_{y=-\infty}^{\infty} \frac{1}{8} (y+2) H(y+2) H(z-y+1) - \frac{1}{4} y H(y) H(z-y+1) \\ &\qquad + \frac{1}{8} (y-2) H(y-2) H(z-y+1) - \frac{1}{8} (y+2) H(y+2) H(z-y-1) \\ &\qquad + \frac{1}{4} y H(y) H(z-y-1) - \frac{1}{8} (y-2) H(y-2) H(z-y-1) dy \end{split}$$

• H(y+2)H(z-y+1) can be non-zero for $z \in [-3,3]$ For $z \in [-3,1]$:

$$\int_{y=-\infty}^{\infty} \frac{1}{8} (y+2)H(y+2)H(z-y+1)dy$$

$$= \int_{y=-2}^{z+1} \frac{1}{8} (y+2)dy$$

$$= \frac{z^2 + 6z + 9}{16}$$

For $z \in [1, 3]$:

$$\int_{y=-\infty}^{\infty} \frac{1}{8} (y+2)H(y+2)H(z-y+1)dy$$

$$= \int_{y=-2}^{2} \frac{1}{8} (y+2)dy$$

$$= \left[\frac{1}{4} y^2 + \frac{1}{4} y \right]_{-2}^{2}$$

$$= \frac{1}{4} (2)^2 + \frac{1}{4} (2) - \frac{1}{4} (-2)^2 - \frac{1}{4} (-2)$$

$$= 1$$

• H(y)H(z-y+1) can be non-zero for $z \in [-1,3]$ For $z \in [-1,1]$

$$\int_{y=-\infty}^{\infty} -\frac{1}{4} y H(y) H(z-y+1) dy$$

$$= \int_{y=0}^{z+1} -\frac{1}{4} y dy$$

$$= -\frac{z^2 + 2z + 1}{8}$$

For $z \in [1, 3]$

$$\int_{y=-\infty}^{\infty} -\frac{1}{4}yH(y)H(z-y+1)dy$$

$$= \int_{y=0}^{2} -\frac{1}{4}ydy$$

$$= -\frac{1}{2}$$

• H(y-2)H(z-y+1) is always zero within the possible range of y.

• H(y+2)H(z-y-1) can be non-zero for $z \in [-1,3]$ For $z \in [-1,3]$

$$\int_{y=-\infty}^{\infty} -\frac{1}{8}(y+2)H(y+2)H(z-y-1)dy$$

$$= \int_{y=-2}^{z-1} -\frac{1}{8}(y+2)dy$$

$$= -\frac{z^2+2z+1}{16}$$

• H(y)H(z-y-1) can be non-zero for $z \in [1,3]$ For $z \in [1,3]$

$$\int_{y=-\infty}^{\infty} +\frac{1}{4}yH(y)H(z-y-1)dy$$

$$= \int_{y=0}^{z-1} -\frac{1}{4}ydy$$

$$= -\frac{z^2 - 2z + 1}{8}$$

• H(y-2)H(z-y-1) is always zero for $y \in [-2,2]$.

Let's put this together in sections:

• For $z \in [-3, -1]$:

$$P(z) = \frac{z^2 + 6z + 9}{16}$$

• For $z \in [-1, 1]$:

$$P(z) = \frac{z^2 + 6z + 9}{16} - \frac{z^2 + 2z + 1}{8} - \frac{z^2 + 2z + 1}{16}$$
$$= -\frac{z^2 - 3}{8}$$

• For $z \in [1, 3]$:

$$P(z) = 1 - \frac{1}{2} + 0 - \frac{z^2 + 2z + 1}{16} + \frac{z^2 - 2z + 1}{8}$$
$$= \frac{z^2 - 6z + 9}{16}$$

• zero elsewhere.

So all together:

$$P(z) = \frac{z^2 + 6z + 9}{16} (H(z+3) - H(z+1))$$
$$-\frac{z^2 - 3}{8} (H(z+1) - H(z-1))$$
$$+\frac{z^2 - 6z + 9}{16} (H(z-1) - H(z-3))$$

I see online there's a simpler version of this that's more general for higher numbers of uniform variables. Was there a better way to approach this? Seems like this approach is valid though, and the function we have is normalized.

4 Suppose that two random variables X_1 and X_2 have a continuous joint distribution for which the joint PDF is as follows: $f(x_1, x_2) = 4x_1x_2$ for $0 < x_1 < 1$ and $0 < x_2 < 1$, = 0 otherwise. Now consider the change of variables $Y_1 = X_1/X_2, Y_2 = X_1X_2$, and let $g(y_1, y_2)$ be the joint PDF of these two variables. Sketch the region in the y_1, y_2 plane for which g is non-zero, and calculate $g(y_1, y_2)$.

To sketch the region, first notice the possible ranges of the variables.

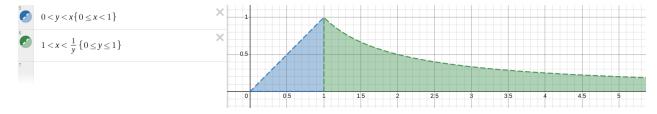
 y_1 can take any value between 0 and infinity. y_2 has a lower bound of zero and a global upper bound of 1.

But consider minimum and maximum values of y_2 for a given y_1 .

If $y_1 \leq 1$, our maximal value will be found by taking $x_2 = 1$ (well arbitrarily close to 1), which means $y_1 = x_1$ which in turn means $y_2 = y_1$ (again in the same arbitrarily close way).

On the other hand if $y_1 > 1$, we can find the value by taking $x_1 = 1 \implies y_2 = x_2 \implies y_2 = \frac{1}{y_1}$.

A sketch of the region where $g \neq 0$ is as follows:



To find $g(y_1, y_2)$ use the Jacobian:

$$g(y_1, y_2) = f(x_1, x_2) \begin{vmatrix} \frac{\partial x_1}{\partial y_1} & \frac{\partial x_1}{\partial y_2} \\ \frac{\partial x_2}{\partial y_1} & \frac{\partial x_2}{\partial y_2} \end{vmatrix}$$

To find these, we'll need $x_1(y_1, y_2), x_2(y_1, y_2)$:

$$y_1 = \frac{x_1}{x_2}$$

$$y_2 = x_1 x_2 \implies x_2 = \frac{y_2}{x_1}$$

$$y_1 = \frac{x_1}{\frac{y_2}{x_1}}$$

$$\implies x_1^2 = y_1 y_2$$

$$x_1 = \sqrt{y_1 y_2}$$

$$x_2 = \sqrt{\frac{y_2}{y_1}}$$

$$g(y_1, y_2) = 4x_1 x_2 \begin{vmatrix} \frac{1}{2} \sqrt{\frac{y_2}{y_1}} & \frac{1}{2} \sqrt{\frac{y_1}{y_2}} \\ -\frac{1}{2} \sqrt{\frac{y_2}{y_1}} \frac{1}{y_1} & \frac{1}{2} \sqrt{\frac{1}{y_1 y_2}} \end{vmatrix}$$

$$= 4y_2 \left| \frac{1}{2} \sqrt{\frac{y_2}{y_1}} \frac{1}{2} \sqrt{\frac{1}{y_1 y_2}} + \frac{1}{2} \sqrt{\frac{y_1}{y_2}} \frac{1}{2} \sqrt{\frac{y_2}{y_1}} \frac{1}{y_1} \right|$$

$$= 4y_2 \left(\frac{1}{4y_1} + \frac{1}{4y_1} \right)$$

$$= \frac{2y_2}{y_1}$$

- 5 Suppose that galactic supernovae obey Poissonian statistics. The mean number of supernovae per century is 1/3.
 - What is the most likely date for the next supernova? Poissonian statistics means we have a pdf of the form $P(k) = \frac{(\lambda t)^k e^{-\lambda t}}{k!}$, where k is the number of supernova observed in time t (measured in centuries), and $\lambda = 1/3$ is the mean number of supernovae in a century. To find the probability that a supernova happens at a particular time, consider the probability that no supernovae occur in time T.

$$P(k=0;T) = e^{-\lambda T}$$

.

So the probability of having a supernova by time t is given by

$$P(1st \text{ supernova within time } T) = 1 - e^{-\lambda T}$$

That is, the probability that the first supernova occurs before T.

$$P(t_s \le T) = 1 - e^{-\lambda T}$$

This is a cumulative probability function, differentiate to get the distribution for $t_s = T$.

$$P(t_s = T) = \lambda e^{-\lambda T}$$

This function is monotonically decreasing, so the most likely date is today, Sept 22 2022.

- What is the probability distribution for the length of the interval between now and the next galactic supernova?
 - It's what we had above (an exponential), the probability of the next supernova happening a time T away from now is given by:

$$P(T) = \lambda e^{-\lambda T}$$

Consider an infinite series of random variables X_i , where each variable is generated from its predecessor according to $X_i = aX_{i-1} + B_i$. Here a is a constant and B_i is a Gaussian random variable with mean m and standard deviation s. If all of the X_i are identically distributed with mean μ and standard deviation σ , then what constraints does this place on a, m, and s? What condition will result in the X_i also being independent from each other? In the case that they are identically distributed but not necessarily independent, derive a formula for the correlation coefficient between X_i and X_{i-j} .

First recall a few things:

- the definition of correlation coefficient: $\rho_{A,B} = \frac{\text{Cov}(A,B)}{\sigma_A \sigma_B}$
- How does scaling affect the mean? $\overline{aX} = a\overline{X}$
- How does scaling affect Variance?

$$\operatorname{Var}(aX) = \overline{(aX)^2} - (\overline{aX})^2 = a^2 \left(\overline{X^2} - (\overline{X})^2\right) = a^2 \operatorname{Var}(X)$$

- How does scaling affect Covariance? $Cov(aX, Y) = \overline{aXY} - \overline{aX}\overline{Y} = a(\overline{XY} - \overline{X}\overline{Y}) = aCov(X, Y)$
- For independent Gaussians, their sum is also a Gaussian, with mean $\mu_{A+B} = \mu_A + \mu_B$ and variance $\sigma_{A+B}^2 = \sigma_A^2 + \sigma_B^2$.

And find a relationship between X_i and X_{i-j} :

$$X_{i} = aX_{i-1} + B_{i}$$

$$= a(aX_{i-2} + B_{i-1}) + B_{i}$$

$$= a^{2}X_{i-2} + aB_{i-1} + B_{i}$$

$$= a^{3}X_{i-3} + a^{2}B_{i-2} + aB_{i-1} + B_{i}$$

$$= a^{4}X_{i-4} + a^{3}B_{i-3} + a^{2}B_{i-2} + a^{1}B_{i-1} + a^{0}B_{i-0}$$

$$\vdots$$

$$X_{i} = a^{j}X_{i-j} + \sum_{k=0}^{j-1} a^{k}B_{i-k}$$

• How are a, m, s constrained?

For the sum and variance of infinitely many B_i to be not-infinite for non-zero m, s, we need $a \in (-1, 1)$.

From the linearity of means, $\mu = a\mu + m$.

And since X_{i-1} , B_i are independent, $\sigma^2 = a^2\sigma^2 + s^2$.

- What's the condition such that the X_i are independent from each other? Well if a = 0 then $X_i = B_i$, just a Gaussian, and I think we can assume all the B_i are independent even though the question doesn't specifically say so.
- Find $\rho_{X_i,X_{i-j}}$ if X_i,X_{i-j} are not independent. Consider the following:

$$Cov((A + B), C) = \overline{(A + B)C} - \overline{A + BC}$$

$$= \overline{AC} + \overline{BC} - (\overline{A} + \overline{B})\overline{C}$$

$$= \overline{AC} - \overline{A}\overline{C} + \overline{BC} - \overline{B}\overline{C}$$

$$= Cov(A, C) + Cov(B, C)$$

If we apply this to our expression for X_i, X_{i-j}

$$Cov(X_{i}, X_{i-j}) = Cov(a^{j}X_{i-j} + \sum_{k=0}^{j-1} a^{k}B_{i-k}, X_{i-j})$$

$$= Cov(a^{j}X_{i-j}, X_{i-j}) + Cov(\sum_{k=0}^{j-1} a^{k}B_{i-k}, X_{i-j})$$

$$= a_{j} Var(X_{i-j}) + Cov(\sum_{k=0}^{j-1} a^{k}B_{i-k}, X_{i-j})$$

$$= a_{j}\sigma^{2} + Cov(\sum_{k=0}^{j-1} a^{k}B_{i-k}, X_{i-j})$$

The B_{i-k} are independent of X_{i-j} (since k < j and X_{i-j} only depends on Bs with a lower index), so the last part vanishes.

$$Cov(X_i, X_{i-j}) = a^j \sigma^2$$

$$\rho_{X_i, X_{i-j}} = \frac{\operatorname{Cov}(X_i, X_{i-j})}{\sigma_{X_i} \sigma_{X_{i-j}}}$$
$$= \frac{\sigma^2 a^j}{\sigma^2}$$
$$= a^j$$