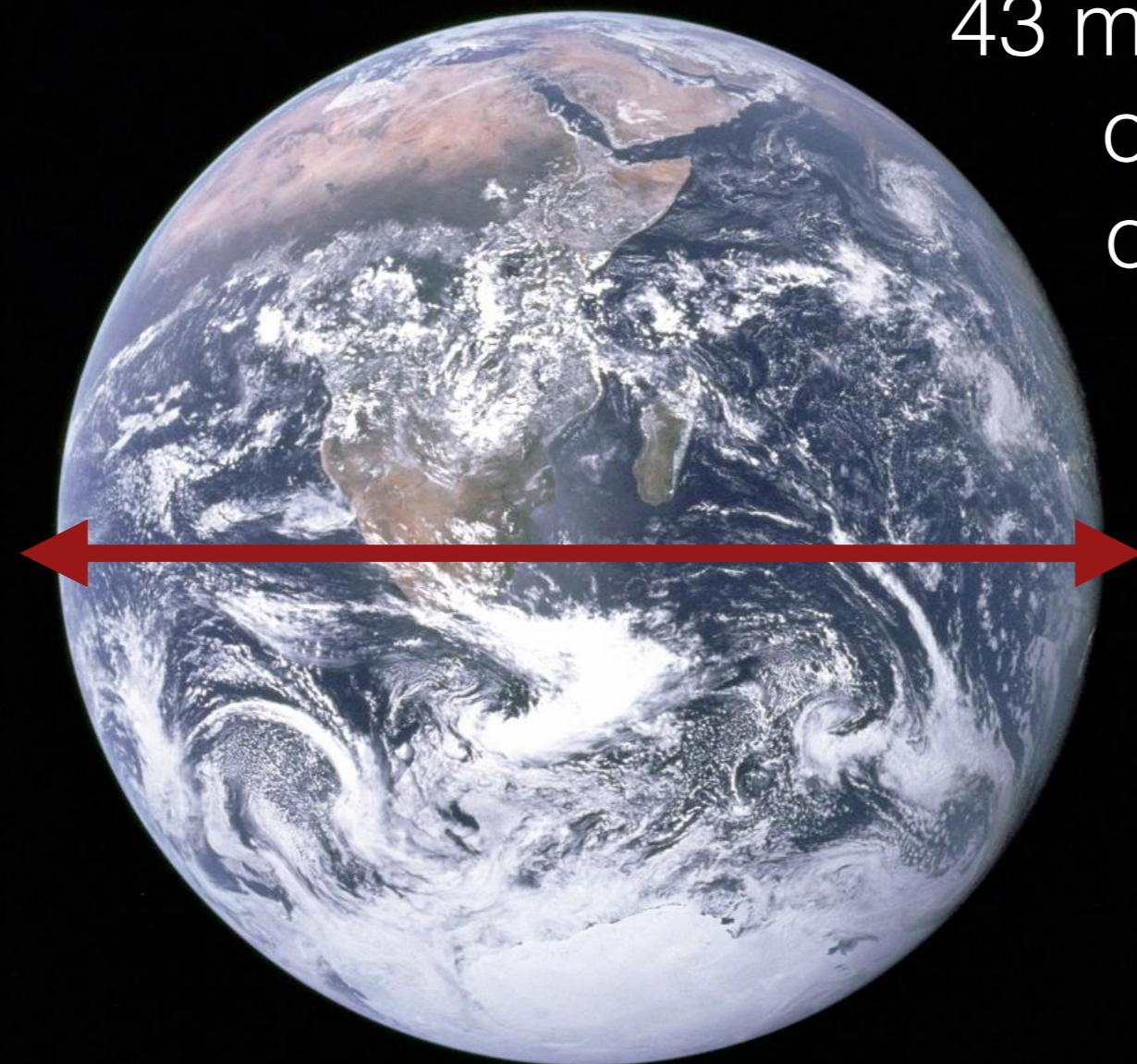


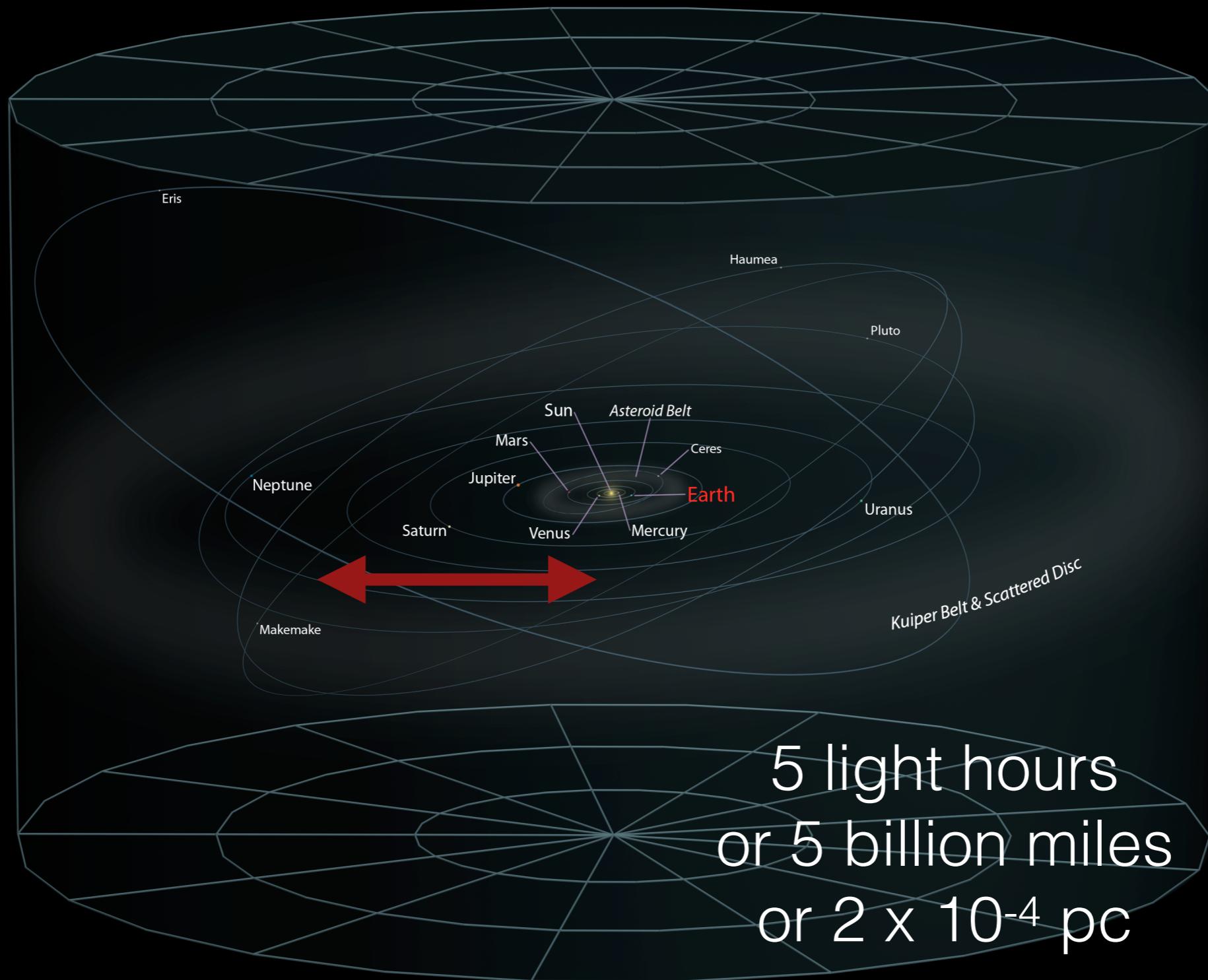
How big is our  
Universe?

# The Earth

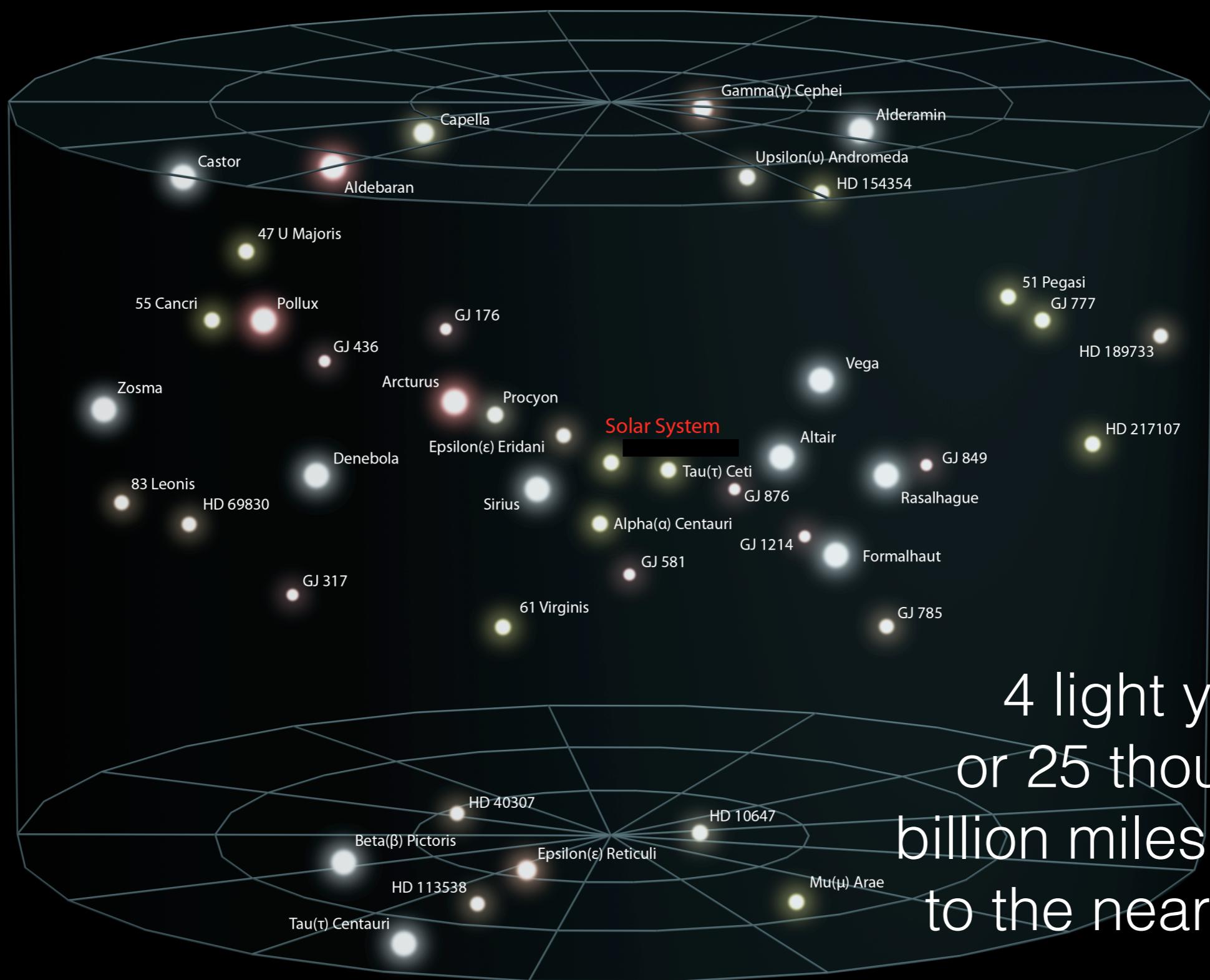


43 milli-light seconds  
or 8000 miles  
or  $4 \times 10^{-7}$  pc

# The Solar System

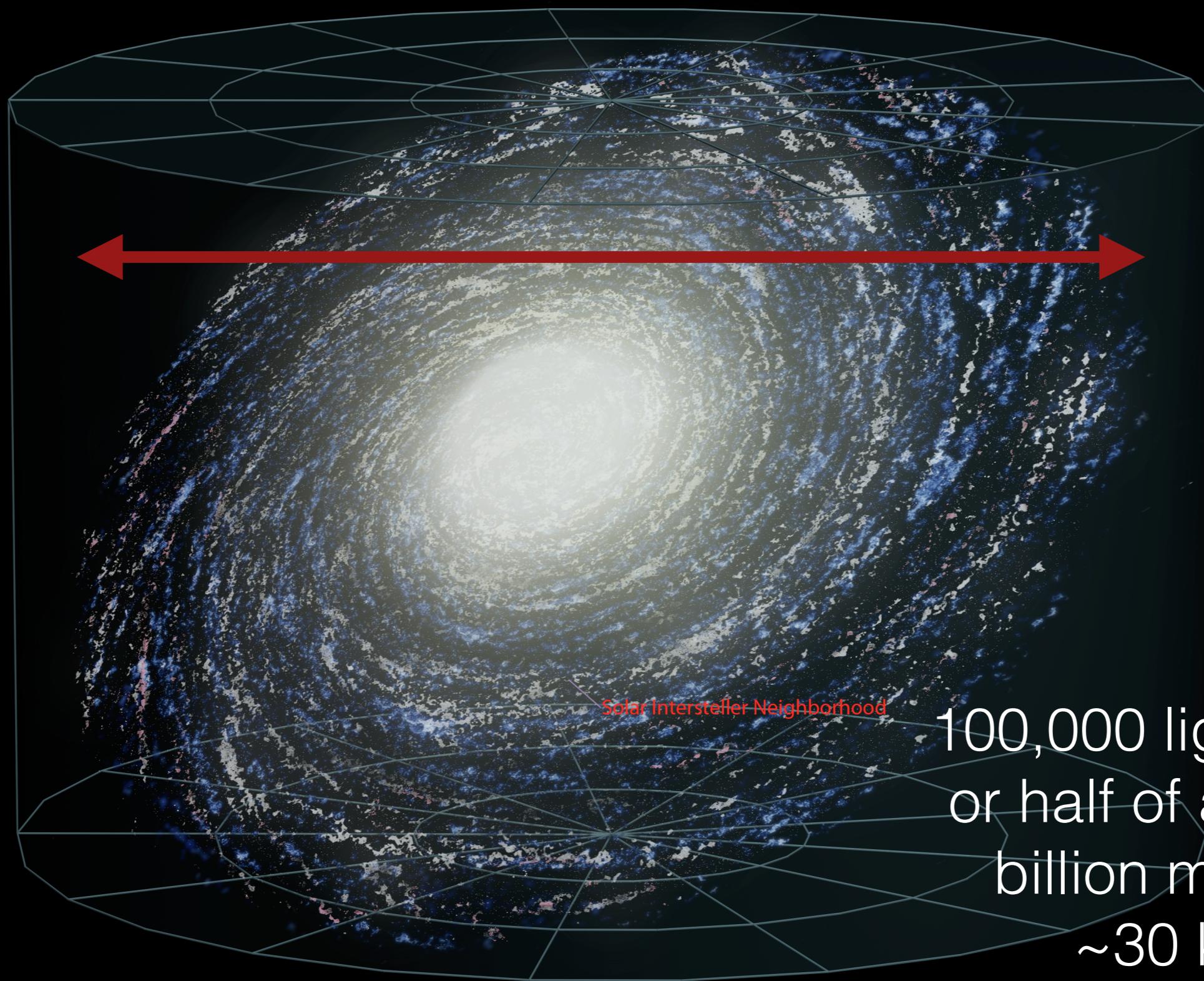


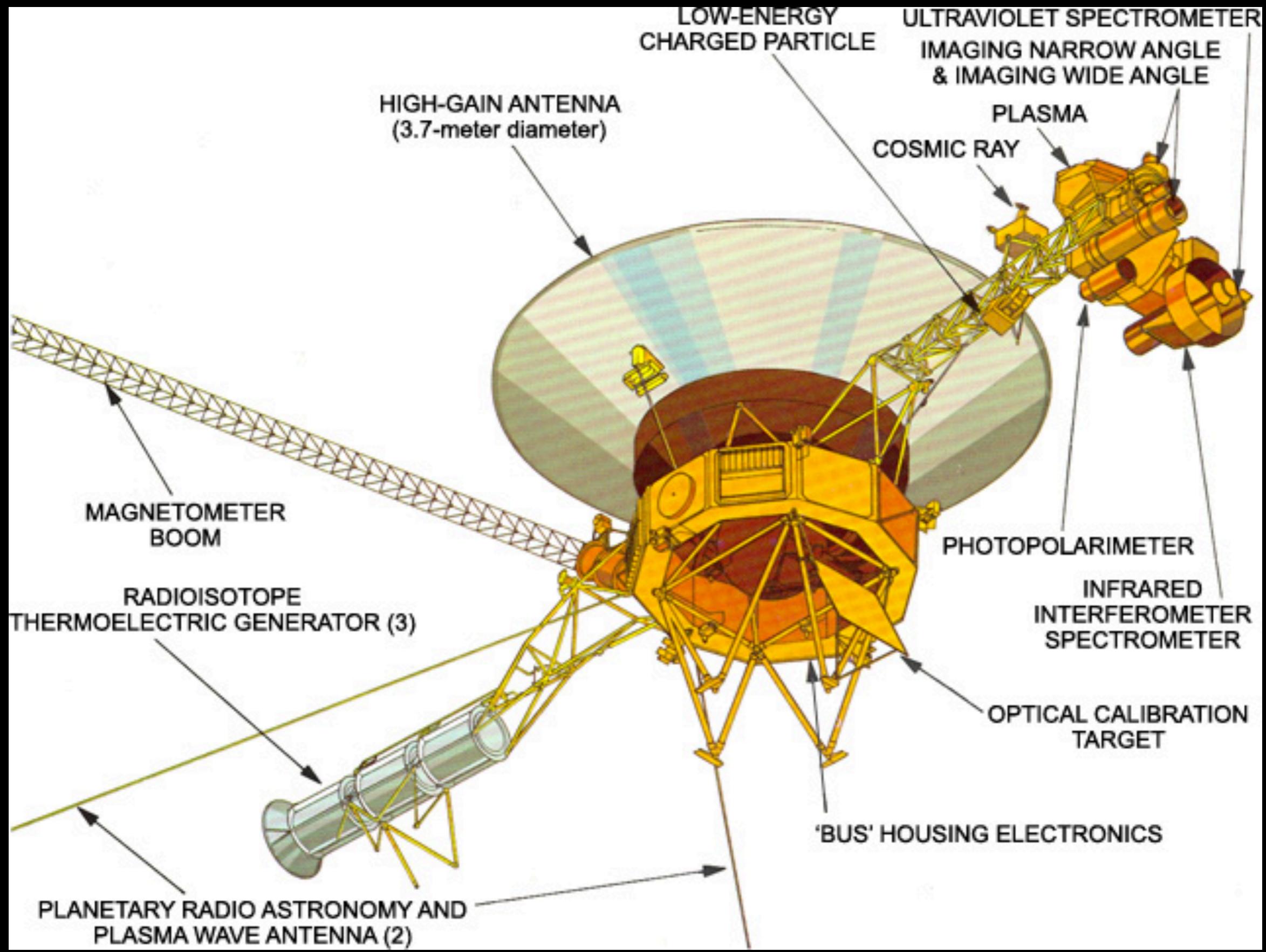
# Nearby Stars

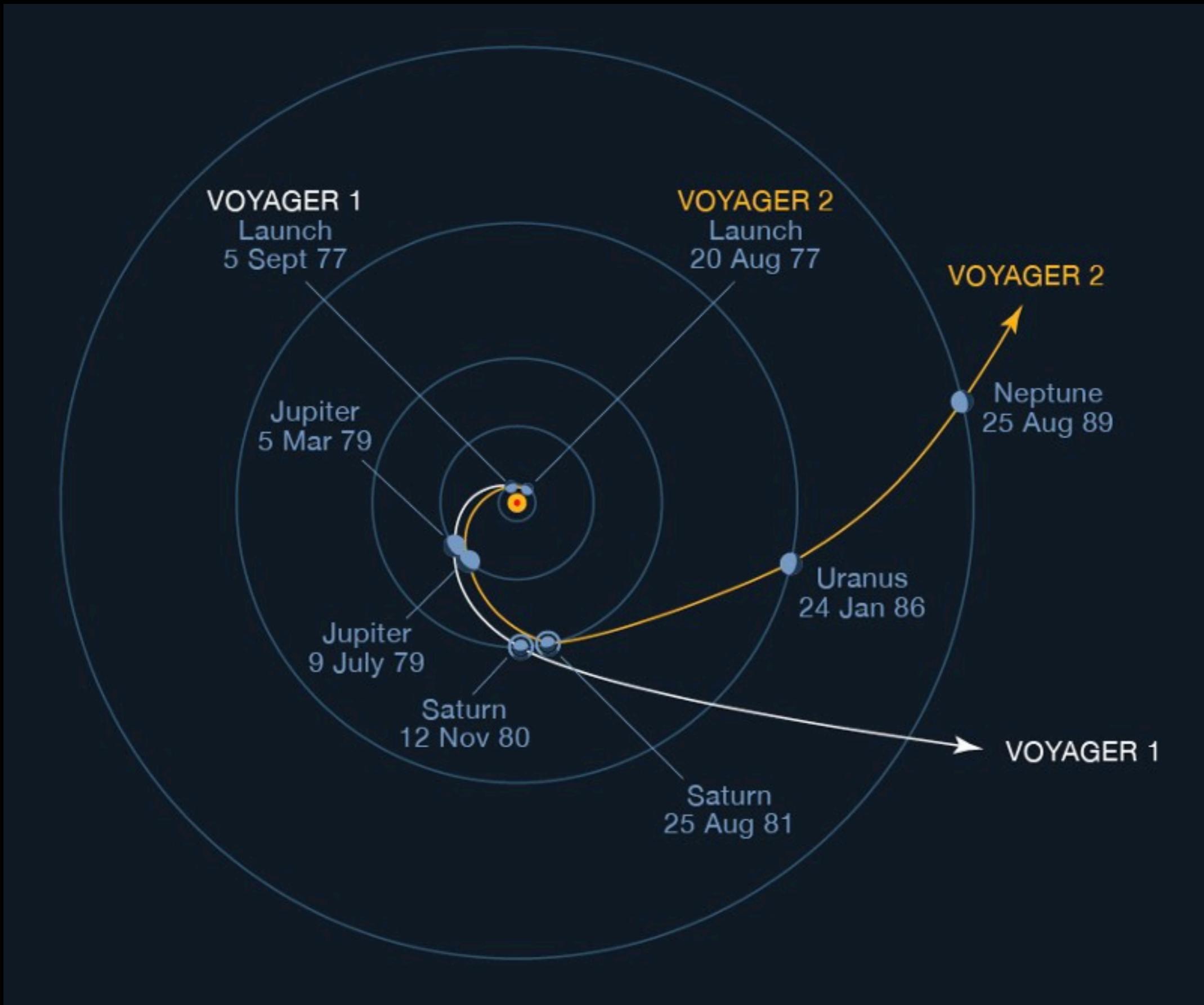


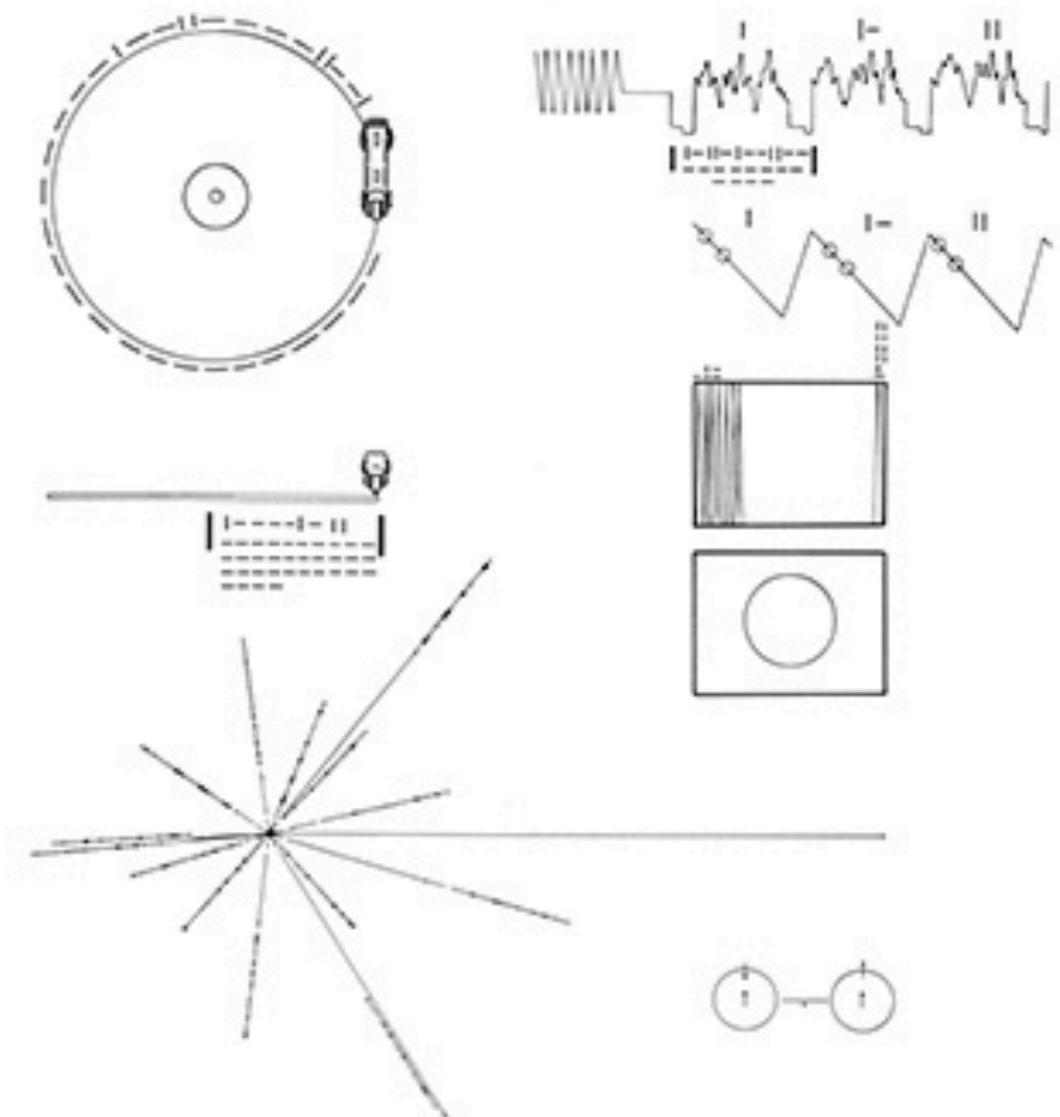
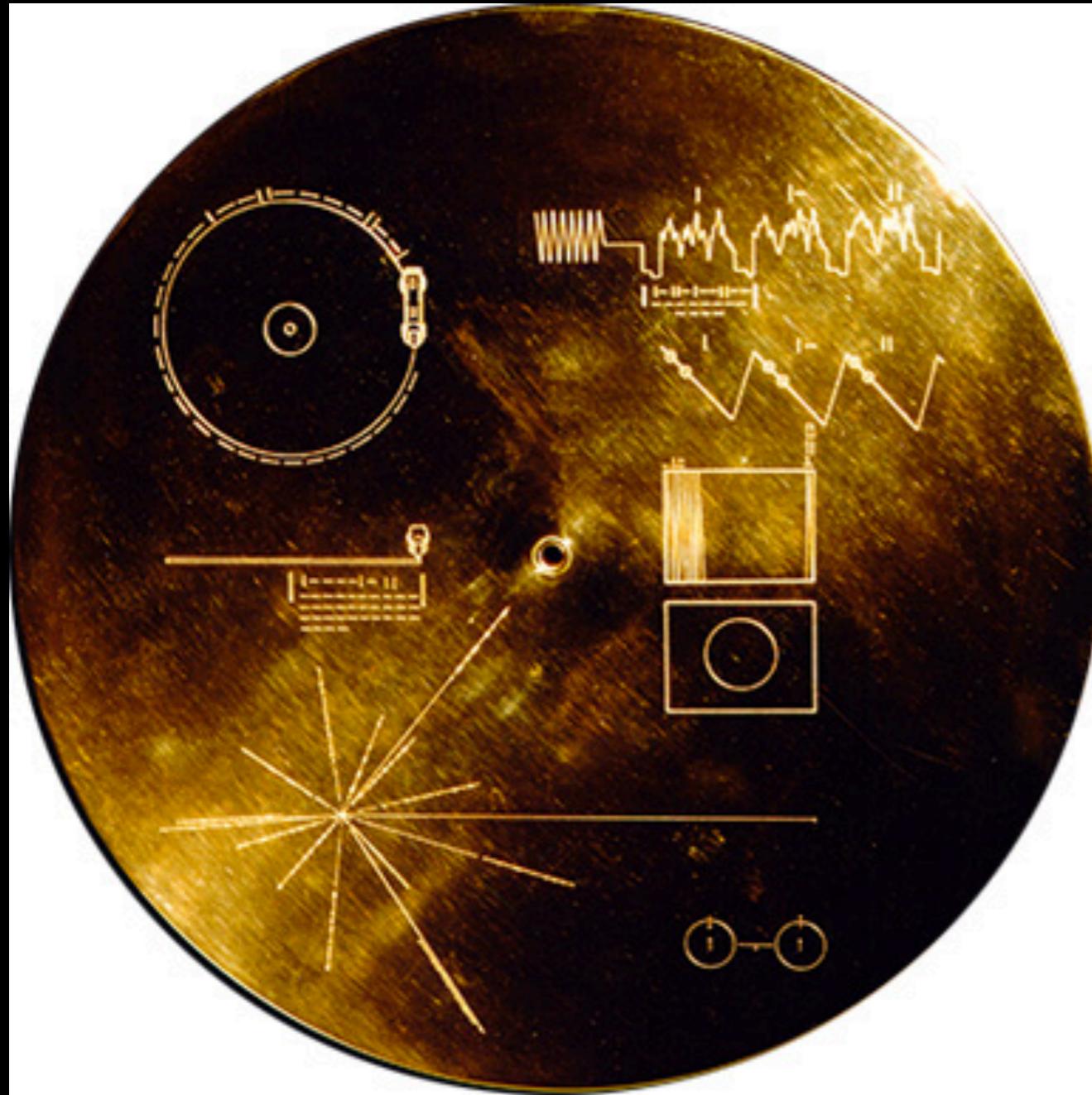
4 light years  
or 25 thousand,  
billion miles or  $\sim 1\text{pc}$   
to the nearest star

# The Milky Way

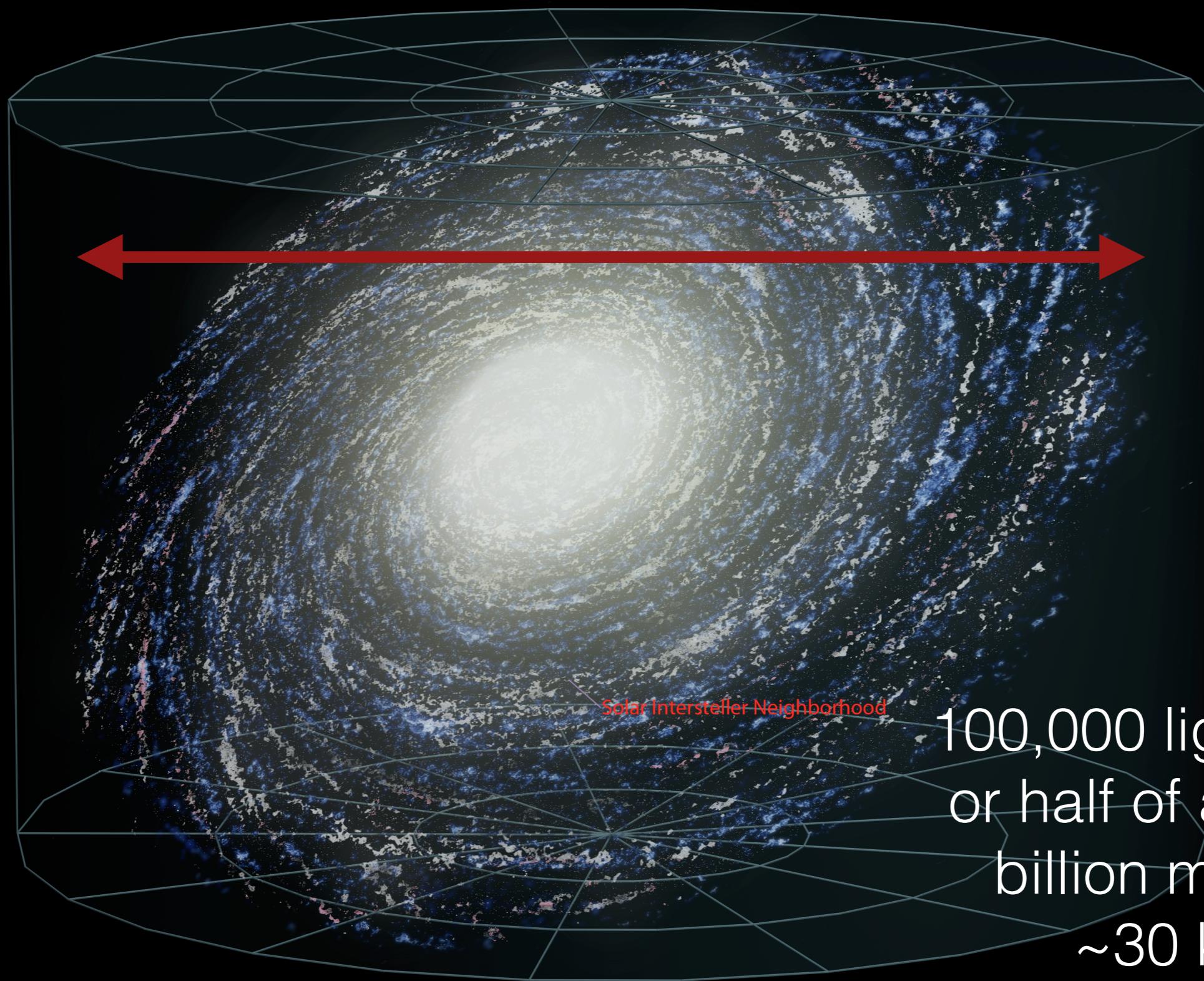




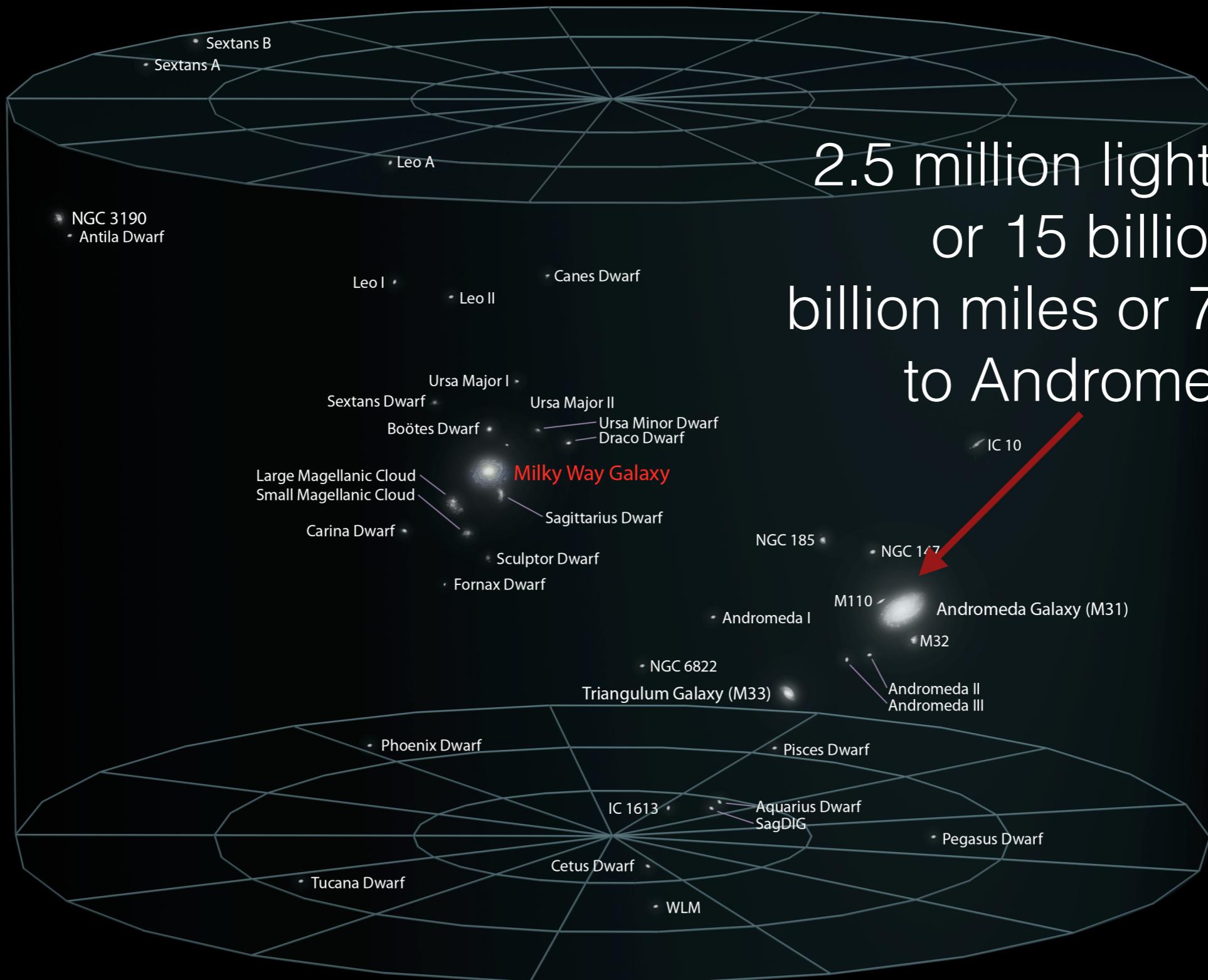




# The Milky Way

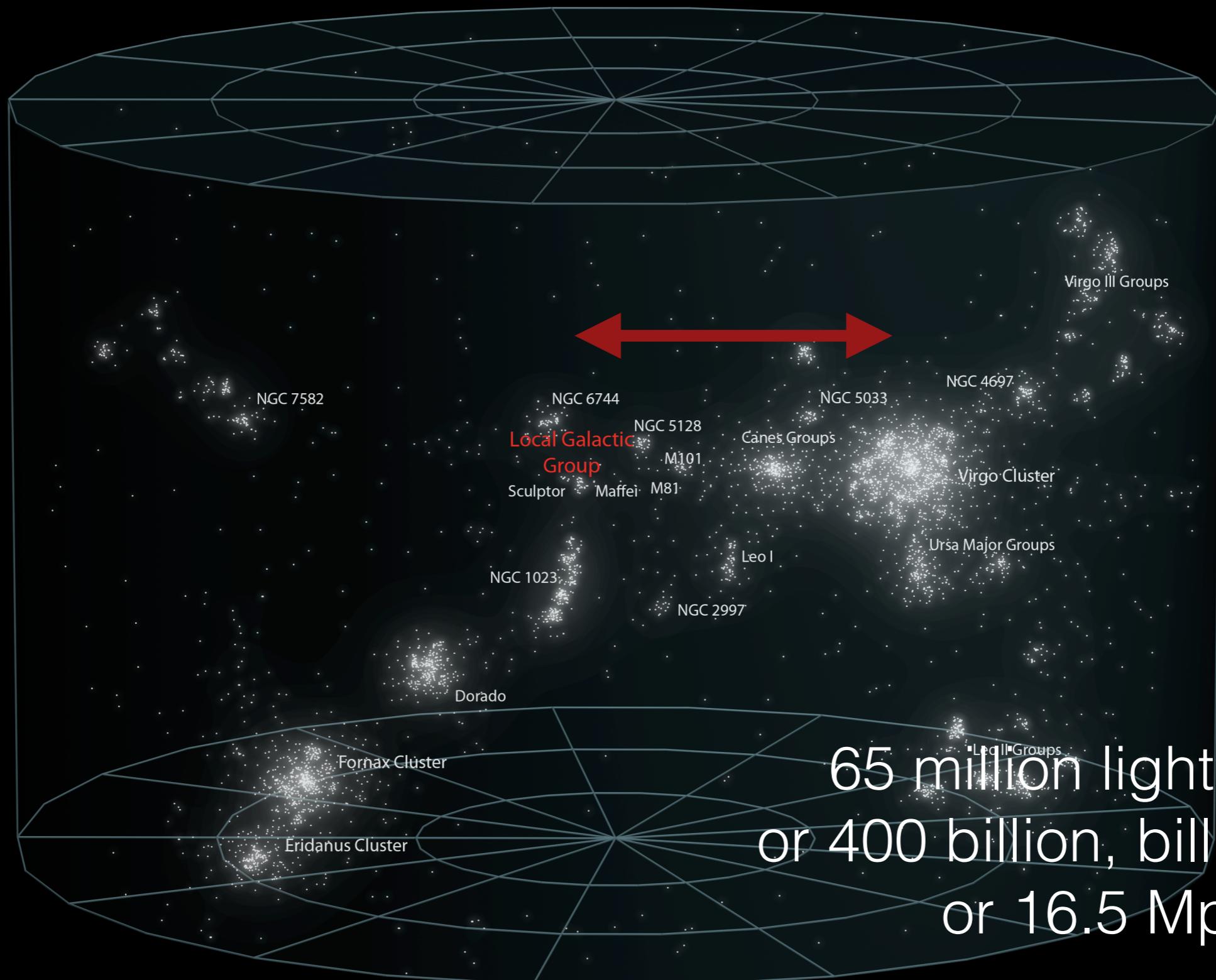


# The Local Group

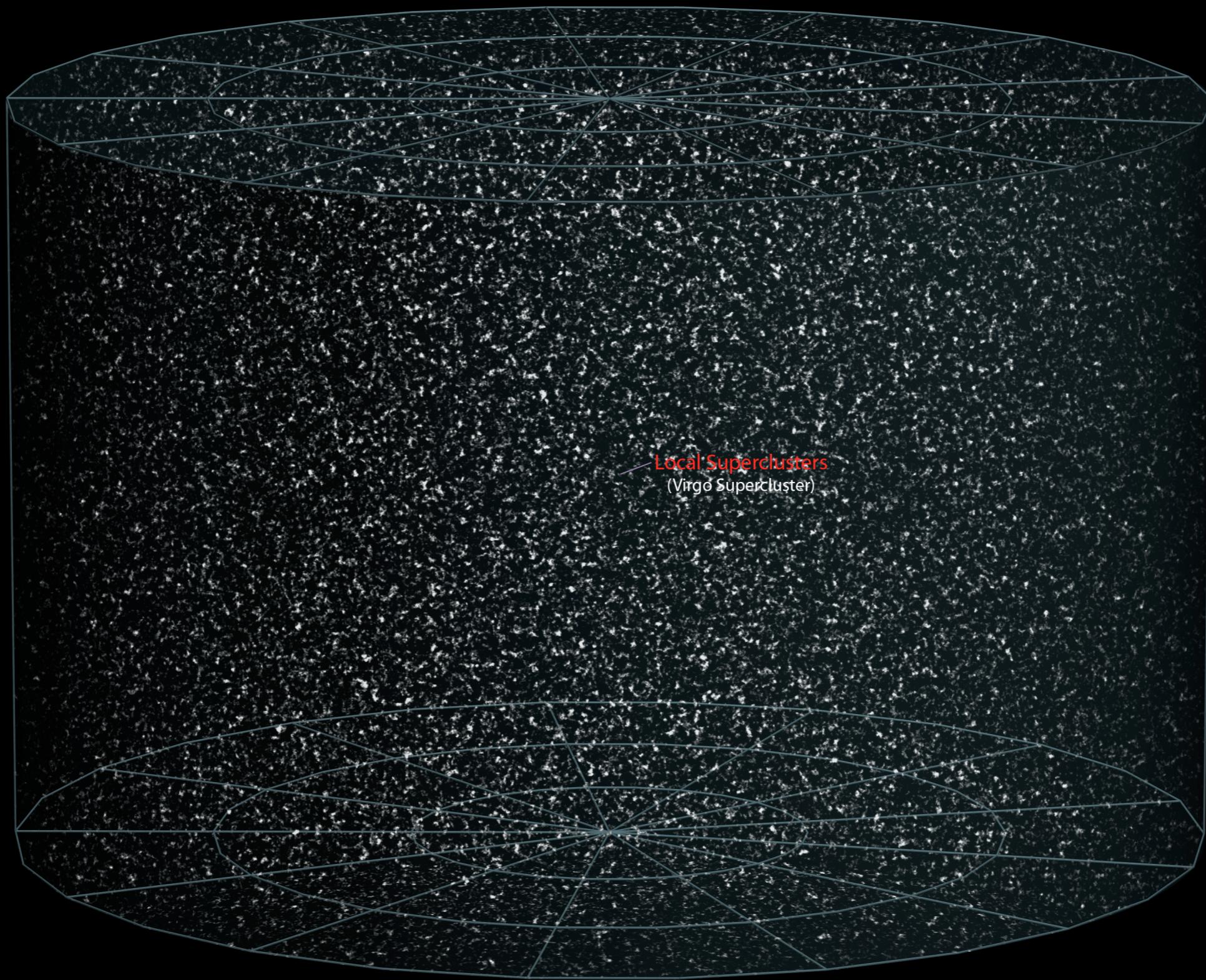


2.5 million light years  
or 15 billion,  
billion miles or 780 kpc  
to Andromeda

# Virgo Supercluster



# We are just one of many







MONGABAY.COM



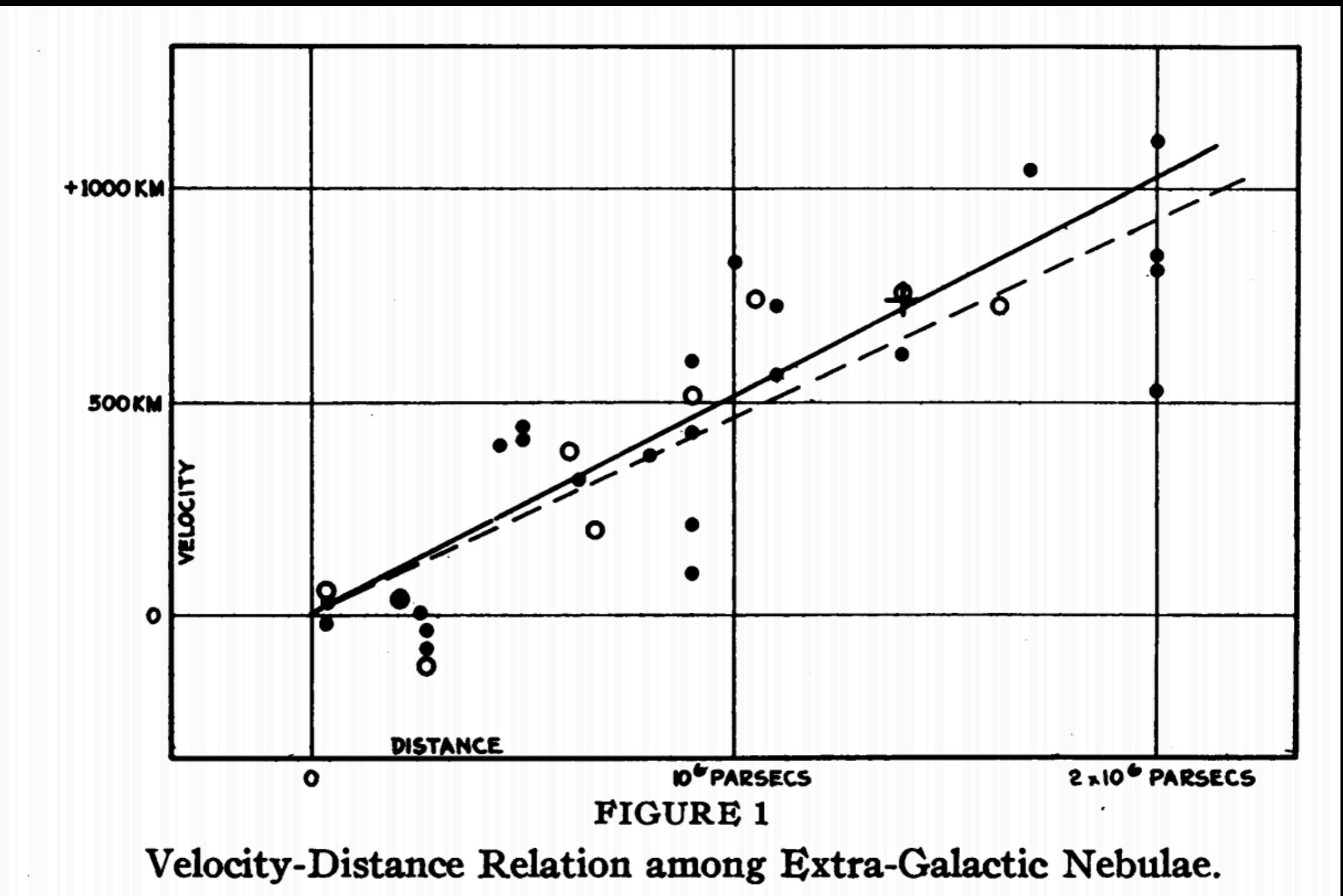
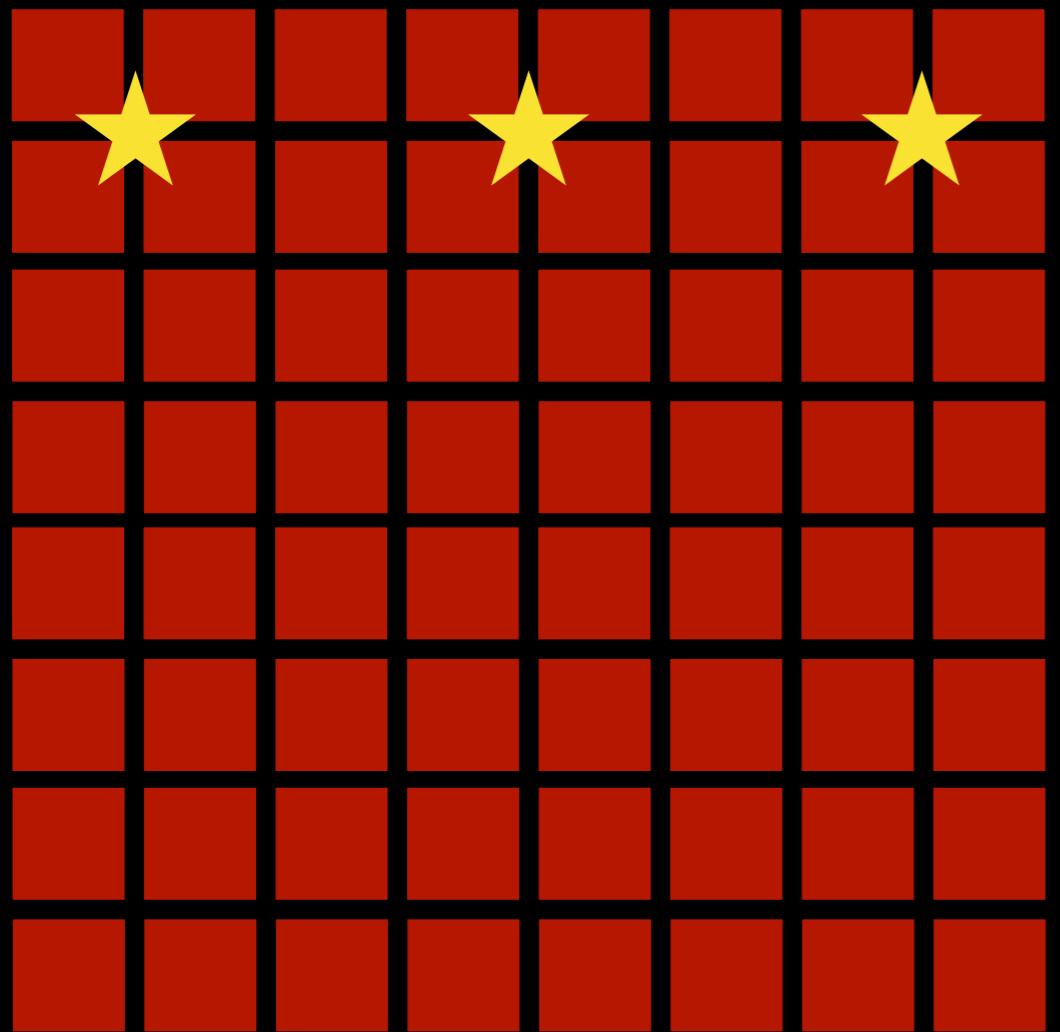
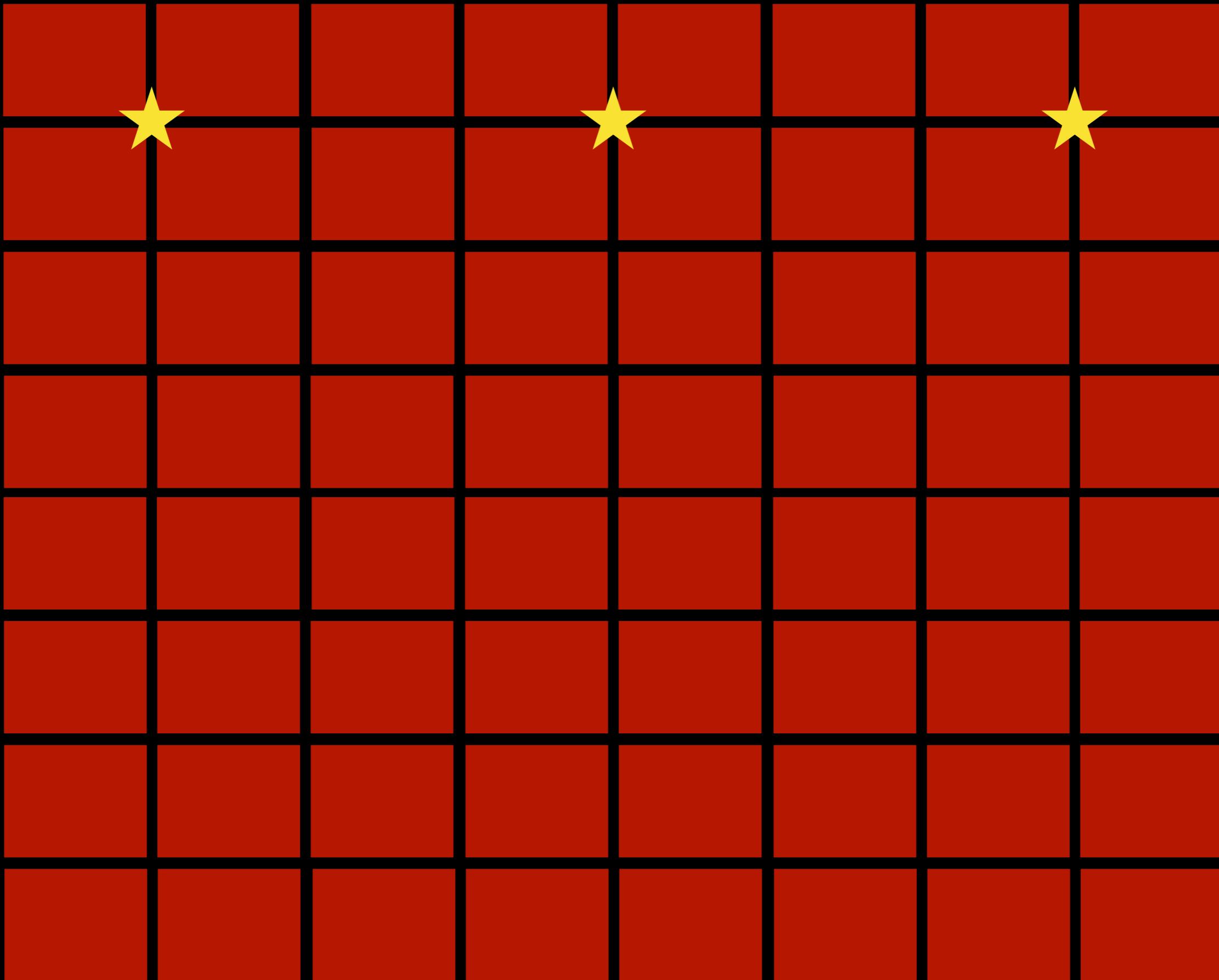


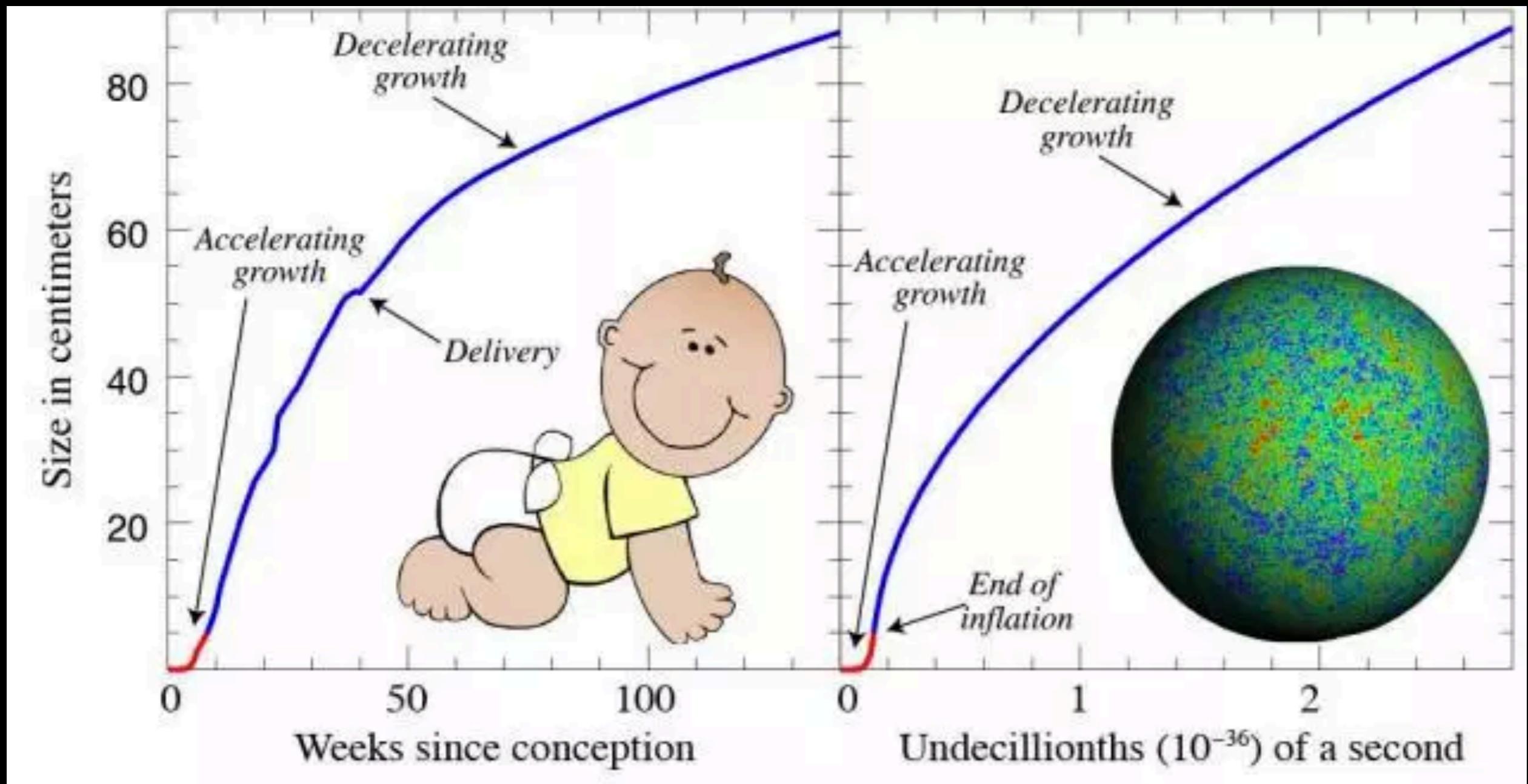
FIGURE 1

Velocity-Distance Relation among Extra-Galactic Nebulae.

Hubble (1929)



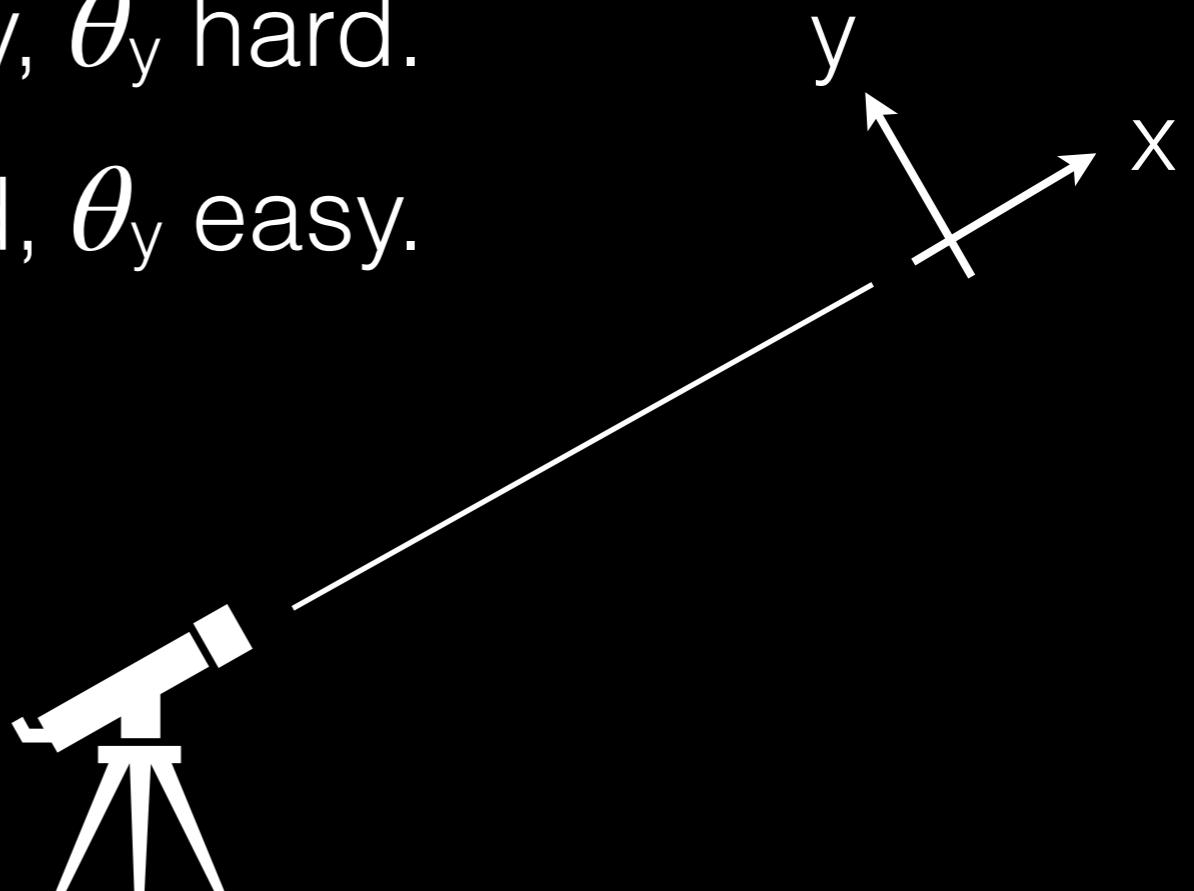




*Our Mathematical Universe* by Max Tegmark

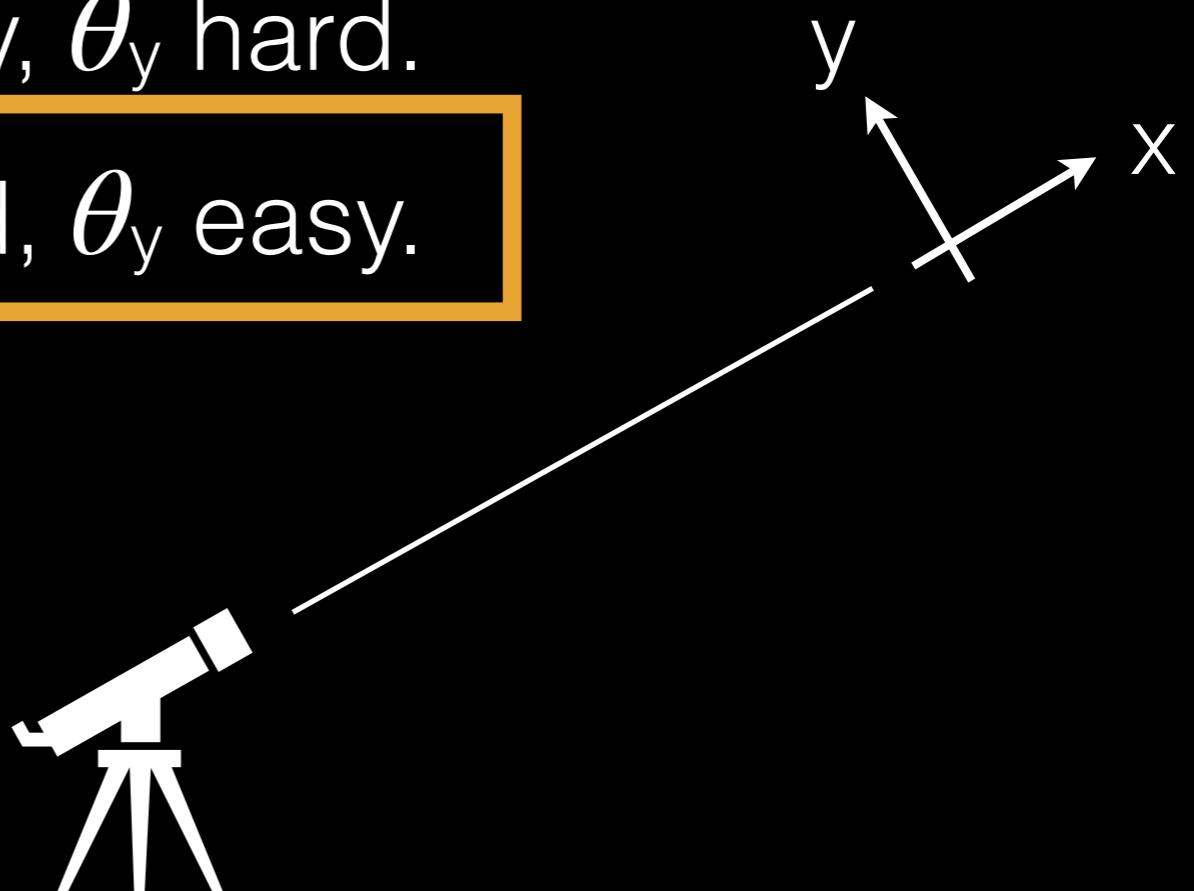
Which of components of a distant galaxy's position and velocity are easy/hard to observe?

- A.  $d_x$  easy,  $\theta_y$  hard;  $v_x$  easy,  $\theta_y$  hard.
- B.  $d_x$  easy,  $\theta_y$  hard;  $v_x$  hard,  $\theta_y$  easy.
- C.  $d_x$  hard,  $\theta_y$  easy;  $v_x$  easy,  $\theta_y$  hard.
- D.  $d_x$  hard,  $\theta_y$  easy;  $v_x$  hard,  $\theta_y$  easy.

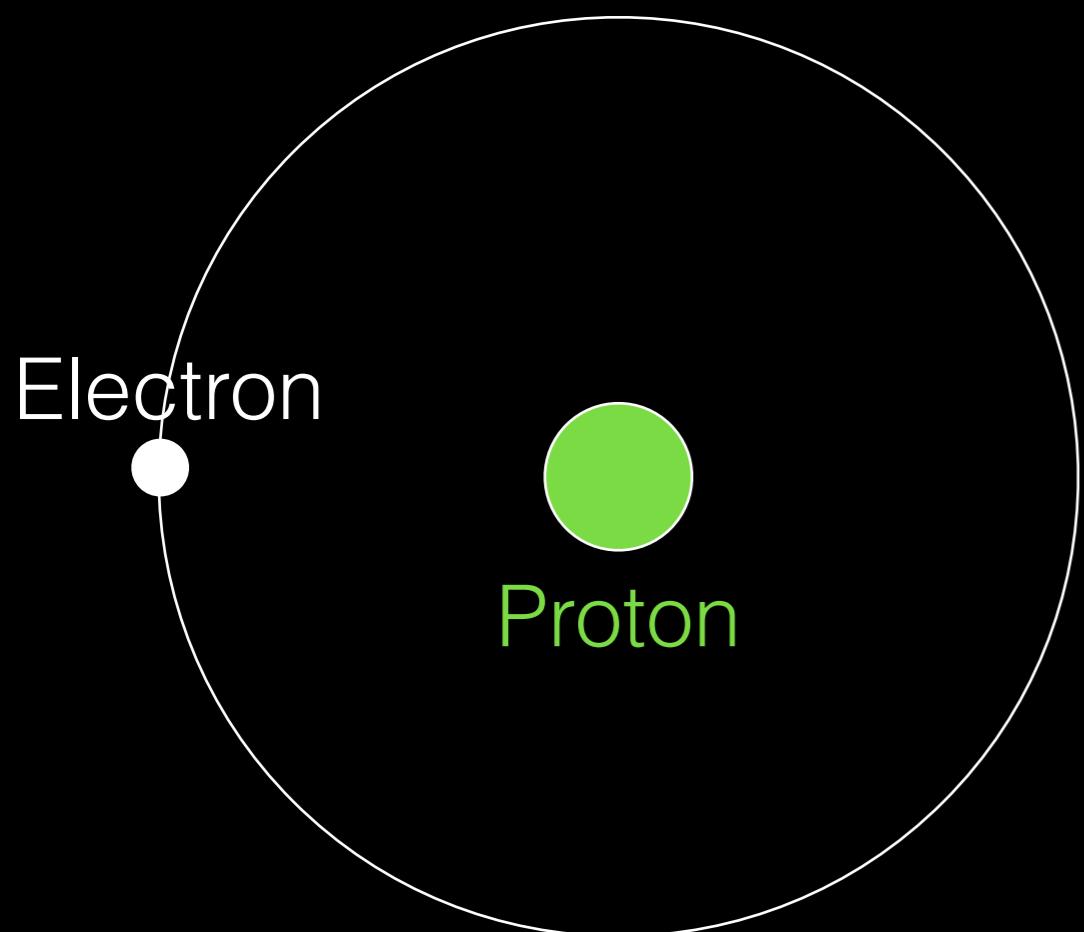


Which of components of a distant galaxy's position and velocity are easy/hard to observe?

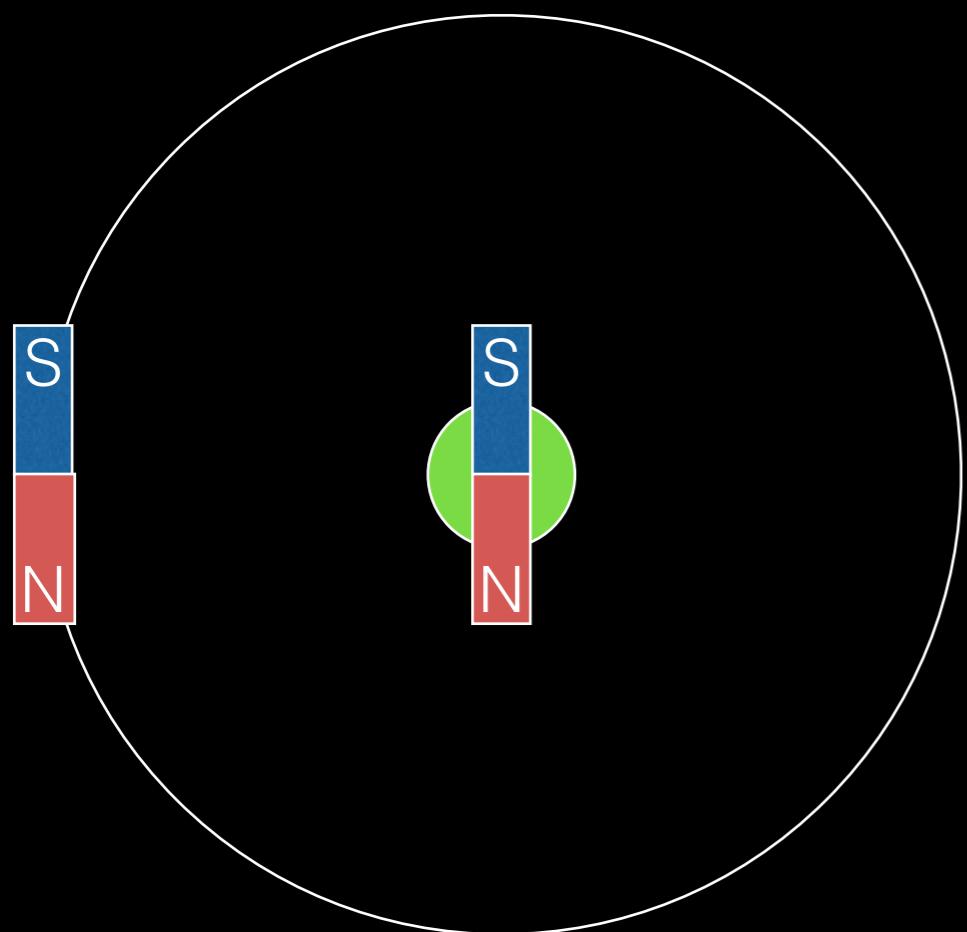
- A.  $d_x$  easy,  $\theta_y$  hard;  $v_x$  easy,  $\theta_y$  hard.
- B.  $d_x$  easy,  $\theta_y$  hard;  $v_x$  hard,  $\theta_y$  easy.
- C.  $d_x$  hard,  $\theta_y$  easy;  $v_x$  easy,  $\theta_y$  hard.
- D.  $d_x$  hard,  $\theta_y$  easy;  $v_x$  hard,  $\theta_y$  easy.



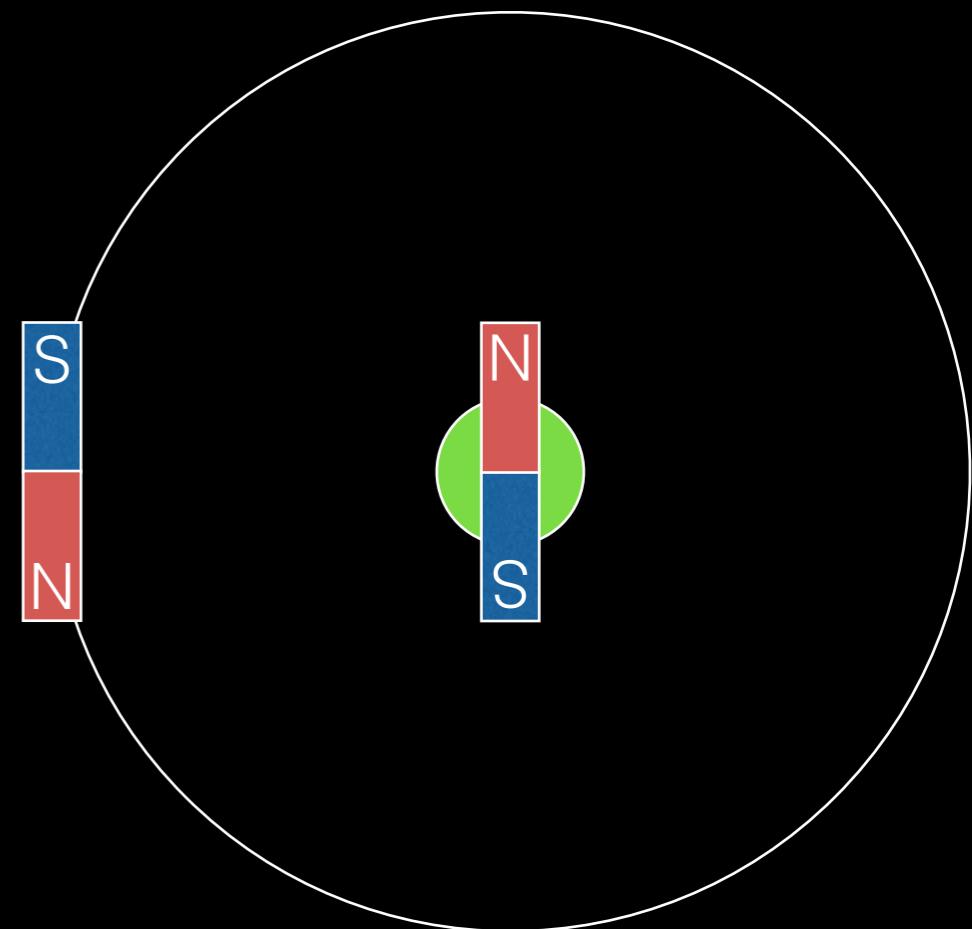
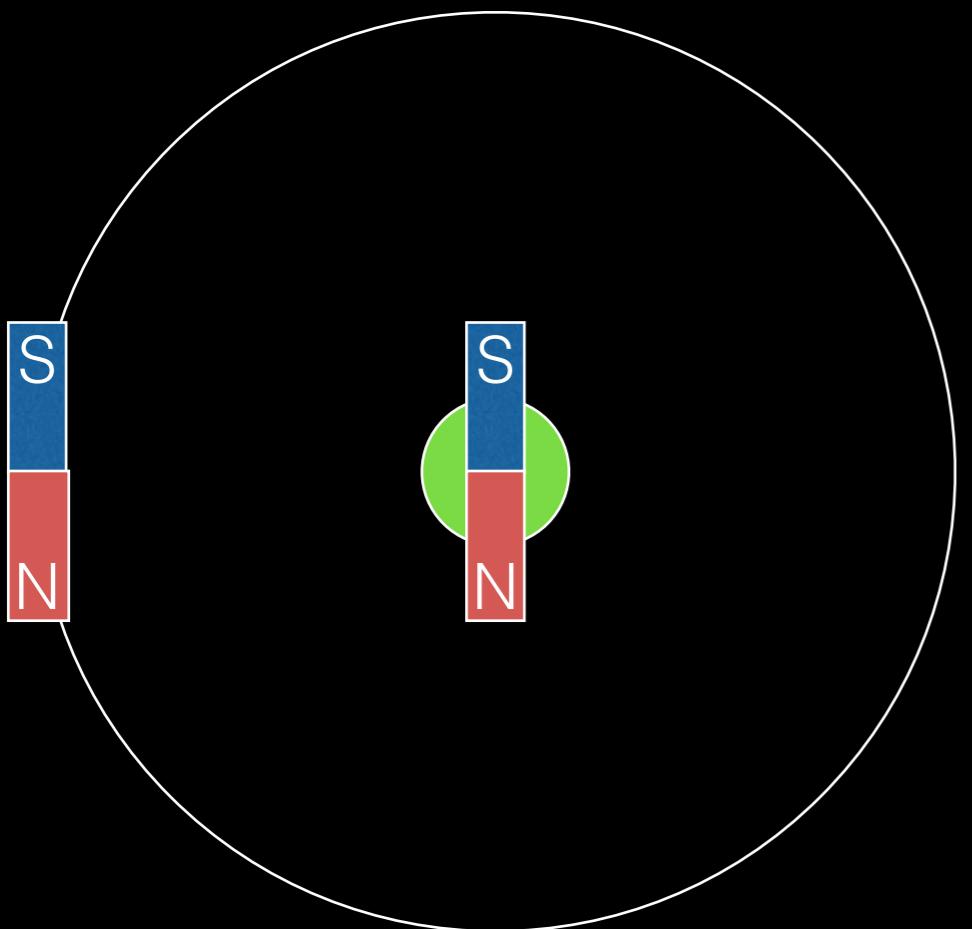
Hydrogen atoms emit radio waves with 21cm wavelengths



Hydrogen atoms emit radio waves with 21cm wavelengths

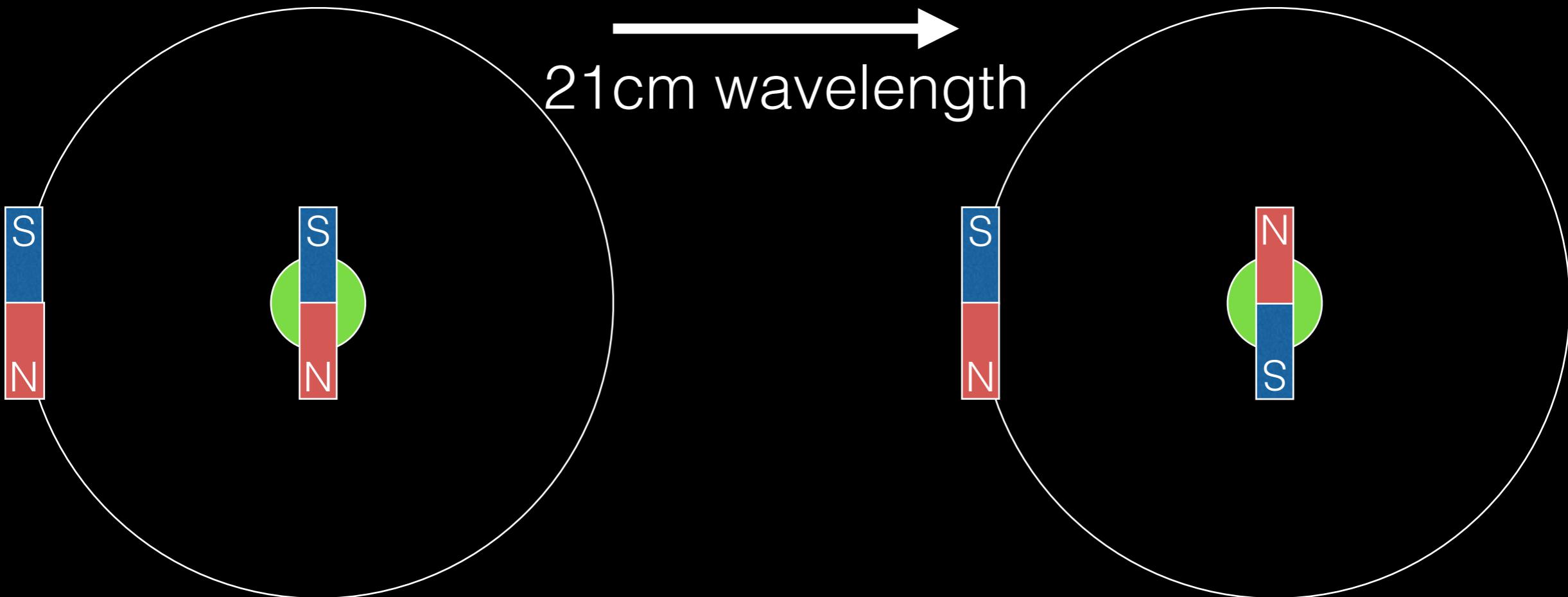


# Hydrogen atoms emit radio waves with 21cm wavelengths

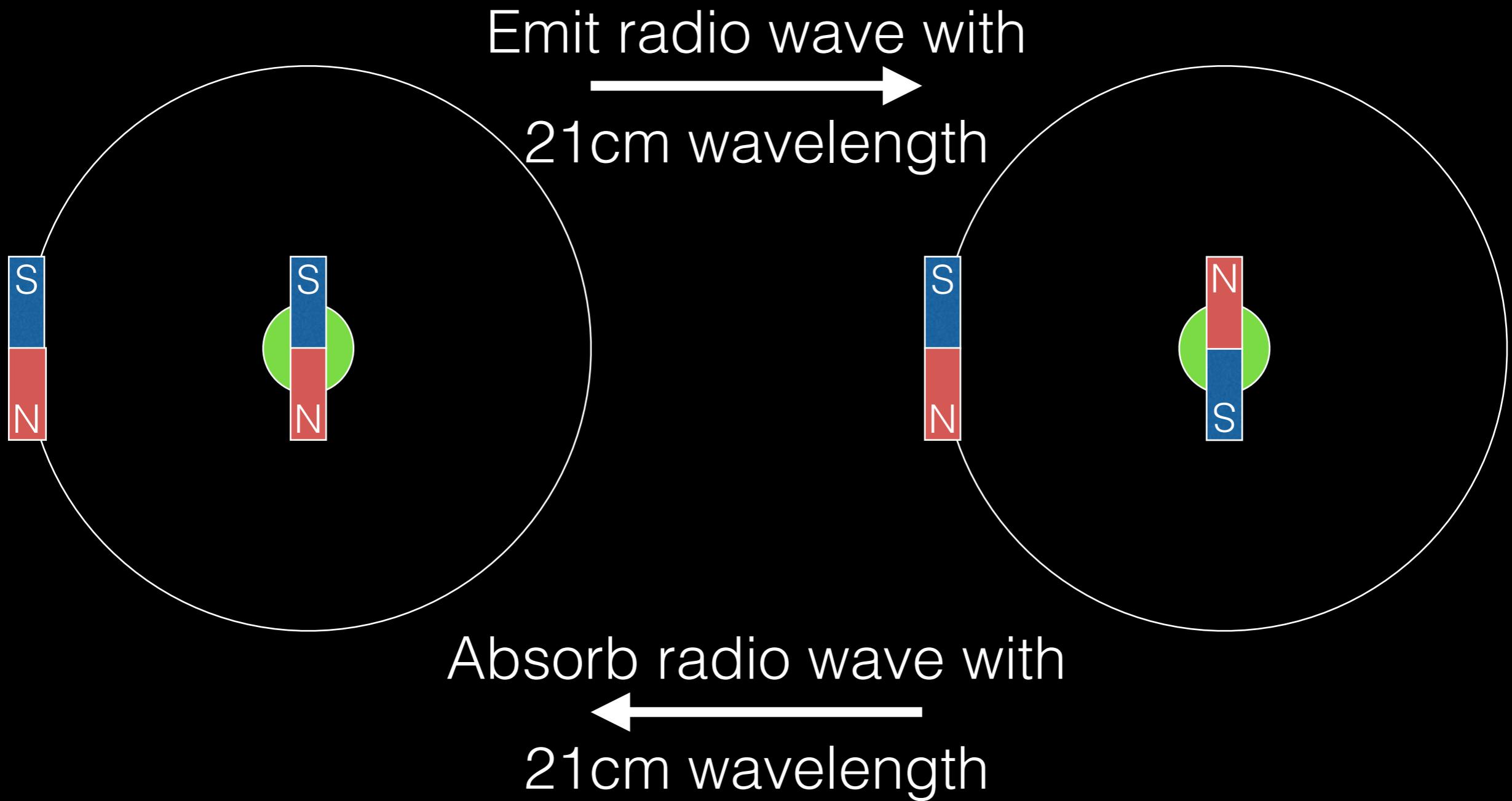


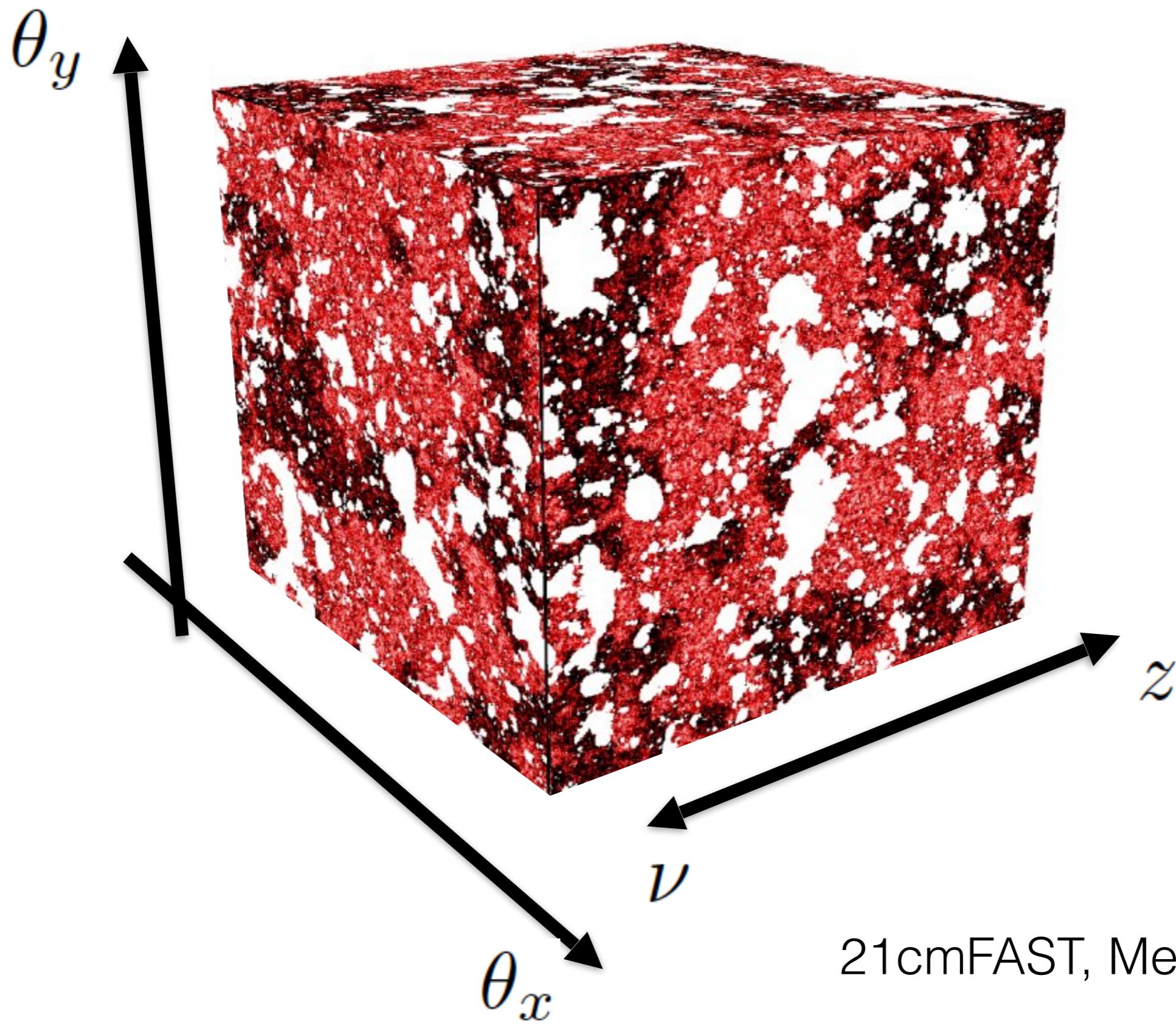
# Hydrogen atoms emit radio waves with 21cm wavelengths

Emit radio wave with  
→  
21cm wavelength



# Hydrogen atoms emit radio waves with 21cm wavelengths





21cmFAST, Mesinger et al.

The night sky is a time  
machine

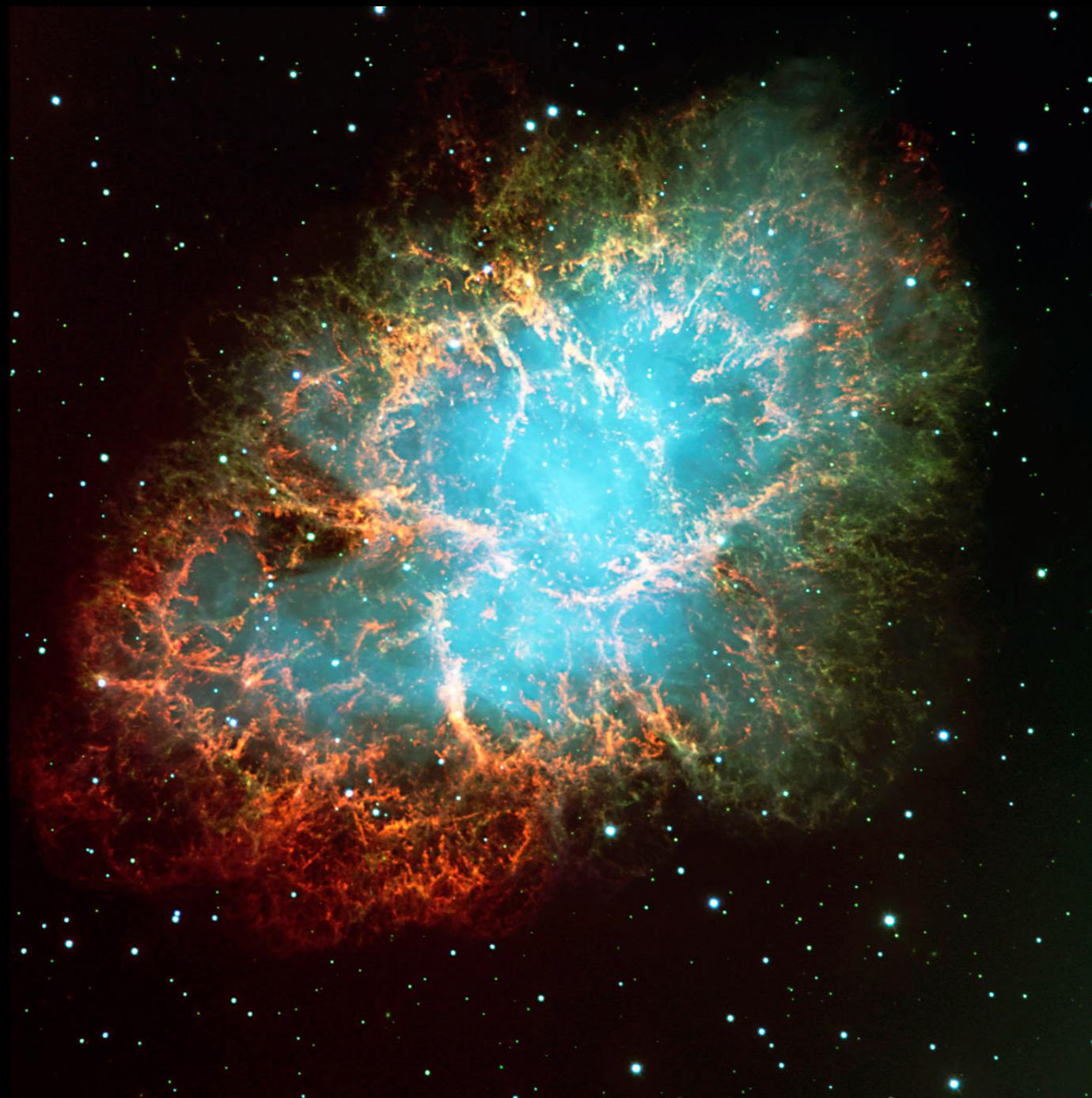
# The night sky is a time machine

- Light coming to us from 4 light years away started traveling towards us 4 years ago.

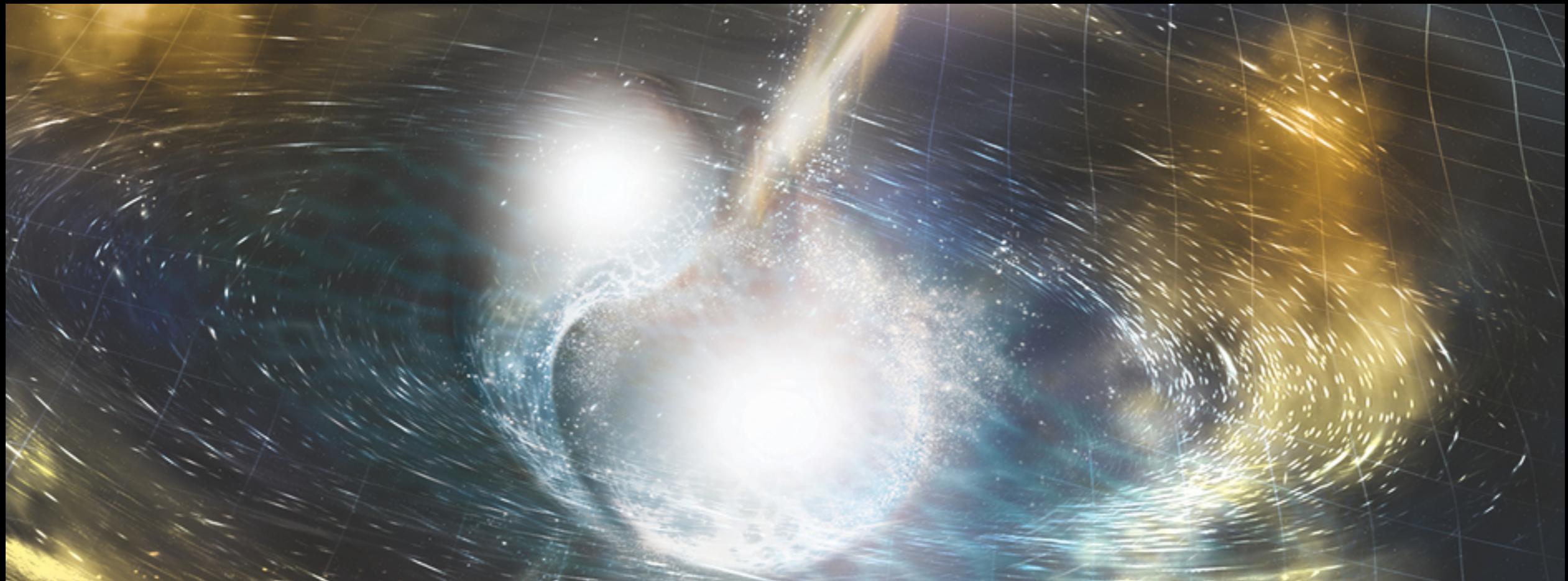
# The night sky is a time machine

- Light coming to us from 4 light years away started traveling towards us 4 years ago.
- Light coming to us from 4000 light years away started traveling towards us 4000 years ago.

We see the Crab Nebula not as it is today, but as it was 6500 years ago



The neutron star merger that was recently detected happened about 120 million years ago!



We see distant galaxies not as they are today, but they were billions of years ago



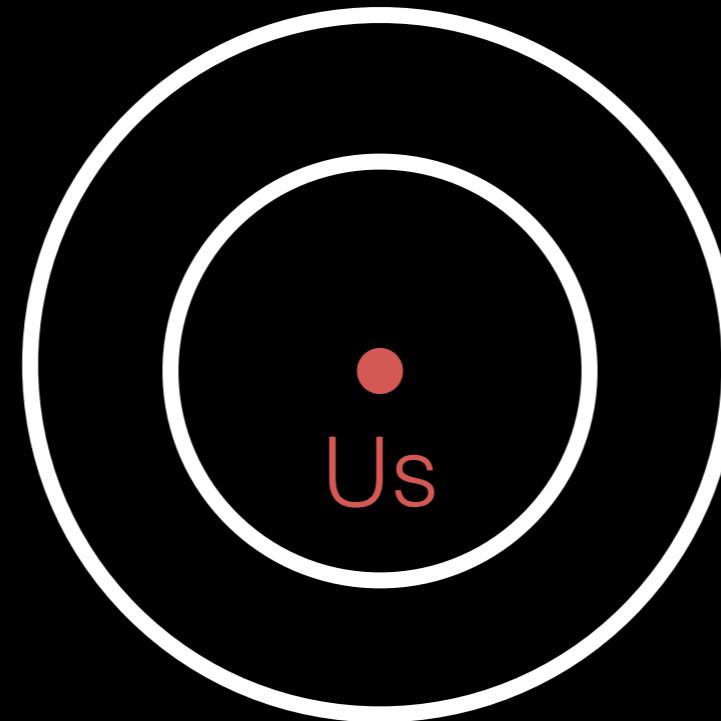
The farther away we look,  
the farther back in time  
we see

•  
Us

10 years ago



20 years ago



How far back have we  
looked? How far back **can**  
we look?

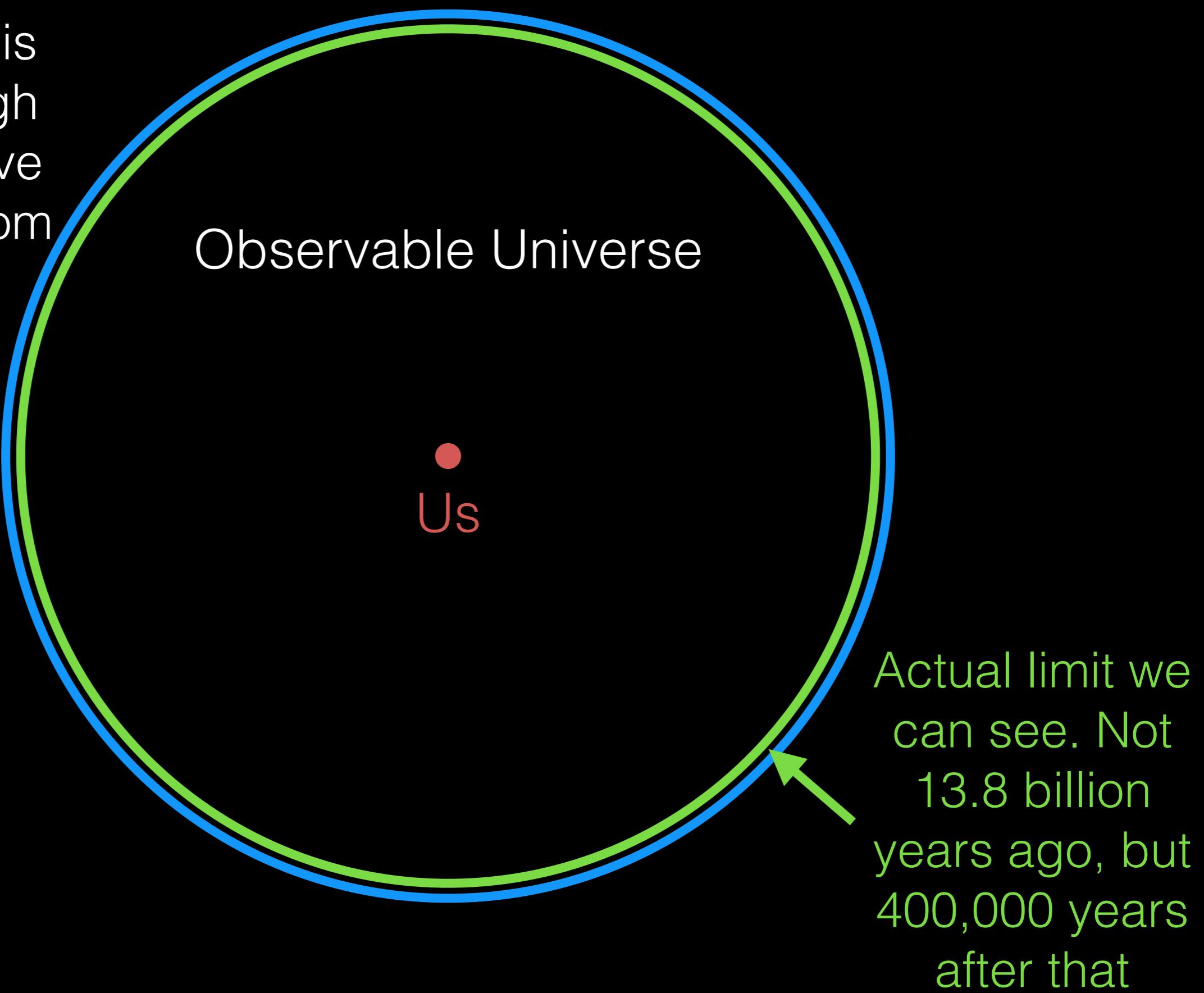
Unobservable  
Universe:  
our Universe is  
not old enough  
for light to have  
reached us from  
out here

13.8 billion years ago

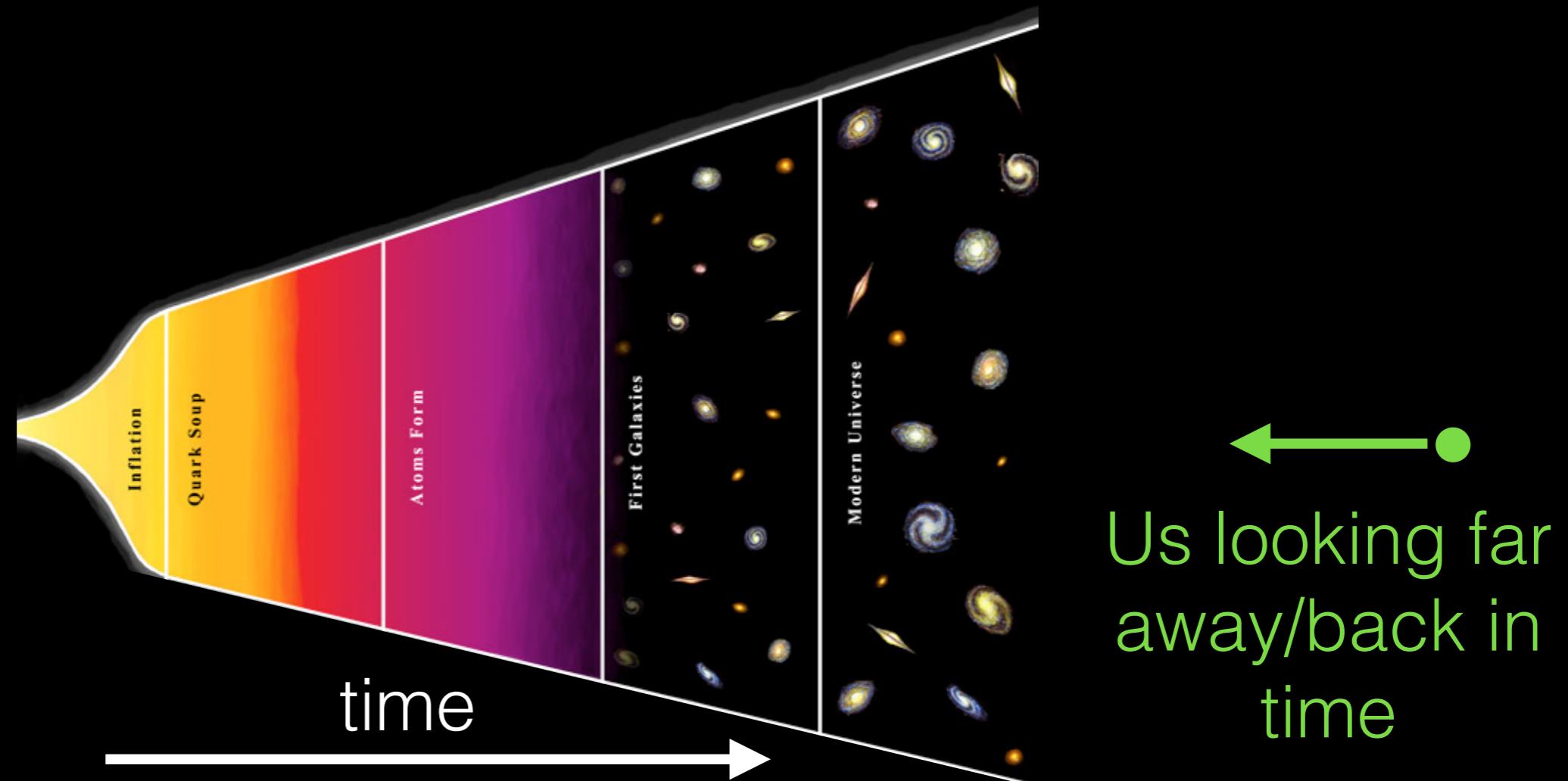
Observable Universe



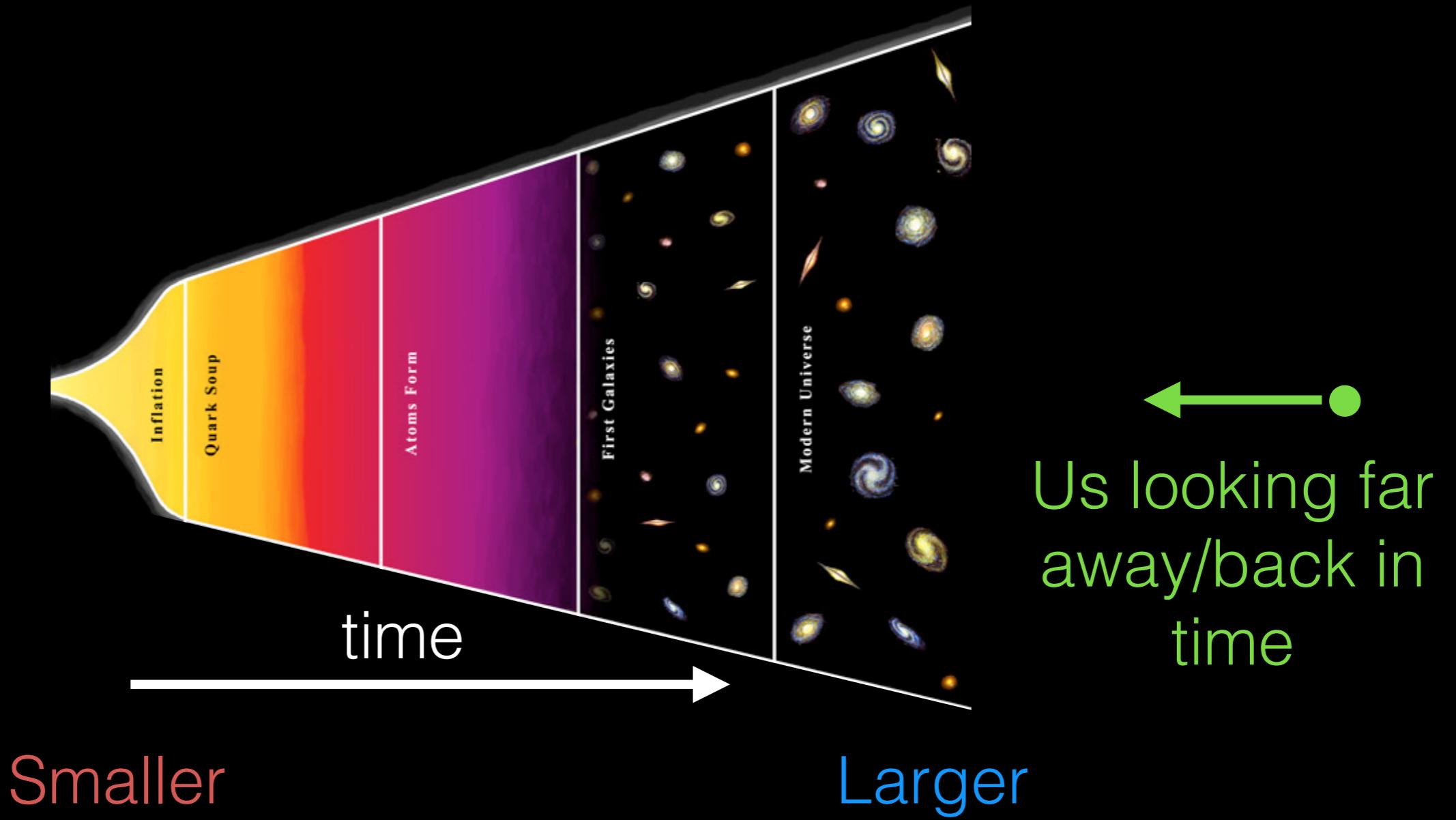
Unobservable  
Universe:  
our Universe is  
not old enough  
for light to have  
reached us from  
out here



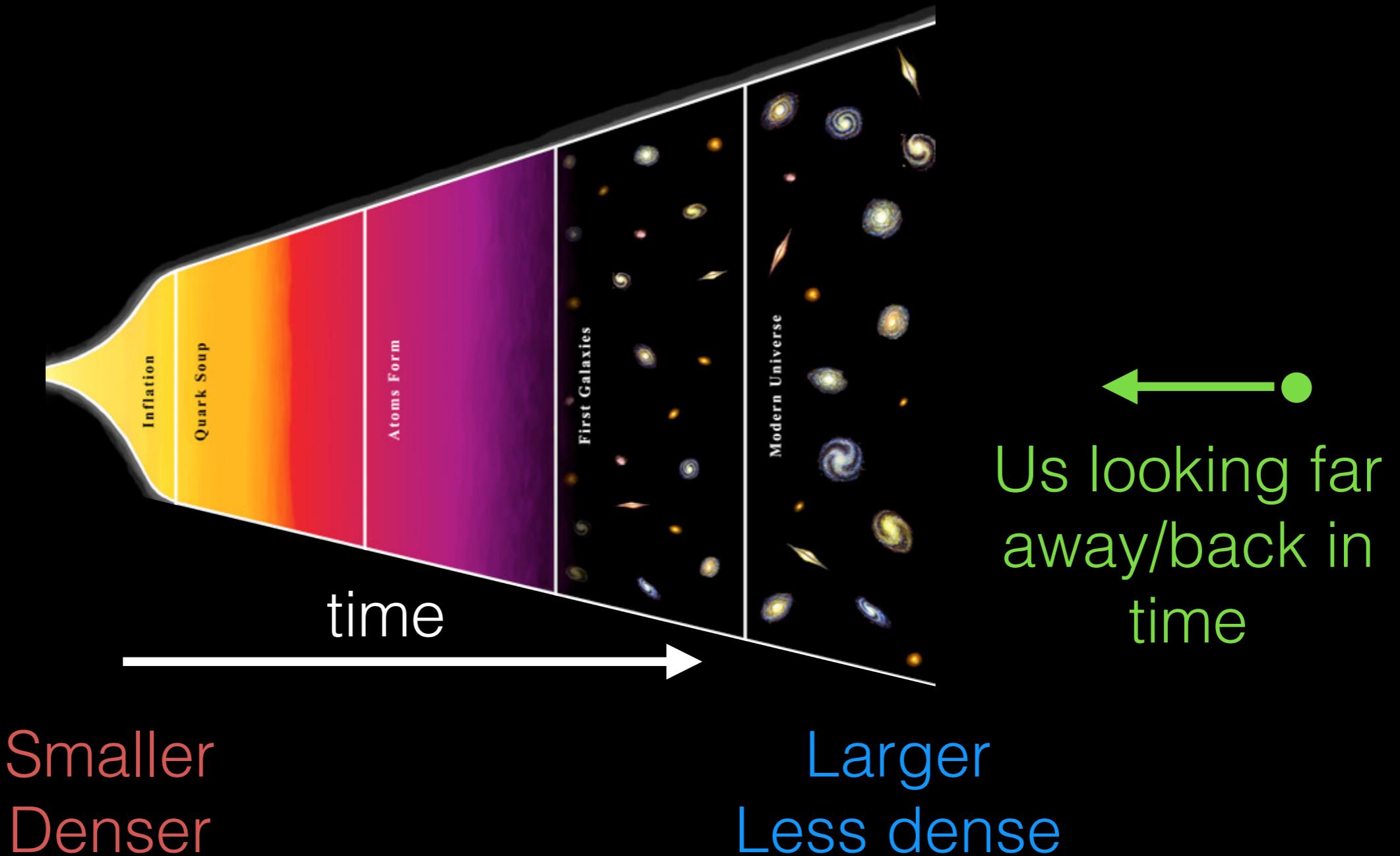
Our Universe is expanding, so as we look back in time, we see a smaller, denser, hotter Universe



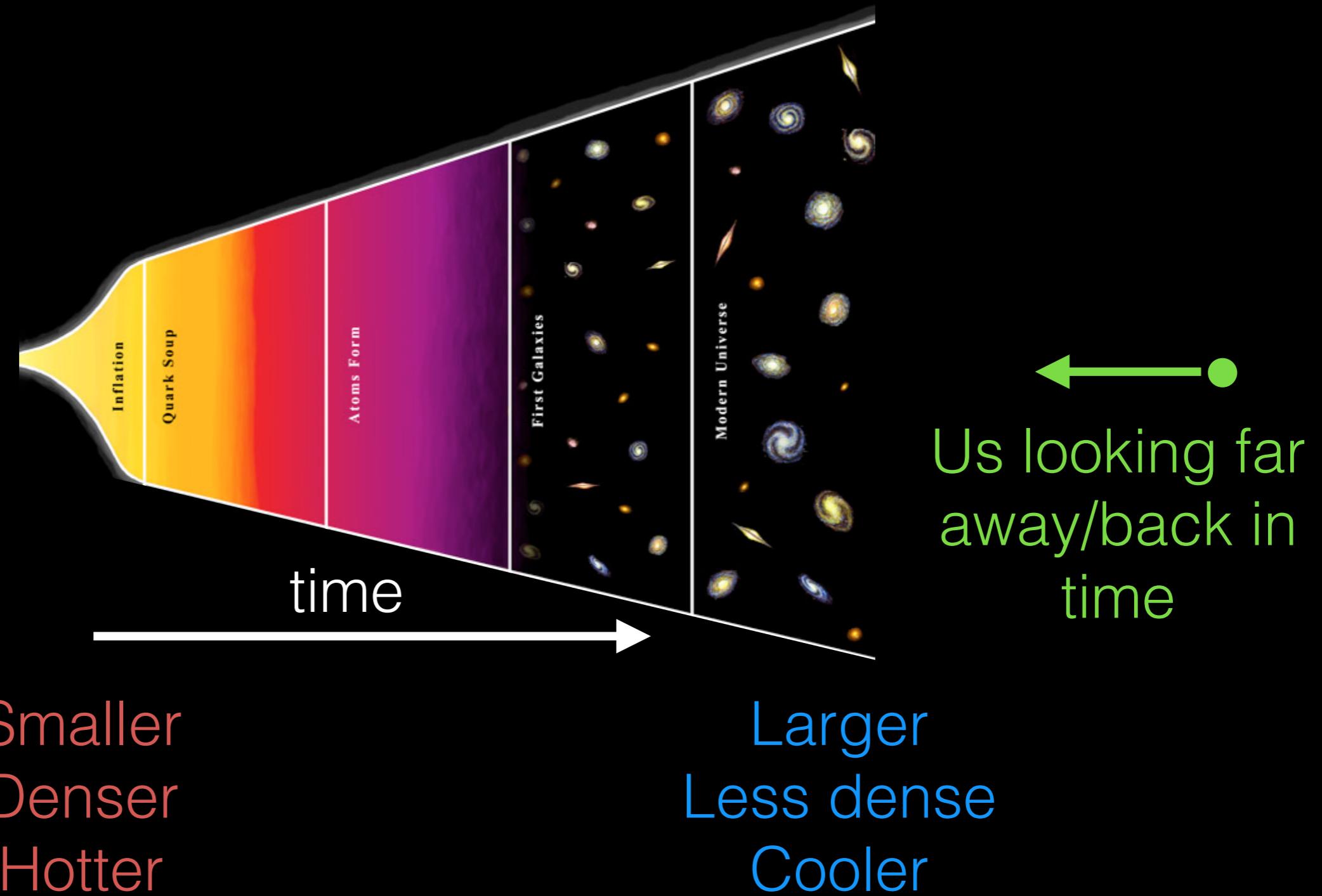
Our Universe is expanding, so as we look back in time, we see a smaller, denser, hotter Universe

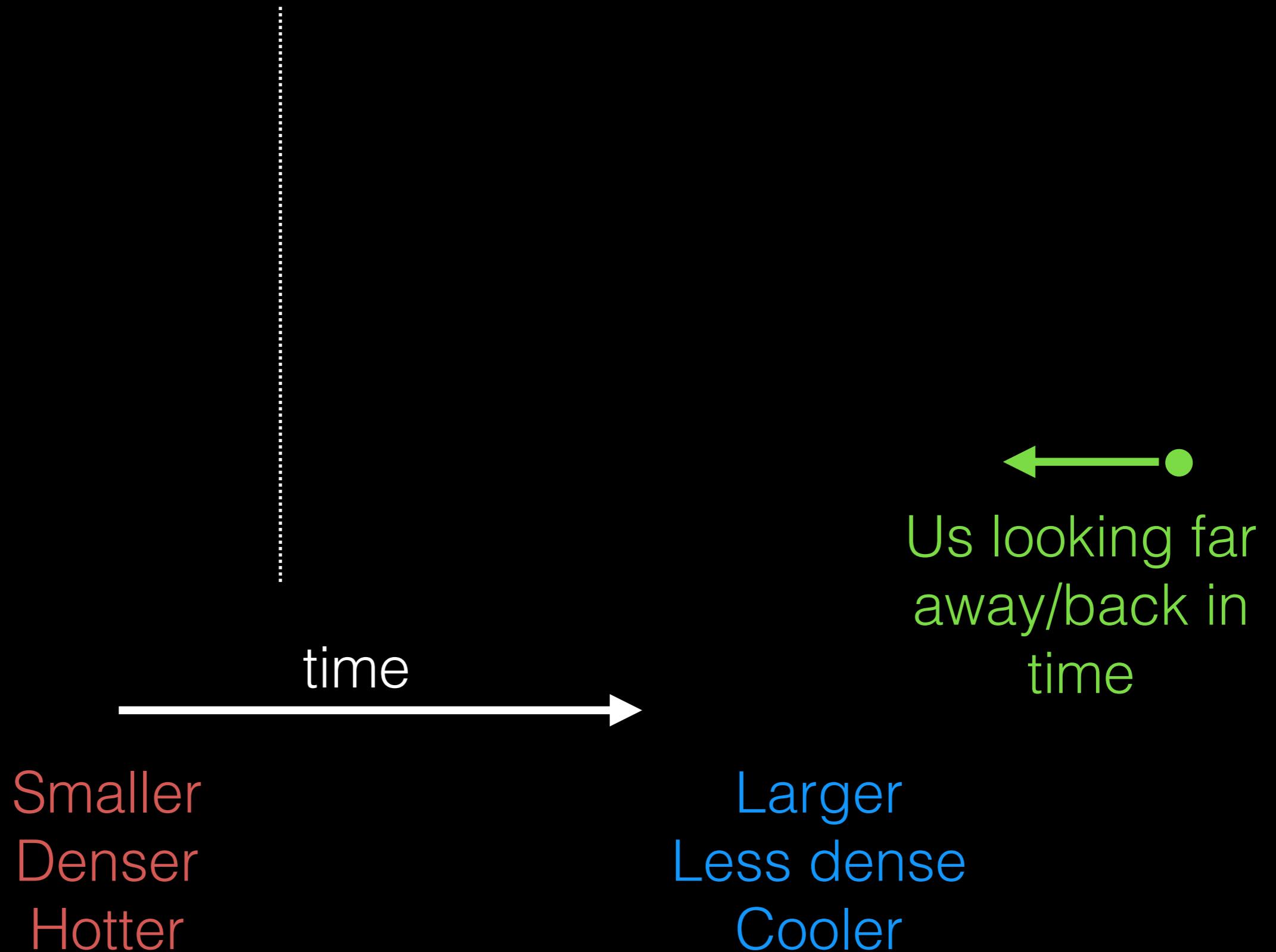


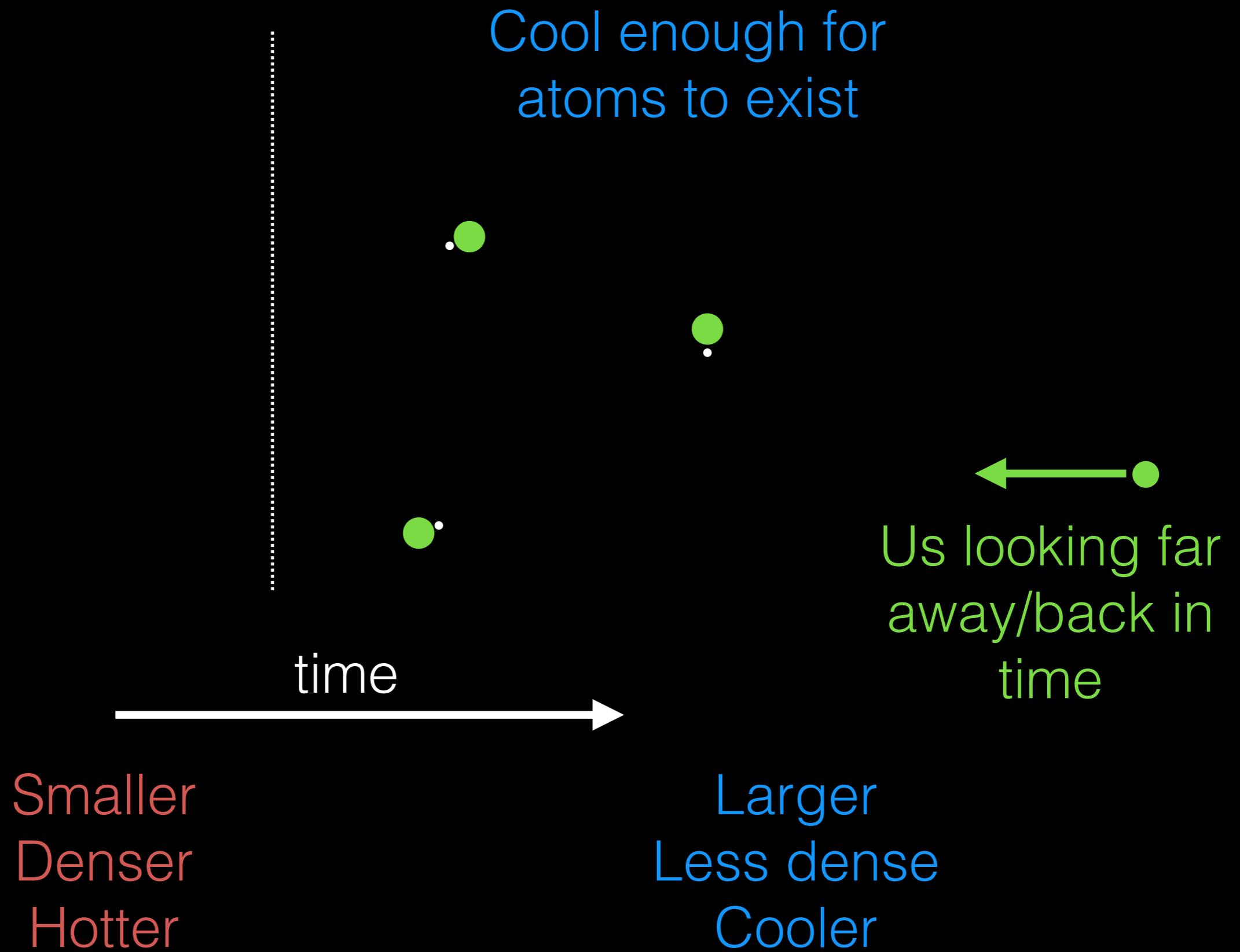
Our Universe is expanding, so as we look back in time, we see a smaller, denser, hotter Universe



Our Universe is expanding, so as we look back in time, we see a smaller, denser, hotter Universe

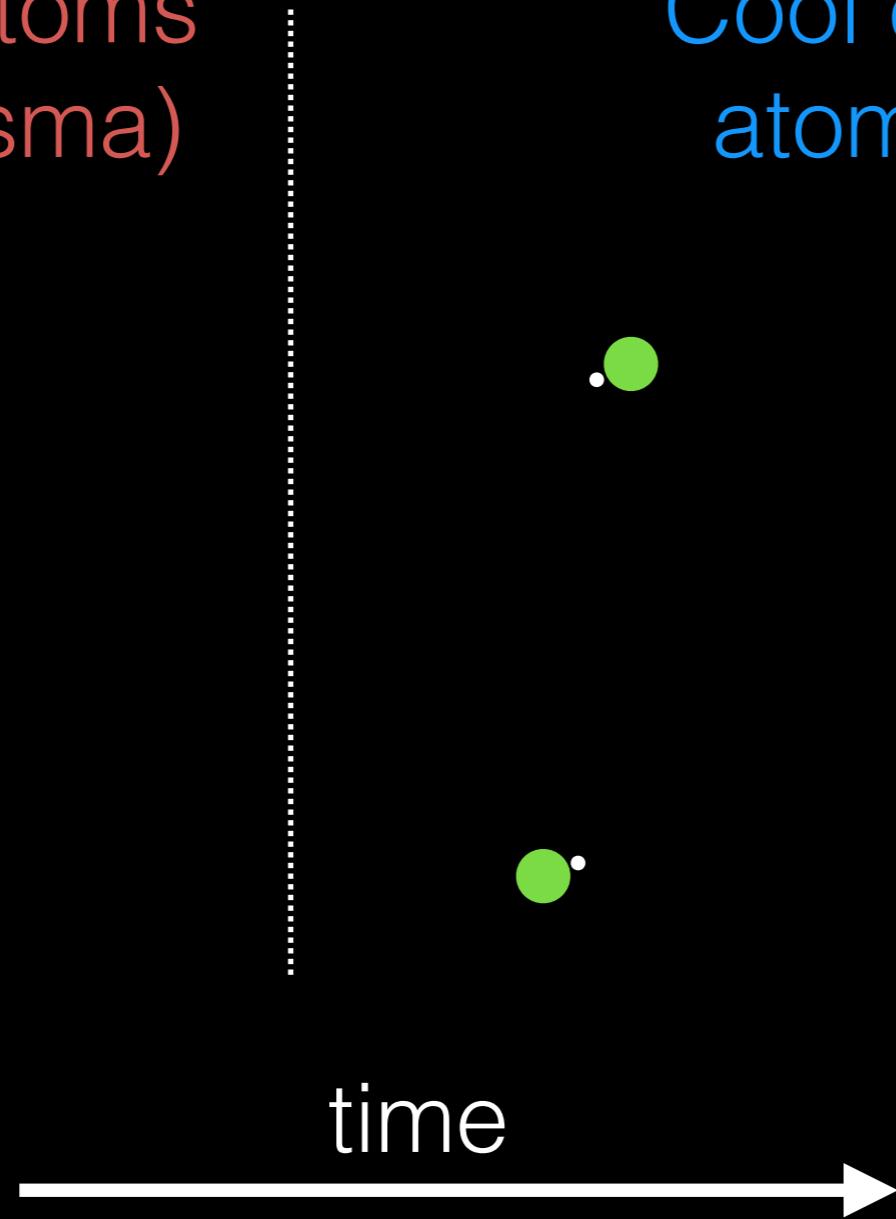






Too hot for atoms  
to form (plasma)

Cool enough for  
atoms to exist



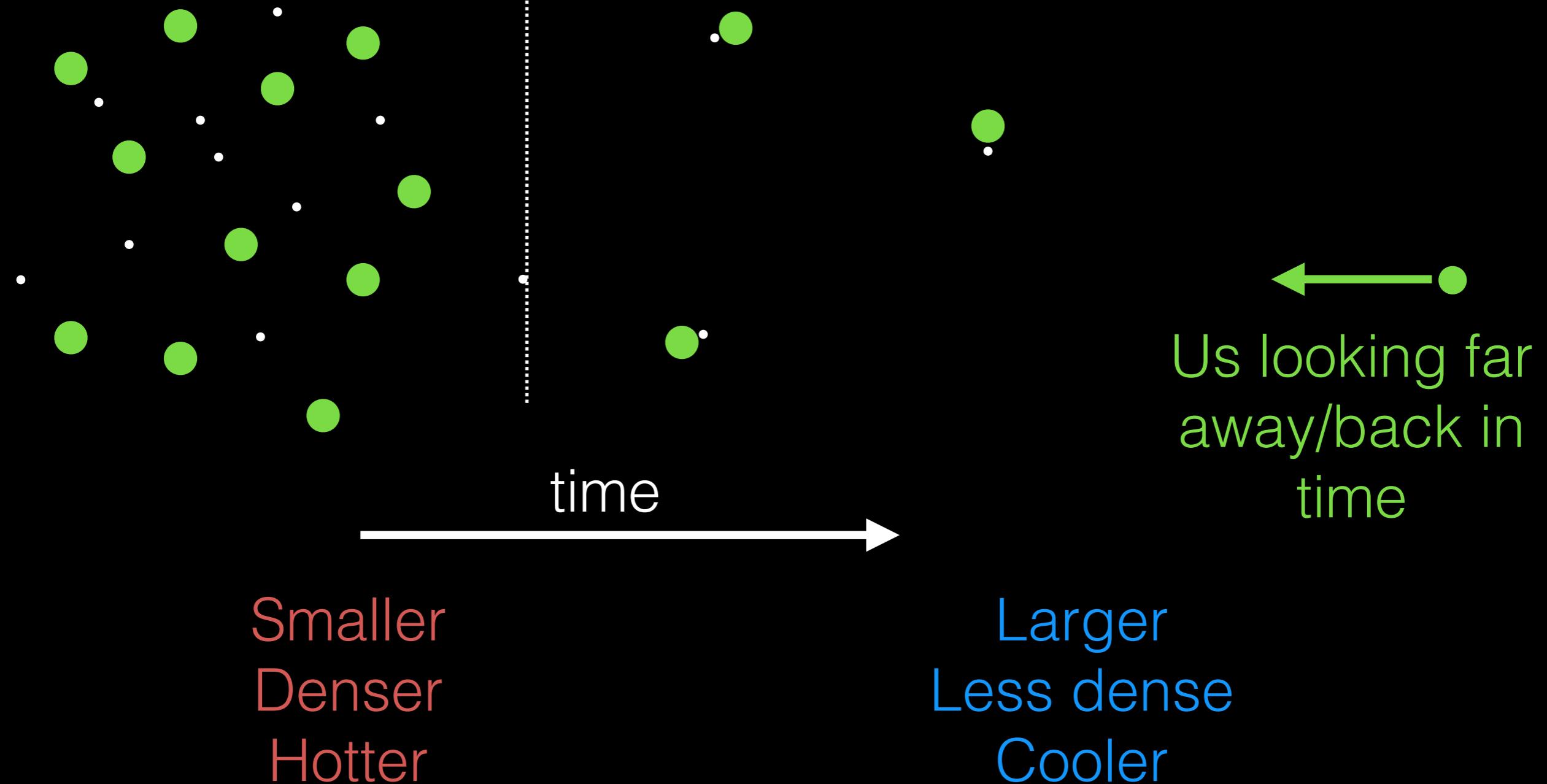
Smaller  
Denser  
Hotter

Larger  
Less dense  
Cooler

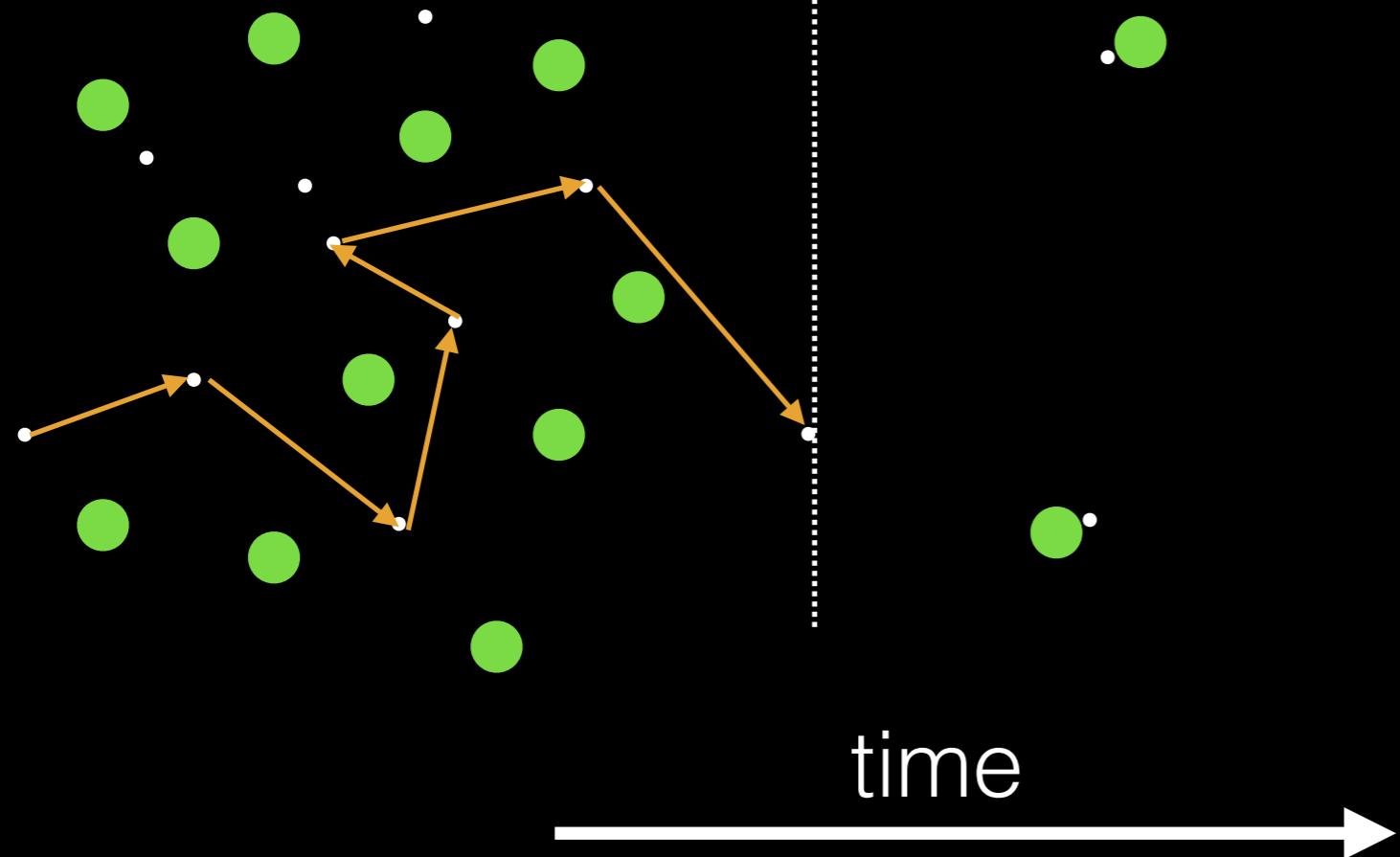
← ●  
Us looking far  
away/back in  
time

Too hot for atoms  
to form (plasma)

Cool enough for  
atoms to exist



Too hot for atoms  
to form (plasma)  
Not transparent



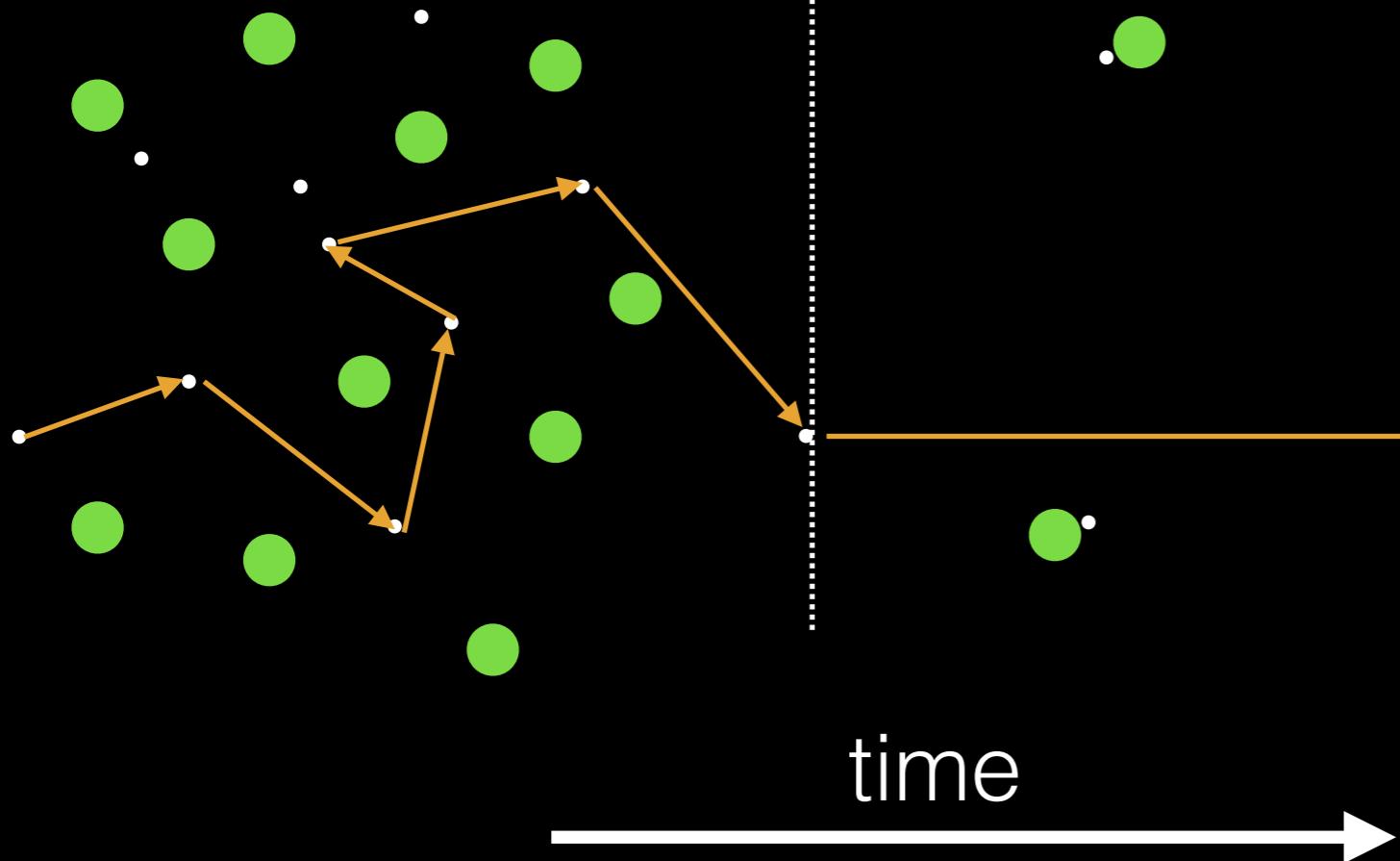
Smaller  
Denser  
Hotter

Larger  
Less dense  
Cooler

Cool enough for  
atoms to exist

Us looking far  
away/back in  
time

Too hot for atoms  
to form (plasma)  
Not transparent

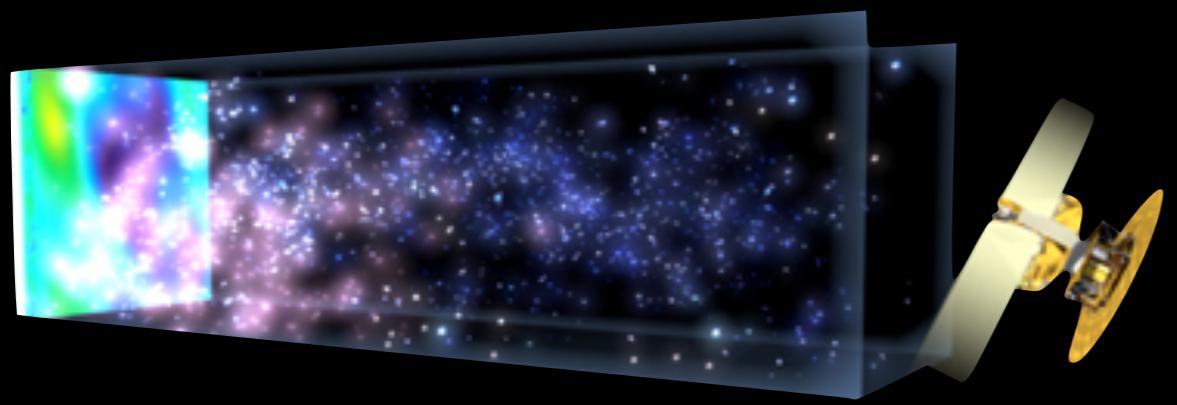


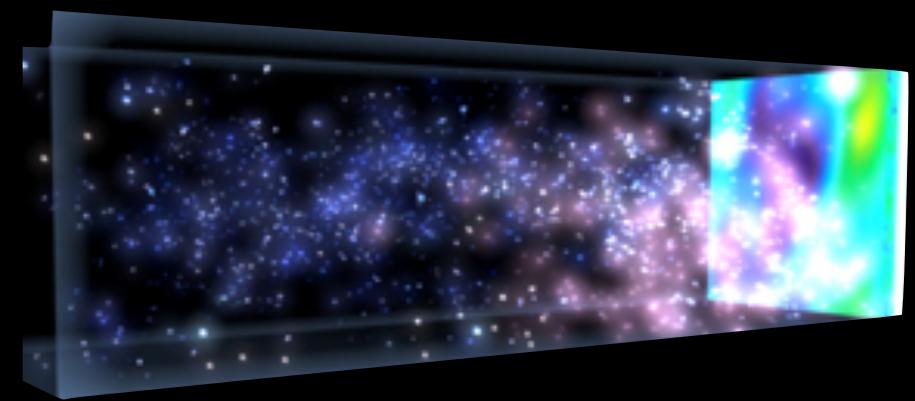
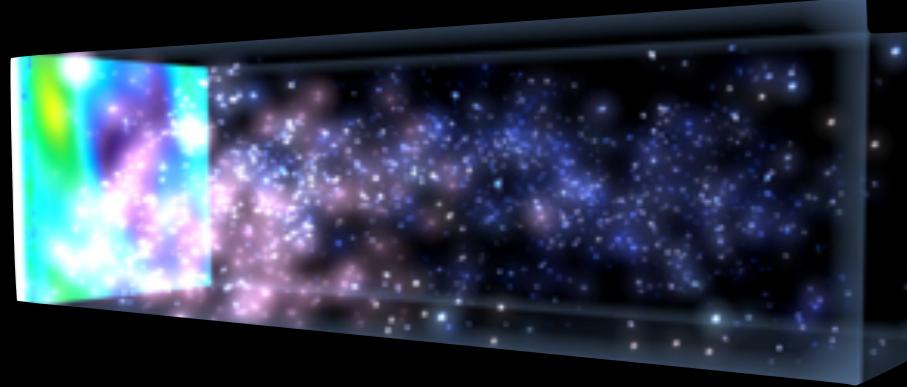
Smaller  
Denser  
Hotter

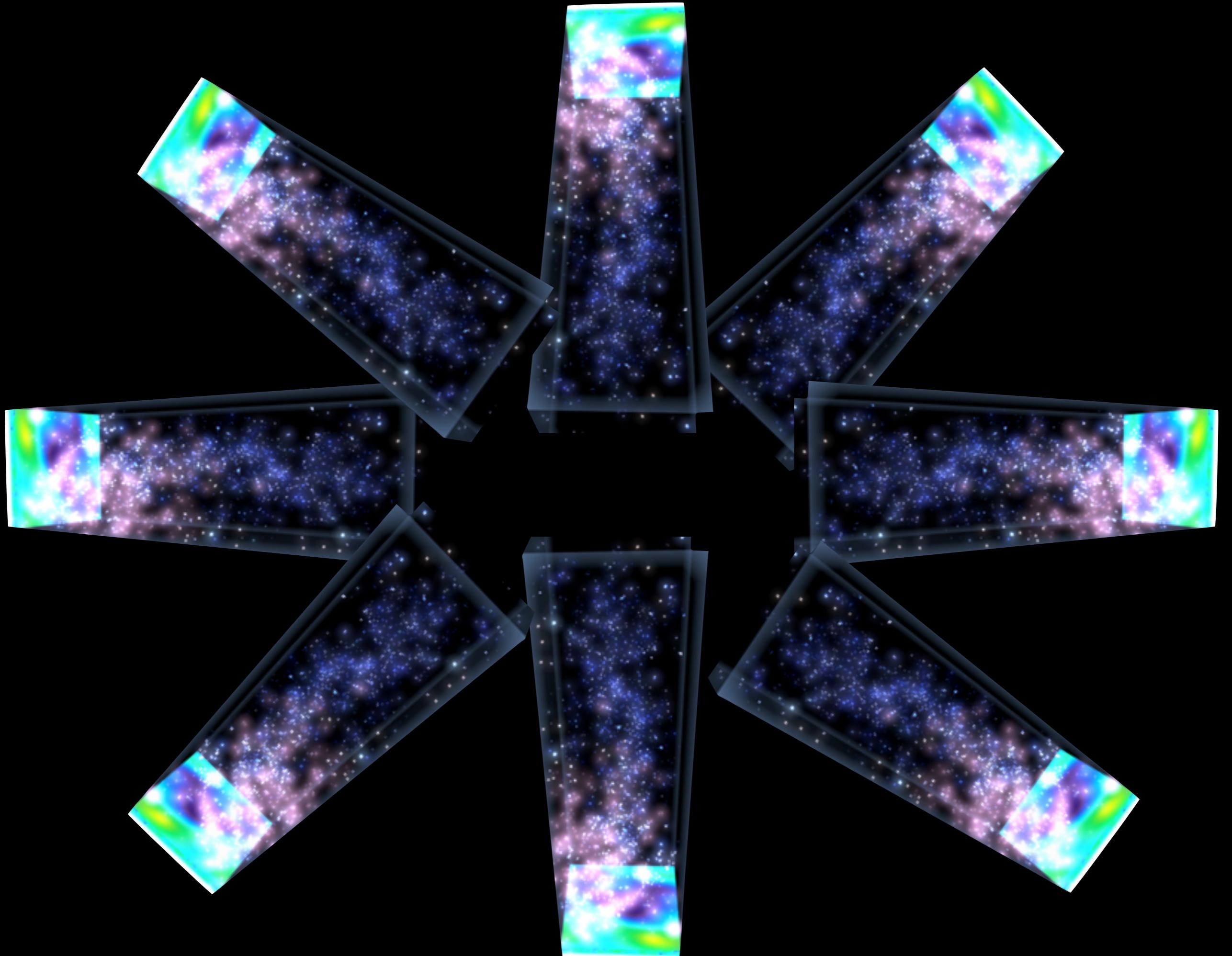
Cool enough for  
atoms to exist  
Transparent

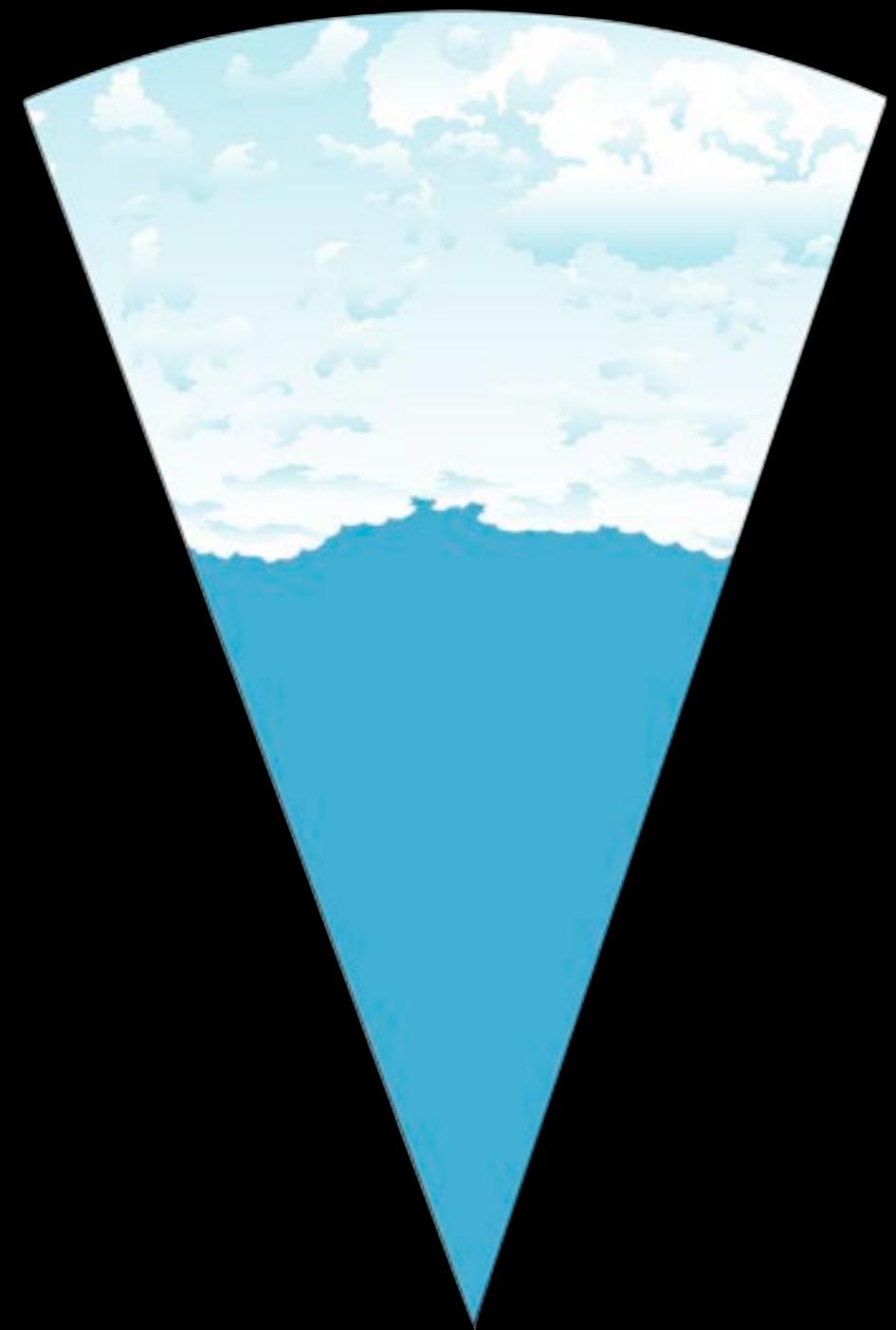
Larger  
Less dense  
Cooler

Us looking far  
away/back in  
time

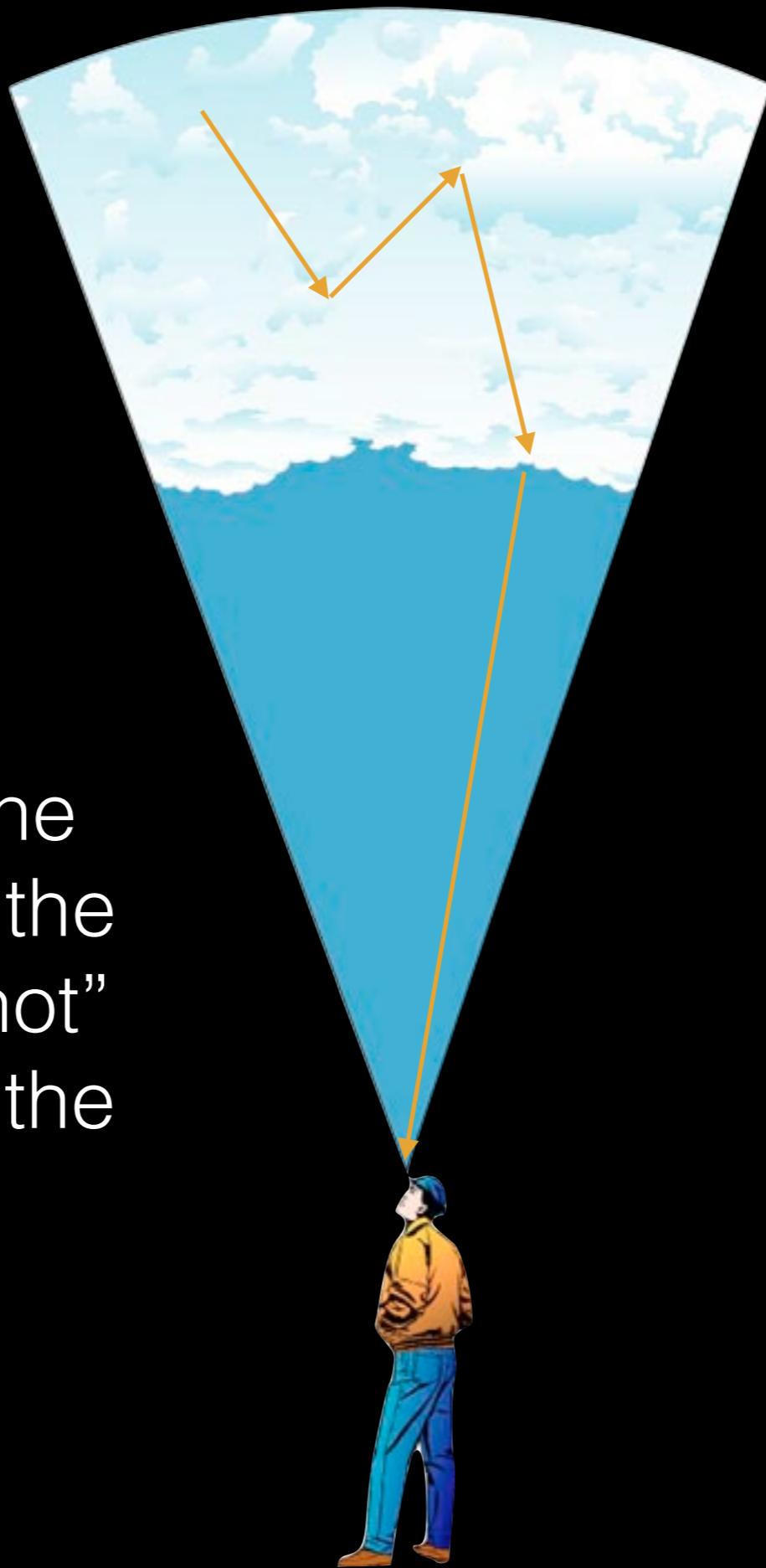




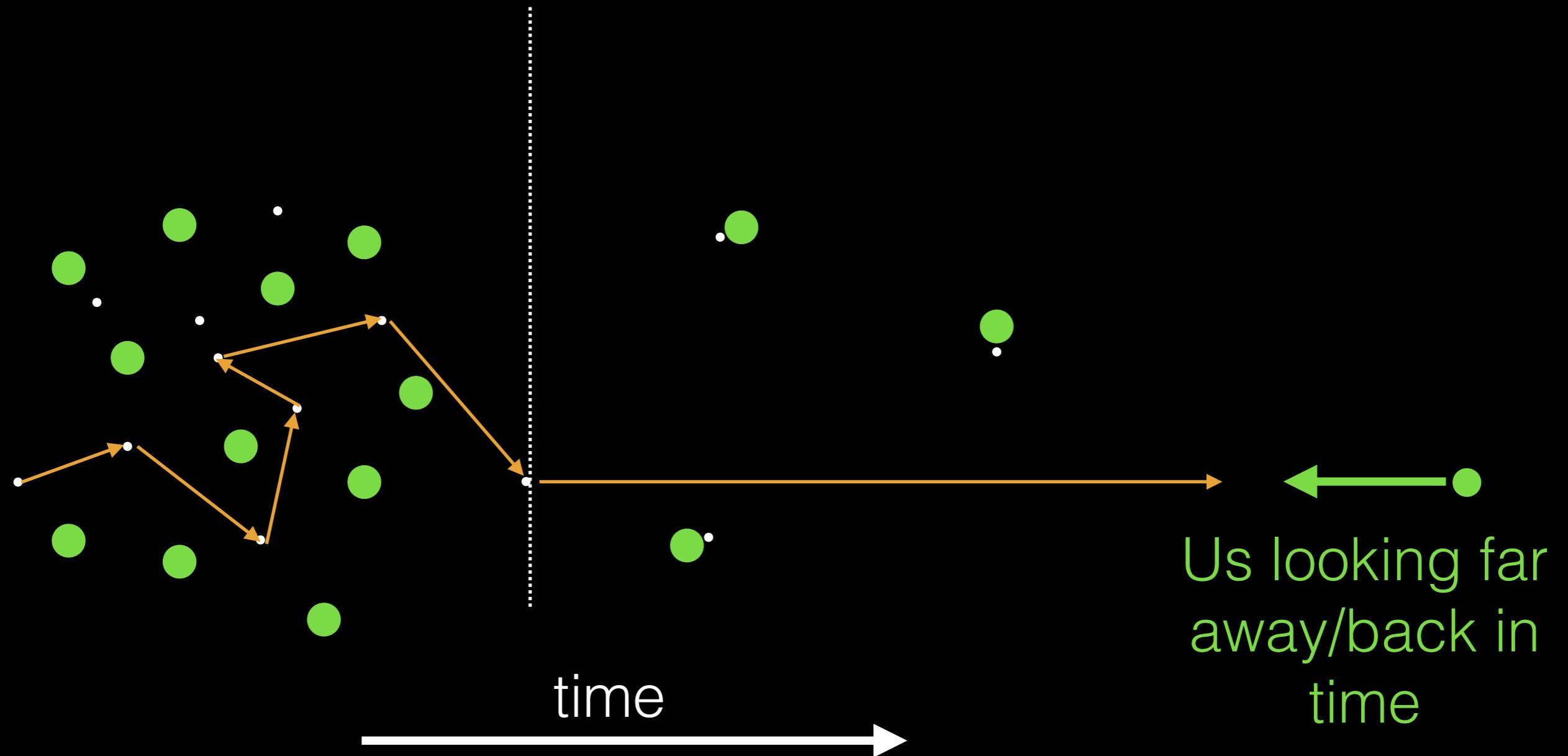




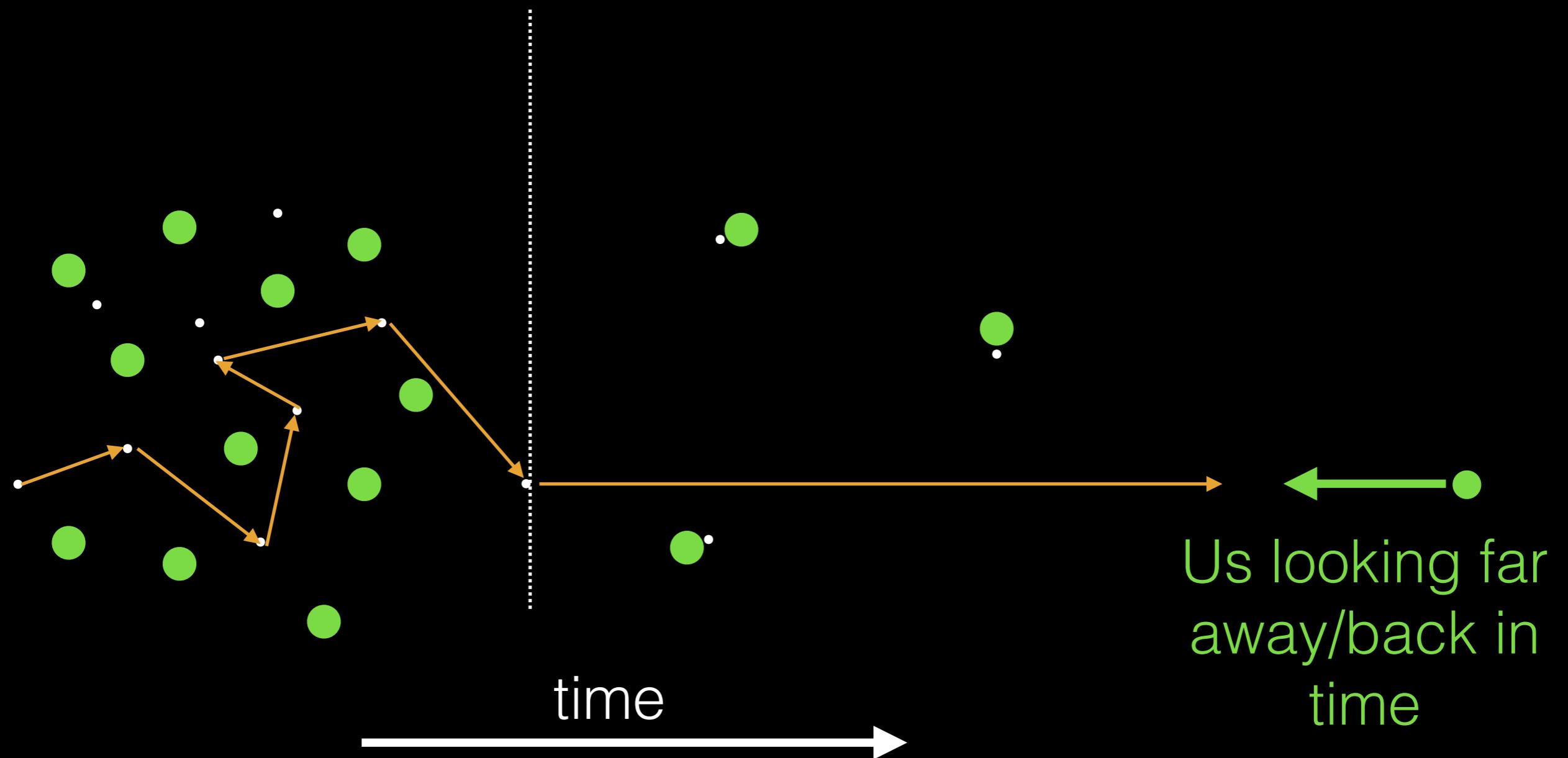
We see an outline  
of the surface of the  
cloud, a “snapshot”  
of the exterior of the  
cloud



400,000 years

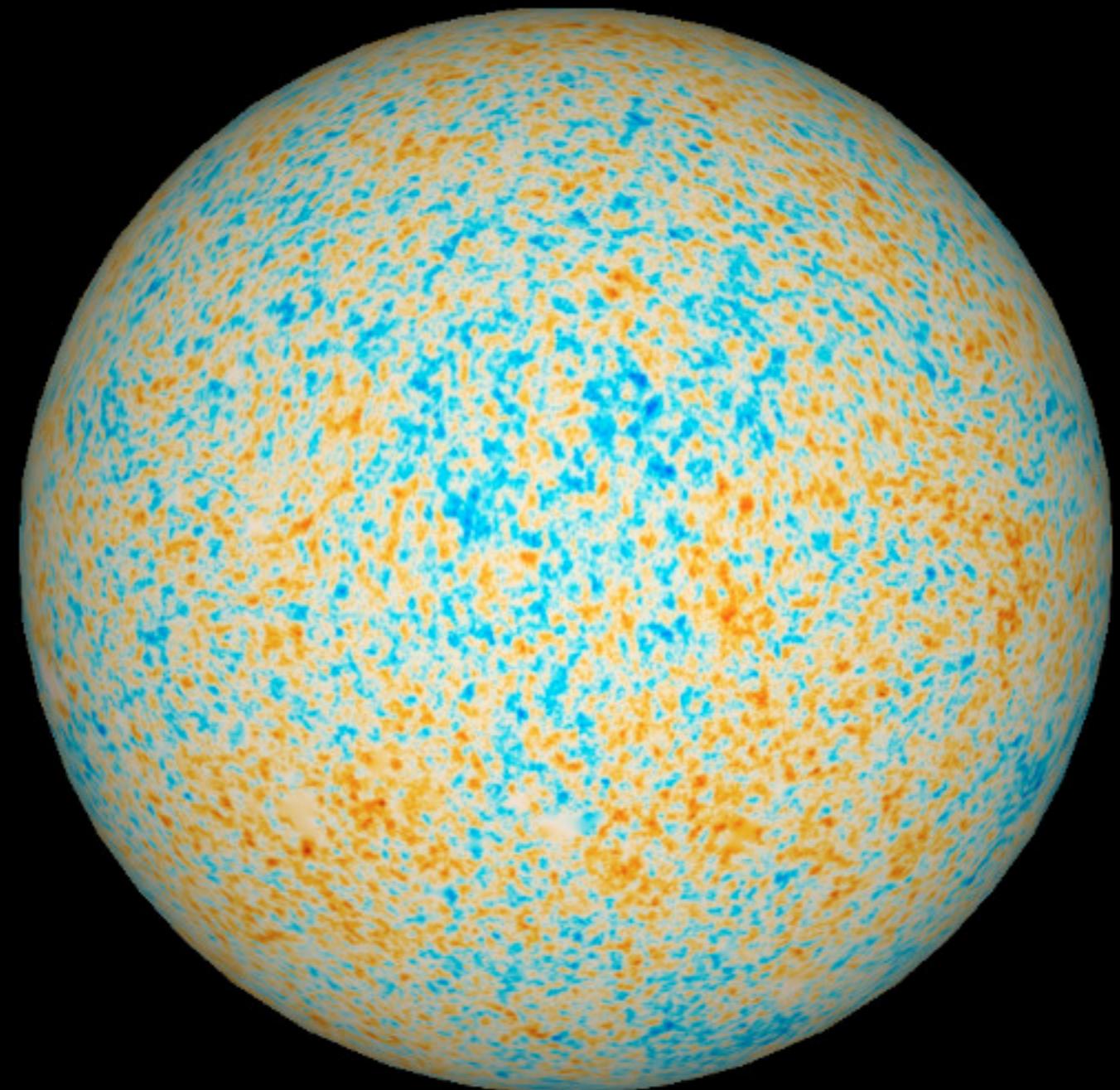


400,000 years



We see a snapshot of our Universe when it was only 400,000 years old. A mere baby!

The Cosmic Microwave Background provides a snapshot of our baby Universe

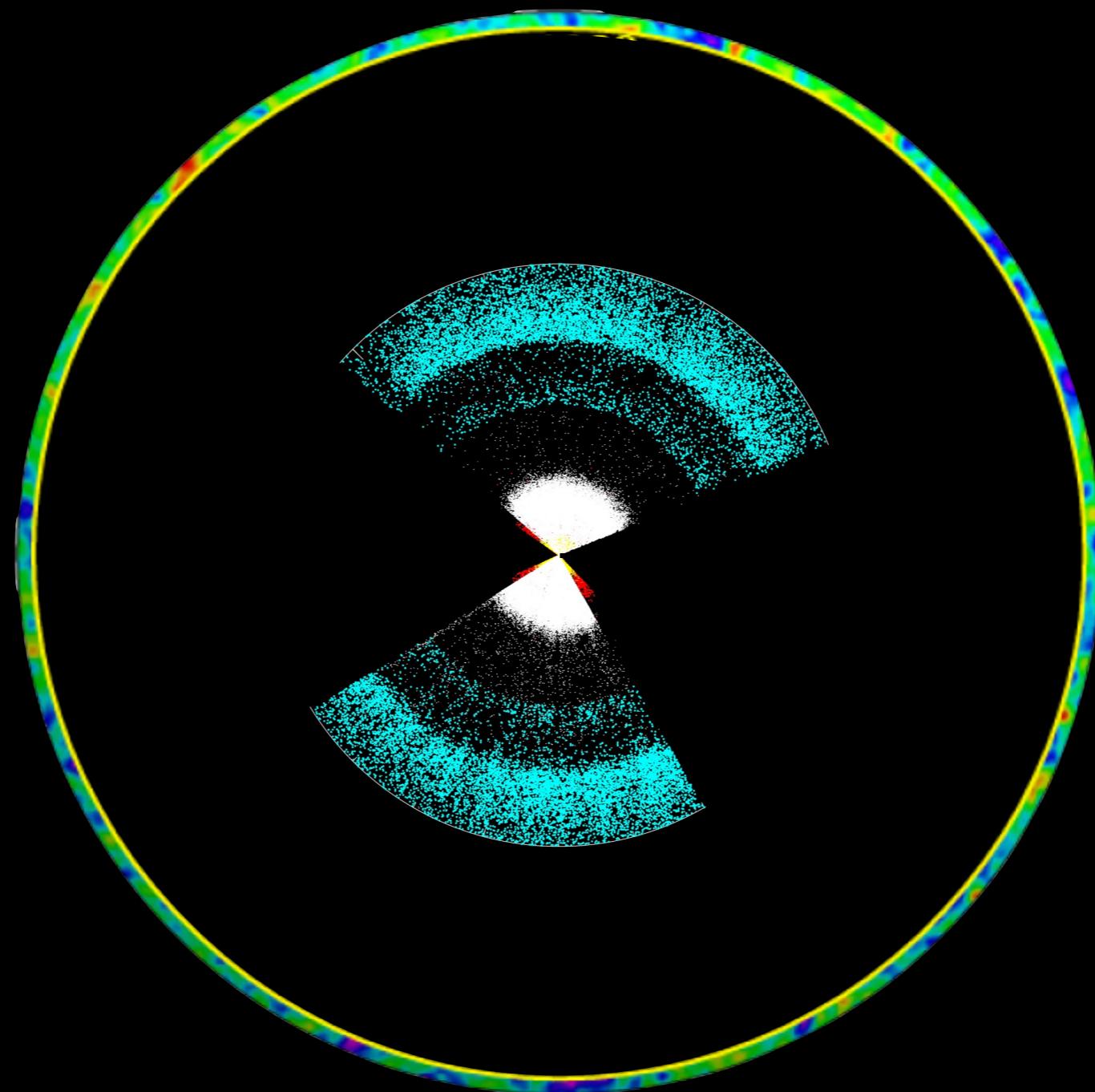




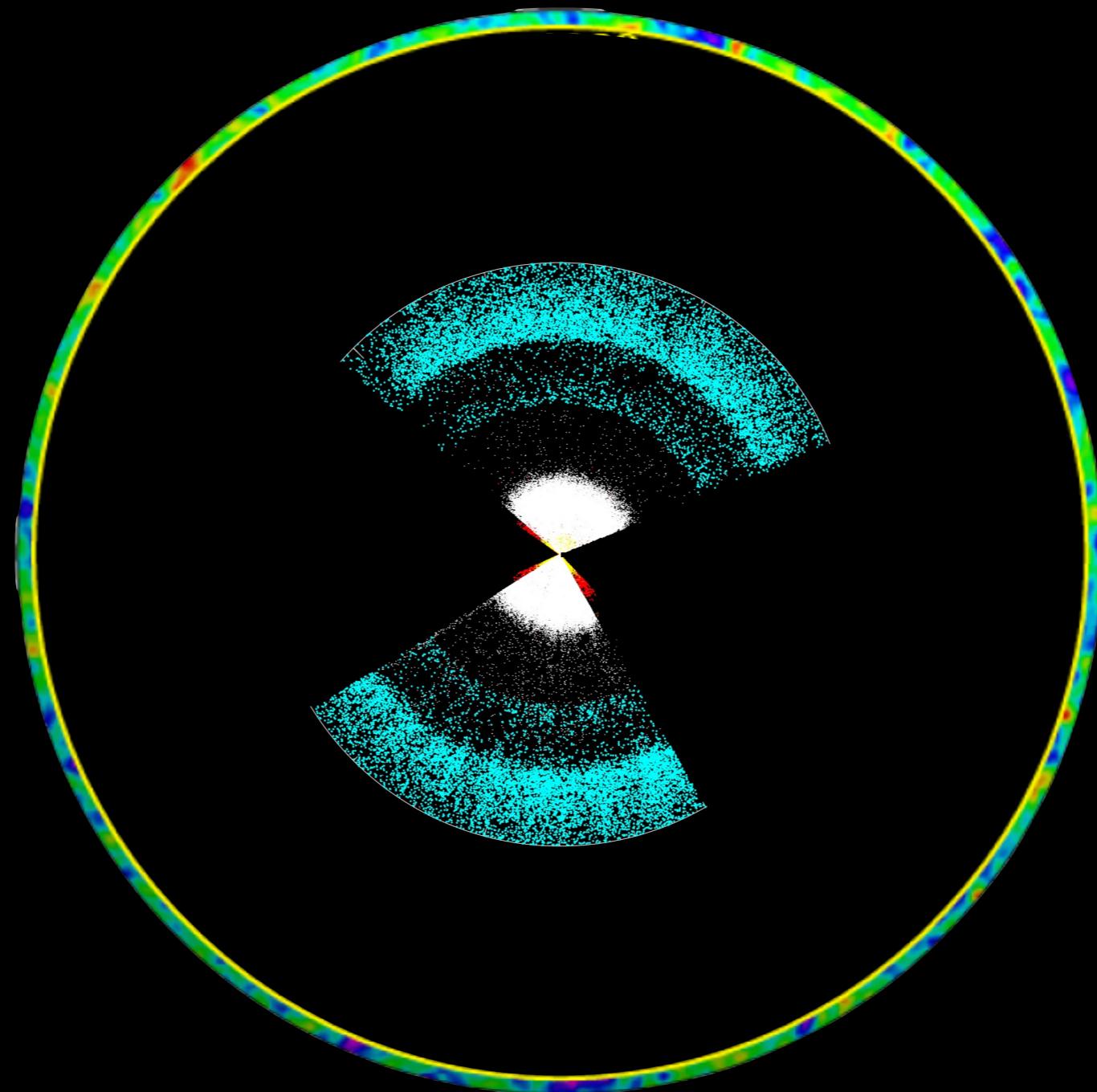
Alvarez et al. (2009)

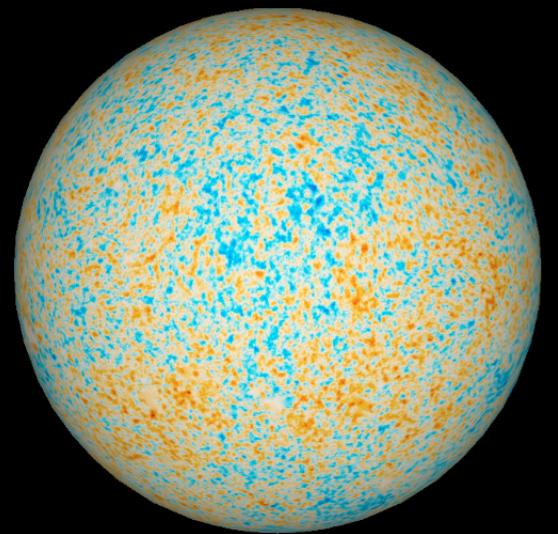
What about the real  
thing?

We have yet to observe most  
of the observable Universe



We have seen our grown-up Universe, our baby Universe, but not our Universe as it was growing up

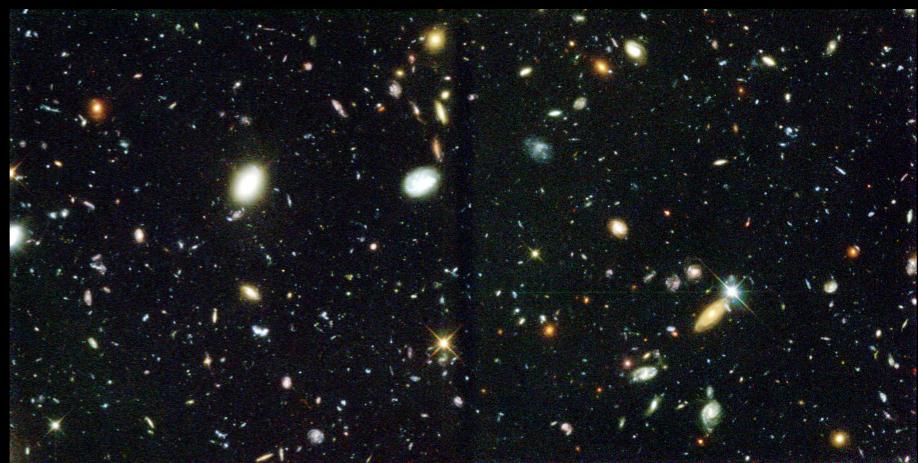




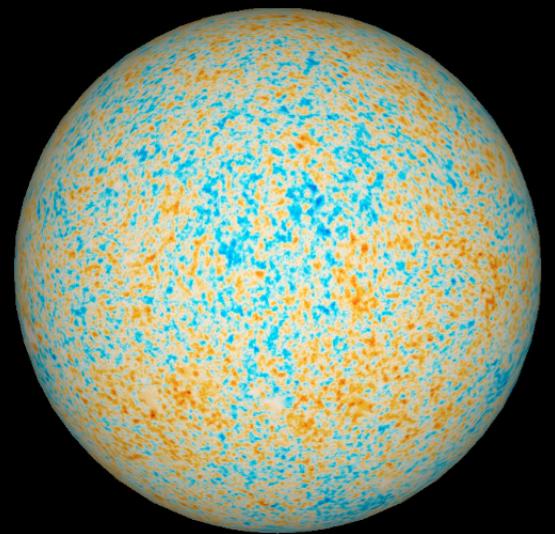
Baby Universe

?

Time

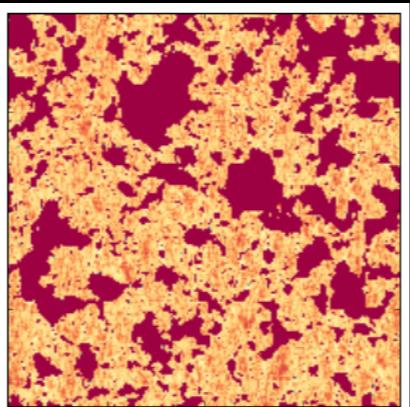


Mature Universe

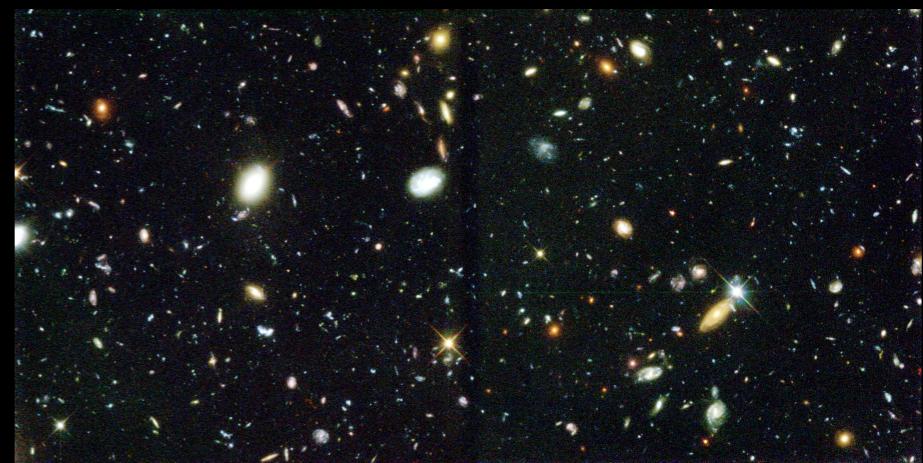


Baby Universe

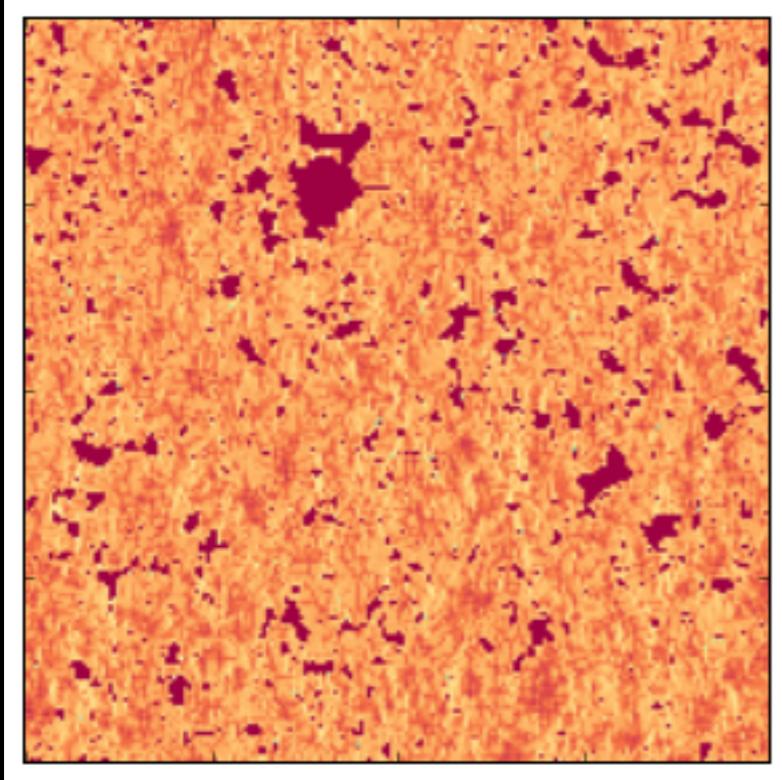
Reionization?



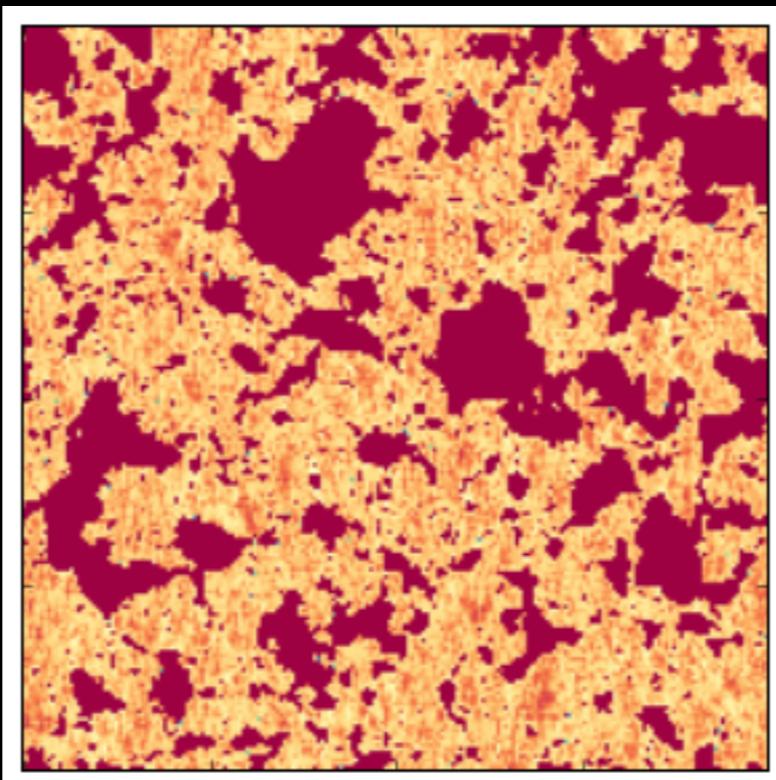
Time



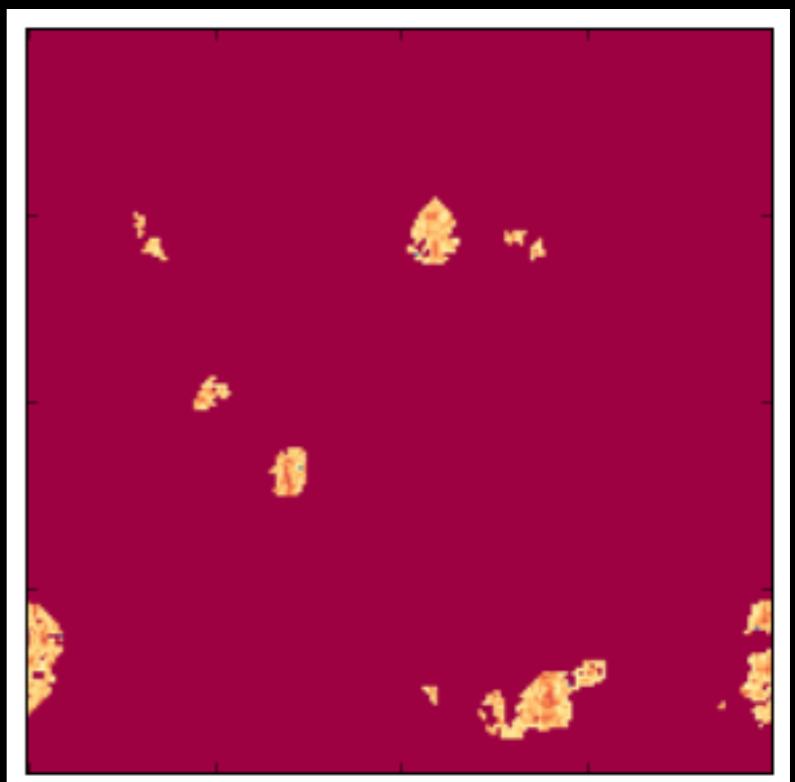
Mature Universe



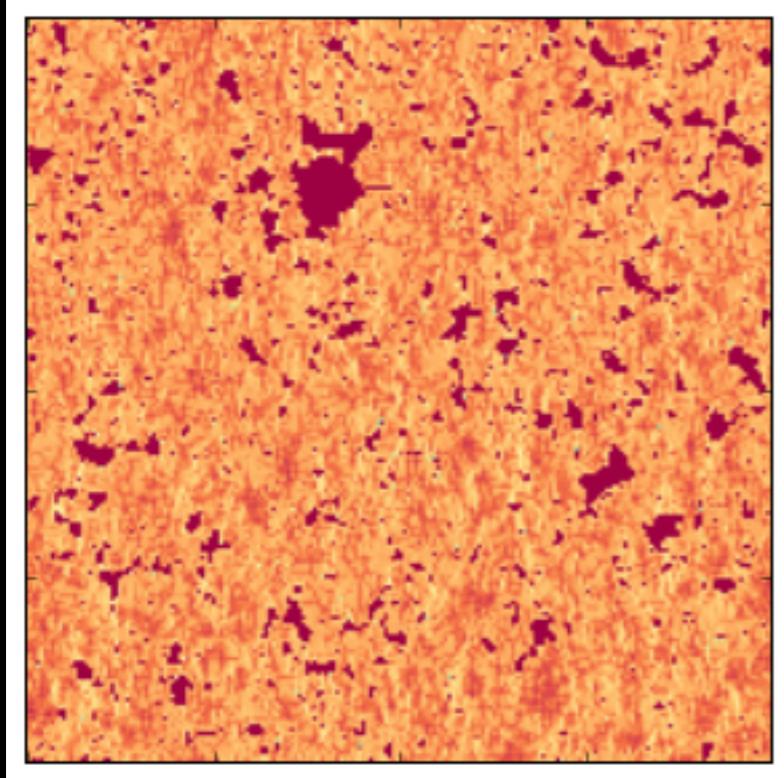
Beginning of  
reionization



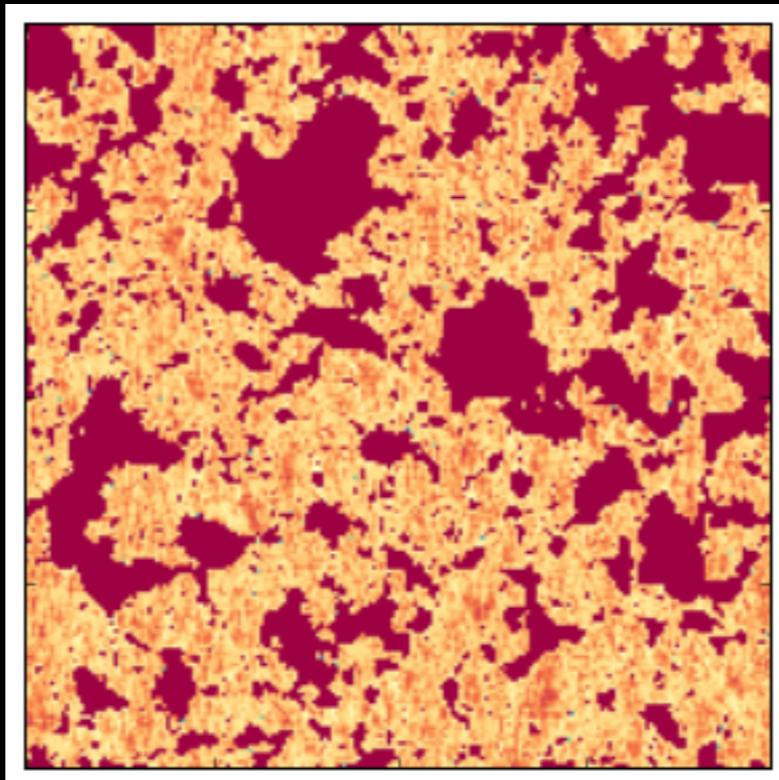
Middle of  
reionization



End of reionization



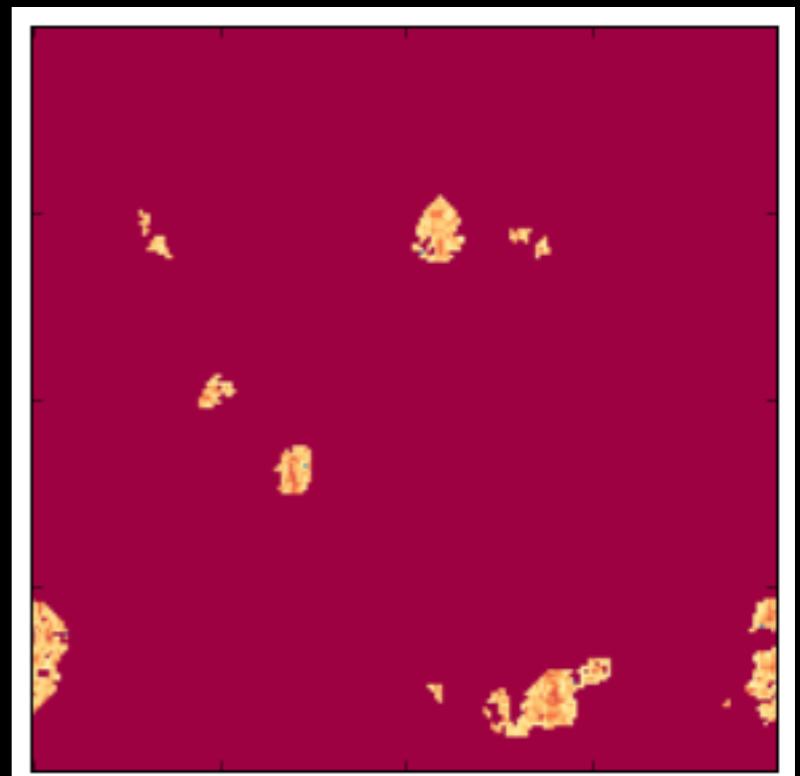
Beginning of  
reionization



Middle of  
reionization

How did the first  
galaxies affect their  
environments?

What was the nature  
of the first galaxies?



End of reionization