

The Radial Acceleration Relation in Rotationally Supported Galaxies

Stacy S. McGaugh and Federico Lelli

*Department of Astronomy, Case Western Reserve University,
10900 Euclid Avenue, Cleveland, OH 44106, USA*

James M. Schombert

Department of Physics, University of Oregon, Eugene, OR 97403, USA
(Dated: September 21, 2016)

+

LA FIN DU MOND? Λ CDM IS FULLY CONSISTENT WITH SPARC ACCELERATION LAW

B.W. KELLER¹ AND J. W. WADSLEY

Department of Physics and Astronomy, McMaster University, Hamilton, Ontario, L8S 4M1, Canada

Spitzer Photometry and Accurate Rotation Curves (SPARC) (McGaugh et al. 2016)

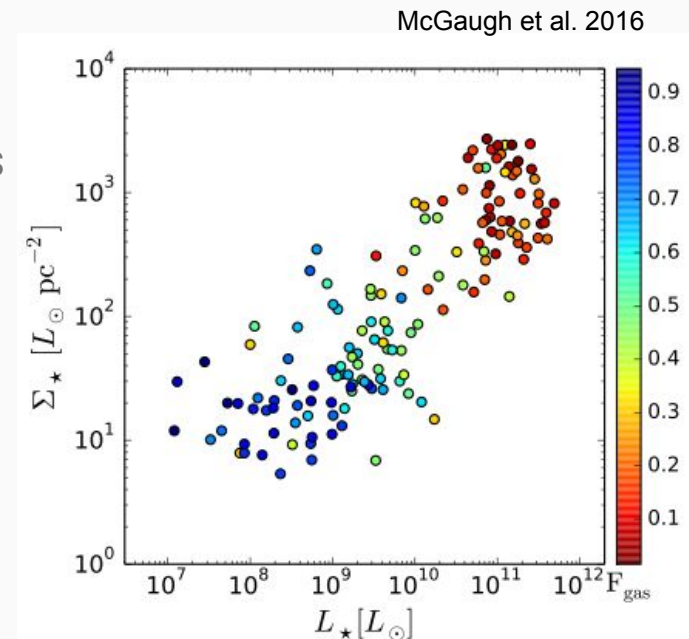
- 175 **rotationally-supported** galaxies (spirals and irregulars)
 - Low redshift
- Near-infrared ($3.6\mu\text{m}$) observations from *Spitzer*
 - Stellar Mass from light
- 21 cm observations
 - Atomic Gas Mass
 - Velocity fields & rotation curves
- Spatially resolved data for gas, stars and rotation curves

Data Quality Control:

- Cut 10 face-on galaxies with $i < 30^\circ$
- Cut 12 galaxies with asymmetric rotation curves
- Total remaining 153 galaxies

SPARC samples:

- Low to high mass
- Low to high surface brightness
- Negligible to dominant gas content



Gravitational Potential of Baryons

- Define baryonic mass as: $M_{\text{bar}} = M_{\star} + M_{\text{gas}}$
- Thus surface density: $\rho_{\text{bar}} = \rho_{\star} + \rho_{\text{gas}}$
- Near-IR data converted to stellar density ρ_{\star}
 - Assumes constant mass-to-light ratio $Y_{\star} \approx 0.5 M_{\odot}/L_{\odot}$ for all rotationally supported galaxies
 - Y_{\star} explored in companion paper; result not sensitive to Y_{\star}
 - By using only one disc value of Y_{\star} , minimize assumptions included in data analysis
 - For excessively bulged galaxies, modeled separate bulge $Y_{\star} \approx 0.7 M_{\odot}/L_{\odot}$ from outer disc
- 21cm data converted to gas density ρ_{HI}
 - Uses hydrogen spin-flip physics, and $\rho_{\text{gas}} = 1.33 \rho_{\text{HI}}$ because of Helium
- Solve Poisson:

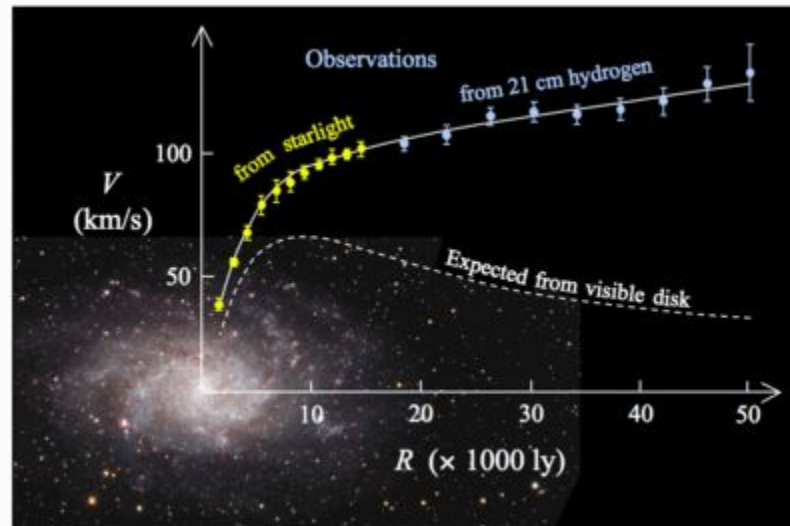
$$\nabla^2 \Phi_{\text{bar}} = 4\pi G \rho_{\text{bar}}$$

- Baryonic acceleration derived only from observed baryons:

$$g_{\text{bar}} = \left| \frac{\partial \Phi_{\text{bar}}}{\partial R} \right|$$

- Presumably, $g_{\text{obs}} = g_{\text{bar}} + g_{\text{DM}}$ if ΛCDM

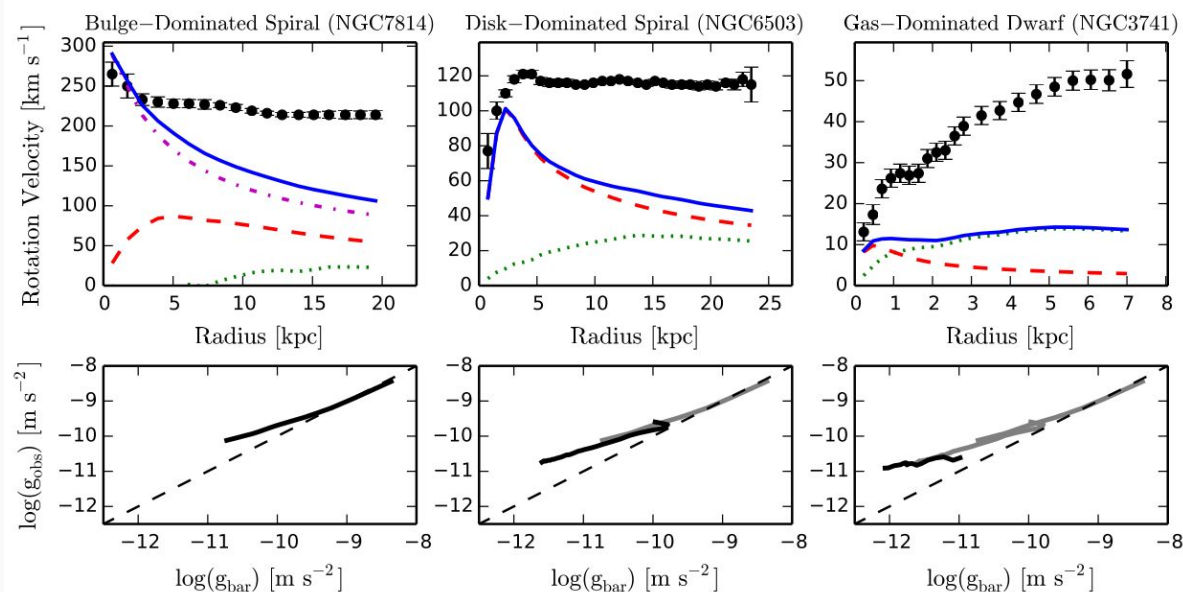
- Rotation Curve Discrepancies
 - Fritz Zwicky 1930s
- Potential Explanations:
 - Dark Matter (DM)
 - Modified Gravity (MOND)
- $g_{\text{obs}} = g_{\text{bar}} + g_{\text{DM}}$



Wikipedia

Results

- Baryonic Mass Model velocity curve: $v_{\text{bar}} = v_{*,\text{disc}} + v_{*,\text{bulge}} + v_{\text{gas}}$
- Observed velocity curves from velocity field data
- Compare derived g_{bar} from stellar + gas mass to observed g_{obs} from rotation curves
- This radial acceleration relation persists for all galaxies of all types in SPARC
- $g_{\text{bar}} - g_{\text{obs}}$ plotted with all galaxies stacked
- Outer regions of high surface brightness galaxies map smoothly to the inner regions of low surface brightness galaxies
- Morphologically agnostic! Same mass discrepancies at same accelerations!



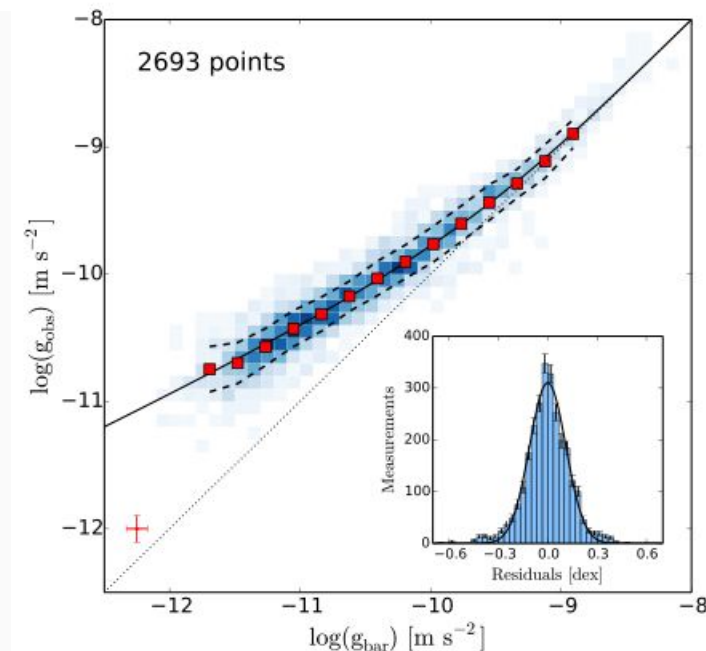
McGaugh et al. 2016

Results

- $g_{\text{obs}} = V^2/R$ plotted against g_{bar} as defined earlier
- ~2700 data points for the 153 galaxies
- Residuals in dex (decimal exponents), just error in \log_{10} units
- **Mean of binned data**
- Dotted = line of unity
- Solid line =

$$g_{\text{obs}} = \mathcal{F}(g_{\text{bar}}) = \frac{g_{\text{bar}}}{1 - e^{-\sqrt{g_{\text{bar}}/g_{\dagger}}}}$$

- Fit param: $g_{\dagger} = (1.20 \pm 0.26) \times 10^{-10} \text{ m/s}^2$
- Uncertainties total at 0.12 dex
 - Υ_* scatter at 0.06 dex
 - Measurement error & calibration are the rest
- **Tight relation, little room for intrinsic scatter**



Discussion

- Observational data leads the theory in this one
- Low intrinsic scatter with all measurement errors
- **SPARC radial acceleration acceleration** empirical with minimal assumptions
- Possible reasons:
 1. The end product of galaxy formation.
 - a. Baryonic feedback, **can be included in DM simulations for comparison**
 2. New dark sector physics that leads to the observed coupling
 - a. DM fluid?
 3. The result of new dynamical laws rather than dark matter (MOND)
 - a. One-to-one relation could mean baryons somehow source the g_{obs}

$$g_{\text{obs}} = \mathcal{F}(g_{\text{bar}}) = \frac{g_{\text{bar}}}{1 - e^{-\sqrt{g_{\text{bar}}/g_{\dagger}}}}$$

~MOND interpolation function, modified gravitational acceleration in **LOW, LOW acceleration regimes**: only reached by galaxies, not seen in solar system/Earth

Since the radial acceleration relation holds across all rotationally-supported morphologies, it is ostensibly a law of nature, a “Kepler’s law for rotating galaxies”

La Fin Du MOND? (Keller & Wadsley 2016)

- Response paper to McGaugh et al. 2016
- Variety of issues currently in Λ CDM:
 - Core/Cusp problem
 - Missing Satellites problem
 - Too Big To Fail problem
- These problems all arise in comparison between DM-only N-body simulations (Λ CDM) and observations
- Issues fixed when baryonic physics included
 - Semi-analytic sub-grid models
- Chief baryonic processes:
 - Black hole feedback
 - Supernovae
- These affect star formation history, thus light/density profile of galaxies

Methodology

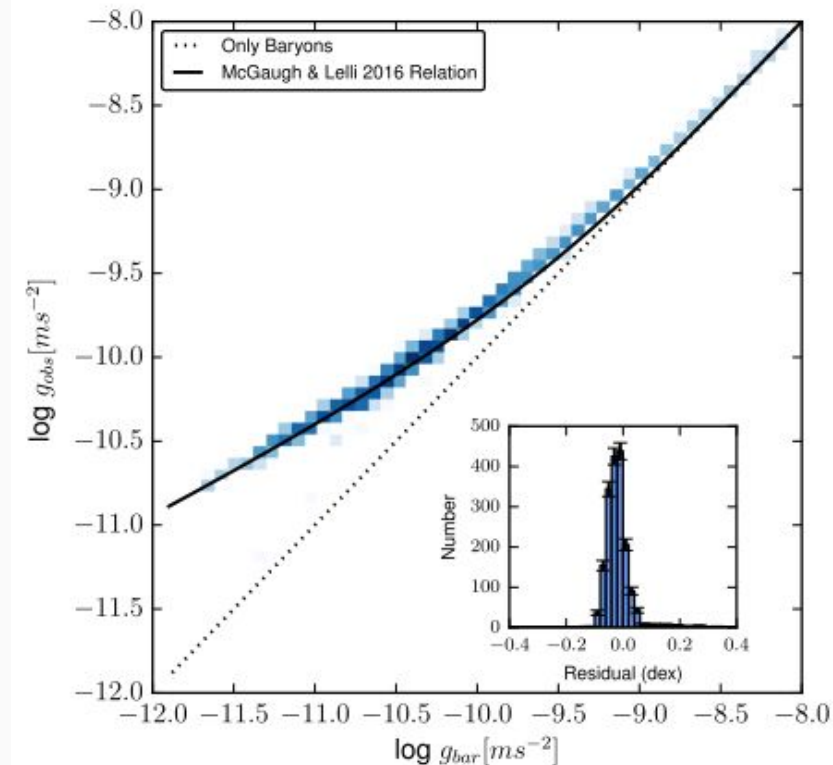
- Conduct same analysis as McGaugh et al. but with Λ CDM+baryonic physics zoomed in DM halos from cosmological simulations
- McMaster Unbiased Galaxy Simulations 2 (MUGS2)
 - WMAP3 Λ CDM cosmology
 - GASOLINE N-body simulator
 - AMIGA halo finder
 - “Physically motivated, first principles model for treating feedback from supernovae”
- Measured rotation curves of simulated galaxies
 - Used DM particle velocities to calculate total angular momentum vector
 - Oriented galaxy such that angular momentum vector perpendicular to x-y plane
- For spatial resolution, divided into 100 annuli each 300 pc thick
- g_{obs} : sum up acceleration contributions from DM, gas, and star particles
- g_{bar} : sum up only contributions from gas and star particles
- Direct summation process identical to numerical solution of Poisson from McGaugh et al.

Results

- $z=0$ simulation datapoints in Hess diagram
- Solid line is the original relation from McGaugh et al.
- Dotted line is line of unity

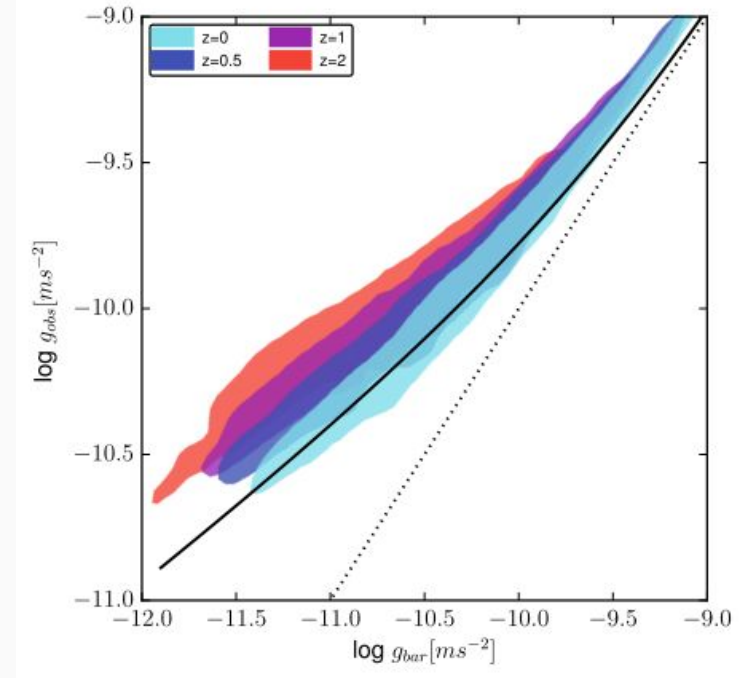
Surprising results:

- Simulations match the empirical relation **better** than the observational data!
- Uncertainty at 0.05 dex
 - Originally estimated at 0.06 dex, agrees!
- Sample of 1800 datapoints
 - $\sim 2/3$ sample size of McGaugh's ~ 2700 points
- Reduced $\chi^2=1.25$



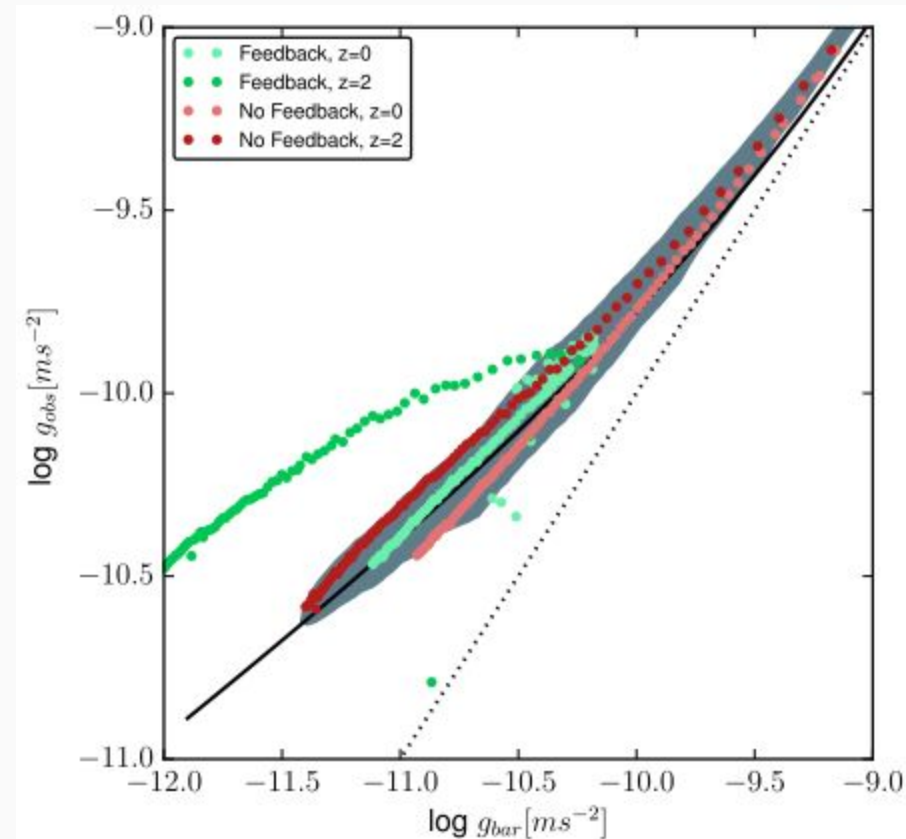
Discussion: No New Physics, Just Λ CDM Hard at Work

- If this relation was due to new, dark sector physics, should show no redshift dependence, because the physics would remain present through universe history
- Would expect redshift dependence if a result of galaxy evolution
- DM simulation snapshots of the same galaxy at different redshifts
 - Sizes scaled by scale factor to remain consistent
- Stellar feedback plays big role in galaxies at $z \sim 2$
 - Supernovae: hot gas outflows
- $z \sim 2$ galaxies depleted of baryons
 - Exactly demonstrated by results!



Discussion: Feedback?

- Simulate same galaxy with and without SNe feedback
- No feedback: no redshift dependence
 - Both $z=0$ and $z=2$ simulation fall within SPARC relation scatter
- Source of redshift dependence must clearly be SNe feedback
- No feedback, only simple dissipative collapse of gas is enough to reproduce SPARC relation



Conclusion

- MOND/SPARC radial acceleration relation can be explained through basic Λ CDM and dissipative gas collapse
- Relation does not hold for all redshifts, since simulations MUST include stellar feedback to accurately represent real galaxies
 - Vigorous feedback in high redshift galaxies flattens $g_{\text{obs}}-g_{\text{bar}}$ relation
- For a universal radial acceleration relation:
 - Would need a fitting parameter $g_+ = g_+(z)$
- **No need for dark sector physics or MOND, just DM!**

Needed:

- High z observational data to look at empirical $g_{\text{obs}}-g_{\text{bar}}$ redshift dependence
- Hope to find redshift dependence as expected
 - If not, Λ CDM may need more work to be done!