

Low-energy extensions of X_{\max} parameterizations

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Based on GAP-2020-058

Motivations

- ▶ Our goal is to broaden the scope of the combined fit in order to cover lower energies down to $E \gtrsim 10^{15}$ eV \rightarrow Galactic-ExtraGalactic transition
- ▶ Existing parametrizations for calculating cosmic ray composition using X_{max} statistics are constrained to energies above $E \gtrsim 10^{17}$ eV (see GAP2020_058)
- ▶ It's crucial to assess whether these current models remain accurate for energies down to 10^{15} eV. Should discrepancies arise, we must consider refining our parametrization models to ensure reliable composition in the knee energy spectrum.

From GAP2020_058:

- ▶ Conex version: version 4.37 for EPOS-LHC and QGSJet II-04, version 7.30 for Sibyll 2.3d
- ▶ Hadronic models used Sibyll 2.3d, EPOS-LHC and QGSJet II-04
- ▶ The log energy range $17 \rightarrow 20$ in 13 fixed $\lg E$ bins with $\Delta \lg E = 0.25$
- ▶ Number of showers 5.4k - 7.7k / bin
- ▶ The X_{\max} used to build the distributions is taken from the **XmxdEdX** branch of the CONEX file
- ▶ Primary nuclei: H, He, N, Si, Ca and Fe

New simulations at CNAF:

- ▶ Conex version: **version 7.60**
- ▶ Hadronic models used Sibyll 2.3d, EPOS-LHC, QGSJet II-04, and **DPM-JET III**
- ▶ The log energy range **$15 \rightarrow 20.5$** in 23 fixed $\lg E$ bins with $\Delta \lg E = 0.25$
- ▶ Number of showers **10k / bin \rightarrow 50k / bin**
- ▶ The X_{\max} used to build the distributions is taken from the **XmxdEdX** branch of the CONEX file
- ▶ Primary nuclei: H, He, N, Si, and Fe

Definitions

▶ Mean:

$$\langle X \rangle = \frac{1}{N} \sum_{i=0}^N X_i \quad (1)$$

▶ Variance:

$$\sigma_X^2 = \frac{1}{N} \sum_{i=0}^N |X_i - \langle X \rangle|^2 \quad (2)$$

▶ Standard Deviation:

$$\sigma_X = \sqrt{\sigma_X^2} \quad (3)$$

▶ Error of the Mean:

$$\epsilon = \frac{\sigma_X}{\sqrt{N}} \quad (4)$$

▶ Error of the Standard Deviation:

$$\rho = \frac{\sigma_X}{\sqrt{2N}} \quad (5)$$

X_{\max} parametrizations

- ▶ We model X_{\max} as a function of $x \equiv \log(E/E_0)$ and $y \equiv \ln A$
- ▶ GAP parametrization (4 free parameters):

$$\begin{aligned}p'_0(y) &= p_0 + \alpha y \\ p'_1(y) &= p_1 + \beta y \\ f(x, y) &= p'_0(y) + p'_1(y)x\end{aligned}$$

which can be re-written as

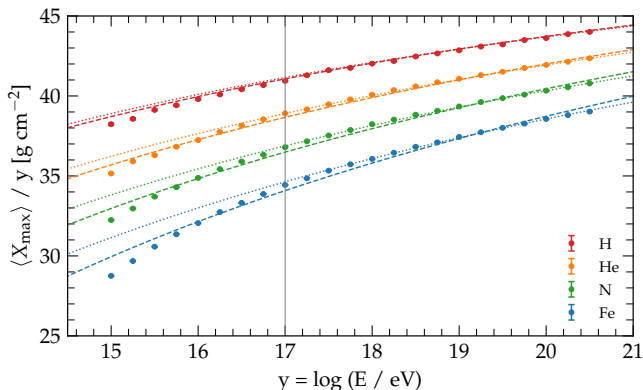
$$f(x, y) = (p_0 + p_1 x) + (\alpha + \beta x)y \equiv f_p(x) + f_E(x)y$$

- ▶ EXT parametrization (6 free parameters):

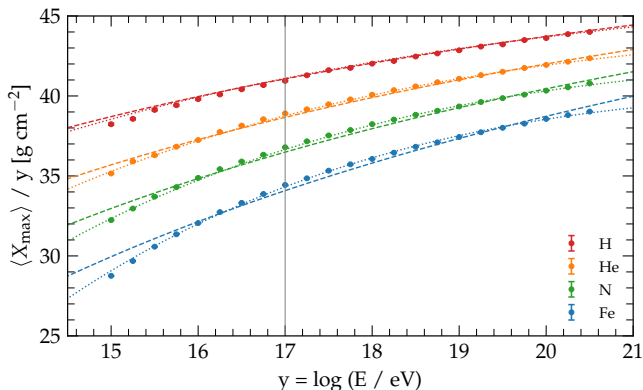
$$\begin{aligned}p'_0(y) &= p_0 + \alpha y \\ p'_1(y) &= p_1 + \beta y \\ p'_2(y) &= p_2 + \gamma y \\ f(x, y) &= p'_0(y) + p'_1(y)x + p'_2(y)x^2\end{aligned}$$

which can be re-written as

$$f(x, y) = (p_0 + p_1 x + p_2 x^2) + (\alpha + \beta x + \gamma x^2)y \equiv f_p(x) + f_E(x)y$$



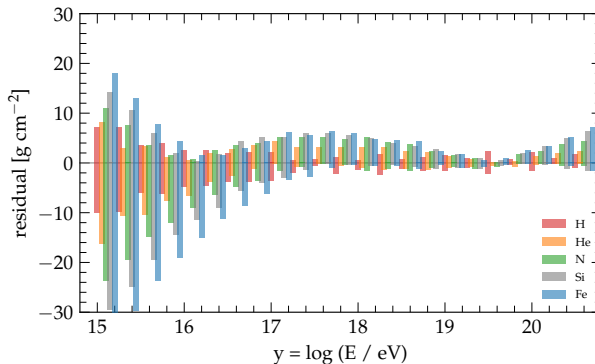
- ▶ Dots: CONEX simulations using Sibyll-23d. Error bars are the **std error of the mean** over N simulations
- ▶ In the y-axis we show X_{\max}/y to emphasize the deviation from the linear evolution
- ▶ Dashed lines: best fit of GAP parametrization assuming $E_{\min} = 10^{15}$ eV
- ▶ Dotted lines: best fit of GAP parametrization assuming $E_{\min} = 10^{17}$ eV



- ▶ Dots: CONEX simulations using Sibyll-23d. Error bars are the **std error of the mean** over N simulations
- ▶ In the y-axis we show X_{\max}/y to emphasize the deviation from the linear evolution
- ▶ Dashed lines: best fit of GAP parametrization assuming $E_{\min} = 10^{15}$ eV
- ▶ Dotted lines: best fit of EXT parametrization assuming $E_{\min} = 10^{15}$ eV

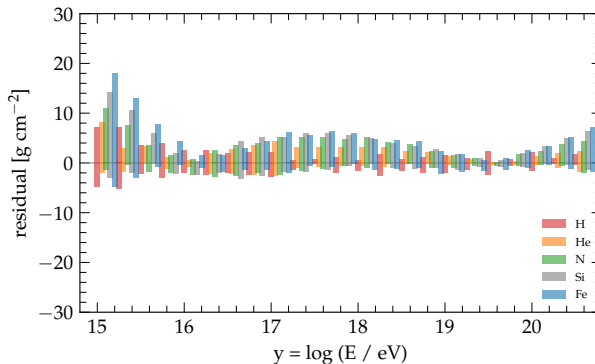
Fit parameters based on Sibyll-2.3D.

	D_0	D_1	D_2	α	β	γ
GAP-full	815.69 ± 0.11	58.76 ± 0.06	-	-26.37 ± 0.03	1.57 ± 0.02	-
GAP-hi	815.83 ± 0.11	58.09 ± 0.11	-	-26.19 ± 0.03	0.67 ± 0.03	-
EXT	816.21 ± 0.12	58.01 ± 0.11	-0.37 ± 0.04	-25.67 ± 0.04	0.56 ± 0.03	-0.46 ± 0.01



- ▶ Positive plane: residuals of the GAP parametrization assuming $E_{\min} = 10^{15}$ eV
- ▶ Negative plane: residuals of the GAP parametrization assuming $E_{\min} = 10^{17}$ eV
- ▶ Mean residuals in g/cm^2 . Si was not included in the fit:

	H	He	N	Si	Fe
GAP-full	2.15	2.45	3.64	4.39	4.94
GAP-hi	3.07	2.95	4.56	5.71	7.11



- ▶ Positive plane: residuals of the GAP parametrization assuming $E_{\min} = 10^{15}$ eV
- ▶ Negative plane: residuals of the EXT parametrization assuming $E_{\min} = 10^{15}$ eV
- ▶ Mean residuals in g/cm^2 . Si was not included in the fit:

	H	He	N	Si	Fe
GAP	2.15	2.45	3.64	4.39	4.94
EXT	1.96	0.98	1.21	1.35	1.37

$\sigma(X_{\max})$ parametrizations

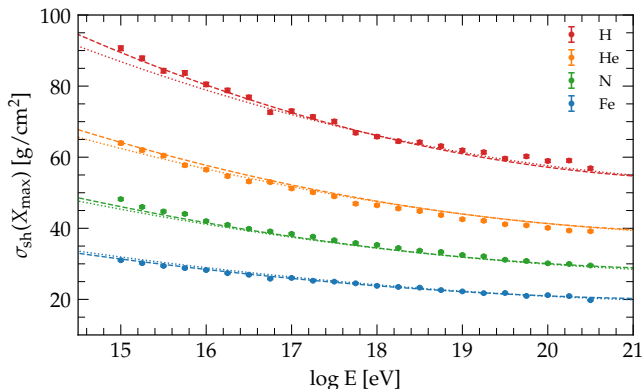
- ▶ We model $\sigma(X_{\max})$ as a function of $x \equiv \log(E/E_0)$ and $y \equiv \ln A$
- ▶ Old parametrization (6 free parameters):

$$a'_0(x) = a_0 + a_1 x$$

$$p(x) = p_0 + p_1 x + p_2 x^2$$

$$f(x, y) = p(x) [1 + a'_0(x)y + b_0 y^2]$$

$\sigma(X_{\max})$ parametrizations



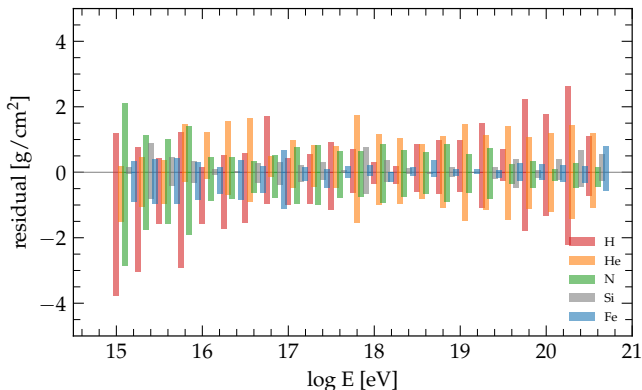
- ▶ Dots: CONEX simulations using Sibyll-23d. Error bars are the **error of the variance** over N simulations
- ▶ Dashed lines: best fit of GAP parametrization assuming $E_{\min} = 10^{15}$ eV
- ▶ Dotted lines: best fit of GAP parametrization assuming $E_{\min} = 10^{17}$ eV

$\sigma^2(X_{\max})$ parametrizations

	p0	p1	p2	a0	a1	b
full	61.0 ± 0.1	-4.5 ± 0.1	0.67 ± 0.02	-0.223 ± 0.001	0.0008 ± 0.0001	0.0161 ± 0.0002
hi	61.3 ± 0.1	-4.3 ± 0.1	0.53 ± 0.07	-0.228 ± 0.001	-0.0002 ± 0.0002	0.0173 ± 0.0003

Fit parameters based on Sibyll-2.3D.

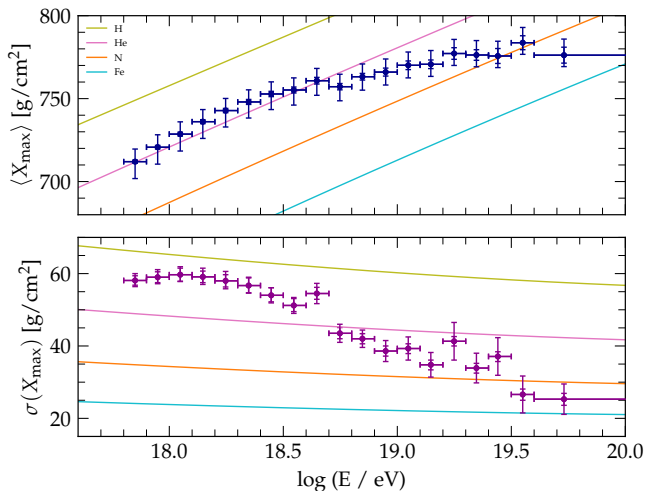
$\sigma(X_{\max})$ residuals



- ▶ Positive plane: residuals of the GAP parametrization assuming $E_{\min} = 10^{15} \text{ eV}$
- ▶ Negative plane: residuals of the GAP parametrization assuming $E_{\min} = 10^{17} \text{ eV}$
- ▶ Mean residuals in g/cm^2 . Si was not included in the fit:

	H	He	N	Si	Fe
GAP full	0.97	1.08	0.68	0.29	0.26
GAP hi	1.34	0.93	0.92	0.25	0.45

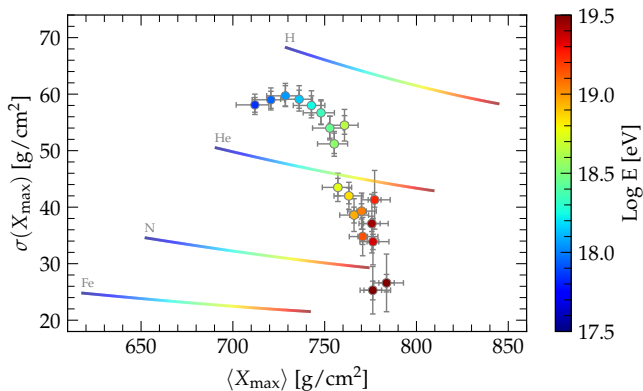
Comparison with Auger data



Based on Sibyll-2.3D.

P.Auger Coll., ICRC 2019, internal use only

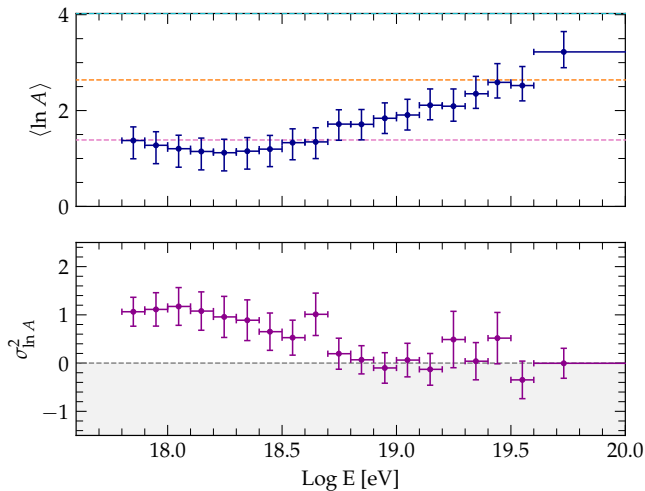
Comparison with Auger data



Based on Sibyll-2.3D.

P.Auger Coll., ICRC 2019, internal use only

Mean $\ln A$ in Auger data



Based on Sibyll-2.3D.

P.Auger Coll., ICRC 2019, internal use only

Variance of $\ln A$ in Auger data

- There are two independent sources of fluctuations: the intrinsic shower-to-shower fluctuations and the $\ln A$ dispersion arising from the mass distribution:

$$\sigma^2(X_{\max}) = \langle \sigma_{\text{sh}}^2 \rangle + \left(\frac{d\langle X_{\max} \rangle}{d \ln A} \right)^2 \sigma_{\ln A}^2 = \langle \sigma_{\text{sh}}^2 \rangle + f_E^2 \sigma_{\ln A}^2 \quad (6)$$

- We assume a parameterization for σ_{sh}^2 as follows

$$\sigma_{\text{sh}}^2(\ln A) = \sigma_p^2 [1 + a \ln A + b(\ln A)^2] \quad (7)$$

therefore

$$\langle \sigma_{\text{sh}}^2 \rangle = \sigma_p^2 [1 + a \langle \ln A \rangle + b \langle (\ln A)^2 \rangle] \quad (8)$$

- After substitution

$$\sigma^2(X_{\max}) = \sigma_p^2 [1 + a \langle \ln A \rangle + b \langle (\ln A)^2 \rangle] + f_E^2 \sigma_{\ln A}^2 \quad (9)$$

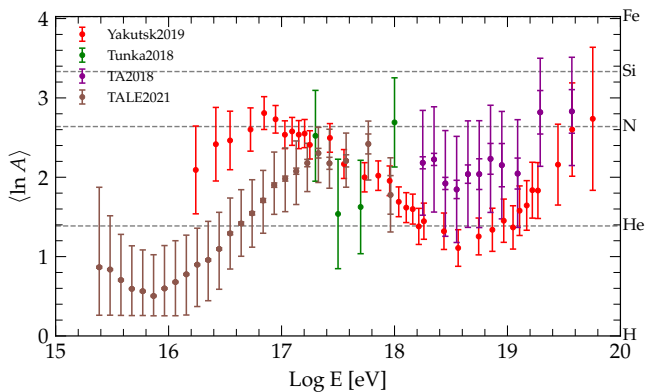
- We apply the definition

$$\langle (\ln A)^2 \rangle = \sigma_{\ln A}^2 + \langle \ln A \rangle^2 \quad (10)$$

- As a result we arrive at

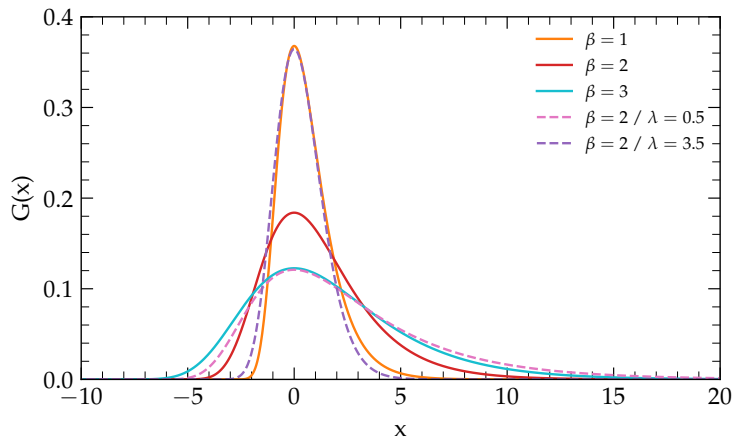
$$\sigma_{\ln A}^2 = \frac{\sigma^2(X_{\max}) - \sigma_{\text{sh}}^2(\langle \ln A \rangle)}{b \sigma_p^2 + f_E^2} \quad (11)$$

Mean $\ln A$ in other datasets



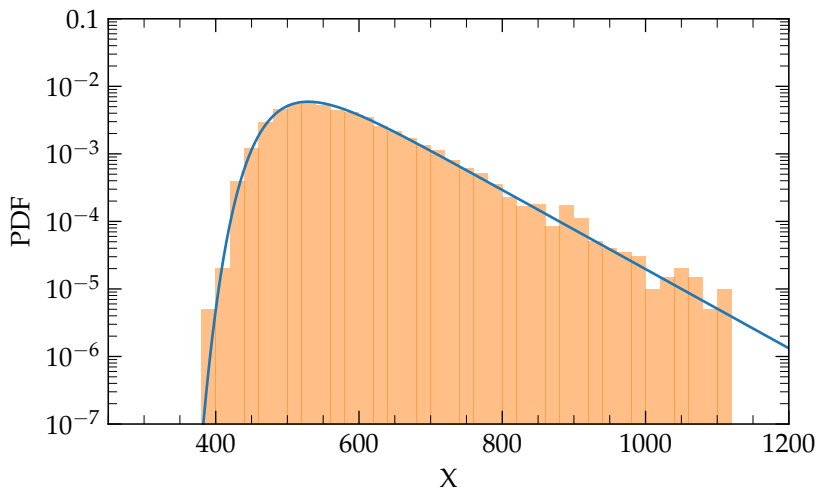
Based on Sibyll-2.3D.

Gumbel function



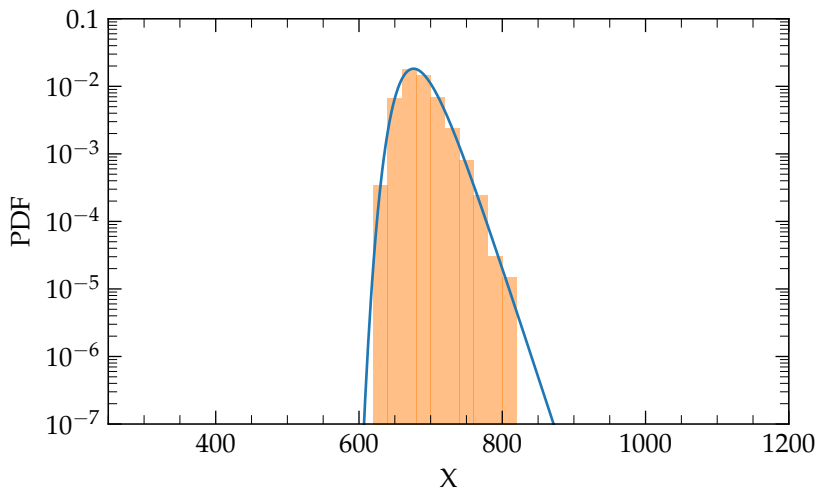
$$G(z) = \frac{1}{\beta} \frac{\lambda^\lambda}{\Gamma(\lambda)} e^{-\lambda(z+e^{-z})}$$

Gumbel function



Mass: H, Energy: 10^{16} eV

Gumbel function



Mass: H, Energy: 10^{20} eV

Parametrizations

- ▷ μ as a function of $x \equiv \log(E/E_0)$ and $y \equiv \ln A$

$$\begin{aligned}p_0^\mu(y) &= a_0^\mu + a_1^\mu y + a_2^\mu y^2 \\p_1^\mu(y) &= b_0^\mu + b_1^\mu y + b_2^\mu y^2 \\p_2^\mu(y) &= c_0^\mu + c_1^\mu y + c_2^\mu y^2 \\\mu(x, y) &= p_0^\mu(y) + p_1^\mu(y)x + p_2^\mu(y)x^2\end{aligned}$$

- ▷ σ as a function of $x \equiv \log(E/E_0)$ and $y \equiv \ln A$

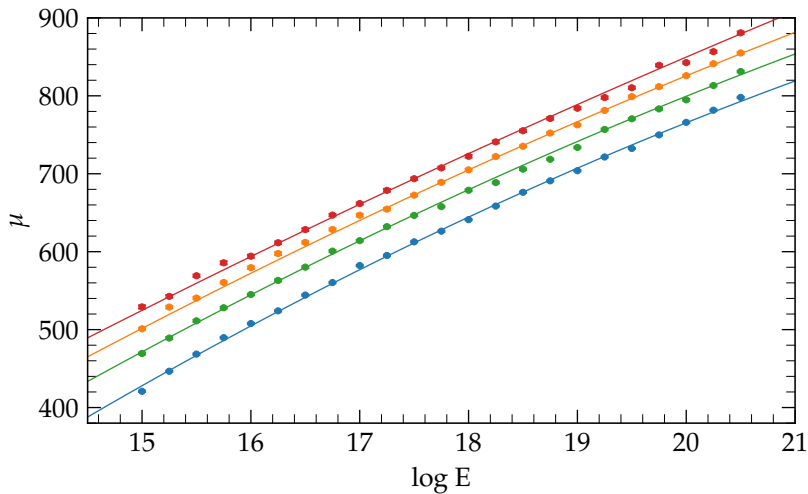
$$\begin{aligned}p_0^\sigma(y) &= a_0^\sigma + a_1^\sigma y + a_2^\sigma y^2 \\p_1^\sigma(y) &= b_0^\sigma + b_1^\sigma y + b_2^\sigma y^2 \\p_2^\sigma(y) &= c_0^\sigma + c_1^\sigma y + c_2^\sigma y^2 \\\sigma(x, y) &= p_0^\sigma(y) + p_1^\sigma(y)x + p_2^\sigma(y)x^2\end{aligned}$$

- ▷ λ as a function of $x \equiv \log(E/E_0)$ and $y \equiv \ln A$

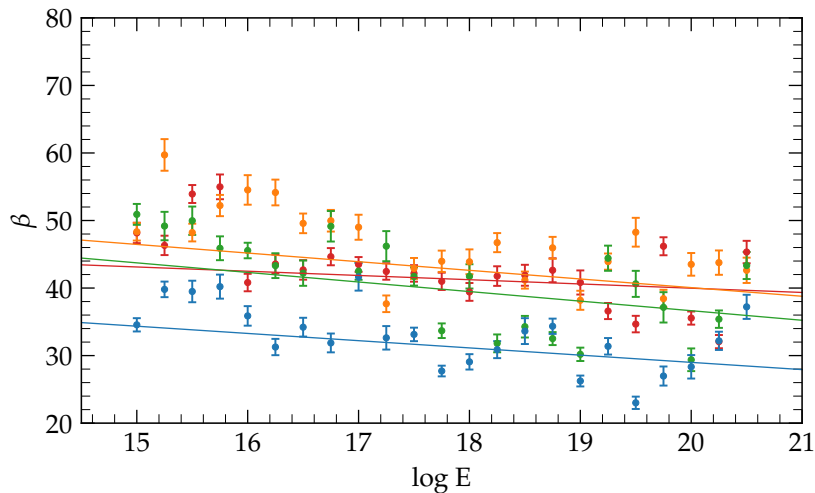
$$\begin{aligned}p_0^\lambda(y) &= a_0^\lambda + a_1^\lambda y + a_2^\lambda y^2 \\p_1^\lambda(y) &= b_0^\lambda + b_1^\lambda y + b_2^\lambda y^2 \\\lambda(x, y) &= p_0^\lambda(y) + p_1^\lambda(y)x + p_2^\lambda(y)x^2\end{aligned}$$

- ▷ 21 free parameters. Cross fingers...

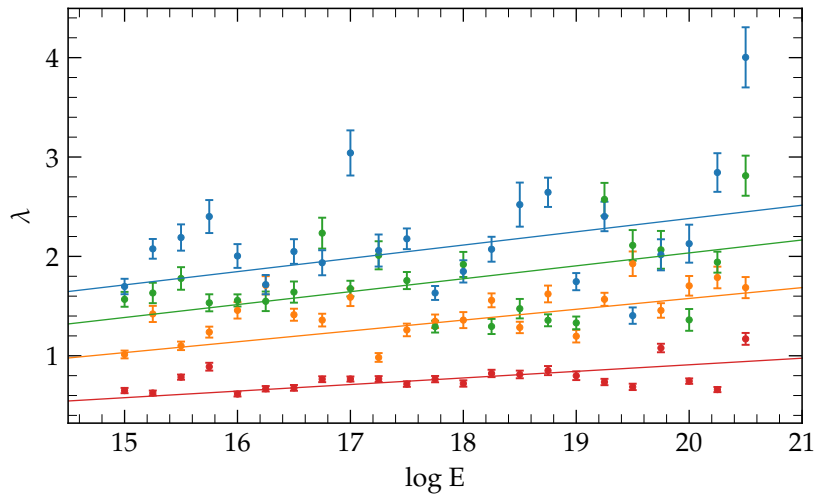
Gumbel function



Gumbel function



Gumbel function



Gumbel function

- ▶ Exponentially Modified Gaussian distribution:
- ▶ Generalized Gumbel distribution:
- ▶ Log-normal distribution:

From Luan B. Arbeletche, Vitor de Souza, Astroparticle Physics 116 (2020) 102389

Next Steps

- ▶ Calculating the parametrization parameters for 4 distinct HIMs → assessing the variations among these models.
- ▶ We plan to augment our dataset of CONEX simulations by 10x. This increase in data volume aims to minimize the scatter in Gumbel parametrizations.
- ▶ We are actively testing additional parametrization methods rather than Gumbel functions.
- ▶ Analysis codes and simulated X_{\max} databases will be public online on GitHub