# QUANTUM MECHANICS NOTES

#### CARSON JAMES

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#### 1. Introduction

## 1.1. Schrödinger Equation.

**Note 1.1.** In the introduction, we keep position general with  $x \in \mathbb{R}^n$  given by  $x = (x_1, \dots, x_n)$ , the usual math notation. The notation in cartesian coordinates changes in three dimensions to r = (x, y, z), the usual physics notation.

**Definition 1.2.** A particle with potential energy V(x,t) is completely described by its **position wavefunction**  $\Psi(x,t)$ , which satisfies the **Schrödinger equation**:

$$i\hbar\frac{\partial}{\partial t}\Psi = -\frac{\hbar^2}{2m}\Delta\Psi + V\Psi$$

**Interpretation 1.3.** We interpret  $|\Psi(x,t)|^2$  to be the **probability density** for the position, x, of the particle at time t. Therefore, we require that for each  $t \in \mathbb{R}$ ,

$$\int_{\mathbb{R}^n} \Psi(x,t)^* \Psi(x,t) dx = 1$$

## 1.2. Operators.

**Definition 1.4.** We define the  $j^{th}$  position and momentum coordinate operators  $X_j, P_j$ , (in position space) by

$$X_i \Psi(x,t) = x_i \Psi(x,t)$$

and

$$P_{j}\Psi(x,t) = -i\hbar \frac{\partial}{\partial x_{j}}\Psi(x,t)$$

We define the **position** and **momentum** operators, X and P, by

$$X = (X_1, X_2, \cdots, X_n)$$

and

$$P = (P_1, P_2, \cdots, P_n)$$

We denote  $P \cdot P$  by  $P^2$ . Note that

$$P^2 = -\hbar^2 \Delta$$

If the partical has potential energy V(x,t), we define the **Hamiltonian** operator, H, by

$$H = \frac{P^2}{2m} + V$$

Thus the Schrödinger equation reads

$$i\hbar \frac{\partial}{\partial t} \Psi = H\Psi$$

Note 1.5. If the potential energy doesn't depend on time, we may write

$$H = \frac{P^2}{2m} + V(X)$$

meaning Hamiltonian only depends on the position and momentum operators, X and P. For the rest of these notes, we assume that the potential energy V does not depend on time.

**Definition 1.6.** Let A and B be operators. Then B is said to be the **adjoint** of A if for each  $\Psi_1$ ,  $\Psi_2$ ,

$$\langle \Psi_1 | A \Psi_2 \rangle = \langle B \Psi_1 | \Psi_2 \rangle$$

i.e.

$$\int_{\mathbb{R}^n} \Psi_1^*(A\Psi_2) dx = \int_{\mathbb{R}^n} (B\Psi_1)^* \Psi_2 dx$$

If B is the adjoint of A, we write

$$B = A^{\dagger}$$

Exercise 1.7. Let A be an operator, then

- (1) for each  $\Psi_1, \Psi_2, \langle A\Psi_1 | \Psi_2 \rangle = \langle \Psi_1 | A^{\dagger} \Psi_2 \rangle$
- $(2) (A^{\dagger})^{\dagger} = A$

*Proof.* (1) For wavefunctions  $\Psi_1$ ,  $\Psi_2$ , we have

$$\begin{split} \langle A\Psi_1|\Psi_2\rangle &= \langle \Psi_2|A\Psi_1\rangle^* \\ &= \langle A^\dagger\Psi_2|\Psi_1\rangle^* \quad \text{(by definition)} \\ &= \langle \Psi_1|A^\dagger\Psi_2\rangle \end{split}$$

(2) For each  $\Psi_1, \Psi_2$ , we have that

$$\langle A\Psi_1 | \Psi_2 \rangle = \langle \Psi_1 | A^{\dagger} \Psi_2 \rangle$$
$$= \langle (A^{\dagger})^{\dagger} \Psi_1 | \Psi_2 \rangle$$

This implies that for each  $\Psi_1, \Psi_2$ ,

$$\langle [A - (A^{\dagger})^{\dagger}] \Psi_1, \Psi_2 \rangle = 0$$

Therefore for each  $\Psi_1$ ,

$$\left[A - (A^{\dagger})^{\dagger}\right]\Psi_1 = 0$$

Hence  $\langle A - (A^{\dagger})^{\dagger} = 0$  and  $A = (A^{\dagger})^{\dagger}$ .

**Definition 1.8.** An linear operator Q is **self-adjoint** if

$$Q = Q^{\dagger}$$

**Interpretation 1.9.** For each measurable, observable quantity  $\hat{Q}$ , there is a self-adjoint operator Q whose eigenvalues are the possible measurment values and whose eigenfunctions are the possible states of the system at measurment.

**Exercise 1.10.** The operators  $X_i$ ,  $P_i$  and H are self adjoint.

Hint: for H, use Green's second identity.

*Proof.* Since  $x_i$  is real, clearly

$$\langle \Psi_1 | X_i \Psi_2 \rangle = \langle X_i \Psi_1 | \Psi_2 \rangle$$

Similarly, we have that

$$\langle \Psi_1 | P_j \Psi_2 \rangle = \int_{\mathbb{R}^n} \Psi_1^* \left( -i\hbar \frac{\partial}{\partial x_j} \Psi_2 \right) dx$$

$$= -i\hbar \int_{\mathbb{R}^n} \Psi_1^* \left( \frac{\partial}{\partial x_j} \Psi_2 \right) dx$$

$$= i\hbar \int_{\mathbb{R}_n} \left( \frac{\partial}{\partial x_j} \Psi_1^* \right) \Psi_2 dx \qquad \text{(integration by parts)}$$

$$= \int_{\mathbb{R}^n} \left( -i\hbar \frac{\partial}{\partial x_j} \Psi_1 \right)^* \Psi_2 dx$$

$$= \langle P\Psi_1 | \Psi_2 \rangle$$

Finally

$$\begin{split} \langle \Psi_1 | H \Psi_2 \rangle - \langle H \Psi_1 | \Psi_2 \rangle &= \int_{\mathbb{R}^n} \Psi_1^* \bigg( -\frac{\hbar^2}{2m} \Delta \Psi_2 + V \Psi_2 \bigg) dx - \int_{\mathbb{R}^n} \bigg( -\frac{\hbar^2}{2m} \Delta \Psi_1 + V \Psi_1 \bigg)^* \Psi_2 dx \\ &= \frac{\hbar^2}{2m} \int_{\mathbb{R}^n} (\Delta \Psi_1^*) \Psi_2 - \Psi_1^* (\Delta \Psi_2) dx \\ &= 0 \qquad \text{(Green's second identity)} \end{split}$$

Exercise 1.11. Let Q be a self-adjoint operator. Then

(1) the eigenvalues of Q are real.

(2) the eigenfunctions of Q corresponding to distinct eigenvalues are orthogonal. Proof.

(1) Let  $\lambda$  be an eigenvalue of Q with corresponding eigenfunction  $\Psi$ . Then

$$\lambda \langle \Psi | \Psi \rangle = \langle \Psi | Q \Psi \rangle$$
$$= \langle Q \Psi | \Psi \rangle$$
$$= \lambda^* \langle \Psi | \Psi \rangle$$

Thus  $\lambda = \lambda^*$  and is real

(2) Let  $\lambda_1$  and  $\lambda_2$  be eigenvalues of Q with corresponding eigenfunctions  $\Psi_1$  and  $\Psi_2$ . Suppose that  $\lambda_1 \neq \lambda_2$ . Then

$$\lambda_2 \langle \Psi_1 | \Psi_2 \rangle = \langle \Psi_1 | Q \Psi_2 \rangle$$

$$= \langle Q \Psi_1 | \Psi_2 \rangle$$

$$= \lambda_1 \langle \Psi_1 | \Psi_2 \rangle$$

So  $(\lambda_2 - \lambda_1)\langle \Psi_1 | \Psi_2 \rangle = 0$ . Which implies that  $\langle \Psi_1 | \Psi_2 \rangle = 0$ 

**Definition 1.12.** Let A and B be operators. The **commutator** of A and B, [A, B], is defined by

$$[A, B] = AB - BA$$

Exercise 1.13. We have  $[X_i, P_j] = \delta_{i,j} i\hbar$ .

*Proof.* For a position wave function  $\Psi$ ,

$$\begin{split} [X_j,P_j]\Psi(x,t) &= [x_j,-i\hbar\frac{\partial}{\partial x_j}]\Psi(x,t) \\ &= (-i\hbar)\left[x_j\frac{\partial}{\partial x_j}\Psi(x,t) - \frac{\partial}{\partial x_j}x_j\Psi(x,t)\right] \\ &= (-i\hbar)\left[x_j\frac{\partial}{\partial x_j}\Psi(x,t) - \Psi(x,t) - x_j\frac{\partial}{\partial x_j}\Psi(x,t)\right] \\ &= i\hbar\Psi(x,t) \end{split}$$

Hence  $[X_j, P_j] = i\hbar$ 

For  $i \neq j$ ,

$$X_{i}P_{j}\Psi(x,t) = \frac{\partial}{\partial x_{j}}x_{i}\Psi(x,t)$$
$$= -i\hbar x_{i}\frac{\partial}{\partial x_{j}}\Psi(x,t)$$
$$= P_{j}X_{i}\Psi(x,t)$$

So

$$[X_i, P_j] = 0$$

**Exercise 1.14.** Let A, B and C be operators, then [AB, C] = A[B, C] + [A, C]B

*Proof.* We have

$$[AB, C] = ABC - CAB$$

$$= ABC - ACB + ACB - CAB$$

$$= A(BC - CB) + (AC - CA)B$$

$$= A[B, C] + [A, C]B$$

## 1.3. Continuity Equation.

**Exercise 1.15.** If V is real and  $\Psi$  satisfies the Schrödinger equation, then

$$i\hbar \frac{\partial}{\partial t} \Psi^* = -H\Psi^*$$

*Proof.* We have that

$$\begin{split} i\hbar\frac{\partial}{\partial t}\Psi^* &= \left(-i\hbar\frac{\partial}{\partial t}\Psi\right)^* \\ &= \left(-\left[-\frac{\hbar^2}{2m}\Delta\Psi + V\Psi\right]\right)^* \\ &= -\left[-\frac{\hbar^2}{2m}\Delta\Psi^* + V\Psi^*\right] \\ &= -H\Psi^* \end{split}$$

Exercise 1.16. We have that

$$\frac{\partial}{\partial t}(\Psi^*\Psi) + \frac{\hbar}{2mi}\nabla\cdot \left[\Psi^*(\nabla\Psi) - (\nabla\Psi^*)\Psi\right] = 0$$

Proof.

$$\begin{split} \frac{\partial}{\partial t}(\Psi^*\Psi) &= \left(\frac{\partial}{\partial t}\Psi^*\right)\Psi + \Psi^*\left(\frac{\partial}{\partial t}\Psi\right) \\ &= \left(\frac{\hbar}{2mi}(\Delta\Psi^*)\Psi - \frac{1}{i\hbar}V\Psi^*\Psi\right) + \left(-\frac{\hbar}{2mi}\Psi^*(\Delta\Psi) + \frac{1}{i\hbar}V\Psi^*\Psi\right) \\ &= \frac{\hbar}{2mi}\bigg[(\Delta\Psi^*)\Psi - \Psi^*(\Delta\Psi)\bigg] \\ &= -\frac{\hbar}{2mi}\bigg[\Psi^*(\Delta\Psi) - (\Delta\Psi^*)\Psi\bigg] \\ &= -\frac{\hbar}{2mi}\nabla\cdot\bigg[\Psi^*(\nabla\Psi) - (\nabla\Psi^*)\Psi\bigg] \end{split}$$

Therefore

$$\frac{\partial}{\partial t}(\Psi^*\Psi) + \frac{\hbar}{2mi}\nabla\cdot \left[\Psi^*(\nabla\Psi) - (\nabla\Psi^*)\Psi\right] = 0$$

**Definition 1.17.** We define the **probability current density**, j, of the particle to be

$$j = \frac{\hbar}{2mi} \left[ \Psi^*(\nabla \Psi) - (\nabla \Psi^*) \Psi \right]$$

## 1.4. Position and Momentum Space.

**Definition 1.18.** We define the **momentum wavefunction**,  $\Phi$ , of the particle to be the Fourier transform of the position wavefunction:

$$\begin{split} \Phi(p,t) &= F[\Psi](p,t) \\ &= \frac{1}{(2\pi\hbar)^{n/2}} \int_{\mathbb{P}^n} \Psi(x,t) e^{-i\frac{p\cdot x}{\hbar}} dx \end{split}$$

Note 1.19. We recall the following facts about Fourier transforms:

(1)
$$\Phi(p,t) = \frac{1}{(2\pi\hbar)^{n/2}} \int_{\mathbb{R}^n} \Psi(x,t) e^{-i\frac{p\cdot x}{\hbar}} dx$$
and
$$\Psi(x,t) = \frac{1}{(2\pi\hbar)^{n/2}} \int_{\mathbb{R}^n} \Phi(p,t) e^{i\frac{p\cdot x}{\hbar}} dp$$
(2)
$$F\left[\frac{\partial}{\partial x_j}\Psi\right] = \frac{ip_j}{\hbar} F[\Psi]$$
and
$$F^{-1}\left[\frac{\partial}{\partial p_j}\Phi\right] = -\frac{ix_j}{\hbar} F[\Psi]$$
(3)
$$\int_{\mathbb{R}^n} \Psi_1^* \Psi_2 dx = \int_{\mathbb{R}^n} F[\Psi_1]^* F[\Psi_2] dx$$

**Note 1.20.** Let Q(X, P) be a self-adjoint operator. Then the properties of the Fourier transform inmply that:

$$Q(X,P) = \begin{cases} Q(x,-i\hbar\nabla) & (position\ space) \\ Q(i\hbar\nabla,p) & (momentum\ space) \end{cases}$$

**Exercise 1.21.** If  $\Psi$  satisfies the Schrödinger equation, then  $\Phi$  satisfies

$$i\hbar \frac{\partial}{\partial t} \Phi = \frac{p^2}{2m} \Phi + V(i\hbar \nabla) \Phi$$

*Proof.* Starting with the Schrödinger equation, we have

$$i\hbar \frac{\partial}{\partial t} \Psi = \left[ \frac{P^2}{2m} + V(X) \right] \Psi$$
$$= \left[ \frac{-\hbar^2}{2m} \Delta + V(x) \right] \Psi \qquad \text{(position space)}$$

Taking Fourier transforms of both sides, we see that

$$\begin{split} i\hbar\frac{\partial}{\partial t}\Phi &= \left[\frac{P^2}{2m} + V(X)\right]\Phi \\ &= \left[\frac{p^2}{2m} + V(i\hbar\nabla)\right]\Phi \qquad \text{(position space)} \end{split}$$

**Interpretation 1.22.** We interpret  $|\Phi(p,t)|^2$  to be the probability density for the momentum, p, of the particle at time t.

Note 1.23. For a self-adjoint operator Q(X,P), the expected value of Q, is given by

$$\langle Q \rangle = \begin{cases} \langle \Psi(x,t) | Q(x,-i\hbar\nabla)\Psi(x,t) \rangle & (position \ space) \\ \langle \Phi(p,t) | Q(i\hbar\nabla,p)\Phi(p,t) \rangle & (momentum \ space) \end{cases}$$

### 1.5. Stationary States.

**Definition 1.24.** When the potential energy V doesn't depend on time, we look for solutions to the Schrödinger equation of the form

$$\Psi(x,t) = \psi(x)\varphi(t)$$

With a closer look, we find that

(1) 
$$H\psi = E\psi$$

(2) 
$$\varphi(t) = e^{-i\frac{E}{\hbar}t}$$

Statement (1) is referred to as the **time-independent Schrödinger equation**. Eigenfuntions of the Hamiltonian operator are called **stationary states**. If the possible eigenvalues for the Hamiltonian operator are discreet  $(E_n)_{n\in\mathbb{N}}$  with stationary states  $(\psi_n)_{n\in\mathbb{N}}$ , then the general solution to the Schrödinger equation is

$$\Psi(x,t) = \sum_{n \in \mathbb{N}} c_n \psi_n(x) e^{-i\frac{E_n}{\hbar}t}$$

where

$$c_n = \int_{\mathbb{R}^n} \psi_n^*(x) \Psi(x, 0) dx$$

**Definition 1.25.** An energy eigenvalue  $E_n$  of H is said to have a **degeneracy of degree** k if it corresponds to k orthonormal stationary states.

**Note 1.26.** If the energy eigenvalues  $(E_n)_{n\in\mathbb{N}}$  have degeneracies of degrees  $(k_n)_{n\in\mathbb{N}}$  with corresponding orthonormal stationary states  $(\psi_{n,j})_{j=1}^{k_n}$  and

$$\Psi(x,t) = \sum_{n \in \mathbb{N}} \sum_{j=1}^{k_n} c_{n,j} \psi_{n,j}(x) e^{-i\frac{E_n}{\hbar}t}$$

Then the probability of measuring the energy  $E_n$  is

$$\mathbb{P}(E_n) = \sum_{j=1}^{k_n} |c_{n,j}|^2$$

**Definition 1.27.** If the spectrum of the Hamiltonian is discreet, the stationary state with the least energy is called the **ground state**. The stationary states that are not the ground state are called **excited states**.

#### 2. Fundamental Examples in One Dimension

## 2.1. The Infinite Square Well.

**Definition 2.1.** The infinite square well is defined by the potential

$$V(x) = \begin{cases} \infty & x \in I_1 = (-\infty, a] \\ 0 & x \in I_2 = (0, a) \\ \infty & x \in I_3 = [a, \infty) \end{cases}$$

Exercise 2.2. By starting with a finite potental well and letting the height of the well go to infinity, show that the stationary states and their energies are given by

$$\psi_n(x) = \begin{cases} \sqrt{\frac{2}{a}} \sin(\frac{n\pi}{a}x) & x \in (0, a) \\ 0 & x \notin (0, a) \end{cases}$$

and

$$E_n = \frac{n^2 \pi^2 \hbar^2}{2ma^2}$$

*Proof.* Define

$$V_{\alpha}(x) = \begin{cases} \alpha & x \in I_1 \\ 0 & x \in I_2 \\ \alpha & x \in I_3 \end{cases}$$

For the potential energy  $V_{\alpha}$ , in sections  $I_1, I_3$  the Schrödinger equation may be written as

$$\frac{\mathrm{d}^2 \psi}{\mathrm{d}x^2} = \frac{2m}{\hbar^2} (\alpha - E)\psi$$

Assuming  $\alpha > E$ , we may write  $l = \frac{\sqrt{2m(\alpha - E)}}{\hbar}$  and substitute to get

$$\frac{\mathrm{d}^2 \psi}{\mathrm{d}x^2} = l^2 \psi$$

Thus in region  $I_1$ ,  $\psi_1(x) = Ae^{lx} + Be^{-lx}$  and in region  $I_3$ ,  $\psi_3(x) = Fe^{lx} + Ge^{-lx}$ . Since  $e^{-lx}$  blows up as  $x \to -\infty$ , B = 0. Since  $e^{lx}$  blows up as  $x \to \infty$ , F = 0.

In section  $I_2$ , the Schrödinger equation may be written as

$$\frac{\mathrm{d}^2 \psi}{\mathrm{d}x^2} = -\frac{2mE}{\hbar^2}\psi$$

We write  $k = \frac{\sqrt{2mE}}{\hbar}$  and substitute to get

$$\frac{\mathrm{d}^2 \psi}{\mathrm{d}x^2} = -k^2 \psi$$

Hence in region  $I_2$ ,  $\psi_2(x) = C \sin(kx) + D \cos(kx)$ .

So far we have

$$\psi_{\alpha}(x) = \begin{cases} Ae^{lx} & x \in I_1 \\ C\sin(kx) + D\cos(kx) & x \in I_2 \\ Ge^{-lx} & x \in I_3 \end{cases}$$

To find possible wavefunctions  $\psi$  for the infinite potential, we let  $\alpha \to \infty$ . As  $\alpha \to \infty$ , we have that  $l \to \infty$ . Hence  $\psi_1 \to 0$  and  $\psi_3 \to 0$ . So for the infinite potential,

$$\psi(x) = \begin{cases} C\sin(kx) + D\cos(kx) & x \in (0, a) \\ 0 & x \notin (0, a) \end{cases}$$

By continuity at the points x = 0 and x = a, we see that  $0 = C\sin(0) + D\cos(0)$  which implies that D = 0 and  $0 = C\sin(ka)$  which yields various solutions

$$k_n = \frac{n\pi}{a} \quad n \in \mathbb{Z}$$

To avoid non-normalizable solutions or linearly dependent solutions, we restrict  $n \in \mathbb{N}$ . Our energies are then

$$E_n = \frac{\hbar^2 k_n^2}{2m} = \frac{\hbar^2 n^2 \pi^2}{2ma^2} \quad n \in \mathbb{N}$$

and (after normalizing) our stationary states are

$$\psi_n(x) = \begin{cases} \sqrt{\frac{2}{a}} \sin(\frac{n\pi}{a}x) & x \in (0, a) \\ 0 & x \notin (0, a) \end{cases}$$

#### 2.2. The Harmonic Oscillator.

**Definition 2.3.** The **harmonic oscillator** in one dimension is defined by the potential energy:

$$V(x) = \frac{1}{2}m\omega^2 x^2$$

We define the **lowering operator**,  $a_{-}$ , by

$$a_{-} = \frac{1}{\sqrt{2\hbar m\omega}} \left( m\omega X + iP \right)$$

and we define the **raising operator**,  $a_+$ , by

$$a_{+} = \frac{1}{\sqrt{2\hbar m\omega}} \left( m\omega X - iP \right)$$

Exercise 2.4. The adjoint of the lowering operator is the raising operator:

$$(a_-)^{\dagger} = a_+$$

*Proof.* Let  $\Psi_1$ ,  $\Psi_2$  be wavefunctions. Since X, P are self-adjoint, we have that

$$\begin{split} \langle \Psi_1 | a_- \Psi_2 \rangle &= \frac{1}{\sqrt{2\hbar m \omega}} \langle \Psi_1 | (m \omega X + i P) \Psi_2 \rangle \\ &= \frac{1}{\sqrt{2\hbar m \omega}} \bigg[ m \omega \langle \Psi_1 | X \Psi_2 \rangle + i \langle \Psi_1 | P \Psi_2 \rangle \bigg] \\ &= \frac{1}{\sqrt{2\hbar m \omega}} \bigg[ \langle m \omega X \Psi_1 | \Psi_2 \rangle + \langle -i P \Psi_1 | \Psi_2 \rangle \bigg] \\ &= \frac{1}{\sqrt{2\hbar m \omega}} \langle (m \omega X - i P) \Psi_1 | \Psi_2 \rangle \\ &= \langle a_+ \Psi_1 | \Psi_2 \rangle \end{split}$$

Exercise 2.5. We have that

(1) 
$$a_{-}a_{+} = \frac{1}{\hbar\omega}H + \frac{1}{2}$$
  
(2)  $a_{+}a_{-} = \frac{1}{\hbar\omega}H - \frac{1}{2}$   
(3)  $[a_{-}, a_{+}] = 1$ 

(3) 
$$[a_-, a_+] = 1$$

Proof. (1)

$$a_{-}a_{+} = \frac{1}{2\hbar m\omega} (m\omega X + iP) (m\omega X - iP)$$

$$= \frac{1}{2\hbar m\omega} \left[ (m^{2}\omega^{2}X^{2} + P^{2}) - m\omega i (XP - PX) \right]$$

$$= \frac{1}{\hbar\omega} (\frac{1}{2m}P^{2} + \frac{1}{2}m\omega^{2}X^{2}) - \frac{i}{2\hbar} [X, P]$$

$$= \frac{1}{\hbar\omega} H + \frac{1}{2}$$

- (2) Similar
- (3) Trivial

Exercise 2.6. If  $H\psi = E\psi$ , then

(1) 
$$Ha_-\psi = (E - \hbar\omega)a_-\psi$$

(2) 
$$Ha_+\psi = (E + \hbar\omega)a_+\psi$$

Proof.

$$\begin{split} Ha_-\psi &= \hbar\omega \bigg(a_-a_+ - \frac{1}{2}\bigg)a\psi \\ &= \hbar\omega \bigg(a_-a_+a_- - \frac{1}{2}a_-\bigg)\psi \\ &= \hbar\omega a_- \bigg(a_+a_- - \frac{1}{2}\bigg)\psi \\ &= \hbar\omega a_- \bigg(a_+a_- + \frac{1}{2} - 1\bigg)\psi \\ &= \hbar\omega a_- \bigg(\frac{1}{\hbar\omega}H - 1\bigg)\psi \\ &= a_-H\psi - \hbar\omega a_-\psi \\ &= (E - \hbar\omega)a_-\psi \end{split}$$

# (2) Similar

**Interpretation 2.7.** The lowering operator "lowers" a stationary state  $\psi$  with energy E to a stationary state  $a_-\psi$  with energy  $E - \hbar\omega$  and the raising operator "raises" a stationary state  $\psi$  with energy E to a stationary state  $a_+\psi$  with energy  $E + \hbar\omega$ .

**Definition 2.8.** Since the zero function is a solution to the time-independent Schrödinger equation, we define the ground state,  $\psi_0$  of the harmonic oscillator to be the stationary state that satisfies  $a_-\psi_0 = 0$ . The excited states  $\psi_n$ , for  $n \ge 1$ , are obtained by applying the rasing operator n times and then normalizing.

Exercise 2.9. We have that

(1) 
$$\psi_0 = \left(\frac{m\omega}{\pi\hbar}\right)^{\frac{1}{4}} e^{-\frac{m\omega}{2\hbar}x^2}$$

$$(2) E_0 = \frac{1}{2}\hbar\omega$$

(3) 
$$\psi_n = c_n(a_+)^n \psi_0 \quad (for some \ constant \ c_n)$$

$$(4) E_n = \hbar\omega(n + \frac{1}{2})$$

Proof.

(1) The simple differential equation  $a_{-}\psi_{0} = 0$  has the solution

$$\psi_0 = Ae^{-\frac{m\omega}{2\hbar}x^2}$$

Thus

$$|\psi_0|^2 = |A|^2 e^{-\frac{m\omega}{\hbar}x^2}$$

If we normalize this function, we obtain

$$\psi_0 = \left(\frac{m\omega}{\pi\hbar}\right)^{\frac{1}{4}} e^{-\frac{m\omega}{2\hbar}x^2}$$

(2) It is tedious but straightforward to show that

$$H\psi_0 = \frac{1}{2}\hbar\omega\psi_0$$

- (3) Clear by definition.
- (4) Clear by previous exercise.

### Exercise 2.10.

(1) 
$$\psi_{n+1} = \frac{1}{\sqrt{n+1}} a_+ \psi_n$$

(2) 
$$\psi_{n-1} = \frac{1}{\sqrt{n}} a_- \psi_n$$

(1)  $\psi_{n+1} = \frac{1}{\sqrt{n+1}} a_+ \psi_n$ (2)  $\psi_{n-1} = \frac{1}{\sqrt{n}} a_- \psi_n$ Hint: use the adjoint-ness of  $a_-$  and  $a_+$ 

Proof.

(1)

$$a_{-}a_{+}\psi_{n} = \left(\frac{1}{\hbar\omega}H + \frac{1}{2}\right)\psi_{n}$$
$$= \frac{1}{\hbar\omega}E_{n}\psi_{n} + \frac{1}{2}\psi_{n}$$
$$= (n+1)\psi_{n}$$

Since  $\psi_{n+1} = ca_+\psi_n$ , we have that

$$1 = \langle \psi_{n+1} | \psi_{n+1} \rangle$$

$$= \langle ca_{+} \psi_{n} | ca_{+} \psi_{n} \rangle$$

$$= |c|^{2} \langle a_{+} \psi_{n} | a_{+} \psi_{n} \rangle$$

$$= |c|^{2} \langle a_{-} a_{+} \psi_{n} | \psi_{n} \rangle$$

$$= |c|^{2} \langle (n+1) \psi_{n} | \psi_{n} \rangle$$

$$= |c|^{2} (n+1) \langle \psi_{n} | \psi_{n} \rangle$$

$$= |c|^{2} (n+1)$$

So 
$$c = \frac{1}{\sqrt{n+1}}$$

(2) Similar to (1).

**Exercise 2.11.** The  $n^{th}$  stationary state is given by  $\psi_n = \frac{1}{\sqrt{n!}} (a_+)^n \psi_0$ *Proof.* Clear by induction.

Exercise 2.12. Show that

(1) 
$$\psi_1(x) = \left(\frac{4m^3\omega^3}{\hbar^3\pi}\right)xe^{-\frac{m\omega}{2\hbar}x^2}$$
  
(2)  $E_1 = \frac{3}{2}\hbar\omega$ 

$$(2) E_1 = \frac{3}{2}\hbar\omega$$

*Proof.* Straightforward.

**Exercise 2.13.** If particle one is in state  $\psi_0$  at time t=0, then the momentum wave function is

$$\Phi(p,t) = \left(\frac{1}{m\omega\pi\hbar}\right)^{\frac{1}{4}} e^{-\frac{1}{2m\omega\hbar}p^2} e^{-i\frac{\omega}{2}t}$$

*Proof.* By assumption

$$\Psi(x,t) = \psi_0(x)e^{-i\frac{\omega}{2}t}$$

Thus

$$\Phi(p,t) = \frac{1}{\sqrt{2\pi\hbar}} \int_{\mathbb{R}} \Psi(x,t) e^{-i\frac{px}{\hbar}} dx$$

The rest is straightforward.

### 3. Fundamental Examples in Three Dimensions

# 3.1. Spherical Harmonic Oscillator (Cartesian Coordinates).

**Definition 3.1.** The spherical harmonic oscillator (in cartesian coordinates) is defined by the potential energy

$$V(x, y, z) = x^2 + y^2 + z^2$$

Exercise 3.2. In cartesian coordinates, the the stationary states of the harmonic oscillator are given by

$$\psi_{n_x,n_y,n_z}(x,y,z) = \psi_{n_x}(x)\psi_{n_y}(y)\psi_{n_z}(z)$$

with energies

$$E_{n_x,n_y,n_z} = \hbar\omega \left(n_x + n_y + n_z + \frac{3}{2}\right)$$

where  $\psi_{n_x}, \psi_{n_y}, \psi_{n_z}$  are stationary states for the one dimensional harmonic oscillator.

*Proof.* We look for solutions of the form  $\psi(x, y, z) = \psi_x(x)\psi_y(y)\psi_z(z)$ . Plugging this into the time-independent Schrödinger equation, we get

$$-\frac{\hbar^2}{2m} \left[ \frac{\partial^2 \psi_x}{\partial x^2} \psi_y \psi_z + \psi_x \frac{\partial^2 \psi_y}{\partial y^2} \psi_z + \psi_x \psi_y \frac{\partial^2 \psi_z}{\partial z^2} \right] + \frac{1}{2} m \omega^2 (x^2 + y^2 + x^2) \psi = E \psi$$

Dividing both sides by  $\psi$  and rearranging, we obtain

$$\left(-\frac{\hbar^2}{2m}\frac{\partial^2\psi_x}{\partial x^2}\frac{1}{\psi_x} + \frac{1}{2}m\omega^2x^2\right) + \left(-\frac{\hbar^2}{2m}\frac{\partial^2\psi_y}{\partial y^2}\frac{1}{\psi_y} + \frac{1}{2}m\omega^2y^2\right) + \left(-\frac{\hbar^2}{2m}\frac{\partial^2\psi_z}{\partial z^2}\frac{1}{\psi_z} + \frac{1}{2}m\omega^2z^2\right) = E$$

Thus each part is constant and we may write

$$-\frac{\hbar^2}{2m}\frac{\partial^2 \psi_x}{\partial x^2} + \frac{1}{2}m\omega^2 x^2 \psi_x = E_x \psi_x$$
$$-\frac{\hbar^2}{2m}\frac{\partial^2 \psi_y}{\partial y^2} + \frac{1}{2}m\omega^2 y^2 \psi_y = E_y \psi_y$$
$$-\frac{\hbar^2}{2m}\frac{\partial^2 \psi_z}{\partial z^2} + \frac{1}{2}m\omega^2 z^2 \psi_z = E_z \psi_z$$

So we have three one-dimensional harmonic oscillators and we have

$$\psi_x = \psi_{n_x} = \frac{1}{\sqrt{n_x!}} (a_+)^{n_x} \psi_0 \text{ and } E_x = E_{n_x} = \hbar \omega \left( n_x + \frac{1}{2} \right)$$
 $\psi_y = \psi_{n_y} = \frac{1}{\sqrt{n_y!}} (a_+)^{n_y} \psi_0 \text{ and } E_y = E_{n_y} = \hbar \omega \left( n_y + \frac{1}{2} \right)$ 
 $\psi_z = \psi_{n_z} = \frac{1}{\sqrt{n_z!}} (a_+)^{n_z} \psi_0 \text{ and } E_z = E_{n_z} = \hbar \omega \left( n_z + \frac{1}{2} \right)$ 

Thus

$$\psi = \psi_{n_x,n_y,n_z}(x,y,z) = \psi_{n_x}(x)\psi_{n_y}(y)\psi_{n_z}(z)$$

with energy

$$E = E_{n_x, n_y, n_z} = \hbar\omega \left(n_x + n_y + n_z + \frac{3}{2}\right)$$

**Exercise 3.3.** Show that the degree of degeneracy of  $E_n$  is

$$deg(E_n) = \binom{n+2}{2}$$

*Proof.* Stars and bars

Interpretation 3.4. The energies of the three-dimensional harmonic oscillator are given by  $E_n = \hbar\omega\left(n + \frac{3}{2}\right)$  which correspond to  $\binom{n+2}{2}$  stationary states.

### 3.2. Spherical Coordinates.

**Definition 3.5.** We now set n = 3, and work with spherical coordinates  $(r, \theta, \phi)$  where r is the distance in from the origin,  $0 \le \theta \le \pi$  is the angle with initial side on the positive z-axis, and  $0 \le \phi < 2\pi$  is the angle in the x-y plane with initial side on the positive x-axis going towards the positive y-axis.

**Proposition 3.6.** In spherical coordinates, the time independent Schrödinger equation becomes

$$-\frac{\hbar^2}{2m} \left[ \frac{1}{r^2} \frac{\partial}{\partial r} \left( r^2 \frac{\partial}{\partial r} \right) + \frac{1}{r^2 \sin \theta} \frac{\partial}{\partial \theta} \left( \sin \theta \frac{\partial}{\partial \theta} \right) + \frac{1}{r^2 \sin^2 \theta} \frac{\partial^2}{\partial \phi^2} \right] \psi + V \psi = E \psi$$

**Definition 3.7.** If the potential energy V only depends on r, then we can solve for stationary solutions of the form  $\psi(r, \theta, \phi) = R(r), Y(\theta, \phi)$ . It results that there is some constant l such that

(1) 
$$\frac{1}{R}\frac{d}{dr}r^{2}\frac{dR}{dr} - \frac{2m}{\hbar^{2}}r^{2}(V-E) = l(l+1)$$

(2) 
$$\frac{1}{Y} \left[ \frac{1}{\sin \theta} \frac{\partial}{\partial \theta} \left( \sin \theta \frac{\partial Y}{\partial \theta} \right) + \frac{1}{\sin^2 \theta} \frac{\partial^2 Y}{\partial \phi^2} \right] = -l(l+1)$$

The number l is called the azimuthal quantum number, equation (1) is called the radial equation and equation (2) is called the angular equation.

**Definition 3.8.** We can look for solutions to the angular equation of the form  $Y(\theta, \phi) = \Theta(\theta)\Phi(\phi)$ . It results that there is some constant m such that

(1) 
$$\frac{1}{\Theta}\sin\theta \frac{d}{d\theta} \left(\sin\theta \frac{d\Theta}{d\theta}\right) + l(l+1)\sin^2\theta = m^2$$

$$\frac{1}{\Phi} \frac{d^2 \Phi}{d\phi^2} = -m^2$$

Equation (2) has the solution

$$\Phi(\phi) = e^{im\phi}$$

Since  $(r, \theta, \phi)$  is the same point in space as  $(r, \theta, \phi + 2\pi)$ , we require that  $\Phi(\phi) = \Phi(\phi + 2\pi)$ . This implies that  $m \in \mathbb{Z}$ . The integer m is called the **magnetic quantum number**.

If  $l \in \mathbb{N}_0$  and  $m \leq l$ , then equation (1) has the solution

$$\Theta(\theta) = AP_l^m(\cos\theta)$$

where  $P_l^m$  is the **associated Legendre** function given by

$$P_l^m(x) = (1 - x^2)^{\frac{|m|}{2}} \left(\frac{d}{dx}\right)^{|m|} P_l(x)$$

and  $P_l(x)$  is the  $l^{th}$  Legendre polynomial defined by

$$P_l(x) = \frac{1}{2^l l!} \left(\frac{d}{dx}\right)^l (x^2 - 1)^l$$

The angular function  $Y_l^m(\theta,\phi) = A_l^m P_l^m(\cos\theta) e^{im\phi}$  may be normalized by setting

$$A_l^m = \epsilon \sqrt{\frac{(2l+1)}{4\pi} \frac{(l-|m|)!}{(l+|m|)!}}$$

where

$$\epsilon = \begin{cases} (-1)^m & m \ge 0\\ 1 & m < 0 \end{cases}$$

The normalized angular functions are called **spherical harmonics**.

Exercise 3.9. Compute some spherical harmonics.

**Definition 3.10.** If we make the substitution u(r) = rR(r), we may rewrite the radial equation as

$$-\frac{\hbar^2}{2m}\frac{d^2u}{dr^2} + \left[V + \frac{\hbar^2}{2m}\frac{l(l+1)}{r^2}\right]u = Eu$$

which looks like the one dimensional Schrödinger equation. The function

$$V + \frac{\hbar^2}{2m} \frac{l(l+1)}{r^2}$$

 $is \ called \ the \ \textbf{\it effective potential}.$ 

- 3.3. The Infinite Spherical Box.
- 3.4. The Hydrogen Atom.

# 3.5. Spherical Harmonic Oscillator (Spherical Coordinates).

**Definition 3.11.** The spherical harmonic oscillator (in spherical coordinates) is defined by the potential energy

$$V(r) = r^2$$

**Exercise 3.12.** Making the substitution  $\kappa = \frac{\sqrt{2mE}}{\hbar}$ , we can rewrite the radial equation for the harmonic oscillator as

$$\frac{1}{\kappa^2} \frac{d^2 u}{dr^2} = \left[ \frac{\hbar^2 \omega^2 (\kappa r)^2}{2^2 E^2} + \frac{l(l+1)}{(\kappa r)^2} - 1 \right] u$$

*Proof.* Straightforward

**Exercise 3.13.** Making the substitution  $\rho = \kappa r$  and  $\rho_0 = \frac{\hbar \omega}{2E}$ , we can rewrite the radial equation as

$$\frac{1}{\kappa^2} \frac{\mathrm{d}^2 u}{\mathrm{d}r^2} = \left[ \rho_0^2 \rho^2 + \frac{l(l+1)}{\rho^2} - 1 \right] u$$

*Proof.* Straightforward.

Exercise 3.14. We have

$$\frac{\mathrm{d}^2 u}{\mathrm{d}\rho^2} = \frac{1}{\kappa^2} \frac{\mathrm{d}^2 u}{\mathrm{d}r^2}$$

and thus we may rewrite the radial equation as

$$\frac{\mathrm{d}^2 u}{\mathrm{d}\rho^2} = \left[\rho_0^2 \rho^2 + \frac{l(l+1)}{\rho^2} - 1\right] u$$

*Proof.* Straightforward by chain-rule.

Exercise 3.15. As  $\rho \to \infty$ ,  $u \approx e^{-\frac{\rho_0}{2}\rho^2}$ 

*Proof.* As  $\rho \to \infty$ ,

$$\frac{\mathrm{d}^2 u}{\mathrm{d}\rho^2} \approx \rho_0^2 \rho^2 u$$

Trying the function  $u(\rho) = e^{-\frac{\rho_0}{2}\rho^2}$ , we see that

$$\frac{\mathrm{d}^2 u}{\mathrm{d}\rho^2} = (\rho_0^2 \rho^2 - \rho_0) e^{-\frac{\rho_0}{2}\rho^2}$$
$$\approx \rho_0^2 \rho^2 e^{-\frac{\rho_0}{2}\rho^2} \quad (\text{as } \rho \to \infty)$$
$$= \rho_0^2 \rho^2 u$$

Exercise 3.16. As  $\rho \to 0$ ,  $u \approx \rho^{l+1}$ 

*Proof.* As  $\rho \to 0$ ,

$$\frac{\mathrm{d}^2 u}{\mathrm{d}\rho^2} \approx \frac{l(l+1)}{\rho^2} u$$

Trying the function  $u(\rho) = \rho^{l+1}$ , we see that

$$\frac{\mathrm{d}^2 u}{\mathrm{d}\rho^2} = l(l+1)\rho^{l-1}$$
$$= \frac{l(l+1)}{\rho^2}u$$

**Note 3.17.** We can now, "glue" these functions together with a third unknown function  $v(\rho)$  to obtain the prototype solution

$$u(\rho) = \rho^{l+1} e^{-\frac{\rho_0}{2}\rho^2} v(\rho)$$

**Exercise 3.18.** Suppose that for some nice function  $v(\rho)$ ,

$$u(\rho) = \rho^{l+1} e^{-\frac{\rho_0}{2}\rho^2} v(\rho)$$

Then computing  $\frac{d^2u}{d\rho^2}$  and plugging into the radial equation and simplifying, we obtain the relation

$$\rho \frac{\mathrm{d}^2 v}{\mathrm{d}\rho^2} + 2(l+1-\rho_0\rho^2) \frac{\mathrm{d}v}{\mathrm{d}\rho} + \rho(1-\rho_0(2l+3))v = 0$$

*Proof.* Very tedious but straightforward.

**Exercise 3.19.** If  $v(\rho)$  can be represented by a power series

$$v(\rho) = \sum_{j=0}^{\infty} c_j \rho^j$$

then plugging in  $v(\rho)$  into the previous relation combining like terms and solving for the coefficients yields the relations

$$c_1 = 0$$

and

$$c_{j+2} = \left[ \frac{\rho_0(2j+2l+3)-1}{(j+2)(j+2l+3)} \right] c_j \quad j \ge 0$$

This implies that for each odd j,  $c_j = 0$ .

*Proof.* Tedious but straightforward.

**Exercise 3.20.** If for each  $j \geq 0$ ,  $c_{2j} \neq 0$ , then v behaves asymptotically like  $e^{\rho_0 \rho^2}$ . Thus  $u(\rho)$  behaves asymptotically like  $\rho^{l+1}e^{\frac{\rho_0}{2}\rho^2}$ . This implies that R(r) is not normalizable. Therefore there exists  $j_{max} \geq 0$  such that  $c_{2j_{max}+2} = 0$  and  $v(\rho)$  is a polynomial of degree  $2j_{max}$  and consists of only even powers of  $\rho$ .

*Proof.* As  $j \to \infty$ ,  $c_{j+2} \approx \frac{2\rho_0}{j} c_j$ . Hence  $v(\rho)$  behaves asymptotically like

$$\sum_{j=0}^{\infty} \frac{2^{j} \rho_{0}^{j}}{\prod_{k=1}^{j} 2k} \rho^{2j} = \sum_{j=0}^{\infty} \frac{(\sqrt{\rho_{0}} \rho)^{2j}}{j!}$$
$$= e^{(\sqrt{\rho_{0}} \rho)^{2}}$$
$$= e^{\rho_{0} \rho^{2}}$$

Exercise 3.21. The energies allowed for this system are

$$E_n = \hbar\omega \left(n + \frac{3}{2}\right) \quad n \in \mathbb{N}_0$$

*Proof.* Using the recursion relation found earlier, we have that

$$0 = \left[ \frac{\rho_0(2j_{max} + 2l + 3) - 1}{(j_{max} + 2)(j_{max} + 2l + 3)} \right] c_{j_{max}}$$

This implies that

$$0 = \rho_0(2j_{max} + 2l + 3) - 1$$

and so

$$\frac{1}{\rho_0} = 2j_{max} + 2l + 3$$

Using the fact that  $\rho_0 = \frac{\hbar \omega}{2E}$ , we solve for E to obtain

$$E = \hbar\omega \left( j_{max} + l + \frac{3}{2} \right)$$

Since  $j_{max}$  and l may be any non-negative integers, we introduce a non-negative integer  $n = j_{max} + l$  and index the allowed energies as

$$E_n = \hbar\omega \left(n + \frac{3}{2}\right) \quad n \in \mathbb{N}_0$$

3.6. Orbital Angular Momentum.

**Definition 3.22.** Extrapolating from the classical formula for angular momentum, we define the **orbital angular momentum operator** L, of a particle by

$$L = R \times P$$

so that

$$L_x = YP_z - ZP_y$$
  

$$L_y = ZP_x - XP_z$$
  

$$L_z = XP_y - YP_x$$

and

$$L^2 = L_x^2 + L_y^2 + L_z^2$$

Exercise 3.23. We have that

- $(1) [L_x, L_y] = i\hbar L_z$
- $\begin{array}{l} (2) \ [L_y, L_z] = i\hbar L_x \\ (3) \ [L_z, L_x] = i\hbar L_y \end{array}$

Proof.

$$\begin{split} [L_x, L_y] &= (Y P_z - Z P_y) (Z P_x - X P_z) - (Z P_x - X P_z) (Y P_z - Z P_y) \\ &= Y P_x (P_z Z - Z P_z) + X P_y (Z P_z - P_z Z) \\ &= (X P_y - Y P_x) [Z . P_z] \\ &= i \hbar L_z \end{split}$$

- (2) Similar
- (3) Similar

### Exercise 3.24.

(1) 
$$[L^2, L_x] = 0$$

(2) 
$$[L^2, L_y] = 0$$

(3) 
$$[L^2, L_z] = 0$$

Proof.

(1)

$$[L^{2}, L_{x}] = [L_{x}^{2}, L_{x}] + [L_{y}^{2}, L_{x}] + [L_{z}^{2}, L_{x}]$$

$$= (L_{y}[L_{y}, L_{x}] + [L_{y}, L_{x}]L_{y}) + (L_{z}[L_{z}, L_{x}] + [L_{z}, L_{x}]L_{z})$$

$$= -i\hbar(L_{y}L_{z} + L_{z}L_{y}) + i\hbar(L_{z}L_{y} + L_{y}L_{z})$$

$$= 0$$

- (2) Similar.
- (3) Similar.

**Exercise 3.25.** The operators  $L_x, L_y$  and  $L_z$  are self-adjoint.

*Proof.* Let  $\Psi_1, \Psi_2$  be wave functions. Since  $X_i$  and  $P_j$  are self-adjoint and commute for  $i \neq j$ , we have that

$$\begin{split} \langle \Psi_1 | L_x \Psi_2 \rangle &= \langle \Psi_1 | Y P_z \Psi_2 \rangle - \langle \Psi_1 | Z P_y \Psi_2 \rangle \\ &= \langle P_z Y \Psi_1 | \Psi_2 \rangle - \langle P_y Z \Psi_1 | \Psi_2 \rangle \\ &= \langle Y P_z \Psi_1 | \Psi_2 \rangle - \langle Z P_y \Psi_1 | \Psi_2 \rangle \\ &= \langle L_x \Psi_1 | \Psi_2 \rangle \end{split}$$

So  $L_x$  is self-adjoint. The case is similar for  $L_y$  and  $L_z$ 

**Definition 3.26.** We define the raising operator  $L_+$  and lowering operator  $L_-$  by

$$L_+ = L_x + iL_y$$
 and  $L_- = L_x - iL_y$ 

Exercise 3.27.

$$[L^2, L_+] = [L^2, L_-] = 0$$

Proof. Trivial.  $\Box$ 

Exercise 3.28. The lowering operator is the adjoint of the raising operator:

$$L_{-} = (L_{+})^{\dagger}$$

*Proof.* Let  $\Psi_1, \Psi_2$  be wavefunctions. Then

$$\begin{split} \langle \Psi_1 | L_+ \Psi_2 \rangle &= \langle \Psi_1 | L_x \Psi_2 \rangle + i \langle \Psi_1 | L_y \Psi_2 \rangle \\ &= \langle L_x \Psi_1 | \Psi_2 \rangle + i \langle L_y \Psi_1 | \Psi_2 \rangle \\ &= \langle L_x \Psi_1 | \Psi_2 \rangle + \langle -i L_y \Psi_1 | \Psi_2 \rangle \\ &= \langle (L_x - i L_y) \Psi_1 | \Psi_2 \rangle \\ &= \langle L_- \Psi_1 | \Psi_2 \rangle \end{split}$$

Hence  $L_{-} = (L_{+})^{\dagger}$ .

Exercise 3.29. We have

(1) 
$$[L_z, L_+] = \hbar L_+$$
  
(2)  $[L_z, L_-] = -\hbar L_-$ 

Proof.

(1)

$$[L_z, L_+] = [L_z, L_x] + i[L_z, L_y]$$
$$= i\hbar L_y + \hbar L_x$$
$$= \hbar L_+$$

(2) Similar.

Exercise 3.30. We have

$$(1) L^2 = L_+ L_- + L_z^2 - \hbar L_z$$

(1) 
$$L^2 = L_+ L_- + L_z^2 - \hbar L_z$$
  
(2)  $L^2 = L_- L_+ + L_z^2 + \hbar L_z$ 

Proof.

(1)

$$L_{+}L_{-} = (L_{x} + iL_{y})(L_{x} - iL_{y})$$

$$= L_{x}^{2} - i(L_{x}L_{y} - L_{y}L_{x}) + L_{y}^{2}$$

$$= L_{x}^{2} + L_{y}^{2} - i[L_{x}, L_{y}]$$

$$= L^{2} - L_{z}^{2} + \hbar L_{z}$$

Therefore

$$L^2 = L_+ L_- + L_z^2 - \hbar L_z$$

(2) Similar

Exercise 3.31. Suppose that f is simultaneously an eigenfunction of  $L^2$  with eigenvalue  $\lambda$ and an eigenfunction of  $L_z$  with eigenvalue  $\mu$ . Then

- (1)  $L_+f$  is simultaneously an eigenfunction of  $L^2$  with eigenvalue  $\lambda$  and an eigenfunction of  $L_z$  with eigenvalue  $\mu + \hbar$
- (2)  $L_+f$  is simultaneously an eigenfunction of  $L^2$  with eigenvalue  $\lambda$  and an eigenfunction of  $L_z$  with eigenvalue  $\mu - \hbar$

Proof.

(1) First we have

$$L^{2}L_{+}f = L_{+}L^{2}f$$
$$= L_{+}\lambda f$$
$$= \lambda L_{+}f$$

Second we see that

$$L_z L_+ f = \left[ L_+ L_z + \left( L_z L_+ - L_+ L_z \right) \right] f$$

$$= \left( L_+ L_z + \left[ L_z, L_+ \right] \right) f$$

$$= \left( \mu L_+ + \hbar L_+ \right) f$$

$$= \left( \mu + \hbar \right) L_+ f$$

(2) Similar.

**Interpretation 3.32.** The rasing/lowering operators change the state of a particle from simultaneous eigenstate to simultaneous eigenstate and increase/decrease the z-component of the orbital angular momentum of a particle by  $\hbar$ , but does not change the total orbitle angular momentum of the particle.

Note 3.33. If we repeatedly applied the increasing/decreasing operator to a simultaneous eigenstate of a particle, we would arrive at a top/bottom eigenstate since there is a finite amount of total angular momentum.

**Definition 3.34.** We define the **top state**  $f_t$  to be the simultaneous eigenstate of  $L^2$  and  $L_z$  such that

$$L_+ f_t = 0$$

and we define the **bottom state**  $f_b$  to be the simultaneous eigenstate of  $L^2$  and  $L_z$  such that

$$L_- f_b = 0$$

**Exercise 3.35.** Let  $\hbar l$  and  $\hbar l'$  be the eigenvalues of  $L_z$  for  $f_t$  and  $f_b$  respectively. Then

- (1)  $\hbar^2 l(l+1)$  and  $\hbar^2 l'(l'-1)$  are the eigenvalue of  $L^2$  for  $f_t$  and  $f_b$  respectively
- (2) l' = -l
- (3)  $l \in \frac{1}{2}\mathbb{N}_0$

Proof.

(1)

$$L^{2}f_{t} = (L_{-}L_{+} + L_{z}^{2} + \hbar L_{z})f_{t}$$
$$= \hbar^{2}l^{2}f_{t} + \hbar^{2}lf_{t}$$
$$= \hbar^{2}l(l+1)$$

The case is similar for l'.

(2) For some  $N \in \mathbb{N}_0$  and constant C, we have  $f_b = CL_-^N f_t$ . Thus

$$\begin{split} \hbar^2 l'(l'-1)f_b &= L^2 f_b \\ &= L^2 C L_-^N f_t \\ &= C L_-^N L^2 f_t \\ &= C L_-^N \hbar^2 l(l+1) f_t \\ &= \hbar^2 l(l+1) f_b \end{split}$$

So l'(l'-1) = l(l+1). By completing the square, we see that the only two ways that this equation is satisfied is if l' = l+1 or l' = -l. The first case is not possible since it would imply that the orbital angular momentum for a particle is greater in the state  $f_b$  than in the state  $f_t$ . So l' = -l.

(3) Since for some  $N \in \mathbb{N}_0$  and constant C,  $L_b = CL_-^N f_t$ , a previous exercise implies that

$$-\hbar l f_b = L_z f_b$$

$$= L_z C L_-^N f_t$$

$$= \hbar (l - N) C L_-^N f_t$$

$$= \hbar (l - N) f_b$$

So -l = l - N and  $l = \frac{N}{2}$ . Thus  $l \in \frac{1}{2} \mathbb{N}_0$ .

**Definition 3.36.** For  $m=-l,-l+1,\cdots l-1,l$  define  $f_l^m$  to be the simultaneous eigenstate of  $L^2$  and  $L_z$  given by

$$f_l^m = \langle L_-^{l-m} f_t | L_-^{l-m} f_t \rangle^{-\frac{1}{2}} L_-^{l-m} f_t$$

Exercise 3.37. We have that

- (1) the eigenvalue of  $L^2$  corresponding to  $f_l^m$  is  $\hbar^2 l(l+1)$
- (2) the eigenvalue of  $L_z$  corresponding to  $f_l^m$  is  $\hbar m$ .

*Proof.* Straightforward (and kind of already did it in the last exercise).  $\Box$ 

Proposition 3.38.