

# NOTE 257 –

Creating a Useful Glish Client (C), Glen Langston, NRAO-GB

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## 1 Summary

C programs are easily added to the `glish` command prompt in AIPS++, by creating `glish` clients. We present an example client which calculates radio source flux densities based on the Ott *et al.* 1994 and Baars *et al.* 1987 models for a set of reference sources. This example is intended for C programmers who wish to add new capabilities to `glish` and AIPS++. This example can be modified for “easy” incorporation of stand alone functions.

This document describes the components needed to create a `glish` client. Two types of interfaces between C and `glish` are shown, 1) passing strings and 2) arbitrary sized double precision arrays.

The steps for construction, testing and execution are given. In appendices, the `glish` and C++ code is listed. To view the entire C code functions, down-load the tar file from the web. This document and the code are on the web at: <http://www.gb.nrao.edu/~glangsto/aips++/glishClient>

This document is an update of the examples created by Rick Fisher in 1997. Rick’s document is useful as an example of how any type of Glish variable can be passed between `glish`, C, Fortran and C++. See: <http://www.gb.nrao.edu/~rfisher>

## 2 Components

These Glish clients are run as a part of an AIPS++, DISH or vanilla `glish` session. The clients are produced from a number of components listed below:

<code>ottClient.g</code>	Glish scripts interfacing to the C++ wrapper around the observers C program. This script loads the <code>glish</code>
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procedure. The scripts handle data typing and runs a test. A Unix shell interface to the program `ott` is presented. session to confirm the configuration.

<code>ottClient.cc</code>	C++ wrapper that identifies Glish events and calls the C functions.
<code>ottFluxes.c</code>	C function to calculate the Ott <i>et al.</i> 1994 source model flux densities.
<code>ott.c</code>	Stand alone C function to calculate the Ott <i>et al.</i> 1994 model. This function also used <code>ottFluxes.c</code>
<code>MakeOtt</code>	Make file that is configured for the local AIPS++ installation. This file will probably require modification for successful installation of the <code>ottClient</code>

### 3 Test Execution of `ott`

After creating the client and stand-alone program, test `ott` with different inputs. With only the frequency (MHz) argument, `ott` program prints the flux densities for all sources for a input frequency (MHz). Output flux densities are in Janskies.

```
% ott 1408
Calculating flux densities for frequency 1408.000000 MHz
  CASA    2092.0688+/- 160.548
  3C286    14.6208+/-   0.940
  3C48     16.3005+/-   1.192
  3C147    22.0340+/-   1.384
  3C138     8.4819+/-   0.848
1934-638   16.2053+/-   1.621
  3C405   2706.6796+/- 429.630
  DR21     26.6606+/-   3.627
 NGC7027    7.9449+/-   3.972
  3C295    22.2488+/-   1.676
  3C123    47.3906+/-   4.739
  3C161    18.6043+/-   1.860
  3C218    42.7130+/-   4.271
  3C227     7.6057+/-   0.380
 3C249.1    2.2779+/-   0.228
```

VIRA	203.1464+/-	10.157
3C309.1	7.4350+/-	0.520
3C348	46.9291+/-	2.346
3C353	56.3030+/-	2.815
	1.0000+/-	1.000

Epoch 2002DEC02 Cas-A Flux density based on 22.92 year decrease  
 Cas-A decrease factor: 0.9907/yr => 0.8081 total factor  
 CASA 1690.5828+/- 129.737

The `ott` program will also calculate the flux density for a single source at a single frequency. This capability is used in `ottClient.g`, function `ottShell()` to call a C program from `glsh`.

```
% ott -s 3C286 1408
14.6208
```

## 4 Construction

To create `ott`, the stand alone program, and `ottClient`, the `glsh` client, download the TAR file `ott.tar`. Extract the components, edit the `MakeOtt` makefile and create the program in the following steps.

```
% mkdir ottClient          # create a new directory
% cd ottClient              # go to the directory
% wget http://www.gb.nrao.edu/~glangsto/aips++/ott.tar #use wget or ftp
% tar -xvf ott.tar          # separate the components
% emacs MakeOtt             # edit the makefile for local paths
% source /home/aips++/stable/aipsinit.sh    # setup the aips environment
% make                      # create the programs
```

The tar-ball `ott.tar` also contains the executables for RedHat Linux. The `make` step can be skipped if `ott` works properly at the command line prompt. The individual components of `ott.tar` are at: <http://www.gb.nrao.edu/simglangsto/aips>

## 5 Execution of `ottClient.g`

The first step in using the C clients in `glsh` is loading the programs and `glsh` scripts into the local `aips++` directory.

Next, the shell environment variables must be configured for the local `aips++` installation. This is done with a Unix script. Consult with your `aips++` expert to find the script.

The client is tested during loading in the `glsh` session. An example session is:

```
% bash          # start the bash shell
% source /home/aips++/stable/aipsinit.sh
AIPS++ Version 1.8 Build 321
% glsh -l ottClient.g
reading .glshrc_local from /home/aips++/stable2
```

Test Execution of Ott functions:

```
ottSourceName ( '3C295')      := T
freqs              := [2e+08 3e+09 4e+10 5e+11]  (Hz)
ottFluxes( freqs)           := [76.2198 11.0079 0.379414 0.00340961]  (Jy)
ottShell( '3C295', freqs)    := [76.2198 11.0079 0.379414 0.00341]  (Jy)
Glish version 2.7.
- exit
```

The function `ottShell()` works fine for a few flux density values. However its performance is too slow for an array larger than a thousand or so frequencies. In this case the function `ottFluxes()` is far superior.

## 6 Documentation

The program `ott` prints help if no input arguments are provided, as show below:

```
% ott
```

```
ott: prints Ott et al 1994 and Baars et al 1977 flux densities.
usage: ott [-s <sourceName>] [-d <epoch>] frequencyMHz
where [-s <sourceName>] is the optional source name.
where <frequencyMHz>    frequency for model flux densities
where [<epoch>]          optional date string for calculating Cas-A decay
                        string is in YYmmmmDD format (ie 99feb04 or 05apr01)
```

The models of Ott, M. Witzel, A., Quirrenbach, A., Kirchbaum, T.P. Standke, K., J., Schalinski, C. J., and Hummel, C.A., 1994 *Astronomy and Astrophysics*, Vol 284 pg 331 and Baars, Genzel, Pauliny-Toth and Witzel 1977, *Astronomy and Astrophysics*, Vol 61, page 99 are used.

The Ott et al models are good for 1408 to 23000 MHz for most sources For DR21, the Baars et al 1977 model is used. For Cas-A, the current date is used to calculate flux density Based on the Baars epoch 1980 model for decline in flux densities Cas-A Model:  $0.97 \pm 0.04 - 0.30 \pm 0.04 \log(\text{freq./GHz})$  percent decrease/year

The Error estimates are approximate, using values in the text and

interpolating. The default is 5 percent.

Sources Modeled:

CASA	111-2	2323+588
3C286	1328+307	1331+305
3C48	0134+329	0137+331
3C147	0538+498	0542+498
3C138	0518+165	0521+166
1934-638		
3C405	CYGA	1957+406
DR21	2037+421	2039+423
NGC7027	2105+420	2107+422
3C295	1409+524	1411+522
3C123	0433+295	0437+296
3C161	0624-058	0627-058
3C218	0915-119	0918-120
3C227	0945+077	0947+074
3C249.1	1100+772	1104+769
VIRA	1228+127	1230+123
3C309.1	1458+718	1459+716
3C348	1648+051	1651+049
3C353	1717-009	1720-009

## 7 References

- 1 Baars, Genzel, Pauliny-Toth and Witzel (1977), Astronomy and Astrophysics, Vol. 61, pg. 99.
- 2 Ott, M. Witzel, A., Quirrenbach, A., Kirchbaum, T.P. Standke, K., J., Schalinski, C. J., and Hummel, C.A., (1994) Astronomy and Astrophysics, Vol. 284, pg 331.

## 8 Appendix: ottClient.g

Below is a listing of the ottClient glish script used to interface glish to the C functions.

```
#File ottClient.g, version 1.3, released 02/12/02 at 11:26:08
# retrieved by SCCS 02/12/02 at 11:26:18
#Glish event wrapper functions for calling C functions from Glish
#HISTORY
```

```

# 021202 GIL change ottFlux to ottShell
# 021122 GIL minor initial version based on ex_client.g
# 021115 GIL initial version based on ex_client.g
# 021114 GIL update for minor changes to glish
# 970803 JRF Initial version very well documented at
#          http://www.gb.nrao.edu/~rfisher/Glish/ex\_client.html

global ottSource := '3C286'

ottFluxes := function ( valu )
# ottFluxes takes an array of frequencies (Hz) and returns an array of
# flux densities (Jy). Must first set the source name with ottSourceName()
{
    freqFluxes := as_double( valu)          /* transfer frequencies */
    oFluxes := sf->ottSetFlux( freqFluxes);

    return (oFluxes)
} /* end of ottFluxes() */

ottSourceName := function ( sourceName )
{ #store source name as global with error checking
  # INPUTS:  sourceName      string name of source "CASA" or "3C286" etc

  global ottSource;

  ottSource := as_string( sourceName);

  global sf := client('ottClient');
  dummyValue := sf->ottSetSource( ottSource); /* now set the source name */

  return(T)
} /* end of ottSourceName() */

ottShell := function ( sourceName, frequencyVector)
{ # ottFlux gets a single ott et al flux via a commandline interface to ott.
  # INPUTS:  sourceName      string name of source "CASA" or "3C286" etc
  #          frequencyVector Array of frequencies (Hz)
  # OUTPUT:  fluxVector      Array of flux densities (Jy)
  # This function is 100 times slower than ottFluxes()

```

```

frequencyShape := shape( frequencyVector);

if ( length( frequencyShape) > 1) {
    print "Frequency vector must be one dimensional!"
    return 0;
}
n := frequencyShape[1];           # get number of frequencies
fluxVector := array( 0, n);       # create output array

frequencyMHz := as_double( frequencyVector) * 0.000001; #from Hz to MHz */
for (i in (1:n)) {                # for all frequencies
    shellString := sprintf( "ott -s %s %f", sourceName, frequencyMHz[i]);
    # execute the shell string and convert to double
    fluxVector[i] := as_double( shell( shellString));
} # end for all frequencies
return fluxVector;
} # end of ottShell

freqs := [ 2e8, 3e9, 4e10, 5e11];      # set the test frequencies

print ''
print 'Test Execution of Ott functions:'
print 'ottSourceName ( \'3C295\')      :=', ottSourceName( '3C295');
print 'Freqs                          :=', freqs, "(Hz)";
print 'ottFluxes( freqs)              :=', ottFluxes( freqs), "(Jy)";
print 'ottShell( \'3C295\', freqs)     :=', ottShell( '3C295', freqs),"(Jy)";

```

## 9 Appendix: ottClient.cc

The C++ wrapper main program to the C functions is listed below. The functions performing the calculations are in ottFluxes.c.

```

/* File ottClient.cc, version 1.2, released 02/12/02 at 11:42:52
   retrieved by SCCS 02/12/02 at 11:43:21

```

C++ client code wrapper for C functions to calculate the Ott et al fluxes.

HISTORY

021202 GIL remove un-used segments of the code.



021118 GIL add some comments  
 021115 GIL Initial version based on ex\_client.cc  
 021114 GIL Removed support for complex functions  
 970803 JRF Initial version, well documented at  
[http://www.gb.nrao.edu/~rfisher/Glish/ex\\_client.html](http://www.gb.nrao.edu/~rfisher/Glish/ex_client.html)

#### DESCRIPTION

Rick Fisher created a very nice set of example code that allows an observer to quickly create a interface from C to glish.  
 This function impliments the C++ wrapper to a simple C program to calculate the Ott et al 1994 flux densities for a set of reference sources.

There are two steps in the process: First is setting the source name for which the values are calculated. The second step is providing an array of frequencies (Hz), for which the flux densities are required.  
 \*/

```
#include <stdio.h>
#include <string.h>
#include <math.h>
#include "Glish/Client.h"
```

```
// Declare all of the C/C++ functions that you are going to use or put
// the main() function at the end.
```

```
void setFluxes(Client &c, GlishEvent *e);
void setSource(Client &c, GlishEvent *e);
```

```
extern "C" {
    char * setOttSourceName( char * source);
    char * setOttSourceFluxes( long n, double values[]);
} /* end of C declares */
```

```
int main (int argc, char **argv) {
    // This creates a required client object.
    Client c(argc, argv);
    // The client can be invoked with arguments, but we'll bypass that
    // complication.
    if (argc > 1) {
        printf ("Usage: cl := client('ex_client')\n");
        return 1;
    } /* end if an argument */
```



```
// vvvvvvvvvvvvvv Your code substituted below here. vvvvvvvvvvvvvv
    for (i = 0; i < array_length; i++)
        freqFluxes[i] = dataValues[i];

    setOttSourceFluxes( array_length, freqFluxes);
    for (i = 0; i < array_length; i++) /* transfer out */
        return_value[i] = freqFluxes[i];
// ~~~~~ Your code substituted above here. ~~~~~
    delete freqFluxes;
} /* end else if data of expected type */
Value *rep = new Value(return_value, array_length, COPY_ARRAY);
c.Reply(rep);
delete return_value;
Unref(rep);
} /* end of setFluxes() */

#define MAXSIZE 100 /* set string max size */
void setSource(Client &c, GlishEvent *e) {
    // setSource() takes the source name event and transfers the name to the C
    // functions.
    Value *val = e->Val();
    char return_value[MAXSIZE] = "", sourceName[MAXSIZE] = "", * errMsg = '\0';

    if (val->Type() != TYPE_STRING) {
        printf("String type value expected from 'reverse_string'\n");
    } else {
        char *received_value = val->StringVal();
        strncpy( sourceName, received_value, MAXSIZE);
        sourceName[MAXSIZE-1] = '\0';
        // The call of val->StringVal() implicitly allocates memory
        // for the string which must be deleted to avoid a memory leak.
        delete received_value;
// vvvvvvvvvvvvvv Your code substituted below here. vvvvvvvvvvvvvv
        errMsg = setOttSourceName( sourceName);
        if (errMsg != '\0') /* check for error strings */
            strcpy( return_value, "!!!! ");
        strcat( return_value, sourceName);
// ~~~~~ Your code substituted above here. ~~~~~
    }
    Value *rep = new Value(return_value);
}
```

```

        c.Reply(rep);
        Unref(rep);
    } /* end of setSource() */

```

## 10 Appendix: MakeOtt

The makefile, MakeOtt used to create the clients is listed below. The tar-ball ott.tar contains a symbolic link from MakeOtt to makefile.

```

#Make file for C++ wrapper to C code for Ott et al 1994 flux densities.
#HISTORY
# 021122 GIL Add make file for stand alone program ott
# 021118 GIL Update for configuration changes in GB.
# 970101 JRF Initial version
#DESCRIPTION
#This make file creates two programs
#ottClient: glish client allowing a glish session to call C functions
#ott:      stand alone program (linux or solaris) calc Ott etal

# at the ATNF
#BASE_DIR = /aips++
# at Green Bank
#BASE_DIR = /aips++/test
BASE_DIR = /home/aips++/stable

#Now use Linux not solaris
#LIBRARIES = -L$(BASE_DIR)/sun4sol_gnu/lib
LIBRARIES = -L$(BASE_DIR)/linux/lib

INCLUDES = -I$(BASE_DIR)/code/aips/glish/include -I./xlib

#this make file makes two programs
EXECUTABLES = ottClient ott

all: $(EXECUTABLES)

#selected code taken un-modified from the VLBA antenna control code and
#OVLBI tracking station.
VLBASTRINGS = str2mjd.o str2rad.o stripWhite.o time2str.o mjd2str.o \
srclist.o cvrtuc.o dateObs2DMjd.o today2mjd.o ottFluxes.o

```

```

#the VLBA code environment includes definitions of C structures not
#used here. This stub replaces the include with the standard definitions
vlb.h :
ln -s STDDEFS.H vlb.h

#The VLBA code contains two great general purpose includes of constant values
#Math constants and the definitions of TRUE, FALSE etc.
$(VLBASTRINGS) : MATHCNST.H STDDEFS.H vlb.h $(@:.o=.c)

#now describe how to compile
CC=gcc
CFLAGS= -O -Wall -g $(INCLUDES) $(LIBRARIES)

COMPILE.c= $(CC) $(CFLAGS) $(CPPFLAGS) -c

#rule for converting .c to a .o file
.c.o :
@ $(RM) $@
$(COMPILE.c) $<

ottClient.o : ottClient.cc
g++ $(INCLUDES) -c ottClient.cc

ottClient : ottClient.o $(VLBASTRINGS) $(OTTCLIENTC)
g++ -o ottClient $(OTTCLIENTC) $(VLBASTRINGS) ottClient.o \
$(LIBRARIES) -lglish -lsos -lnpd -lm

ott : ott.o $(VLBASTRINGS) $(OTTCLIENTC)
$(CC) $(CFLAGS) -o ott $(OTTCLIENTC) $(VLBASTRINGS) ott.o -lm

```