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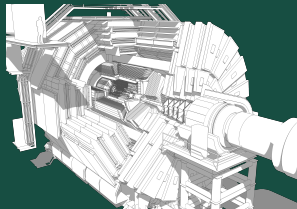
Simplest Fermion Vector-Like Portal Dark Matter model:

Search in the compressed mass region at the CMS experiment

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Introduction



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Dark Matter (DM) constitutes one of the main unsolved problems in fundamental physics. Ever since it was proposed to explain the rotation curves of galaxies, some other astronomical observations left little doubt of its existence.

Whatever DM is, the Standard Model (SM) is not able to produce a candidate that have, at the present time in the universe, stability and also interact very little or not at all with known matter, proven properties of DM.



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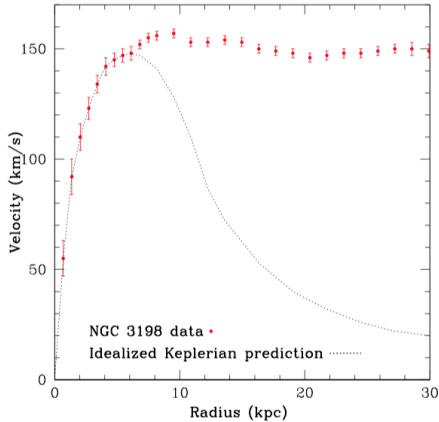


Figure 1: Measured rotational velocities of HI regions in NGC 3198 compared to an idealized Keplerian behavior [Astron. and Astrophys. 223, 47-60]

Bullet cluster



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Figure 2: The Bullet cluster, the result of a subcluster (the “bullet”) colliding with the larger galaxy cluster 1E 0657-56 [arXiv:1711.02117]



Motivation



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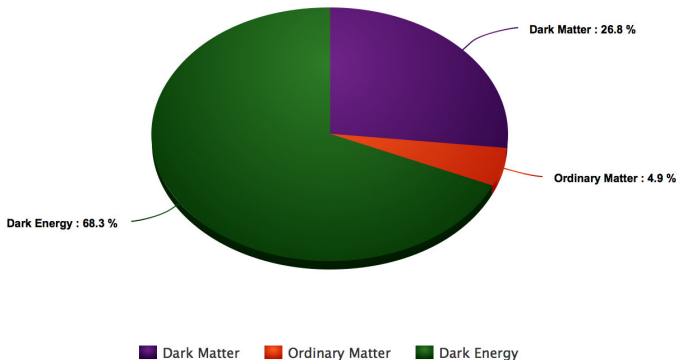


Figure 3: Simplified plot showing the estimate abundances, of the known components of the universe.



Detection Channels



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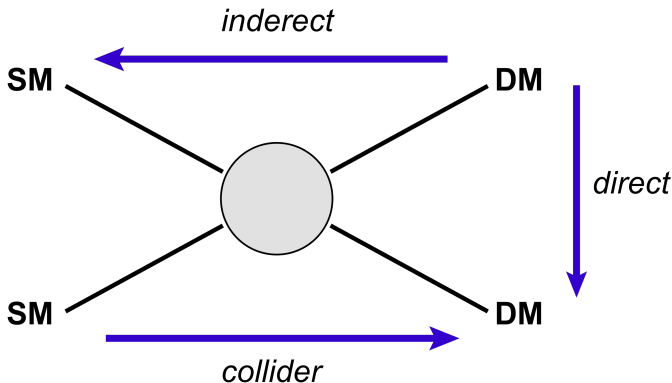
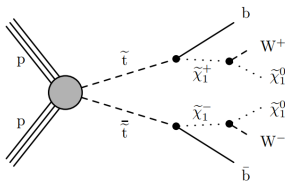


Figure 4: Schematic showing the possible dark matter detection channels.

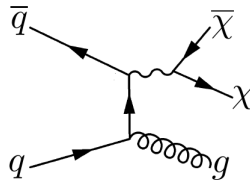


Signals

- 1 Cascades of heavier particles to the DM candidate and SM, as in (a).
- 2 DM production in association with other SM particles, like (b).



(a)



(b)



Model



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The VLF model introduce a vector-like charged fermion also $SU(2)_L$ singlet, and a scalar particle. The free Lagrangian reads

$$\mathcal{L} = \mathcal{L}_{SM} + m_F \bar{F} F + (Y_\ell S \bar{F} \ell_R + \text{h.c.}) + V(S, H) ;$$

m_F is the singlet fermion mass parameter, ℓ_R are the SM right-handed lepton fields, Y_ℓ are the Yukawa couplings. The contribution to the scalar potential is given by

$$V(S, H) = \frac{m_S^2}{2} S^2 + \frac{\lambda_S}{4} S^4 + \lambda_{SH} S^2 |H|^2 ,$$

with scalar mass parameter m_S and quartic couplings λ_S and λ_{SH} .



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The kinetic terms for the vector-like charged fermion reads as follows

$$\mathcal{L} \bar{F}(D_\mu \gamma^\mu)F, \quad (1)$$

consequently the simplest production and decay process at the LHC is pair-production of vector-like fermions, followed by their subsequent decay to a lepton and the DM scalar.

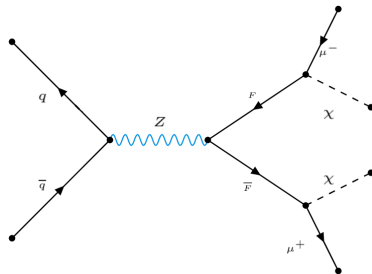


Figure 6: Pair-production of vector-like fermions ($pp \rightarrow F^- F^+$), in a Drell-Yan process, followed by their decay into a lepton and the DM particle ($F^- \rightarrow \ell^- S + \text{h.c.}, \ell = e, \mu, \tau$)



Model



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We focus on scenarios where the vector-like portal for DM annihilation is dominant and, therefore, we set initially.

$$\lambda_{SH} = 0;$$

we assume as well, that the DM candidate does not couples to the electron.

$$Y_e = 0;$$

the remaining parameters, Y_μ, Y_τ , m_S and m_F are allowed to vary freely. Furthermore we are searching into the compress mass regime i.e.

$$\Delta m = m_F - m_S \lesssim 50 \text{ GeV}$$

One-muon + jet channel



The process shown in Figure 6, leads to the signature opposite sign leptons plus missing energy however,

- for small $\Delta m(F - S)$ the probability that one or both leptons are not detected in the collider increases.
- $pp \rightarrow F^+ F^- j$ provides a jet which can be used as a trigger.

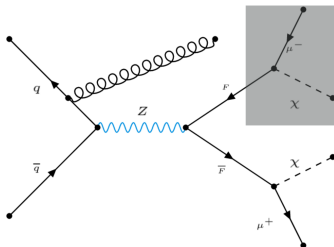


Figure 7: Pair-production of vector-like fermions, followed by their decay into a lepton and the DM particle, one missing lepton and a ISR jet

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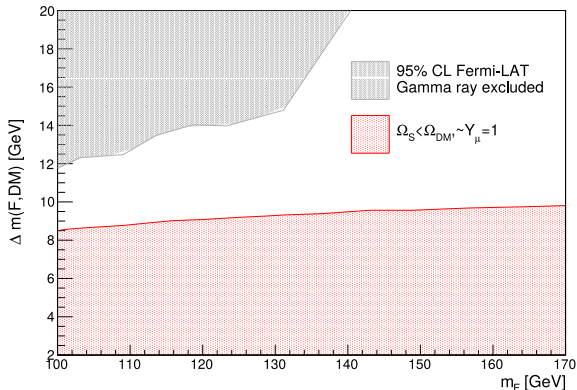


Figure 8: m_F vs Δm . The gray upper region is excluded by Fermi-LAT H.E.E.S, while the red area is where the relic density is not satisfied by the model.



The backgrounds process for this search are the single top, the di-boson WZ , and $W + jets$. The last one being the most important with production cross section three orders larger than the others (see Figure 9).

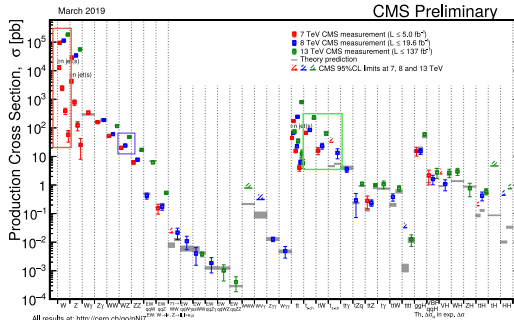


Figure 9: Summary of the cross section measurements of Standard Model processes [<https://twiki.cern.ch/twiki/bin/view/CMSPublic/PhysicsResultsCombined>].



Samples Generation



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The Monte Carlo(MC) collision samples for the signal and the BackGrounds (BG) were generated using a combination of MadGraph (version 2.5.5) for the event generation, Pythia (version 8.233) for the hadronization and Delphes (version 3.4.1) for the detector effect emulation.



Background Tuning



To generate the background we use the tunes extract for the data from CMS at the reference [CMS-PAS-GEN-17-001]

PYTHIA parameter	Value
PDF Set	NNPDF3.1
$\alpha_S(M_Z)$	0.118
SPACESHOWER:RAPIDITYORDER	on
MULTIPARTONINTERACTIONS:ECMREF [GeV]	7000
α_S^{ISR}	0.118/NLO
α_S^{FSR}	0.118/NLO
α_S^{MPI}	0.118/NLO
α_S^{ME}	0.118/NLO
MULTIPARTONINTERACTIONS:PT0REF [GeV]	1.41
MULTIPARTONINTERACTIONS:ECMPOW	0.03344
MULTIPARTONINTERACTIONS:CORERADIUS	0.7634
MULTIPARTONINTERACTIONS:COREFRACTION	0.63
COLORRECONNECTION:RANGE	5.176
χ^2/dof	1.04

Table 1: Pythia 8 parameter values

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After the tuning we use Rivet[] over the analysis [Phys-RevD.96.072005] to compare

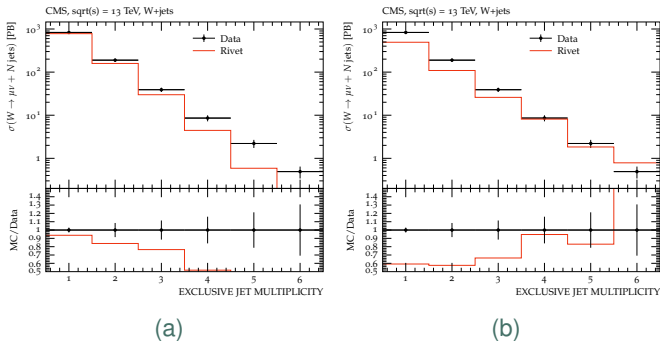


Figure 10: Differential cross section measurement for the exclusive jet multiplicities with tune (a) and without tune (b)



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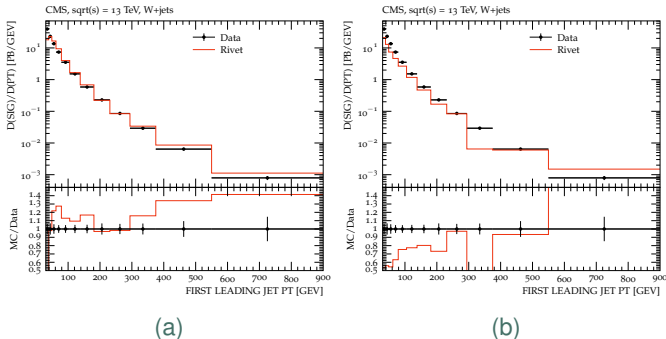


Figure 11: Differential cross section measurement for the exclusive jet multiplicities with tune (a) and with out tune (b)



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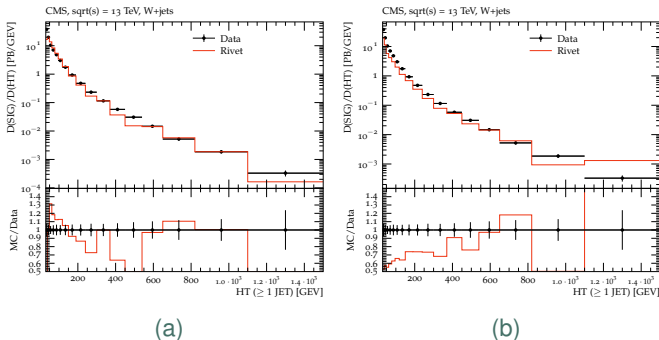


Figure 12: Differential cross section measurement for the jets HT for at least 1 jet with tune (a) and with out tune (b)



Selection criteria



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The values for the cuts were defined through an optimization process based on both the significance \mathcal{Z} (defined in the 2) and the efficiency of the cut (defined in the 3) of the signal Vs BG

$$\mathcal{Z} = \frac{S}{\sqrt{S+B}}, \quad (2)$$

where S and B are the yields for Signal and BG. We normalize all MC to 100 fb^{-1} luminosity and the signal mass point $m_F = 145 \text{ GeV}$ vs $\Delta m = 10 \text{ GeV}$

$$e_{ff}(\chi) = \frac{\chi_{AC}}{\chi_{BC}}, \quad (3)$$

where χ can be signal or BG, and the subscripts BC and AC stands for "Before Cuts" and "After Cuts".

m_T Selection



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Most signal (bottom-right figure) events are around 50 GeV, but for all BG the peak of events is more towards 80 GeV, especially the $W + Jets$. That is why we chose to make a cut of the type **Less-Than**.

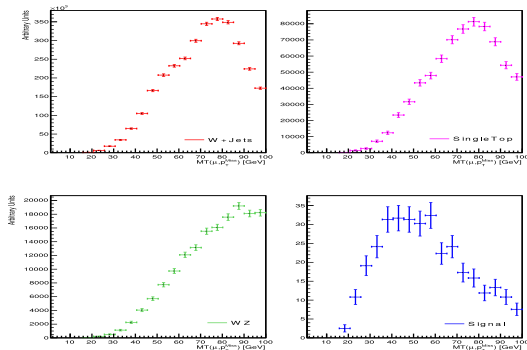


Figure 13: $m_T(\mu, p_T^{miss})$ histograms for the signal and each background sample. The bin-width is 5 GeV. Only the basic selection has been applied.



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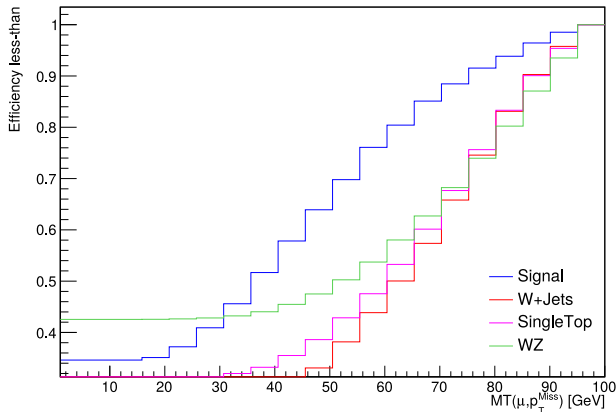


Figure 14: $M_T(\mu, p_T^{\text{miss}})$ Efficiency plots for the signal and each background sample. The main BG is less than 1% before 40 GeV, although the signal is more than 50%.



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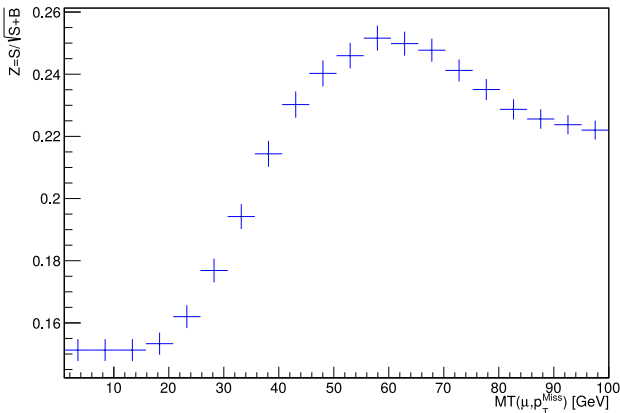


Figure 15: Significance of the cut (type less-than) for $m_T(\mu, p_T^{Miss})$. The maximum of Z is reached for 60 GeV, but to take out the majority of the $W + jet$ BG events, the cut at 40 GeV is preferred.

$p_T(\mu)$ Selection



Most signal events are concentrated at low values of p_T as it was expected. It is chosen a less-than cut .

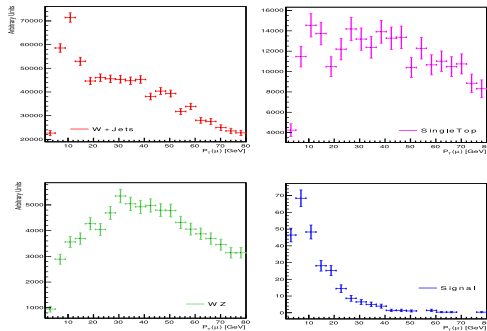


Figure 16: $p_T(\mu)$ histograms for the signal and each background sample. The bin-width is 5 GeV. Besides the basic selection, the cut $M_T < 40$ GeV was made.



$p_T(\mu)$ Selection



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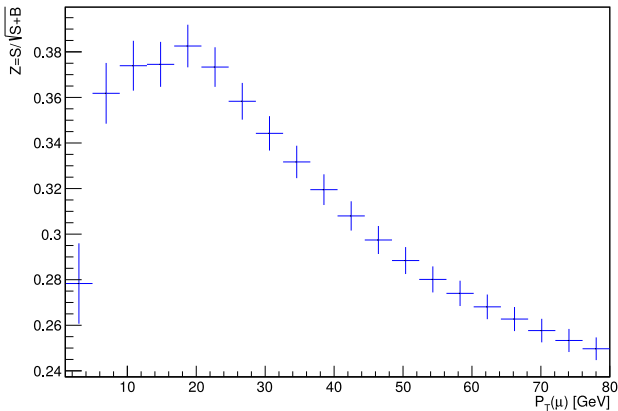


Figure 17: Significance of the cut (type less-than) for $p_T(\mu)$. The Z curve reach the maximum at $p_T(\mu) < 20$ GeV and falls fast for higher values.



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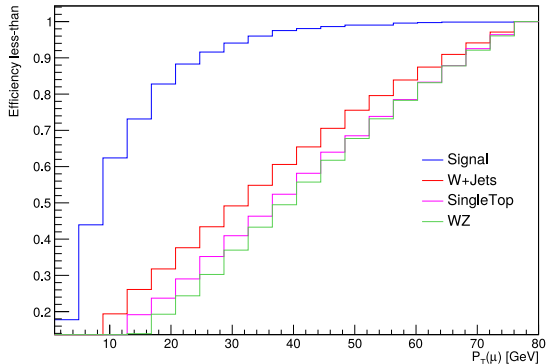


Figure 18: $p_T(\mu)$ efficiency plots for the signal and each background sample. For the cut suggested by Figure 17, the signal reach more than 80% of efficiency, while the main BG is less than 40%, confirming the cut based on the significance.

$N(Jet)$ Selection



The signal, W+jets and WZ BG have a similar behavior, but in order to reduce the Single Top events a less-than cut can be used.

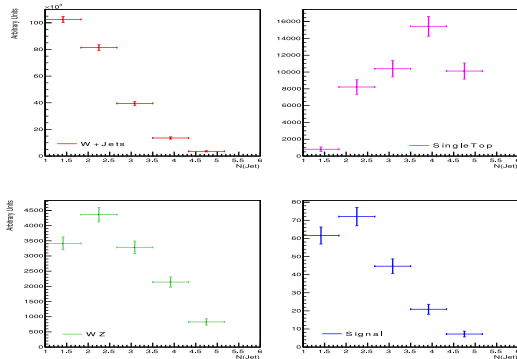


Figure 19: $N(Jet)$ histograms. The cuts applied are: the basic selection, $M_T < 40$ GeV and $p_T(\mu,) < 20$ GeV.



$N(Jet)$ Selection



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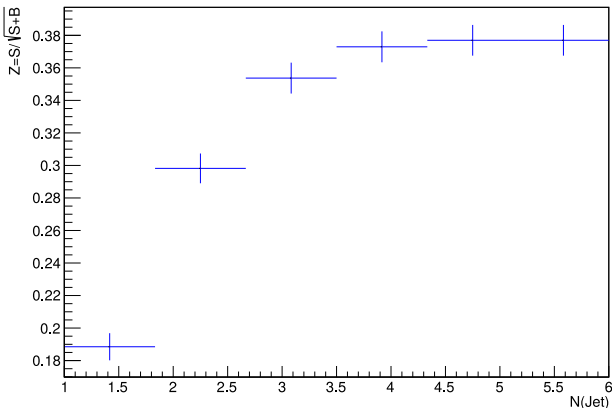


Figure 20: The Significance for $N(Jet)$ reach its maximum at $N(Jet) < 4$.



$N(Jet)$ Selection



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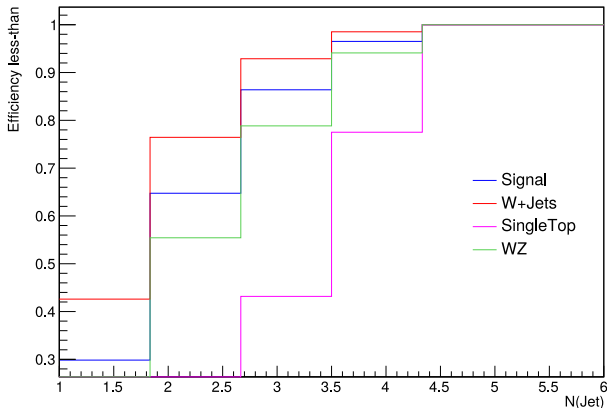


Figure 21: The $N(Jet)$ efficiencies. For the signal at $N(Jet) < 4$ the efficiency is almost 100% while for the main BG is less than 80%.



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In the **Table 2** are summarized the cuts made in order to maximize the significance of the signal, along with the accumulated efficiency after cut.

	W+Jets		Single Top		WZ		TOTAL
XSection (Pb)	3740.26291		288.17		22.82		
	Eff _a	N	Eff _a	N	Eff _a	N	N
$\sigma \times \mathcal{L}$	-	374,026,291.00	-	28,817,000.00	-	2,282,000.00	405,125,291.00
MC	-	6,250,050.00	-	319,006.00	-	173,715.00	6,742,771.00
Basics	1.42%	5,305,450.73	4.34%	1,250,490.66	29.90%	682,295.18	7,238,236.57
$M_T(p_T(\mu), p_T^{miss}) < 40 \text{ GeV}$	0.32%	1,183,471.55	1.17%	336,853.44	5.40%	123,338.22	1,643,663.21
$p_T(\mu) < 20 \text{ GeV}$	0.07%	267,441.89	0.18%	52,032.26	0.63%	14,463.25	333,937.40
$N(Jets) < 4$	0.07%	247,513.77	0.07%	19,421.76	0.48%	11,060.90	277,996.43
$p_T(Jet_{lead}) > 50 \text{ GeV}$	0.07%	247,214.55	0.07%	19,331.42	0.48%	11,060.90	277,606.87
Weight	59.84		90.33		13.14		

Table 2: Weights for the MC samples, based on the expected events, number of MC samples generated for the study, weights for the samples assuming a 100fb^{-1} of luminosity, the number of events, and the accumulated efficiency after each cut.



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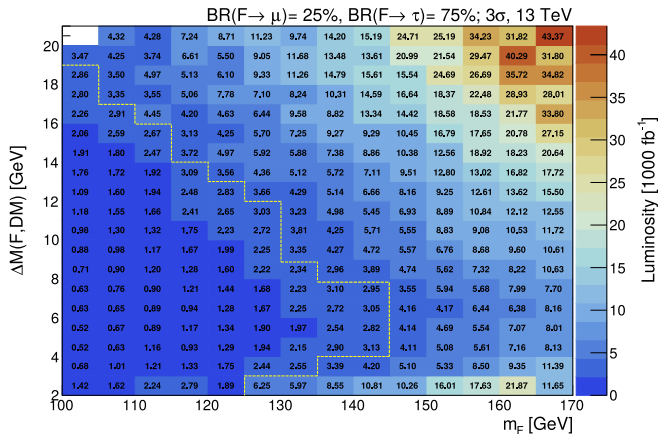


Figure 22: 3σ reach for m_F vs Δm parameter space for various luminosity in 1000 fb^{-1} . The Yellow line is 3000 fb^{-1}



Exclusion Potential



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The exclusion potential of the study is the region of parameters of the model in which at 3000 fb^{-1} the model reaches 3σ . part of this region is still unexplored (see **Figure 8**), the intersection between the regions in **Figures 8** and **22** is shown in the next slide (Figure 23).

The exclusion is carried out assuming that what is going to be observed is the yields due to the SM alone.



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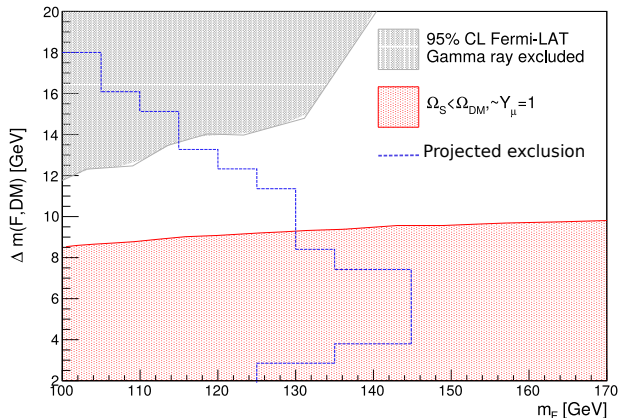


Figure 23: Exclusion potential of the model. The region left of the yellow line is where the model reach 3σ of significance at 3000 fb^{-1} of luminosity



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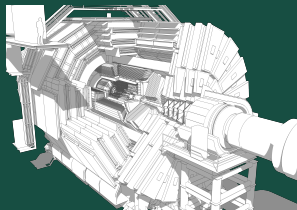
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- 1 Although the model have been analyzed. there is still the region of the parameter space ($\Delta m = m_F - m_S \lesssim 50$ GeV) with exclusion potential.
- 2 The full dark matter content may be explained either by freeze-out at higher redshifts, or by another dark matter content.
- 3 We can obtain exclusion sensitivity over a large region of parameter space has not yet been covered by any other search.

Thank You.

C.Salazar





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Gamma-ray from the Galactic center measured Fermi-LAT Collaboration could be explained by these type of models.

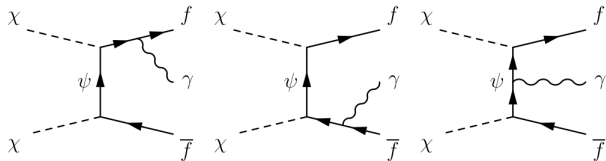


Figure 24: Internal Bremsstrahlung processes of (real) scalar DM. From [Phys Rev Lett. 2013 Aug 30;111(9):091301.]



- Search for top squark pair production in pp collisions at $\sqrt{s} = 13$ TeV using single lepton events (arXiv:1706.04402) 25
- Search for top squarks decaying via four-body or chargino-mediated modes in single-lepton final states in proton-proton collisions at $\sqrt{s} = 13$ TeV (arXiv:1805.05784)

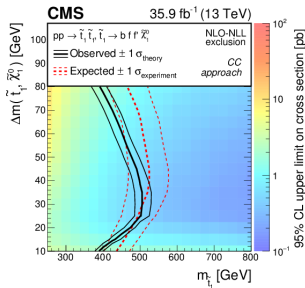


Figure 25: Exclusion limit at 95% for the four-body decay of the top squark as a function of $m(\tilde{t})$ and δm