

# On string vacua without supersymmetry

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## Lessons from SUSY breaking

naturalness + ~~SUSY~~  $\longrightarrow$  high scales

- Support from swampland (Cribiori, Lüst, Scalisi, 2021)

static + ~~SUSY~~  $\longrightarrow$  instabilities

...but our universe *expands*

- Dynamics: a way out?

## Game plan

→ Non-SUSY strings

→ Flux vacua...

→ ...and brane dynamics

→ Holographic interpretation

→ Phenomenology?

# Building a vantage point: string-scale SUSY breaking

Heterotic model (Alvarez-Gaume, Ginsparg, Moore, Vafa, 1986)

$$E_8 \times E_8 \rightarrow SO(16) \times SO(16)$$

→ ~~SUSY~~:  $\Lambda_{\text{quantum}} > 0$

Orientifold models: IIB or 0B

$$O9 + 32 \overline{D9}$$

→  $USp(32)$  (Sugimoto, 1999) or  $U(32)$  (Sagnotti, 1995)

## Brane SUSY breaking

SUSY ✓ in 10d bulk + SUSY ✗ on branes (Antoniadis, Dudas, Sagnotti, 1999)

Low-energy EFT: goldstino + couplings (Dudas, Mourad; Pradisi, Riccioni, 2000)

O9 + 32  $\overline{\text{D9}}$ : residual tension  $V(\phi) = T e^{\gamma \phi}$

No tachyons but back-reaction!

## Low-energy description

$$\mathcal{S}_{\text{eff}} = \int d^{10}x \sqrt{-g} \left( \mathcal{R} - \frac{1}{2}(\partial\phi)^2 - V(\phi) - \frac{e^{\alpha\phi}}{12} H_3^2 + \dots \right)$$

## Taming back-reaction

**IR** ingredients: electric and magnetic fluxes ( $H_3 = dC_2$  or  $dB_2$ )

**UV** ingredients: p-branes (Dudas, Mourad, Sagnotti, 2001)

# Compactifications

Maximally symmetric space-time  $X$

$$ds_{10}^2 = e^{2u(y)} ds_X^2 + ds_Y^2$$

No-go theorem (IB, Lanza, 2020)

$$\Lambda_X \propto \left(1 - (q-1) \frac{\gamma}{\alpha}\right) \int_Y dy \sqrt{g_Y} e^{2cu(y)} V(\phi)$$

$|\partial\mathcal{V}| \geq \mathcal{O}(1) \mathcal{V}$  : dS conjecture holds!

## ...AdS compactifications (Mourad, Sagnotti, 2016)

- Constant dilaton
- $\text{AdS}_3 \times \mathbb{S}^7$  (orientifolds),  $\text{AdS}_7 \times \mathbb{S}^3$  (heterotic)
- electric vs magnetic flux  $N$  of  $H_3 = dC_2$  or  $dB_2$

$$N \gg 1 \quad \longrightarrow \quad e^\phi, (\alpha' \mathcal{R}) \ll 1$$



## Orientifold models: $\text{AdS}_3 \times \mathbb{S}^7$

Parameters:  $V = T e^{\frac{3}{2}\phi}$ , coupling  $\alpha = 1$  to R-R 3-form

$$\text{residual tension: } T = 2k_{10}^2 \times \begin{cases} 64 T_{\text{D9}} & USp(32) \\ 32 T_{\text{D9}} & U(32) \end{cases}$$

electric flux

$$N = \int_{\mathbb{S}^7} \star e^\phi H_3$$

(super)gravity regime

$$L_3, R_7 \propto N^{3/16} \quad e^\phi \propto N^{-1/4}$$

## Heterotic model: $\text{AdS}_7 \times \mathbb{S}^3$

Parameters:  $V = \Lambda e^{\frac{5}{2}\phi}$ , coupling  $\alpha = -1$  to NS-NS 3-form

1-loop vacuum energy:  $\Lambda = (\text{modular integral}) = \frac{\mathcal{O}(1)}{\alpha'}$

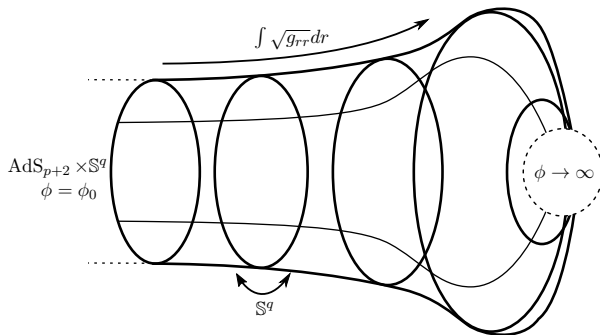
magnetic flux

$$N = \int_{\mathbb{S}^3} H_3$$

(super)gravity regime

$$L_7, R_3 \propto N^{5/8} \quad e^\phi \propto N^{-1/2}$$

## A picture made of branes: near-horizon and pinch-off (Antonelli, IB, 2019)



$SO(1, p) \times SO(q)$  symmetry  $\rightarrow \phi(r), v(r), b(r)$

$$ds^2 = e^{\frac{2}{p+1}v - \frac{2q}{p}b} dx_{1,p}^2 + e^{2v - \frac{2q}{p}b} dr^2 + e^{2b} R_0^2 d\Omega_q^2,$$

$$H_{p+2} = \frac{\textcolor{red}{N}}{e^{\alpha\phi} (R_0 e^b)^q} e^{2v - \frac{q}{p}(p+2)b} d^{p+1}x \wedge dr$$

## Perturbative instabilities (IB, Mourad, Sagnotti, 2018)

AdS vacua: **unstable scalar KK modes** (BF-bound violations)

Orientifolds  $\longrightarrow$   $\ell = 2, 3, 4$   $\text{AdS}_3 \times \mathbb{S}^7$

Heterotic  $\longrightarrow$   $\ell = 1$   $\text{AdS}_7 \times \mathbb{S}^3$

Workarounds:

$\mathbb{S}^q \longrightarrow \mathcal{M}_q$  ...need spectrum of  $\Delta_{\mathcal{M}_q}$

Orbifold? ...ok for  $\mathbb{S}^3$ , hard for  $\mathbb{S}^7$

## Non-perturbative instabilities: brane picture (Antonelli, IB, 2019)

AdS vacua  $\rightarrow$  *flux tunneling* (Brown, Teitelboim, 1987-1988), (Blanco-Pillado, Schwartz-Perlov, Vilenkin, 2009)

$$\mathcal{E}_{\text{vac}} \propto -N^{-3} \quad \text{or} \quad -N^{-2} \quad N \longrightarrow N - \delta N : \text{ out of EFT}$$

Instantons  $\longleftrightarrow$  branes (D1 or NS5)? *right charge & dim.*

AdS  $\rightarrow$  **near-horizon** of brane stack...

$\rightarrow$  *brane-antibrane* nucleation

$$S_{\text{brane}}^E = \left[ \tau_p \text{Area} - \frac{N\mu_p}{e^{\alpha\phi} R^q} \text{Vol} \right]_{\text{extremum}} = B_{\text{CdL}}$$

## Consistency: the right branes

$$S_{\text{brane}}^E = \tau_p \Omega_{p+1} L^{p+1} \left[ \frac{1}{(\beta^2 - 1)^{\frac{p+1}{2}}} - \frac{p+1}{2} \beta \int_0^{\frac{1}{\beta^2-1}} \frac{x^{\frac{p}{2}}}{\sqrt{1+x}} dx \right]$$

### Consistency:

- *existence*: nucl. parameter  $\beta \equiv v_0 \left( \frac{\mu_p}{\tau_p} \right) g_s^{-\frac{\alpha}{2}} > 1$
- *semi-classical*:  $\beta = \mathcal{O}(N^0) \longrightarrow \tau_p = T_p g_s^{-\frac{\alpha}{2}}$

$$\tau_p^{\text{string}} = \frac{T_p}{g_s^\sigma}$$

$$\sigma = 1 + \frac{1}{2} \alpha^{\text{string}}_{\text{electric}}$$

## After tunneling: Lorentzian evolution

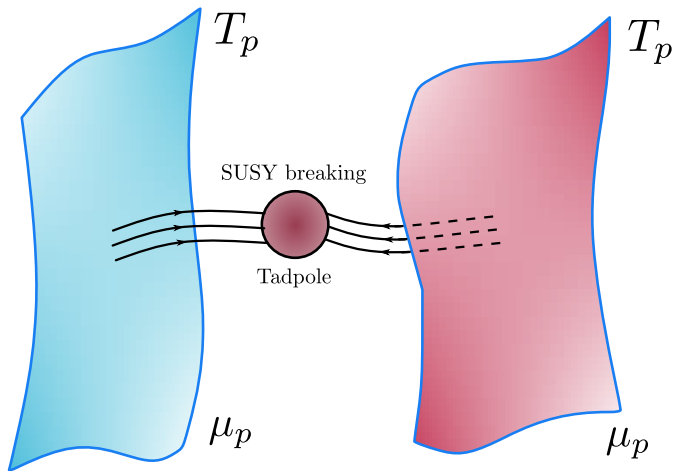
probe  $p/\bar{p}$ -brane in (Poincaré) AdS throat at pos.  $Z$ :

$$V_{\text{probe}} = \tau_p \left( \frac{L}{Z} \right)^{p+1} \left( 1 \pm v_0 \frac{\mu_p}{T_p} \right)$$

for our string models:  $v_0 > 1$ !

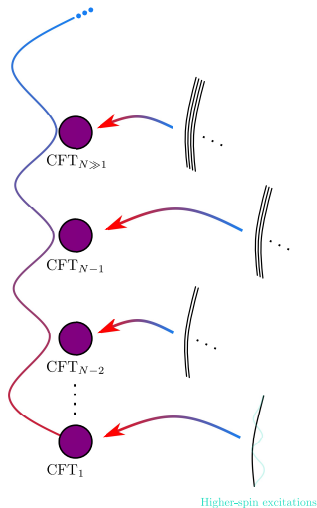
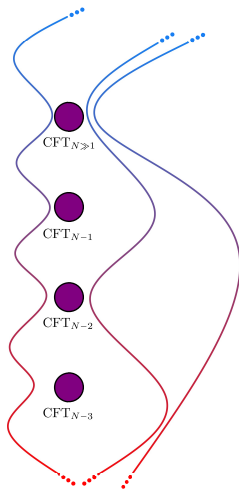
**WGC** :  $\nearrow \left( \frac{\text{charge}}{\text{tension}} \right)_{\text{eff}} \nearrow$

→ these non-SUSY branes **feel the right forces**

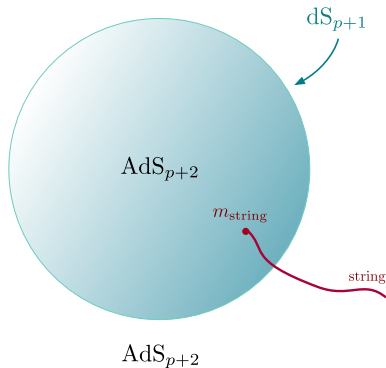




# Holographic interpretation (Antonelli, Basile, IB, 2018)



# Riding the bubble: de Sitter cosmology



## World-volume geometry

$$ds_{\text{bubble}}^2 = -dt^2 + a^2(t) d\Omega_p^2$$

## Junction conditions

Friedmann equations for bubble radius

$$\left(\frac{\dot{a}}{a}\right)^2 = -\frac{1}{a^2} + \frac{\beta^2 - 1}{L^2}$$

Brane bending  $\longrightarrow$  **Einstein eqs.** (Banerjee, Danielsson, Dibitetto, Giri, Schillo, 2019)

Brane action vs Einstein-Hilbert:  $M_{\text{dS}}^{1-p} = \beta (\beta + 1) \frac{M_{\text{AdS}}^{-p}}{\delta L_{\text{AdS}}}$

$$\frac{\mathcal{E}_{\text{dS}}}{M_{\text{dS}}^{p+1}} \propto N^{1 - \frac{\gamma \left(1 + \frac{q}{p}\right)}{(q-1)\gamma - \alpha}} \ll 1$$

Extra stuff: matter, radiation, curvature corrections...

## Outlook

- ~~SUSY~~ brane interactions  $\longrightarrow$  matching in **various regimes** ✓
  - $\rightarrow$  *Probe, amplitudes, near-horizon geometry, gauge theory*
- **Swampland implications** ✓
  - $\rightarrow$  *Weak gravity, distance conjecture, emergent strings, dS vacua...*
- **Non-extremal branes:** **back-reaction** & **fine-tuning**
- Explore **dS brane-world** EFT (*"Swampland on the brane"?*)

## Take-home message

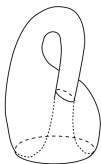
(Natural) SUSY breaking  $\longrightarrow$  Dynamics  $\longrightarrow$  Phenomenology?

**Backup slides**

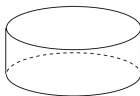
## Heterotic torus amplitude

$$\mathcal{T}_{SO(16) \times SO(16)} = \int_{\mathcal{F}} \frac{d^2\tau}{\tau_2^6} \frac{1}{|\eta(\tau)|^{16}} \left[ O_8 \overline{(V_{16} C_{16} + C_{16} V_{16})} + V_8 \overline{(O_{16} O_{16} + S_{16} S_{16})} \right. \\ \left. - S_8 \overline{(O_{16} S_{16} + S_{16} O_{16})} - C_8 \overline{(V_{16} V_{16} + C_{16} C_{16})} \right]$$

- No massless states (level matching)
- Gravitational sector + adjoint vector
- $(128, 1) \oplus (1, 128)$  spinor
- $(16, 16)$  spinor



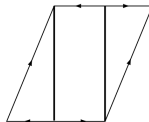
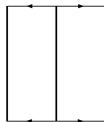
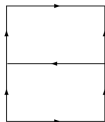
Klein bottle



Annulus



Möbius strip





## Orientifold amplitudes

$$\mathcal{K} = \frac{1}{2} \int_0^\infty \frac{d\tau_2}{\tau_2^6} \frac{(V_8 - S_8)(2i\tau_2)}{\eta^8(2i\tau_2)}, \quad \mathcal{A} = \frac{N^2}{2} \int_0^\infty \frac{d\tau_2}{\tau_2^6} \frac{(V_8 - S_8)\left(\frac{i\tau_2}{2}\right)}{\eta^8\left(\frac{i\tau_2}{2}\right)}$$

Orientifold of IIB  $\longrightarrow$  Klein bottle, annulus, **Möbius strip** ( $N = 32$ )

$$\mathcal{M} = \frac{\varepsilon N}{2} \int_0^\infty \frac{d\tau_2}{\tau_2^6} \frac{\left(\widehat{V}_8 - \widehat{S}_8\right)\left(\frac{i\tau_2}{2} + \frac{1}{2}\right)}{\widehat{\eta}^8\left(\frac{i\tau_2}{2} + \frac{1}{2}\right)}, \quad \mathcal{M}_{\text{BSB}} = \frac{\varepsilon N}{2} \int_0^\infty \frac{d\tau_2}{\tau_2^6} \frac{\left(\widehat{V}_8 + \widehat{S}_8\right)\left(\frac{i\tau_2}{2} + \frac{1}{2}\right)}{\widehat{\eta}^8\left(\frac{i\tau_2}{2} + \frac{1}{2}\right)}$$

# Tadpoles and back-reaction

## String-scale back-reaction from tadpoles

### Workarounds or alternatives?

- Non-tachyonic compactification of **tachyonic** models (Faraggi, Matyas, Percival, 2019-2021)
- **Misaligned SUSY** (Dienes, 1994-2001) (Cribiori, Parameswaran, Tonioni, Wrase, 2020)
- Suppression of vacuum energy (Dienes, 1990) (Kachru, Kumar, Silverstein, 1999) (Angelantonj, Cardella, 2004)
  - *Exponential*: (Abel, Dienes, Mavroudi, Stewart, 2015-2017)
- **New** expansion parameters (**fluxes**)

## Perturbations and mixings

**Linearized analysis:** AdS tensors + angular momenta  $\ell$  (IB, Mourad, Sagnotti, 2018)

→ **Tensors:** no mixing → stable ✓

→ **Vectors:** mixing, still stable ✓  $\delta g_{\mu i}$   $\delta B_{\mu i}$

→ **Scalars:** Einstein eqs. → 2 constraints!

$$\delta\phi \quad \delta B_2 = \star_3 d\mathcal{B}$$

$$\delta g_{\mu\nu} = \mathcal{A} g_{\mu\nu}^{(0)} \quad \delta g_{ij} = \mathcal{C} g_{ij}^{(0)} \quad \delta g_{\mu i} = \nabla_\mu \nabla_i \mathcal{D}$$

## Linearized scalar equations: orientifold case

$$L_3^2 \square A - \left[ 4 + 3\sigma_3 + \frac{\ell_3}{3} (\sigma_3 - 1) \right] A + \frac{7}{2} \alpha \sigma_3 \delta\phi - \frac{\ell_3}{2} (\sigma_3 - 1) B = 0$$

$$L_3^2 \square \delta\phi + 2\alpha\sigma_3 A - \left[ 2\alpha^2\sigma_3 + \tau_3 + \frac{\ell_3}{3} (\sigma_3 - 1) \right] \delta\phi + \alpha \frac{\ell_3}{3} (\sigma_3 - 1) B = 0$$

$$L_3^2 \square B - 8\sigma_3 A + 4\alpha\sigma_3 \delta\phi - \frac{\ell_3}{3} (\sigma_3 - 1) B = 0$$

where:

$$\ell_3 = \ell(\ell + 6) \quad \sigma_3 = 1 + 3 \frac{L_3^2}{R_7^2} \quad \tau_3 = L_3^2 V_0''$$

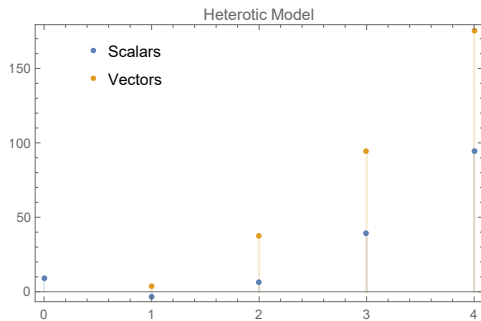
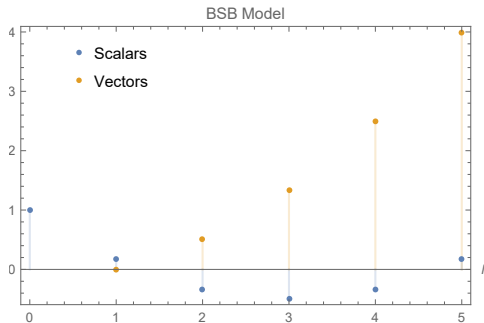
## Linearized scalar equations: heterotic case

$$\begin{aligned}L_7^2 \square A - [\ell_7 (\sigma_7 - 3) + 5 \sigma_7 + 12] A + \frac{5}{2} \alpha \sigma_7 \delta\phi - \frac{3 \ell_7}{2} (\sigma_7 - 3) B &= 0 \\L_7^2 \square \delta\phi + 6 \alpha \sigma_7 A - [2 \alpha^2 \sigma_7 + \tau_7 + \ell_7 (\sigma_7 - 3)] \delta\phi + \alpha \ell_7 (\sigma_7 - 3) B &= 0 \\L_7^2 \square B - 8 \sigma_7 A + 4 \alpha \sigma_7 \delta\phi - \ell_7 (\sigma_7 - 3) B &= 0\end{aligned}$$

where:

$$\ell_7 = \ell (\ell + 2) \quad \sigma_7 = 3 + \frac{L_7^2}{R_3^2} \quad \tau_7 = L_7^2 V_0''$$

## Results: violations of BF bounds



Orbifolds: *can get rid of unstable modes...*  $\longrightarrow$  *vacuum bubbles?* (Horowitz, Orgera, Polchinski, 2008)

## Non-perturbative instabilities: flux tunneling

- Gravity in  $D = p + 2 + q$  dims + fluxes:  $\mathbb{S}^q$  reduction

$$ds^2 = R^{-\frac{2q}{p}} ds_{p+2}^2 + R^2 d\Omega_q^2$$

- Reduced action:  $(p + 2)$ -Einstein frame

$$\mathcal{S}_{p+2} = \frac{1}{2\kappa_{p+2}^2} \int d^{p+2}x \sqrt{-g_{p+2}} \left( \mathcal{R}_{p+2} - 2\Lambda R^{-\frac{2q}{p}} \right)$$

$$\text{vacuum energy} \longrightarrow \mathcal{E}_{\text{vac}} \propto -R^{-\frac{2q}{p}-2}$$

$\mathcal{E}_{\text{vac}}$  depends on flux...

...higher-dim. instantons, flux transitions (Blanco-Pillado, Schwartz-Perlov, Vilenkin, 2009)

## Many branes: background geometry

$SO(1,p) \times SO(q)$  symmetry:  $\phi(r)$ ,  $v(r)$ ,  $b(r)$

$$ds^2 = e^{\frac{2}{p+1}v - \frac{2q}{p}b} dx_{1,p}^2 + e^{2v - \frac{2q}{p}b} dr^2 + e^{2b} R_0^2 d\Omega_q^2$$
$$H_{p+2} = c e^{2v - \frac{q}{p}(p+2)b} d^{p+1}x \wedge dr, \quad c \equiv \frac{N}{e^{\alpha\phi} (R_0 e^b)^q}$$

**(Constrained) Toda-like system:**  $(+, -, +)$  kinetic term w/ potential

$$U = -T e^{\gamma\phi + 2v - \frac{2q}{p}b} - \frac{n^2}{2R_0^{2q}} e^{-\alpha\phi + 2v - \frac{2q(p+1)}{p}b} + \frac{q(q-1)}{R_0^2} e^{2v - \frac{2(D-2)}{p}b}$$



## Geometry: near-horizon

Recover original  $\text{AdS}_{p+2} \times \mathbb{S}^q$  with  $[r < 0]$

$$\phi = \phi_0, \quad e^v = \frac{L}{p+1} \left( \frac{R}{R_0} \right)^{-\frac{q}{p}} \frac{1}{-r}, \quad e^b = \frac{R}{R_0}$$

**Radial perturbations:**  $\delta\phi, \delta v, \delta b \propto (-r)^\lambda$

$$\{\lambda\}_{\text{orient}} = \left\{ -1, \frac{1 \pm \sqrt{13}}{2}, \frac{1 \pm \sqrt{5}}{2} \right\}, \quad \{\lambda\}_{\text{het}} = \left\{ -1, \pm 2\sqrt{\frac{2}{3}}, 1 \pm 2\sqrt{\frac{2}{3}} \right\}$$

→ 2 extremality-breaking deformations

[2 asymptotic fine-tunings?]

## Geometry: “far-horizon”

Away from branes [ $r > 0$ ]: assume  $U \sim U_T = -T e^{\gamma\phi+2v-\frac{2q}{p}b}$

Solutions as  $r \rightarrow \infty$ :  $\phi, v, b \propto y(r) + \text{subleading}$

$$y'' \sim \hat{T} e^{\Omega y + L r}$$

$$\frac{1}{2} \Omega y'^2 + L y' \sim \hat{T} e^{\Omega y + L r} - M$$

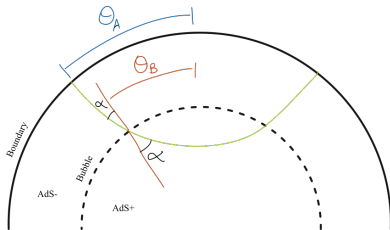
where  $\Omega = \frac{D-2}{8} \gamma^2 - \frac{2(D-1)}{D-2} = \frac{D-2}{8} (\gamma^2 - \gamma_{\text{crit}}^2)$  (IB, Mourad, Sagnotti, 2018)

→ Orientifolds:  $\phi, v, b \propto r^2$  (due to  $\Omega = 0$ )

→ Heterotic:  $\phi, v, b \propto \log(r_0 - r)$

## Our check: bubble entanglement entropy (in 3d)

$$ds_{\pm}^2 = L_{\pm}^2 \left( -\cosh^2 \rho_{\pm} d\tau_{\pm}^2 + d\rho_{\pm}^2 + \sinh^2 \rho_{\pm} d\phi_{\pm}^2 \right)$$

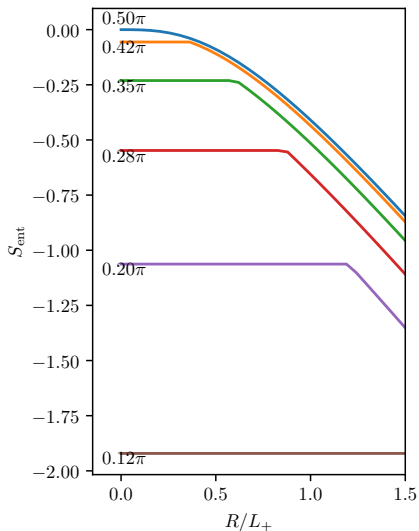


**Thin-wall:** geodesic is hyperbolic polygonal

$$\begin{aligned} \text{length} = & 2L_+ \Lambda + 2L_+ \log(\cosh \rho_+ - \sinh \rho_+ \cos(\theta_B - \theta_A)) \\ & + L_- \cosh^{-1}(\cosh^2 \rho_- - \sinh^2 \rho_- \cos(2\theta_B)) + \mathcal{O}(\Lambda^{-1}) \end{aligned}$$

→ Angle at bubble  $\theta_B$ : **no-kink condition**

## c-functions and “bubble RG”



Finite part of  $S_{\text{ent}}(\theta_A, R)$   
decreases with **bubble radius**:  
candidate  **$c$ -function**!

### Check: trace anomaly

(Henningson, Skenderis, 1998)

$$\langle T^\mu{}_\mu \rangle = \frac{c_{\text{ent}}(R)}{12} \mathcal{R}$$

**Moving the bubble:**  
*can use integral geometry*

(Antonelli, IB, Bombini, 2018)

# Moving the bubble: a primer of integral geometry

**Crofton theorem:** *space of all lines*  $\mathcal{K}$  (Crofton, 1968)

$$\text{length}(\gamma) = \frac{1}{4} \int_{\mathcal{K}} n_{\gamma, \kappa} \omega(\kappa) \leftarrow \text{Crofton form}$$

Asymptotically  $\mathbb{H}^2$  slices

$$\omega_{\text{bubble}} = \Omega \omega_{\mathbb{H}^2}$$

moving bubble via isometry

$\longrightarrow \Omega$  scalar!

