

Question 1:

- Astronomical unit (AU) - average distance between earth and the Sun.

$$1AU = 1.496 \times 10^8$$

(to 4 s.f)

- Solar radius (R_{\odot})

$$1R_{\odot} = 6.955 \times 10^5 \text{ km}$$

(to 4 s.f) \odot is used to notate quantities related to the Sun.

- Light-year (ly), distance light travels in a year through empty space.

$$1ly = 9.461 \times 10^{12} \text{ km}$$

(to 4s.f)

- Parsec (pc)

$$1pc = 3.086 \times 10^{13} \text{ km}$$

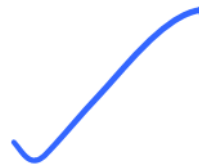
(to 4s.f)

- kilo = 10^3

- mega = 10^6

- giga = 10^9

About 5 mins to read and make notes.

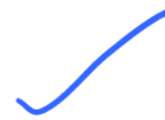
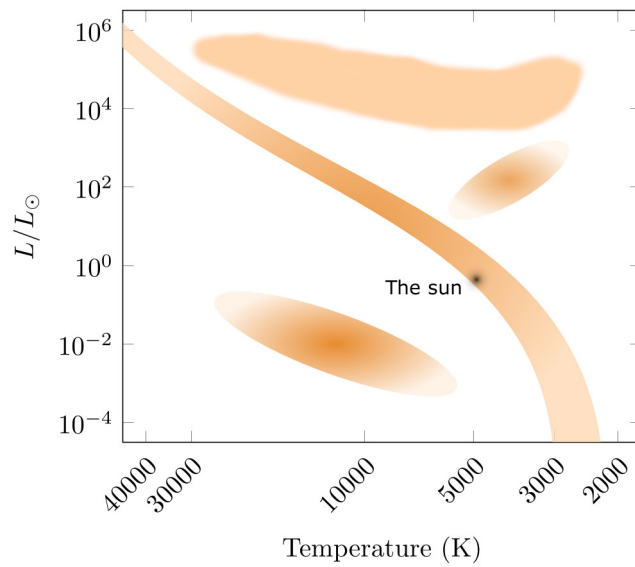


Good

Question 2:

$$\theta^2 = (\Delta\alpha)^2(\cos\delta)^2 + (\Delta\delta)^2$$

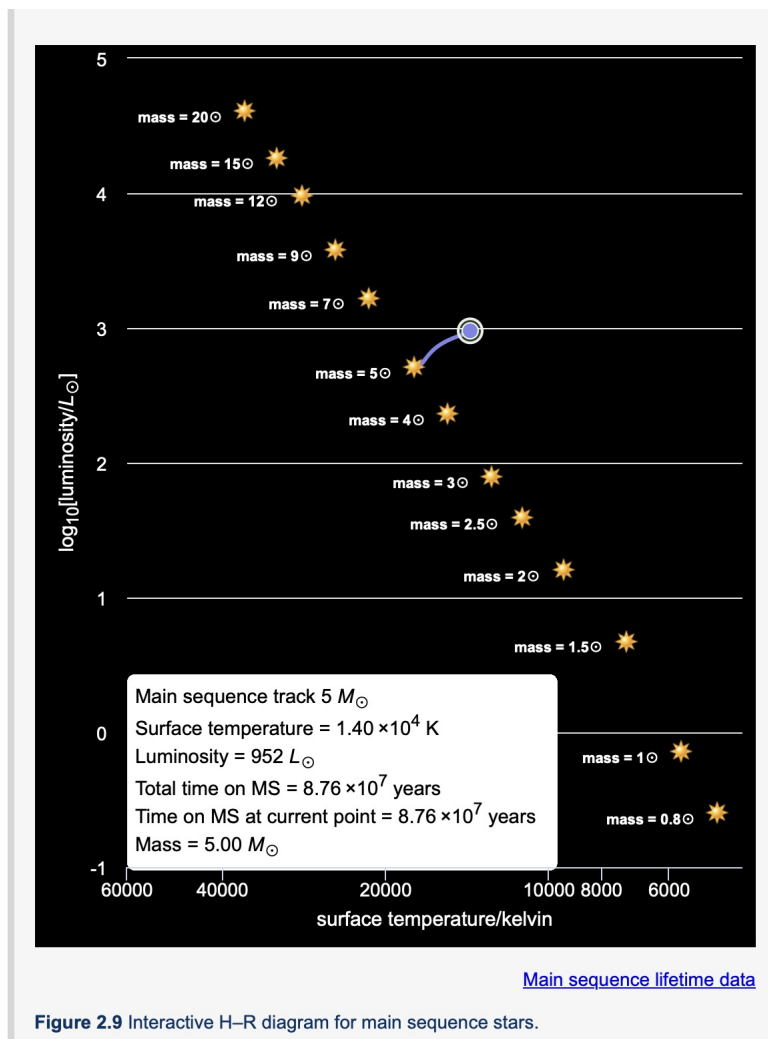
Excellent use of equation editor, spot on.

Question 3:**Part (a):**

Remember to give all your images and diagrams a title, and reference too if appropriate.

Part (b):

Also give a title and reference - what are we looking at?

Part (c):

Investigating the properties of stars on the main sequence, Copyright
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Question 4:

Star	Apparent magnitude(m)	Distance (d/pc)
Alpha Cygni	1.25	800
Beta Cygni	2.93	130
Gamma Cygni	2.23	560
Delta Cygni	2.87	51
Epsilon Cygni	2.48	22

**Good.**

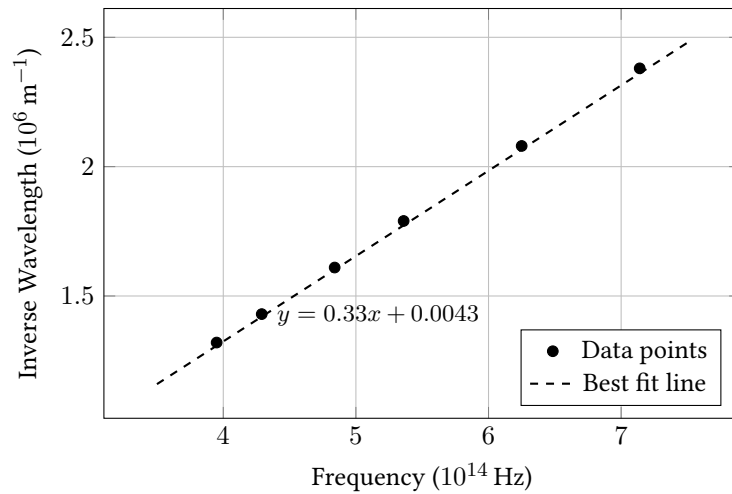
Question 5:

Figure 1: Plot of inverse wave-length versus frequency with a best fit line

paulallen@Pauls – AirTMA00
python3line_of_best_fit.py
Slope : 0.3325256100466257
Intercept : 0.004284972035984325



Okay, but better to swap your axes here, as the gradient then equals the speed of light

$$\begin{aligned}
 c &= \lambda f \\
 f &= \frac{c}{\lambda} \\
 \frac{1}{\lambda} &= \frac{f}{c}
 \end{aligned}
 \tag{1}$$

Using;

$$c \approx \frac{1}{\text{gradient}} \approx 3.01 \times 10^8 \text{ m s}^{-1}$$

Grand Total: 100/100