

1 Local wind regime induced by giant linear dunes:
2 comparison of ERA5-Land ~~re-analysis~~ reanalysis
3 with surface measurements

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10 **Abstract**

11 Emergence and growth of sand dunes results from the dynamic interaction
12 between topography, wind flow and sediment transport. While feedbacks be-
13 tween these variables are well studied at the scale of a single ~~and relatively~~
14 ~~small~~ dune, the average effect of a periodic ~~large-scale~~ dune pattern on atmo-
15 ~~small~~ atmospheric flows remains poorly constrained, due to a lack of data in major sand
16 seas. Here, we compare ~~field-local~~ measurements of surface ~~wind-data-winds~~
17 to the predictions of the ERA5-Land climate reanalysis at four locations in
18 Namibia, ~~including within the giant-dune within and outside the giant linear~~
19 ~~dune~~ field of the Namib sand sea. In the desert plains to the north of the sand
20 sea, observations and predictions agree well. This is also the case in the inter-
21 dune areas of the sand sea, ~~except for the weak winds blowing at night, which~~
22 ~~exhibit additional components during the day. During the night, however, an~~

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23 additional wind component aligned with the giant dune orientation, which
24 are not measured, in contrast to the easterly wind predicted by the ERA5-
25 Land reanalysis. We quantify these similarities and differences and provide a
26 physical understanding of the relevant aerodynamical regimes to relate them
27 link these discrepancies, with wind deviation and velocity attenuation larger
28 than 50° and 60 %, to the daily cycle of the turbulent atmospheric boundary
29 layer over a dune pattern of given wavelength. We conclude by identifying the
30 conditions under which the ERA5 Land reanalysis data can reliably be used
31 to study dune morphodynamics. We also propose that, in multidirectional
32 wind regimes, deflections of specific winds by complex topography, and to the
33 associated flow regimes. During the night, a shallow boundary layer induces a
34 flow confinement associated with a strong streamline compression above the
35 giant dunes, resulting in large flow deviations, especially for the lower
36 winds. During the day, the flow confinement is reduced by a thicker boundary
37 layer and higher wind velocities, and the feedback of the giant dunes on the
38 atmospheric flow is negligible. We finally propose that this mechanism and the
39 resulting wind deflections by the giant dunes could explain the occurrence of
40 smaller-scale secondary dune patterns with elongating along a different orienta-
41 tion compared to the primary structures between which they develop.

42 **Keywords** Atmospheric boundary layer · Sand dunes · Flow over hills

43 1 Introduction

44 The description of turbulent flows over complex topography is relevant for a
45 large variety of different environmental systems (Finnigan et al. 2020)(Sherman 1978; Walmsley et al. 1982; Baines 1995;
46 . For example, the flow over hills is of primary interest for wind power, me-
47 teorological and air pollution phenomena (Taylor et al. 1987). The prop-
48 erties of these flows are also key to the understanding of geophysical phe-
49 nomena, including the formation of wind-driven waves on the ocean surface
50 (Sullivan and McWilliams 2010), dissolution bedforms (Claudin et al. 2017)
51 (Claudin et al. 2017; Guérin et al. 2020), or sedimentary ripples and dunes
52 (Charru et al. 2013; Courrech du Pont 2015)(Bagnold 1941; Charru et al. 2013; Courrech du Pont 2015)
53 . Importantly, the troposphere presents a vertical structure, with a lower con-
54 vective boundary layer, of typical kilometer-scale thickness, capped by a stably
55 stratified region (Stull 1988). The largest topographic obstacles, such as moun-
56 tains, can therefore interact with this upper region and lead to internal wave
57 generation or significant wind disturbances, such as lee-side downslope winds
58 (Durran 1990).

59 ~~Focusing on the wind close to the surface, two related~~ Compared to hills and
60 ~~mountains, aeolian sand dunes offer idealized elevation profiles for the study~~
61 ~~of atmospheric turbulent flow over topographies, due to their smooth shape,~~
62 ~~free of canopies. Besides, dunes provide a rather wide range of scales, from~~
63 ~~decameters to kilometers, and very often come in a fairly regular pattern, which~~
64 ~~further simplifies the flow structure analysis. Past studies have highlighted~~
65 ~~two important topographic feedbacks on the windflow over dunes can be~~
66 ~~commented on separately~~
wind flow close to the dune/hill surface. First is the
67 effect on wind speed, with documented flow acceleration on upwind slopes
68 (Weaver and Wiggs 2011) and deceleration on downwind slopes (Baddock et al.
69 2007), where the speed-up factor is essentially proportional to the obstacle as-
70 pect ratio (Jackson and Hunt 1975). Importantly, the velocity maximum is typ-
71 ically shifted upwind of the obstacle crest (Jackson and Hunt 1975; Claudin et al. 2013)
72 . This behaviour has been theoretically predicted by means of asymptotic anal-
73 ysis of a neutrally stratified boundary-layer flow over an obstacle of vanishing
74 aspect ratio (Jackson and Hunt 1975; Mason and Sykes 1979; Sykes 1980; Hunt et al. 1988; Belcher and J.C.R. 1998)
75 (Jackson and Hunt 1975; Mason and Sykes 1979; Sykes 1980; Hunt et al. 1988; Belcher and J.C.R. 1998; Kroy et al. 2000)
76 . Experiments in flumes (Zilker et al. 1977; Zilker and Hanratty 1979; Fred-
77 erick and Hanratty 1988; Poggi et al. 2007; Bristow et al. 2022), in wind
78 tunnels (Gong and Ibbetson 1989; Finnigan et al. 1990; Gong et al. 1996)
79 and in field conditions at all scales (Taylor and Teunissen 1987; Claudin et al.
80 2013; Fernando et al. 2019; Lü et al. 2021), have also documented this ef-
81 fect. Interestingly, a similar behaviour exists for the pressure perturbation,
82 but with a slight downwind shift for the pressure minimum (Claudin et al.
83 2021). The second effect, much less studied, is the flow deflection that oc-
84 curs when the incident wind direction is not perpendicular to the ridge crest.
85 While predicted to be small (less than 10°) in the linear regime valid for
86 shallow topography (Gadal et al. 2019), significant flow steering has been
87 reported in the field on the downwind side of steep enough obstacles, such

as mountain ranges (Kim et al. 2000; Lewis et al. 2008; Fernando et al. 2019), well-developed sand dunes (Walker et al. 2009; Hesp et al. 2015; Walker et al. 2017; Smith et al. 2017; de Winter et al. (Tsoar and Yaalon 1983; Sweet and Kocurek 1990; Walker and Nickling 2002; Smith et al. 2017) and in particular coastal foredunes (e.g. Hunter et al. (1983), Rasmussen (1989), Walker et al. (2006), Walker et al. (2009), Hesp et al. (2015), Walker et al. (2017), de Winter et al. (2020)), mountain ranges (Kim et al. 2000; Lewis et al. 2008; Fernando et al. 2019), and valley topographies (Wiggs et al. 2002; Garvey et al. 2005).

For practical reasons, wind measurement Wind measurements over sand dunes has been mainly performed over small bedforms, typically a few meters high (corresponding to several tens of meters long) (e.g. Mulligan (1988) Hesp et al. (1989), Lancaster et al. (1996), Mckenna Neuman et al. (1997), Sauermann et al. (2003), Andreotti et al. (2002), Walker and Nickling (2002), Weaver and Wiggs (2011)). Giant dunes For practical reasons, fewer studies performed similar measurements on giant dunes (Hayholm and Kocurek 1988), with kilometer-scale wavelengths and heights of tens of meters, are more difficult to investigate although for several reasons they. However, such large dunes provide a choice configuration for the study of turbulent flows over a complex topography. First, one expects larger wind disturbances for larger obstacles. Secondly, their large size makes can make them interact with the vertical structure of the atmosphere (Andreotti et al. 2009). Third, they usually form large patterns in sand seas and thus behave as rather clean periodic perturbations, in contrast with isolated dunes. Finally, because the morphodynamics of aeolian bedforms are is strongly dependent on the local wind regime (Livingstone and Warren 2019), one can expect to see the consequences of windflow disturbance by large dunes on neighbouring small dunes (Brookfield 1977; Ewing et al. 2006). A similar effect is observed on the properties of impact ripple patterns due to the presence of dunes (Howard 1977; Hood et al. 2021) (Howard 1977; Hood et al. 2021).

Atmospheric flows have been much studied at the desert-scale with climate reanalyses based on global atmospheric models (Blumberg and Greeley 1996; Livingstone et al. 2010; Ashkenazy et al. 2012; Blumberg and Greeley 1996; Livingstone et al. 2010; Ashkenazy et al. 2012; Jolivet et al. 2021; Hu et al. 2021; Gunn et al. 2021), such as ERA-40, ERA-Interim or ERA-5 ERA5 (Uppala et al. 2005; Dee et al. 2011; Hersbach et al. 2020). However, the spatial resolution of these reanalyses (tens of kilometers) of these reanalyses implies average quantities that do not resolve the smaller scales of interest, which range from individual dunes to small mountains (Livingstone et al. 2010). Recently, the release of ERA5-Land has partly resolved this limitation by providing up to 70 years of hourly wind predictions at a 9 km spatial resolution (Muñoz-Sabater et al. 2021). However, its validity remains to be studied, especially in remote desert areas where assimilation of measured data is very low.

In this work, we compare local wind speeds and directions measured by meteorological stations at four different locations inside and north of the giant-dune giant linear dune field of the Namib sand sea to the regional predictions of the ERA5-Land climate reanalysis. Where the meteorological stations are surrounded by a relatively flat environment, we show that local measurements and regional predictions agree well. The agreement is also good in the

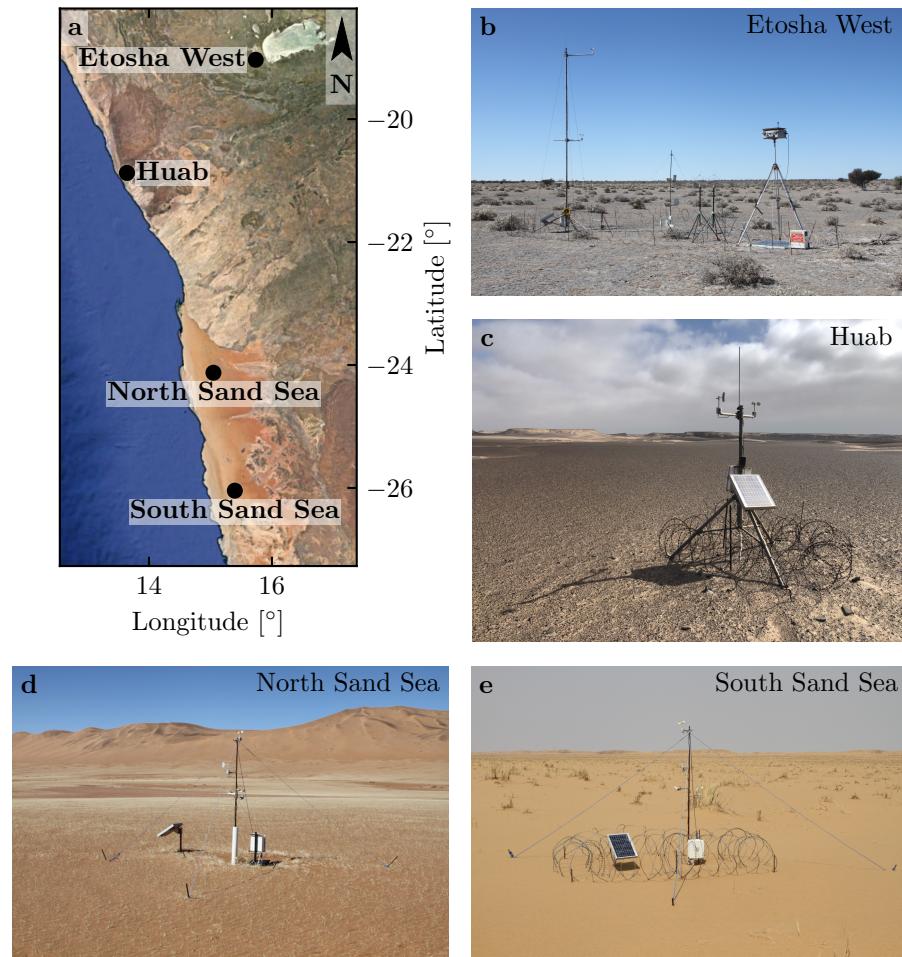


Fig. 1 Wind data used in this study. Studied field sites. **a:** Location of the different sites in Namibia. **b–e:** Satellite images. Photographs of these different environments (Google-Earth, Maxar Technologies, CNES/Airbus). In each subplot, the left and right wind roses represent the data from the ERA5-Land climate reanalysis and the local wind stations, respectively. Note that the bars show the direction towards which the wind blows. The black dots show the location of local wind meteorological stations.

134 interdune areas of the sand sea, except for some weak winds blowing at night,
 135 which exhibit an additional component aligned with the giant dune orientation.
 136 These winds are not predicted by the ERA5-Land reanalysis (section 2).
 137 Further, we are able to link the magnitude of these differences to the circadian
 138 cycle of the atmospheric boundary layer (section 3). Finally, we draw implica-
 139 tions for the wind disturbances on smaller-scale dunes (section 4), suggesting
 140 a possible origin for crossing dunes.

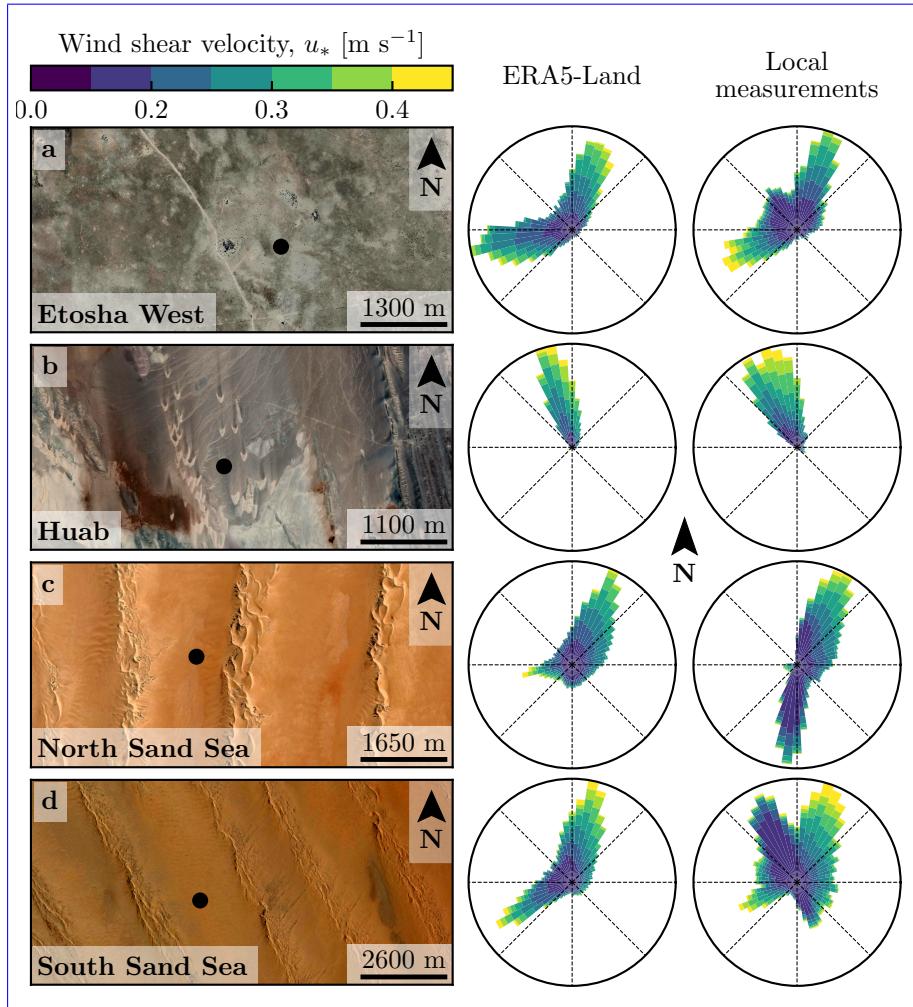


Fig. 2 Wind data used in this study. Satellite images of the different environments (Google-Earth, Maxar Technologies, CNES/Airbus) are shown on the left. The black dots show the location of the wind measurements stations. On the right of the photos, the corresponding wind roses representing the data from the ERA5-Land climate reanalysis and the local wind stations are displayed. Note: the graphical convention for the wind roses is that the bars show the direction towards which the wind blows (see color bar for velocity scale).

141 2 Wind regimes across the Namib Sand Sea

142 We measured the wind regime at four different locations in Namibia, repre-
 143 sentative of various arid environments across the Namib desert (Fig. 1, [Online](#)
 144 [Resource](#) Fig. 42). The Etosha West station was located at the Adamax wa-
 145 terhole to the west of Etosha Pan in northern Namibia, in a sparsely vegetated

area. The Huab station was near the coast on a hyper-arid flat gravel plain lying north the ephemeral Huab river. Here, barchan dunes up to a few meters in height develop from the sediment blowing out of the river valley (Nield et al. 2017; Hesp and Hastings 1998). These two stations were both located in relatively flat environments. In contrast, the North Sand Sea and South Sand Sea stations were located in the interdunes between linear dunes with kilometer-scale wavelengths, hectometer-scale heights and superimposed patterns. In this section, we describe and compare winds from local measurements and climate reanalysis predictions.

155 2.1 Wind and elevation data

156 At each meteorological station (Fig. 1), wind speed and direction were sampled every 10 minutes using cup anemometers (Vector Instruments A100-LK) 157 and wind vanes (Vector Instruments W200-P) at ~~heights which varied a single~~ 158 ~~height, which was~~ between 2 m and 3 m depending on the station. The available 159 period of measurements at each station ranged from 1 to 5 discontinuous 160 years distributed between 2012 and 2020 (Online Resource Fig. S24). We 161 checked that at least one complete seasonal cycle was available for each 162 station. Regional winds were extracted at the same locations and periods from 163 the ERA5-Land dataset, which is a replay at a smaller spatial resolution of 164 ERA5, the latest climate reanalysis from the ECMWF (Hersbach et al. 2020; 165 Muñoz-Sabater et al. 2021). This dataset provided hourly predictions of the 166 10-m wind velocity and direction at a spatial resolution of $0.1^\circ \times 0.1^\circ$ ($\simeq 9$ km 167 in Namibia).

168 To enable direct comparison, the local wind measurements were averaged 169 into 1-hr bins centered on the temporal scale of the ERA5-Land estimates 170 (Online Resource Fig. S3S2). As the wind velocities of both datasets were pro- 171 vided at different heights, we converted them into shear velocities u_* (Online 172 Resource section 1), characteristic of the turbulent wind profile. Wind roses 173 in Fig. 1(b-e) show the resulting wind data.

174 Dune properties were computed using autocorrelation on the 30-m Digital 175 Elevation Models (DEMs) of the shuttle radar topography mission (Farr et al. 176 2007). For the North and South Sand Sea stations, we obtain, respectively, 177 orientations of 85° and 125° with respect to the North, wavelengths of 2.6 km 178 and 2.3 km and amplitudes (or half-heights) of 45 m and 20 m (Online Resource 179 Fig. S5-S4 for more details). This agrees with direct measurements made on 180 site.

182 2.2 Comparison of local and regional winds

183 The measured and predicted wind regimes are shown in Fig. 12. In the Namib, 184 the regional wind patterns are essentially controlled by the sea breeze, result- 185 ing in strong northward components (sometimes slightly deviated by the large

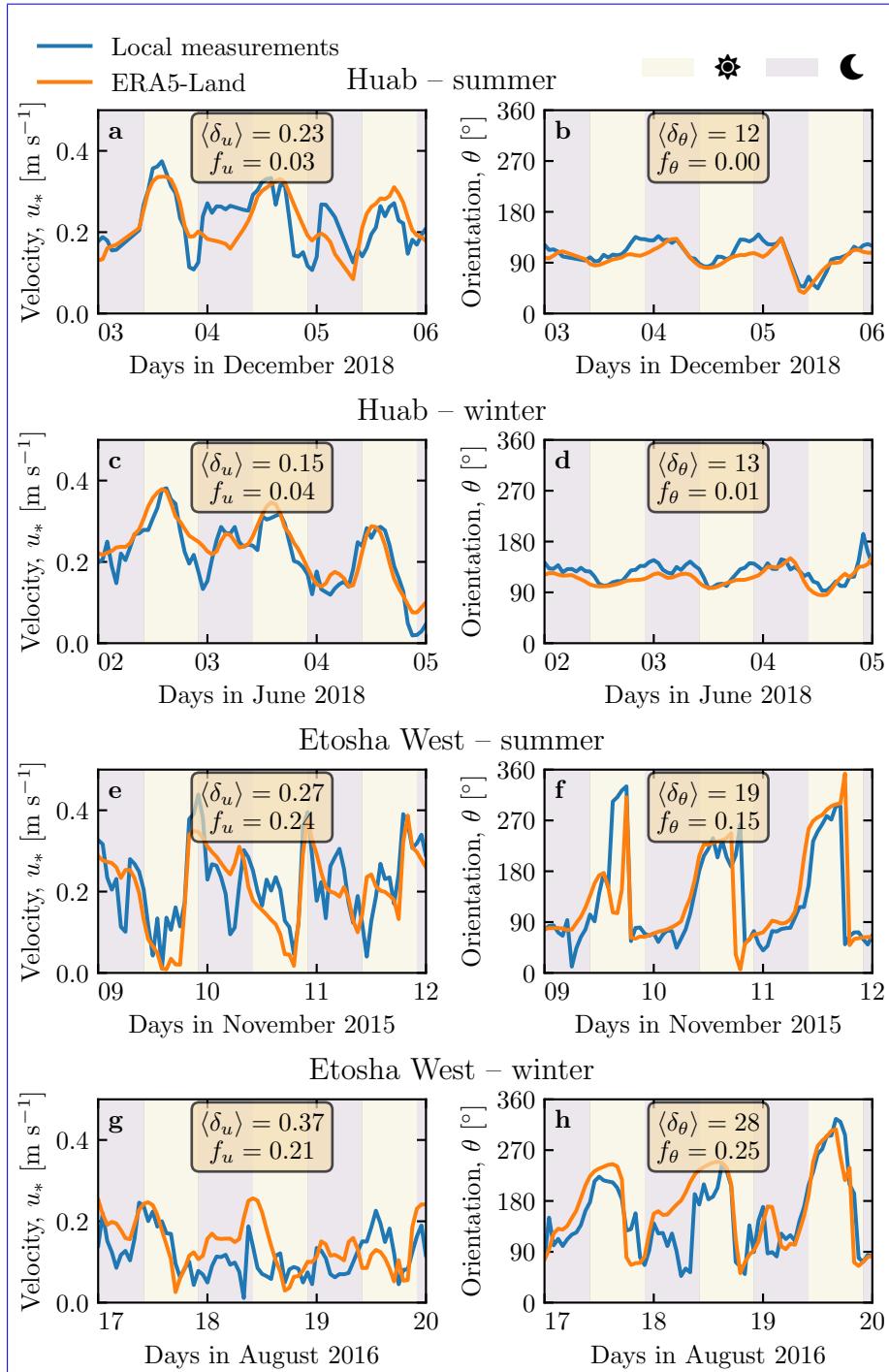


Fig. 3 Temporal comparison between the wind data coming from the ERA5-Land climate reanalysis (orange lines) and from the local measurements (blue lines). Coloured swatches indicate day (between 1000–10.00 UTC and 2200–22.00 UTC) and night (before 1000–10.00 UTC or after 2200–22.00 UTC). Numbers in legends indicate the average flow deflection δ_θ and relative wind modulation δ_u over the displayed period (see section 3.2 for their definitions), as well as the percentage f_θ and f_u of occurrence of extreme events ($\delta_\theta > 50^\circ$, $|\delta_u| > 0.6$). **a–b:** Etosha West–Huab station in summer. **b–c:** Etosha West–Huab station in winter. **d–e:** North–Sand–Sea–Etosha West station in summer. **f–g:** North–Sand–Sea–Etosha West station in winter. Time series of the two other stations are shown in Online Resource Fig. S65.

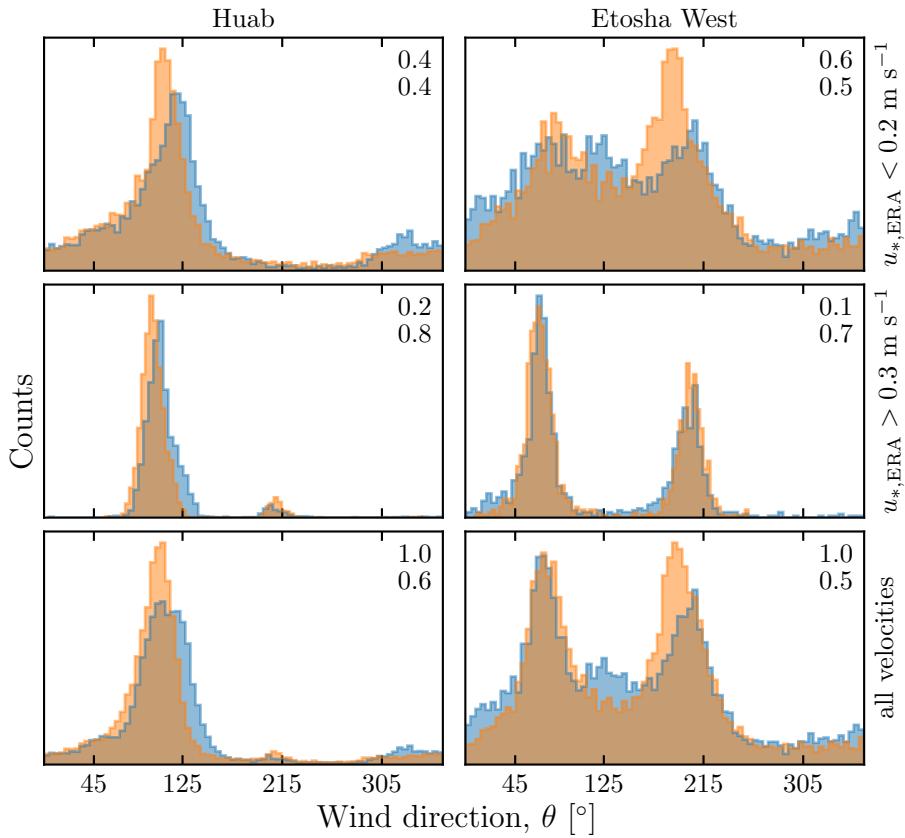


Fig. 4 Distributions of wind direction at Huab and Etosha West stations for the ERA5-Land climate reanalysis (orange) and the local measurements (blue). In each subplot, both distributions are plotted from the same time steps, selected for different ranges of the wind wind velocity (rows) in the ERA5-Land dataset. The numbers at the upper right corners give the percentage of time steps selected in each sub-range (top), as well as the percentage of them corresponding to the day – defined between 10.00 UTC and 22.00 UTC (bottom).

scale topography) present in all regional wind roses (Lancaster 1985). These daytime winds are dominant during the period October-March (Fig. 23f and Online Resource Fig. S64f). During April-September, an additional (and often nocturnal) easterly component can also be recorded, induced by the combination of katabatic winds forming in the mountains, and infrequent ‘berg’ winds, which are responsible for the high wind velocities observed (Lancaster et al. 1984). The frequency of these easterly components decreases from inland to the coast. As a result, bidirectional wind regimes within the Namib Sand Sea and at the west Etosha site (Fig. 1b,d,e2a,c,d) and a unidirectional wind regime on the coast at the outlet of the Huab River (Fig. 4e2b) are observed.

In the case of the Etosha West and Huab stations, the time series of wind speed and direction from the regional predictions quantitatively match those

corresponding to the local measurements ([Fig. 2a–d](#)) [Figs. 3, 4](#) and Online Resource [Figs. S6a–d, S7, S9](#)[Fig. S5](#)). For the North Sand Sea and South Sand Sea stations within the giant [linear](#) dune field, we observe that this agreement is also good, but limited to the October-March time period ([Fig. 2e–h](#) and [Online Resource Fig. S6e–h](#)[4a, b, e, f](#)). However, the field-measured wind roses exhibit additional wind components aligned with the [giant](#)-dune orientation, as evidenced on the satellite images ([Fig. 12c,d](#)).

More precisely, during the April-September period, the local and regional winds in the interdune match during daytime only, i.e when the southerly/-southwesterly sea breeze dominates ([Figs. 2e,f](#) and [3](#), [Online Resource Fig. S8](#)[5c,d,g,h](#) and [6](#)). In the late afternoon and during the night, when the easterly ‘berg’ and katabatic winds blow, measurements and predictions differ. In this case, the angular wind distribution of the local measurements exhibits two additional modes corresponding to reversing winds aligned with the [giant](#)-dune orientation ([purple frame in Fig. 36](#), [Online Resource Figs. S8 and S10](#)[Fig. S6](#)). This deviation is also associated with a general attenuation of the wind strength ([Online Resource Fig. S11](#)[S7](#)). Remarkably, all these figures show that these wind reorientation and attenuation processes occur only at low velocities of the regional wind, typically for $u_{*, ERA} \lesssim 0.2 \text{ m s}^{-1}$ $u_{*, ERA5-Land} \lesssim 0.2 \text{ m s}^{-1}$. For shear velocities larger than $u_{*, ERA} \approx 0.3 \text{ m s}^{-1}$ $u_{*, ERA5-Land} \approx 0.3 \text{ m s}^{-1}$, the wind reorientation is not apparent. Finally, for intermediate shear velocities, both situations of wind flow reoriented along the dune crest and not reoriented can be successively observed ([Online Resource Fig. S10](#)[S6](#)). Importantly, these values are not precise thresholds ([and certainly not related to the threshold for sediment transport](#)), but indicative of a crossover between regimes, whose physical interpretation is discussed in the next section.

224 3 Influence of wind speed and circadian cycle on the atmospheric 225 boundary layer

The wind deflection induced by [linear](#)-dunes has previously been related to the incident angle between wind direction and crest orientation, with a maximum deflection evident for incident angles between 30° and 70° (Walker et al. 2009; Hesp et al. 2015). In the data analysed here, the most deflected wind at both the North and South Sand Sea stations is seen to be where the incident angle is perpendicular to the giant dunes ([Figs. 4](#) and [3](#), [Online Resource Fig. S8](#)[2](#) and [6](#)). It therefore appears that in our case, the incident wind angle is not the dominant control on maximum wind deflection. Further, and as shown in [Fig. 36](#), winds of high and low velocities show contrasting behaviour in characteristics of deflection. This suggests a change in hydrodynamical regime between the winds. In this section, we discuss the relevant parameters associated with the dynamical mechanisms that govern the interactions between the atmospheric boundary layer flow and giant dune topographies. This analysis allows us to provide a physics-based interpretation of our measured wind data.

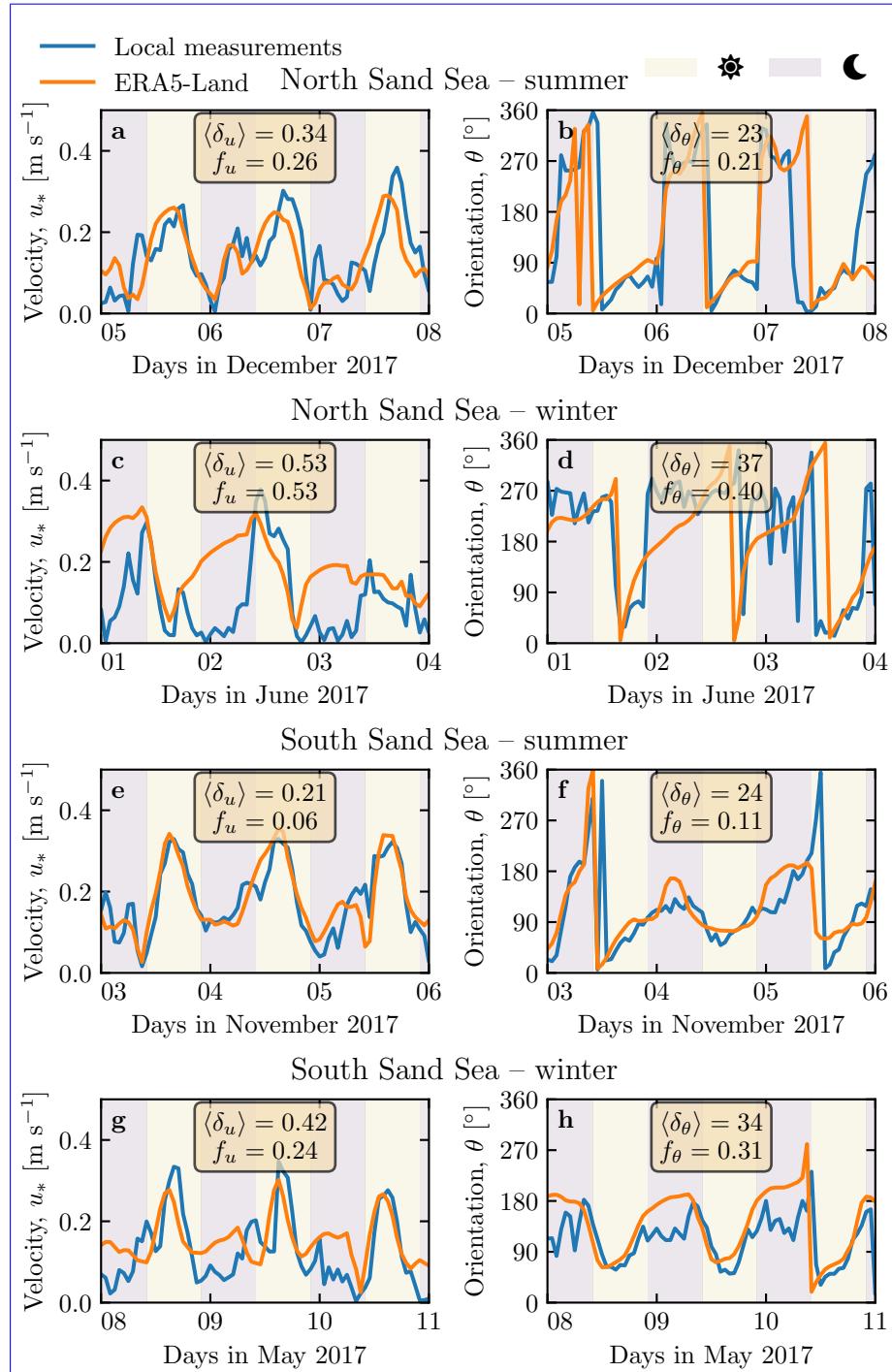


Fig. 5 Distributions of wind direction at the Same as Fig. 3 for North Sand Sea Station for the ERA5 Land climate reanalysis (orange) and the local measurements station in summer (bluea–b). In each subplot, both distributions are plotted from the same time steps, selected with constraints on the wind direction (columns) and/or wind velocity (rows) of the ERA5 Land dataset. The grey vertical dashed lines indicate the dune orientation. The numbers at the top right give the percentage of time steps selected North Sand Sea station in each sub-range, as well as the percentage corresponding to the daytime winter (between 1000 UTC and 2200 UTCb–c). The purple frame highlights the regime (low wind velocities, nocturnal easterly wind) South Sand Sea station in which the data from both datasets differ. A similar figure can be obtained for the North summer (d–e) and South Sand Sea station in winter (Online Resource Fig. S8f–g).

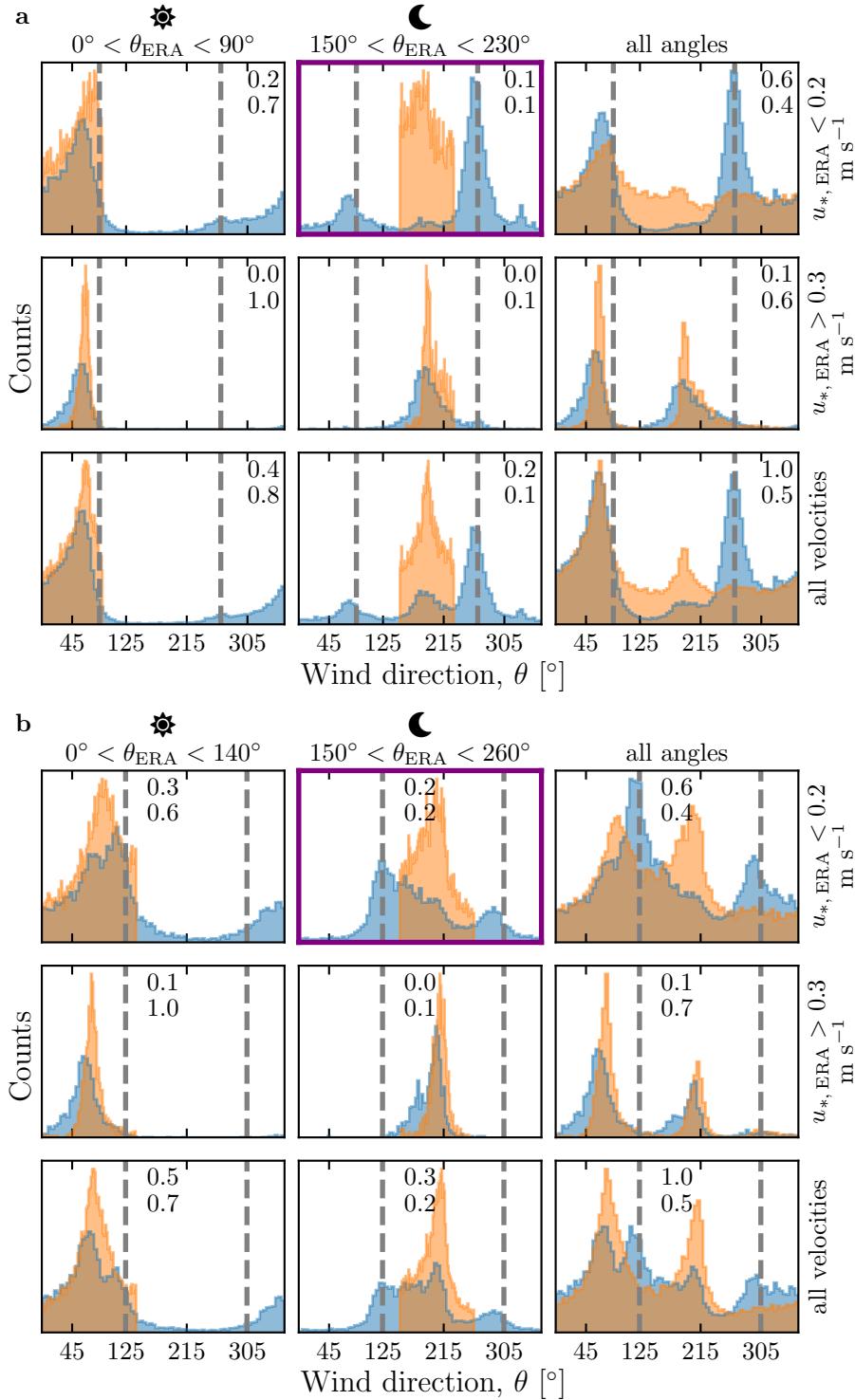


Fig. 6 Same as Fig. 4 but for North Sand Sea (a) and South Sand Sea (b) stations. Here, subplots correspond to different ranges for the wind direction (columns) and wind velocity (rows) of the ERA5-Land dataset. The grey vertical dashed lines indicate the main dune orientation. In contrast with observations at the Huab and Etosha West stations (Fig. 4), histograms do not match well at low wind velocities, and the purple frame highlights the regime (low wind velocities, nocturnal easterly wind) in which the data from both datasets differ most.

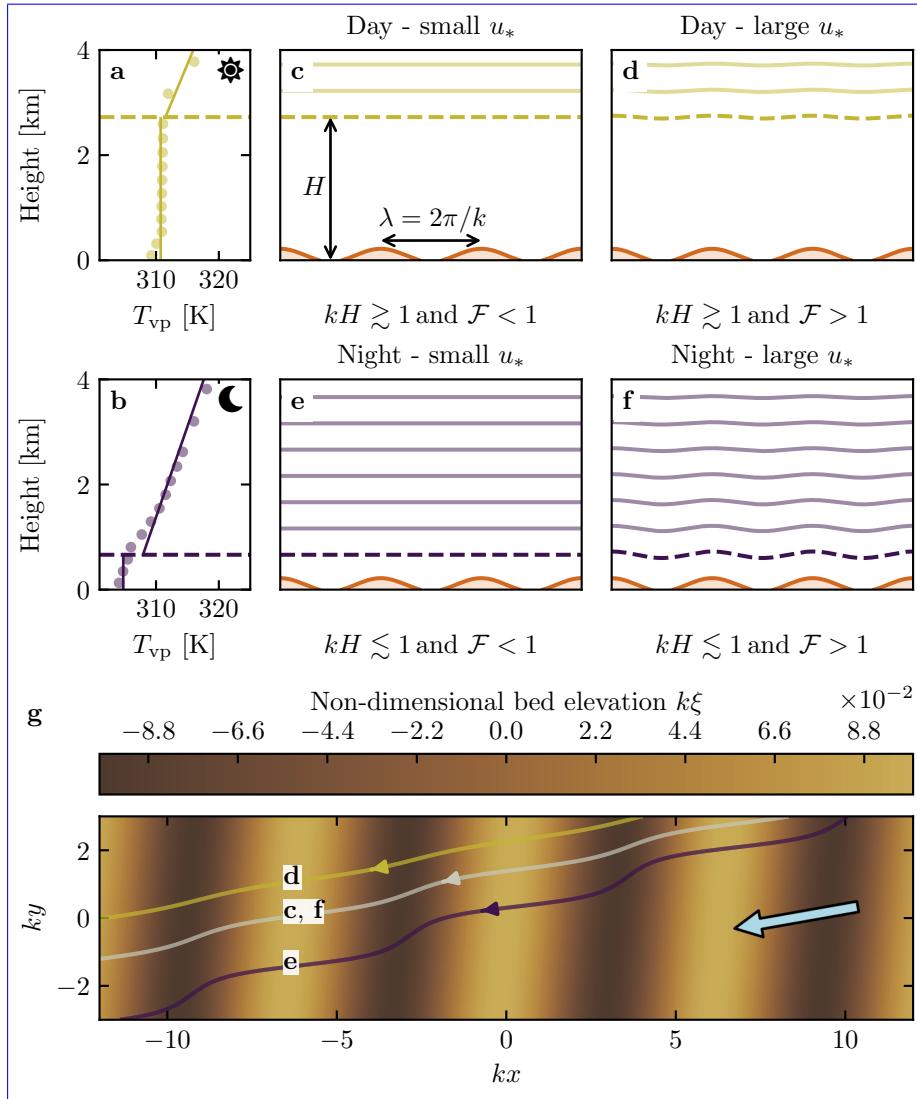


Fig. 7 a–b: Vertical profiles of the virtual potential temperature T_{vp} at two different time steps (day - 03/11/2015 - 1200–12.00 UTC, night - 01/13/2013 - 0900–09.00 UTC) at the North Sand Sea station. Dots: data from the ERA5 reanalysis. Dashed lines: boundary layer height given by the ERA5 reanalysis (Online Resource section 2). Plain lines: vertical (boundary layer) and linear (free atmosphere) fits to estimate the stratification properties. c–f: Sketches representing the interaction between the giant dunes and the atmospheric flow for different meteorological conditions. g: Streamlines over a sinusoidal topography $\xi(x, y)$ qualitatively representing the effect of low, medium and strong flow confinement, in relation to the above panels (see Appendix 1 for more details). The blue arrow indicates the undisturbed wind direction.

240 3.1 Flow over a modulated bed

241 Taking as a reference the turbulent flow over a flat bed, the general framework
 242 of our study is understanding and describing the flow response to a bed mod-
 243 ulation (e.g. a giant dune). Without loss of generality, we can consider in this
 244 context an idealised bed elevation in the form of parallel sinusoidal ridges, with
 245 wavelength λ (or wavenumber $k = 2\pi/\lambda$) and amplitude ξ_0 , and where the ref-
 246 erence flow direction makes a given incident angle with respect to the ridge
 247 crest (Andreotti et al. 2012). Part of this response, on which we focus here,
 248 is the flow deflection by the ridges. In a simplified way, it can be understood
 249 from the Bernoulli principle (Hesp et al. 2015): as the flow approaches the
 250 ridge crest, the compression of the streamlines results in larger flow velocities,
 251 and thus lower pressures (Rubin and Hunter 1987)(Jackson and Hunt 1975).
 252 An incident flow oblique to the ridge is then deflected towards lower pressure
 253 zones, i.e towards the crest. Turbulent dissipation tends to increase this effect
 254 downstream, resulting in wind deflection along the crest in the lee side (Gadal
 255 et al. 2019).

256 Flow confinement below a capping surface, which enhances streamline com-
 257 pression, has a strong effect on the hydrodynamic response and typically in-
 258 creases flow deflection. This is the case for bedforms forming in open channel
 259 flows such as rivers (Fourrière et al. 2010; Unsworth et al. 2018)(Kennedy 1963; Chang and Simons 1970; Mizumura 1995
 260 . This is also relevant for aeolian dunes as they evolve in the turbulent atmo-
 261 spheric boundary layer (ABL) capped by the stratified free atmosphere (FA)
 262 (Andreotti et al. 2009). Two main mechanisms, associated with dimensionless
 263 numbers must then be considered (Fig. 47). First, topographic obstacles typi-
 264 cally disturb the flow over a characteristic height similar to their length. As
 265 flow confinement is characterised by a thickness H , the interaction between
 266 the dunes and the wind in the ABL is well captured by the parameter kH .
 267 The height H is directly related to the radiative fluxes at sensitive heat flux
 268 from the Earth surface. It is typically on the order of a kilometre, but sig-
 269 nificantly varies with the circadian and seasonal cycles. Emerging and small
 270 dunes, with wavelengths in the range 20 to 100 m, are not affected by the
 271 flow confinement, corresponding to $kH \gg 1$. For giant dunes with kilometre
 272 kilometer-scale wavelengths, however, their interaction with the FA is can be
 273 significant (Andreotti et al. 2009). This translates into a parameter kH in the
 274 range 0.02–5, depending on the moment of the day and the season. A second
 275 important mechanism is associated with the existence of a thin intermediate
 276 so-called capping layer between the ABL and the FA. It is characterised by a
 277 density jump $\Delta\rho$, which controls the ‘rigidity’ of this interface, i.e. how much
 278 its deformation affects streamline compression. This is usually quantified using
 279 the Froude number (Vosper 2004; Stull 2006; Sheridan and Vosper 2006; Hunt
 280 et al. 2006; Jiang 2014):

$$\mathcal{F} = \frac{U}{\sqrt{\frac{\Delta\rho}{\rho_0} g H}}, \quad (1)$$

where U is the wind velocity at the top of the ABL and ρ_0 its average density. The intensity of the stratification, i.e. the amplitude of the gradient $|\partial_z \rho|$ in the FA, also impacts its ability to deform the capping layer under the presence of an underlying obstacle, and thus affects the influence of flow confinement. This can be quantified using the internal Froude number (Vosper 2004; Stull 2006; Sheridan and Vosper 2006; Hunt et al. 2006; Jiang 2014) $\mathcal{F}_1 = kU/N$, where $N = \sqrt{-g\partial_z \rho/\rho_0}$ is the Brunt-Väisälä frequency (Stull 1988). Both Froude numbers have in practice the same qualitative effect on flow confinement (a smaller Froude corresponding to a stiffer interface), and we shall restrict the main discussion to \mathcal{F} only.

With this theoretical framework in mind, and in the context of the measured wind data in the North and South Sand Sea stations, the smallest wind disturbances are expected to occur during the day, when the ABL depth is the largest and comparable to the dune wavelength ($kH \gtrsim 1$), which corresponds to a weak confinement situation (Fig. 47c,d). In contrast, large wind disturbances are expected to occur during the night, when the confinement is mainly induced by a shallow ABL (Fig. 47e). However, this strong confinement can be somewhat reduced in the case of strong winds, corresponding to large values of the Froude number and a less ‘rigid’ interface (Fig. 47f). This is in qualitative agreement with the transition from deflected to non-deflected winds related to low and high velocities observed in our data (Sec. 2.2).

3.2 Data distribution in the flow regimes

We can go one step further and analyse how our data quantitatively spread over the different regimes discussed above. For that purpose, one needs to compute kH and \mathcal{F} from the time series. H , U and the other atmospheric parameters can be deduced from the various vertical profiles (temperature, humidity) available in the ERA5 climate reanalysis (Online Resource section 2). We quantify the flow deflection δ_θ as the minimal angle between the wind orientations comparing the local measurements and the regional predictions. We also compute the relative velocity modulation as

$$\delta_u = \frac{\frac{u_*^{\text{ERA}} - u_*^{\text{station}}}{u_*^{\text{ERA}}} \frac{u_*^{\text{ERA5-Land}} - u_*^{\text{Local mes.}}}{u_*^{\text{ERA5-Land}}}}{\frac{u_*^{\text{ERA}}}{u_*^{\text{ERA5-Land}}}}. \quad (2)$$

These two quantities are represented as maps in the plane (\mathcal{F} , kH) (Fig. 58a,b), and one can clearly identify different regions in these graphs. Small wind disturbances (small δ_θ and δ_u) are located in the top-right part of the diagrams, corresponding to a regime with low-interaction as well as low-confinement (kH and \mathcal{F} large enough, Fig. 47d). Lower values of kH (stronger interaction) or of Froude number (stronger confinement) both lead to an increase in wind disturbances, both in terms of orientation and velocity. Below a crossover value $kH \simeq 0.3$, wind disturbance is less sensitive to the \mathcal{F} -value. This is probably due to enhanced non-linear effects linked to flow modulation by the obstacle

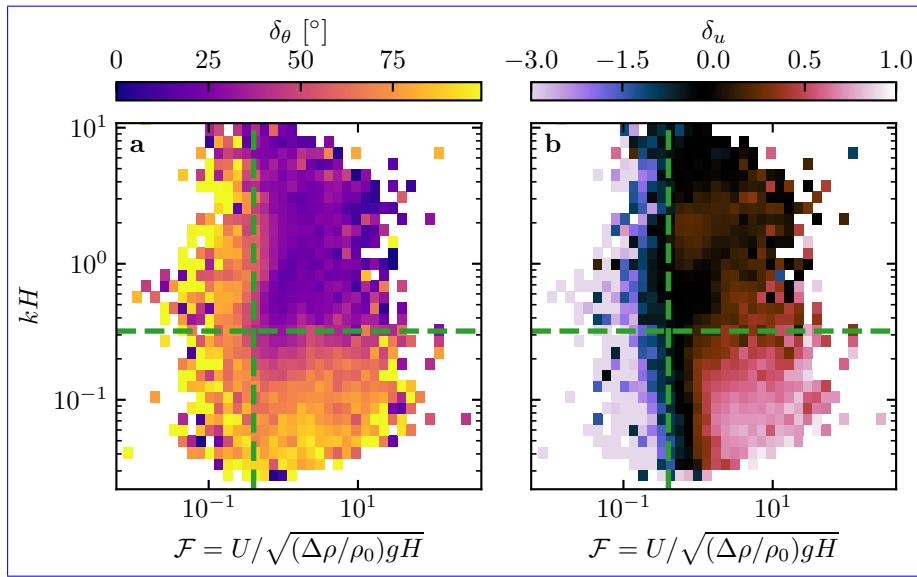


Fig. 8 Regime diagrams of the wind deviation δ_θ (**a**) and relative attenuation/amplification δ_u (**b**) in the space (\mathcal{F}, kH) , containing the data from both the North Sand Sea and South Sand Sea stations. **Green** The green dashed lines empirically delimit the different regimes. The point density in each bin of the diagrams is shown in Online Resource Fig. S14–S10 – 95% of the data occur in the range $-1 < \delta u < 1$. Similar regime diagrams in the spaces (\mathcal{F}_I, kH) and $(\mathcal{F}_I, \mathcal{F})$ are shown in Online Resource Fig. S12–S11.

when confinement is strong (e.g. wakes and flow recirculations). The Froude number also controls a transition from damped to amplified wind velocities in the interdune, with a crossover around $\mathcal{F} \approx 0.4$ (Fig. 58b). Such an amplification is rather unexpected. Checking the occurrence of the corresponding data, it appears that these amplifications are associated with the southerly sea breeze, and occur dominantly during the October-March period, when the other easterly wind is not present (Online Resource Fig. S12a–b). Furthermore, they occur less frequently during the afternoon, and more frequently at the end of the day (Online Resource Fig. S12c). This effect may be linked to a change in the flow behaviour in the lee side of the obstacles but further measurements are needed in order to assess the different possibilities (Baines 1995; Vosper 2004).

It is important to discuss the sensitivity of the results with respect to the choice of the hydrodynamic roughnesses (see Online Resource section 4). In fact, the only quantities dependent on this choice are those which involve the amplitude of the velocities: wind shear velocities, Froude number \mathcal{F} and relative velocity modulation δ_u . Those associated with wind direction are independent of this choice. Considering the possible range of realistic roughnesses values, the uncertainty on velocities estimated using the law of the wall is at most 30%. A similar maximum uncertainty applies to the Froude number. This uncertainty also propagates to δ_u , for which Figure S14 shows that the

choice of roughness has little influence of its temporal variations, even if it can induce a global increase or decrease of its values. As such, the choice of z_0 will not qualitatively affect the overall aspect of the regime diagram presented in Figure 8b. It may only change the crossover value of δ_u at which the transition between regimes is observed (dashed green lines in that figure). Our conclusions are thus robust with respect to the somewhat arbitrary choice of the hydrodynamic roughnesses in the use of the data from ERA5-Land reanalysis.

349 4 Discussion and conclusion

The feedback of the giant dunes on the wind flow has important implications for smaller scales bedforms. As illustrated in Fig. 69, small linear dunes (~ 50 m wide) are often present in the 1–2 km interdune between giant linear dunes in the Namib Sand Sea (Livingstone et al. 2010). These smaller dunes do not exhibit the same orientation as the large ones, and are sometimes named ‘crossing dunes’ (Chandler et al. 2022). Whilst differences between large and small scale dune patterns are observed ubiquitously, they are largely usually attributed to the presence of two different dune growth mechanisms, leading to two different dune patterns (orientations and/or morphologies) for the same wind regime (Courrech du Pont et al. 2014; Runyon et al. 2017; Lü et al. 2017; Song et al. 2019; Gadal et al. 2020; Hu et al. 2021). Here, however, our arguments enable the development of differing orientations for the small and giant linear dunes whilst also imposing governed by the same dune growth mechanism (elongating mode). Figure 69 shows how the orientations for the small and giant dunes can be derived from the locally measured and regionally predicted winds respectively (red arrows in Fig. 69). These predictions require a specification for to specify the threshold of eolian aeolian sand transport. Importantly, its value expressed as (a shear velocity $u_{th} \approx 0.15 \text{ ms}^{-1}$ estimated $u_{th} \approx 0.15 \text{ m s}^{-1}$) is reached in the deflected wind regime already winds which are deflected – recall that the larger winds are not deflected, see Fig. 6. The feedback of the giant dunes on the wind described in this study thus provides a potential explanation for the existence of these small linear dunes elongating across the interdune, a dynamic which has remained unresolved to date. These crossing dunes could provide additional constraints for the inference of local winds from bedforms, similarly to that currently performed on Mars using ripple orientations (Liu and Zimbelman 2015; Hood et al. 2021). Further work is needed to investigate these processes in more detail, including measurements of sediment transport and flow on the top of dunes.

This study presents the evidence that wind flow patterns around giant dunes are influenced by the atmospheric boundary layer, particularly during nocturnal conditions. It leaves However, we do not address here the question of the limitation of their pattern coarsening, and leave open the debate as to whether the size of giant dunes is limited controlled by the depth of this layer (Andreotti et al. 2009), in contrast to an unconstrained dune growth;

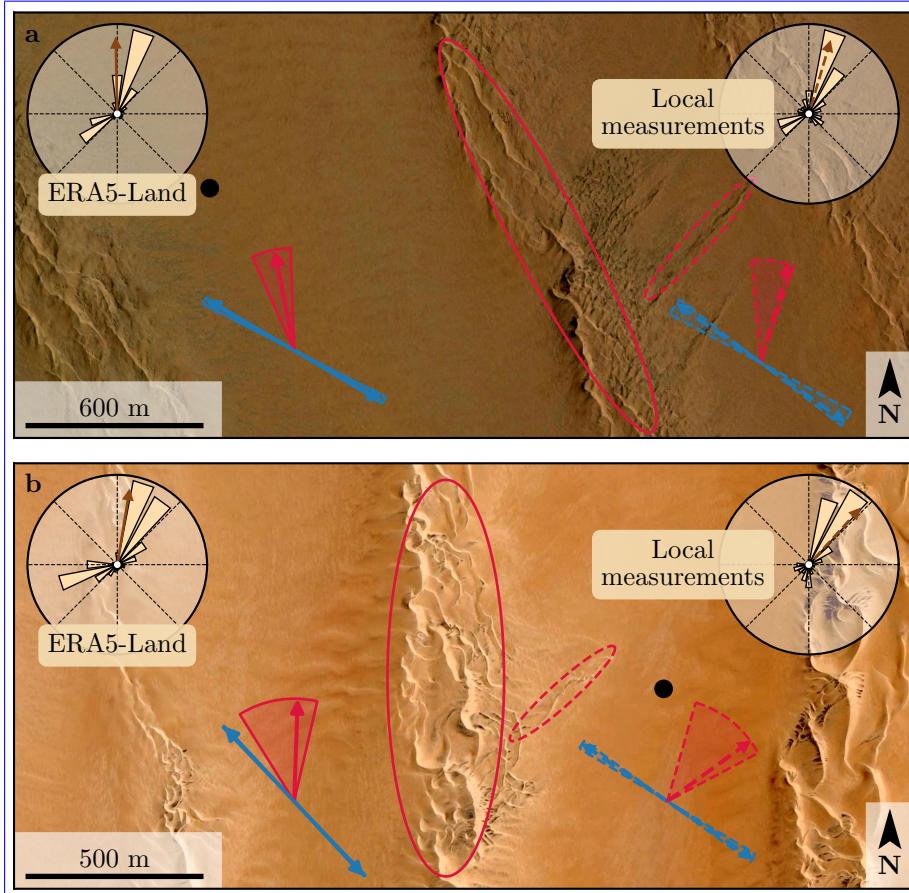


Fig. 9 Implications for smaller scale patterns in (a) the South Sand Sea and (b) North Sand Sea. The ellipses indicate the different types of elongating dunes, at large (plain line) and small (dashed line) scales. Dune orientation are predicted using the model of Courrech du Pont et al. (2014) from the sand flux angular distributions, shown here (roses and) along with the resultant transport direction (brown arrow) for typical values (grain size 180 μm , flux-up ratio of 1.6). Code for corresponding arrows: use of ERA5-Land data (plain line), use of local measurements (dashed line), prediction for the bed instability growth mechanism (blue), prediction for the elongation growth mechanism (red). Wedges show the uncertainty on the orientation calculation when parameters are varied. The black dots indicate the location of the meteorological stations with respect to in the dunes interdune. See Appendix 2 for additional details.

384 sediment supply limited and ever-slower with size (Gunn et al. 2021a) growth
 385 with size (Werner and Kocurek 1999; Gunn et al. 2021a). More field evidence
 386 is definitively needed from additional dune fields, but this mechanism would
 387 then allow for the inference of the ABL depth from giant bedform wavelengths
 388 where measurements are not feasible or available, such as Titan (Lorenz et al.
 389 2010).

390 To conclude on conditions under which the ERA5-Land reanalysis data can
391 reliably be used to study dune morphodynamics, we summarise the compari-
392 son of local (direct measurements) and regional (climate reanalysis) wind data
393 [as follows](#). In flat areas, the agreement between the two confirms the ability of
394 the ERA5-Land climate reanalysis to predict the wind regime down to scales
395 ~ 10 km, i.e the model grid. When smaller scale topographies are present
396 (giant dunes in our case), locally measured winds can significantly differ from
397 the regionally predicted ones. This is the case when the disturbances induced
398 by the dunes interact with the lower part of the ABL vertical structure, which
399 presents circadian variations. During the day, when the capping layer is typi-
400 cally high, this interaction is small, and the ERA5-Land predictions are also
401 quantitatively consistent with the local data. During the night, however, the
402 presence of a shallow atmospheric boundary layer induces a strong confine-
403 ment of the flow, and is associated with large wind deflection by the dunes.
404 Importantly, we find that this effect can be counterbalanced for large wind
405 velocities, which are capable of deforming the capping layer, thus decreasing
406 the influence of the confinement.

407 The theoretical computation of the wind disturbances induced by sinu-
408 soidal ridges under flow confinement has been performed in the linear limit
409 (Andreotti et al. 2009, 2012), i.e. when the aspect [ratio](#)-[ratio](#) of these ridges
410 is small ($k\xi_0 \ll 1$). These models are able to qualitatively reproduce the ob-
411 served wind deflection (Appendix 1, Online Resource Figs. ?? and ??S11 and
412 S13), and thus provide the physical support for the interpretation we propose
413 here based on hydrodynamic regimes. However, these models cannot quanti-
414 tatively predict the magnitude of [these](#)-[our](#) observations, probably due to the
415 presence of expected non-linearities in high confinement situations linked to
416 strong flow modulations. Besides, these linear calculations only predict wind
417 attenuation in the interdune, in contrast with the observed enhanced veloci-
418 ties associated with particular evening winds from the South during the period
419 October-March (Online Resource Fig. ??S12). Some other models predict dif-
420 ferent spatial flow structures in response to a modulated topography, such as
421 lee waves and rotors (Baines 1995; Vosper 2004). However, our measurements
422 are located at a single point in the interdune, [so we are](#) and [we are thus](#) unable
423 to explore these types of responses. Data at different places along and across
424 the ridges are needed to investigate and possibly map such flow structures,
425 and for further comparisons with the models.

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430 All data used in this study can be found in Gadale et al. (2022). Note that it contains
431 modified Copernicus Climate Change Service Information (2021). Neither the European
432 Commission nor ECMWF is responsible for any use that may be made of the Copernicus
433 Information or Data it contains. [Fully documented codes used](#) [Documented codes used in this](#)
434 [study](#) to analyse this [study](#)-[data](#) are available at <https://github.com/Cgadal/GiantDunes>
435 (will be made public upon acceptance of this manuscript for publication).

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Appendix 1: Linear theory of wind response to topographic perturbation

Following the work of Fourrière et al. (2010), Andreotti et al. (2012) and Andreotti et al. (2009), we briefly describe in this appendix the framework for the linear response of a turbulent flow to a topographic perturbation of small aspect ratio. As a general bed elevation can be decomposed into Fourier modes, we focus here on a sinusoidal topography:

$$\xi = \xi_0 \cos [k (\cos(\alpha)y - \sin(\alpha)x)], \quad (3)$$

which is also a good approximation for the giant dunes observed in the North Sand Sea and South Sand Sea Station (Fig. 42 and Online Resource Fig. S5S4). Here, x and y are the streamwise and spanwise coordinates, $k = 2\pi/\lambda$ the wavenumber of the sinusoidal perturbation, α its crest orientation with respect to the x -direction (anticlockwise) and ξ_0 its amplitude. The two components of the basal shear stress $\tau = \rho_0 u_* \mathbf{u}_*$, constant in the flat bottom reference case, can then be generically written as:

$$\tau_x = \tau_0 \left(1 + k \xi_0 \sqrt{\mathcal{A}_x^2 + \mathcal{B}_x^2} \cos [k (\cos(\alpha)y - \sin(\alpha)x) + \phi_x] \right), \quad (4)$$

$$\tau_y = \tau_0 k \xi_0 \sqrt{\mathcal{A}_y^2 + \mathcal{B}_y^2} \cos [k (\cos(\alpha)y - \sin(\alpha)x) + \phi_y], \quad (5)$$

where τ_0 is the reference basal shear stress on a flat bed. We have defined the phase $\phi_{x,y} = \tan^{-1}(\mathcal{B}_{x,y}/\mathcal{A}_{x,y})$ from the in-phase and in-quadrature hydrodynamical coefficients $\mathcal{A}_{x,y}$ and $\mathcal{B}_{x,y}$. They are functions of k and of the flow conditions, i.e the bottom roughness, the vertical flow structure and the incident flow direction, and the theoretical framework developed in the above cited papers proposes methods to compute them in the linear regime.

⁴⁷³ Following Andreotti et al. (2012), the effect of the incident wind direction
⁴⁷⁴ can be approximated by the following expressions:

$$\mathcal{A}_x = \mathcal{A}_0 \sin^2 \alpha, \quad (6)$$

$$\mathcal{B}_x = \mathcal{B}_0 \sin^2 \alpha, \quad (7)$$

$$\mathcal{A}_y = -\frac{1}{2} \mathcal{A}_0 \cos \alpha \sin \alpha, \quad (8)$$

$$\mathcal{B}_y = -\frac{1}{2} \mathcal{B}_0 \cos \alpha \sin \alpha, \quad (9)$$

⁴⁷⁵ where \mathcal{A}_0 and \mathcal{B}_0 are now two coefficients independent of the dune orientation
⁴⁷⁶ α , corresponding to the transverse case ($\alpha = 90^\circ$). ~~For~~^{In the case of} a fully
⁴⁷⁷ turbulent boundary layer capped by a stratified atmosphere, these coefficients
⁴⁷⁸ depend on kH , kz_0 , \mathcal{F} and \mathcal{F}_I (Andreotti et al. 2009). ~~In this study~~^{For their}
⁴⁷⁹ ~~computation~~, we assume ~~here~~ a constant hydrodynamic roughness $z_0 \simeq 1$ mm
⁴⁸⁰ (Online Resource section 1). For the considered giant dunes, this leads to
⁴⁸¹ $kz_0 \simeq 2 \cdot 10^{-6}$, as their wavelength is $\lambda \simeq 2.4$ km (or $k \simeq 2 \cdot 10^{-3} \text{ m}^{-1}$). Values
⁴⁸² of z_0 extracted from field data indeed typically fall between 0.1 mm and 10
⁴⁸³ mm (Sherman and Farrell 2008; Field and Pelletier 2018). Importantly, \mathcal{A}_0
⁴⁸⁴ and \mathcal{B}_0 do not vary much in the corresponding range of kz_0 (Fourrière et al.
⁴⁸⁵ 2010), and the results presented here are robust with respect to this choice.

⁴⁸⁶ With capping layer height and Froude numbers computed from the ERA5-
⁴⁸⁷ Land time series, the corresponding \mathcal{A}_0 and \mathcal{B}_0 can be deduced, as displayed
⁴⁸⁸ in Online Resource Fig. ??S13. Interestingly, it shows similar regimes as in
⁴⁸⁹ the diagrams of Fig. ??S13 and Online Resource Fig. ??S11a,b, supporting the
⁴⁹⁰ underlying physics. However, the agreement is qualitative only. Further, the
⁴⁹¹ linearity assumption of the theoretical framework requires $(|\tau| - \tau_0) / \tau_0 \ll 1$,
⁴⁹² which translates into $k\xi \sqrt{\mathcal{A}_0^2 + \mathcal{B}_0^2} \ll 1$. In our case, the giant dune morphol-
⁴⁹³ ogy gives $k\xi_0 \simeq 0.1$, which means that one quits the regime of validity of the
⁴⁹⁴ linear theory when the coefficient modulus $\sqrt{\mathcal{A}_0^2 + \mathcal{B}_0^2}$ becomes larger than a
⁴⁹⁵ few units. In accordance with the theoretical expectations, these coefficients
⁴⁹⁶ present values on the order of unity ($\mathcal{A}_0 \simeq 3$ and $\mathcal{B}_0 \simeq 1$) in unconfined situations
⁴⁹⁷ (Claudin et al. 2013; Lü et al. 2021). In contrast and as illustrated in
⁴⁹⁸ Online Resource Fig. ??S13a,b, larger values are predicted in case of strong
⁴⁹⁹ confinement, which does not allow us to proceed to further quantitative com-
⁵⁰⁰ parison with the data.

⁵⁰¹ Finally, the linear model is also able to reproduce the enhancement of
⁵⁰² the flow deflection over the sinusoidal ridges when $\sqrt{\mathcal{A}_0^2 + \mathcal{B}_0^2}$ is increased
⁵⁰³ (Online Resource Fig. ??S13). Here, using $k\xi_0 \simeq 0.1$ to be representative of
⁵⁰⁴ the amplitude of the giant dunes at the North Sand Sea station, the coefficient
⁵⁰⁵ modulus is bounded to 10.

506 **Appendix 2: Sediment transport and dune morphodynamics**

507 We summarise in this appendix the sediment transport and dune morphodynamics
 508 theoretical framework leading to the prediction of sand fluxes and dune
 509 orientations from wind data.

510 *Sediment transport* — The prediction of sand fluxes from wind data has been
 511 a long standing issue in aeolian geomorphological studies (Fryberger and Dean
 512 1979; Pearce and Walker 2005; Sherman and Li 2012; Shen et al. 2019). Based
 513 on laboratory studies in wind tunnels (Rasmussen et al. 1996; Iversen and
 514 Rasmussen 1999; Creyssels et al. 2009; Ho et al. 2011), as well as physical
 515 considerations (Ungar and Haff 1987; Andreotti 2004; Durán et al. 2011; Pähzt
 516 and Durán 2020), it has been shown that the steady saturated saltation flux
 517 over a flat sand bed depends linearly on the shear stress:

$$\frac{q_{\text{sat}}}{Q} = \Omega \sqrt{\Theta_{\text{th}}} (\Theta - \Theta_{\text{th}}), \quad (10)$$

518 where Ω is a proportionality constant, $Q = d\sqrt{(\rho_s - \rho_0)gd/\rho_0}$ is a character-
 519 istic flux, $\Theta = \rho_0 u_*^2 / (\rho_s - \rho_0)gd$ the Shields number, and Θ_{th} its threshold
 520 value below which saltation vanishes. $\rho_s = 2.6 \text{ g cm}^{-3}$ and $d = 180 \mu\text{m}$ are
 521 the grain density and diameter, and g is the gravitational acceleration. The
 522 shear velocity, and consequently the Shields number as well as the sediment
 523 flux, are time dependent.

524 Recently, Pähzt and Durán (2020) suggested an additional quadratic term
 525 in Shields to account for grain-grain interactions within the transport layer at
 526 strong wind velocities:

$$\frac{q_{\text{sat}, t}}{Q} = \frac{2\sqrt{\Theta_{\text{th}}}}{\kappa\mu} (\Theta - \Theta_{\text{th}}) \left(1 + \frac{C_M}{\mu} [\Theta - \Theta_{\text{th}}] \right), \quad (11)$$

527 where $\kappa = 0.4$ is the von Kármán constant, $C_M \simeq 1.7$ a constant and $\mu \simeq 0.6$ is
 528 a friction coefficient, taken to be the avalanche slope of the granular material.
 529 The fit of this law to the experimental data of Creyssels et al. (2009) and
 530 Ho et al. (2011) gives $\Theta_{\text{th}} = 0.0035$. The fit of Eq. 10 on these same data
 531 similarly gives $\Omega \simeq 8$ and $\Theta_{\text{th}} = 0.005$. The sand flux angular distributions
 532 and the dune orientations in Fig. 6–9 are calculated using this law (11). We
 533 have checked that using the ordinary linear relationship (10) instead does not
 534 change the predicted dune orientations by more than a few degrees.

535 *Dune orientations* — Dune orientations are predicted with the dimensional
 536 model of Courrech du Pont et al. (2014), from the sand flux time series com-
 537 puted with the above transport law. Two orientations are possible depending
 538 on the mechanism dominating the dune growth: elongation or bed instabil-
 539 ity. The orientation α corresponding the bed instability is then the one that
 540 maximises the following growth rate (Rubin and Hunter 1987):

$$\sigma \propto \frac{1}{H_d W_d T} \int_0^T q_{\text{crest}} |\sin(\theta - \alpha)| dt, \quad (12)$$

541 where θ is the wind orientation measured with respect to the same reference
 542 as α , and H_d and W_d are dimensional constants respectively representing the
 543 dune height and width. The integral runs over a time T , which must be repre-
 544 sentative of the characteristic period of the wind regime. The flux at the crest
 545 is expressed as:

$$q_{\text{crest}} = q_{\text{sat}} [1 + \gamma |\sin(\theta - \alpha)|], \quad (13)$$

546 where the flux-up ratio γ has been calibrated to 1.6 using field studies, under-
 547 water laboratory experiments and numerical simulations. Predictions of the
 548 linear analysis of Gadal et al. (2019) and Delorme et al. (2020) give similar
 549 results.

550 Similarly, the dune orientation corresponding to the elongation mechanism
 551 is the one that verifies:

$$\tan(\alpha) = \frac{\langle q_{\text{crest}}(\alpha) \mathbf{e}_\theta \rangle \cdot \mathbf{e}_{WE}}{\langle q_{\text{crest}}(\alpha) \mathbf{e}_\theta \rangle \cdot \mathbf{e}_{SN}}, \quad (14)$$

552 where $\langle \cdot \rangle$ denotes a vectorial time average. The unitary vectors \mathbf{e}_{WE} , \mathbf{e}_{SN} and
 553 \mathbf{e}_θ are in the West-East, South-North and wind directions, respectively.

554 The resulting computed dune orientations, blue and red arrows in Fig. 69,
 555 then depend on a certain number of parameters (grain properties, flux-up ratio,
 556 etc.), for which we take typical values for aeolian sandy deserts. Due to the lack
 557 of measurements in the studied places, some uncertainties can be expected. We
 558 therefore run a sensitivity test by calculating the dune orientations for grain
 559 diameters ranging from 100 μm to 400 μm and for a speed-up ratio between
 560 0.1 and 10 (wedges in Fig. 69).

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907 **Local wind regime induced by giant linear dunes**
 908 — Supplementary Material —

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915 **1. Shear velocity and calibration of the hydrodynamical roughness**

916 As the regionally predicted and locally measured velocities are available at
 917 different heights, we can not compare them directly. We therefore convert all
 918 velocities into shear velocities u_* , characteristic ~~of the turbulent~~ ~~the turbulent~~
 919 logarithmic velocity profile (Spalding 1961; Stull 1988):

$$\frac{u(z)}{u_*} = \frac{1}{\kappa} \ln \left(\frac{z}{z_0} \right), \quad (15)$$

920 where z is the vertical coordinate, $\kappa = 0.4$ the von Kármán constant and z_0 the
 921 hydrodynamic roughness. Note that, strickly speaking, this logarithmic profile
 922 is valid for a neutrally stratified ABL only. Vertical density gradients occuring
 923 in other conditions may thus induce large discrepancies (Monin and Obukhov 1954; Garratt 1994; Dyer 1974)
 924 . However, as our wind measurements are in the flow region close enough to
 925 the surface, where these effects are negligible, this logarithmic wind profile
 926 remains a fairly good approximation in all conditions (Gunn et al. 2021b).
 927 Several measurements of hydrodynamic roughnesses are available (Raupach
 928 1992; Bauer et al. 1992; Brown et al. 2008; Nield et al. 2014). In the absence
 929 of sediment transport, it is governed by the geometric features of the bed
 930 (Flack and Schultz 2010; Pelletier and Field 2016). When aeolian saltation
 931 occurs, it is rather controlled by the altitude of Bagnold's focal point (Durán
 932 et al. 2011; Valance et al. 2015), which depends on the wind velocity and grain
 933 properties (Sherman and Farrell 2008; Zhang et al. 2016; Field and Pelletier
 934 2018). Whether associated with geometric features or with sediment transport,
 935 its typical order of magnitude is the millimetre scale on sandy surfaces.

936 We do not have precise velocity vertical profiles to be able to deduce an
 937 accurate value of z_0 in the various environments of the meteorological stations
 938 (vegetated, arid, sandy). Our approach is to rather select the hydrodynamic
 939 roughness which allows for the best possible matching between the regionally
 940 predicted and locally measured winds, i.e. minimising the relative difference δ
 941 between the wind vectors of both datasets:

$$\delta = \frac{\sqrt{\langle \| \mathbf{u}_{*,\text{era}} - \mathbf{u}_{*,\text{station}} \|^2 \rangle}}{\sqrt{\langle \| \mathbf{u}_{*,\text{era}} \|^2 \rangle \langle \| \mathbf{u}_{*,\text{station}} \|^2 \rangle}}, \quad (16)$$

where $\langle \cdot \rangle$ denotes time average. This parameter is computed for values of z_0 in ERA5-Land analysis ranging from 10^{-5} m to 10^{-2} m for the four different stations. Note that for the North Sand Sea and South Sand Sea stations, where the giant dunes feedback presumably affect the wind, we take into account the non-deflected winds only in the calculation of δ (with a 15° tolerance).

As shown in Online Resource Fig. S4S3, the minimum values of δ in the space $(z_0^{\text{ERA5Land}}, z_0^{\text{local}})$ form a line. We thus set the roughness in the ERA5-Land analysis to the typical value $z_0 = 10^{-3}$ m, and deduce the corresponding ones for the local stations. It leads to 2.7, 0.8, 0.1 and 0.5 mm for the Etosha West, North Sand Sea, Huab and South Sand Sea stations, respectively. Importantly, this approach somewhat impacts the calculation of the shear velocities, but not that of the wind directions. As such, most of our conclusions are independent of such a choice. However, it may affect the magnitude of the wind velocity attenuation/amplification in flow confinement situations.

2. Computation of the ABL characteristics

The estimation of the non-dimensional numbers associated with the ABL requires the computation of representative meteorological quantities. In arid areas, the vertical structure of the atmosphere can be approximated by a well mixed convective boundary layer of height H , topped by the stratified free atmosphere (Stull 1988; Shao 2008). In this context, one usually introduces the virtual potential temperature T_{vp} , which is a constant T_0 inside the boundary layer, and increases linearly in the FA (Online Resource Fig. S12S8a):

$$T_{\text{vp}}(z) = \begin{cases} T_0 & \text{for } z \leq H, \\ T_0 \left(1 + \frac{\Delta T_{\text{vp}}}{T_0} + \frac{N^2}{g}(z - H) \right) & \text{for } z \geq H, \end{cases} \quad (17)$$

where ΔT_{vp} is the temperature discontinuity at the capping layer and $N = \sqrt{g\partial_z T_{\text{vp}}/T_0}$ is the Brunt-Väisälä frequency, characteristic of the stratification. Note that, under the usual Boussinesq approximation, temperature and air density variations are simply related by $\delta T_{\text{vp}}/T_0 \simeq -\delta\rho/\rho_0$ (see Online Resource of Andreotti et al. (2009)), so that N can equivalently be defined from the density gradient as next to Eq. 4(1).

The ERA5 dataset provides vertical profiles of the geopotential ϕ , the actual temperature T and the specific humidity η at given pressure levels P . The vertical coordinate is then calculated as:

$$z = \frac{\phi R_t}{g R_t - \phi}, \quad (18)$$

where $R_t = 6371229$ m is the reference Earth radius and $g = 9.81$ m s $^{-2}$ is the gravitational acceleration. One also computes the virtual potential temperature as:

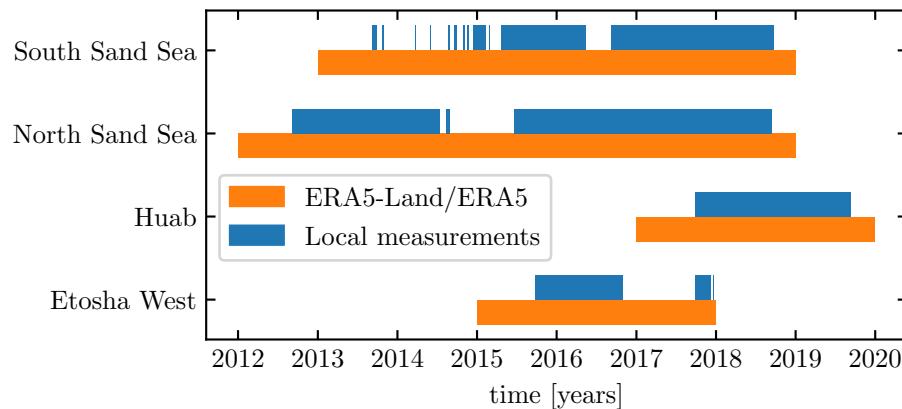
$$T_{\text{vp}} = T \left[1 + \left(\frac{M_d}{M_w} - 1 \right) \eta \right] \left(\frac{P_0}{P} \right)^{R/C_p}, \quad (19)$$

where $P_0 = 10^5$ Pa is the standard pressure, $R = 8.31$ J/K is the ideal gas constant, $C_p \simeq 29.1$ J/K is the air molar heat capacity, and $M_w = 0.018$ kg/Mol and $M_d = 0.029$ kg/Mol are the molecular masses of water and dry air respectively. The specific humidity is related to the vapour pressure p_w as

$$\eta = \frac{\frac{M_w}{M_d} p_w}{p - \left(1 - \frac{M_w}{M_d}\right) p_w}. \quad (20)$$

The ERA5 dataset also provides an estimate of the ABL depth H , based on the behaviour of the Richardson vertical profile. This dimensionless number is defined as the ratio of buoyancy and flow shear terms, and can be expressed as $\text{Ri} = N^2 / (\partial_z u)^2$. It vanishes in the lower well-mixed layer where T_{vp} is constant, and increases in the stratified FA. Following the method and calibration of Vogelegang and Holtlag (1996); Seidel et al. (2012), the value $\text{Ri}(z) \simeq 0.25$ has been shown to be a good empirical criterion to give $z \simeq H$ within a precision varying from 50% for the shallower ABL (e.g. at night) to 20% for situations of stronger convection.

Examples of vertical profiles of the virtual potential temperature deduced from ERA5 are shown in Online Resource Fig. S12S8a. For each of them, an average temperature is computed below the ABL depth ($z < H$), and a linear function is fitted above, allowing us to extract the temperature jump ΔT_{vp} . Importantly, some profiles display a vertical structure that cannot be approximated by the simple form (17) used here (Online Resource Fig. S12S8b). In practice, we removed from the analysis all of those leading to the unphysical case $\Delta T_{\text{vp}} < 0$. We have noticed that these ‘ill-processed’ profiles dominantly occur in winter and are evenly spread across the hours of the day. Importantly, they represent $\simeq 12\%$ of the data only (Online Resource Fig. S12S8c,d), and we are thus confident that this data treatment does not affect our conclusions.



Gant chart representing the valid time steps for the two data sets, for all stations.

Fig. S1 Photographs of the meteorological stations. **a:** South Sand Sea station. **b–e:** North Sand Sea station. **d:** Huab station. **e:** Etosha West station.

Gant chart representing the valid time steps for the two data sets, for all stations.

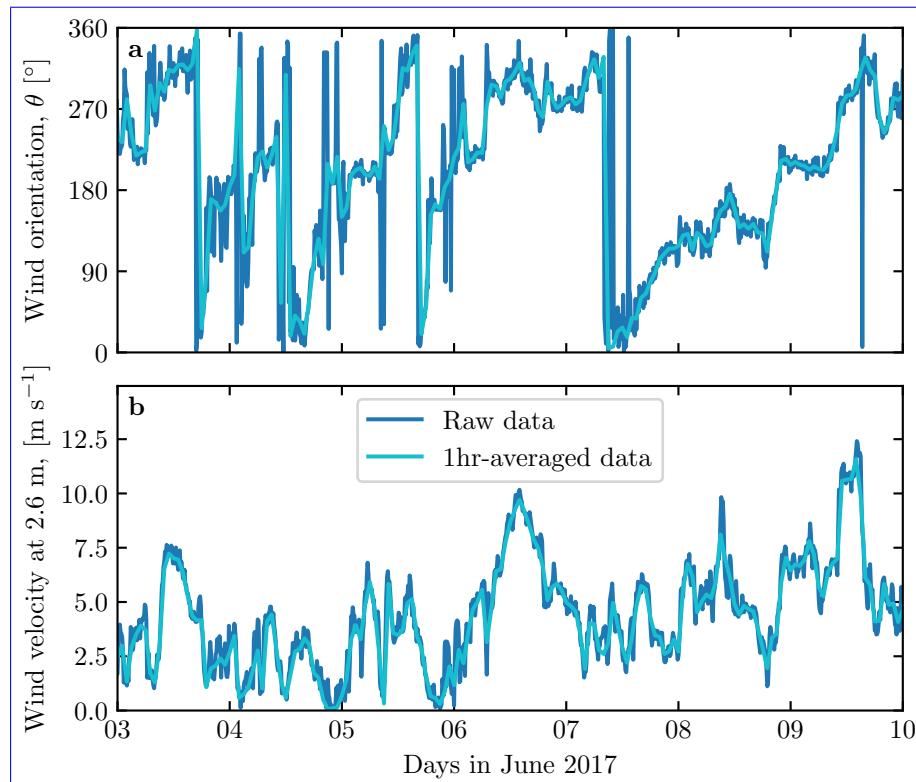


Fig. S2 Comparison between raw local and hourly-averaged wind measurements; **a:** comparison between raw (blue) and hourly-averaged (light blue) data for from South Sand Sea station. **a:** wind direction. **b:** wind velocity at height 2.6 m.

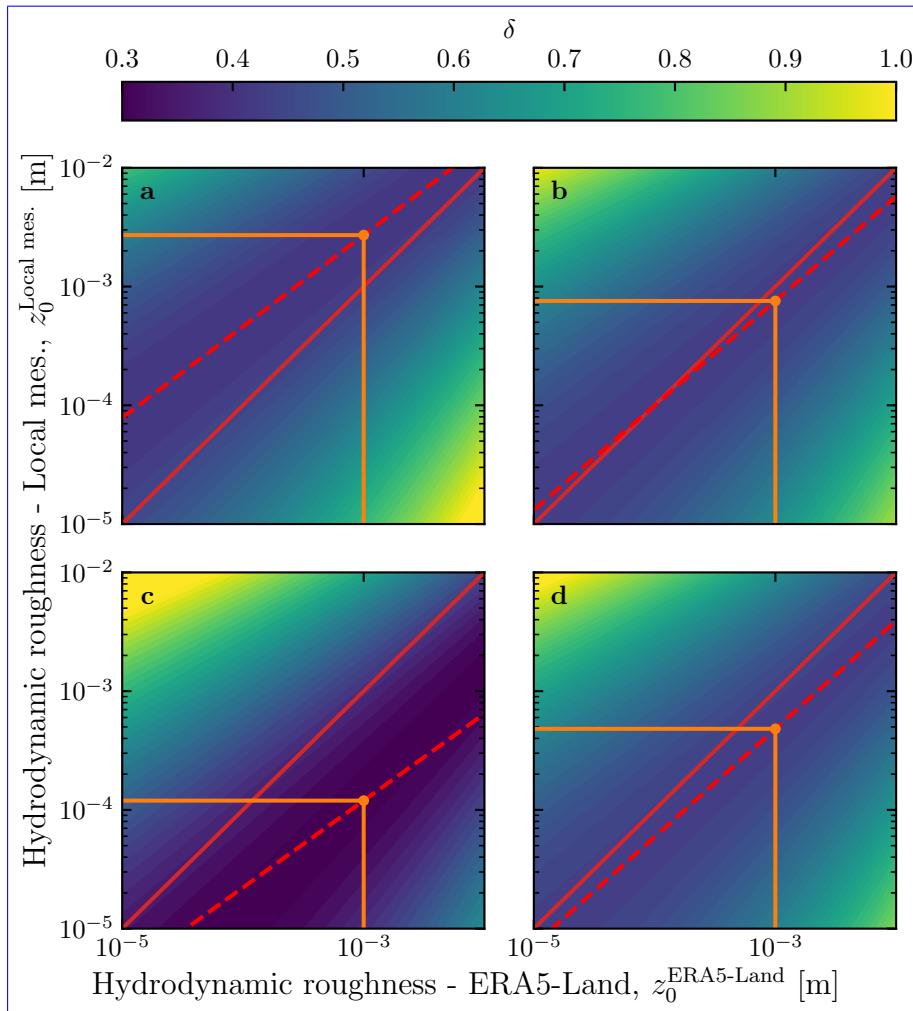


Fig. S3 Calibration of hydrodynamic roughness. The parameter δ (Eq. 16) quantifying the difference between local and predicted winds is shown as a function of the hydrodynamic roughnesses chosen for the ERA5-Land and for local winds, for the (a) Etosha West, (b) North Sand Sea, (c) Huab and (d) South Sand Sea stations. The red dashed and plain lines show the minima of δ and the identity line, respectively. The orange lines and dots highlight the chosen hydrodynamic roughnesses for the local winds deduced from setting $\underline{z_0^{\text{ERA5Land}}} = 1 \underline{z_0^{\text{ERA5Land}}} = 1$ mm.

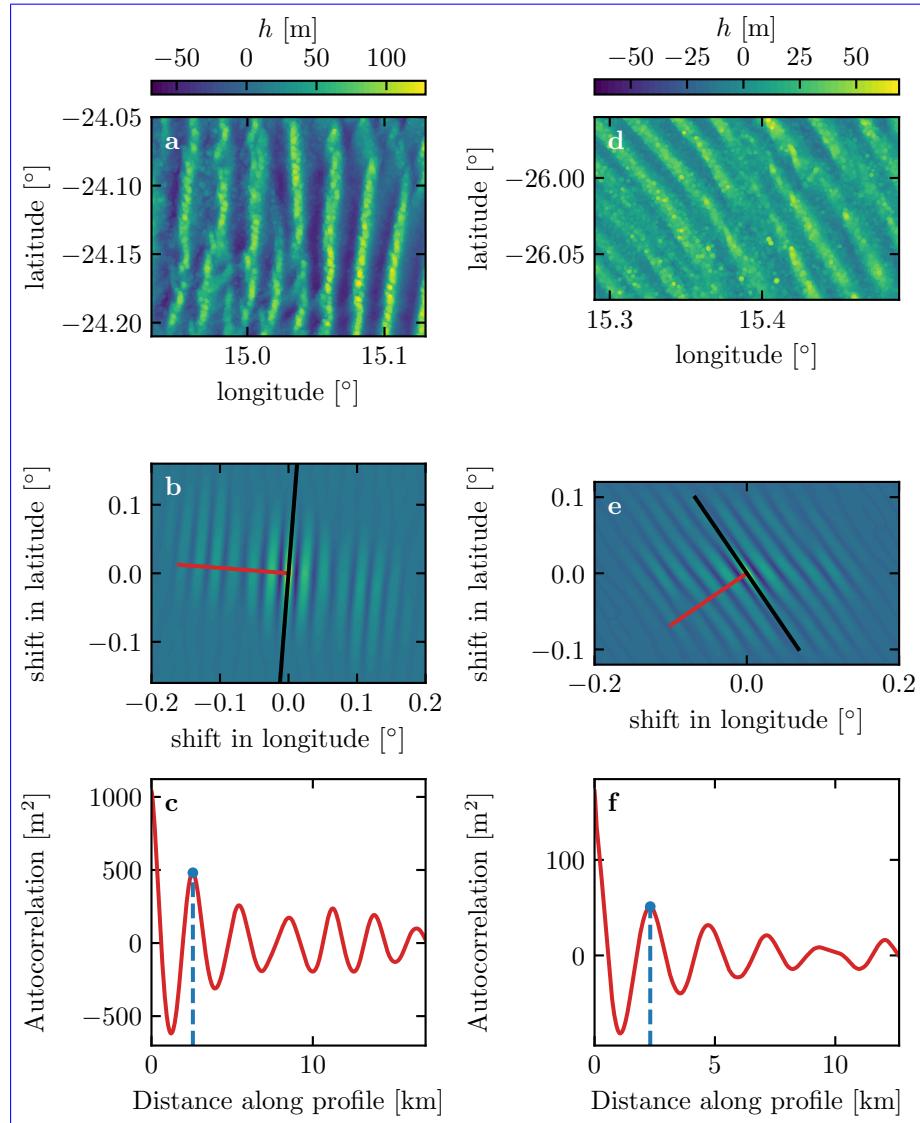


Fig. S4 Analysis of the DEMs of the North Sand Sea (left column – panels **a**, **b**, **c**) and South Sand Sea (right column – panels **d**, **e**, **f**) stations. **a–d**: Bed elevation detrended by a fitted second order polynomial base-line. **b–e**: Autocorrelation matrix shown in **color scale**. The black line shows the detected dune orientation, and the red line represents the autocorrelation profile along which the dune wavelength is calculated, displayed in **c–f**. The blue lines and dots show the first peak of the autocorrelation profile, whose abscissa gives the characteristic wavelength of the dune pattern.

Temporal comparison between the wind data coming from the ERA5-Land climate reanalysis (orange lines) and from the local measurements (blue lines). Coloured swathes indicate day (between 1000 UTC and 2200 UTC) and night (before 1000 UTC or after 2200 UTC). **a–b**: Huab station in summer. **b–c**: Huab station in winter. **d–e**: South Sand Sea station in summer. **f–g**: South Sand Sea station in winter.

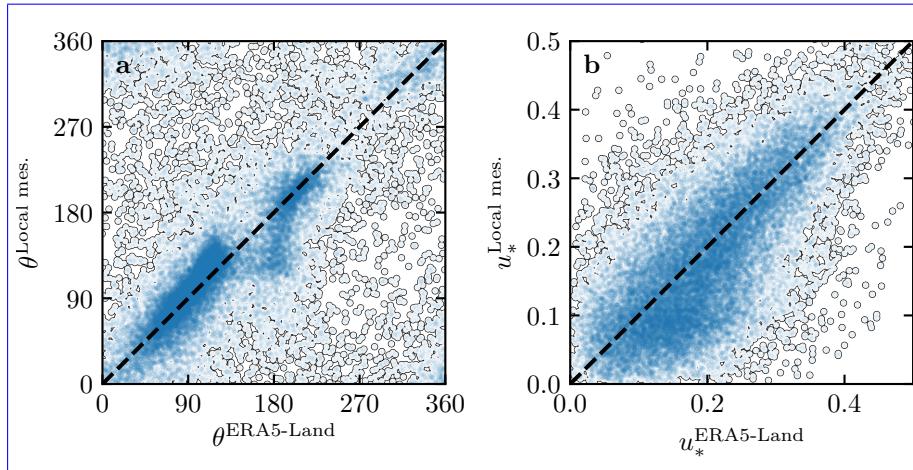


Fig. S5 Statistical comparison of the wind orientation (a) and velocity (b) between the ERA5-Land dataset and the local measurements for the Huab and Etosha West stations. Data point clustering around identity lines (dashed and black) provide evidence for agreement of the two sets.

Distributions of wind direction at the South Sand Sea Station for the ERA5-Land climate reanalysis (orange) and the local measurements (blue)—equivalent of Fig. 3. In each subplot, both distributions are plotted from the same time steps, selected with constraints on the wind direction (columns) and/or wind velocity (rows) in the ERA5-Land dataset. The gray vertical dashed lines indicate the dune orientation. The numbers at the top centre give the percentage of time steps selected in each sub-range, as well as the percentage of them corresponding to the day (between 1000 UTC and 2200 UTC). The purple frame highlights the regime (small wind velocities, nocturnal summer wind) during which the wind data from both datasets differ.

Distributions of wind direction at the Huab and Etosha West stations for the ERA5-Land climate reanalysis (orange) and the local measurements (blue). In each subplot, both distributions are plotted from the same time steps, selected with constraints on the wind velocity (rows) in the ERA5-Land dataset. The numbers at the top centre give the percentage of time steps selected in each sub-range, as well as the percentage of them corresponding to the day (between 1000 UTC and 2200 UTC). Compared to the North and South Namib stations (Fig. 3 and Fig. S8), histograms match for high and low velocities.

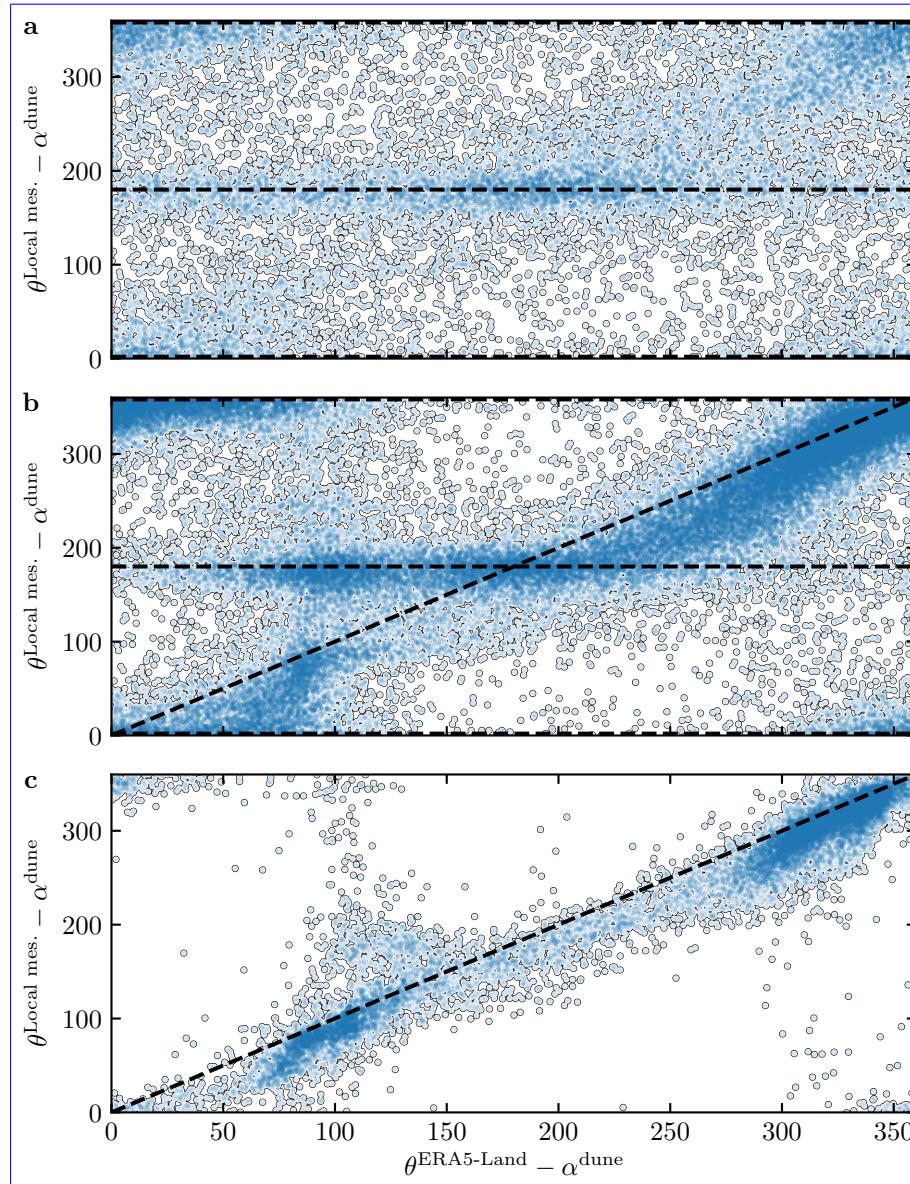


Fig. S6 Statistical comparison of the wind orientation between the ERA5-Land dataset and the local measurements for the South Sand Sea and North Sand Sea stations, for different velocity ranges. a: $u_{*,\text{ERA}} < 0.1 \text{ m s}^{-1}$ $u^{\text{ERA5-Land}} < 0.1 \text{ m s}^{-1}$. b: $0.1 < u_{*,\text{ERA}} \leq 0.25 \text{ m s}^{-1}$ $0.1 < u^{\text{ERA5-Land}} \leq 0.25 \text{ m s}^{-1}$. c: $u_{*,\text{ERA}} \geq 0.25 \text{ m s}^{-1}$ $u^{\text{ERA5-Land}} \geq 0.25 \text{ m s}^{-1}$. The measured dune orientations are subtracted to the wind orientation, which allows us to plot both stations on the same graph. Black dashed lines indicate locally measured orientations aligned with the dune crests (here 0° , 180° and 360° – panels a, b), as well as the identity lines (panels b, c).

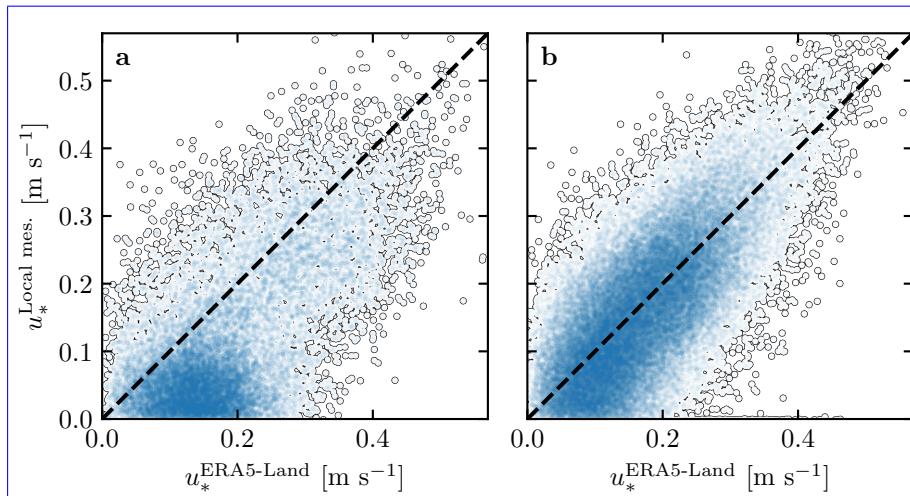


Fig. S7 Statistical comparison of the wind velocity between the ERA5-Land dataset and the local measurements for the South Sand Sea and North Sand Sea stations. **a:** Nocturnal summer easterly wind. **b:** Diurnal southerly wind. Black dashed lines are identity lines. The angle ranges used to select diurnal and nocturnal summer winds are the same as those in [FigFigs. 3–4](#) and [Online Resourcee FigFigs. S86 of the main article](#).

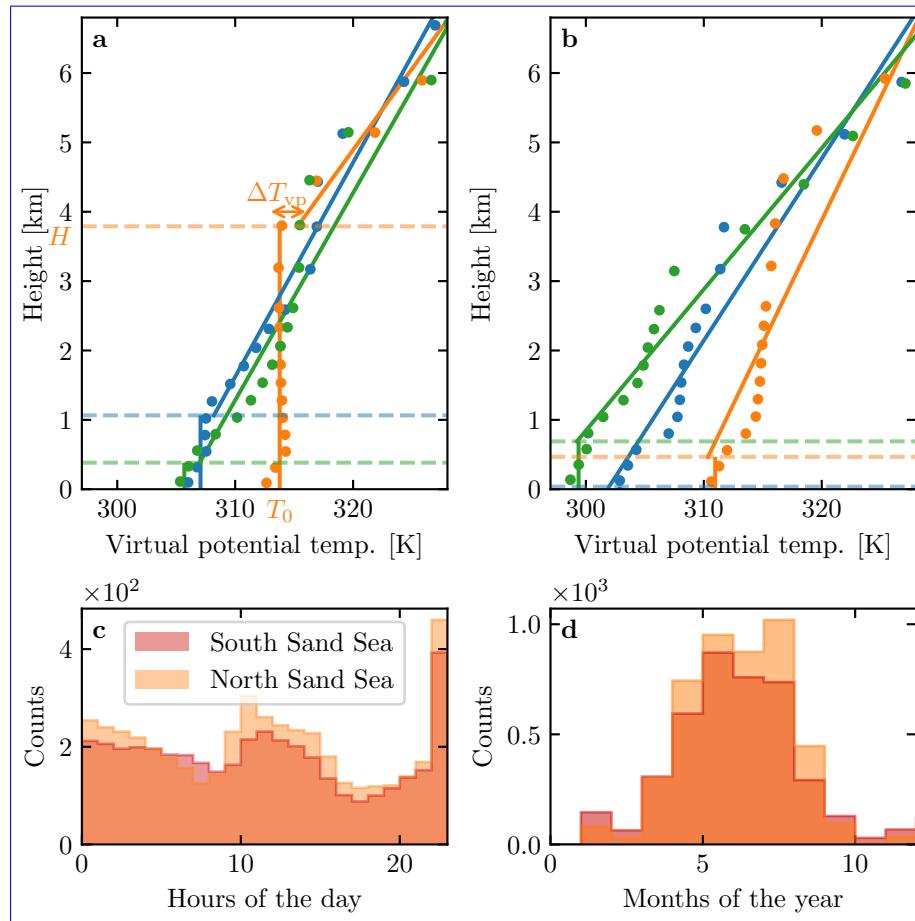


Fig. S8 a: Vertical profiles of the virtual potential temperature at three different times (blue: 29/11/2012 - 11:00–11:00 UTC, orange: 21/03/2017 - 12:00–12:00 UTC, green: 21/03/2017 - 20:00–20:00 UTC) at the South Sand Sea station. Dots: data from the ERA5 reanalysis. Dashed lines: boundary layer height given by the ERA5 reanalysis. Plain lines: vertical (ABL) and linear (FA) fits to estimate the quantities displayed in Online Resource Fig. S13S9. b: Examples of ill-processed vertical profiles at three different times (blue: 2/12/2013 - 23:00–23:00 UTC, orange: 20/03/2017 - 00:00–00:00 UTC, green: 14/07/2017 - 14:00–14:00 UTC) at the South Sand Sea station. c: Hourly distribution of ill-processed vertical profiles dat South (orange) and North (light orange) Sand Sea station: Monthly distribution of ill-processed vertical profiles hourly (c) and monthly (d) counts.

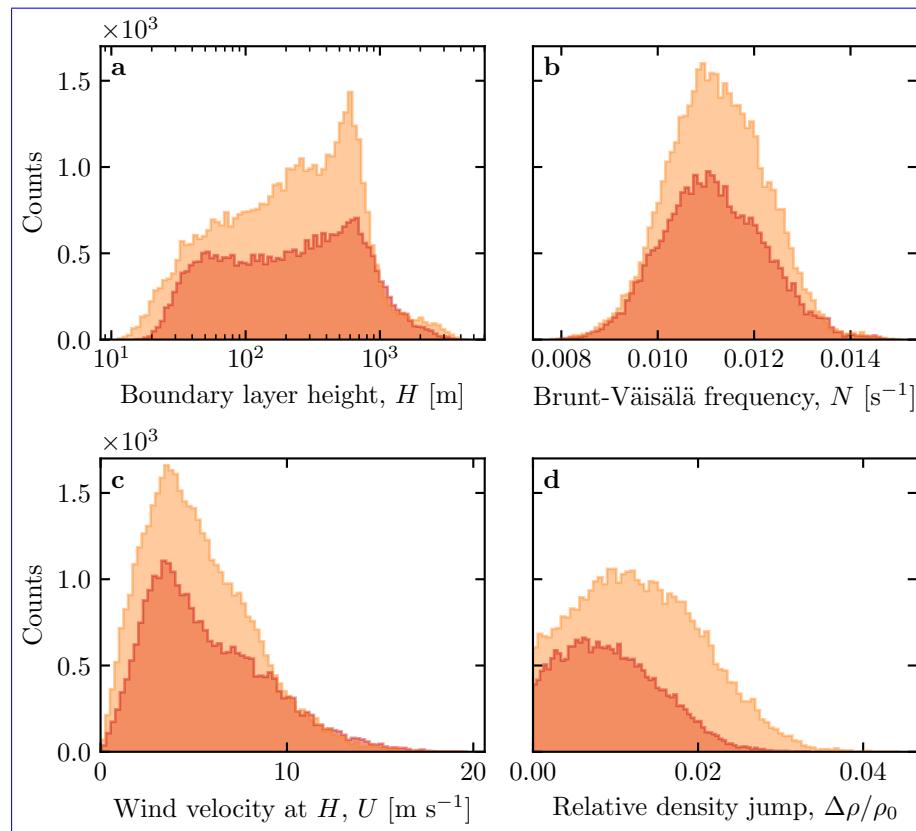


Fig. S9 Distributions of the meteorological parameters resulting from the processing of the ERA5-Land data for the South Sand Sea (**blueorange**) and the North Sand Sea (**light orange**) stations.

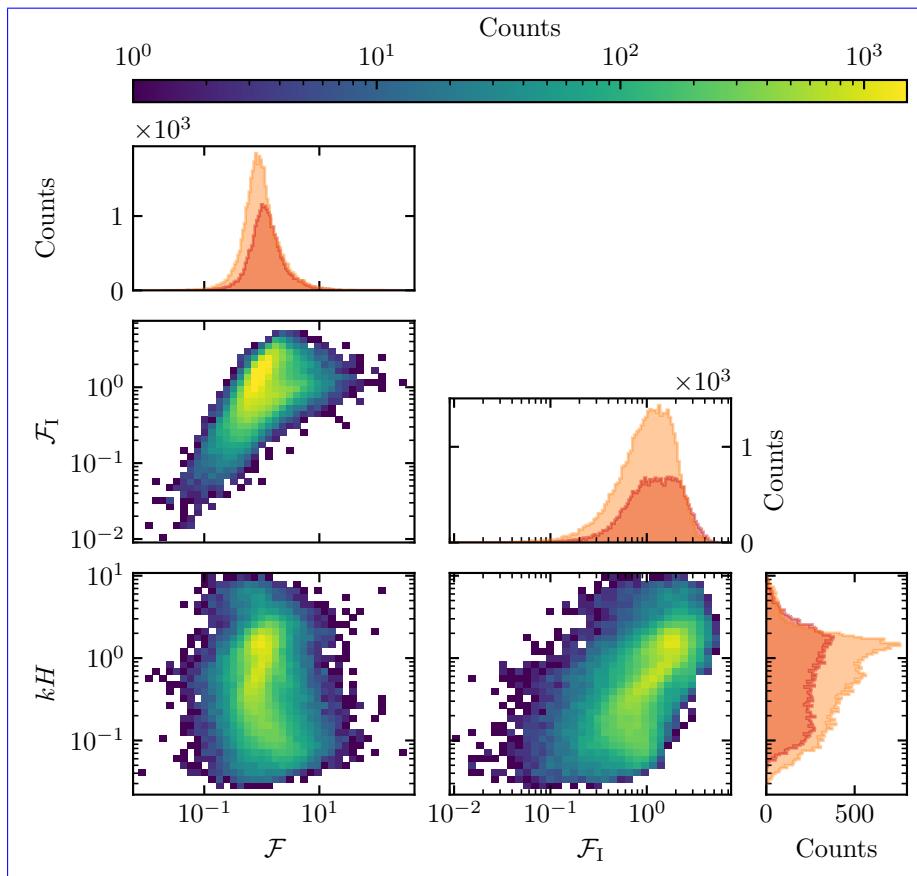


Fig. S10 Non-dimensional parameters distributions. For the marginal distributions, the light orange corresponds to the South Sand Sea station, and the blue to the North Sand Sea station.

Regime diagrams of the wind deviation δ_θ and relative attenuation/amplification δ_u in the spaces (\mathcal{F}_I, kH) and $(\mathcal{F}_I, \mathcal{F})$, containing the data from both the North Sand Sea and South Sand Sea stations. Green dashed lines empirically delimit the different regimes. The point density in each bin of the diagrams is shown in Online Resource Fig. S14 – 95% of the data occur in the range $-1 < \delta u < 1$. The similar regime diagrams in the space (\mathcal{F}, kH) are shown in Fig. 5.

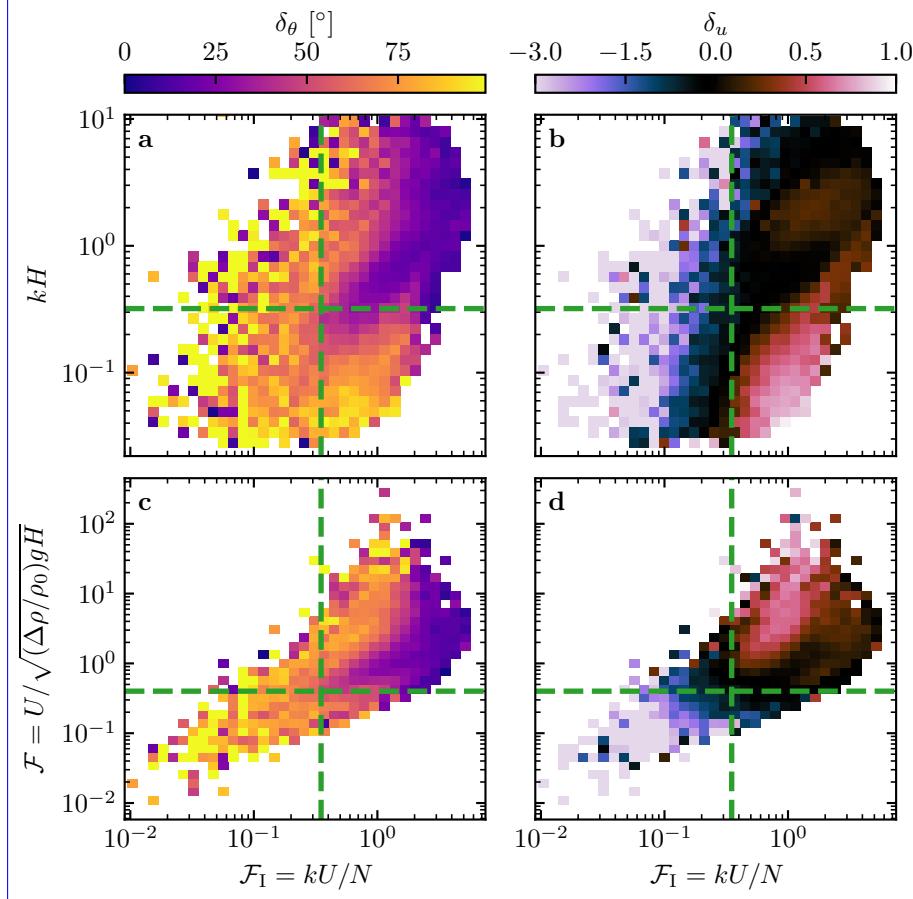


Fig. S11 Regime diagrams of the wind deviation δ_θ and relative attenuation/amplification δ_u in the spaces (\mathcal{F}_I, kH) and $(\mathcal{F}_I, \mathcal{F})$, containing the data from both the North Sand Sea and South Sand Sea stations. Green dashed lines empirically delimit the different regimes. The point density in each bin of the diagrams is shown in Online Resource Fig. S10 – 95% of the data occur in the range $-1 < \delta u < 1$. The similar regime diagrams in the space (\mathcal{F}, kH) are shown in Fig. 8.

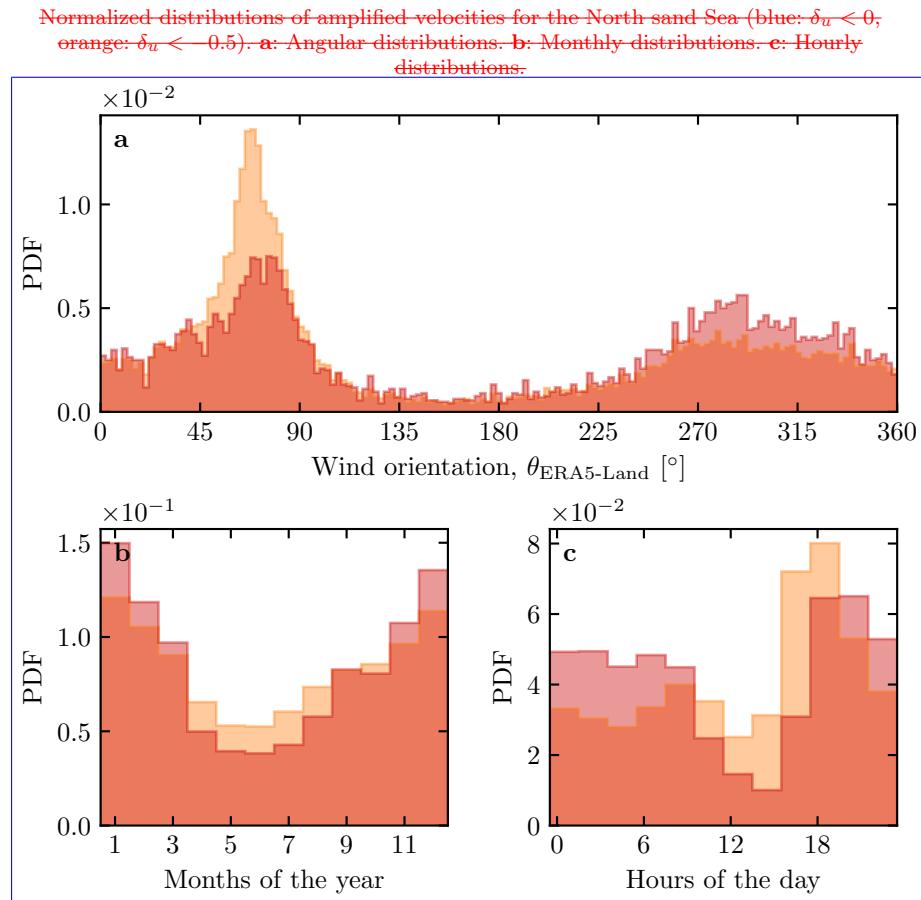


Fig. S12 Normalized distributions of amplified velocities for the North Sea (light orange: $\delta_u < 0$, orange: $\delta_u < -0.5$). **a:** Angular distributions. **b:** Monthly distributions. **c:** Hourly distributions.

Computation of the flow disturbance with the linear model of Andreotti et al. (2009). **a** and **b** Magnitude of the hydrodynamic coefficients A_0 and B_0 , calculated from the values of the non-dimensional numbers corresponding to the ERA5 Land time series presented in Figs. 4 and 5. **c** Shear velocity streamlines over sinusoidal ridges of amplitude $k\xi_0 = 0.1$ and for increasing values of $\sqrt{A_0^2 + B_0^2}$. From the upper to the lower streamline, values of $(kH, \mathcal{F}, \mathcal{F}_I, A_0, B_0, \sqrt{A_0^2 + B_0^2})$ are (1.9, 0.6, 1.5, 3.4, 1.0, 3.5), (1.5, 0.3, 0.4, 4.8, 1.4, 5.0), (0.1, 3.5, 1.0, 8.6, 0.1, 8.6), (0.5, 0.05, 0.04, 9.6, 2.5, 9.9).

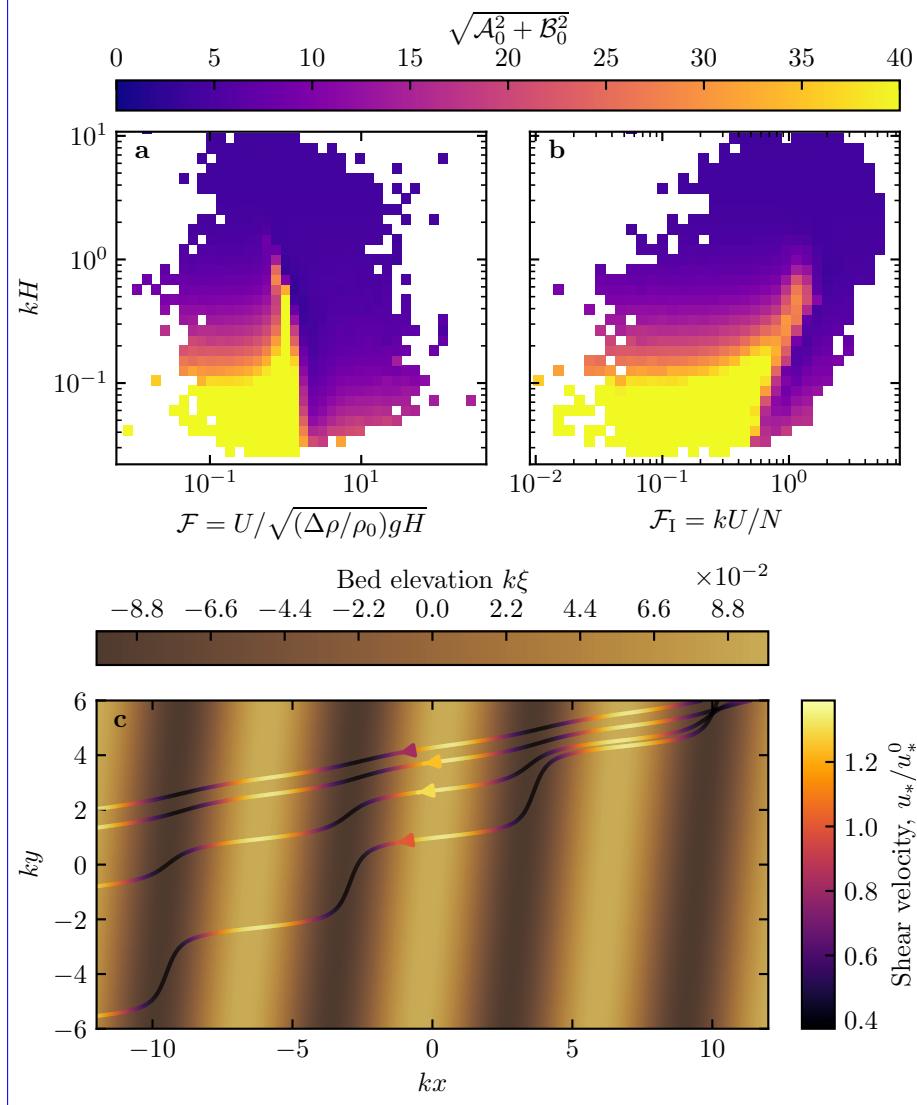


Fig. S13 Computation of the flow disturbance with the linear model of Andreotti et al. (2009). **a–b:** Magnitude of the hydrodynamic coefficients A_0 and B_0 , calculated from the time series of the non-dimensional numbers corresponding to the ERA5-Land wind data and ERA5 data on vertical pressure levels. **c:** Shear velocity streamlines over sinusoidal ridges of amplitude $k\xi_0 = 0.1$ and for increasing values of $\sqrt{A_0^2 + B_0^2}$. From the upper to the lower streamline, values of $(kH, \mathcal{F}, \mathcal{F}_I, A_0, B_0, \sqrt{A_0^2 + B_0^2})$ are (1.9, 0.6, 1.5, 3.4, 1.0, 3.5), (1.5, 0.3, 0.4, 4.8, 1.4, 5.0), (0.1, 3.5, 1.0, 8.6, 0.1, 8.6), (0.5, 0.05, 0.04, 9.6, 2.5, 9.9).

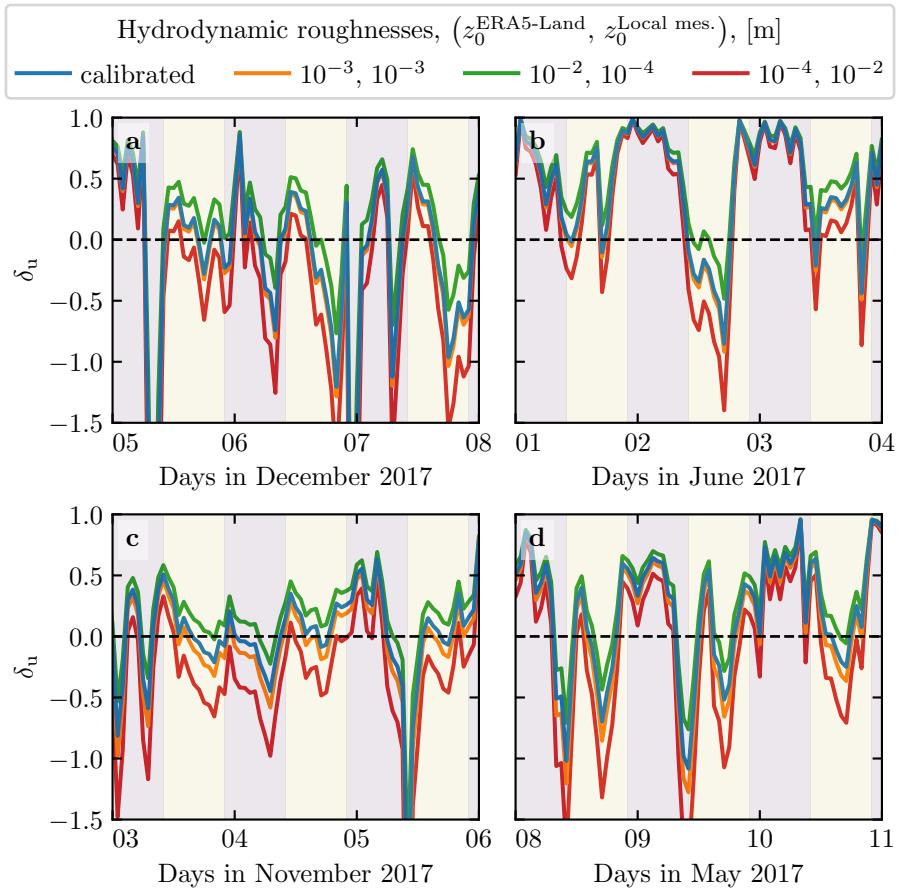


Fig. S14 Time series of the relative velocity disturbance δ_u corresponding to Fig. 5, for different values of the hydrodynamic roughnesses. **a:** North Sand Sea – summer. **b:** North Sand Sea – winter. **c:** South Sand Sea – summer. **d:** South Sand Sea – winter. Note that δ_θ is independent of the choice of $z_0^{\text{ERA5-Land}}$ and $z_0^{\text{Local mes.}}$.