

## Noether's theorem

Consider a continuous transformation on field  $\varphi$ :

$$\varphi(x^\mu) \rightarrow \varphi'(x^\mu) = \varphi(x^\mu) + a \Delta \varphi(x)$$

with  $a$  an infinitesimal parameter

This is a symmetry if the equations of motion (the Euler-Lagrange eqs.) remain invariant. This is the case if  $\mathcal{L}$  is invariant up to a divergence

$$\mathcal{L}(x^\mu) \rightarrow \mathcal{L}(x^\mu) + a \partial_\mu J^\mu(x^\mu) = \mathcal{L}(x^\mu) + a \Delta \mathcal{L}(x^\mu)$$

Then

$$\begin{aligned} a \Delta \mathcal{L} &= \frac{\partial \mathcal{L}}{\partial \varphi} (a \Delta \varphi) + \frac{\partial \mathcal{L}}{\partial (\partial_\mu \varphi)} \partial_\mu (a \Delta \varphi) = a \frac{\partial \mathcal{L}}{\partial \varphi} \Delta \varphi + a \partial_\mu \left( \frac{\partial \mathcal{L}}{\partial (\partial_\mu \varphi)} \Delta \varphi \right) \\ &\quad - a \partial_\mu \left( \frac{\partial \mathcal{L}}{\partial (\partial_\mu \varphi)} \right) \Delta \varphi \\ &= -a \left[ \underbrace{\partial_\mu \left( \frac{\partial \mathcal{L}}{\partial (\partial_\mu \varphi)} \right) - \frac{\partial \mathcal{L}}{\partial \varphi}}_{=0} \right] \Delta \varphi + a \partial_\mu \left( \frac{\partial \mathcal{L}}{\partial (\partial_\mu \varphi)} \Delta \varphi \right) \end{aligned}$$

$$\text{Thus } \partial_\mu J^\mu(x^\mu) = \partial_\mu \left( \frac{\partial \mathcal{L}}{\partial (\partial_\mu \varphi)} \Delta \varphi \right) \Rightarrow \partial_\mu \left[ \frac{\partial \mathcal{L}}{\partial (\partial_\mu \varphi)} \Delta \varphi - J^\mu \right] = 0$$

$$\Rightarrow \partial_\mu j^\mu = 0 \quad \text{where} \quad j^\mu = \frac{\partial \mathcal{L}}{\partial (\partial_\mu \varphi)} \Delta \varphi - J^\mu$$

So current  $j^\mu$  is conserved

$$\text{Charge } Q = \int j^0 d^3x$$

Consider, next, spacetime transformations

Infinitesimal translation  $x^\mu \rightarrow x^\mu + a^\mu$

$$\text{Then } \varphi(x) \rightarrow \varphi(x^\mu + a^\mu) = \varphi(x^\mu) + a^\mu \partial_\mu \varphi(x)$$

Also  $\mathcal{L} \rightarrow \mathcal{L} + a^\mu \partial_\mu \mathcal{L} = \mathcal{L} + g^{\mu\nu} a_\nu \partial_\mu \mathcal{L} = \mathcal{L} + a_\nu \partial_\mu (g^{\mu\nu} \mathcal{L})$   
( $\mathcal{L}$  is scalar)

Since also  $\mathcal{L} \rightarrow \mathcal{L} + a^\nu \partial_\nu \mathcal{L}$  then  $J^\mu_\nu = g^\mu_\nu \mathcal{L}$  or  $J^{\mu\nu} = g^{\mu\nu} \mathcal{L}$

Then we have four conserved currents  $T^{\mu\nu} = \frac{\partial \mathcal{L}}{\partial(\partial_\mu \varphi)} \partial^\nu \varphi - g^{\mu\nu} \mathcal{L}$   
 $T^{\mu\nu}$  is the energy-momentum tensor or stress-energy tensor

The Hamiltonian  $H$  is  $H = \int T^{00} d^3x = \int \mathcal{H} d^3x$  (with  $\mathcal{H}$  the Hamiltonian density)  
(energy)

and it is the conserved quantity associated with time translations.

Also the physical momenta  $p^i = \int T^{0i} d^3x$   
are the conserved quantities associated with spatial translations

Note that stress-energy tensors appear in field equations of general relativity

$$R^{\mu\nu} - \frac{1}{2} g^{\mu\nu} R = -8\pi G T^{\mu\nu}$$

$\downarrow$  Ricci tensor       $\downarrow$  Ricci scalar

Real scalar field (Klein-Gordon field)  $\mathcal{L} = \frac{1}{2} \partial_\mu \varphi \partial^\mu \varphi - \frac{1}{2} m^2 \varphi^2$

$$T^{00} = \frac{\partial \mathcal{L}}{\partial(\partial_0 \varphi)} \partial^0 \varphi - g^{00} \mathcal{L} = \partial^0 \varphi \partial^0 \varphi - \frac{1}{2} \partial_\mu \varphi \partial^\mu \varphi + \frac{1}{2} m^2 \varphi^2$$

So the field energy is  $H = \int T^{00} d^3x = \frac{1}{2} \int ((\partial^0 \varphi)^2 + \vec{\nabla} \varphi \cdot \vec{\nabla} \varphi + m^2 \varphi^2) d^3x$

which is positive definite so not affected by negative-energy problem

Conjugate (or canonical) momentum  $\pi(x^\mu) = \frac{\partial \mathcal{L}}{\partial \dot{\varphi}} = \dot{\varphi}(x^\mu)$

Also Hamiltonian density  $\mathcal{H}$  is given by

$$\mathcal{H} = \pi \dot{\varphi} - \mathcal{L} = \dot{\varphi} \dot{\varphi} - \mathcal{L} = T^{00} \text{ consistent with above expressions}$$

Fourier expansion for the field  $\varphi$ :

$$\varphi(x^\mu) = \int \frac{d^4 p}{(2\pi)^4} 2\pi \delta(p^2 - m^2) \theta(p^0) \left[ a(p^\mu) e^{-i p^\mu x_\mu} + a^\dagger(p^\mu) e^{i p^\mu x_\mu} \right] (2p^0)^{1/2}$$

$$\Rightarrow \varphi(x) = \int \frac{d^3 p}{(2\pi)^3 (2p^0)^{1/2}} \left[ a(p) e^{-i p \cdot x} + a^\dagger(p) e^{i p \cdot x} \right]$$

$$\text{where } p^0 = (\vec{p}^2 + m^2)^{1/2}$$

Field  $\varphi$  is a Hermitian operator  $\varphi^\dagger = \varphi$



Real scalar field  $\varphi$

"second quantization"

Equal-time commutation relations

$$[\varphi(\vec{x}, t), \pi(\vec{y}, t)] = i \delta^3(\vec{x} - \vec{y})$$

$$[\varphi(\vec{x}, t), \varphi(\vec{y}, t)] = 0 \text{ and } [\pi(\vec{x}, t), \pi(\vec{y}, t)] = 0$$

(these are the analogs of  $[\hat{x}, \hat{p}_x] = i$ ,  $[\hat{x}, \hat{x}] = 0$ ,  $[\hat{p}_x, \hat{p}_x] = 0$ )

Inverse Fourier transform to get operators  $a(p)$  and  $a^\dagger(p)$ :

$$a(p) = \int d^3x (2\pi)^{-3/2} i [e^{ip \cdot x} \partial_0 \varphi(x) - (\partial_0 e^{ip \cdot x}) \varphi(x)]$$

$$\text{and } a^\dagger(p) = \int d^3x (2\pi)^{-3/2} i [\varphi(x) \partial_0 e^{-ip \cdot x} - (\partial_0 \varphi(x)) e^{-ip \cdot x}]$$

$$\text{Then } [a(p), a^\dagger(p')] = (2\pi)^3 \delta^3(\vec{p} - \vec{p}')$$

$$[a(p), a(p')] = 0 \text{ and } [a^\dagger(p), a^\dagger(p')] = 0$$

Construct operator  $N(p) = a^\dagger(p) a(p)$  with eigenkets  $|n(p)\rangle$  and eigenvalues  $n(p)$ :  $N(p) |n(p)\rangle = n(p) |n(p)\rangle$

$$\text{We have } [N(p), a(p)] = -a(p) \text{ and } [N(p), a^\dagger(p)] = a^\dagger(p)$$

(compare with simple harmonic oscillator in quantum mechanics)