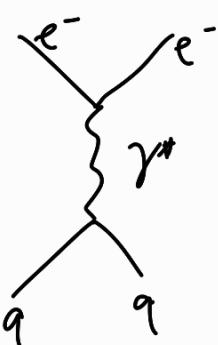
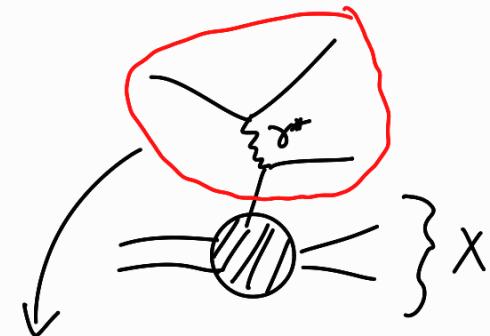
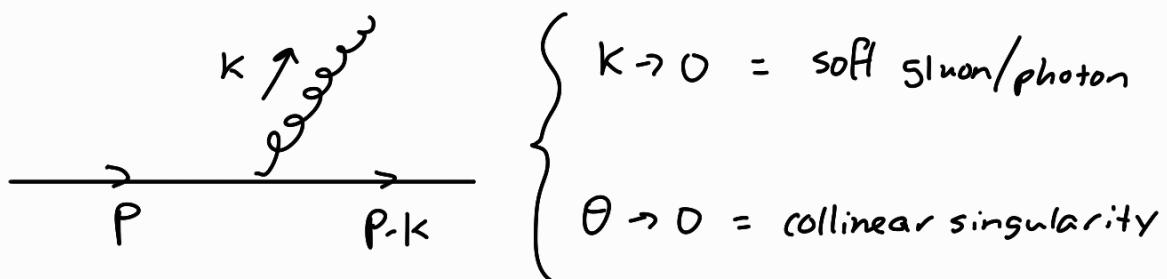


# Renormalization

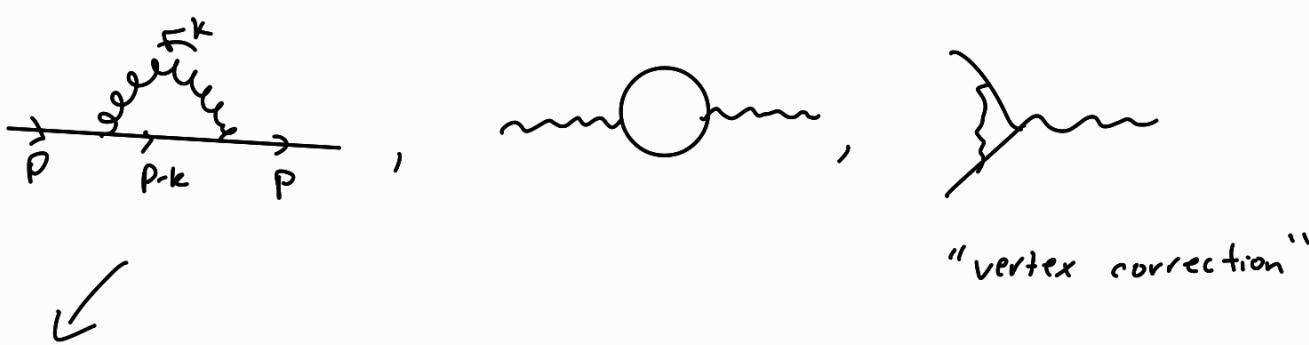
- Deep inelastic scattering at tree level:



- no gluon, no QCD anything
- $\rightarrow 30\%$  off of experimental value



- these arise due to detector limitations and other things
- other "corrections":

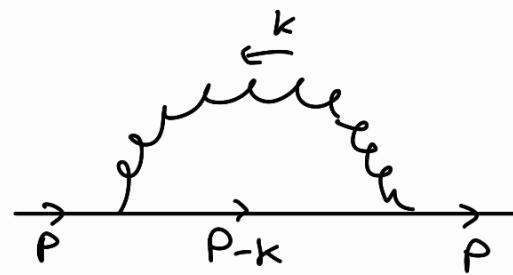


· involves an  $\int \frac{d^4 k}{(2\pi)^4}$

- however, such integrals diverge for certain energies
- we can "manipulate, absorb, etc." these diagrams to, remarkably, avoid the infinities

- there is also a process called "regularization" that is a non-trivial mathematical magic
  - involves turning integral into a function that we can do the limit of
- 

## Self Energy Diagram for Quark or Electron:



$\rightarrow \Sigma_{ij}(p)$  is this self-energy

$\rightarrow$  The full (including all self energies)  $\equiv \tilde{S}_{ij}(p)$

Can be shown that

$$\tilde{S}_{ij} = \frac{\delta_{ij}}{m - p - \Sigma(p)}$$

$$w/ \quad \Sigma_{ij}(p) = \delta_{ij} \Sigma(p)$$

$$w/ \quad \tilde{S}_{ij} = i \int d^4x e^{-ip \cdot x} \langle 0 | T \bar{\psi}_i(x) \bar{\psi}_j(0) | 0 \rangle_c$$

- Fourier transform
- subscript c means "connected"
- series expand 1st  $\tilde{S}_{ij}$  in terms of  $\Sigma(p)$ , the "1-particle irreducible" diagram (1PI)

$$\tilde{S}_{ij}(p) = \delta_{ij} \left\{ \tilde{S}_0(p) + \tilde{S}_0(p) \Sigma(p) \tilde{S}_0(p) + \left( \tilde{S}_0 \Sigma(p) \tilde{S}_0 \right)^2 + \dots \right\}$$



- we interpret it by original propagator, so

$$\tilde{S}_{ij}(p) = \frac{1}{1PI} + \frac{\text{gluon}}{1PI} + \frac{\text{gluon gluon}}{\text{not } 1PI}$$

- these 1PI cannot be cut to create 2 new diagram

first step for renorm/regularization is finding these 1PI's.

Feynman rules for QCD:

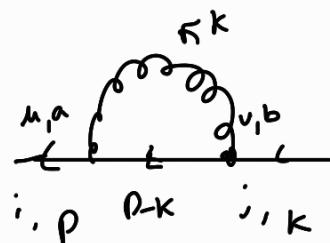
$$\overrightarrow{\text{line}} \rightarrow \frac{1}{m-p}, \quad \overset{a}{\overbrace{\text{line}}} \overset{b}{\overbrace{\text{line}}} \rightarrow \frac{\delta_{ab}}{k^2} d^{uv}(k)$$

$$\text{loop} \rightarrow \int \frac{d^4 k}{(2\pi)^4 i} \quad d^{uv}(k) = g^{uv} \text{ for Feynman gauge}$$

$$= g^{uv} - (1-\alpha) \frac{k_u k_v}{k^2} \text{ for others}$$

$$\overrightarrow{\text{line}} \overset{a, u}{\text{line}} \rightarrow g_s \gamma^u T_{ij}^a \quad \begin{matrix} \leftarrow \text{adjoint rep. 1-8 (the algebra)} \\ \leftarrow \text{fund. rep 1-3} \end{matrix}$$

how:



start w/  $\int \frac{d^4 k}{(2\pi)^4 i} g_s T_{il}^a \gamma^u \quad \text{, and a loop } \int \frac{d^4 k}{(2\pi)^4 i}$

$$\boxed{\int \frac{d^4 k}{(2\pi)^4 i} g_s T_{il}^a \gamma^u \left[ \frac{\delta_{ln}}{m - (p - k)} \right] g_s T_{nj}^b \gamma^v \cdot \frac{\delta_{ab}}{k^2} d^{uv}(k)}$$

$\downarrow$   
fermion propagator      gluon propagator

\* a quark can't transform in between, l must be equal to n

- looking at  $SU(3)$  generators:

$$\rightarrow \sum_{ij}(\rho) = \delta_{ab} \delta_{ln} T_{il}^a T_{nj}^b = (T^a_i T^a_j)_{ij} = \delta_{ij} C_F,$$

- $C_F = \frac{N_c^2 - 1}{N_c}$ ,  $N_c$  = num of colors; for  $SU(3)_c$ ,  $N_c = 3$

. Casimir operator of the fundamental representation

- $C_F = \frac{8}{3}$

$$\Rightarrow \sum_{ij}(\rho) = g_s^2 \delta_{ij} C_F \int \frac{d^4 k}{(2\pi)^4 i} \gamma^u \left[ \frac{1}{m - p + k} \right] \gamma^v \cdot \frac{g_{uv}}{k^2} = \delta_{ij} \sum(\rho)$$

- in Feynman Gauge,  $\alpha=1$  so  $d^{uv} = g^{uv}$

$$\Rightarrow \Sigma(p) = g_s^2 C_F \int \frac{d^4 k}{(2\pi)^4} \gamma^\mu \left[ \frac{1}{m - p + k} \right] \gamma_\mu \cdot \frac{1}{k^2}$$

- we can do  $\frac{1}{p} = \frac{p}{p^2}$  (mult top/bottom by  $p$ )

$$\rightarrow = g_s^2 C_F \int \frac{d^4 k}{(2\pi)^4} \gamma^\mu \frac{m - p + k}{m^2 - (p+k)^2} V_\mu \cdot \frac{1}{k^2}$$

· For large  $k$ , and gammas irrelevant:

$$\sim \int \frac{d^4 k}{(2\pi)^4} \frac{k}{k^2 \cdot k^2}$$

$$\sim \int d^4 k \frac{1}{k^3}$$

This is volume in 4D, but this is in large  $k$  behavior, so this diverges ("linearly divergent")

$$\lim_{k \rightarrow \infty} K$$

- In 'Cutoff' regularization, we just cut the integral:

$$\int \frac{d^4 k}{(2\pi)^4}$$

- In Pauli-Villars: replace propagator with

$$\frac{1}{m^2 - k^2} - \frac{1}{M^2 - k^2} = \frac{M^2 - m^2}{(m^2 - k^2)(M^2 - k^2)} \quad \text{which reduces to}$$

regular propagator in  $M \rightarrow \infty$

. better for Lorentz invariance than 'Cutoff', but fails for gauge theories with massive gauge bosons

- Analytical Regularization

$$\frac{1}{m^2 - k^2} \rightarrow \frac{1}{(m^2 - k^2)^\alpha}, \quad \alpha \in \mathbb{C} \text{ w/ } \operatorname{Re}(\alpha) > 1$$

. helpful to prove renormalizability

. but violates gauge invariance

## Lattice Regularization:

- discretize Minkowski spacetime
- short-distance contribution in configuration space which corresponds to the large momentum contribution is eliminated

## Dimensional Regularization (DR):

- spoils NO symmetries!
- maps integral into series of analytical terms in terms of dimension:

$$\int \frac{d^4 k}{(2\pi)^4} i \rightarrow \int \frac{d^D k}{(2\pi)^D} i$$

- try for  $D < 4$ .

## 8/28 : More Regularization

Green's Functions: quantum expectation value of a time-ordered product of fields

$$\hat{\phi}(x) = \langle 0 | T[\phi(x_1) \dots \phi(x_n)] | 0 \rangle$$

- "correlator"

- Simplest ex: Dirac Propagator:

$$\langle 0 | \bar{\psi}_a(x) \bar{\psi}_b(y) | 0 \rangle$$

going from point x to y

- or  $\langle 0 | \bar{\psi}_b(y) \bar{\psi}_a(x) | 0 \rangle$

- We know solutions of Dirac EQ:

$$\bar{\psi}_a(x) = \int \frac{d^3 p}{(2\pi)^3} \frac{1}{\sqrt{2\varepsilon_p}} \sum_s \left( a^{(s)} u^{(s)}(p) e^{-ip \cdot x} + b^{(s)\dagger} v^{(s)}(p) e^{ip \cdot x} \right)$$

$a$ 's /  $b$ 's become operators in 2<sup>nd</sup> quantization: satisfy anti-commutation relations

- we can now calc. the two amplitudes above; can group them together  $\rightarrow$  "retarded" Green's function

$$S_R^{ab}(x-y) = \Theta(x^0 - y^0) \text{L} \langle \{ \gamma_a(x), \gamma_b(y) \} \rangle \text{L}$$

$$S_R(x-y) = (i\not D - m) D_R(x-y)$$

~~★ look further at this ★~~

- back to dimensional regularization

$$\int d^4 k \rightarrow \int d^D k$$

- we want  $D=4$  and to express our divergent integrals as an analytic function of  $D$

- 1<sup>st</sup> consequence:  $m=0, \dots, 3$ ,  $m=0, \dots, D-1$

. now,  $g^{\mu\nu} g_{\mu\nu} = D$ .

$$\{ \gamma^m, \gamma^\nu \} = 2g^{m\nu}, \quad \gamma^m \gamma_m = D$$

$$\gamma^m \gamma_\nu \gamma_m = (2-D)\gamma_\nu$$

. in our integral, we had a  $(2\pi)^4$ , but this is a bit funny

. what we do is

$$\int \frac{1}{(2\pi)^4} \rightarrow \int \frac{1}{(2\pi)^D}$$

but the "measure" in  $D$ -dim must recover  $\frac{1}{(2\pi)^4}$  for  $D=4$

. 2<sup>nd</sup> consequence:  $\text{Tr}[\gamma^m \gamma^\nu] = 2^{1/2} g^{m\nu}$  for  $D$  even

. correctly reduces to  $4g^{m\nu}$  for  $D=4$ .

. this gets messy; what we do instead is keep  $D$ -dim measure, but use 4-dim traces

$$\rightarrow \int \frac{d^D k}{(2\pi)^D} \quad \& \quad \text{Tr}[\gamma^m \gamma^\nu] = 2g^{m\nu}$$

- now we have conventions; can tackle self energy
- for simplicity,  $m \rightarrow 0$ , so

geller

self energy

$$\rightarrow \Sigma(p) = g_s^2 C_F(2-D) \int_{(2\pi)^D} \frac{d^D k}{k^2 (k-p)^2}$$

- for  $D < 3$ , this converges for sure. but, what do we do? what is  $d^D x$ ?
- We introduce "Feynman Parametrization":

$$\begin{aligned} \frac{1}{k^2(k-p)} &\rightarrow \frac{1}{AB} = \int_0^1 \frac{dx}{\{xA + (1-x)B\}^2} \quad u = x(A-B) + B \\ &\quad du = (A-B)dx \quad \frac{du}{A-B} = dx \\ &= \frac{1}{B-A} \int_A^B \frac{\frac{du}{A-B}}{u^2} = \frac{1}{A-B} \left[ \frac{1}{u} \right]_A^B \\ &= \frac{1}{A-B} \left[ \frac{1}{B} - \frac{1}{A} \right] \\ &= \frac{1}{A-B} \left[ \frac{A-B}{AB} \right] = \frac{1}{AB} \quad \checkmark \end{aligned}$$

- this can be generalized to any number of factors?

- Now:

$$\Sigma(p) = g_s^2 C_F(2-D) \int_{(2\pi)^D} \frac{d^D k}{(2\pi)^D} (k-p) \int_0^1 \frac{dx}{[x(k-p)^2 + (1-x)k^2]^2}$$

- again, integral converges for  $D < 3$

- apparently we can swap the two integral signs?

$$\Sigma(p) = g_s^2 C_F(2-D) \int_{(2\pi)^D} \frac{d^D k}{(2\pi)^D} \int_0^1 \frac{k-p}{[(k-xp)^2 + x(1-x)p^2]^2}$$

- since this preserves translational symmetry, we can shift the momentum:  $k' = k-xp$

$$\Sigma(p) = g_s^2 C_F(2-D) \int_0^1 dx \int_{(2\pi)^D} \frac{d^D k'}{[(k'-x(1-x)p)^2 + x(1-x)p^2]^2}$$

- now we have the sum of 2 functions. the first is a function of  $k'$ , but we see it's odd: evaluates to zero w/ symmetric limits

$$\rightarrow \Sigma(p) = g_s^2 C_F(D-2) p \int_0^1 dx (1-x) \int_{(2\pi)^D} \frac{d^D k'}{[k'^2 + x(1-x)p^2]^2}$$

- now we do a "Wick Rotation" (again, DR preserves all symmetries)

- in Minkowski space,  $k'^2 = k''^2 - |\vec{k}'|^2$

- go to Euclidean space where we want  $-k'^2 - |\vec{k}'|^2$

- give it an  $i!$  basically a complex rotation

$$k'_0 = i\bar{k}_0, \quad \bar{k} \in \mathbb{R}$$

$$\bar{k}' = \bar{\vec{k}}$$

$$\text{now, } k'^2 = -K_0^2 - |\bar{k}|^2$$

$$\text{and } d^D k \rightarrow i d^D \bar{k}$$

<sup>10,</sup>

$$\Sigma(p) = g_s^2 C_F(D-2) \rho \int_0^1 dx (1-x) \int \frac{d^D \bar{k}}{(2\pi)^D} \frac{1}{(\bar{k}^2 + L)^2}$$

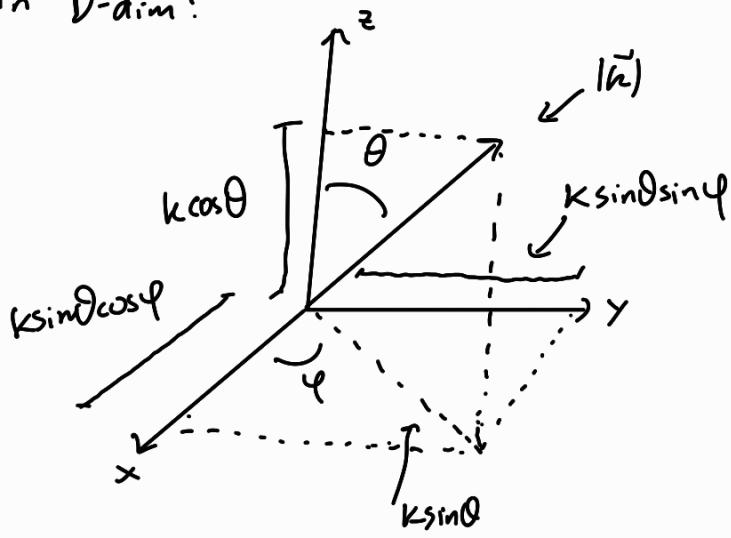
$$\text{where } L = -x(1-x)\rho^2$$

- Now it's Euclidean: we know it's non-singular for  $L > 0$
- . so, we want to keep  $\rho^2 < 0$ : "space-like"

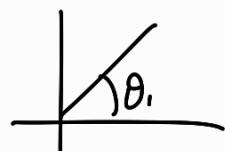
in 3-D, we do spherical coords:

$$d^3 \bar{k} = d\omega_3 |\bar{k}|^2 d|\bar{k}|$$

similar here! In D-dim:



1)  $S_1$   
one angle



$$\rightarrow \begin{cases} x^1 = \cos \theta, \\ x^2 = \sin \theta, \end{cases}$$

$$|\bar{k}| = 1$$

2)  $S_2$   
two angles

$$\begin{cases} x^1 = \cos \theta_1, \\ x^2 = \sin \theta_1 \cos \theta_1, \\ x^3 = \sin \theta_1 \sin \theta_1 \end{cases}$$

here, project to 3-D

3)  $S_3$  (hypersphere!)  
three angles  $\theta_1, \theta_2, \theta_3$

$$\begin{cases} x^1 = \cos \theta_1, \\ x^2 = \sin \theta_1 \cos \theta_2, \\ x^3 = \sin \theta_1 \sin \theta_2 \cos \theta_3, \\ x^4 = \sin \theta_1 \sin \theta_2 \sin \theta_3 \end{cases} \quad \begin{matrix} \leftarrow & \text{then to 2-D} \\ \leftarrow & \end{matrix}$$

4)  $S^5$

Four angles

$$\left\{ \begin{array}{l} x_1 = \cos\theta_1 \\ x_2 = \sin\theta_1 \cos\theta_2 \\ x_3 = \sin\theta_1 \sin\theta_2 \cos\theta_3 \\ x_4 = \sin\theta_1 \sin\theta_2 \sin\theta_3 \cos\theta_4 \\ x_5 = \sin\theta_1 \sin\theta_2 \sin\theta_3 \sin\theta_4 \end{array} \right.$$

We can tell the pattern now:

$S^D$  D-1 angles

$$\left\{ \begin{array}{l} x_1 = \cos\theta_1 \\ x_2 = \sin\theta_1 \cos\theta_2 \\ \vdots \\ x_{D-1} = \sin\theta_1 \dots \sin\theta_{D-2} \cos\theta_{D-1} \\ x_D = \sin\theta_1 \dots \sin\theta_{D-1} \end{array} \right.$$

- now:  $\underline{k}_0 = |\underline{k}| \cos\theta_1$

:

$$\underline{k}_{D-1} = |\underline{k}| \sin\theta_1 \dots \sin\theta_{D-1}$$

and  $d^D \underline{k} = d\Omega_D |\underline{k}|^{D-1} d|\underline{k}|$

$$d\Omega_D = \prod_{l=1}^{D-1} \sin^{D-1-l} \theta_l d\theta_l$$

$\rightarrow \int d\Omega_D = \text{Ansatz}$

HW: calc:

$$\int_0^\pi d\theta (\sin\theta)^n$$

a/4) Volume element in D dimension:

$$d^D \underline{k} = |\underline{k}|^{D-1} d|\underline{k}| d\Omega_D$$

where

$$d\Omega_D = \prod_{l=1}^{D-1} \sin^{D-1-l} \theta_l d\theta_l$$

now, this is a relation in math, we can say

$$\int \frac{d^D K}{(2\pi)^D} \frac{1}{(K^2 + 2)^2} = \frac{B(\rho/2, 2 - \rho/2)}{(4\pi)^{\rho/2}} \frac{\Gamma(\rho/2 - 2)}{\Gamma(\rho/2)} \quad \textcircled{#1}$$

one representation of  $\Gamma$  is

$$\Gamma(z) = \int_0^\infty t^{z-1} e^{-t} dt \quad \text{for } z \in \mathbb{C} \text{ w/ } \operatorname{Re}[z] > 0$$

$$\text{The "beta" function } B(p, q) = \frac{\Gamma(p)\Gamma(q)}{\Gamma(p+q)}$$

also has an integral representation:

$$B(p, q) = \int_0^\infty \frac{t^{p-1}}{(1+t)^{p+q}} dt$$

to get  $\textcircled{#1}$ , we can get from the Jacobian:

$$\begin{aligned} \int_0^\pi d\Omega_D &= \int_0^\pi d\theta_0 (\sin\theta_0)^{D-2} \dots \int_0^\pi d\theta_{D-2} (\sin\theta_{D-2})^{D-2} \int_0^\pi d\theta_{D-1} \\ &= \frac{2\pi^{\rho/2}}{\Gamma(\rho/2)} \end{aligned}$$

To show this, let's examine unit sphere in D-dim

first, Gaussian integral

$$\begin{aligned} \Gamma_D &= \int_{-\infty}^{\infty} e^{-x^2} dx \\ \rightarrow (\Gamma_D)^D &= \left( \int_{-\infty}^{\infty} e^{-x^2} dx \right)^D \end{aligned}$$

in D-dim, we can take the integral over each independent coord:

$$= \int_{-\infty}^{\infty} e^{-\sum_{i=1}^D x_i^2} d^D x$$

< polar coords again!  $d^D x = x^{D-1} e^{-x^2} dr d\Omega_D$ ,

$$\text{where } x \equiv \sqrt{x_1^2 + x_2^2 + \dots + x_D^2}$$

$$\text{so } -\sum_{i=1}^D x_i^2 = -x^2, \text{ so}$$

$$\rightarrow \int d\Omega_D \int x^{D-1} e^{-x^2} dx$$

$$= \int d\omega_D \underbrace{\frac{1}{2} \int_0^\infty (x^2)^{D/2-1} e^{-x^2} dx^2}_{\text{well-known}} \\ = \frac{1}{2} \Gamma(D/2)$$

so  $(\sqrt{m})^D = \frac{1}{2} \Gamma(D/2) \int d\omega_D$

or  $\int d\omega_D = \frac{2m^{D/2}}{\Gamma(D/2)}$

$D$	$\Gamma(D/2)$	$\int d\omega_D$
1	$\sqrt{\pi}$	2
2	1	$2m$
3	$\pi^{1/2}$	$4\pi$
4	1	$2m^2$
$\vdots$	$\vdots$	$\vdots$

At last, we return to our quark self-energy.

$$\sum(p) = g_s^2 C_F (D-2) \cancel{\int} \frac{\Gamma(2-D/2)}{(4\pi)^{D/2}} (-p^2)^{D/2-2}$$

$$\times \int_0^1 dx x^{D/2-2} (1-x)^{D/2-1}$$

can express integral in terms of Beta

$$\sum(p) = \frac{2C_F g_s^2}{(4\pi)^{D/2}} \cancel{\int} (-p^2)^{D/2-2} (D-1) B\left(\frac{D}{2}, \frac{D}{2}\right) \Gamma\left(2 - \frac{D}{2}\right)$$

- analytic fn of  $D$ ; valid for  $D > 3 \& p^2 < 0$ .
- want to "analytic continuation" in a region where  $D \& p$  are arbitrary complex numbers.
- First:  $\Gamma(2 - \frac{D}{2}) = \int_0^\infty t^{2-D/2-1} e^{-t} dt$  [from  $\int_0^\infty t^{z-1} e^{-t} dt$ ]
- . observe:  $D = 4, 6, 8, \dots$  are poles; lead to  $\frac{1}{t}$ , which can't do 0 lower limit
- . for  $D = 4$ :  $\Gamma(0) = \int_0^\infty \frac{e^{-t}}{t} dt \rightarrow \infty$ ; diverges...
- Second: there is a "branch cut" on the positive real axis in the  $p^2$  plane
- . imagine  $(-p^2)^\alpha$  for small  $\alpha$  - taylor exp:

$$= 1 + \ln(-p^2)\alpha + \frac{1}{2!} \ln(-p^2)\alpha^2 + \dots, \text{ we have a log!}$$

•  $p$  can be  $\pm$ ; approaching zero from either side gives diff infities.

• for  $D=4$ , we'd have  $S(p) \sim \frac{c_F g_S}{(4\pi)^D} \frac{2}{4-D} p^4$

↖ !!

• Now,  $S = \int L dt = \int \mathcal{L} d^4x \rightarrow \mathcal{L}$  has [energy], so  $S$  has [E.t].

This is same unit as  $\hbar$ , which is angular momentum.

In natural units,  $\hbar=c=1$ . So we need this to be satisfied.

In  $D$ -dim:  $S = \int \mathcal{L} d^Dx$ . for  $\hbar=c=1$ , we'll have  $\dim[\mathcal{L}] = 0$

$\mathcal{L}$   
called "mass dimension"

- Compton wavelength:  $\lambda = \underline{\lambda^0(1-\cos\theta)}$ , where  $\lambda = \frac{\hbar}{mc} \rightarrow \frac{1}{m}$  (mass),

so  $[\lambda] = [m]^{-1}$

• Now we look inside the  $\mathcal{L}$  and see the mass dimensions, because we must have that  $[\mathcal{L}] = 0$ .

• Let's examine

$$\cancel{x}^{\mu, \alpha} i g_S \bar{\psi} T^\alpha \gamma^\mu \psi A_\mu^\alpha$$

$\dim[g_S] = ?$

$\dim[\bar{\psi}] = \dim[\psi] =$

$\dim[A_\mu^\alpha] =$

• first,

$$\mathcal{L} = a_1 \frac{1}{4} F^{\mu\nu} F_{\mu\nu} + a_2 \bar{\psi} \gamma^\mu \partial_\mu \psi + a_3 \bar{\psi} \gamma^\mu \psi$$

$$+ a_4 \bar{\psi} \sigma^{\mu\nu} F_{\mu\nu} \psi + a_5 F^{\mu\nu} F_\nu^\lambda F_{\lambda\mu} + \dots$$

- These are all the bilinear covariants

• we restrict them by imposing Lorentz/Gauge invariance

- looking at field strength tensor:  $F^{\mu\nu} = 2^\mu A^\nu - 2^\nu A^\mu$ .

$$\rightarrow \bar{F}^{\mu\nu} F_{\mu\nu} \propto [\partial^\mu A^\nu \partial_\mu A_\nu] = 0, \text{ then } [A_\mu] = D-2.$$

since  $[2^\mu] = 1$ , so

$$[A_\mu] = \frac{D-2}{2}.$$

- now we can look at  $\bar{\psi} \not{A} \psi \quad [\bar{\psi} \gamma^\mu \partial_\mu \psi] = 0$

$$\gamma \text{ unitless}, [D_\mu] = 1 \rightarrow [\gamma^2][2] = 0$$

$$\rightarrow [\gamma] = \frac{D-1}{2}$$

$$\text{now, } [g_s] + 2[\gamma] + [A_\mu] = 0$$

$$[g_s] + (D-1) + \frac{D-2}{2} = 0$$

$$[g_s] = 2 - \frac{D}{2}$$

- our "constant" now dependent on dimension
- related to factorization/renormalization scale  $\mu$

$$\text{let } g_s = g_0 \mu^{2-D/2}$$

$\uparrow$   
dimensionless, now

Now, we can introduce a new constant  $\epsilon$ , and let

$$D = 4 - 2\epsilon,$$

$$g_0 \epsilon = \frac{1-0}{2} = 2^{-D/2}, \quad g_s = g_0 \mu^\epsilon$$

In our self-energy:

$$\Sigma(p) = \frac{g_0^2 c_F}{(4\pi)^2} p \left( \frac{-p^2}{4\pi m^2} \right)^{-\epsilon} (1-\epsilon) \beta(1-\epsilon, 1-\epsilon) \Gamma(\epsilon)$$

now we perform a Laurent expansion (basically a Taylor expansion for complex)

$$\Sigma(p) = \frac{g_0^2}{(4\pi)^2} c_F p \left( \frac{1}{\epsilon} - \gamma_E + 1 - \ln \left( \frac{-p^2}{4\pi m^2} \right) \right) + O(\epsilon),$$

$$\gamma_E = \text{Euler-Mascheroni constant} = 0.57721$$

$$\text{where we used } \Gamma(\epsilon) = \frac{1}{\epsilon} - \gamma_E + O(\epsilon),$$

$$\text{and we can approximate } (1-\epsilon) \beta(1-\epsilon, 1-\epsilon) = \Delta$$

$$\text{- we have now mapped the singularities as poles } \left( \frac{1}{\epsilon} \right)$$

- One important note: we didn't use  $\gamma^5$ ... in D-dim,  $\gamma^5$  is weird and requires a special treatment (it cannot be defined uniquely explicitly)

- Another note: DR conventions

1) D-dim spacetime metric is always mostly minus

2)  $\text{Tr}[1] = 4$  (in space of  $\gamma^5$ 's)

3)  $\int \frac{d^D k}{(2\pi)^D}$  defines the integral "measure"

4) only thing we can say now is that  $\gamma^5$  is an object that satisfies  $\{\gamma^5, \gamma^\mu\} = 0$

### Fermion self-energy correction in an arbitrary covariant gauge

- we chose  $\alpha=1$  before; going back to before this step:

$$\Sigma(p) = g_s^2 C_F \int \frac{d^D k}{(2\pi)^D} \frac{1}{i} \frac{1}{k^2 (k-p)^2} \times \left\{ \gamma_\mu (k-p) \gamma^\mu - (1-\alpha) \underbrace{\frac{k(k-p)k}{k^2}}_{\text{calculated this already}} \right\}$$

let's call  $\tilde{\Sigma}(p) = \Sigma_1(p) - (1-\alpha) \Sigma_2(p)$ , where  $\Sigma_1(p)$  is what we already found;

$$\Sigma_1(p) = g_s^2 C_F \int \frac{d^D k}{(2\pi)^D} \frac{k(k-p)k}{i(k^2)^2 (k-p)^2}$$

- now, we need a more general case for Feynman parameterization:

$$\frac{1}{AB^2} = 2 \int_0^1 dx \frac{(1-x)}{\{xA + (1-x)B\}^3}, \quad B = k^2, \quad A = (k-p)^2$$

- we choose  $L = -x(1-x)p^2$  like before

suitable

- At this point we massage denominator to introduce a shift in  $k$ :

$$k \rightarrow k' = k - xp$$

- doing this, we can show that

$$\frac{1}{[x(k-p)^2 + (1-x)k^2]^3} \rightarrow \frac{1}{[-(k-xp)^2 - x(1-x)p^2]^2}$$

$\overbrace{k}^T$        $\overbrace{L}^U$

$$\text{so, } \Sigma_2(p) = 2g_s^2 C_F \int_0^1 dx (1-x) \int \frac{d^D k}{(2\pi)^D} \frac{1}{i} \frac{k(k-p)k}{[-(k-xp)^2 + L]^3}$$

then, doing  $k \rightarrow k'$ :

$$= 2g_s^2 C_F \int_0^1 dx (1-x) \int \frac{d^D k'}{(2\pi)^D} \underbrace{\frac{(1-x) k' \not{p} k' - 2x k'^2 \not{p} - x L \not{p}}{[-k'^2 + L]^3}}_{\sim k'^2}$$

- now, recall  $g^{\mu\nu} g_{\mu\nu} = 0$ ; we can isolate  $\not{p}$ , we have something like  $k' \not{p} k' \sim p^\mu k'^\mu k'^\nu$ . then our integral is of the form

$$\sim \int d^D k' k'_m k'_v \cdot f(k'^2)$$

$$\sim \frac{D}{D} \int \sim \rightarrow \frac{g^{\mu\nu} g_{\mu\nu}}{D} \int d^D k' k'_m k'_v$$

$$\rightarrow \frac{g^{\mu\nu}}{D} \int d^D k' k'^2 f(k'^2)$$

So,

$$\Sigma_2(p) = 2g_s^2 C_F p \int_0^1 dx (1-x) \int \frac{d^D k'}{(2\pi)^D} i \frac{1}{[-k'^2 + L]^3} \left\{ \left( \frac{2(1-x)}{D} - 1-x \right) k'^2 - x L \right\}$$

want to again perform Wick rotation to get to Euclidean space:

$$\begin{aligned} \Sigma_2(p) &= 2g_s^2 C_F p \int_0^1 dx (1-x) \int \frac{d^D k}{(2\pi)^D} \\ &\times \left\{ \frac{1}{[k^2 + L]^3} \left( \frac{2(1-x)}{D} - 1-2x \right) - \frac{1}{[k^2 + L]^2} (2(1-x) - 1-x) \right\} \end{aligned}$$

then we can use the generalization of a previous result:

$$\int \frac{d^D k}{(2\pi)^D} \frac{1}{[k^2 + L]^a} = \frac{\Gamma(a - D/2)}{(4\pi)^{D/2} \Gamma(a)} \cdot L^{D/2 - a},$$

where  $a \in \mathbb{C}$ ,  $\operatorname{Re}[a] > 0$

- we also need 3 properties:

$$\beta(p, q) = \frac{\Gamma(p) \Gamma(q)}{\Gamma(p+q)}$$

$$\Gamma(z+1) = z \Gamma(z) \quad \text{for some } z \in \mathbb{C}$$

$$B(p, q) = \int_0^1 dx \cdot x^{p-1} (1-x)^{q-1}$$

- w/ this, we can show

$$\Sigma_2(p) = \frac{2 g_s^2 C_F}{(4\pi)^{D/2}} \rho(-p^2)^{D/2-2} (D-1) B\left(\frac{D}{2}, \frac{D}{2}\right) \Gamma\left(2 - \frac{D}{2}\right)$$

- same as  $\Sigma_1$ !

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- w/ this, we have calculated  $\Sigma(p)$ :

$$\Sigma(p) = \frac{2 \alpha C_F g_s^2}{(4\pi)^{D/2}} \rho(-p^2)^{D/2-2} (D-1) B\left(\frac{D}{2}, \frac{D}{2}\right) \Gamma\left(z - \frac{D}{2}\right)$$

now, Laurent Expansion in  $\varepsilon$ :

$$\Sigma(p) = \frac{\alpha g_s^2}{(4\pi)^2} C_F p \left[ \frac{1}{\varepsilon} - \gamma_E + 1 - \ln\left(-\frac{p^2}{4\pi m^2}\right) \right] + O(\varepsilon)$$

- this was regularization; now to Renormalization

### Renormalizational Schemes (preliminaries)

Def: redefinition of mass/coupling constant together w/ a readjustment of the normalization of the Green's function

- renormalization is not unique
- divergent pieces in the Green's fn's cannot be uniquely defined
- there is an ambiguity in the final piece of the Green's fn.
- Q: How to remove this ambiguity?
- specify how the divergent piece is defined so that we are able to systematically subtract it.
- this particular subtraction procedure is called a renormalization scheme.
- different schemes are related by a finite renormalization

### Example:

- let's consider  $\Sigma(p)$  we just calculated.

$$\Sigma(p) = \frac{\alpha g_s^2}{(4\pi)^2} C_F \left[ \frac{1}{\varepsilon} - \gamma_E + 1 - \ln\left(\frac{-p^2}{4\pi\mu^2}\right) \right] + O(\varepsilon)$$

if we substitute this into

$$\tilde{S}_{ij}(p) = \frac{-\delta_{ij}}{p + \Sigma(p)}$$

we first define:

$$\sigma(p^2) = \frac{\alpha g_{os}^2}{(4\pi)^2} C_F \left[ \frac{1}{\varepsilon} - \gamma_E + 1 - \ln\left(\frac{-p^2}{4\pi\mu^2}\right) \right] + O(g_{os}^4) + O(\varepsilon)$$

$$\rightarrow S_{ij}(p) = \frac{-\delta_{ij}}{p} \frac{1}{1 + \sigma(p^2)}$$

now,  $\tilde{S}_{ij}$  has a pole at  $p=0$

also, our massless quark prescription will remain; it does not get changed by our renormalization

this will be the case for all orders.

we will use a multiplicative factor called  $Z_2$  (historical reasons  $\rightarrow$  it's the "quark field renormalization constant")

Our renormalized propagator is:

$$\tilde{S}_{R,ij}(p) = Z_2^{-1} \tilde{S}_{ij}(p)$$

$Z_2$  is a fn of  $g_0 \rightarrow$  can expand:

$$Z_2 = 1 - z_2 + O(g_{os}^4),$$

where (lowercase)  $z_2$  contains a  $g_{os}^2$  and divergent piece.

Substituting:

$$\tilde{S}_{R,ij}(p) = \frac{-\delta_{ij}}{p} \cdot \frac{1}{1 + \sigma(p^2) - z_2}$$

now, we should also renormalize  $g_{os}$ , but it turns out, it will have no effect at this perturbative order.

now:  $\tilde{S}_{R,ij}(p)$  should be free of divergence, so  $\sigma(p^2) - z_2$  must be finite and the divergences in  $\sigma(p^2)$  must be canceled by those in  $z_2$ .

- we can now define  $z_2$  to fulfill this. Of course, then,  $z_2$  can be determined up to a finite constant, hence why we said that it's not unique.
- so, we need some other procedure/requirement to set up the scheme.
- the first of which we will examine is called the:

### On-shell Subtraction

- $z_2$  is determined on the mass shell of quarks by imposing the condition

$$\tilde{S}_{R,ii}(p) \sim \frac{\delta_{ij}}{m-p} \text{ for } p \sim m$$

for our case w/  $m=0$ , then  $z_2 = \sigma(0)$ . but  $\sigma(0)$  is not well defined, so we don't use it.

### Off-shell Subtraction

- at some unphysical scale (hence off-shell) of  $p^2$  say  $p^2 = -\lambda^2$  with  $-\lambda^2 > 0$  and we require that  $\tilde{S}_{R,ii}(p)$  is of the form of the free massless propagator:

$$\tilde{S}_{R,ii}(p) \sim -\frac{\delta_{ij}}{p} \quad \text{for } p^2 \sim -\lambda^2$$

- this determines  $z_2$  such that

$$z_2 = \sigma(-\lambda^2) = \frac{\alpha g_{0s}^2}{(4\pi)^2} C_F \left[ \frac{1}{\epsilon} - \gamma_E + 1 - \ln\left(\frac{\lambda^2}{4\pi m^2}\right) \right]$$

so that

$$\tilde{S}_{R,ii}(p) = -\frac{\delta_{ij}}{p} \left[ 1 - \frac{\alpha g_{0s}^2}{(4\pi)^2} C_F \ln\left(\frac{-p^2}{\lambda^2}\right) \right]^{-1}$$

- $\epsilon$  pole vanishes ( $1/\epsilon$  is reabsorbed)!

- This scheme is also called "momentum-space subtraction scheme"

### Minimal Subtraction Scheme ('t Hooft)

- very specific to DR  $\rightarrow$  only eliminate the  $\frac{1}{\epsilon}$  pole that makes  $z_2$  very economical; often used in QCD (a modified ver.)
- Requirement imposes  $z_2 \propto \frac{g_{0s}^2}{(4\pi)^2} C_F \frac{1}{\epsilon}$
- therefore,

$$\tilde{S}_{\text{R},ii}(\rho) = -\frac{\delta_{ij}}{\rho} \left\{ 1 - \frac{\alpha g_{os}^2}{(4\pi)^2} F \left[ \gamma_E - 1 + \ln \left( \frac{-\rho^2}{4\pi\mu^2} \right) \right] \right\}^{-1}$$

- advantage is a simpler expression for  $\tilde{Z}_2$ , but leads to more complicated Green's fn.
- can be converted to off-shell subtraction by

$$\gamma^2 = 4\pi e^{1-\gamma_E} \mu^2$$

Standard: Modified Minimal Subtraction Scheme ( $\overline{\text{MS}}$ ; "ms bar")  
from normal ms

- the  $\gamma_E - 1 + \ln(4\pi)$  shows up everywhere and is very annoying!
- we want to get rid of it.

$$\rightarrow \tilde{Z}_2 = \frac{\alpha g_{os}^2}{(4\pi)^2} F \left( \frac{1}{\epsilon} - \gamma_E + \ln(4\pi) \right)$$

then,

$$\tilde{S}_{\text{R},ii}(\rho) = -\frac{\delta_{ij}}{\rho} \left\{ 1 - \frac{\alpha g_{os}^2}{(4\pi)^2} F \left[ -1 + \ln \left( \frac{-\rho^2}{m^2} \right) \right] \right\}$$

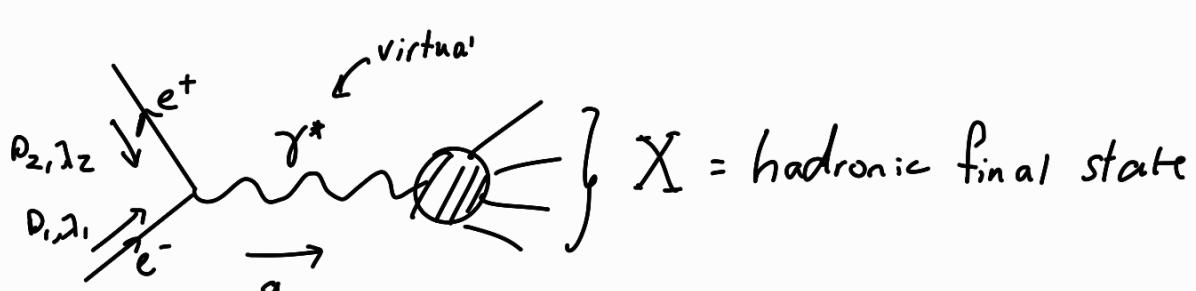
Real quick: general Feynman Parametrization

$$\prod_{i=1}^n \frac{1}{A_i^{\alpha_i}} = \frac{\Gamma(\alpha)}{\prod_{i=1}^n \Gamma(\alpha_i)} \int_0^1 \left( \prod_{i=1}^n dx_i x_i^{\alpha_i-1} \right) \frac{\delta(1-\lambda)}{\left( \sum_{i=1}^n x_i A_i \right)^\alpha}$$

$$\text{where } \alpha = \sum_{i=1}^n \alpha_i$$

$$x = \sum_{i=1}^n x_i$$

Electron/Positron Annihilation :  $e^- + e^+ \rightarrow X$



The Feynman amplitude:  $\langle \bar{X} | T | e^+ e^- \rangle$ ,  $T$  is similar to scattering matrix; it's a transition operator.

Don't know how to blob, but we can start w/ L.H.S. of diagram

$$-ie \bar{V}^{(\lambda_2)}(p_2) \gamma^\mu u^{(\lambda_1)}(p_1) \cdot \frac{-i}{q^2} \cdot \langle \bar{X} | (-ie) J_\mu(o) | 0 \rangle ,$$

$J_\mu(x)$  is some em current w/ quarks... but why a current?

Recall:  $\mathcal{L}_{QED} = (-e \bar{\psi} \gamma^\mu \psi + e j^\mu) A_\mu$

$\stackrel{T}{\uparrow}$                              $\uparrow$   
well behaved:  $e^+/e^-$                             this is the more "arbitrary" part

- we have to square the amplitude. To do this, we need 2 properties:

• completeness:  $\sum_x |x\rangle \langle x| = \mathbb{1}$

• translational invariance:  $J_\mu(x) = e^{i \hat{p} \cdot x} J_\mu(0) e^{-i \hat{p} \cdot x}$

$\hat{p}$  has  $\hat{p}^\mu |x\rangle = p_x^\mu |x\rangle$ , and

$$e^{i \hat{p} \cdot x} |x\rangle = e^{i p_x \cdot x} |x\rangle$$

• now, generic cross section formula:

$$\sigma = \frac{1}{2s} \cdot \frac{1}{4} \sum_{\lambda_1, \lambda_2} \sum_x (2m)^4 \delta^4(p_x - q) |\langle \bar{X} | \hat{T} | e^+ e^- \rangle|^2$$

(generally) =  $\frac{1}{k(s)} \cdot \frac{1}{(2J_1 + 1)(2J_2 + 1)} \sum_{\lambda_1, \lambda_2, m_1, m_2 \dots}$

$$\times \int \frac{d^3 k_i}{(2\pi)^3 2k_i^0} (2m)^4 \delta^4 \left( \sum_{j=1}^n k_j - p_1 - p_2 \right) r \sqrt{\langle k_1 m_1, k_2 m_2 \dots | \hat{T} | p_1 \lambda_1, p_2 \lambda_2, \dots \rangle^2},$$

where  $k(s)$  is "flux factor":  $k(s) = \sqrt{[s - (m_1 + m_2)^2][s - (m_1 - m_2)^2]}$

- we use the original one; substituting our exp for  $\langle \bar{X} | \hat{T} | e^+ e^- \rangle$ , we get

$$\sigma = \frac{e^4}{2s^3} \lambda^{\mu\nu} W_{\mu\nu},$$

$$\lambda^{\mu\nu} = p_1^\mu p_2^\nu + p_1^\nu p_2^\mu - \frac{q^2}{2} g^{\mu\nu}; \text{ leptonic tensor}$$

$$W^{\mu\nu} = \sum_x (2m)^4 \delta^4(p_x - q) \langle 0 | J_\mu(0) | x \rangle \langle x | J_\nu(0) | 0 \rangle; \text{ hadronic tensor}$$

$$= \sum_x \int e^{-i(p_x - q) \cdot x} \langle 0 | J_\mu(0) | x \rangle \langle x | J_\nu(0) | 0 \rangle d^4 x$$

exponential is just a number, so we can move around

$$\begin{aligned}
 &= \sum_x \int e^{iq \cdot x} e^{-ip_x \cdot x} e^{i0x} \sim \\
 &= \sum_x \int e^{iq \cdot x} \langle 0 | e^{i0x} J_m(0) e^{-ip_x \cdot x} | x \rangle \langle x | J_n(0) | 0 \rangle d^4x \\
 &= \sum_x \int e^{iq \cdot x} \langle 0 | J_m(x) | x \rangle \langle x | J_n(0) | 0 \rangle d^4x \\
 &\quad \text{↑ completeness}
 \end{aligned}$$

$$\rightarrow = \sum_x \int e^{iq \cdot x} \langle 0 | J_m(x) J_n(0) | 0 \rangle d^4x$$

we can show that  $\int e^{iq \cdot x} \langle 0 | J_n(0) J_m(x) | 0 \rangle d^4x = 0$ , then

$$W^{mn} = \sum_x \int e^{iq \cdot x} \langle 0 | [J_m(x), J_n(0)] | 0 \rangle$$

10/9  
 $\star$  in general: we have something like  
 $\langle p | \underbrace{e^{ip \cdot x} J_m(0)}_{J_m(x)} e^{-ip_x \cdot x} | x \rangle$ ,  
but here  $p=0$ , so  
 $\langle 0 | \underbrace{J_m(0)}_{J_m(x)} e^{-ip_y \cdot y} | x \rangle$

we will now show  $\int e^{iq \cdot x} \langle 0 | J_n(0) J_m(x) | 0 \rangle d^4x = 0$

- let's reinsert:  $\sum_x \int e^{iq \cdot x} \langle 0 | J_n(0) | x \rangle \langle x | J_m(x) | 0 \rangle d^4x$

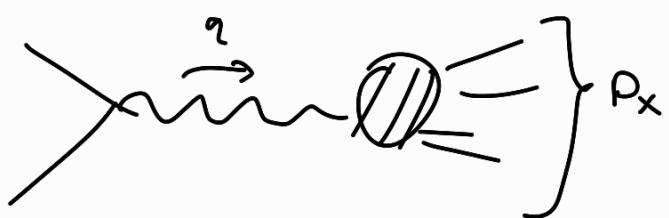
- now,  $\langle x | J_m(x) | 0 \rangle = \langle x | e^{\hat{i}\vec{p}_x \cdot x} J_m(0) | 0 \rangle$ , and

$$\langle x | e^{\hat{i}\vec{p}_x \cdot x} = \langle x | e^{i\vec{p}_y \cdot x} \text{ (a number, so we move)}$$

$$\rightarrow \sum_x \int e^{i(q+q_y) \cdot x} \langle 0 | J_n(0) | x \rangle \langle x | J_m(0) | 0 \rangle d^4x$$

$$= \sum_x (2\pi)^4 \delta^4(q + q_y) \langle 0 | J_n(0) J_m(0) | 0 \rangle$$

Recall:



- we choose our reference frame such that  $q^\mu = (q^0, \vec{q})$  (C.O.M. frame)
- now,  $q^2 = (p_1^2 + p_2^2) = 2p_1 \cdot p_2 = s > 0$
- in this frame,  $p_1^\mu = (p_1^0, \vec{p}_1, p_2)$ ,  $p_2^\mu = (p_2^0, \vec{p}_2, -p_2)$ ,

so  $q^m = (p_1^0 + p_2^0, \vec{0})$ , so  $q^2 \approx 2p_1^0 p_2^0 = 0 \dots$  so both

$p_x$  must be +, but delta function doesn't encompass this, so it's zero.

- so:  $\omega_{\mu\nu} = \int e^{iq \cdot x} \langle 0 | [J_\mu(x), J_\nu(0)] | 0 \rangle d^4x$

- in high-energy regime,  $q^0 \rightarrow \infty$ , but this leads to wild oscillations

- so only small values of  $x$  are allowed to keep this well behaved.

- since the commutator (operator  $J_\mu(x)$ ) is a fn of  $x$ , we can now expand it perturbatively

- in order to satisfy causality,  $[J_\mu(x), J_\nu(0)] = 0$  for  $x^2 < 0$ , so we must have  $x^2 \geq 0$ .

- now we use Lorentz-invariance/current-conservation to parameterize  $\omega_{\mu\nu}(q)$  into all second rank tensors involving  $q$ :

$$\omega_{\mu\nu} = A q_\mu q_\nu + B q^2 g_{\mu\nu}$$

applying these invariance/conservations:

$$\omega_{\mu\nu} = (q_\mu q_\nu - q^2 g_{\mu\nu}) \frac{1}{2s^3} \ell^{mn} W(q^2)$$

back to our cross sec:  $\sigma = \frac{e^4}{2s^3} \ell^{mn} \omega_{\mu\nu} :$

$$= \frac{4\pi\alpha_{em}^2}{3s} W(s)$$

For  $e^- + e^+ \rightarrow \mu^- + \mu^+$ , we had

$$\sigma_{nn} = \frac{4\pi\alpha_{em}^2}{3s},$$

no  $W(q^2)$ , because no inner structure, just a vertex.

We can take ratios to cancel the prefactor:

$$R = \frac{\sigma}{\sigma_{nn}} = W(s),$$

$R \equiv$  Prell Ratio

- we can write  $R$  this way:

$$R = -\frac{2\pi}{q^2} \int e^{iq \cdot x} \langle 0 | [J_\mu(x), J_\nu(0)] | 0 \rangle$$

•  $R$  is a fn of  $s$ , the coupling  $g$  (generically), and some "renormalization scale"  $\mu$  which is associated w/ where we cut our perturbative series off:  $R(s, g, \mu^2) = R\left(\frac{s}{\mu^2}, g\right)$ .

- according to renormalization group theory,  $R$  must satisfy

$$\left[ \mu \frac{\partial}{\partial \mu} + B(g) \frac{\partial}{\partial g} \right] R\left(\frac{s}{\mu^2}, g\right) = 0$$

- this total derivative being zero indicates a desirable stability for  $R$

- the general solution of this is

$$R\left(\frac{s}{\mu^2}, g\right) = R(1, \bar{g}(s)),$$

where we group energy/scale into coupling using  $B$ -fn;

$$\frac{d\bar{g}}{dt} = B(g); \quad \bar{g}(\ln(s/\mu^2)) = g, \quad \text{where}$$

$$t = \frac{1}{2} \ln\left(\frac{s}{\mu^2}\right)$$

now,

$$R\left(\frac{s}{\mu^2}, g\right) = \sum_i Q_i^2 \left[ 1 + a\left(\frac{s}{\mu^2}\right) g^2 + b\left(\frac{s}{\mu^2}\right) g^4 + \dots \right]$$

$\in$  all quark flavors/gens

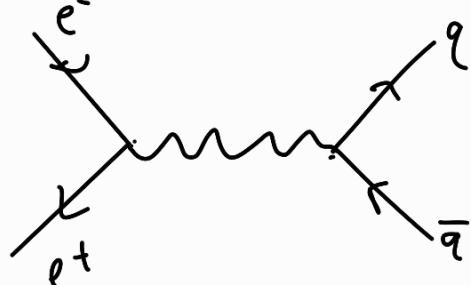
where the coefficients have  $\ln(s/\mu^2)$  dependence... bad!

the new one:

$$R(1, \bar{g}(s)) = \sum_i Q_i^2 \left[ 1 + a(1) \bar{g}(s)^2 + b(1) \bar{g}(s)^4 + \dots \right]$$

now, coefficients are good, and pushing  $s$  up,  $\bar{g}$  goes smaller!

- Born-level :



$$\sigma_B = \frac{4\pi \alpha_{em}^2}{s}$$

- calculating the virtual NLO diagrams gives a term like

$$\sigma_B(m), \text{ which when added w/}$$

$Z^2 \sigma_B$  from our renormalization cancels the divergences

$$F_V = \frac{1}{2\pi} \int \frac{d^3 \vec{k}_1}{(2\pi)^3 2k_{10}} \frac{d^3 \vec{k}_2}{(2\pi)^3 2k_{20}} (2\pi)^4 \delta^4(k_1 + k_2 - p_1 - p_2) F_V,$$

where

$$F_V = \left( \sum_i Q_i \right)^2 \frac{e^2}{q^4} \text{Tr}[p_2 \gamma^\mu p_1 \gamma^\nu] \text{Tr}[k_1 \Lambda_m k_2 \gamma_V],$$

with  $\Lambda_m$  being the vertex correction at 1-loop

- w/ our  $m=0$  prescription, we get IR divergences with  $\frac{1}{p^2}$  stuff.

• we go back to

$$\sum(p) = g_s^2 C_F (D-2) \int_0^1 dx \int \frac{d^D k'}{(2\pi)^D i} \frac{-(1-x)k'}{\sum [k'^2 + x(1-x)p^2]^2}$$

for  $p^2 \rightarrow 0$  ( $m \rightarrow 0$ ),  $k' \neq 0$ :

$$= g_s^2 C_F (2-D) \underbrace{\int_0^1 dx (1-x)}_{\frac{1}{2}} \int \frac{d^D k'}{(2\pi)^D i} \frac{1}{(k')^4}$$

- turns out  $k'$  bit is zero

$$-\text{in context of DR: } \int \frac{d^D k}{(-k^2)^\alpha} = 0 \quad \text{where } \alpha > 0$$

• doing Wick rotation  $\rightarrow -k^2 = \bar{k}^2$ ,

• then doing angular integration

$$\int \frac{d^D \bar{k}}{(-\bar{k}^2)^\alpha} = \frac{i\pi^{D/2}}{\Gamma(D/2)} \int_0^\infty (\bar{k}^2)^{D/2-\alpha-1} d\bar{k}^2$$

- we now recognize if  $D > 2\alpha$ , we get a UV divergence,

if  $D < 2\alpha$ , we get an IR divergence

- what we will do is split the integral into  $[0, 1]$ ,  $[1, \infty]$ , s.t. we can integrate in both ranges more nicely, then do some complex analysis stuff at 1.

• here,  $\bar{k}^2 > 1^2 \rightarrow \text{UV} ; \bar{k}^2 < 1^2 \rightarrow \text{IR}$

$$S_{01} = \frac{i\pi^{D/2}}{\Gamma(D/2)} \left[ \int_0^{\Lambda^2} (\bar{k}^2)^{D/2-\alpha-1} d\bar{k}^2 + \int_{\Lambda^2}^{\infty} (\bar{k})^{D/2-\alpha-1} d\bar{k}^2 \right]$$

↑  
convergent  
for  $D > 2\alpha$

con. for  $D < 2\alpha$

- we will call first  $D_i$  for IR, second  $D_u$  for UV

$$= i \frac{\pi^{D/2}}{\Gamma(D/2)} \left[ \frac{\Lambda^{D_i - 2\alpha}}{\frac{1}{2}D_i - \alpha} - \frac{\Lambda^{D_u - 2\alpha}}{\frac{1}{2}D_u - \alpha} \right], \text{ which is divergent if } D_i = D_u = 2\alpha$$

- now we use analytical continuation to move the domain into complex plane, such that the denominators are only divergent in real part  $\rightarrow$  we just need a small imaginary part

$$\rightarrow \frac{\Lambda^{D_i/2 - \alpha - 1}}{\frac{1}{2}D_i - \alpha + i\epsilon} - \frac{\Lambda^{D_u/2 - \alpha - 1}}{\frac{1}{2}D_u - \alpha + i\epsilon}$$

- now we can do the  $D_i = D_u \rightarrow 2\alpha$  so that they no longer diverge; instead they cancel

- thus, w/ analytical continuation, we have shown that in the context of DZ,

$$\int \frac{d^D k}{(-k^2)^\alpha} = 0.$$

Back to

$$\zeta(p) \Big|_{p^2=0} = g_s^2 C_F (D-2) \not{p} \cdot \frac{1}{2} \int \frac{d^D k'}{(2\pi)^D} \frac{1}{(-k^2)^2}$$

$$= \frac{g_s^2}{(4\pi)^2} C_F (D-2) \not{p} \left[ \frac{1}{\epsilon} - \frac{1}{\epsilon} \right]$$

- we keep them separate for now to distinguish their origin  
they conspire later

- for vertex correction:

$$\Lambda_m = \gamma_m \frac{g_{os}^2}{8\pi^2} C_F \left( \frac{4m^2}{-q^2} \right)^\varepsilon$$

$$\times \Gamma(1+\varepsilon) B(-\varepsilon, 2-\varepsilon) \left[ \frac{1}{\varepsilon} - \frac{2}{\varepsilon^2} - 2 \right],$$

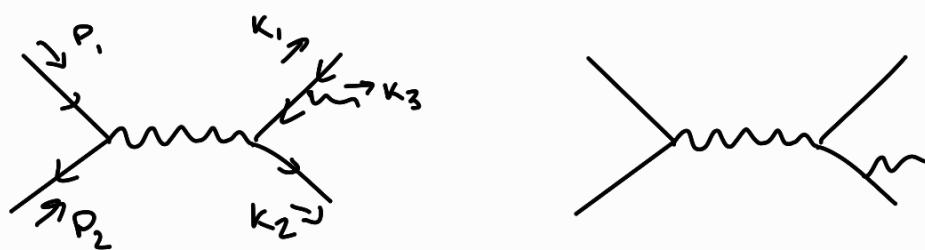
$m^2$  is the scale to make  $g$  dimensionless

$$\tilde{\sigma}_v = \sigma_v + (z_2^2 - 1)\bar{\sigma}_\theta$$

$$= A_v \bar{\sigma}_\theta, \quad \text{where}$$

$$A_v = \frac{\alpha_s}{\pi} C_F \left( \frac{4\pi m^2}{s} \right)^\varepsilon \frac{\cos(\pi\varepsilon)}{\Gamma(1-\varepsilon)} \left[ -\frac{1}{\varepsilon^2} - \frac{3}{2\varepsilon} - 4 + O(\varepsilon) \right]$$

### Real emission corrections



Fermi Golden Rule:

$$\tilde{\sigma}_R = \frac{1}{8s} \int \prod_{i=1}^3 \frac{d^{D-1}k_i}{(2\pi)^{D-1}} \frac{1}{2k_{i,0}} (2\pi)^D \delta^{(D)} \left( \sum_{i=1}^3 k_i - p_1 - p_2 \right) F_R,$$

$$F_R = - \sum_i Q_i^2 \frac{e^2}{q^2} g_s^2 \langle \Gamma \text{Tr} [ \gamma_2 \gamma^m \gamma_\nu \gamma^v ] \rangle^L, \text{ leptonic part,} \\ \times \text{Tr} [ \gamma_1 S_{\lambda M} \gamma_2 S_\nu^2 ] \sim G_{\mu\nu}, \text{ "quark" part,}$$

$$S_{\mu\nu} = \gamma_m \cdot \frac{-1}{k_1 + k_3} \gamma_\nu + \gamma_\nu \cdot \frac{1}{k_2 + k_3} \gamma_m$$

- leptonic part:

$$L^{\mu\nu} = 4 \left[ P_1^\mu P_2^\nu + P_1^\nu P_2^\mu - \frac{q^2}{2} g^{\mu\nu} \right] \text{ (after trace)}$$

- can collect phase space parts together!

$$I_{\mu\nu} = \int \prod_{i=1}^3 \frac{d^D k_i}{2k_{i,0}} \delta^{(D)}\left(\sum_{i=1}^3 k_i - q\right) G_{\mu\nu}$$

$$\sigma_R = \frac{-e^4 g_s^2}{8s(2\pi)^{2D-3}} \cdot \frac{1}{q^4} \sum_i Q_i^2 \mathcal{L}^{\mu\nu} I_{\mu\nu}$$

- Now,  $I_{\mu\nu} = I_{\mu\nu}(a)$ , and it must satisfy current conservation
- We know  $\partial_\mu J^\mu = 0$ . Here,  $J^\mu \rightarrow I_{\mu\nu}$ , since that contains our current information
- The conjugate for  $Q_\mu$  is  $q^\mu$  in position space, so we must have

$$q^\mu I_{\mu\nu} = 0$$

We can parametrize  $I_{\mu\nu}$  like so:

$$I_{\mu\nu}(a) = F(a) \left[ \frac{q_\mu q_\nu}{q^2} - g_{\mu\nu} \right]$$

$$\text{Inversely, then, } F(a) = \frac{-g^{\mu\nu} I_{\mu\nu}}{D-1}$$

$$\text{Now, } \mathcal{L}^{\mu\nu} I_{\mu\nu} = \frac{(D-2)}{(D-1)} q^2 g^{\mu\nu} I_{\mu\nu},$$

or, we can have

$$g^{\mu\nu} G_{\mu\nu} = -8(1-\epsilon) \frac{x_1^2 + x_2^2 - \epsilon x_3^2}{(1-x_1)(1-x_2)},$$

$$\text{where } \epsilon = \frac{4-D}{2} \text{ and } x_i = \frac{2k_i \cdot q}{q^2}$$

With all this,

$$g^{\mu\nu} I_{\mu\nu} = \frac{\pi(s\pi)^{1-2\epsilon}}{4\Gamma(2-2\epsilon)} \int_0^1 \prod_{i=1}^3 (1-x_i)^{-\epsilon} dx_i \delta\left(2 - \sum_{i=1}^3 x_i\right) g^{\mu\nu} G_{\mu\nu}$$

*side hint:*  $q^2 = q \cdot (k_1 + k_2 + k_3) \rightarrow$  can see  $x_i$  from here

Thus,

$$\sigma_R = \left( \sum_i Q_i \right) \alpha_m^2 \alpha_s^2 C_F \frac{2}{3} \left( \frac{4\pi\mu}{s} \right)^{2\epsilon} \frac{(1-\epsilon)^2}{(3-2\epsilon)\Gamma(2-2\epsilon)} K,$$

$$\text{where } K = \int_0^1 \prod_{i=1}^3 (1-x_i)^{-\epsilon} dx_i \delta\left(2 - \sum_{i=1}^3 x_i\right) \frac{x_1^2 + x_2^2 - \epsilon x_3^2}{(1-x_1)(1-x_2)}$$

- We do the integral:

$$K = \left( \frac{4}{\varepsilon^2} - \frac{12}{\varepsilon} + 10 - 4C \right) B(1-\varepsilon, 2-2\varepsilon) \Gamma(1-\varepsilon, 1-\varepsilon) + O(\varepsilon)$$

What we want to do is have  $\sigma_R = A_R \sigma_B$

thus, we need  $\sigma_B$  in D-im for consistency:

$$\rightarrow \sigma_B = \frac{4\pi \alpha_s^2}{3s} \left( \sum_i Q_i^2 \right) \left( \frac{4\pi}{s} \right)^\varepsilon \frac{3(1-\varepsilon)\Gamma(2-\varepsilon)}{(3-2\varepsilon)\Gamma(2-2\varepsilon)}$$

- Now we can just do a ratio to pull out  $\sigma_B$ :  $\frac{\sigma_R}{\sigma_B} = A_R$

$$A_R = \frac{\alpha_s}{\pi} C_F \left( \frac{4\pi \mu^2}{s} \right)^\varepsilon \frac{\cos(\pi\varepsilon)}{\Gamma(1-\varepsilon)} \left[ \frac{1}{\varepsilon^2} + \frac{3}{2\varepsilon} + \frac{19}{4} + O(\varepsilon) \right]$$

$$\text{Now, } r = (1 + A_U + A_R) \sigma_B$$

Recalling Ar:

$$A_R = \frac{\alpha_s}{\pi} C_F \left( \frac{4\pi \mu^2}{s} \right)^\varepsilon \frac{\cos(\pi\varepsilon)}{\Gamma(1-\varepsilon)} \left[ -\frac{1}{\varepsilon^2} - \frac{3}{2\varepsilon} - 4 + O(\varepsilon) \right]$$

- Our poles cancel!

$$\sigma = \left[ 1 + \frac{3}{4} C_F \frac{\alpha_s}{\pi} \right] \sigma_B$$

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