If we neglect the wass of the quarks

$$\nabla_{V} = \frac{1}{85} \int \frac{d^{3}k_{1}}{(2\pi)^{3}k_{10}} \frac{d^{3}k_{2}}{(2\pi)^{3}2k_{20}} (2\pi)^{4} \delta^{(4)}_{(k_{1}+k_{2}-k_{1}-k_{2})} F_{V}$$

Ly 1-loop vertex

* We can use the Feynman gauge for this calculation

Mq =0 >> mass singularity >> infrared divergences

Let's review the vure normalited 1-loop self-energy contrabilition for mg=0 in the Freynman gauge

$$\frac{\mathcal{L}(p)}{\mathcal{L}(p)} = \frac{1}{9} \left(\frac{1}{9} \operatorname{cos} \left(\frac{1}{9} \operatorname{cos$$

More general case in DR

$$\int \frac{d^{3}k}{(-k^{2})^{x}} = 0 \qquad x > 0$$

with a Wick Rotation are obtain

$$\int \frac{d^{2}k}{(-k^{2})^{\alpha}} = i \frac{\pi}{\Gamma(D/2)} \int_{0}^{+\infty} (k^{2})^{D/2 - \alpha - 1} dk^{2}$$

$$K^2 = -k^2$$

D>2x -> votraviolet d'vergence UV-div

D<2x -> Infrared divergence IR-div

. No matte metical maning full region in D.

We found in pravious cases that the integral made sense for D<3. We are going to use the same approach here and speit the integration posts in It in two parts:

$$\begin{cases} \mathbb{R}^2 > \Lambda^2 & \text{intraviolat part} \\ \mathbb{R}^2 < \Lambda^2 & \text{infraeed part} \end{cases}$$

$$\int \frac{d^{2}k}{(-k^{2})^{\alpha}} = i \frac{\pi^{3/2}}{T(0/2)} \left[\int_{0}^{(k^{2})^{3/2-\alpha-1}} dk^{2} + \int_{0}^{+\infty} (k^{2})^{3/2-\alpha-1} dk^{2} \right]$$

1st integreel convergent for D>2x 2nd integreel convergente for D<2x

D = regulator for IR and UV divergences.

We call $D_I = D$ in the 1st integral and $D_{I\!\!P} = D$ in the 2nd. Integrating, we obtain

$$\int \frac{d^{D} k}{(-k^{2})^{\alpha}} = i \frac{\sqrt{D_{1} - 2\alpha}}{\sqrt{2D_{1} - \alpha}} - \frac{\sqrt{D_{1} - 2\alpha}}{\sqrt{2D_{1} - \alpha}}$$

The two terms have poles at $D_I = D_U = 2x$

DI-2x

NDI-2x

1/2 DI-X

1/2 DU-X

Can be cantinved analitically and the

Constraints DI > 2x and Du < 2x

NDI-2x

1/2 DI-X+iE' - NDU-X+iE

Can be removed.

If DI and Du are identified after this analytic continuation, then when DI = Du on the right a hand side of the equation is zero. =>

$$\int \frac{d^{3}k}{(-k)^{\alpha}} = 0$$

This result is justified within DR.

Therefore, if we cook at $\Sigma(p)$ for $p^2 = 0$ $\Sigma(p) = g_{so}^2 G(D-2) \not = \int_0^1 dx (1-x) \int_0^1 d^3 k! \frac{1}{(2\pi)^2 i} \frac{1}{(2$ after we have removed ([dok ku f(ki) =0) the linear terms in ku,

 $\sum_{p=0}^{2} (p) = g_{0s}^{2} C_{F}(D-2) / \int_{(2\pi)^{0}i}^{d} \frac{d^{2}k}{k^{2}} d^{2}k$

This is like the case we analyzed before with $\alpha=2$

 $\sum_{i} (f) \Big|_{p^2 = 0} = \frac{9^2}{4\pi} \frac{G}{2} \frac{F}{E} + \frac{1}{E} \frac{1}{E} = \frac{1}{12} \frac{9^2}{4\pi} \frac{G}{2} \left(\frac{1}{E} - \frac{1}{E^2} \right)$

where & and & are equal to the (4-D)/2 param.

+We do not set E'= E here because we want to distinguish between the two divergences

*The jutegral for My can be treeated in the same

 $\Lambda_{\mu} = \chi_{\mu} \frac{g_{os}^{2}}{8\pi^{2}} G \left(\frac{4\pi\mu^{2}}{-q^{2}}\right)^{\epsilon} T(1+\epsilon) B(1-\epsilon, 2-\epsilon) \left(\frac{1}{\epsilon'} - \frac{2}{\epsilon^{2}} - 2\right)$

μ? - mess scale to make g dimension less.

9 = \$1 + \$2 The knowledge of the singular structure of 5(\$), allows us to determine the field remormalization const. Z2

53)

The remerkable task of renormalization is. accomplished in this case:

$$A_{V} = \frac{1}{\pi} G \left(\frac{4\pi \mu^{2}}{S} \right)^{\epsilon} \frac{\cos(\pi \epsilon)}{7(1-\epsilon)} \left[-\frac{1}{\epsilon^{2}} - \frac{3}{2\epsilon} - 4 + O(\epsilon) \right]$$

where
$$d_s = \frac{9^2}{4\pi}$$

Real emission corrections

$$T_{R} = \frac{1}{8s} \int_{1=1}^{3} \frac{d^{D-1}k_{i}}{(2\pi)^{D-1}2k_{i}} (2\pi)^{D} \delta^{(D)} \left(\sum_{i=1}^{3} k_{i} - k_{i} - k_{2} \right) T_{R}$$