

## Photon field quantization in Lorenz gauge

$$\partial_\mu A^\mu = 0, \pi^\mu = \frac{\partial \mathcal{L}}{\partial \dot{A}_\mu}$$

Try commutation relations  $[A^\mu(\vec{x}, t), \pi^r(\vec{y}, t)] = i g^{\mu r} \delta^3(\vec{x} - \vec{y})$   
(Covariant quantization)  $[A^\mu(\vec{x}, t), A^r(\vec{y}, t)] = 0, [\pi^\mu(\vec{x}, t), \pi^r(\vec{y}, t)] = 0$

However there is a problem:  $\pi^0$  still vanishes so the first equation cannot hold for  $A^0, \pi^0$

so we need to change the Lagrangian  $\rightarrow$  add a term that will still give Maxwell equations for Lorenz constraint  $\partial_\mu A^\mu = 0$

try  $\mathcal{L} = -\frac{1}{4} F^{\mu\nu} F_{\mu\nu} - \frac{1}{2} (\partial_\mu A^\mu)^2$  where  $-\frac{1}{2} (\partial_\mu A^\mu)^2$  is gauge-fixing term

$$\text{Then } \partial_\mu \left( \frac{\partial \mathcal{L}}{\partial (\partial_\mu A_r)} \right) = \frac{\partial \mathcal{L}}{\partial A_r} \Rightarrow \partial_\mu (-\partial^\mu A^r + \partial^r A^\mu - g^{\mu r} \partial_\rho A^\rho) = 0 \Rightarrow \partial_\mu \partial^\mu A^r = 0$$

Klein-Gordon eq.

A more general gauge-fixing term would be  $-\frac{\xi}{2} (\partial_\mu A^\mu)^2$

The choice  $\xi = 1$  is called the Feynman gauge.

Now  $\pi^0 = \frac{\partial \mathcal{L}}{\partial \dot{A}_0} = -\partial_\mu A^\mu$  still vanishes in Lorenz gauge

Resolution: treat the Lorenz condition not as an operator identity but as a vanishing expectation value for states:  $\langle \psi | \partial_\mu A^\mu | \psi \rangle = 0$

Then  $\pi^0 = -\partial_\mu A^\mu$  can satisfy commutation relations

Lorenz - gauge  $A^\mu(x) = \int \frac{d^3p}{(2\pi)^3 (2p^0)^{1/2}} \sum_{\lambda=0}^3 \left[ \epsilon^\mu_{(\lambda)}(p) a_{(\lambda)}(p) e^{-ip \cdot x} + \epsilon^{\mu*}_{(\lambda)}(p) a^\dagger_{(\lambda)}(p) e^{ip \cdot x} \right]$

Four independent solutions ( $\lambda=0,1,2,3$ ) for  $\epsilon^\mu$   
 In the frame  $p^\mu = (p, 0, 0, p)$  we choose as a basis

$$\epsilon^{(0)} = \begin{pmatrix} 1 \\ 0 \\ 0 \\ 0 \end{pmatrix}, \quad \epsilon^{(1)} = \begin{pmatrix} 0 \\ 1 \\ 0 \\ 0 \end{pmatrix}, \quad \epsilon^{(2)} = \begin{pmatrix} 0 \\ 0 \\ 1 \\ 0 \end{pmatrix}, \quad \epsilon^{(3)} = \begin{pmatrix} 0 \\ 0 \\ 0 \\ 1 \end{pmatrix}$$

We see that  $p^\mu \epsilon_\mu^{(1)} = 0$  and  $p^\mu \epsilon_\mu^{(2)} = 0$  so these are transverse - physical - photons  
 but  $p^\mu \epsilon_\mu^{(0)} \neq 0$  and  $p^\mu \epsilon_\mu^{(3)} \neq 0$   $\epsilon^{(0)}$  are timelike photons and  $\epsilon^{(3)}$  are longitudinal/  
 $\epsilon^{(0)}$  and  $\epsilon^{(3)}$  are unphysical

conjugate momentum  $\pi^\mu = \frac{\partial \mathcal{L}}{\partial \dot{A}_\mu} = F^{\mu 0} - g^{\mu 0} \partial_r A^r \Rightarrow \pi^0 = -\dot{A}^0 + \vec{\nabla} \cdot \vec{A}$  and  $\pi^i = 2\dot{A}^i - \dot{A}^i$

Then commutation relations for  $a(p)$  and  $a^\dagger(p)$  become

$$[a^{(\lambda)}(p), a^{(\lambda')\dagger}(p')] = -(2\pi)^3 \delta^3(\vec{p} - \vec{p}') g^{\lambda\lambda'} \quad \text{where } \lambda, \lambda' = 0, 1, 2, 3$$

$$\text{and } [a^{(\lambda)}(p), a^{(\lambda')}] = 0, \quad [a^{(\lambda)\dagger}(p), a^{(\lambda')\dagger}(p')] = 0$$

It can be shown that the contributions of the timelike and longitudinal photons cancel out in the Hamiltonian  
 $\rightarrow$  only the physical transverse photons contribute



In discussing gauge transformations  $\psi(x) \rightarrow e^{i\theta(x)} \psi(x)$ ,  $A_\mu \rightarrow A_\mu - \frac{1}{g} \partial_\mu \theta$   
 we introduced the covariant derivative  $D_\mu = \partial_\mu + iq A_\mu$   
 Note that under gauge transformation

$$D_\mu \psi \rightarrow \left[ \partial_\mu + iq \left( A_\mu - \frac{1}{g} \partial_\mu \theta \right) \right] (e^{i\theta} \psi) = \cancel{i \partial_\mu \theta e^{i\theta} \psi} + e^{i\theta} \partial_\mu \psi + iq A_\mu e^{i\theta} \psi - \cancel{i \partial_\mu \theta e^{i\theta} \psi} \\ = e^{i\theta} (\partial_\mu + iq A_\mu) \psi = e^{i\theta} D_\mu \psi$$

Invariance under  $\psi \rightarrow e^{i\theta} \psi$  is called  $U(1)$  gauge invariance

In general can consider  $\psi \rightarrow U \psi$  where  $U$  is a unitary matrix  $U^\dagger U = 1$

In the case of QED, with  $U(1)$ ,  $U = e^{i\theta}$  is a unitary  $1 \times 1$  matrix

This generalizes to higher groups  $U(N)$  for  $N \times N$  matrices

If we impose  $\det U = 1$  (subtract overall phase) then  $SU(N)$  group (special unitary)

In general  $U = e^{iH}$  with  $H$  Hermitian:  $H^\dagger = H$

$SU(2)$  for weak interactions  $H = \vec{\sigma} \cdot \vec{\theta}$  where  $\vec{\sigma}$  are the Pauli spin matrices  
 So  $\psi \rightarrow e^{i\vec{\sigma} \cdot \vec{\theta}} \psi$   $SU(2)$  transformation

$SU(3)$  for strong interactions  
Quantum Chromodynamics  $H = \frac{1}{2} \lambda^a \theta^a$  where  $\lambda^a$  are eight Gell-Mann matrices  
 So  $\psi \rightarrow e^{i \frac{1}{2} \lambda^a \theta^a} \psi$   $SU(3)$  transformation

In general, for  $SU(N)$ :  $H = T^a \theta^a$  and  $U = e^{iT^a \theta^a}$

where  $T^a$  are the generators of the Lie group

We have  $[T^a, T^b] = i f^{abc} T^c$  where  $f^{abc}$  are the structure constants

Then gauge field tensor (generalization of  $F^{\mu\nu}$  for photons) is

$$G_a^{\mu\nu} = \partial^\mu A_a^\nu - \partial^\nu A_a^\mu + g f_{abc} A_b^\mu A_c^\nu$$

The last term shows that  
e.g. gluons self-interact

Gravity as a gauge theory (general relativity)

$A^\mu$  corresponds to  $\Gamma_{\lambda\nu}^\mu = \frac{1}{2}(\partial_\lambda g_{\nu\mu} + \partial_\nu g_{\lambda\mu} - \partial_\mu g_{\lambda\nu})$  or Christoffel symbol  
gauge transformation coordinate transformation connection coefficient

and  $D_\mu \psi = \partial_\mu \psi + iq A_\mu \psi$  corresponds to  $D_\nu V^\mu = \partial_\nu V^\mu + \Gamma_{\lambda\nu}^\mu V^\lambda$

Finally,  $G_a^{\mu\nu}$  corresponds to Riemann curvature tensor  $R_{\lambda\mu\nu}^\sigma$

$$\text{with } R_{\lambda\mu\nu}^\sigma = \partial_\nu \Gamma_{\lambda\mu}^\sigma - \partial_\mu \Gamma_{\lambda\nu}^\sigma + \Gamma_{\lambda\mu}^\rho \Gamma_{\rho\nu}^\sigma - \Gamma_{\lambda\nu}^\rho \Gamma_{\rho\mu}^\sigma$$

Einstein equations  
of general relativity

$$R^{\mu\nu} - \frac{1}{2} g^{\mu\nu} R = -8\pi T^{\mu\nu}$$

with  $R^{\mu\nu}$  the Ricci tensor and  $R$  the Ricci scalar