

(25.9.25) Computational Physics B

class 1

- based on the density of particles to determine the spatial resolution.(grid-based vs particle-based)
 - smoothed particle hydrodynamics (SPH)
 - particles have volume, and we use the density of particles to describe it.

- $$A(\vec{r}) = \sum_i \frac{m_i}{\rho_i} W(\vec{x} - \vec{x}_i, h)$$

- - i: index of the particle
 - m_i : mass of particle i
 - ρ_i : density of particle i
 - h: smoothing length, which means we smooth the mass of the particle to a volume with radius h.
 - W: smoothing kernel, which is a function of distance and smoothing length.
 - change $A_i \rightarrow \rho_i$, then the equation becomes

- $$\rho(\vec{r}) = \sum_i m_i W(\vec{x} - \vec{x}_i, h)$$

- this means the density at point r is contributed by all the particles, and the contribution is determined by the distance between point r and particle i, as well as the smoothing length h.(The kernel function W is the power of the particle mass)
- this method transform the discrete particles to a continuous field, and seems not to contain something about particle interaction.
- when gravity is considered, there may be some problems.
 - when two particles are very close, the gravity between them will be very large, which may lead to numerical instability.
 - so we need to soften the gravity when two particles are very close.

- $$\vec{F} = \frac{Gm_1m_2}{r^2 + \epsilon^2} \hat{r}$$

- ϵ : softening length, which is similar to the smoothing length h.

- this means when the distance between two particles is smaller than ϵ , the gravity will not increase anymore.
- this can avoid numerical problems, but also means we lose some information about small scale structures.

- discretization by frequencies

- spectral method
 - Fourier decomposition

-

$$f(x, t) = \sum_k \hat{f}(t) e^{ikx}$$

transform from the k space to t space.

- if we input too many waves into it, there may exist a "Gibbs spike"
- this method is not suitable for fluid discontinuity, such as shock waves.
- Dedalus -- recommended for spectral method

class 2

- pros and cons

- spectral method

- pros: high resolution, low computational cost
- cons: not suitable for discontinuity(for example, high Mach number flow), only able to solve periodic boundary condition.(idealized geometries)
- application: incompressible turbulence($M \ll 1$), idealized setup(hydrogen atom)

- particle-based

- pros: ODEs, easy to cover large dynamic range(particles flow with the fluid field,high density -> high resolution -> dissolve the physical process)(cosmic web -> dark matter halo -> galaxy -> star formation),
- cons: resolution restricted by smooth length and softening length, low resolution and hard to capture instability(smoothing curve will suppress the instability), disable to include magnetic field($\nabla \cdot B = 0$ is hard to satisfy)(divengence cleaning), so MHD simulation is hard to do.
- application: cosmological simulation, galaxy formation

- grid-based

- pros: high order accuracy, easy to capture shock waves and instabilities(complex physical processes)
- cons: PDEs, hard for coding, high computational cost, hard to cover large dynamic range(even with adaptive mesh refinement, the resolution is still limited by the finest grid)

- application: localized simulation, compressible and incompressible turbulence
- not include plasma and binetic simulation -- not hydrodynamic equations but Maxwell equations
 - particle-in-cell (PIC)(for cosmic rays we use this method)
 - radiative transfer(absorption, reemission, leaping)(can couple with hydrodynamic simulation)
- basic structure of numerical simulations
 - initial valued problem(IVP) -- give the initial condition to see the time revolution
 - boundary valued problem(BVP) -- specify the condition and fix the solution