

## PH 4130/PH 6130 Assignment 5

Deadline **1 March 2018 before 17:00 hrs**

Please show the source code for each of the problems.

1. Extend the analysis done to create astroML figure 4.7 by generating the Gaussian distribution 100 times. In other words, create 10,000 draws from a Gaussian distribution with  $\mathcal{N}(0,1)$  100 times (instead of just once as in Figure 4.7 ). Plot the histogram of Anderson-Darling  $A^2$ , Kolmogorov-Smirnov  $D$ , and Shapiro-Wilks  $W$  test for each of the 100 numerical simulations. (20 points)
2. Download the asteroid dataset from [http://astrostatistics.psu.edu/datasets/asteroid\\_dens.dat](http://astrostatistics.psu.edu/datasets/asteroid_dens.dat). Apply the Shapiro-Wilk test to both the asteroid density values and the natural logarithm of the density values. From the  $p$  values, which of these is closer to a Gaussian distribution? Verify this by plotting histograms of both density and its logarithm and overlaying the best-fit normal distribution (Look up `stats.norm.fit`) (20 points)
3. Download the Hipparcos star catalog from [http://astrostatistics.psu.edu/datasets/HIP\\_star.dat](http://astrostatistics.psu.edu/datasets/HIP_star.dat). Detailed explanation of the columns in this dataset can be found in [http://astrostatistics.psu.edu/datasets/HIP\\_star.html](http://astrostatistics.psu.edu/datasets/HIP_star.html) under “Dataset”. Calculate using two-sample t-test whether the color of the Hyades stars differs from the non-Hyades ones. The Hyades stars have Right Ascension between  $50^\circ$  and  $100^\circ$ , declinations between  $0$  and  $25^\circ$ , proper motion in RA between 90 and 130 mas/year, proper motion in DEC between -60 and -10 mas/year, measurement parallax error less than 5 mas, and color of star less than 0.2 mag. Any other star which does not satisfy any of the above conditions is considered a non-Hyades star. (20 points)
4. The T90 distribution for Beppo-Sax T90 data can be found at <http://www.iith.ac.in/~shantanud/beppoSax.txt>. Apply GMM to  $\log_{10}$  of T90 data and find the optimum number of components using AIC and BIC by plotting AIC/BIC as a function of number of componts (20 points) (Hint: Look at the source code for astroML figure 6.6)