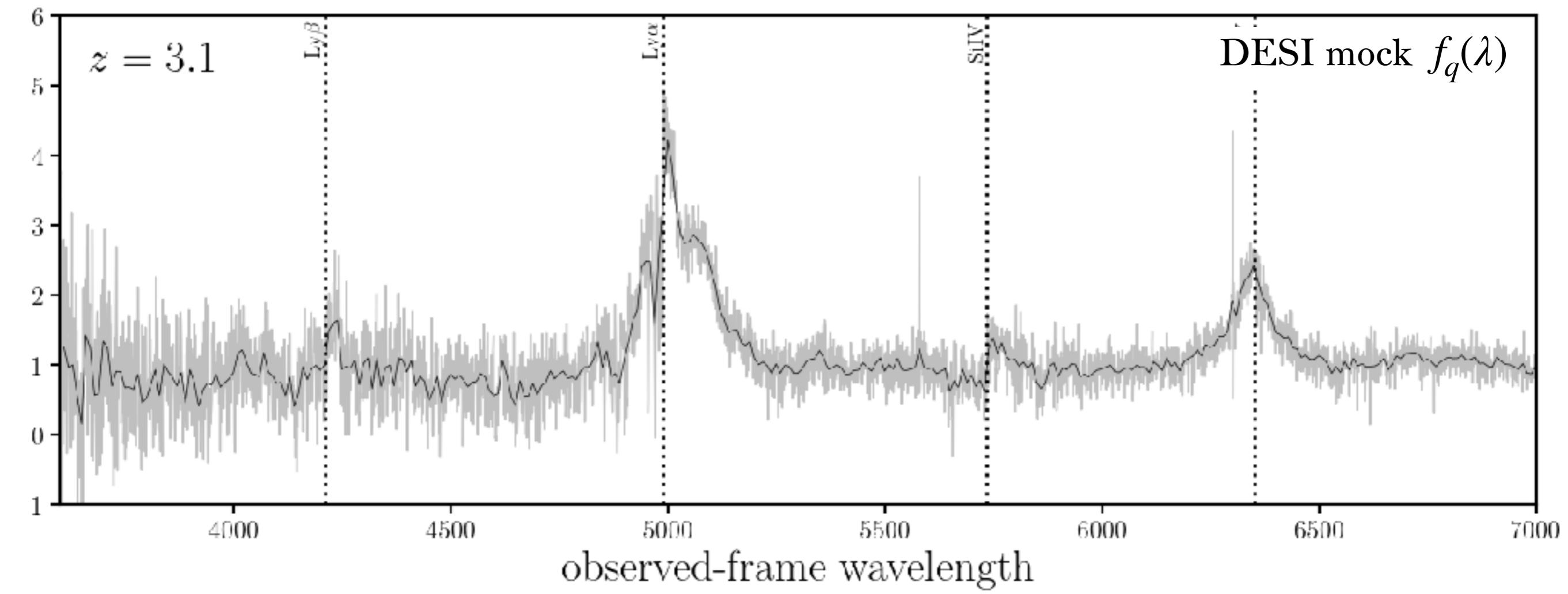




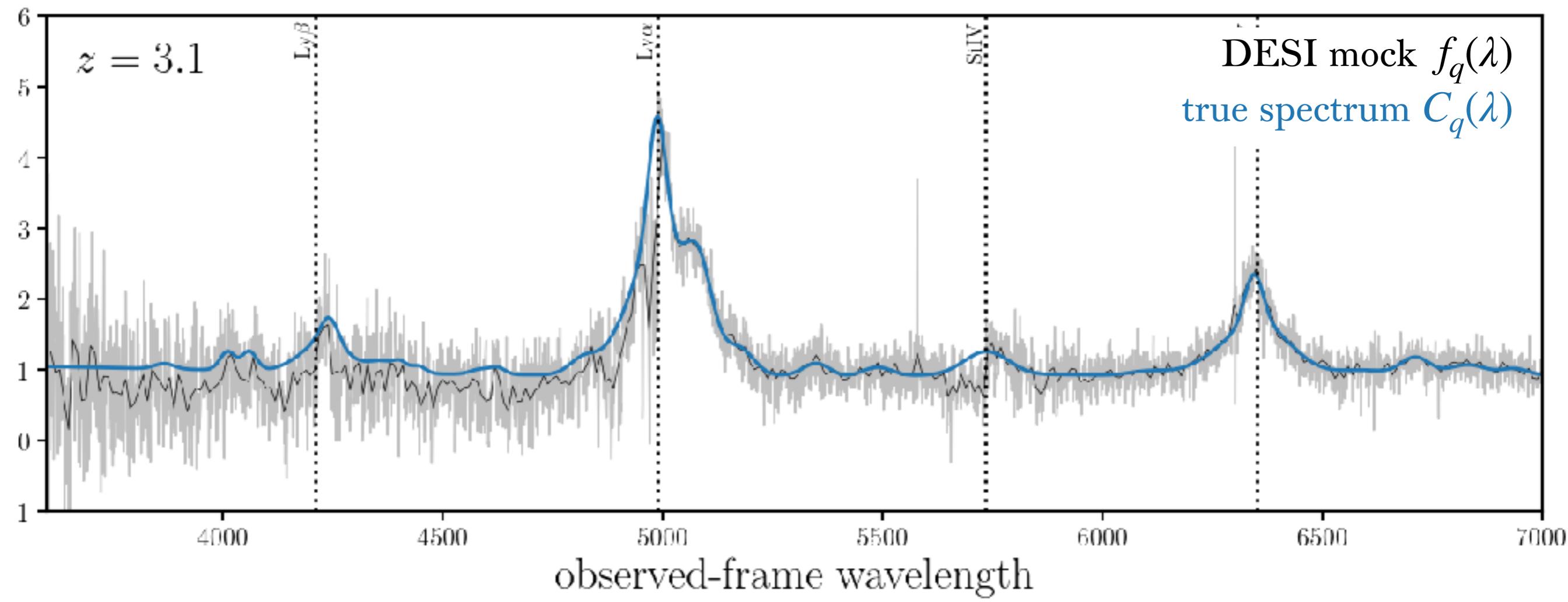
Reconstructing Quasar Spectra and measuring the Ly α Forest

CHANGHOON HAHN

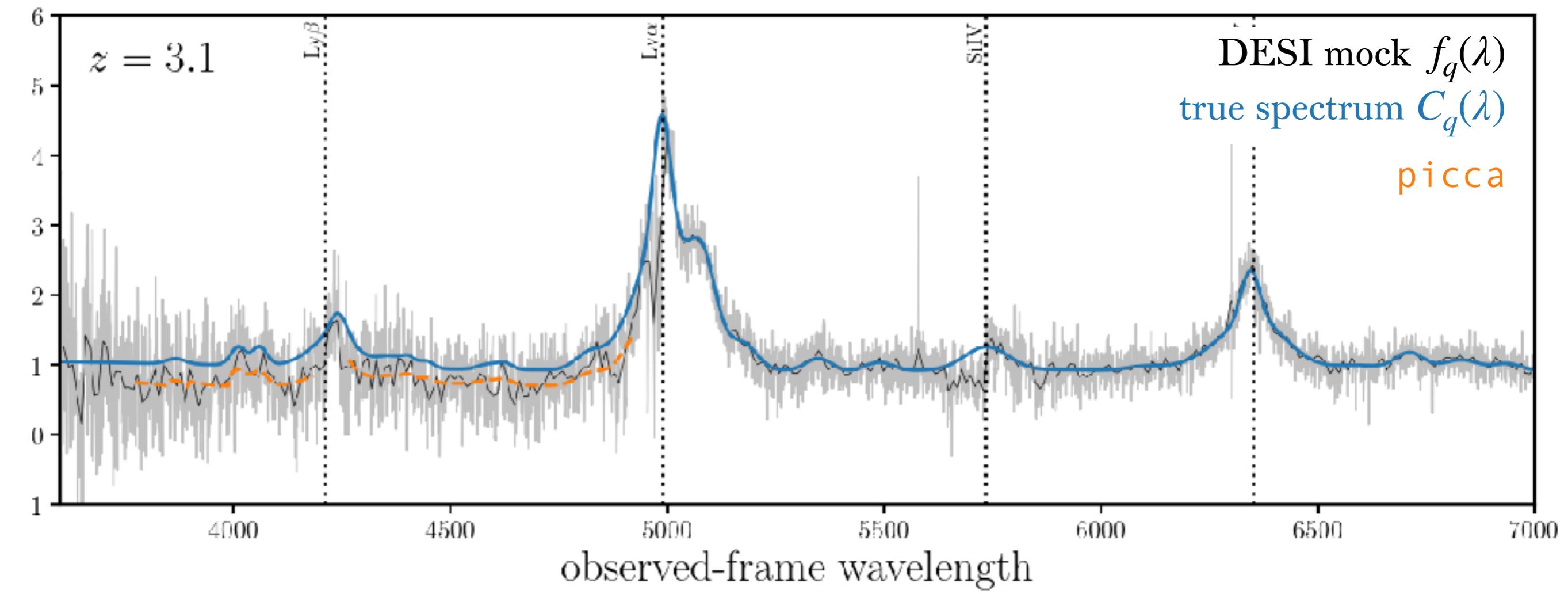
with Satya Gontcho A Gontcho, Peter Melchior,
Abby Bault, Hiram Herrera



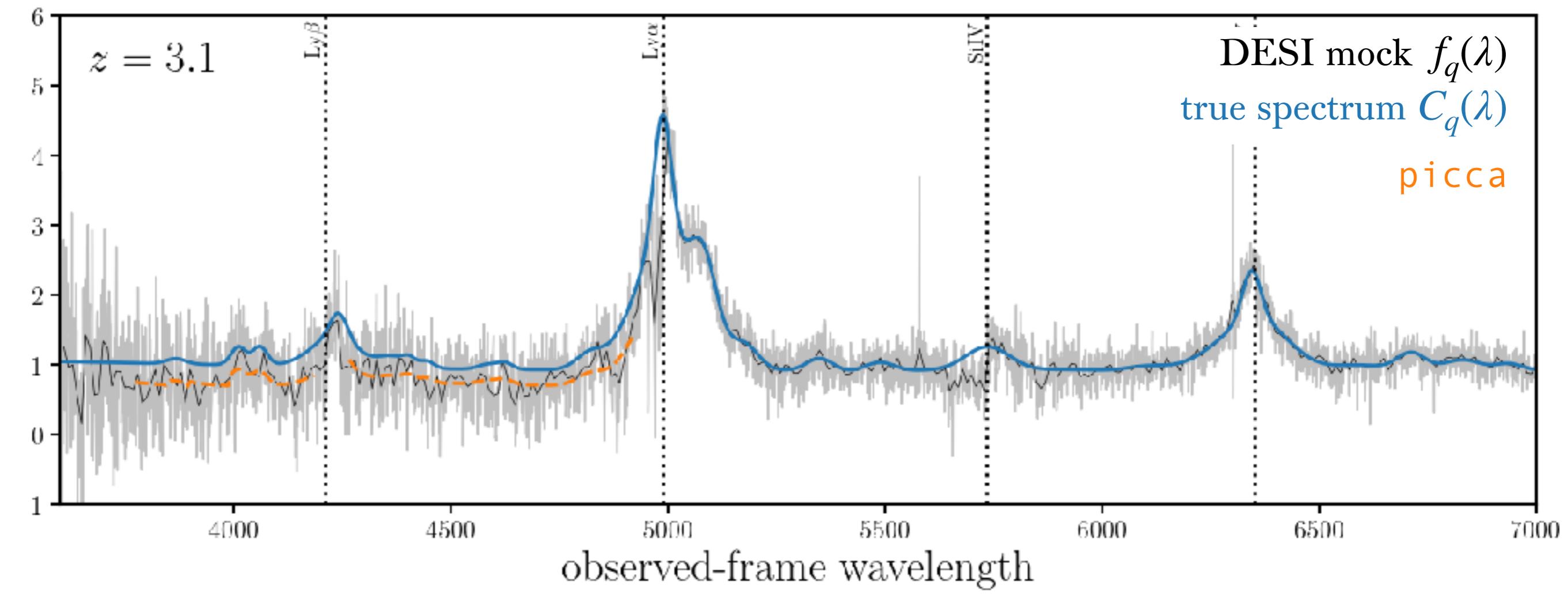
quasar spectra carry the imprint of foreground IGM (e.g. Ly α forest)



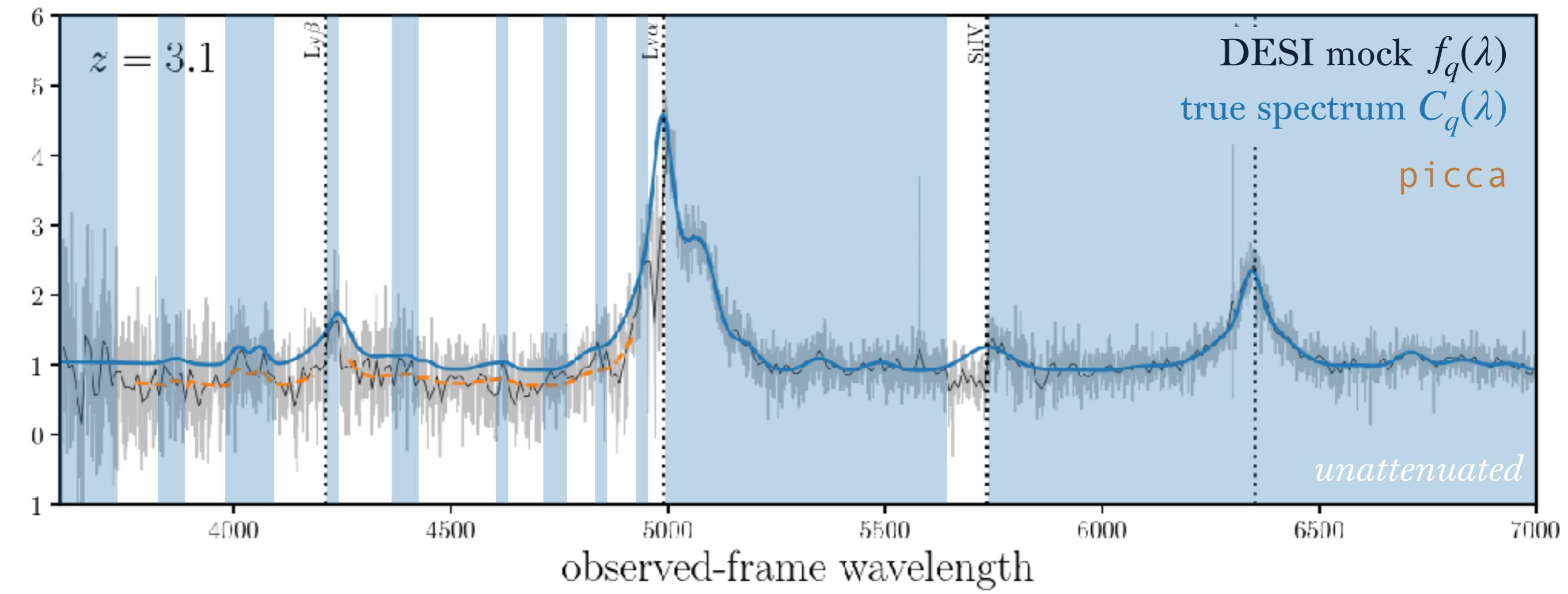
quasar spectra carry the imprint of foreground IGM (e.g. Ly α forest)
but requires estimating the true continuum $C_q(\lambda)$



current state-of-the-art `picca` is *biased*

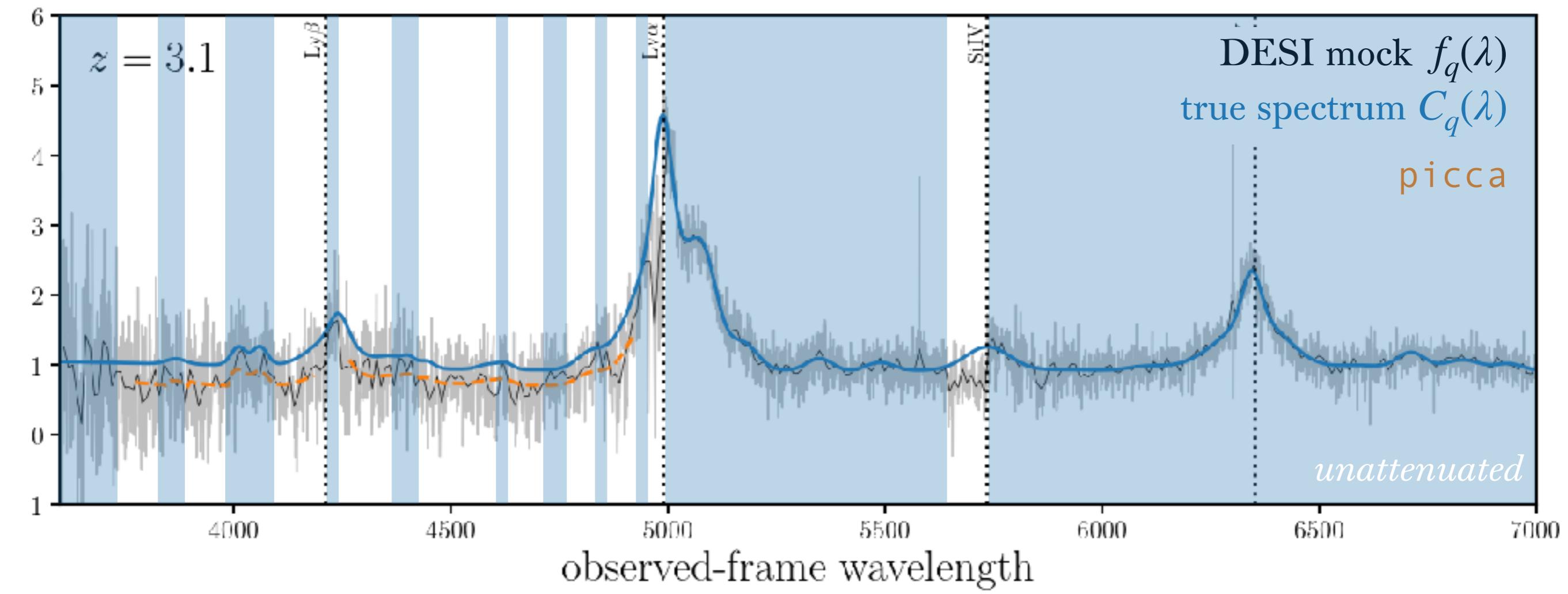


could we *reconstruct* the quasar continuum?



could we *reconstruct* the quasar continuum?

parts of the spectra are unattenuated

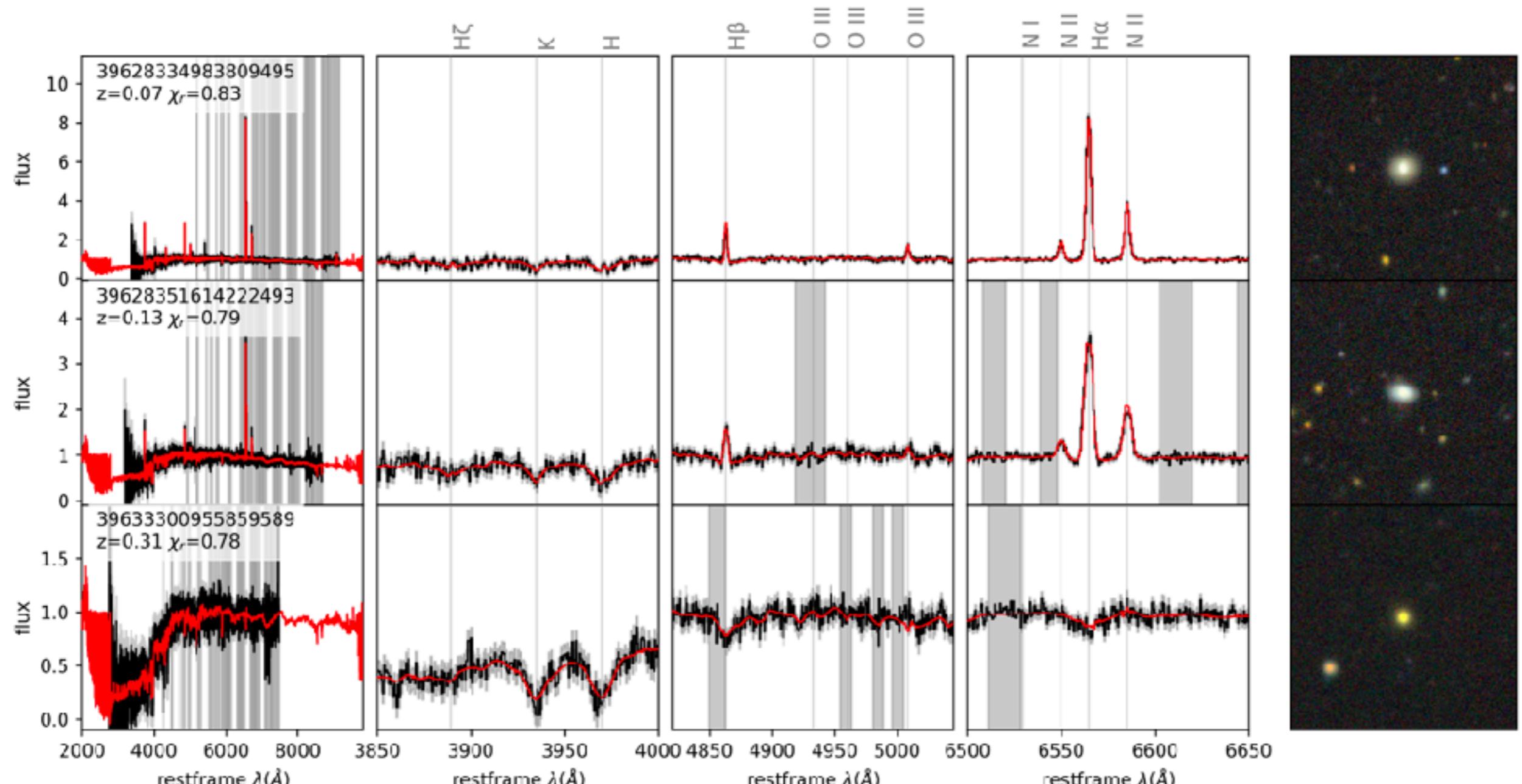


could we *reconstruct* the quasar continuum?

parts of the spectra are unattenuated

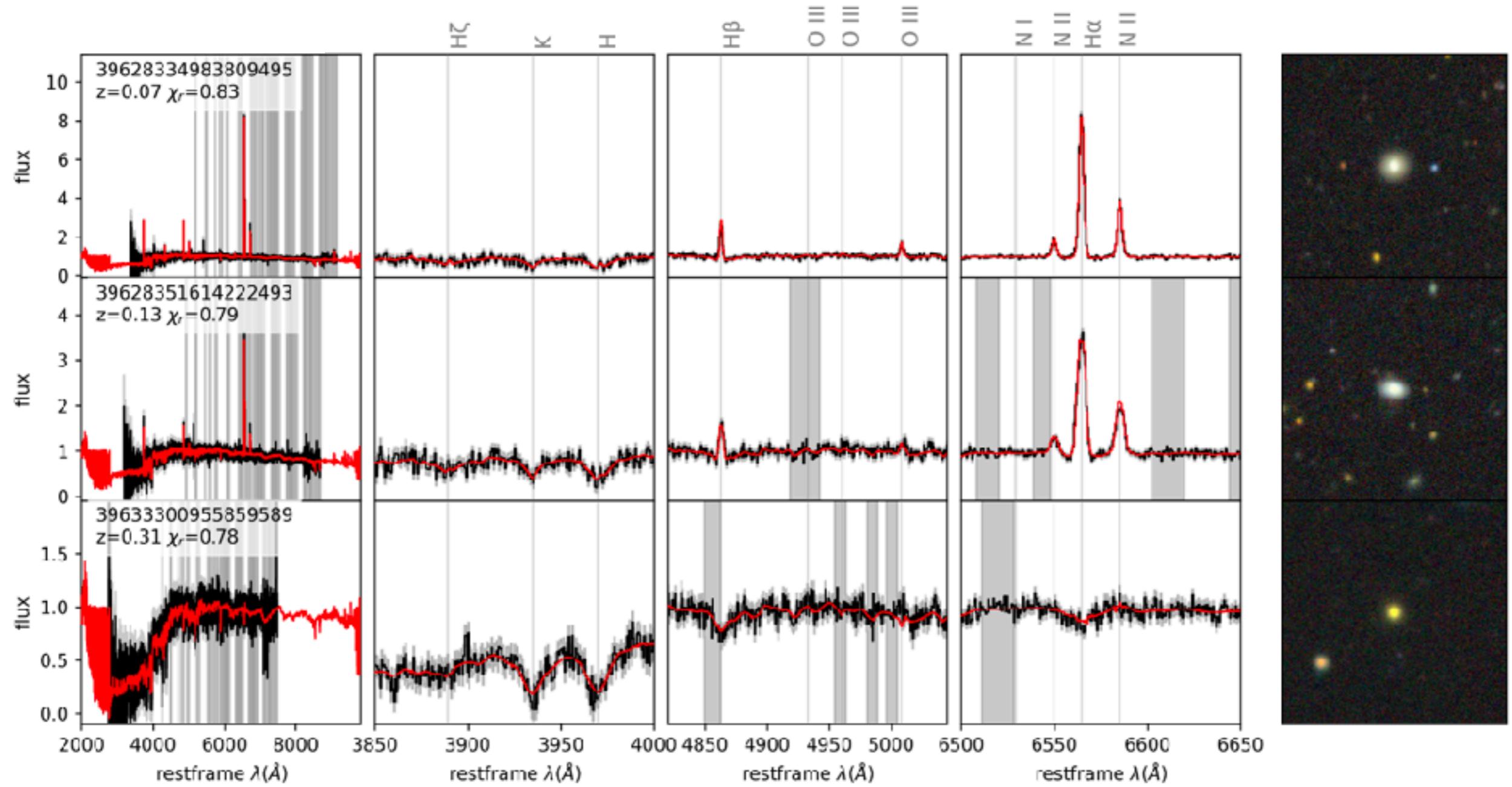
foreground absorption is uncorrelated with the quasar

spectral autoencoders have been very successful in reconstructing galaxy spectra — e.g. **spender**



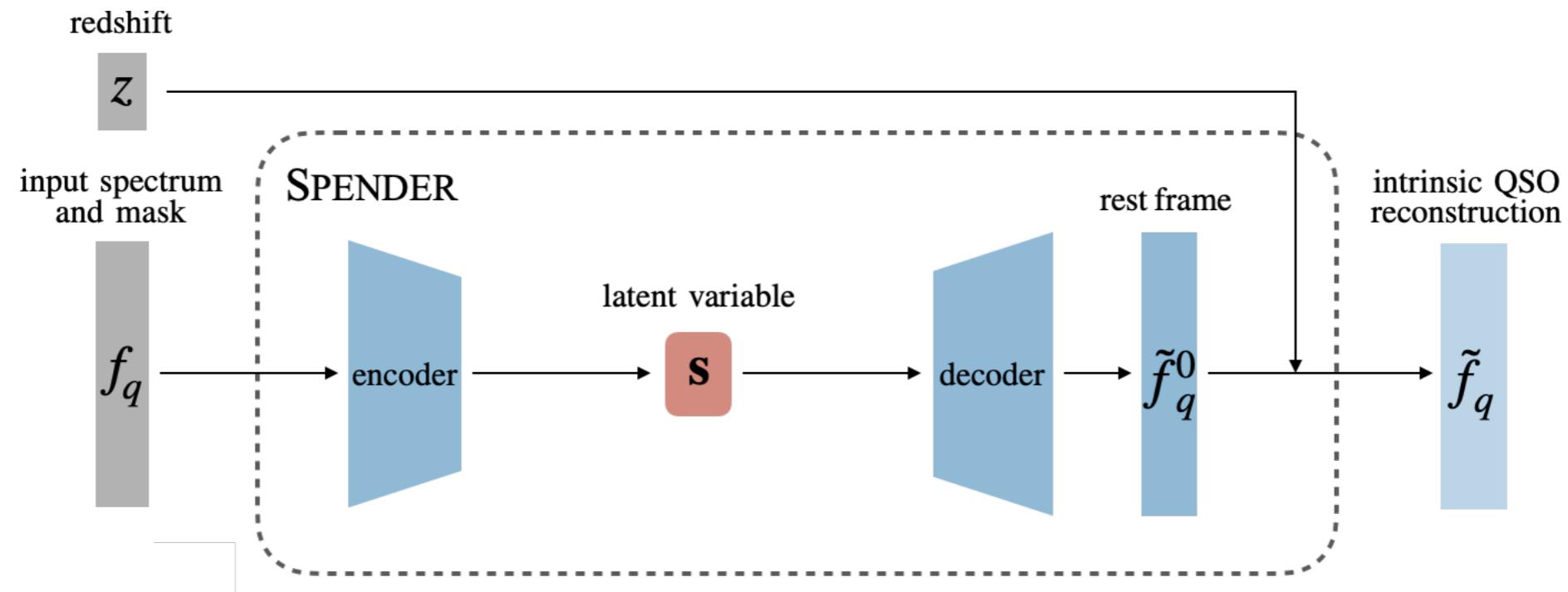
DESI BGS galaxies

spectral autoencoders have been very successful in reconstructing galaxy spectra — e.g. **spender**

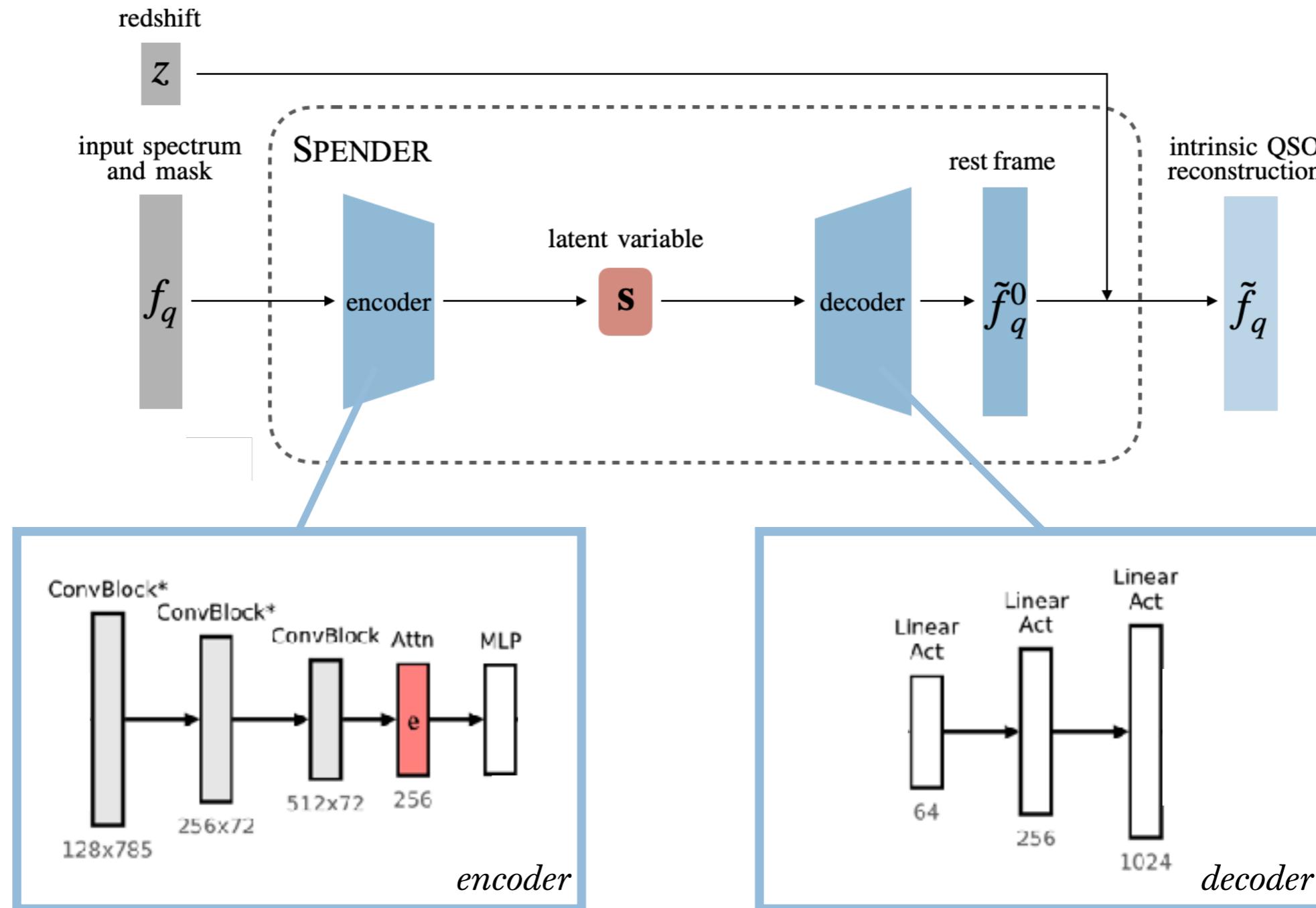


fully *data-driven* (i.e. no training required on simulations)

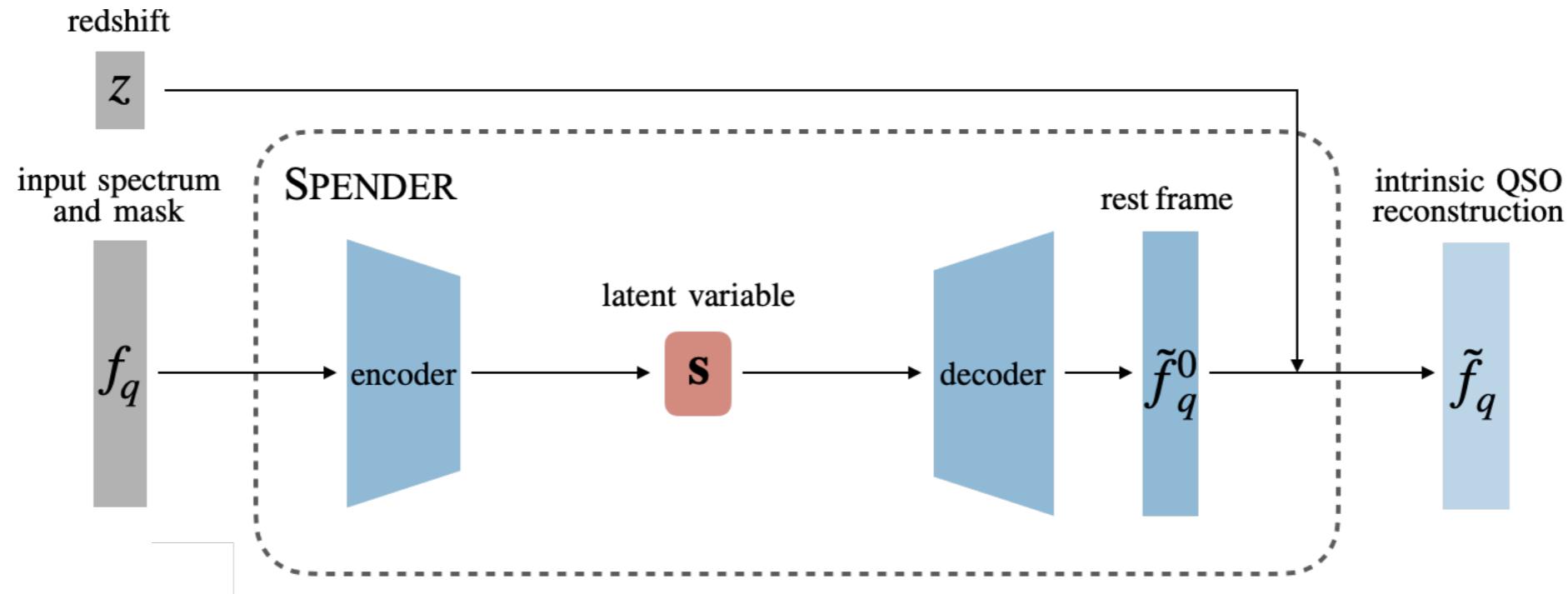
spender: spectral autoencoders learns low-dim representation of spectra in the latent space



spender: spectral autoencoder with attentive convolutional encoder and an explicit redshifting



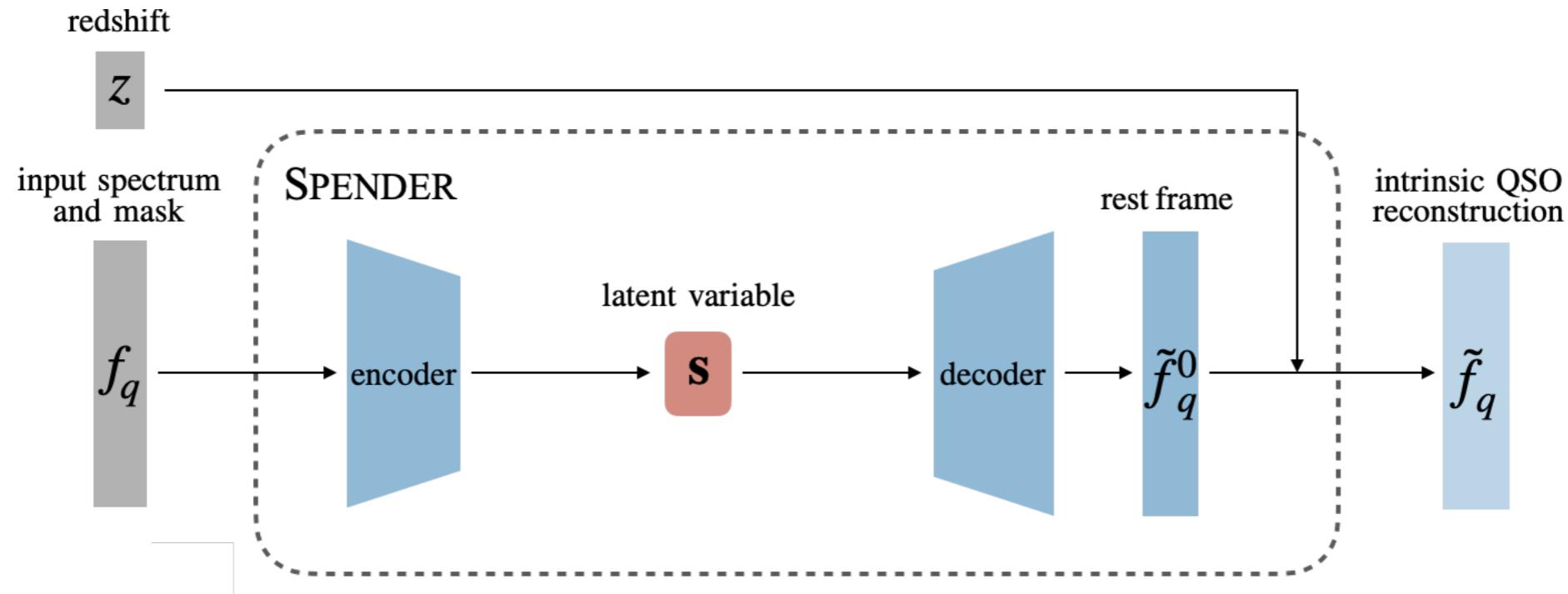
spender: spectral autoencoder with fidelity loss + similarity and consistency losses



$$L_{\text{total}} = L_{\text{fid}} + L_{\text{sim}} + L_{\text{con}}$$

latent distance \propto spectral distance

spender: spectral autoencoder with fidelity loss + similarity and consistency losses

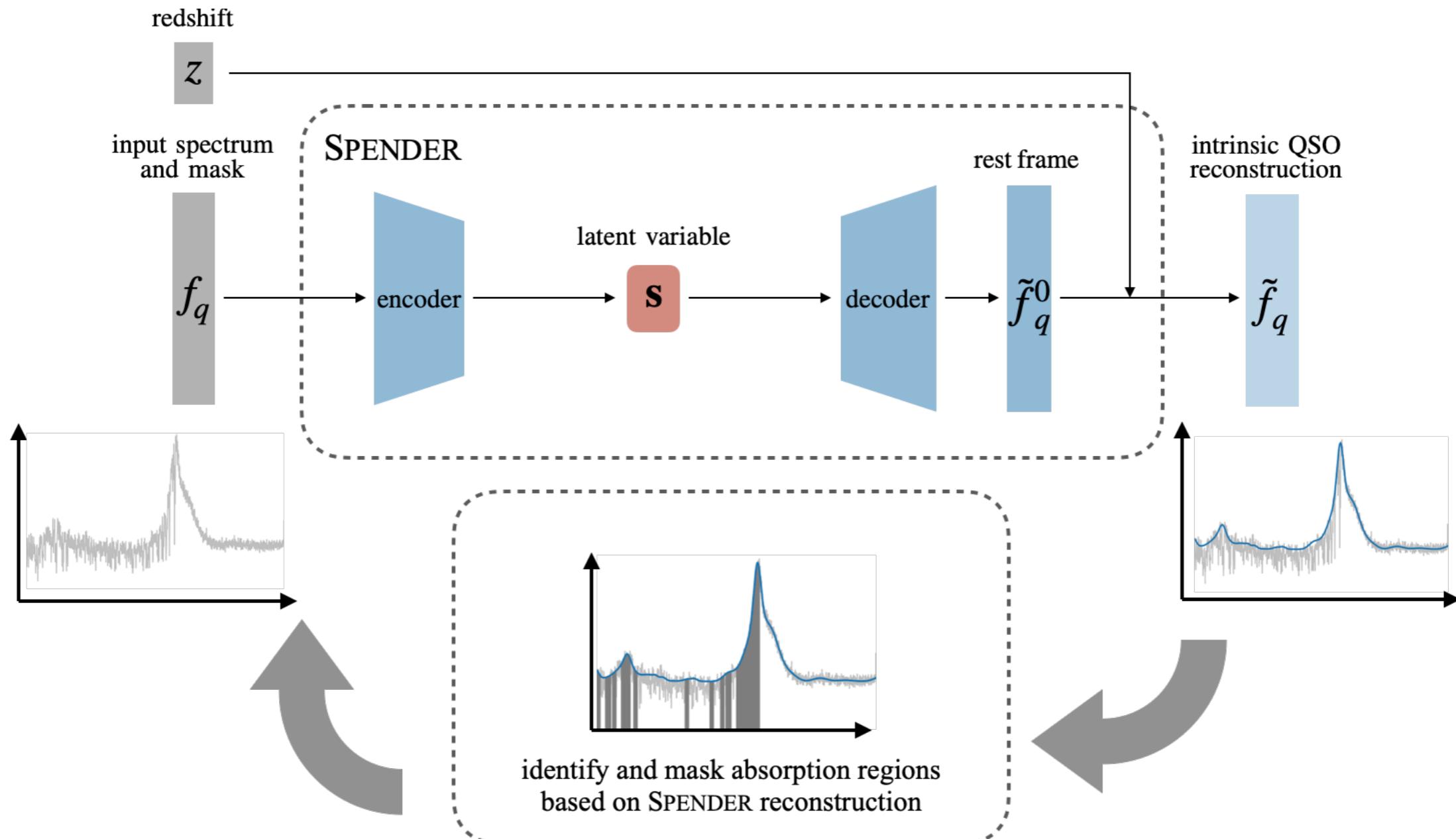


$$L_{\text{total}} = L_{\text{fid}} + L_{\text{sim}} + L_{\text{con}}$$

redshift invariance

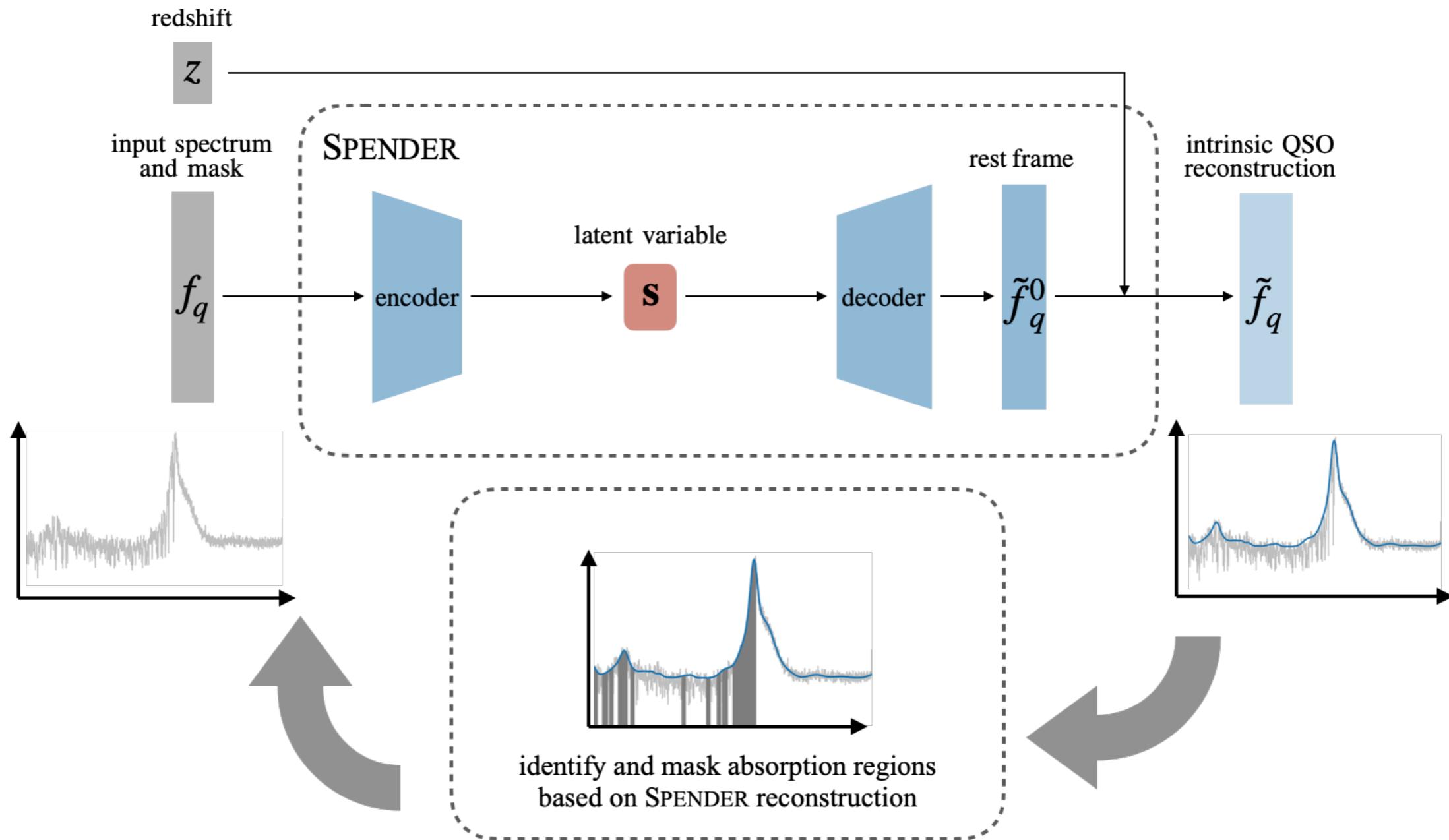


spender + iterative procedure for identifying and masking absorption

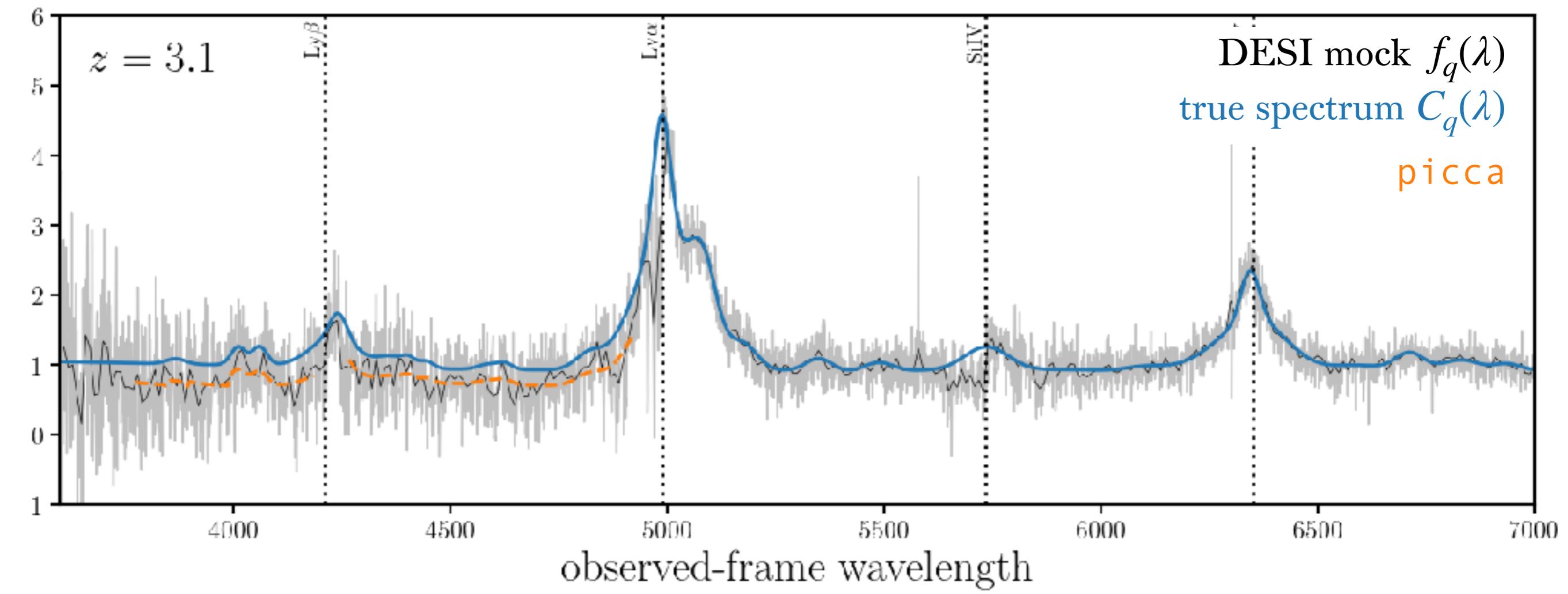


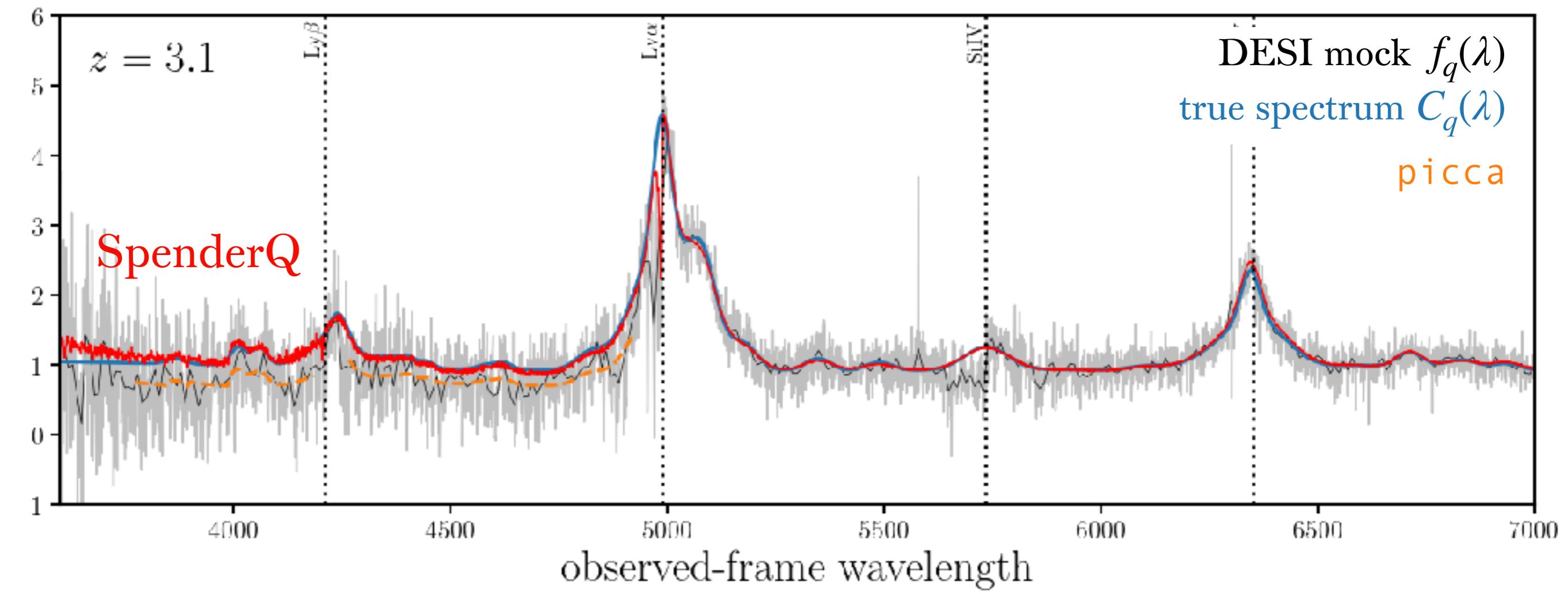


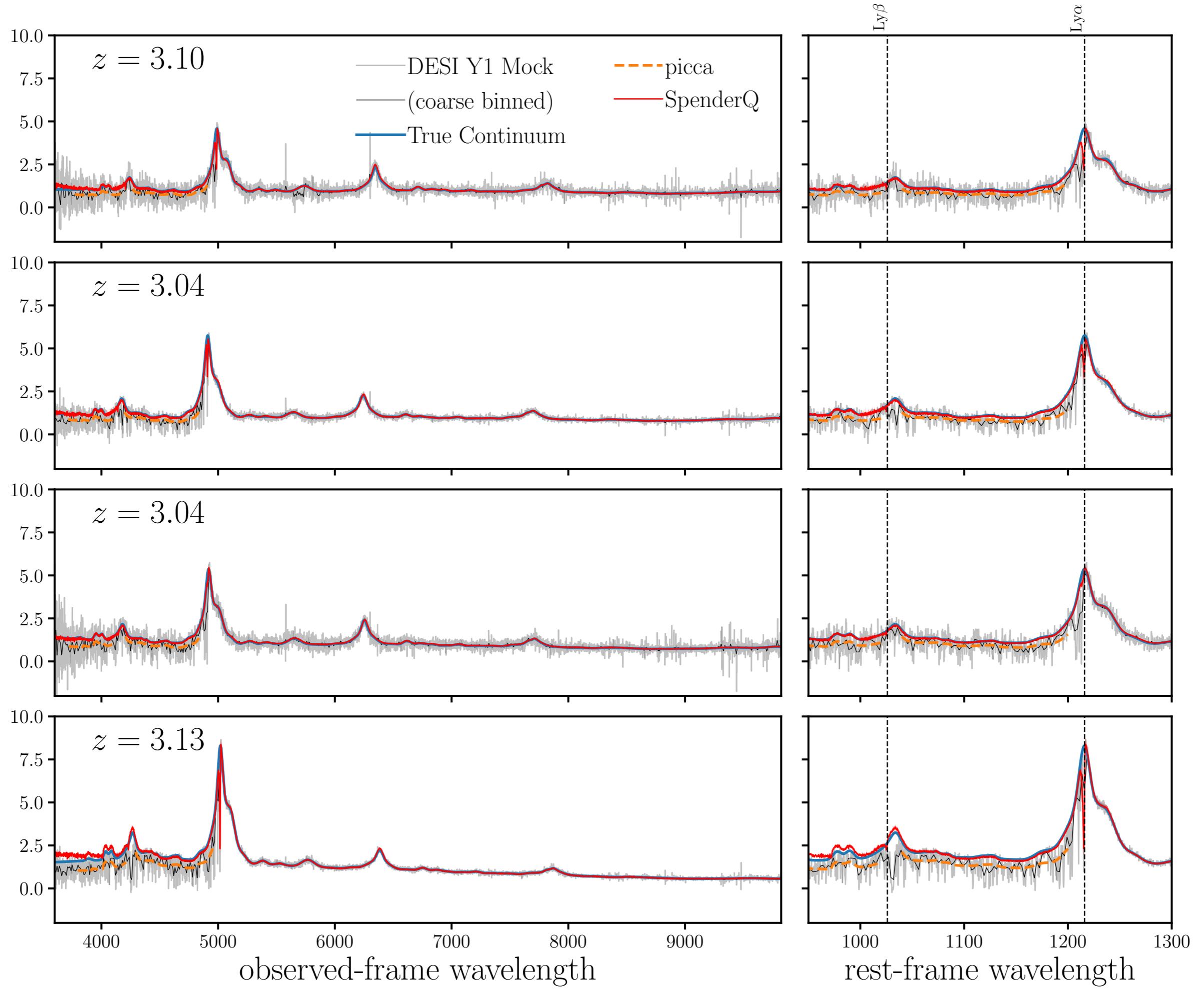
spender + iterative procedure for identifying and masking absorption



SpenderQ will not learn the absorption features since they are uncorrelated with the quasar spectra

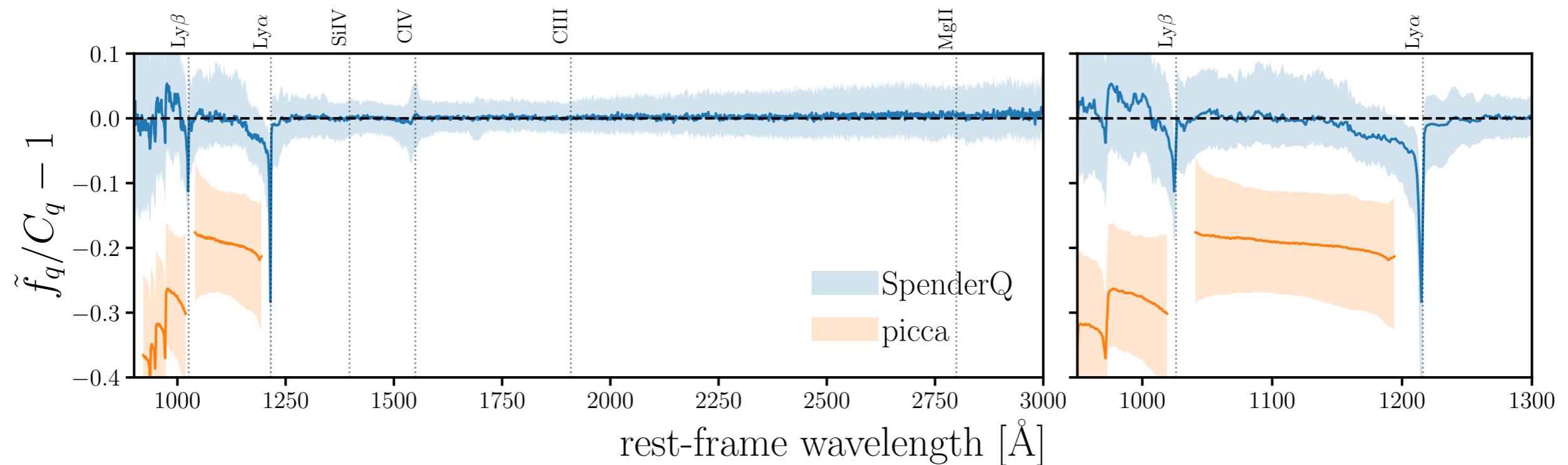




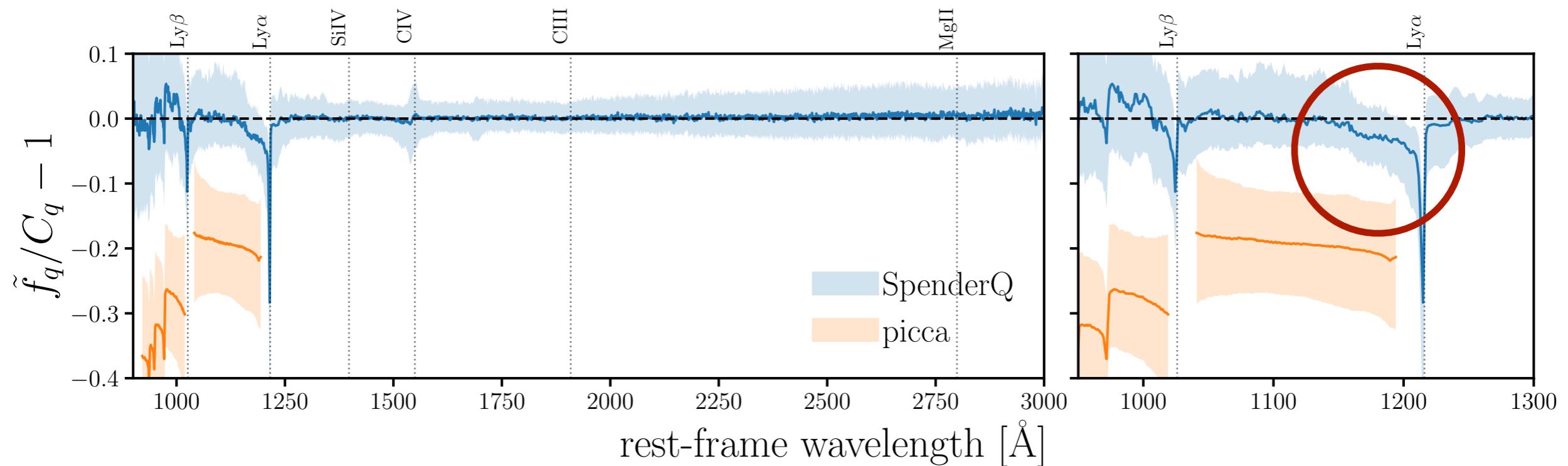


Spender Q validation on DESI Year 1 “*London*” mocks, 400,000 spectra constructed using CoLoRe, Ly α CoLoRe, and quickquasar

SpenderQ provides *percent-level* reconstructions of the quasar spectra
— picca residuals: $\sim 20\%$ in Ly α and $> 30\%$ in Ly β

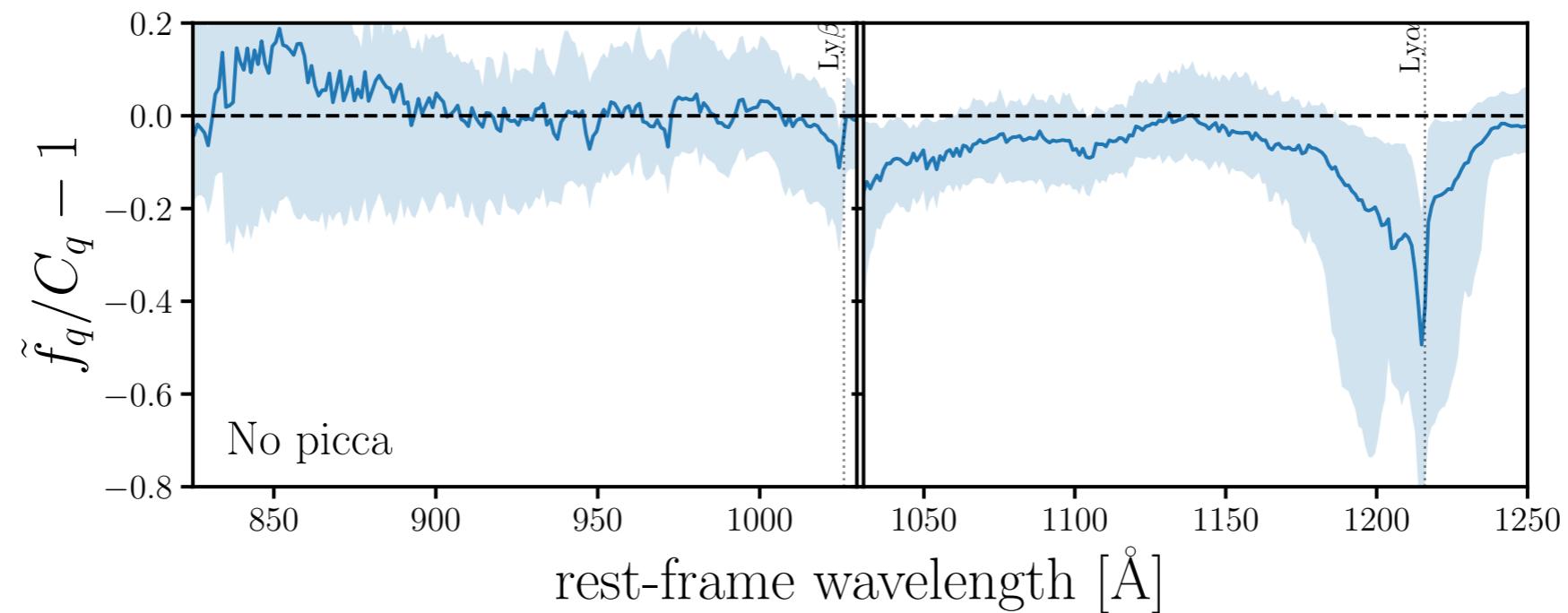


SpenderQ provides *percent-level* reconstructions of the quasar spectra
— picca residuals: $\sim 20\%$ in Ly α and $> 30\%$ in Ly β



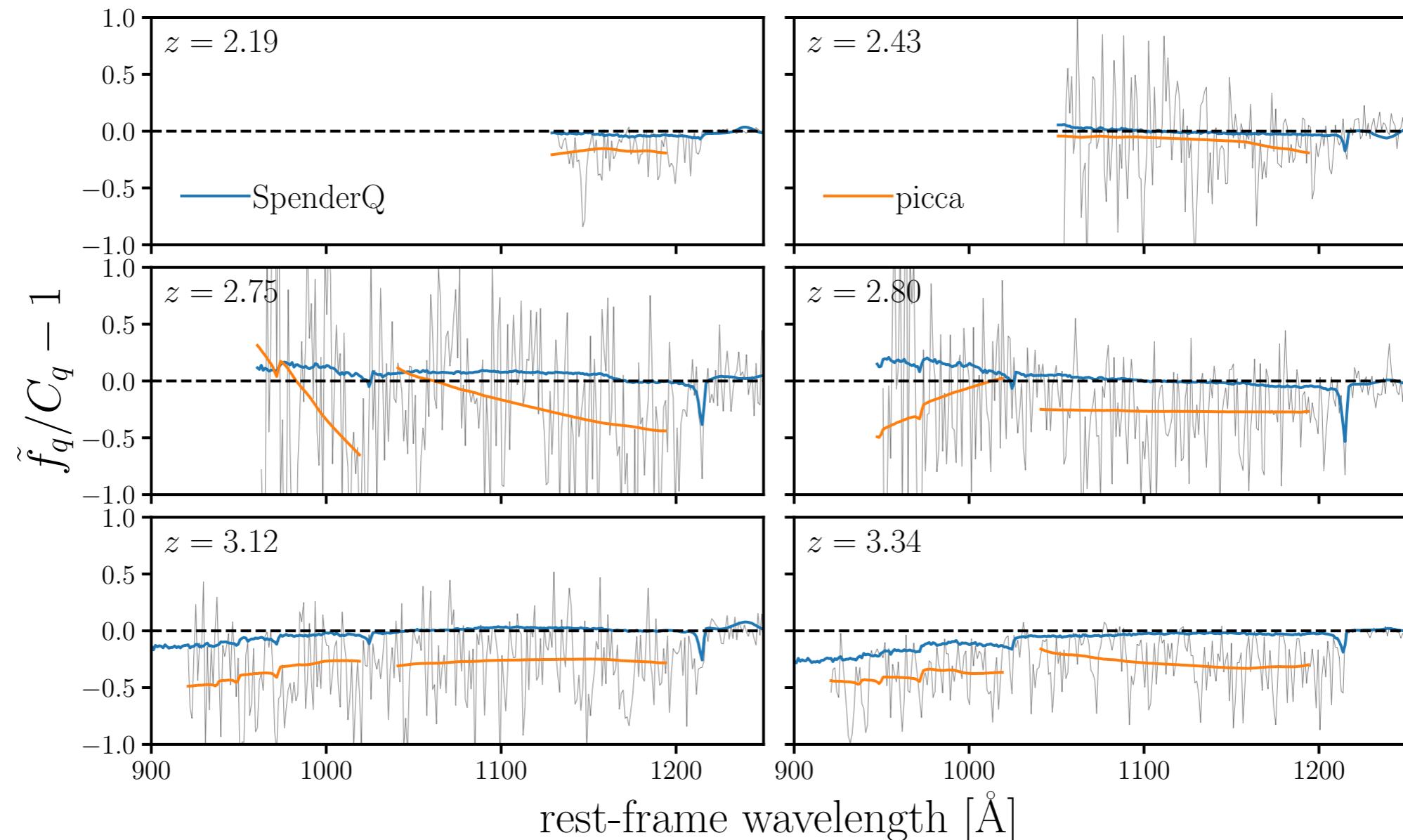
we do not remove spectra flagged with BAL or DLAs in SpenderQ

SpenderQ provides reconstructions for *all* quasars

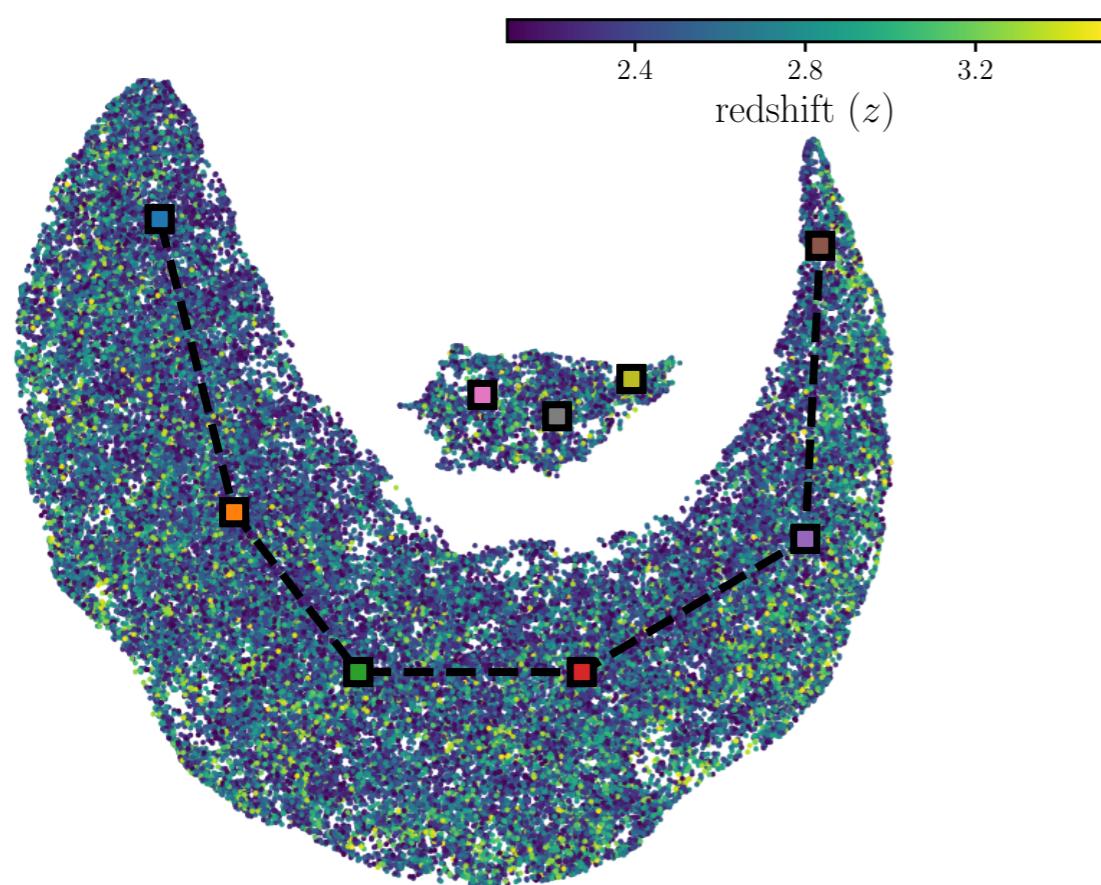


picca requires >150 spectral elements in the Ly α /Ly β region

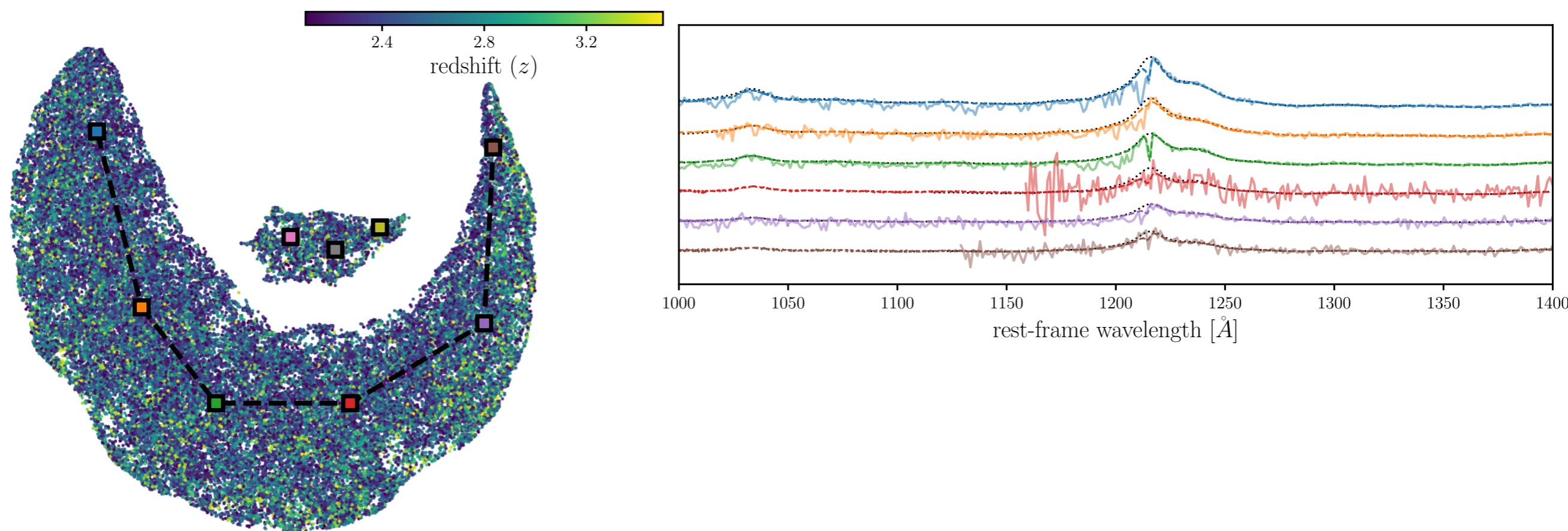
SpenderQ accurately reconstructs the spectrum of *individual* quasars



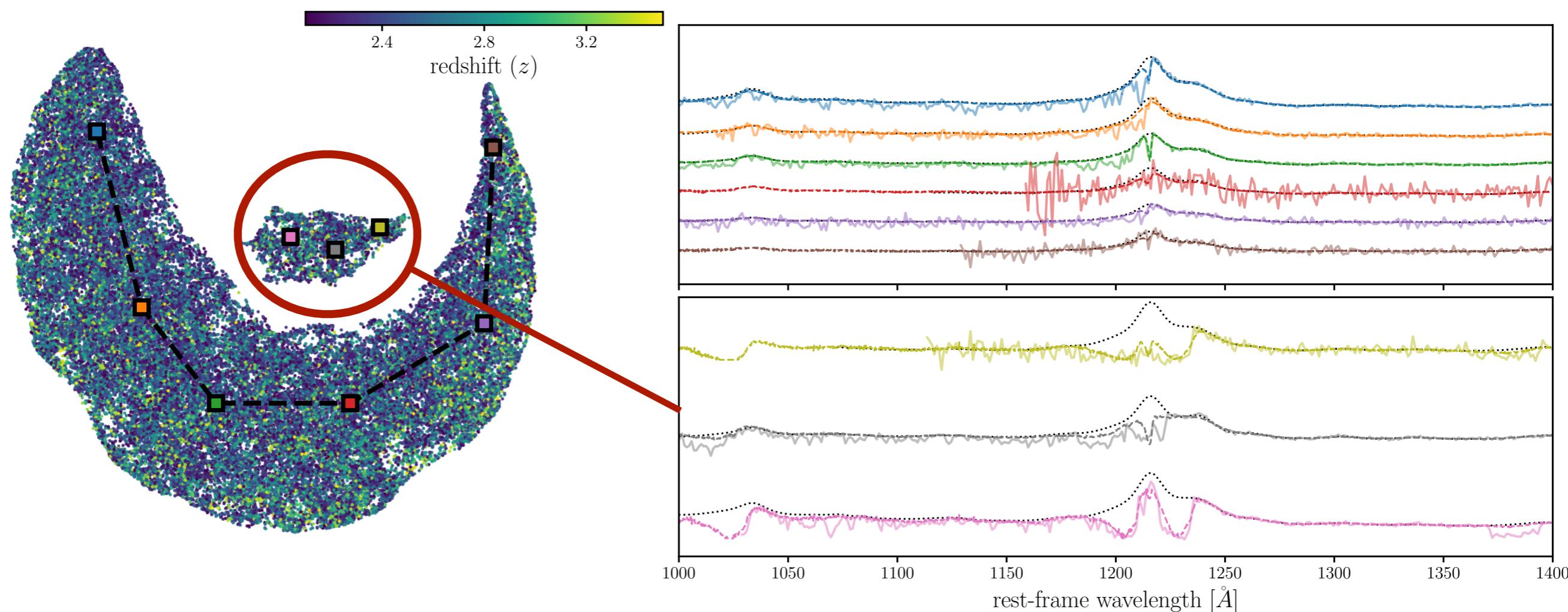
SpenderQ provides a *redshift invariant* and physically informative latent encoding



SpenderQ provides a *redshift invariant* and *physically informative* latent encoding



SpenderQ provides a *redshift invariant* and *physically informative* latent encoding — identifies Broad Absorption Lines (BALs)



SpenderQ applied to DESI Year 1 “*London*” Ly α mocks

Hahn, Gontcho A Gontcho et al (in prep.)

SpenderQ applied to DESI Year 1 “*London*” Ly α mocks

Hahn, Gontcho A Gontcho et al (in prep.)

SpenderQ applied to DESI Early Data Release

Ly α forest “delta” catalog

IGM absorber catalog

reconstructed quasar catalog

latent embedding catalog

Gontcho A Gontcho, Hahn et al (in prep.)

SpenderQ applied to DESI Year 1 “*London*” Ly α mocks

Hahn, Gontcho A Gontcho et al (in prep.)

SpenderQ applied to DESI Early Data Release

Gontcho A Gontcho, Hahn et al (in prep.)

SpenderQ for Ly α clustering analyses — BAO and full-shape analyses

SpenderQ applied to DESI Year 1+ Data Release