

# eMa $\nu$ 1.: Breaking $\sum m_\nu$ Parameter Degeneracies with Three-point Statistics

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## ABSTRACT

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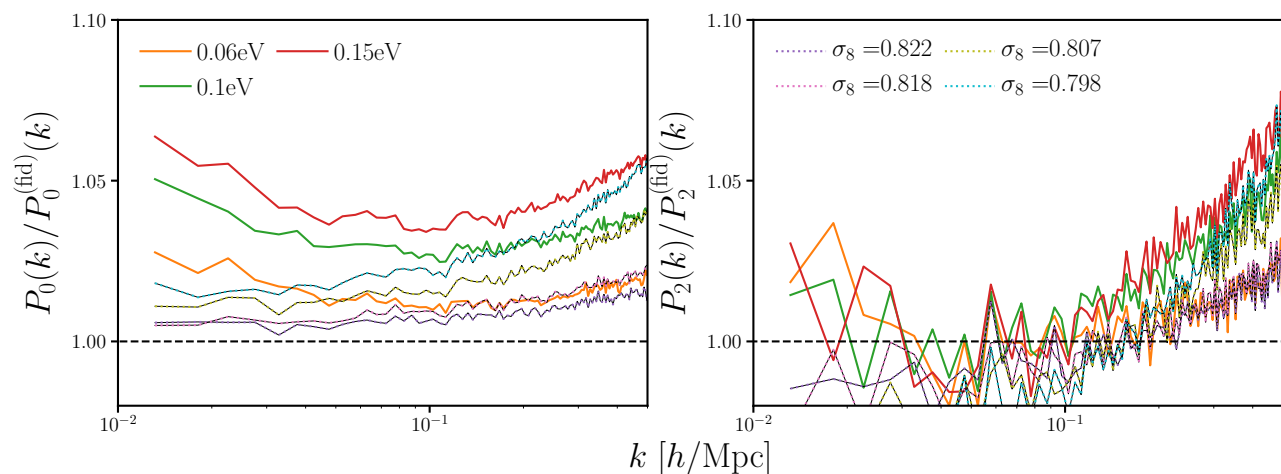
*Keywords:* cosmology: —

## 1. INTRODUCTION

talk about the impact of massive active neutrinos on the matter powerspectrum and how that's detectable with CMB and LSS.

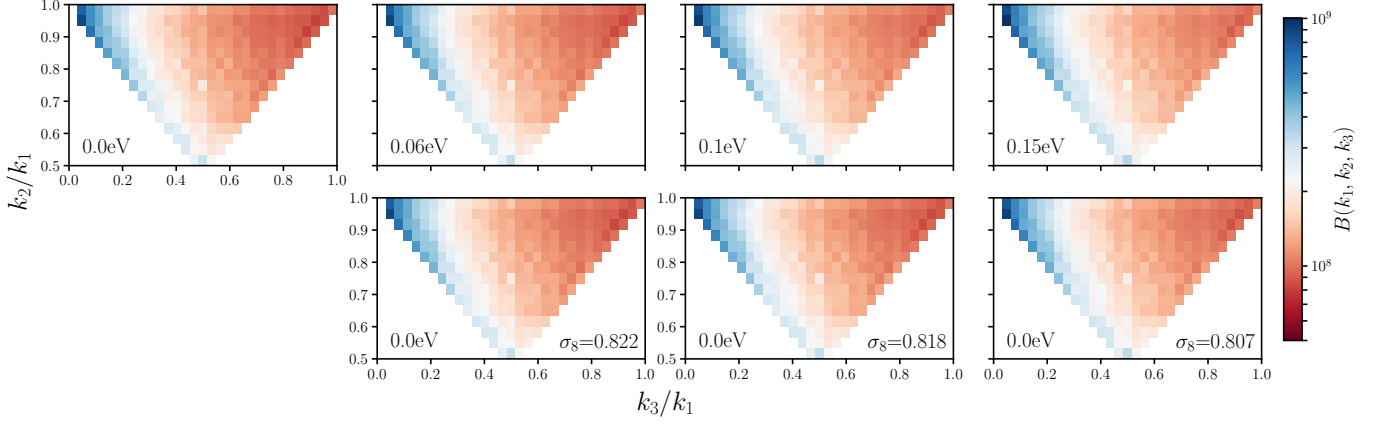
one big roadblock is the degeneracy among  $\tau$   $\sigma_8$  and  $\sum m_\nu$ . short thing about how  $\tau$  is hard to constrain

## 2. HADES SIMULATIONS

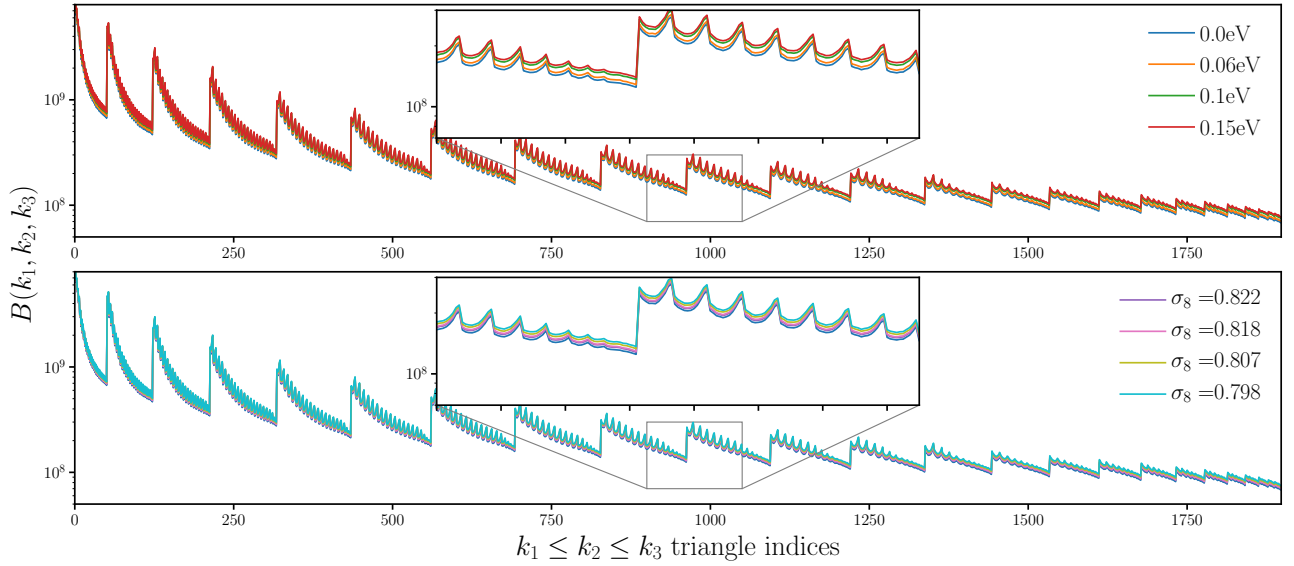


**Figure 1.** Impact of  $\sum m_\nu$  and  $\sigma_8$  on the halo power spectrum monopole and quadrupole.  $\sum m_\nu$  and  $\sigma_8$  produce almost identical effects on halo clustering on small scales ( $k > 0.1 h/\text{Mpc}$ ). This degeneracy can be partially broken through the quadrupole; however,  $\sum m_\nu$  and  $\sigma_8$  produce, within a few percent, almost the same effect on two-point clustering. **CH:** update to coarser binning

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**Figure 2.** The redshift-space halo bispectrum,  $B(k_1, k_2, k_3)$  as a function of triangle configuration shape for  $\sum m_\nu = 0.0, 0.06, 0.10$ , and  $0.15$  eV (top panels) and  $\sigma_8 = 0.822, 0.818, 0.807$ , and  $0.798$  (lower panels). **CH:** details on the triangle configurations and the colormap. We describe the estimator used to calculate  $B(k_1, k_2, k_3)$  in Section 3.



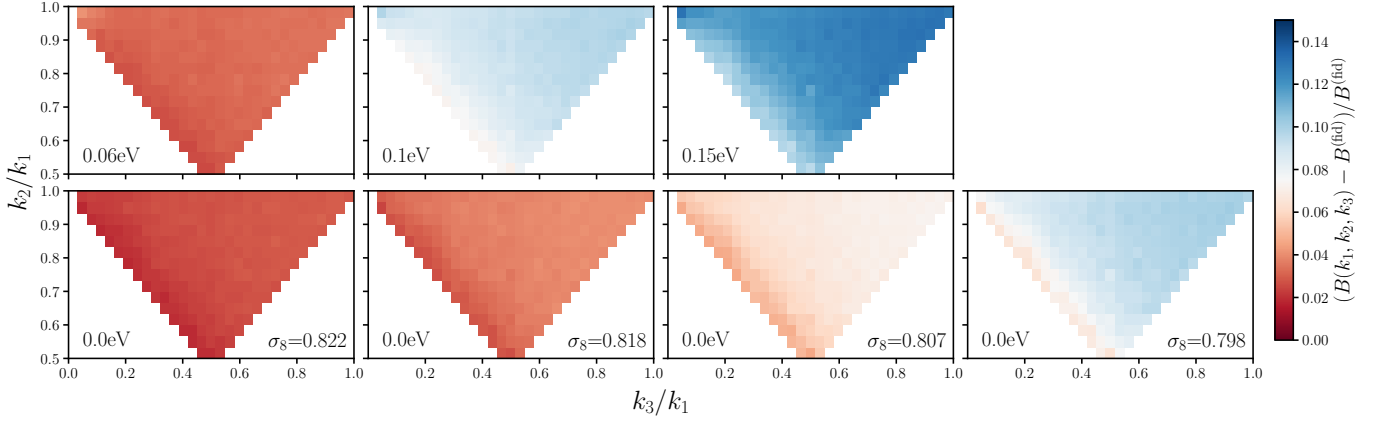
**Figure 3.** The redshift-space halo bispectrum,  $B(k_1, k_2, k_3)$ , as a function of all triangle configurations for  $\sum m_\nu = 0.0, 0.06, 0.10$ , and  $0.15$  eV (top panel) and  $\sigma_8 = 0.822, 0.818, 0.807$ , and  $0.798$  (lower panel). **CH:** details on the ordering of the triangle configurations; also mention how it's roughly scale dependence. We describe the estimator used to calculate  $B(k_1, k_2, k_3)$  in Section 3.

brief description of the hades simulation and the halo catalogs

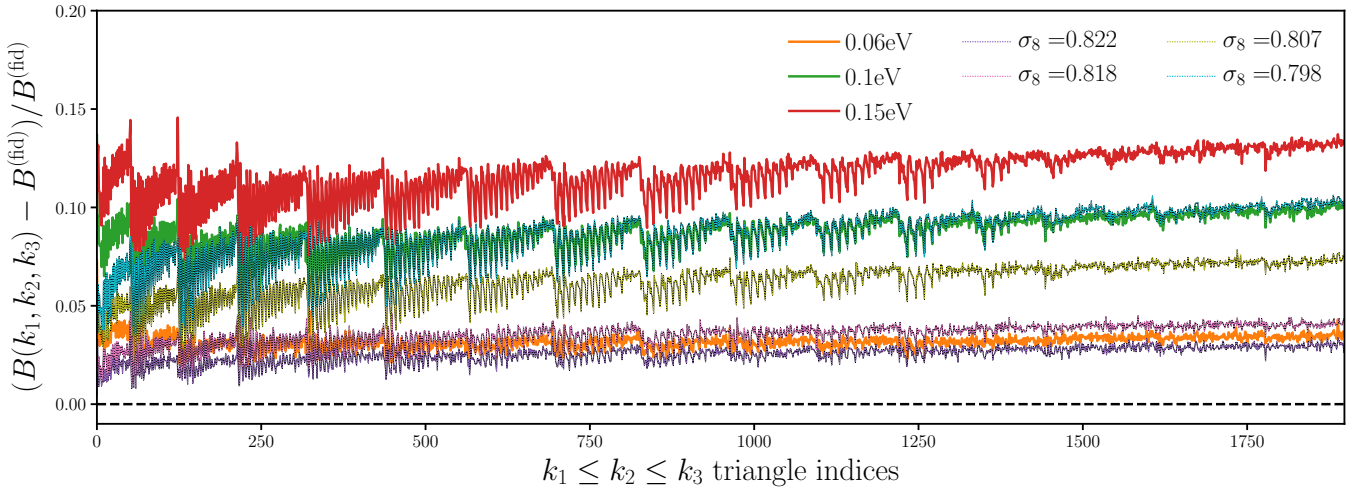
### 3. BISPECTRUM

Brief description of the Scoccimarro et al. bispectrum estimator here

### 4. RESULTS



**Figure 4.** The shape dependence of the  $\sum m_\nu$  and  $\sigma_8$  impact on the redshift-space halo bispectrum,  $\Delta B/B^{(\text{fid})}$ .  $\sum m_\nu = 0.06, 0.10$ , and  $0.15$  eV (top panels; left to right) are aligned with  $\sigma_8 = 0.822, 0.818$ , and  $0.807$  eV (bottom panels; left to right), which produce mostly degenerate imprints on the redshift-space power spectrum. The difference between the top and bottom panels illustrate that  $\sum m_\nu$  induces a significantly different impact on the shape-dependence of the halo bispectrum than  $\sigma_8$ .



**Figure 5.** The impact of  $\sum m_\nu$  and  $\sigma_8$  on the redshift-space halo bispectrum for all triangle configurations:  $\Delta B/B^{(\text{fid})}$ . The impact of  $\sum m_\nu$  differs significantly from the impact of  $\sigma_8$  both in amplitude and scale dependence. For instance,  $\sum m_\nu = 0.15$  eV (red) has a  $\sim 5\%$  stronger impact on the bispectrum than  $\sigma_8 = 0.798$  (cyan dotted), which has little difference in the power spectrum (Figure 1). Combined with the shape-dependence of Figure 4, the contrasting impact of  $\sum m_\nu$  and  $\sigma_8$  on the redshift-space halo bispectrum illustrate that the bispectrum break the degeneracy between  $\sum m_\nu$  and  $\sigma_8$  that degrade constraints from two-point analyses.

## 5. SUMMARY