BGS SV target selection

ChangHoon Hahn (Berkeley)
Omar Ruiz Macias (Durham)
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BGS SV superset target selection criteria

- magnitude limit
- bright object mask
- star-galaxy separation

magnitude limit

The current strategy for the BGS adopts two priority tiers:

Priority 1: $r \leq 19.5$, estimated surface density of $800 \, \mathrm{deg}^{-2}$

Priority 2: $19.5 < r \le 20$, estimated surface density of $600 \, \mathrm{deg}^{-2}$.

for the BGS superset, push the magnitude limit to r < 20.1

DECaLS DR5 Bright stars

SDSS Bright stars

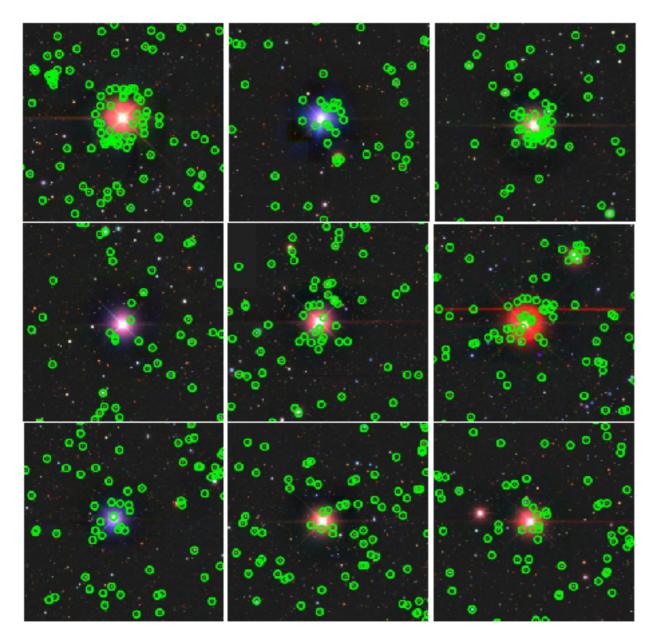


Figure 3.7: Optical images from DECaLS DR5 of the 9 brightest stars from the sample area. The green circles are the BGS objects from DECaLS DR5 with $r \leq 20$. The length of the box is 10 arcmin.

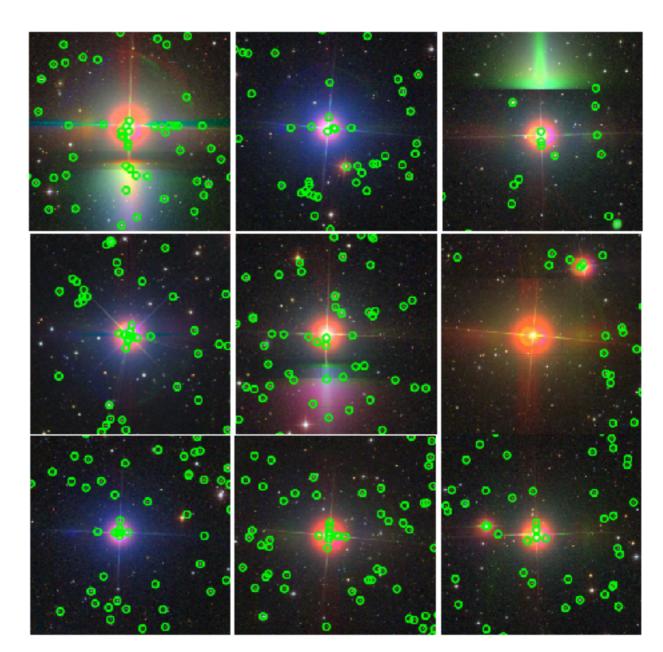
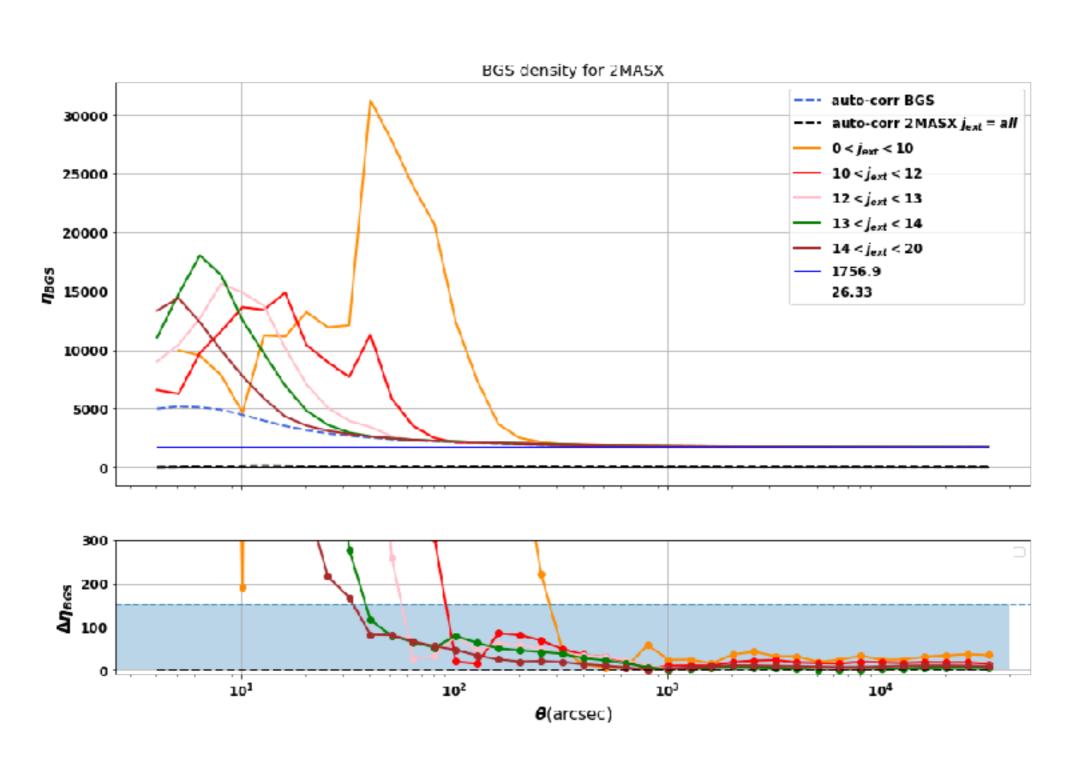


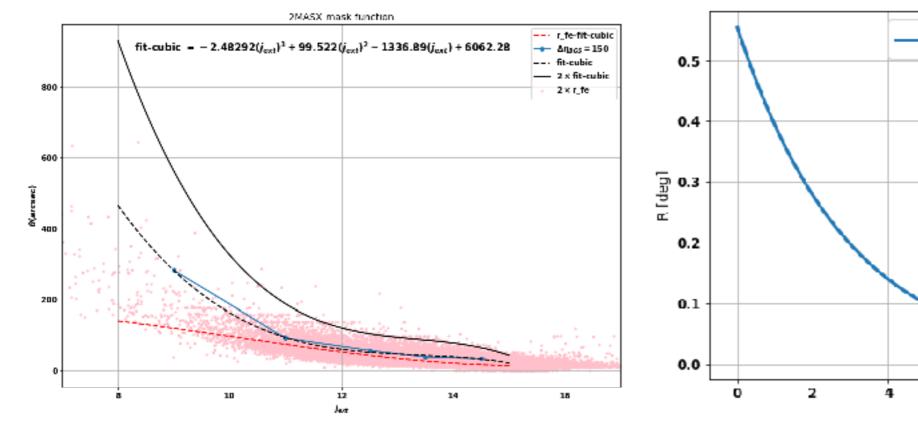
Figure 3.8: Optical images from SDSS of the 9 brightest stars from the sample area. The green circles are the BGS objects from SDSS with $r \le 20$. The length of the box is 10 arcmin.

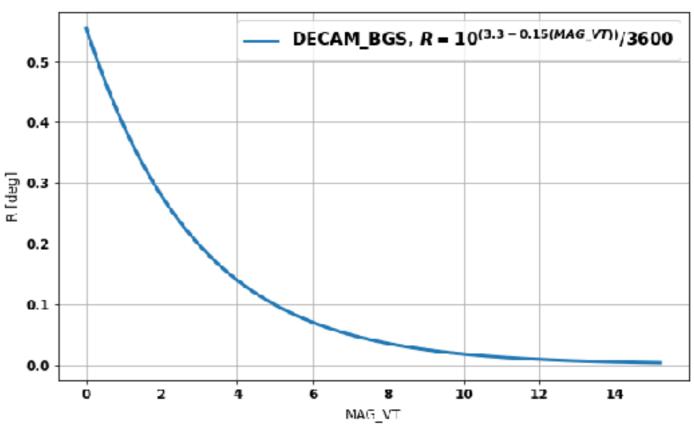
Cross-corr Twomass-DECaLS DR5 BGS



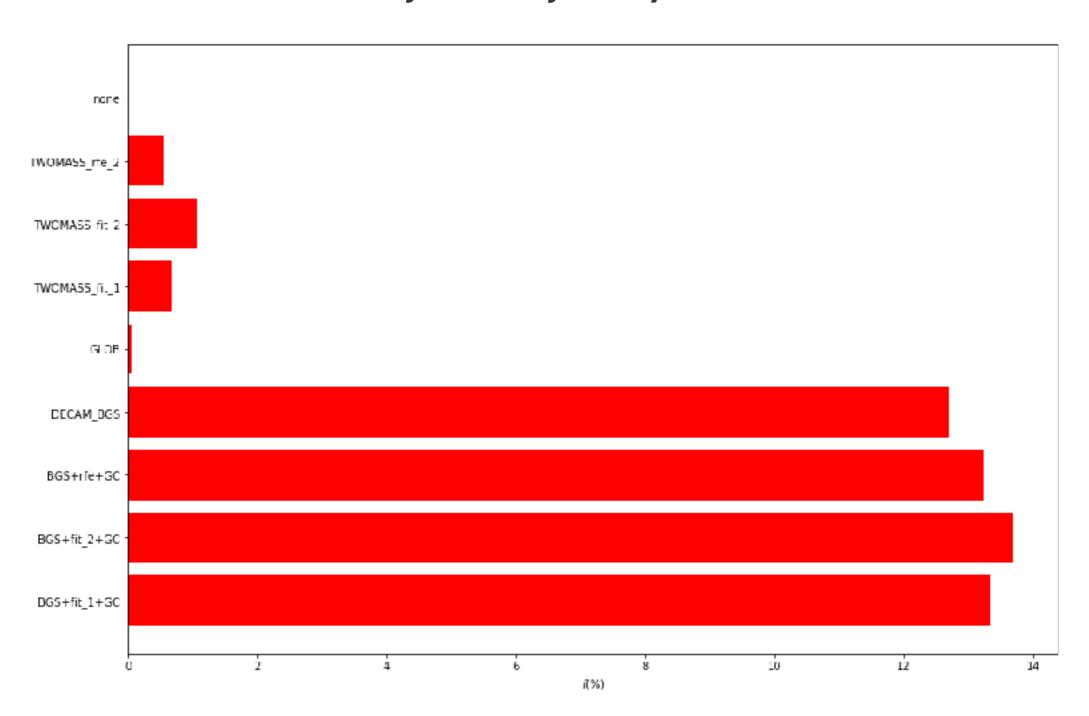
Mask functions for two mass galaxies

Mask function for Tycho2 Stars

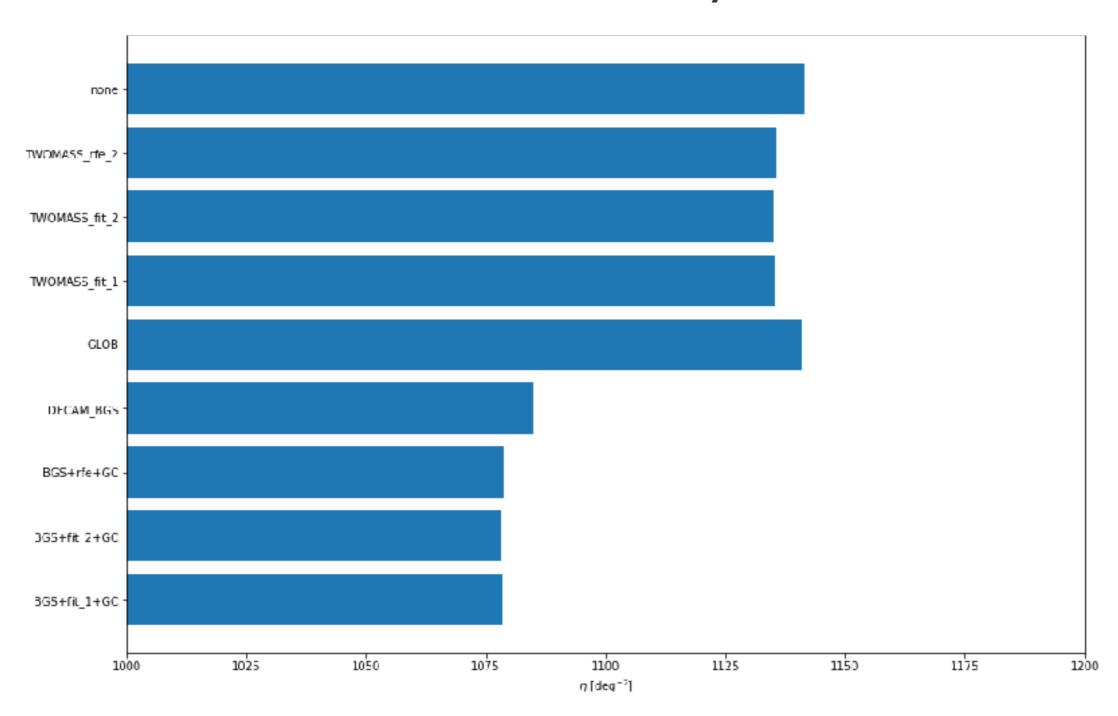




Rejected objects by each mask

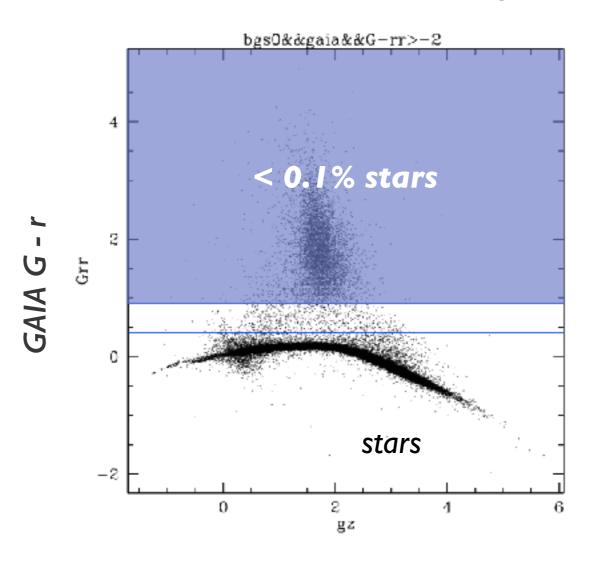


Number density



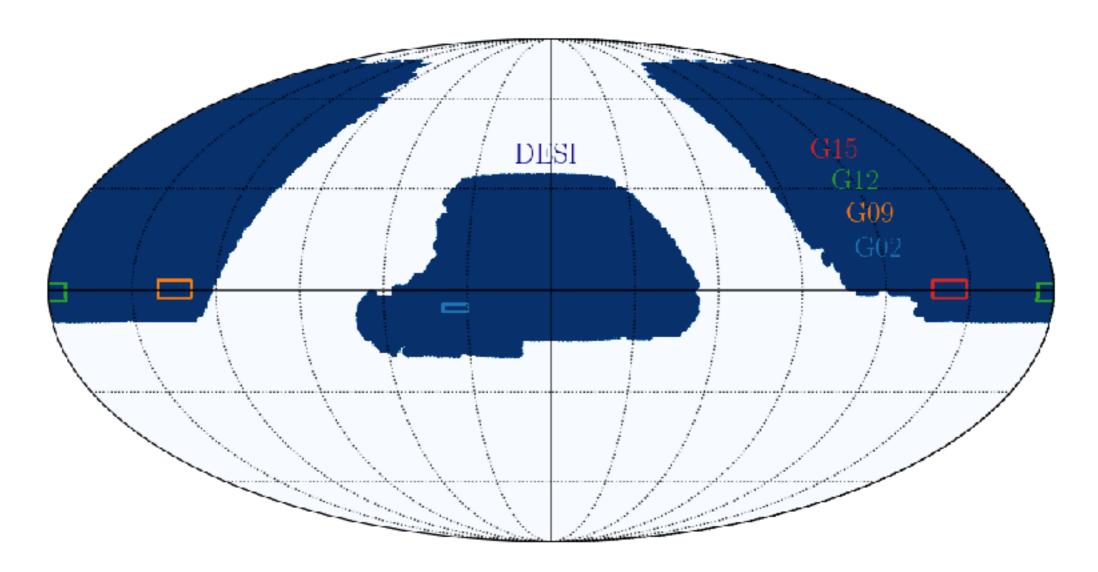
star-galaxy separation





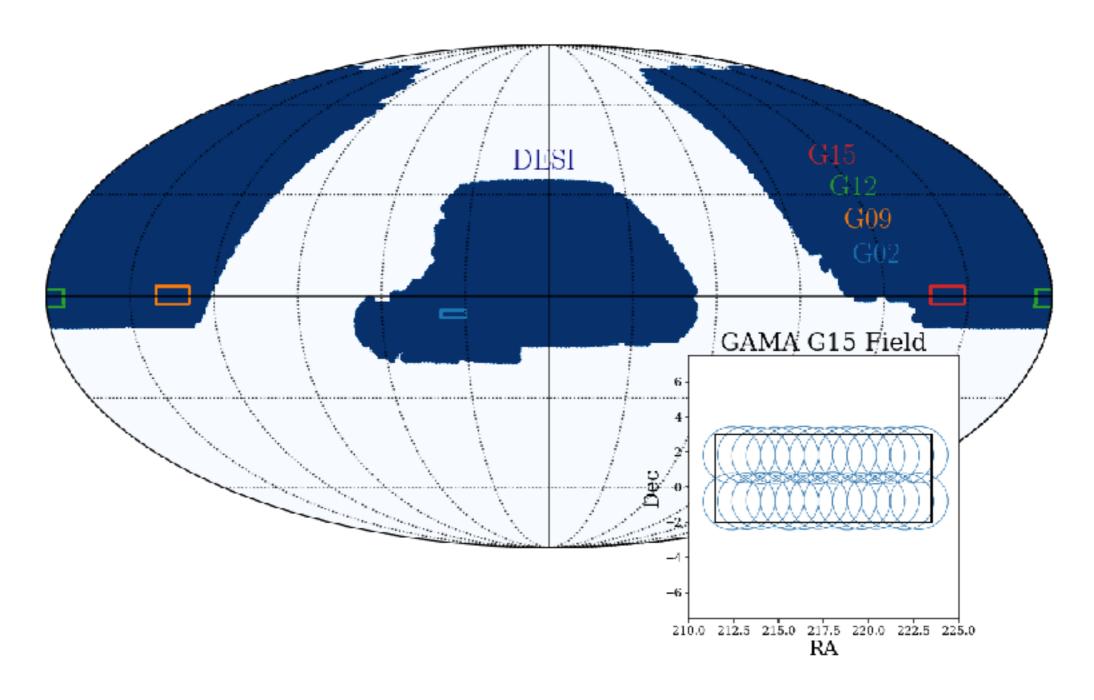
for the BGS superset, allow 10% of objects classified as stars to be targeted to test star-galaxy separation

BGS survey validation

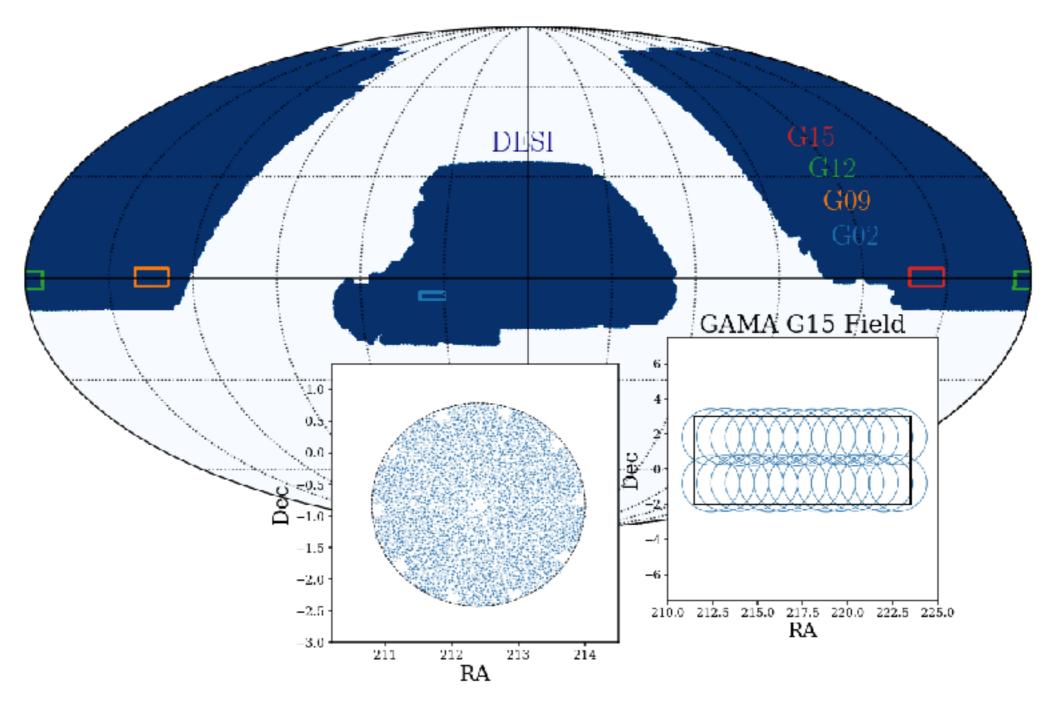


GAMA fields — spectroscopic truth tables

BGS survey validation



BGS survey validation



DECaLS DR5 targets by Adam Myers

fiberassign by Jaime Forero-Romero https://github.com/forero/fiberassign_sv/