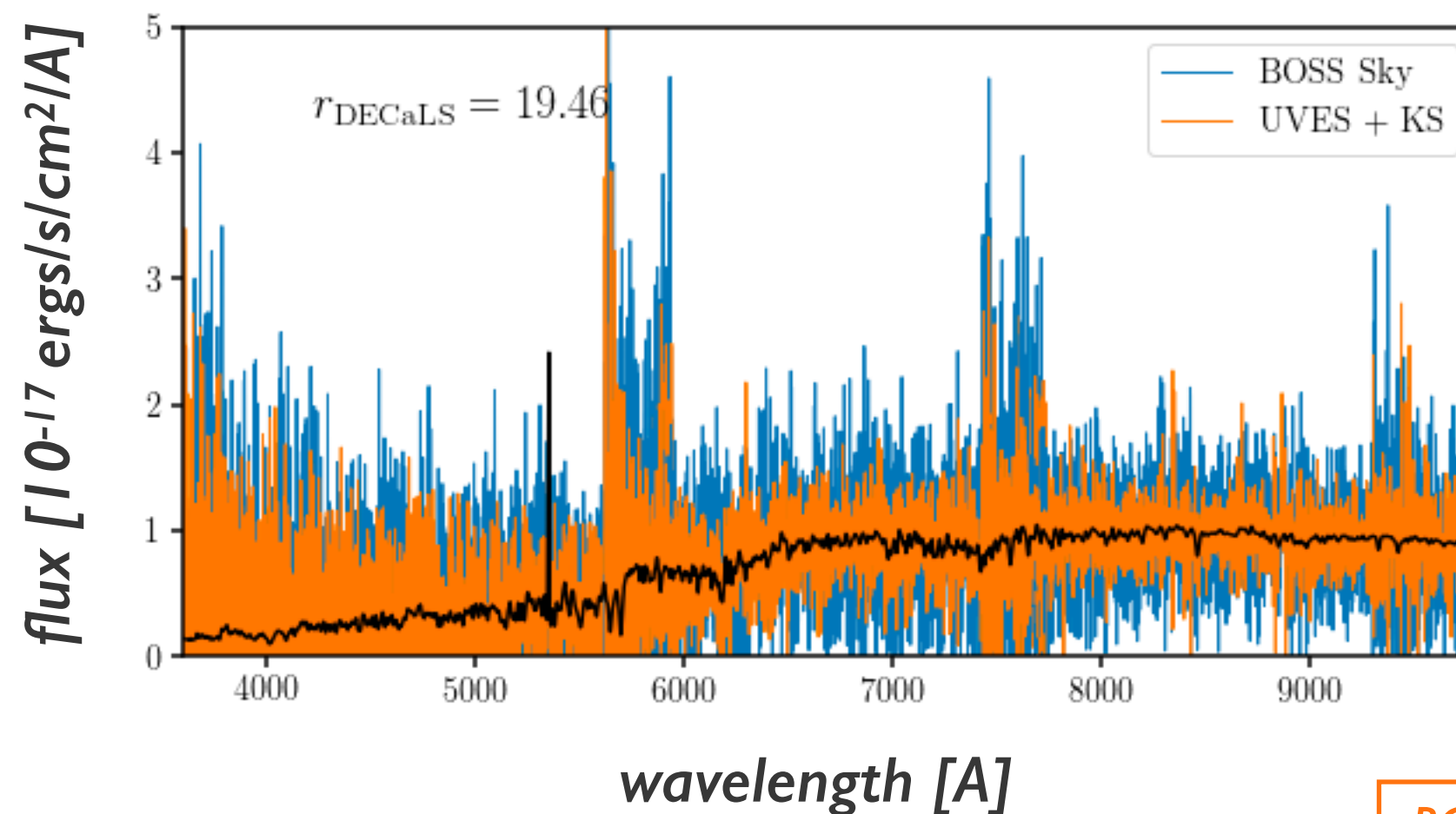


updates on the **bright time sky model**

changhoon hahn

berkeley // postdoc // changhoonhahn@lbl.gov

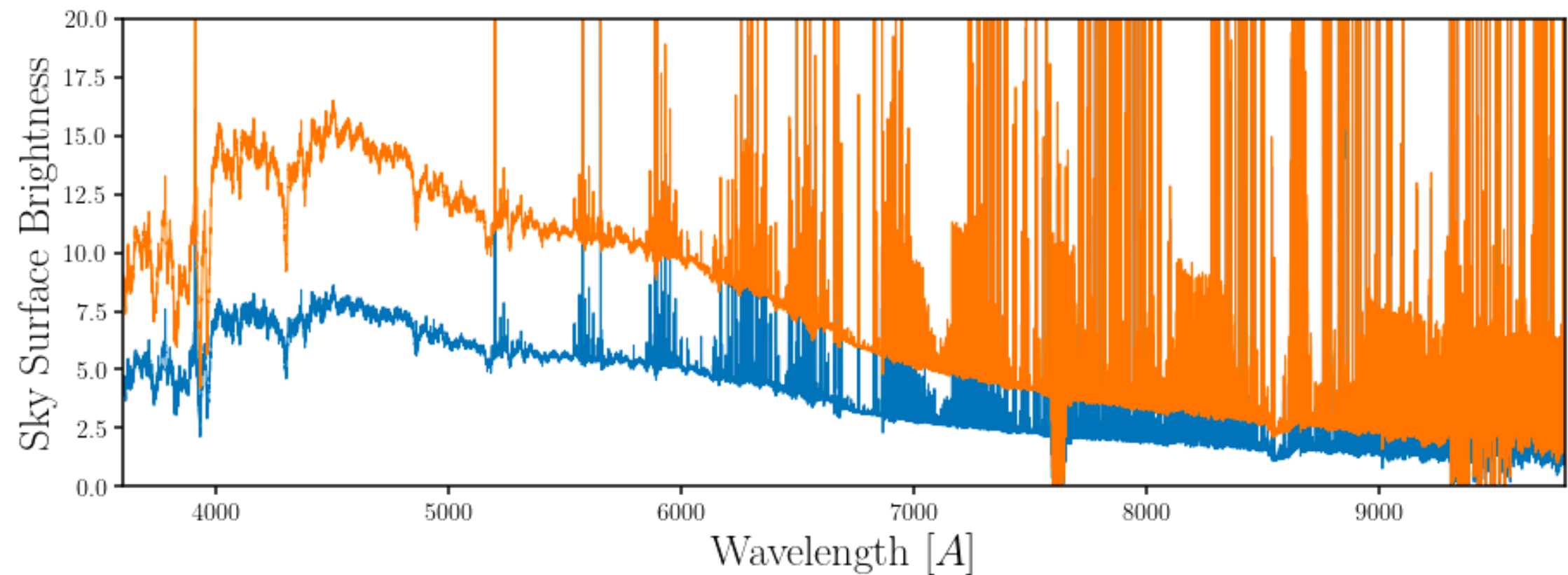
the original sky model (UVES dark sky + Krisciunas & Schaefer 1991)
underestimates the sky brightness



BGS galaxy spectrum
w/ original sky model

BGS galaxy spectrum
w/ sky flux from BOSS fiber

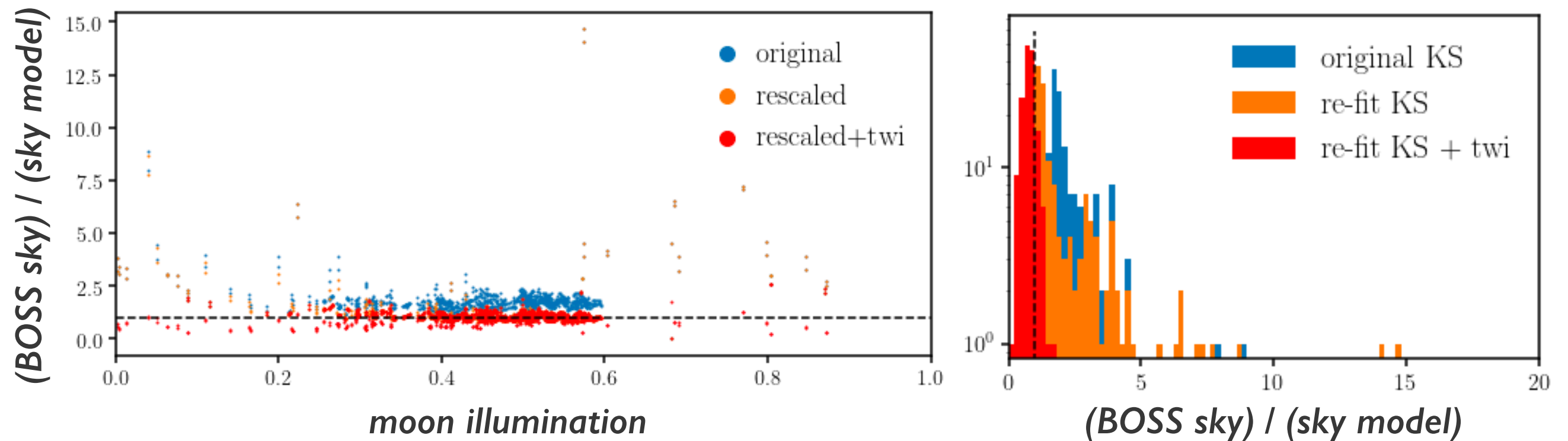
new sky model = (UVES dark sky) + (re-fit Krisciunas & Schaefer 1991) + (twilight)



new sky model

old sky model

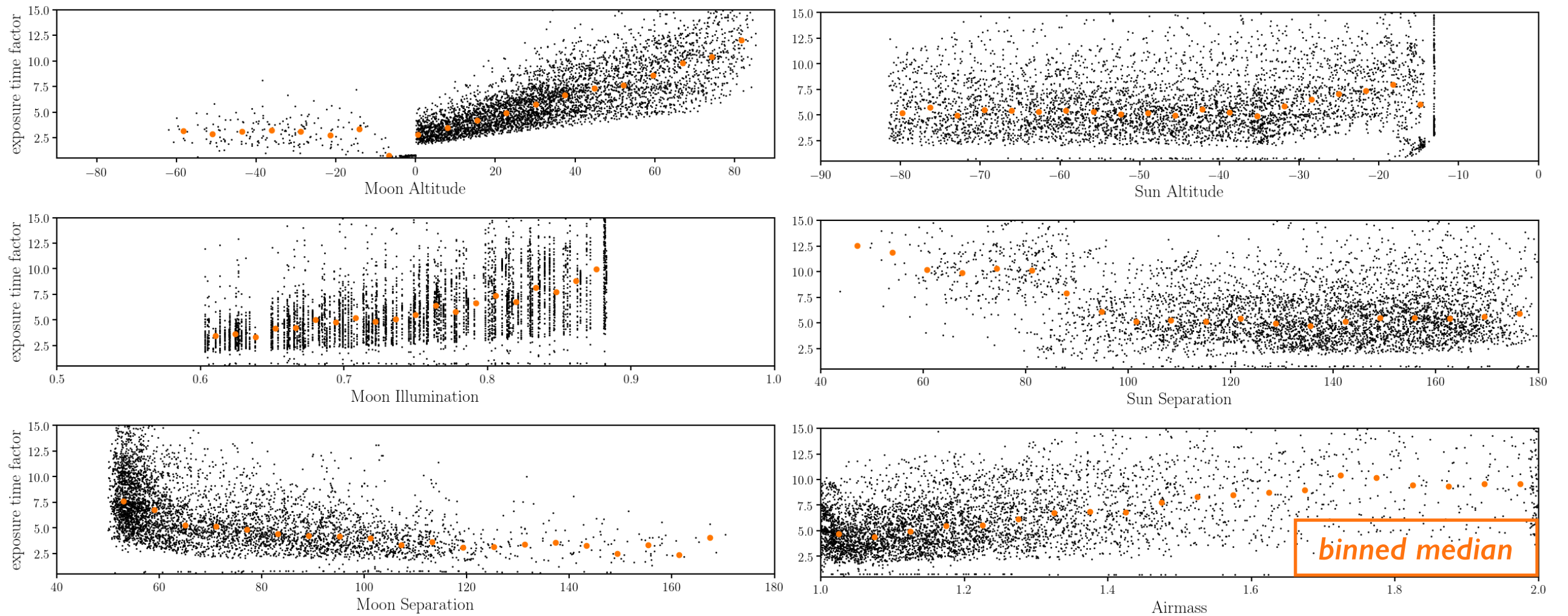
new sky model = (UVES dark sky) + (re-fit Krisciunas & Schaefer 1991) + (twilight)



better reproduces sky flux from BOSS sky fibers and DECam sky exposures

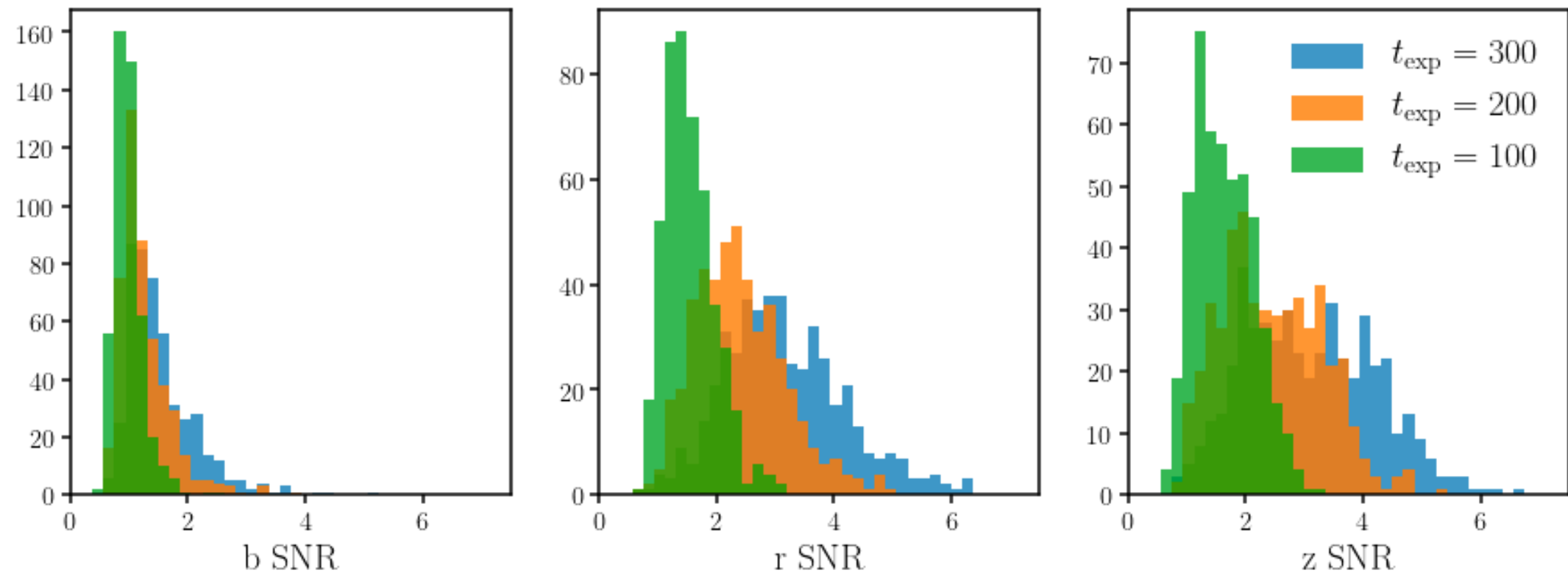
exposure time correction factor from the *new sky model*:

$$\frac{\text{bright exposure time}}{\text{nominal exposure time}} = \frac{\text{bright sky}}{\text{nominal dark sky}}$$



t_{exp} can be estimated for bright times in *surveysim* using
exposure time correction factor Gaussian Process emulator

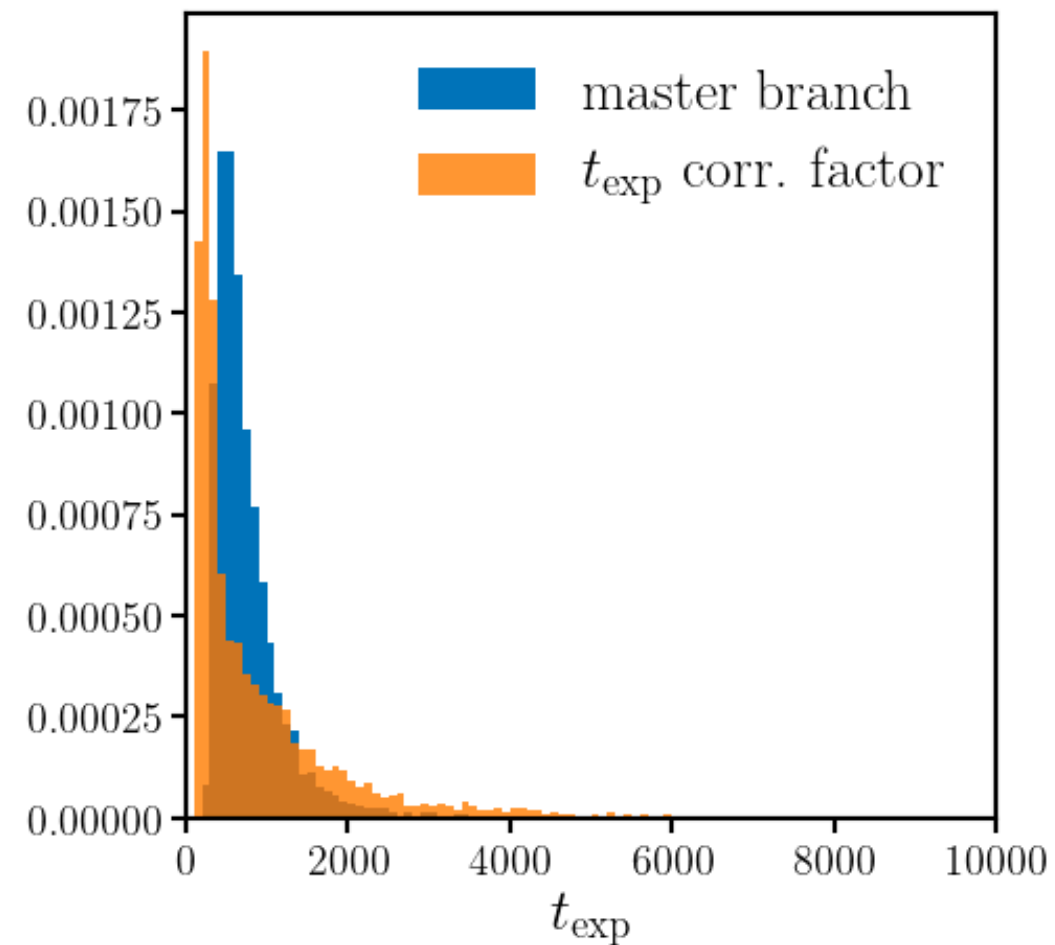
current default nominal $t_{\text{exp}} = 300$ sec but $t_{\text{exp}} \sim 150$ sec may be enough for BGS



SNR distribution of ~ 500 $r \sim 19.5$ galaxy spectra observed during nominal dark sky for $t_{\text{exp}} = 300, 200$, and 100 sec

*assuming redshift completeness is high for $r \sim 19.5$ galaxies with $\text{SNR} \sim 1$

more realistic *surveysim* for **bright time exposures**



original *surveysim*
fixed t_{exp} correction

w/ bright sky t_{exp} correction
nominal $t_{\text{exp}} = 150$ sec

exposure time distribution for bright exposures of *surveysim* outputs