

UltraLight Dark Matter in Simulations: A Chapel-Powered Eigenstate Perspective

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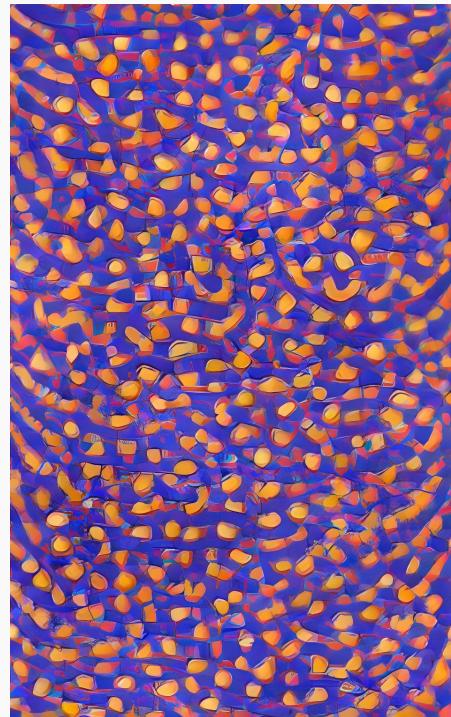
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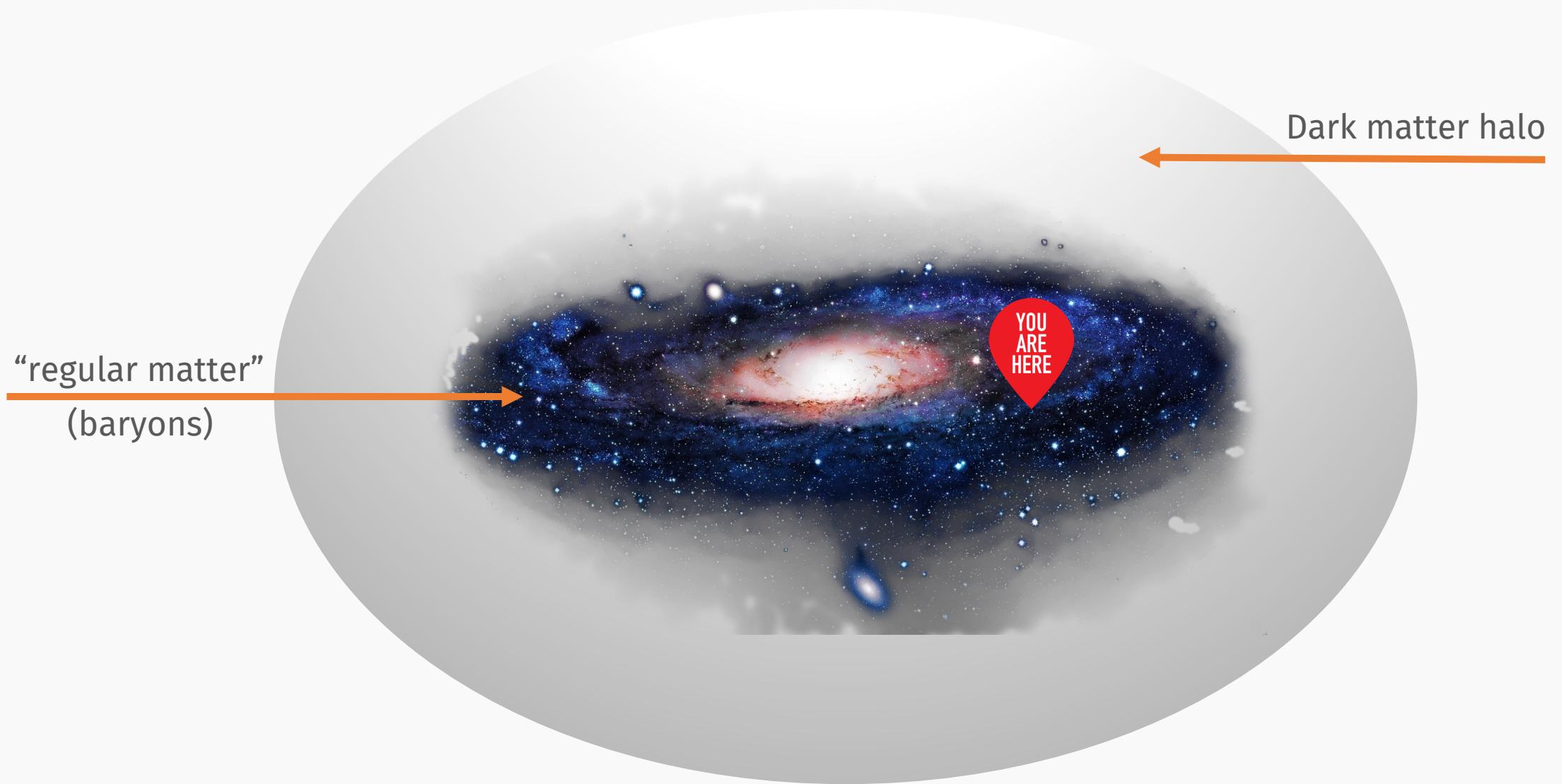
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(UltraLight) Dark Matter

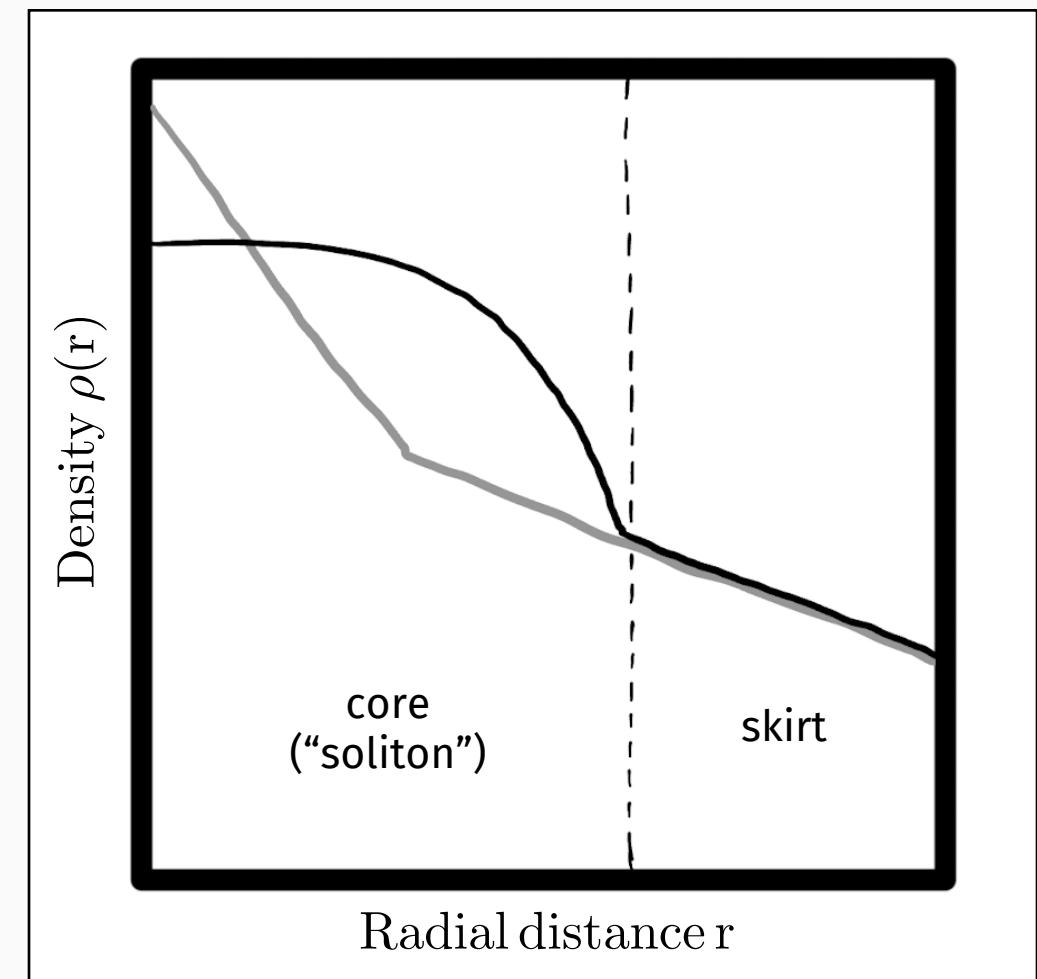
Dark Matter Halos



UltraLight Dark Matter

UltraLight Dark Matter (ULDM):

- is an axion-like scalar boson
- has low mass: $\sim 10^{-22}$ eV
- helps with small-scale problems
 - core-cusp problem (*right*)
 - missing satellites problem
 - too-big-to-fail problem
- cool phenomenology!



CHPLULTRA

The Schrödinger-Poisson System

We need a solver for the following set of equations:

$$i\hbar\dot{\psi} = -\frac{\hbar^2}{2m_a}\nabla^2\psi + m_a\Phi\psi$$

$$\nabla^2\Phi = 4\pi G m_a |\psi|^2$$

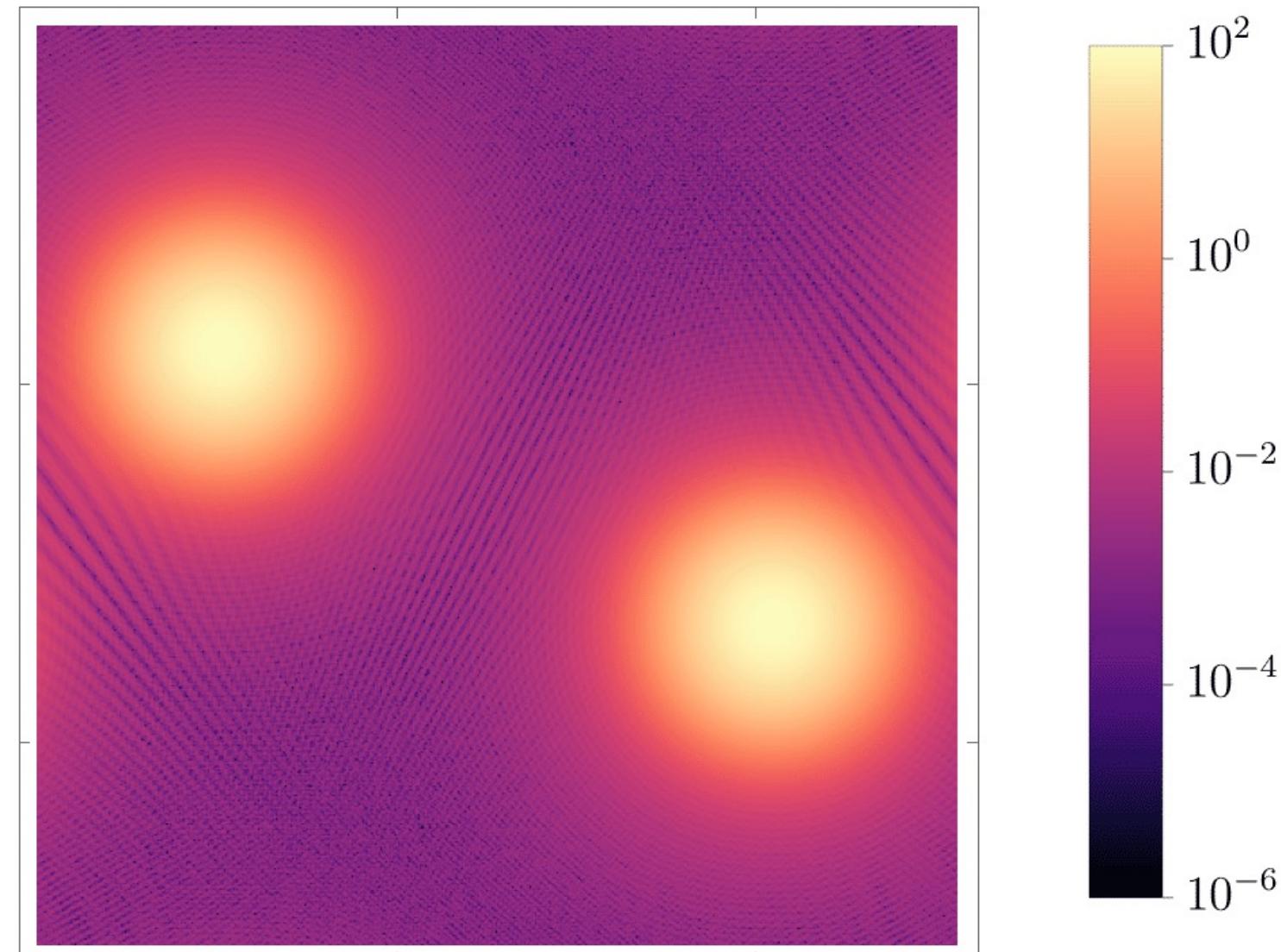
$$\rho = m_a |\psi|^2$$

PyUltraLight vs CHPLULTRA

PyUltraLight:

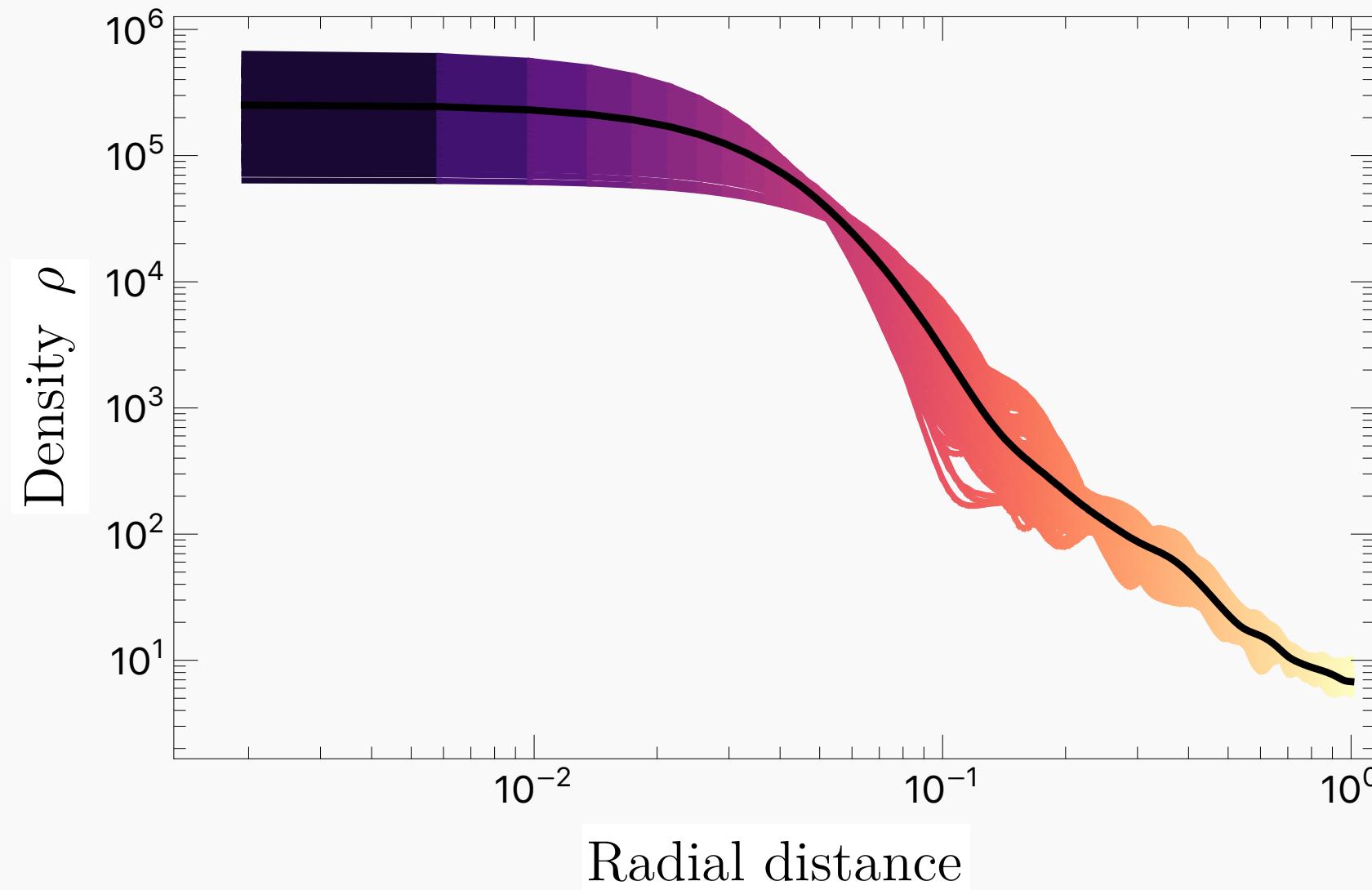
- original implementation in Python
- relies on FFTW wrappers (PyFFTW)
- excellent numerical stability
- parallelized, not distributed
- great for PC use!

A Chapel-powered Halo



A Halo's Eigenstates

The key assumption

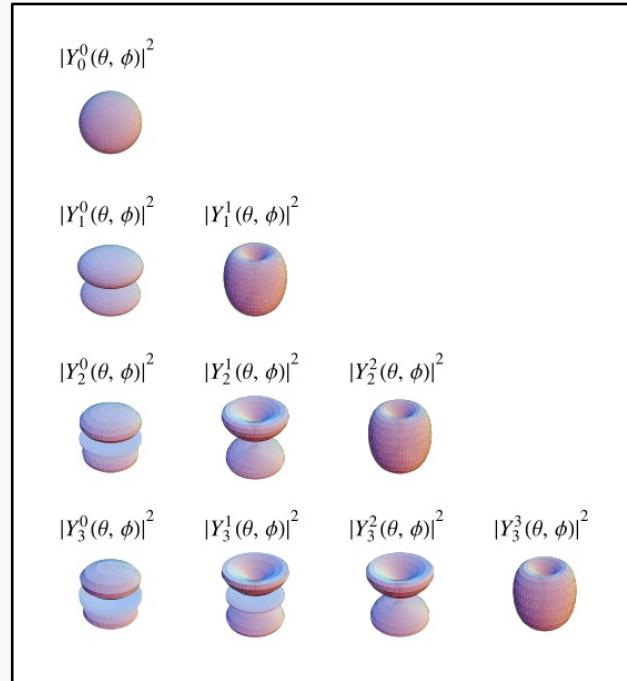
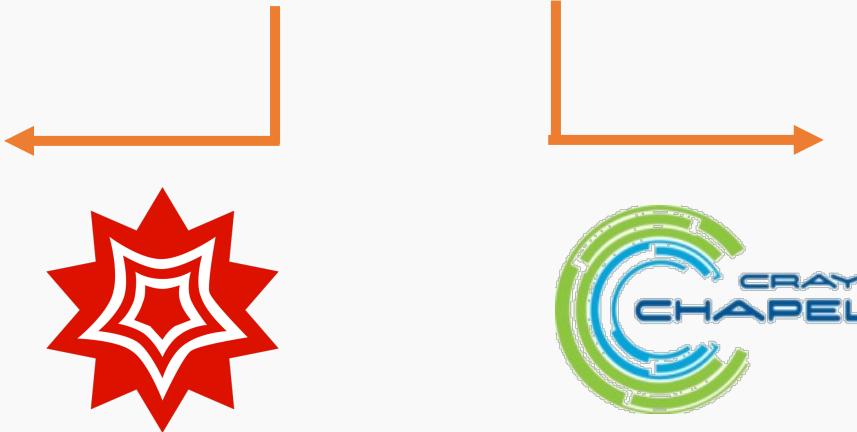
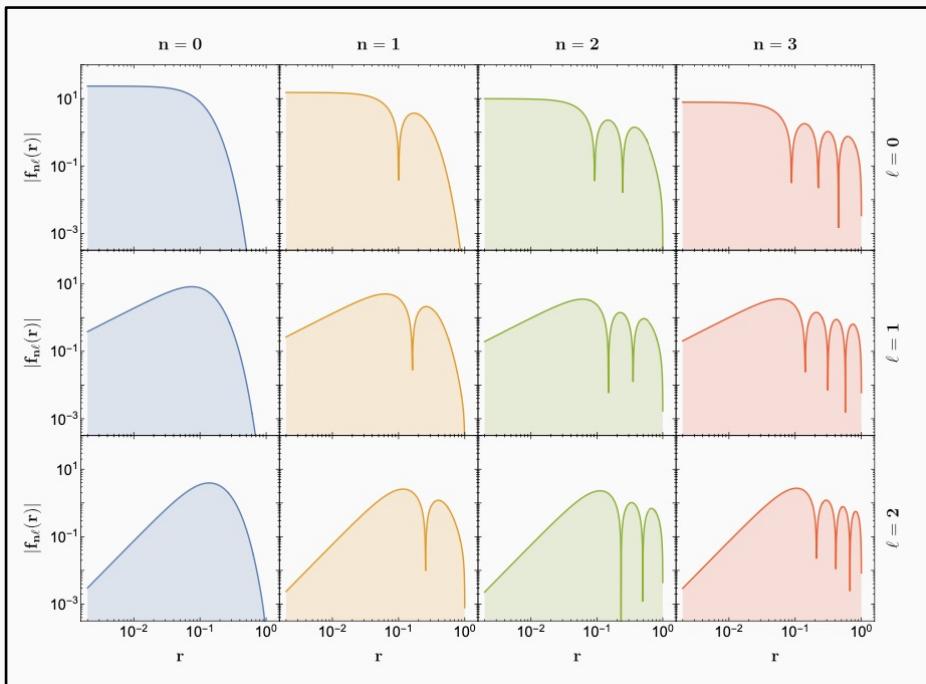


Calculating the Eigenstates

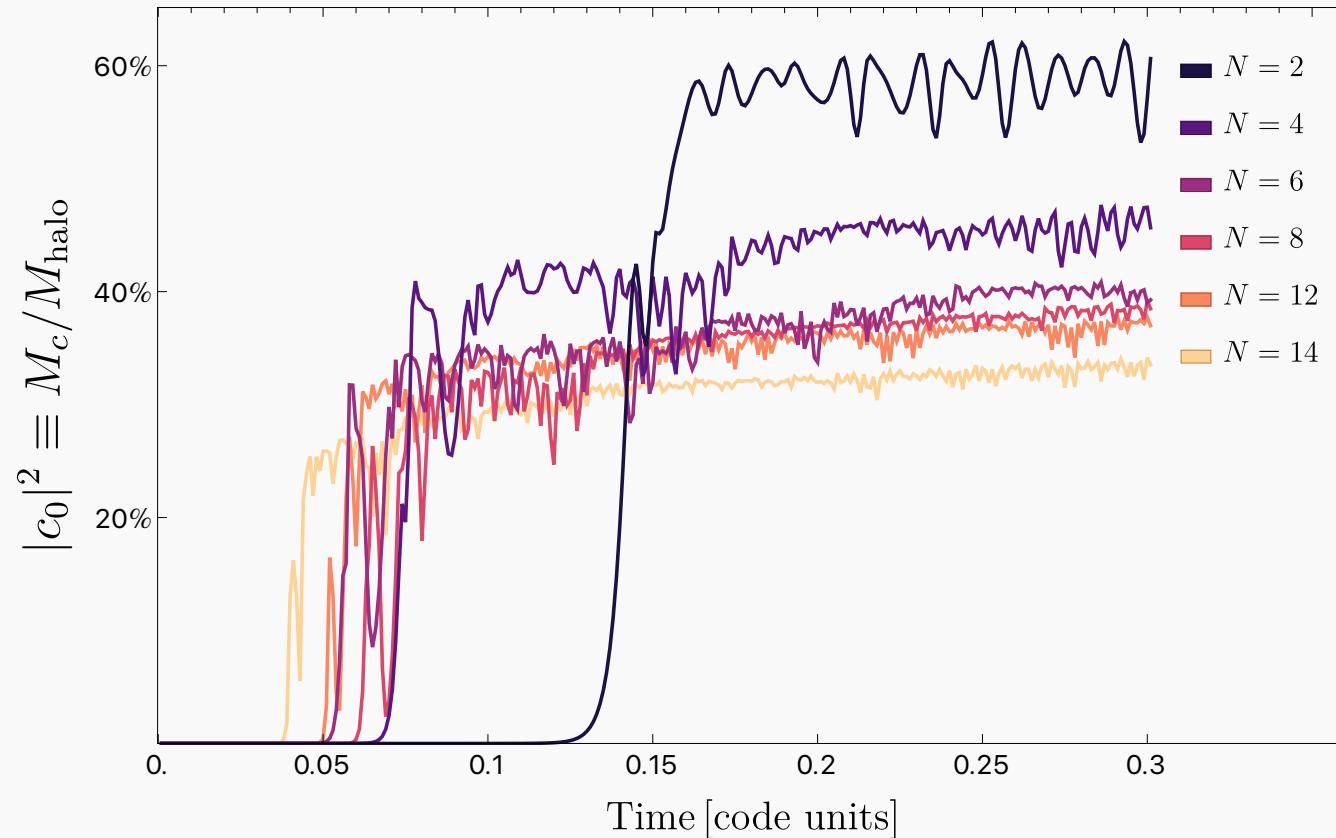
Separable into radial and spherical components:

$$\Psi_{nlm}(r, \phi, \theta) = f_{nl}(r) \times Y_{lm}(\phi, \theta)$$

arxiv: 2109.01920



How Eigenstates Reveal Halo Formation



Conclusions & Future Work

Takeaways

Dark Matter

It's mysterious, but robust computational modeling can help!

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Is fast and efficient thanks to FFTW wrappers and HPC-friendliness of Chapel

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Eigenstates of ULDM halos

A new approach to getting insight into a popular proposed dark matter candidate

Thank you for listening!
Questions?

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