Imperial College London

Notes

IMPERIAL COLLEGE LONDON

DEPARTMENT OF PHYSICS

Advanced Classical Physics

Author:

Chen Huang

Email:

chen.huang23@imperial.ac.uk

Date: November 6, 2023

CONTENTS CONTENTS

Contents

1 Lagrangian Mechanics			n Mechanics	3	
	1.1	Action	Principle	3	
		1.1.1	Fermat's Principle	3	
		1.1.2		4	
		1.1.3		4	
	1.2	Gener	alized coordinates and momenta	5	
	1.3				
		1.3.1	Momentum Conservation	5	
		1.3.2	Energy (or Hamiltonian) Conservation	6	
	1.4	Consti		ϵ	
		1.4.1		ϵ	
		1.4.2		7	
		1.4.3	Example: Helter Skelter	7	
	1.5	Norma		8	
		1.5.1	•	9	
		1.5.2		9	
		1.5.3	Gram-Schmidt Diagonalization		
		1.5.4	Cholesky Decomposition		
		1.5.5	Example: Double Pendulum	2	
	1.6				
		1.6.1	Discrete Global Symmetry		
		1.6.2	Continuous Global Symmetry		
		1.6.3	Noether's theorem		
		1.6.4	Hamiltonian as the Noether Charge for Time Translations 1	6	
2	Hom	niltonia	an 1	_	
4	пан 2.1		tonian Formulation		
	2.1	2.1.1	From Lagrangian to Hamiltonian		
		2.1.1	Hamilton's Equations		
		2.1.2	±		
	2.2		ton's Principle of Least Action		
	2.3			8	
	۵.5	2.3.1			
		2.3.2	Properties of Poisson Brackets		
	2.4				
	'	2.4.1	ical Transformations		

1 Lagrangian Mechanics

1.1 Action Principle

1.1.1 Fermat's Principle

The idea of the 'principle of least action' has its origin in Fermat's principle in optics, according to which light follows the shortest optical path, i.e., the path of shortest time to reach its destination.

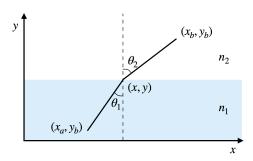


Figure 1: Snell's law of refraction for light passing through media of different indices of refraction.

According to figure (1), consider a light ray from point (x_a, y_a) to (x_b, y_b) , the optical path length follows

$$T(x) = \frac{n_1}{c} \sqrt{(x_a - x)^2 + (y_a - y)^2} + \frac{n_2}{c} \sqrt{(x_b - x)^2 + (y_b - y)^2}$$
 (1)

According to Fermat's principle, we need to find the minimum of this quantity.

$$\frac{\partial T(x)}{\partial x} = \frac{n_1}{c} \frac{x - x_a}{\sqrt{(x_a - x)^2 + (y_a - y)^2}} + \frac{n_2}{c} \frac{x - x_b}{\sqrt{(x_b - x)^2 + (y_b - y)^2}}
= -\frac{n_1}{c} \sin \theta_1 + \frac{n_2}{c} \sin \theta_2 = 0$$
(2)

which shows the Snell's law

$$n_1 \sin \theta_1 = n_2 \sin \theta_2 \tag{3}$$

What happens if the refractive index is a function of space n(x, y)? The Fermat's action becomes

$$T = \int dT = \int \frac{n(x,y)}{c} \sqrt{(dx)^2 + (dy)^2}$$
(4)

Suppose we parameterise the trajectory of the particle by a monotonic parameter λ

$$\mathbf{r}(\lambda) = (x(\lambda), y(\lambda)) \tag{5}$$

such that

$$x(\lambda_a) = x_a, \quad x(\lambda_b) = x_b, \quad y(\lambda_a) = y_a, \quad y(\lambda_b) = y_b$$
 (6)

The Fermat's action is

$$T = \int_{\lambda_a}^{\lambda_b} d\lambda \frac{n(x(\lambda), y(\lambda))}{c} \sqrt{\left(\frac{dx}{d\lambda}\right)^2 + \left(\frac{dy}{d\lambda}\right)^2}$$
 (7)

1.1.2 Euler-Lagrange Equations

In configuration space (the space of coordinates and velocities), the action S is defined as an integral over a function $L(x, \dot{x}, t)$ known as the Lagrangian, as

$$S[x(t)] = \int_{t_a}^{t_b} \mathrm{d}t L(x(t), \dot{x}(t), t) \tag{8}$$

For standard conservative systems the Lagrangian is simply the difference of the kinetic energy T and the potential energy V, i.e.

$$\boxed{L = T - V} \tag{9}$$

The actual physical trajectory is the function x that minimises the action subject to the boundary conditions $x(t_a) = x_a$ and $x(t_b) = x_b$. This perturbation changes the action by an amount δS given by

$$\delta S[x(t)] = \int_{t_a}^{t_b} dt \left[\frac{\partial L}{\partial x} \delta x(t) + \frac{\partial L}{\partial \dot{x}} \delta \dot{x}(t) \right]$$
 (10)

In the second term, we may integrate by parts,

$$\int_{t_a}^{t_b} dt \frac{\partial L}{\partial \dot{x}} \delta \dot{x}(t) = \left[\frac{\partial L}{\partial \dot{x}} \delta x(t) \right]_{t_a}^{t_b} - \int_{t_a}^{t_b} dt \frac{d}{dt} \left(\frac{\partial L}{\partial \dot{x}} \right) \delta x(t) = - \int_{t_a}^{t_b} dt \frac{d}{dt} \left(\frac{\partial L}{\partial \dot{x}} \right) \delta x(t)$$
(11)

Substituting this result into the expression (10) for δS , we have

$$\delta S = \int_{t_a}^{t_b} dt \left[\frac{\partial L}{\partial x} - \frac{d}{dt} \left(\frac{\partial L}{\partial \dot{x}} \right) \right] \delta x(t)$$
 (12)

In order that S be stationary, we require

$$\boxed{\frac{\partial L}{\partial x} - \frac{\mathrm{d}}{\mathrm{d}t} \left(\frac{\partial L}{\partial \dot{x}} \right) = 0}$$
 (13)

This is known as the *Euler-Lagrange equation*, and it is the equation of motion in the Lagrangian formulation of mechanics.

1.1.3 Back to Fermat

In the gauge $\lambda = y$, the total time was given to

$$T = \int_{y_a}^{y_b} dy \frac{n(x(y), y)}{c} \sqrt{\left(\frac{dx}{dy}\right)^2 + 1}$$
 (14)

The Lagrangian is

$$L(x(y), \dot{x}(y), y) = \frac{n(x(y), y)}{c} \sqrt{\left(\frac{\mathrm{d}x}{\mathrm{d}y}\right)^2 + 1}$$
 (15)

The Euler-Lagrange equation for this problem is

$$\frac{\mathrm{d}}{\mathrm{d}y} \left(\frac{\partial L}{\partial \frac{\mathrm{d}x}{\mathrm{d}y}} \right) = \frac{\partial L}{\partial x} \tag{16}$$

which gives

$$\frac{\mathrm{d}}{\mathrm{d}y} \left[n(x,y) \frac{\frac{\mathrm{d}x}{\mathrm{d}y}}{\sqrt{\left(\frac{\mathrm{d}x}{\mathrm{d}y}\right)^2 + 1}} \right] = \frac{\partial n(x,y)}{\partial x} \sqrt{\left(\frac{\mathrm{d}x}{\mathrm{d}y}\right)^2 + 1}$$
(17)

it simplifies a lot in the case where the refractive index is just a function of y. Then

$$\frac{\mathrm{d}}{\mathrm{d}y} \left[n(y) \frac{\frac{\mathrm{d}x}{\mathrm{d}y}}{\sqrt{\left(\frac{\mathrm{d}x}{\mathrm{d}y}\right)^2 + 1}} \right] = 0 \tag{18}$$

which is easily solved as

$$n(y)\frac{\frac{\mathrm{d}x}{\mathrm{d}y}}{\sqrt{\left(\frac{\mathrm{d}x}{\mathrm{d}y}\right)^2 + 1}} = n(y)\frac{\mathrm{d}x}{\sqrt{(\mathrm{d}x)^2 + (\mathrm{d}y)^2}} = n(y)\sin(\theta(y)) = A \tag{19}$$

with A a constant.

1.2 Generalized coordinates and momenta

General coordinate q_i

General conjugate momentum $p_i = \frac{\partial L}{\partial \dot{q}_i}$

$$\delta S = \int dt \left(\sum_{i=1}^{N} \frac{\partial L}{\partial q_i} \delta q_i + \sum_{i=1}^{N} \frac{\partial L}{\partial \dot{q}_i} \delta \dot{q}_i + \frac{\partial L}{\partial t} \right)$$
 (20)

$$\frac{\partial L}{\partial q_i} - \frac{\mathrm{d}}{\mathrm{d}t} \left(\frac{\partial L}{\partial \dot{q}_i} \right) = 0, \quad \forall i = 1, \dots, N$$
(21)

1.3 Conservation Laws

1.3.1 Momentum Conservation

For any generalised coordinate q_i , we define the generalised momentum p_i by

$$p_i = \frac{\partial L}{\partial \dot{q}_i} \tag{22}$$

So we have

$$\frac{\partial L}{\partial q_i} = \frac{\mathrm{d}}{\mathrm{d}t} \left(\frac{\partial L}{\partial \dot{q}_i} \right) = \frac{\mathrm{d}p_i}{\mathrm{d}t}$$
 (23)

Whenever the Lagrangian L does not depend *explicitly* on q_i , the corresponding generalised momentum p_i is conserved

$$\boxed{\frac{\partial L}{\partial q_i} = 0 \quad \Leftrightarrow \quad p_i \text{ is conserved.}}$$
 (24)

1.3.2 Energy (or Hamiltonian) Conservation

First consider the total time derivative of the Lagrangian

$$\frac{\mathrm{d}L}{\mathrm{d}t} = \sum_{i} \frac{\partial L}{\partial q_{i}} \dot{q}_{i} + \sum_{i} \frac{\partial L}{\partial \dot{q}_{i}} \ddot{q}_{i} + \frac{\partial L}{\partial t}$$

$$= \sum_{i} \frac{\mathrm{d}}{\mathrm{d}t} \left(\frac{\partial L}{\partial \dot{q}_{i}} \right) \dot{q}_{i} + \sum_{i} \frac{\partial L}{\partial \dot{q}_{i}} \ddot{q}_{i} + \frac{\partial L}{\partial t}$$

$$= \sum_{i} \frac{\mathrm{d}}{\mathrm{d}t} \left(\frac{\partial L}{\partial \dot{q}_{i}} \dot{q}_{i} \right) + \frac{\partial L}{\partial t}$$
(25)

Rearranging this equation

$$\frac{\partial L}{\partial t} = -\frac{\mathrm{d}}{\mathrm{d}t} \left(\sum_{i} \frac{\partial L}{\partial \dot{q}_{i}} \dot{q}_{i} - L \right) = -\frac{\mathrm{d}}{\mathrm{d}t} \left(\sum_{i} p_{i} \dot{q}_{i} - L \right) = -\frac{\mathrm{d}H}{\mathrm{d}t}$$
 (26)

Whenever the Lagrangian does not depend *explicitly* on t, the Hamiltonian is conserved

$$\frac{\partial L}{\partial t} = 0 \quad \Leftrightarrow \quad H = \sum_{i} p_{i} \dot{q}_{i} - L \text{ is conserved.}$$
 (27)

1.4 Constraints and Number of degrees of Freedom

1.4.1 Kinetic Matrix

Using the chain rule, we have

$$\frac{\mathrm{d}}{\mathrm{d}t} \left(\frac{\partial L}{\partial \dot{q}_i} \right) = \sum_j \frac{\partial^2 L}{\partial \dot{q}_i \partial q_j} \dot{q}_j + \sum_j \frac{\partial^2 L}{\partial \dot{q}_i \partial \dot{q}_j} \ddot{q}_j + \frac{\partial^2 L}{\partial \dot{q}_i \partial t} = \frac{\partial L}{\partial q_i}$$
(28)

We can write in this form

$$\sum_{j} \mathcal{Z}^{ij} \ddot{q}_j + \mathcal{F}^i = 0$$
 (29)

where the kinetic matrix \mathcal{Z}_{ij} and the vector \mathcal{F}_i are both functions of the coordinates

$$\mathcal{Z}^{ij}(q_1, \cdots, q_N, \dot{q}_1, \cdots, \dot{q}_N, t) = \frac{\partial^2 L}{\partial \dot{q}_i \partial \dot{q}_i} = \mathcal{Z}^{ji}$$
(30)

$$\mathcal{F}^{i}(q_{1}, \cdots, q_{N}, \dot{q}_{1}, \cdots, \dot{q}_{N}, t) = \frac{\partial^{2} L}{\partial \dot{q}_{i} \partial q_{j}} \dot{q}_{j} + \frac{\partial^{2} L}{\partial \dot{q}_{i} \partial t} - \frac{\partial L}{\partial q_{i}}$$
(31)

So long as the symmetric matrix \mathcal{Z}_{ij} is non-degenerate, i.e. so long as $\det(\mathcal{Z}) \neq 0$

$$\ddot{q} = -(\mathcal{Z})^{-1}\mathcal{F} \tag{32}$$

There are cases however where the matrix \mathcal{Z} is 'degenerate' and $\det(\mathcal{Z}) = 0$ means that not all coordinates q_i are independent.

1.4.2 Lagrange Multipliers

$$\tilde{L}(q_i, \dot{q}_i, \lambda, t) = L(q_i, \dot{q}_i, t) + \lambda f(q_i, \dot{q}_i, t)$$
(33)

The Euler-Lagrange equation for λ is given by

$$\frac{\delta \tilde{S}}{\delta \lambda} = \frac{\partial \tilde{L}}{\partial \lambda} - \frac{\mathrm{d}}{\mathrm{d}t} \left(\frac{\partial \tilde{L}}{\partial \dot{\lambda}} \right) = f(q_i, \dot{q}_i, t)$$
(34)

and therefore it imposes the constraint f=0 independently of what λ actually is. The Euler-Lagrange equations for the original coordinates are

$$\frac{\partial L}{\partial q_i} - \frac{\mathrm{d}}{\mathrm{d}t} \left(\frac{\partial L}{\partial \dot{q}_i} \right) = -\lambda(t) \frac{\partial f}{\partial q_i}$$
(35)

1.4.3 Example: Helter Skelter

A child of m slides down a helter skelter, the Lagrangian is given by

$$L = \frac{1}{2}m\left(\dot{r}^2 + r^2\dot{\theta}^2 + \dot{z}^2\right) - mgz \tag{36}$$

and

$$z = h - \alpha \theta, \qquad r = \beta \theta$$
 (37)

where α and β are positive constants and h is the height of the helter skelter. We take the angle θ to go from 0 to infinity as the trajectory winds around multiple times. The constraints are

$$C_1 = z - h + \alpha \theta, \qquad C_2 = r - \beta \theta$$
 (38)

1) Method 1: Solve constraints

Put constraints into the Lagrangian

$$L(\theta, t) = \frac{1}{2}m\left(\beta^2\dot{\theta}^2 + \beta^2\theta^2\dot{\theta}^2 + \alpha^2\dot{\theta}^2\right) - mg(h - \alpha\theta)$$
 (39)

The Euler-Lagrange equation for θ is

$$m\frac{\mathrm{d}}{\mathrm{d}t}\left[\left(\alpha^2 + \beta^2\right) + \beta^2\theta^2\right]\dot{\theta} = m\beta^2\theta\dot{\theta}^2 + mg\alpha \tag{40}$$



Figure 2: A helter skelter.

2) Method 2: Work in extended configuration space

$$\tilde{L} = \frac{1}{2}m\left(\dot{r}^2 + r^2\dot{\theta}^2 + \dot{z}^2\right) - mgz + (\lambda_1C_1 + \lambda_2C_2) \tag{41}$$

The Euler-Lagrange equations are

• For r: $mr\dot{\theta}^2 - m\ddot{r} = -\lambda_2$

• For θ : $m\frac{\mathrm{d}}{\mathrm{d}t}\left(r^2\dot{\theta}\right) = \alpha\lambda_1 - \beta\lambda_2$

• For z: $m\ddot{z} + mg = \lambda_1$

• For λ_1 : $z = h - \alpha \theta$

• For λ_2 : $r = \beta \theta$

So we have

$$m\frac{\mathrm{d}}{\mathrm{d}t}\left[\left(\alpha^2 + \beta^2\right) + \beta^2\theta^2\right]\dot{\theta} = m\beta^2\theta\dot{\theta}^2 + mg\alpha \tag{42}$$

1.5 Normal Modes

'Natural' System: We consider the situation where particles can move around their equilibrium positions. We assume that the system is described by N generalised coordinates q_i . We also assume that it is natural, which means that the kinetic energy is a $quadratic\ homogeneous\ function$ of the generalised velocities. We can then write it as

$$T = \frac{1}{2} \sum_{ij} a_{ij}(q_1, \cdots, q_N) \dot{q}_i \dot{q}_j = \frac{1}{2} \dot{\boldsymbol{q}}^T A \dot{\boldsymbol{q}}$$

$$\tag{43}$$

Since the Lagrangian is given by

$$L = T - V(q_1, \cdots, q_N, t) \tag{44}$$

where the potential does not depend on the velocities \dot{q}_i . Then we have

$$\mathcal{Z}^{ij} = \frac{\partial^{2} L}{\partial \dot{q}_{i} \partial \dot{q}_{j}} = \frac{\partial^{2} T}{\partial \dot{q}_{i} \partial \dot{q}_{j}} = \frac{\partial^{2}}{\partial \dot{q}_{i} \partial \dot{q}_{j}} \left(\frac{1}{2} a_{\alpha\beta} \dot{q}_{\alpha} \dot{q}_{\beta} \right)
= \frac{\partial}{\partial \dot{q}_{j}} \left(\frac{1}{2} a_{\alpha\beta} \frac{\partial \dot{q}_{\alpha}}{\partial \dot{q}_{i}} \dot{q}_{\beta} + \frac{1}{2} a_{\alpha\beta} \dot{q}_{\alpha} \frac{\partial \dot{q}_{\beta}}{\partial \dot{q}_{i}} \right)
= \frac{\partial}{\partial \dot{q}_{j}} \left(\frac{1}{2} a_{\alpha\beta} \delta_{i\alpha} \dot{q}_{\beta} + \frac{1}{2} a_{\alpha\beta} \dot{q}_{\alpha} \delta_{i\beta} \right)
= \frac{\partial}{\partial \dot{q}_{j}} \left(\frac{1}{2} a_{i\beta} \dot{q}_{\beta} + \frac{1}{2} a_{\alpha i} \dot{q}_{\alpha} \right)
= \frac{1}{2} a_{ij} \dot{q}_{j} + \frac{1}{2} a_{ji} \dot{q}_{j} = a_{ij}$$
(45)

the kinetic matrix is then given directly by the coefficients a_{ij} .

1.5.1 Equilibrium Points

From the calculation in above, we know that

$$\frac{\partial L}{\partial \dot{q}_i} = a_{ij} \dot{q}_j \tag{46}$$

The Euler-Lagrangian equation

$$\frac{\mathrm{d}}{\mathrm{d}t} \left(\frac{\partial L}{\partial \dot{q}_i} \right) - \frac{\partial L}{\partial q_i} = \frac{\mathrm{d}}{\mathrm{d}t} \left(a_{ij} \dot{q}_j \right) - \frac{\partial T}{\partial q_i} + \frac{\partial V}{\partial q_i}$$

$$= \frac{\mathrm{d}}{\mathrm{d}t} \left(a_{ij} \dot{q}_j \right) - \frac{\partial}{\partial q_i} \left(\frac{1}{2} a_{jk} \dot{q}_j \dot{q}_k \right) + \frac{\partial V}{\partial q_i}$$

$$= a_{ij} \ddot{q}_j - \frac{1}{2} \frac{\partial a_{jk}}{\partial q_i} \dot{q}_j \dot{q}_k + \frac{\partial V}{\partial q_i} = 0$$
(47)

In equilibrium point, \dot{q}_i is a constant, so we require

$$\frac{\partial V}{\partial q_i}\Big|_{q_i=q_{0i}} = 0, \qquad i = 1, 2, \cdots, N$$
 (48)

1.5.2 Small Oscillations

We can write the coordinates as $q_i(t) = q_{0i} + \delta q_i(t)$. The action is given by

$$S = \int dt \left[\frac{1}{2} a_{ij} (\boldsymbol{q}_0) \frac{d}{dt} (q_{0i} + \delta q_i) \frac{d}{dt} (q_{0j} + \delta q_j) - V(\boldsymbol{q}_0 + \delta \boldsymbol{q}) \right]$$

$$= \int dt \left[\frac{1}{2} a_{ij} (\boldsymbol{q}_0) \delta \dot{q}_i \delta \dot{q}_j - \left(V(\boldsymbol{q}_0) + \frac{\partial V}{\partial q_i} \Big|_{\boldsymbol{q}_0} \delta q_i + \frac{1}{2} \frac{\partial^2 V}{\partial q_i \partial q_j} \Big|_{\boldsymbol{q}_0} \delta q_i \delta q_j + \mathcal{O}(\delta q^3) \right) \right]$$

$$= \int dt \left[\frac{1}{2} a_{ij} (\boldsymbol{q}_0) \delta \dot{q}_i \delta \dot{q}_j - \frac{1}{2} \frac{\partial^2 V}{\partial q_i \partial q_j} \Big|_{\boldsymbol{q}_0} \delta q_i \delta q_j \right]$$

$$= \int dt \left[\frac{1}{2} a_{ij} (\boldsymbol{q}_0) \delta \dot{q}_i \delta \dot{q}_j - \frac{1}{2} b_{ij} (\boldsymbol{q}_0) \delta q_i \delta q_j \right] - \int dt V(\boldsymbol{q}_0)$$

$$(49)$$

where

$$b_{ij} = \frac{\partial^2 V}{\partial q_i \partial q_j} \tag{50}$$

The second term is just a constant, so and so for small fluctuations all that matters is the quadratic part, which we refer to as the action for the quadratic fluctuations

$$S_{(2)} = \int dt \left(\frac{1}{2} a_{ij} \delta \dot{q}_i \delta \dot{q}_j - \frac{1}{2} b_{ij} \delta q_i \delta q_j \right)$$

$$= \int dt \left(\frac{1}{2} \delta \dot{\boldsymbol{q}}^T \mathbf{A} \delta \dot{\boldsymbol{q}} - \frac{1}{2} \delta \boldsymbol{q}^T \mathbf{B} \delta \boldsymbol{q} \right)$$
(51)

where the matrix A and B have the components $A_{ij} = a_{ij}$ and $B_{ij} = b_{ij}$.

The Euler-Lagrange equation for the fluctuations is

$$\frac{\mathrm{d}}{\mathrm{d}t} \left(\frac{\partial L_{(2)}}{\partial \delta \dot{q}_i} \right) = \frac{\partial L}{\partial \delta q_i} \tag{52}$$

which is

$$a_{ij}\delta\ddot{q}_{i} = -b_{ij}\delta q_{i}$$
 or $A\delta\ddot{q} = -B\delta q$ (53)

If $det(\mathbf{A}) \neq 0$, then

$$\delta \ddot{q} = -K\delta q \tag{54}$$

with

$$\mathbf{K} = \mathbf{A}^{-1}\mathbf{B} \tag{55}$$

1.5.3 Gram-Schmidt Diagonalization

Since the matrix **A** is symmetric (we may assume this from the outset), we can always diagonalize it using a $N \times N$ orthogonal matrix **O** ($\mathbf{O}^T \mathbf{O} = 1$) so that

$$\mathbf{A}_D = \mathbf{O}^T \mathbf{A} \mathbf{O} \tag{56}$$

with A_D diagonal. The kinetic term

$$T = \frac{1}{2} \delta \dot{\boldsymbol{q}}^T \mathbf{A} \delta \dot{\boldsymbol{q}} = \frac{1}{2} \delta \dot{\boldsymbol{q}}^T \mathbf{O} \mathbf{A}_D \mathbf{O}^T \delta \dot{\boldsymbol{q}} = \frac{1}{2} \delta \dot{\tilde{\boldsymbol{q}}}^T \mathbf{A}_D \delta \dot{\tilde{\boldsymbol{q}}}$$

$$= \frac{1}{2} \delta \dot{\tilde{q}}_i(a_D)_{ij} \delta \dot{\tilde{q}}_j$$

$$= \frac{1}{2} (a_D)_{11} \delta \dot{\tilde{q}}_1^2 + \frac{1}{2} (a_D)_{22} \delta \dot{\tilde{q}}_2^2 + \dots + \frac{1}{2} (a_D)_{NN} \delta \dot{\tilde{q}}_N^2$$
(57)

where

$$\delta \tilde{\boldsymbol{q}} = \mathbf{O}^T \delta \boldsymbol{q} \tag{58}$$

To put the kinetic term in diagonal and normalized form, we require

$$\delta q_i = \sqrt{(a_D)_{ii}^{-1}} \delta Q_i \tag{59}$$

This is the statement that there is a diagonal matrix W whose matrix elements are

$$\mathbf{W}_{ij} = \sqrt{(a_D)_{ii}^{-1}} \delta_{ij} \tag{60}$$

such that

$$\delta \tilde{q} = \mathbf{W} \delta \mathbf{Q} \tag{61}$$

for which

$$\mathbf{W}^T \mathbf{A}_D \mathbf{W} = 1 \tag{62}$$

These two operations amount to saying that there is a matrix S = OW such that

$$\mathbf{S}^T \mathbf{A}_D \mathbf{S} = 1 \tag{63}$$

the action for the fluctuations becomes

$$S_{(2)} = \int dt \left(\frac{1}{2} \delta \dot{\boldsymbol{Q}}^T \mathbf{S}^T \mathbf{A} \mathbf{S} \delta \dot{\boldsymbol{Q}} - \frac{1}{2} \delta \dot{\boldsymbol{Q}}^T \mathbf{S}^T \mathbf{B} \mathbf{S} \delta \dot{\boldsymbol{Q}} \right)$$

$$= \int dt \left(\frac{1}{2} \delta \dot{\boldsymbol{Q}}^T \delta \dot{\boldsymbol{Q}} - \frac{1}{2} \delta \dot{\boldsymbol{Q}}^T \mathbf{k} \delta \dot{\boldsymbol{Q}} \right)$$
(64)

where

$$\mathbf{k} = \mathbf{S}^T \mathbf{B} \mathbf{S} \tag{65}$$

This is the canonically normalized form for the fluctuations.

1.5.4 Cholesky Decomposition

The Cholesky decomposition of a *real Hermitian positive*-definite 1 matrix A, is a decomposition of the form

$$\mathbf{A} = \mathbf{L}\mathbf{L}^T \tag{66}$$

where L is a left triangular matrix with real and positive diagonal entries

$$\mathbf{L} = \begin{pmatrix} \# & 0 & 0 & 0 \\ \# & \# & 0 & 0 \\ \# & \# & \# & 0 \\ \# & \# & \# & \# \end{pmatrix}$$
(67)

So the kinetic term can be written as

$$T = \frac{1}{2}\delta\dot{\boldsymbol{q}}\mathbf{A}\delta\dot{\boldsymbol{q}} = \frac{1}{2}\delta\dot{\boldsymbol{q}}\mathbf{L}\mathbf{L}^T\delta\dot{\boldsymbol{q}} = \frac{1}{2}\delta\dot{\boldsymbol{Q}}^T\delta\dot{\boldsymbol{Q}}$$
(68)

where

$$\delta \mathbf{Q} = \mathbf{L}^T \delta \mathbf{q} \tag{69}$$

¹'Positive' means the eigenvalues of **A** are all positive.

And the action

$$S = \int dt \left(\frac{1}{2} \delta \dot{\boldsymbol{q}} \mathbf{A} \delta \dot{\boldsymbol{q}} - \frac{1}{2} \delta \boldsymbol{q}^T \mathbf{B} \delta \boldsymbol{q} \right)$$

$$= \int dt \left(\frac{1}{2} \delta \dot{\boldsymbol{Q}} \delta \dot{\boldsymbol{Q}} - \frac{1}{2} \delta \boldsymbol{Q}^T \mathbf{k} \delta \boldsymbol{Q} \right)$$

$$= \int dt \left(\frac{1}{2} \delta \dot{Q}_i \delta \dot{Q}_i - \frac{1}{2} k_{ij} \delta Q_i \delta Q_j \right)$$
(70)

where

$$\mathbf{k} = \mathbf{L}^{-1} \mathbf{B} (\mathbf{L}^T)^{-1} \tag{71}$$

The Euler-Lagrange equation for Q_i is given by

$$\delta \ddot{Q}_i = -k_{ij}\delta Q$$
 or $\delta \ddot{\boldsymbol{Q}} = -\mathbf{k}\delta \boldsymbol{Q}$ (72)

We look for solutions of the form

$$\delta \mathbf{Q} = e^{i\omega t} \delta \mathbf{Q}_{\omega} \tag{73}$$

then we have the eigenfunction equation with the matrix k and eigenvalue ω^2 .

$$\mathbf{k}\delta\mathbf{Q}_{\omega} = \omega^2 \delta\mathbf{Q}_{\omega} \tag{74}$$

which means that each normal coordinate $Q_{\omega_{\alpha}}$ oscillates independently of all others with its own normal frequency ω_{α}^2 .

1.5.5 Example: Double Pendulum

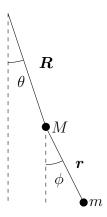


Figure 3: Double pendulum

Consider a double pendulum, with a second pendulum hanging from the first as depicted in Fig.(3). The kinetic energy of the system is

$$T = \frac{1}{2}M\dot{R}^{2} + \frac{1}{2}m(\dot{R} + \dot{r})^{2}$$

$$= \frac{1}{2}M\dot{R}^{2} + \frac{1}{2}m\left(\dot{R}^{2} + \dot{r}^{2} + 2\dot{R}\dot{r}\right)$$

$$= \frac{1}{2}MR^{2}\dot{\theta}^{2} + \frac{1}{2}mR^{2}\dot{\theta}^{2} + \frac{1}{2}mr^{2}\dot{\phi}^{2} + mRr\dot{\theta}\dot{\phi}\cos(\phi - \theta)$$
(75)

The potential energy is

$$V = Mg(-R\cos\theta) + mg(-R\cos\theta - r\cos\phi)$$

= -(M + m)gR\cos\theta - mgr\cos\phi

The equilibrium solution is $\theta_0 = \phi_0 = 0$, so θ and ϕ can be expressed by

$$\theta(t) = \delta\theta(t), \qquad \phi(t) = \delta\phi(t)$$
 (77)

then we have

$$L_{(2)} = \frac{1}{2}(M+m)R^{2}\delta\dot{\theta}^{2} + \frac{1}{2}mr^{2}\delta\dot{\phi}^{2} + mRr\delta\dot{\theta}\delta\dot{\phi} - \frac{1}{2}(M+m)gR\delta^{2}\theta - \frac{1}{2}mgr\delta\phi^{2}$$

$$= \frac{1}{2}MR^{2}\delta\dot{\theta}^{2} + \frac{1}{2}m\left(R\delta\dot{\theta} + r\delta\dot{\phi}\right)^{2} - \frac{1}{2}(M+m)gR\delta^{2}\theta - \frac{1}{2}mgr\delta\phi^{2}$$
(78)

We define that

$$\delta Q_1 = \sqrt{M}R\delta\theta, \qquad \delta Q_2 = \sqrt{m}(R\delta\theta + r\delta\phi)$$
 (79)

so the Lagrangian becomes

$$L_{(2)} = \frac{1}{2}\delta\dot{Q}_1^2 + \frac{1}{2}\delta\dot{Q}_2^2 - \frac{1}{2}\frac{(M+m)g}{MR}\delta Q_1^2 - \frac{1}{2}\frac{mg}{Mr}(\delta Q_1 - \delta Q_2)^2$$
 (80)

and the potential energy is given by

$$V = -\frac{1}{2}\delta \mathbf{Q}^T \mathbf{k} \delta \mathbf{Q} = -\frac{1}{2} \begin{pmatrix} \delta Q_1 & \delta Q_2 \end{pmatrix} \mathbf{k} \begin{pmatrix} \delta Q_1 \\ \delta Q_2 \end{pmatrix}$$
(81)

where

$$\mathbf{k} = \begin{pmatrix} \frac{(M+m)g}{MR} + \frac{mg}{Mr} & -\frac{mg}{Mr} \\ -\frac{mg}{Mr} & \frac{mg}{Mr} \end{pmatrix} = \frac{mg}{Mr} \begin{pmatrix} \frac{(M+m)r}{mR} + 1 & -1 \\ -1 & 1 \end{pmatrix}$$
(82)

To simplify k, we choose R = r and M = m, then we have

$$\mathbf{k} = \frac{g}{R} \begin{pmatrix} 3 & -1 \\ -1 & 1 \end{pmatrix} \tag{83}$$

The eigenvalues ω_1^2 and ω_2^2 satisfies

$$\det(\mathbf{k} - \omega^2 \mathbb{1}) = \frac{g}{R} \left[(3 - \omega^2)(1 - \omega^2) - 1 \right] = 0$$
 (84)

solve this and get

$$\omega_{1,2}^2 = (2 \pm \sqrt{2}) \frac{g}{R} > 0 \tag{85}$$

the associated eigenvectors are $Q_{\pm}=(1\pm\sqrt{2},-1)$. Go back to the original coordinates θ , ϕ , we know that

$$Q_1 \sim \theta, \qquad Q_2 \sim \theta + \phi$$
 (86)

so the normal modes are

$$(\theta, \phi)_{\pm} = (1 \pm \sqrt{2}, -(2 \pm \sqrt{2})) \sim (1, \pm \sqrt{2})$$
 (87)

1.6 Symmetries and Noether's Theorem

Let's parameterize the action by arbitrary coordinates q_i

$$S = \int dt L(q, \dot{q}, t) \tag{88}$$

now we have

$$\delta S = \int dt \left(\frac{\partial L}{\partial q_i} \delta q_i + \frac{\partial L}{\partial \dot{q}_i} \delta \dot{q}_i \right)$$

$$= \left(\frac{\partial L}{\partial \dot{q}_i} \delta q_i \right) \Big|_{t_1}^{t_2} + \int dt \left[\frac{\partial L}{\partial q_i} - \frac{d}{dt} \left(\frac{\partial L}{\partial \dot{q}_i} \right) \right] \delta q_i$$
boundary term
$$= \left(\frac{\partial L}{\partial \dot{q}_i} \delta q_i \right) \Big|_{t_1}^{t_2}$$

$$= \left(\frac{\partial L}{\partial \dot{q}_i} \delta q_i \right) \Big|_{t_2}^{t_2}$$
(89)

Euler-Lagrange equations follow the principle of least action. So $\delta S=0$ up to boundary terms.

1.6.1 Discrete Global Symmetry

Let's consider a simple example, which is a harmonic oscillator

$$S[q(t)] = \int dt \left(\frac{1}{2}m\dot{q}^2 - \frac{1}{2}m\omega^2 q^2\right)$$
(90)

consider the transition q' = -q, and

$$S[q'(t)] = \int dt \left(\frac{1}{2}m\dot{q}'^2 - \frac{1}{2}m\omega^2 q'^2\right) = \int dt \left(\frac{1}{2}m\dot{q}^2 - \frac{1}{2}m\omega^2 q^2\right) = S[q(t)]$$
 (91)

1.6.2 Continuous Global Symmetry

Let's consider a general infinitesimal transform

$$\delta q_i = F_i(q, \dot{q})\delta\lambda$$
 (92)

then the Lagrangian transforms as

$$\delta L(q, \dot{q}, t) = L(q + F\delta\lambda, \dot{q} + \dot{F}\delta\lambda) - L(q, \dot{q})$$

$$= \frac{\partial L}{\partial q_i} F_i \delta\lambda + \frac{\partial L}{\partial \dot{q}_i} \dot{F}_i \delta\lambda$$

$$= \left(\frac{\partial L}{\partial q_i} F_i + \frac{\partial L}{\partial \dot{q}_i} \dot{F}_i\right) \delta\lambda$$
(93)

If this is a symmetry, then this must be a total derivative then

$$\frac{\partial L}{\partial q_i} F_i + \frac{\partial L}{\partial \dot{q}_i} \dot{F}_i = \frac{\mathrm{d}A}{\mathrm{d}t}$$
(94)

and the integration

$$\delta S = \int dt \delta L = \int_{t_1}^{t_2} dt \frac{dA}{dt} \delta \lambda = [A(t_2) - A(t_1)] \delta \lambda = 0$$
 (95)

1.6.3 Noether's theorem

Noether performed a clever trick. She make λ a function of time

$$\delta q_i = F_i(q_i, \dot{q}_i) \delta \lambda(t) \tag{96}$$

now we have

$$\delta S = \int dt \left[\frac{\partial L}{\partial q_i} \delta q_i + \frac{\partial L}{\partial \dot{q}_i} \frac{d}{dt} (\delta q_i) \right]$$

$$= \int dt \left[\frac{\partial L}{\partial q_i} F_i \delta \lambda + \frac{\partial L}{\partial \dot{q}_i} \dot{F}_i \delta \lambda + \frac{\partial L}{\partial \dot{q}_i} F_i \frac{d}{dt} (\delta \lambda) \right]$$

$$= \int dt \left[\frac{dA}{dt} \delta \lambda - \frac{d}{dt} \left(\frac{\partial L}{\partial \dot{q}_i} F_i \right) \delta \lambda \right] + \left(\frac{\partial L}{\partial \dot{q}} F_i \delta \lambda \right) \Big|_{t_1}^{t_2}$$

$$= \int dt \left[-\frac{d}{dt} \left(\frac{\partial L}{\partial \dot{q}_i} F_i - A \right) \delta \lambda \right] + \left(\frac{\partial L}{\partial \dot{q}} F_i \delta \lambda \right) \Big|_{t_1}^{t_2}$$

$$= \int dt \left(-\frac{dG}{dt} \delta \lambda \right) + \left(\frac{\partial L}{\partial \dot{q}} F_i \delta \lambda \right) \Big|_{t_1}^{t_2} = 0$$
(97)

where

$$G = \frac{\partial L}{\partial \dot{q}} F - A \tag{98}$$

so we have

$$\frac{\mathrm{d}G}{\mathrm{d}t} = 0 \quad \text{(up to the boundary term)} \tag{99}$$

Theorem 1

Noether's theorem: For every continuous symmetry, there exist a conserved quantity G conserved in time.

Example

Consider the action

$$S = \int \mathrm{d}t \frac{1}{2} m \dot{q}^2 \tag{100}$$

and the transform $q'=q+\lambda$, i.e., $\delta q=\delta \lambda$ with F=1 and A=0.

$$G = \frac{\partial L}{\partial \dot{q}} F - A = \frac{\partial L}{\partial \dot{q}} = p \tag{101}$$

which means the corresponding generalized momentum component p to be a constant of motion.

1.6.4 Hamiltonian as the Noether Charge for Time Translations

Consider a system which is time translation invariant meaning that the action is invariant (un to boundary terms) under the symmetry $t \to t + \delta t$. Such a transformation induces a change of coordinates

$$\delta q_i(t) = q_i(t+\lambda) - q_i(t) = \dot{q}_i \delta \lambda \tag{102}$$

Similarly the Lagrangian transforms as

$$\delta L = L(q(t+\lambda), \dot{q}(t+\lambda), t+\lambda) - L(q, \dot{q}, t) = \frac{\mathrm{d}L}{\mathrm{d}t} \delta \lambda$$
 (103)

from which we infer

$$A = L \tag{104}$$

Thus the conserved charge implied by Noether's theorem associated with the symmetry of time translation invariance is

$$G = \frac{\partial L}{\partial \dot{q}_i} \dot{q}_i - L = p_i \dot{q}_i - L = H$$
 (105)

which we recognize to be the Hamiltonian.

2 Hamiltonian

2.1 Hamiltonian Formulation

In Lagrangian, we use coordinates and velocities (q, \dot{q}) to parameterise the system, while in Hamiltonian, we use coordinates and momenta (q, p) to parameterise the system.

$$\begin{array}{cccc} \mbox{Lagrangian} & & \mbox{Hamiltonian} \\ L = L(q,\dot{q},t) & \longleftrightarrow & H = H(q,p,t) \\ & & p = \frac{\partial L}{\partial \dot{q}} & \end{array}$$

2.1.1 From Lagrangian to Hamiltonian

The total derivative of Lagrangian can be expressed as

$$dL(q, \dot{q}, t) = \frac{\partial L}{\partial q_i} dq_i + \frac{\partial L}{\partial \dot{q}_i} d\dot{q}_i + \frac{\partial L}{\partial t} dt$$

$$= \frac{\partial L}{\partial q_i} dq_i + p_i d\dot{q}_i + \frac{\partial L}{\partial t} dt$$
(106)

We define the *Hamiltonian function* H by Legendre transformation

$$H = \sum_{i} p_i \dot{q}_i - L \tag{107}$$

2.1.2 Hamilton's Equations

The total derivative of Hamiltonian can be expressed as

$$dH = -\frac{\partial L}{\partial q_i} dq_i + \dot{q}_i dp_i - \frac{\partial L}{\partial t} dt$$

$$= \frac{\partial H}{\partial q_i} dq_i + \frac{\partial H}{\partial p_i} dp_i + \frac{\partial H}{\partial t} dt$$
(108)

Comparing the two expressions, we have

$$\frac{\partial H}{\partial q_i} = -\frac{\partial L}{\partial q_i} = -\dot{p}_i, \qquad \frac{\partial H}{\partial p_i} = \dot{q}_i, \qquad \frac{\partial H}{\partial t} = -\frac{\partial L}{\partial t}$$
(109)

These results are known as Hamilton's equations.

2.1.3 Conservation of the Hamiltonian

We can also calculate the time derivative of the Hamiltonian

$$\frac{\mathrm{d}H}{\mathrm{d}t} = \frac{\partial H}{\partial t} + \sum_{i} \left(\frac{\partial H}{\partial q_{i}} \dot{q}_{i} + \frac{\partial H}{\partial p_{i}} \dot{p}_{i} \right)
= \frac{\partial H}{\partial t} + \sum_{i} \left(\frac{\partial H}{\partial q_{i}} \frac{\partial H}{\partial p_{i}} - \frac{\partial H}{\partial p_{i}} \frac{\partial H}{\partial q_{i}} \right) = \frac{\partial H}{\partial t}$$
(110)

which we summarize below

$$\boxed{\frac{\partial H}{\partial t} = \frac{\mathrm{d}H}{\mathrm{d}t}} \tag{111}$$

We can conclude that if the Hamiltonian has mo explicit time dependence (i.e. $\partial H/\partial t=0$), then the Hamiltonian is conserved.

2.2 Hamilton's Principle of Least Action

The phase space action is

$$S = \int dt \left(\sum_{i} p_{i} \frac{dq_{i}}{dt} - H(q, p, t) \right)$$
 (112)

Let us now perform the variation of this action in which we regard $q_i(t)$ and $p_i(t)$ as independent functions

$$\delta S = \int_{t_1}^{t_2} dt \sum_{i} \left(\delta p_i \frac{dq_i}{dt} + p_i \frac{d\delta q_i}{dt} - \frac{\partial H}{\partial q_i} \delta q_i - \frac{\partial H}{\partial p_i} \delta p_i \right)$$

$$= \sum_{i} p_i \delta q_i \Big|_{t_1}^{t_2} + \int_{t_1}^{t_2} dt \sum_{i} \left[\delta p_i \left(\frac{dq_i}{dt} - \frac{\partial H}{\partial p_i} \right) + \left(-\frac{dp_i}{dt} - \frac{\partial H}{\partial q_i} \right) \delta q_i \right]$$
(113)

Thus demanding that $\delta S = 0$ for variations which vanish as the boundary $\delta q_i(t_1) = \delta q_i(t_2) = 0$, which implies Hamilton's equations

$$\dot{q}_i = \frac{\partial H}{\partial p_i}, \qquad \dot{p}_i = -\frac{\partial H}{\partial q_i}$$
 (114)

2.3 Poisson Brackets

The total time derivative of A(q, p, t) is

$$\frac{\mathrm{d}A}{\mathrm{d}t} = \frac{\partial A}{\partial t} + \sum_{i} \left(\frac{\partial A}{\partial q_{i}} \dot{q}_{i} + \frac{\partial A}{\partial p_{i}} \dot{p}_{i} \right)
= \frac{\partial A}{\partial t} + \sum_{i} \left(\frac{\partial A}{\partial q_{i}} \frac{\partial H}{\partial p_{i}} - \frac{\partial A}{\partial p_{i}} \frac{\partial H}{\partial q_{i}} \right)
= \frac{\partial A}{\partial t} + \{A, H\}$$
(115)

where the quantity

$$\{A, H\} = \sum_{i} \left(\frac{\partial A}{\partial q_i} \frac{\partial H}{\partial p_i} - \frac{\partial A}{\partial p_i} \frac{\partial H}{\partial q_i} \right)$$
 (116)

is known as Poisson bracket. The Poisson brackets for coordinates and momenta

$$\{q_i, q_j\} = \sum_{k} \left(\frac{\partial q_i}{\partial q_k} \frac{\partial q_j}{\partial p_k} - \frac{\partial q_i}{\partial p_k} \frac{\partial q_j}{\partial q_k} \right) = 0$$
 (117)

$$\{p_i, p_j\} = \sum_{k} \left(\frac{\partial p_i}{\partial q_k} \frac{\partial p_j}{\partial p_k} - \frac{\partial p_i}{\partial p_k} \frac{\partial p_j}{\partial q_k} \right) = 0$$
 (118)

$$\{q_i, p_j\} = \sum_{k} \left(\frac{\partial q_i}{\partial q_k} \frac{\partial p_j}{\partial p_k} - \frac{\partial q_i}{\partial p_k} \frac{\partial p_j}{\partial q_k} \right) = \sum_{k} \delta_{ik} \delta_{jk} = \delta_{ij}$$
(119)

We can also write Hamilton's equations in terms of Poisson brackets as

$$\dot{p}_i = \{p_i, H\}, \qquad \dot{q}_i = \{q_i, H\}$$
 (120)

2.3.1 Poisson Brackets and Quantum Mechanics

Compare the Poisson brackets in classical mechanics and the canonical commutation relations in quantum mechanics

Classic	Quantum
$q_i, q_j = 0$	$[\hat{q}_i, \hat{q}_j] = 0$
$\{p_i, p_j\} = 0$	$[\hat{p}_i, \hat{p}_j] = 0$
$\{q_i, p_j\} = \delta_{ij}$	$[\hat{q}_i, \hat{p}_j] = i\hbar \delta_{ij}$

Table 1: The Poisson brackets in classical mechanics and the canonical commutation relations in quantum mechanics

2.3.2 Properties of Poisson Brackets

1. Anti-symmetric

$${A,B} = -{B,A}$$
 (121)

2.

$${A, BC} = {A, B}C + B{A, C}$$
 (122)

3. Jacobi identity

$${A, {B, C}} + {B, {C, A}}, {C, {A, B}} = 0$$
 (123)

2.4 Canonical Transformations

The canonical transformation means a transformation which preserve Poisson brackets. Considering the transformation

$$q \to Q(q, p), \qquad p \to P(q, p)$$
 (124)

they take the same Poisson brackets, i.e.

$$\{f,g\}_{Q,P} = \{f,g\}_{q,p} \tag{125}$$

2.4.1 Infinitesimal Canonical Transformations

The general form of an infinitesimal canonical transformation is

$$\delta q_i = \{q_i, G(q, p, t)\}\delta\lambda, \qquad \delta p_i = \{p_i, G(q, p, t)\}\delta\lambda \tag{126}$$