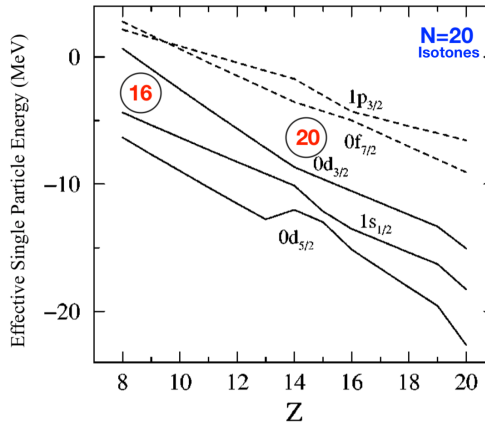


# Shell evolution in nuclei

Matthias Heinz



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[Utsuno et al., 1999]



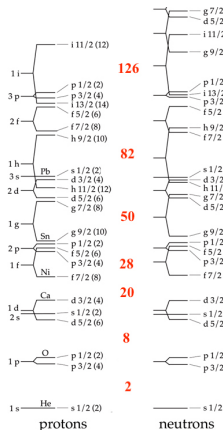
- Overview of the Shell Model
- Motivation
- Disappearance of  $N = 20$  Shell Closure
- Appearance of the  $N = 16$  Shell Closure
- Modern Frontiers for Shell Evolution
- Summary

# Nuclear Shell Model



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- ▶ Mean field and NN correlations
- ▶ Valence space for “active nucleons”
- ▶ Shell gaps and correlations (configuration mixing)
- ▶ Magic nuclei: closed shells for protons and/or neutrons



[Amsler, 2015]

# Shell Evolution

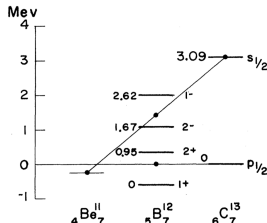


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- ▶ Monopole drift
- ▶ Consider interaction between  $p_{3/2}$  **protons** and the single **neutron**

$$\langle j^2(J=0)j' | V_{1n} + V_{2n} | j^2(J=0)j' \rangle = 2 \sum_{J=|j-j'|}^{j+j'} (2J+1) \langle jj'J | V | jj'J \rangle / \sum_{J=|j-j'|}^{j+j'} (2J+1)$$

- ▶ Spin and isospin dependence
- ▶ Monopole vs. multi-pole correlations

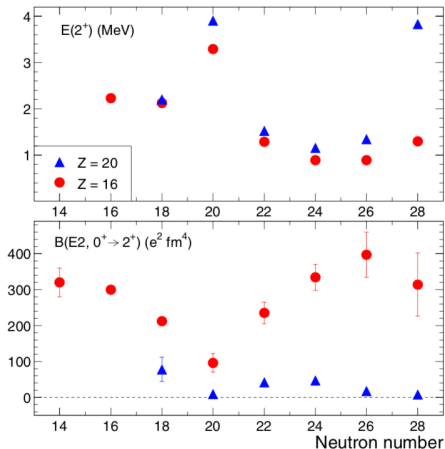


[Talmi and Unna, 1960]

# Properties of “Magic Nuclei”

## Properties of “magic nuclei”:

- ▶ Large binding energy
- ▶ Long lifetime
- ▶ High excitation energy
- ▶ Low  $B(E2)$
- ▶ Spherical  $\rightarrow$  small  $r_{rms}$



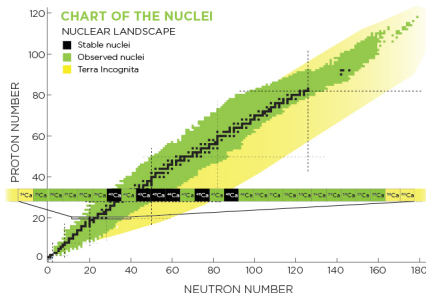
[Alamanos and Gillibert, 2004]

Historical example:

- ▶  $N = 20$  shell disappearance
- ▶  $N = 16$  shell appearance

Experimental challenges:

- ▶ Large range of unstable, bound nuclei
- ▶ Beam intensity quickly drops to 0



[NSCL, 2018]

- ▶ Reduced electric quadrupole transition probability

$$B(E2 : J_i \rightarrow J_f) = \frac{1}{2J_i + 1} |\langle \psi_f | E2 | \psi_i \rangle|^2$$

- ▶ Measured via intermediate Coulomb excitation
  - ▶ Required intensity: ~100 pps
  - ▶ Precision: ~10–15%
- ▶ Alternative: low-energy Coulomb excitation
  - ▶ Required intensity: several thousand pps
- ▶ Alternative: lifetime measurement
  - ▶ Required intensity: several thousand pps

# Coulomb Excitation of $^{32}\text{Mg}$

*Motobayashi et al.*

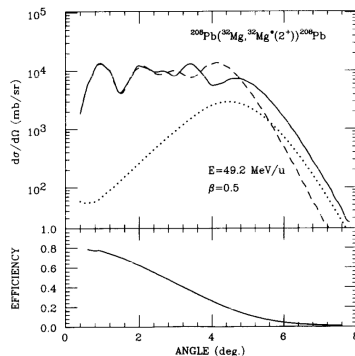


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- ▶  $^{32}\text{Mg}$  beam on  $^{208}\text{Pb}$  target
- ▶ Beam energy 49.2 MeV/u
- ▶ Measure only forward angles
- ▶ Extract  $B(E2)$  from cross section

$$\sigma_{E2} \approx \left(\frac{Ze^2}{\hbar c}\right)^2 \frac{\pi}{e^2 b_{\min}^2} B(E2)$$

- ▶  $\sigma = 91.7 \pm 14.4 \text{ mb}$
- ▶  $B(E2) = 454 \pm 78 \text{ e}^2\text{fm}^4$



[Motobayashi et al., 1995]

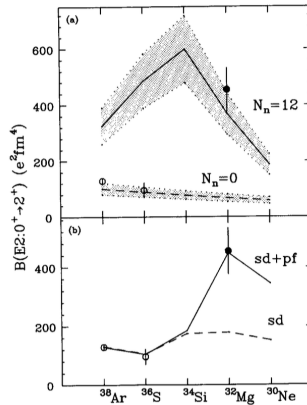


# Measured Value of $B(E2)$

*Motobayashi et al.*



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[Motobayashi et al., 1995]

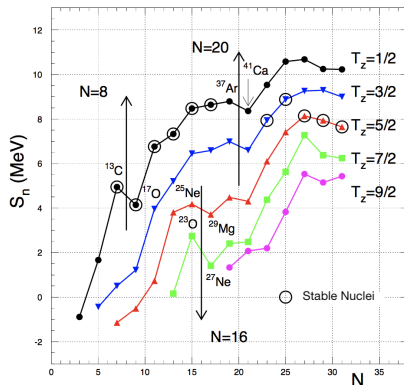
# Observable: $S_n$

*Ozawa et al.*

- ▶ Neutron separation energy
- ▶ Determined by nucleus mass

$$S_n(^{40}\text{Ca}) = M(^{39}\text{Ca}) + M(n) - M(^{40}\text{Ca})$$

- ▶ Mass measured by Penning traps and time-of-flight methods
- ▶ Penning traps:
  - ▶ Required intensity: ~10 pps
  - ▶ Precision: ~1 keV
- ▶ Time-of-flight:
  - ▶ Required intensity: ~0.1 pps
  - ▶ Precision: ~100 keV



Data: [Ozawa et al., 2000]  
Base figure: [Obertelli, 2018]

## Observable: $E_x(2^+)$



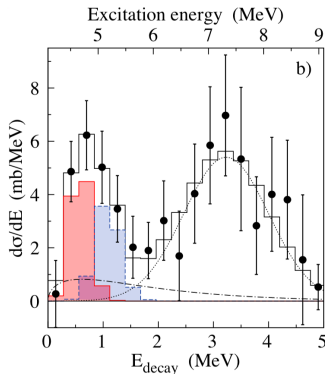
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- ▶ Excitation energy to  $J^P = 2^+$  state
- ▶ First excited state is  $2^+$  in almost all even-even nuclei
- ▶ Near stability, measured via  $\beta$ -decay
- ▶ For exotic nuclei, measured via in-beam gamma spectroscopy
  - ▶ Required intensity:  $\sim 10$  pps
  - ▶ Precision:  $\sim 1$ – $10$  keV

# Measuring $E_x(2^+)$ of $^{24}\text{O}$

*Tshoo et al.*

- ▶ Proton inelastic scattering
- ▶  $^{24}\text{O}(p, p')^{24}\text{O}^* \rightarrow ^{23}\text{O} + n$
- ▶ Excited states are unbound
- ▶ Measure  $^{23}\text{O}$  and  $n$  invariant mass
- ▶ Determine decay energy
  - ▶  $E_{\text{decay}} = 0.56 \pm 0.05 \text{ MeV}$
- ▶ Compute excitation energy
- ▶ Need to know  $S_n$ 
  - ▶  $S_n = 4.09 \pm 0.13 \text{ MeV}$
- ▶  $E_x(2^+) = 4.65 \pm 0.14 \text{ MeV}$



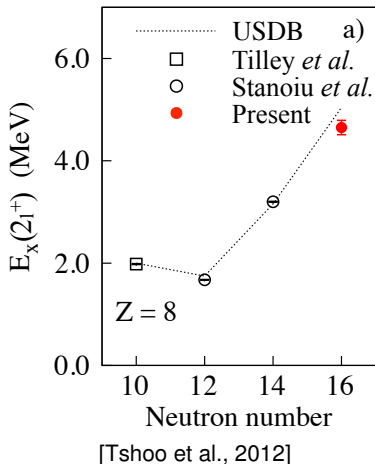
[Tshoo et al., 2012]

# Measuring $E_x(2^+)$ of $^{24}\text{O}$

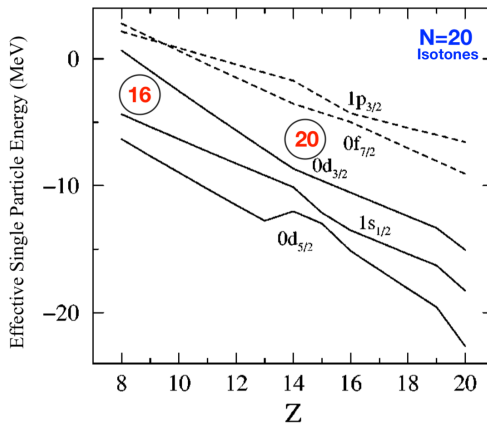
*Tshoo et al.*



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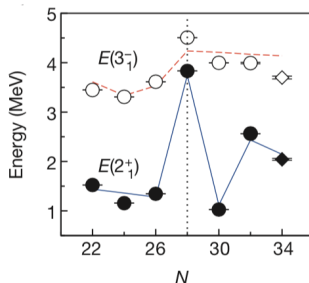
## $N = 20$ and $N = 16$ Recap



[Utsuno et al., 1999]

# Modern Frontiers for Shell Evolution

- ▶ Are  $N = 32$  and  $N = 34$  “magic numbers” for  $Ca$ ?
- ▶ What are the properties of  $^{78}Ni$  and  $^{100}Sn$ ?
- ▶ What are shell closures for super heavy elements?



[Steppenbeck et al., 2013]



- ▶ Shells evolve and shell closures are not universal across the nuclear landscape
- ▶ Observable signs of shell closure:
  - ▶ Low  $B(E2)$
  - ▶ Large binding energy
  - ▶ High  $E(2^+)$
  - ▶ Small  $r_{rms}$
  - ▶ Long lifetime
- ▶ Still many unexplored regions
- ▶ New facilities currently being constructed to explore these effects





Alamanos, N. and Gillibert, A. (2004).  
*Selected Topics in Reaction Studies with Exotic Nuclei*.  
Springer-Verlag, Berlin Heidelberg.



Amsler, C. (2015).  
The nuclear shell model.  
In *Nuclear and Particle Physics*, 2053-2563, pages 6–1 to 6–12. IOP  
Publishing.



Caurier, E., Nowacki, F., and Poves, A. (2014).  
Merging of the islands of inversion at  $N = 20$  and  $N = 28$ .  
*Phys. Rev.*, C90(1):014302.



Dobaczewski, J., Hamamoto, I., Nazarewicz, W., and Sheikh, J. A. (1994).  
Nuclear shell structure at particle drip lines.  
*Phys. Rev. Lett.*, 72:981–984.



Motobayashi, T. et al. (1995).  
Large deformation of the very neutron-rich nucleus Mg-32 from  
intermediate-energy Coulomb excitation.  
*Phys. Lett.*, B346:9–14.



NSCL (2018).  
<https://www.nscl.msu.edu/public/science/index.html>.



Obertelli, A. (2018).  
personal communication.



Otsuka, T., Fujimoto, R., Utsuno, Y., Brown, B. A., Honma, M., and Mizusaki, T. (2001).

Magic numbers in exotic nuclei and spin isospin properties of  $nn$  interaction.

*Phys. Rev. Lett.*, 87:082502.



Otsuka, T., Matsuo, T., and Abe, D. (2006).

Mean Field with Tensor Force and Shell Structure of Exotic Nuclei.

*Phys. Rev. Lett.*, 97:162501.







Otsuka, T. and Suzuki, T. (2013).

Three-Body Forces and Neutron-Rich Exotic Nuclei.

*Few Body Syst.*, 54(7-10):891–896.



-  Ozawa, A., Kobayashi, T., Suzuki, K., Yoshida, K., and Tanihata, I. (2000).  
A new magic number,  $N=16$ , Near to the neutron drip-line.  
*Phys. Rev. Lett.*, 84:5493–5495.
-  Steppenbeck, D. et al. (2013).  
Evidence for a new nuclear ‘magic number’ from the level structure of  $^{54}\text{Ca}$ .  
*Nature*, 502(7470):207–210.
-  Talmi, I. and Unna, I. (1960).  
Order of Levels in the Shell Model and Spin of  $\text{Be-11}$ .  
*Phys. Rev. Lett.*, 4:469–470.
-  Tshoo, K. et al. (2012).  
The  $N = 16$  spherical shell closure in  $^{24}\text{O}$ .  
*Phys. Rev. Lett.*, 109:022501.



Utsuno, Y., Otsuka, T., Mizusaki, T., and Honma, M. (1999).

Varying shell gap and deformation in N approximately 20 unstable nuclei studied by the Monte Carlo shell model.

*Phys. Rev.*, C60:054315.



Wimmer, K. et al. (2010).

Discovery of the shape coexisting  $0^+$  state in  $^{32}\text{Mg}$  by a two neutron transfer reaction.

*Phys. Rev. Lett.*, 105:252501.