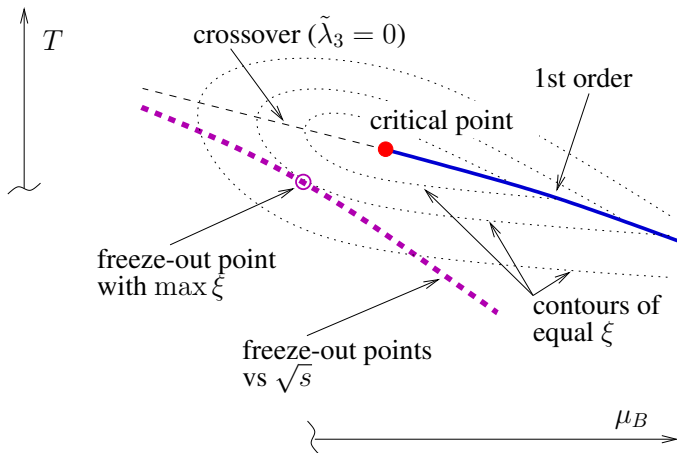


QCD critical point searches at STAR

Matthias Heinz



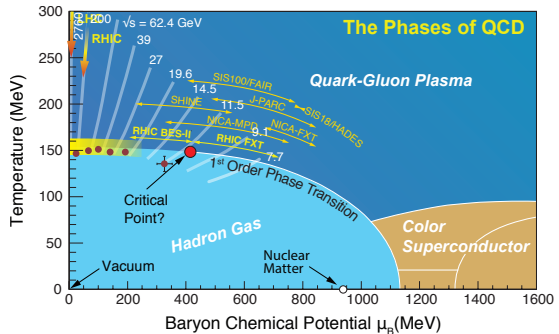
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Stephanov 2009

The RHIC beam energy scan

- ▶ Quark-gluon plasma
- ▶ Hadron gas
- ▶ First-order phase transition and critical point



Caines 2017



- ▶ Critical mode σ which develops infinite correlation length ξ near critical point
- ▶ Treat σ as a classical field
- ▶ Note that for the zero momentum mode $\sigma_0 = \int d^3x \sigma(x) / V$, there are cumulants:

$$\kappa_2 = \langle \sigma_0^2 \rangle = \frac{T}{V} \xi^2$$

$$\kappa_3 = \langle \sigma_0^3 \rangle \sim \frac{T}{V} \xi^6$$

$$\kappa_4 = \langle \sigma_0^4 \rangle_c \sim \frac{T}{V} \xi^8$$

- ▶ Higher-order cumulants are especially sensitive to the increase in the correlation length



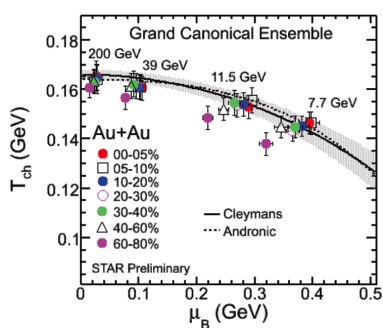
- ▶ Can adopt similar approach to distributions of conserved quantities in experiment
 - ▶ Net baryon number B
 - ▶ Net charge Q
 - ▶ Net strangeness S
- ▶ Connect to theory via susceptibility χ

$$\chi_B^{(n)} = \frac{1}{VT^3} C_{n,B}$$

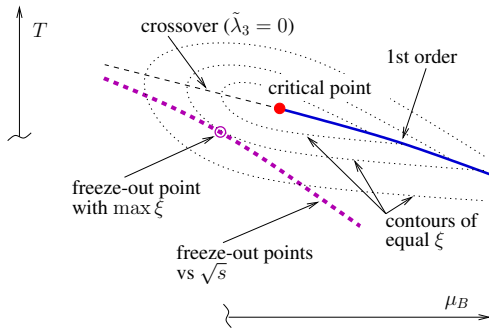
- ▶ Measurements of event-by-event fluctuations of conserved quantities allow us to observe critical point

Overall approach

- ▶ Lower $\sqrt{s_{NN}}$ probes larger μ_B
- ▶ At beam energy scan energies, collision centrality affects μ_B as well

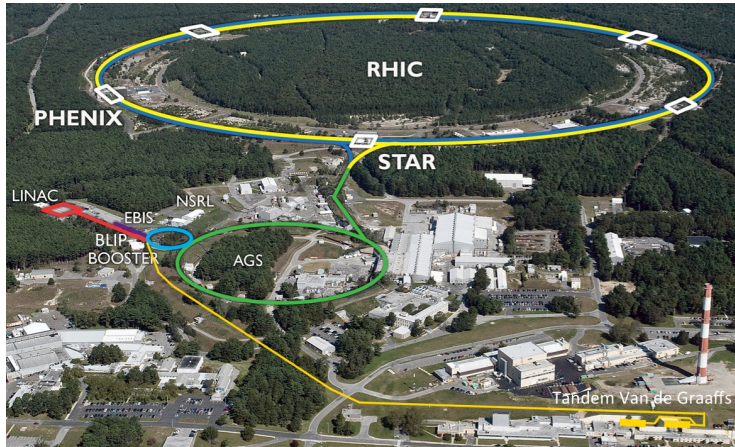


STAR (Schmah et al.) 2013



Stephanov 2009

The STAR detector

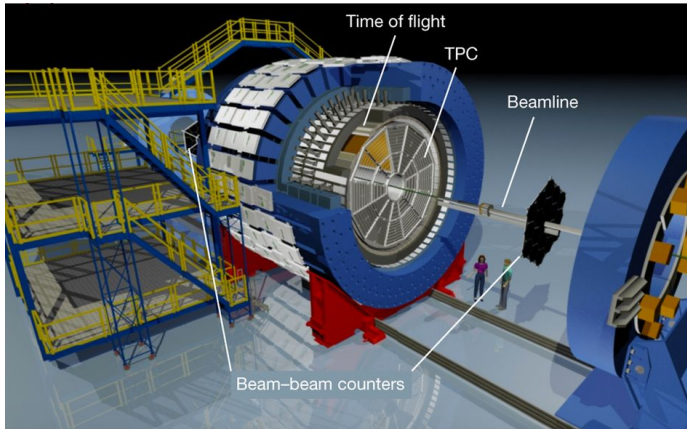


BNL 2019

The STAR detector



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STAR (Adamczyk et al.) 2017



- ▶ Consider net proton number ΔN_p as proxy for net baryon number
- ▶ Can also consider as proxy for net charge number

$$\Delta N_p = N_p - N_{\bar{p}}$$

- ▶ Au+Au collisions
- ▶ Center-of-mass energies: 7.7, 11.5, 19.6, 27, 39, 62.4, and 200 GeV
- ▶ Transverse momentum range: $0.4 < p_T < 2$ GeV/c
- ▶ Pseudorapidity acceptance: $|\eta| < 1.0$

- ▶ Minimum centrality bin size is single value for particle multiplicity
- ▶ Typically work in terms of ranges, like 0-5%
- ▶ Need to correctly assemble cumulant value within a larger centrality bin

$$C_n = \frac{\sum_{r=N_1}^{N_2} n_r C_n^r}{\sum_{r=N_1}^{N_2} n_r} = \sum_{r=N_1}^{N_2} \omega_r C_n^r$$

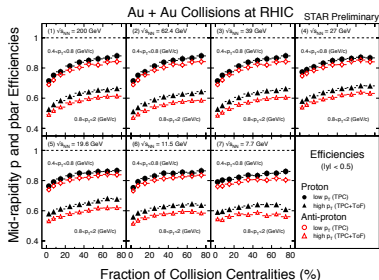
- ▶ Corrected cumulants can be used to compute cumulant ratios
- ▶ Propagation of statistical errors is straightforward
- ▶ Question: Why not apply the same treatment to cumulant ratios?

$$\frac{\sum_{r=N_1}^{N_2} \omega_r C_n^r}{\sum_{r=N_1}^{N_2} \omega_r C_m^r} \neq \sum_{r=N_1}^{N_2} \omega_r \frac{C_n^r}{C_m^r}$$



- ▶ Care needs to be taken when determining centrality
- ▶ In general, more particles used in centrality allows for better resolution
- ▶ However, we need to avoid autocorrelation effects by determining centrality with the same particles used later in the analysis
- ▶ Introduce new centrality definition:
 - ▶ For net-proton analyses, determine centrality with charged kaon and pion multiplicities in $|\eta| < 1$
 - ▶ For net-charge analyses (not discussed here), use particles in $0.5 < |\eta| < 1$ and do analysis with particles in $|\eta| < 0.5$

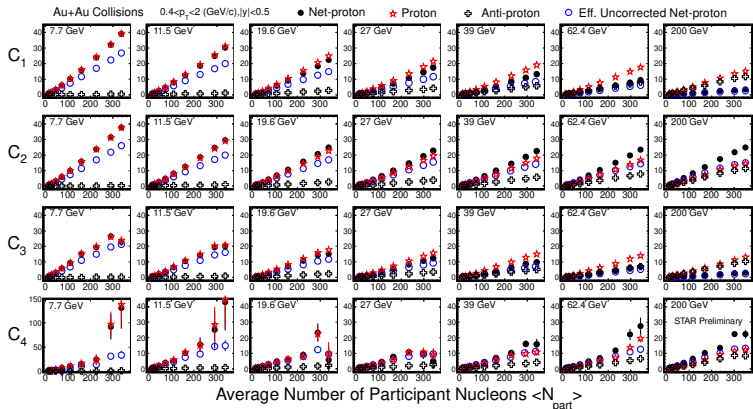
Efficiency correction and statistical error estimation



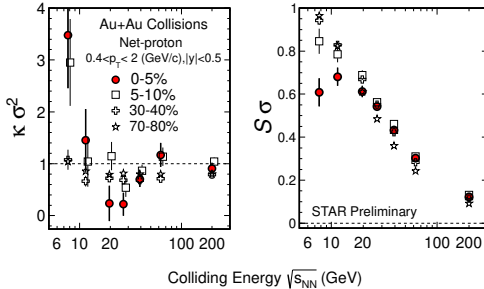
STAR (Luo et al.) 2015

- ▶ Need to account for detector efficiency
- ▶ Particle identification is handled differently for low and high p_T
 - ▶ For $0.4 < p_T < 0.8$ GeV/c, only the TPC is used
 - ▶ For $0.8 < p_T < 2.0$ GeV/c, both TOF and TPC are used
- ▶ Estimate statistical error along with efficiency correction (Delta theorem)
- ▶ Should take place just before doing centrality bin width correction

Observed cumulants up to fourth order



Cumulant ratios



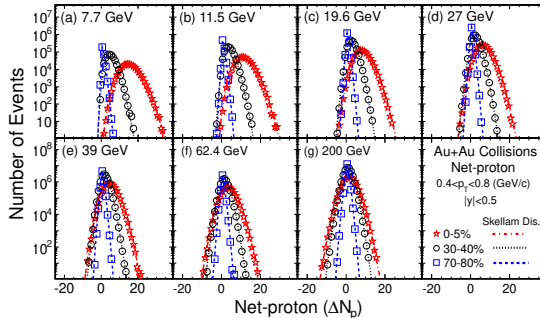
STAR (Luo et al.) 2015

$$\kappa \sigma^2 = C_4 / C_2$$

$$S \sigma = C_3 / C_2$$

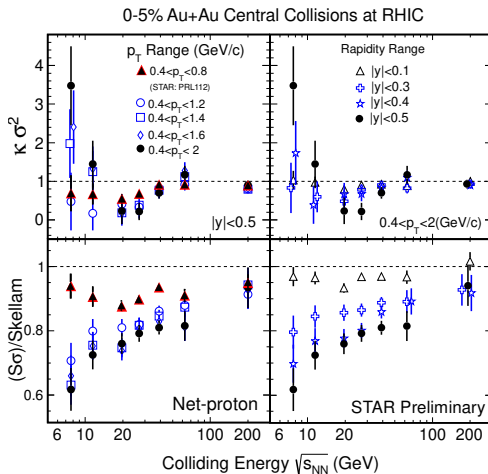
- Form ratios of cumulants to cancel volume and temperature dependence
- S is skewness
- κ is kurtosis

Skellam distributions



STAR (Adamczyk et al.) 2014

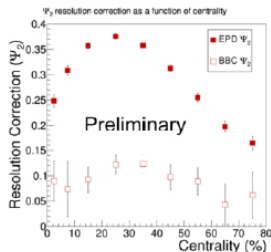
- ▶ Assume protons and anti-protons are distributed as if sampled from independent Poisson distributions
- ▶ Represents thermal statistical fluctuations of net-proton number



STAR (Luo et al.) 2015

Beam energy scan II

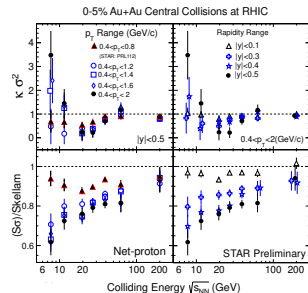
- ▶ Run at $\sqrt{s_N N} = 9.1$ GeV in addition to BES-I energies
- ▶ Energies below 7.7 GeV via fixed-target program
- ▶ Detector upgrades:
 - ▶ iTPC: inner time projection chamber with wider acceptance $|\eta| < 1.5$ and higher resolution
 - ▶ EPD: event plane detector to better determine event plane and offer improved centrality determination



STAR (Yang et al.) 2019

Summary

- ▶ Key ideas:
 - ▶ Look to higher cumulants of distributions of conserved quantities
 - ▶ Use ratios to cancel dependence on quantities other than ξ
 - ▶ Look for non-monotonic behavior in ratios
- ▶ Results seem suggestive, but not conclusive
- ▶ Would like theoretical predictions for susceptibility ratios
- ▶ Need better statistics \rightarrow BES-II



STAR (Luo et al.) 2015



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