

PHY 301

Quantum Mechanics

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Contents

1 Introduction

Quantum: quantities can vary by discrete amounts.

Mechanics: study of motion.

1.1 Franklin's Oil-Drop Experiment

Spilled a spoonful (2 ml) of oil on the surface of a lake and extended to about 2000 m², but not more. This experiment shows the **existent** and **size** of atoms.

$$V = Sh$$
$$h = \frac{V}{S} \sim \frac{2 \times 10^{-6}}{2 \times 10^3} \text{ m} \sim 10^{-9} \text{ m}$$

Atomic and molecular scales are **nanometric**

1.2 From Classical Mechanics to Quantum Mechanics

In classical mechanics, position is a function of time. **Deterministic**.

In quantum mechanics, the position of a particle is a random variable. **Probabilistic**.

2 Wave Function

2.1 Definition in 1-D space

Definition 2.1 (Wave Function). *For a small particle living in a **one-dimensional space**, the wave function Ψ is a complex-valued function of space and time:*

$$\Psi : \mathbf{R} \times \mathbf{R} \rightarrow \mathbb{C}$$

$$(x, t) \mapsto \Psi(x, t) \in \mathbb{C}$$

Remark 2.1. $|\Psi(x, t)|^2$ is the *p.d.f.* of finding the particle in position x at time t .

Remark 2.2. $\int_a^b |\Psi(x, t)|^2 dx$ is the *c.d.f* of finding the particle between position $[a, b]$ at time t .

Remark 2.3. *Integration is over **space**, t is a **parameter**.*

2.2 Mean and variance of the position

These two statistics are expressed as integrals over the entire space. They are **deterministic functions of time**. Given wave function Ψ , then we have:

$$\begin{aligned}\langle x \rangle(t) &= \int_{-\infty}^{+\infty} x |\Psi(x, t)|^2 dx \\ \langle Var(x) \rangle(t) &= \int_{-\infty}^{+\infty} (x - \langle x \rangle(t))^2 |\Psi(x, t)|^2 dx \\ [\Psi] &= \frac{1}{\sqrt{L}}\end{aligned}$$

2.3 Example: probability density of position for classical object

...

3 The Schrödinger equation

One dimension's Schrödinger equation for wave wave function Ψ ($x \in \mathbf{R}$ is a space coordinate, t is time, V is a **real-valued potential**) is

$$\boxed{i\hbar \frac{\partial \Psi(x, t)}{\partial t} = -\frac{\hbar^2}{2m} \frac{\partial^2}{\partial x^2} \Psi(x, t) + V(x) \Psi(x, t)} \quad (1)$$

Complex conjugate form is

$$\boxed{-i\hbar \frac{\partial \Psi^*(x, t)}{\partial t} = -\frac{\hbar^2}{2m} \frac{\partial^2}{\partial x^2} \Psi^*(x, t) + V(x) \Psi^*(x, t)} \quad (2)$$

Planck constant \hbar is

$$\hbar = \frac{h}{2\pi} \simeq 1.05 \times 10^{-34} \text{ kg} \cdot \text{m}^2 \cdot \text{s}^{-1}$$

Assumption 3.1. Ψ and all its derivatives are smooth and go to zero when $|x|$ goes to infinity, faster than any negative power of x .

3.1 Normalization

Due to its definition, the wave function has to be normalized at all time. That is, for all t , we have

$$\boxed{\int_{-\infty}^{\infty} \Psi^*(x, t) \Psi(x, t) dx = 1} \quad (3)$$

Remark 3.1. *Not all solutions of the Schrödinger equation are wave function.*

Theorem 3.1. *A normalized wave function stays normalized.*

Proof. For a normalized wave function at time $t = 0$:

$$\int_{-\infty}^{\infty} \Psi^*(x, 0) \Psi(x, 0) dx = 1$$

consider $t > 0$:

$$\begin{aligned}
& \frac{d}{dt} \left(\int_{-\infty}^{\infty} \Psi^*(x, t) \Psi(x, t) dx \right) \\
&= \int_{-\infty}^{+\infty} \frac{\partial \Psi^*(x, t)}{\partial t} \Psi(x, t) dx + \int_{-\infty}^{+\infty} \Psi^*(x, t) \frac{\partial \Psi(x, t)}{\partial t} dx \\
&= \frac{1}{i\hbar} \int_{-\infty}^{+\infty} \left(+\frac{\hbar^2}{2m} \frac{\partial^2 \Psi^*(x, t)}{\partial x^2} - V(x) \Psi^*(x, t) \right) \Psi(x, t) dx + \\
&\quad \frac{1}{i\hbar} \int_{-\infty}^{+\infty} \Psi^*(x, t) \left(-\frac{\hbar^2}{2m} \frac{\partial^2 \Psi(x, t)}{\partial x^2} + V(x) \Psi(x, t) \right) dx \\
&= \frac{\hbar}{2mi} \left(\left[\Psi \frac{\partial}{\partial x} \Psi^* \right]_{-\infty}^{+\infty} - \int \frac{\partial}{\partial x} \Psi^* \frac{\partial}{\partial x} \Psi dx \right. \\
&\quad \left. - \left[\Psi^* \frac{\partial}{\partial x} \Psi \right]_{-\infty}^{+\infty} + \int \frac{\partial}{\partial x} \Psi \frac{\partial}{\partial x} \Psi^* dx \right) \\
&= 0
\end{aligned}$$

by Assumption ???. Hence, Ψ is normalized at all time. □

3.2 Linear momentum

We can not know actual position of a particle, hence we use the expectation of position $\langle x \rangle(t)$ instead.

$$\boxed{\langle x \rangle(t) = \int_{-\infty}^{+\infty} \Psi^*(x, t) x \Psi(x, t) dx} \tag{4}$$

the velocity of a particle is

$$\boxed{\langle V \rangle = \frac{d\langle x \rangle}{dt}} \tag{5}$$

the average linear momentum is

$$\boxed{\langle p \rangle = \int_{-\infty}^{+\infty} \Psi^*(x, t) \left(-i\hbar \frac{\partial}{\partial x} \right) \Psi(x, t) dx} \tag{6}$$

Proof.

$$\begin{aligned}\langle p \rangle &= m \frac{d\langle x \rangle}{dt} \\&= m \int_{-\infty}^{\infty} \frac{d\Psi^*(x, t)}{dt} x \Psi + \Psi^*(x, t) x \frac{d\Psi(x, t)}{dt} dx \\&= \frac{m}{i\hbar} \int_{-\infty}^{\infty} \left(\frac{\hbar^2}{2m} \frac{\partial^2 \Psi^*(x, t)}{\partial x^2} - V(x) \Psi^*(x, t) \right) x \Psi(x, t) + \left(-\frac{\hbar^2}{2m} \frac{\partial^2 \Psi(x, t)}{\partial x^2} + V(x) \Psi(x, t) \right) x \Psi^*(x, t) dx \\&= \frac{\hbar}{2i} \int_{-\infty}^{\infty} \frac{\partial^2 \Psi^*(x, t)}{\partial x^2} x \Psi(x, t) - \frac{\partial^2 \Psi(x, t)}{\partial x^2} x \Psi^*(x, t) dx \\&= \frac{\hbar}{2i} \int_{-\infty}^{\infty} x \frac{\partial}{\partial x} \left(\frac{\partial \Psi^*(x, t)}{\partial x} \Psi(x, t) - \frac{\partial \Psi(x, t)}{\partial x} \Psi^*(x, t) \right) dx \\&= \frac{\hbar}{2i} \left(\left[x \left(\frac{\partial \Psi^*(x, t)}{\partial x} \Psi(x, t) - \frac{\partial \Psi(x, t)}{\partial x} \Psi^*(x, t) \right) \right]_{-\infty}^{+\infty} - \int_{-\infty}^{\infty} \frac{\partial \Psi^*(x, t)}{\partial x} \Psi(x, t) - \frac{\partial \Psi(x, t)}{\partial x} \Psi^*(x, t) dx \right) \\&= \frac{\hbar}{i} \int_{-\infty}^{\infty} \frac{\partial \Psi(x, t)}{\partial x} \Psi^*(x, t) dx \\&= \int_{-\infty}^{\infty} \Psi^*(x, t) \left(-i\hbar \frac{\partial}{\partial x} \right) \Psi(x, t) dx\end{aligned}$$

□

3.3 From the Schrödinger equation to the Newton law

Ehrenfest's theorem

$$\frac{d\langle p \rangle}{dt} = -\langle V'(x) \rangle \quad (7)$$

obey classical laws?

Proof. ...

□

3.4 Correspondence principle

\hbar to 0?

Energy = kinetic energy + potential energy

Position operator ...

Linear-momentum operator ...

Hamiltonian operator ...

3.5 Separable solutions: time-independent solutions

Consider one kind of solutions of Schrödinger equation in the form:

$$\Psi(x, t) := \psi(x)\phi(t)$$

the Schrödinger equation becomes:

$$i\hbar\psi(x)\phi'(t) = -\frac{\hbar^2}{2m}\psi''(x)\phi(t) + V(x)\psi(x)\phi(t) \quad (8)$$

$$i\hbar\frac{\phi'(t)}{\phi(t)} = -\frac{\hbar^2}{2m}\frac{\psi''(x)}{\psi(x)} + V(x) \quad (9)$$

Eq.?? is satisfied for all values of the independent variables x and t , hence there exists a **constant number** E that Eq.?? = E for both sides. We have

$$i\hbar\frac{\phi'(t)}{\phi(t)} = E \quad (10)$$

$$-\frac{\hbar^2}{2m}\frac{\psi''(x)}{\psi(x)} + V(x) = E \quad (11)$$

Time-dependent factor Eq.??

$$\begin{aligned} \phi'(t) &= \frac{E\phi(t)}{i\hbar} \\ \phi(t) &= \phi(0)e^{\frac{E}{i\hbar}t} \end{aligned} \quad (12)$$

Space-dependent factor Eq.??

$$\begin{aligned} -\frac{\hbar^2}{2m}\psi''(x) + V(x)\psi(x) &= E\psi(x) \\ \hat{H}\psi(x) &= E\psi(x) \end{aligned} \quad (13)$$

Eq.?? is called the ***time-independent Schrödinger equation***.

There are reasons about why introduce time-independent Schrödinger equation: 1. 2.

Remark 3.2. E is a constant number.

Remark 3.3. E is a real number.

Remark 3.4. $\exp(-\frac{iEt}{\hbar})^* = \exp(\frac{iEt}{\hbar})$ when $E \in R$

Proof. ... □

Remark 3.5. $E \geq \min(V(x))$

Proof. ... □

3.6 Example: the infinite square well

Suppose

$$V(x) = \begin{cases} 0, & \text{if } 0 \leq x \leq a \\ \infty, & \text{otherwise} \end{cases}$$

then $\psi(x) = 0$ when $x \notin [0, a]$. When $x \in [0, a]$, we have time-independent Schrödinger $\psi(x)$ satisfied

$$\begin{aligned} -\frac{\hbar^2}{2m} \frac{d^2\psi}{dx^2} &= E\psi \\ \frac{d^2\psi}{dx^2} &= -k^2\psi \end{aligned} \tag{14}$$

where

$$k \equiv \frac{\sqrt{2mE}}{\hbar} \tag{15}$$

Eq.?? is the classical simple harmonic oscillator equation with general solution

$$\psi(x) = A \sin kx + B \cos kx \tag{16}$$

where A and B are constant.

Continuity of $\psi(x)$ requires boundary condition

$$\psi(0) = \psi(a) = 0$$

by $\psi(0) = 0$, we have $B = 0$, $\psi(x) = A \sin kx$

by $\psi(a) = 0$, we have trivial (non-normalizable) solution $A = 0$ or $\sin kx = 0$,

$ka = 0, \pm\pi, \pm2\pi, \pm3\pi, \dots$, $k \leq 0$ is useless. Hence the distinct solutions are

$$k_n = \frac{n\pi}{a}, \quad \text{with } n = 1, 2, 3, \dots$$

According to Eq.??, the corresponding values of E is

$$E_n = \frac{\hbar^2 k_n^2}{2m} = \frac{n^2 \pi^2 \hbar^2}{2ma^2}$$

To find A , we normalize $\Psi(x, t)$

$$\begin{aligned}\int_{-\infty}^{\infty} |\Psi(x, t)|^2 dx &= 1 \\ \int_{-\infty}^{\infty} \psi(x)^2 dx &= 1 \\ \int_{-\infty}^{\infty} |A|^2 \sin^2(kx) dx &= 1 \\ |A|^2 &= \frac{2}{a} \\ A &= \sqrt{\frac{2}{a}}\end{aligned}$$

Finally, we have solutions of time-independent Schrödinger Eq.??

$$\psi_n(x) = \sqrt{\frac{2}{a}} \sin\left(\frac{n\pi}{a}x\right)$$

with corresponding energy

$$E_n = \frac{n^2 \pi^2 \hbar^2}{2ma^2}$$

Consequence I If the wave function at $t = 0$ is a normalized eigenfunction $\psi(x)$ (ie. $n = 1$), we integrate the Schrödinger equation by introducing a phase factor, we have

$$\Psi(x, t) = \exp\left(-\frac{iEt}{\hbar}\right) \psi(x), \forall x, \forall t.$$

$|\Psi(x, t)|^2$ and then $\langle x \rangle$ are independent on time t . The wave function is stationary.

Consequence II If the wave function at $t = 0$ is a sum of eigenfunction $\psi(x)$ with $c_n \in \mathbb{C}$, we integrate the Schrödinger equation by introducing phase factors, we have

$$\Psi(x, t) = A \exp\left(-\frac{iE_1 t}{\hbar}\right) \psi_1(x) + B \exp\left(-\frac{iE_2 t}{\hbar}\right) \psi_2(x), A \in \mathbb{C}, \forall x, \forall t$$

if $E_1 \neq E_2$, then $|\Psi(x, t)|^2$ and then $\langle x \rangle$ are dependent on time t . The wave function is a periodic function of time.

3.7 The free particle: time-dependent solutions

When $V(x) = 0$, we have

$$\begin{aligned}i\hbar \frac{\phi'(t)}{\phi(t)} &= E \\ -\frac{\hbar^2}{2m} \frac{\psi''(x)}{\psi(x)} &= E\end{aligned}$$

by Eq.??, Eq.?. let $k \equiv \frac{\sqrt{2mE}}{\hbar} > 0$, we get general solution

$$\Psi(x, t) = Ae^{ik(x - \frac{\hbar k}{2m}t)} + Be^{-ik(x + \frac{\hbar k}{2m}t)}$$

let $k \equiv \pm \frac{\sqrt{2mE}}{\hbar}$ the solution becomes

$$\Psi_k(x, t) = Ae^{i(kx - \frac{\hbar k^2}{2m}t)} \quad (17)$$

Eq.?? has a space period λ and a time period T .

Eq.?? is not normalizable.

4 de Broglie Waves

$E < V_0$ and $\hat{H}\psi = E\psi$. V is constant on three pieces of R with function

$$V(x) = \begin{cases} -V_0 & \text{for } -a \leq x \leq a \\ 0 & \text{for } |x| > a \end{cases}$$

The solutions of the time-independent Schrödinger equation is

$$\psi(x) = \begin{cases} e^{ikx} + Ae^{-ikx} & \text{for } x < -a \\ Be^{iqx} + Ce^{-iqx} & \text{for } -a < x \leq a \\ De^{ikx} & \text{for } x > a \end{cases}$$

When $x < 0$,

$$\frac{d^2\psi(x)}{dx^2} = -k^2\psi(x)$$

$$\psi(x) = 1$$

when $0 < x \leq a$,

$$\frac{d^2\psi(x)}{dx^2} = \frac{2m}{\hbar^2}(V_0 - E)\psi(x) \quad (18)$$

$$\frac{d^2\psi(x)}{dx^2} = \mathcal{K}^2\psi(x) \quad (19)$$

$$\psi(x) = B \quad (20)$$

as

$$\boxed{\mathcal{K} = \frac{\sqrt{2m(V_0 - E)}}{\hbar}, \mathcal{K} > 0} \quad (21)$$

when $x > a$,

$$\frac{d^2\psi(x)}{dx^2} = -k^2\psi(x)$$

$$\psi(x) = A \sin(kx) + iB \cos(kx)$$

WAVE PERIOD

4.1 The tunnel effect

ss

5 Uncertainty Principle

5.1 Hilbert space $L^2(\mathbf{R})$

Definition 5.1 (Inner product). *For a pair of functions f and g of functions of one real variable with complex values, i.e., $f, g : \mathbf{R} \rightarrow \mathbf{C}$. The inner product of f and g is*

$$\langle f | g \rangle = \int_{-\infty}^{+\infty} f^*(x)g(x)dx$$

Property 5.1. $\langle f | g \rangle = \langle g | f \rangle^*$

Proof. ... □

Property 5.2. *The inner product is \mathbf{C} -linear on the right and anti- \mathbf{C} -linear on the left. For functions f, g, h and complex numbers λ, μ , we have*

$$\begin{aligned}\langle f | \lambda g + \mu h \rangle &= \lambda \langle f | g \rangle + \mu \langle f | h \rangle \\ \langle \lambda f + \mu g | h \rangle &= \lambda^* \langle f | h \rangle + \mu^* \langle g | h \rangle\end{aligned}$$

Proof. ... □

Definition 5.2 (Hilbert space). *The space of square-integrable functions, denoted by $L^2(\mathbf{R})$, is the set of functions that have finite inner-product with themselves:*

$$L^2(\mathbf{R}) = \{f : \mathbf{R} \rightarrow \mathbf{C}, \langle f | f \rangle < \infty\}$$

at any time t . The square root of the inner product of $f \in L^2(\mathbf{R})$ with itself denoted by

$$\|f\| = \sqrt{\langle f | f \rangle}$$

is called the $L^2(\mathbf{R})$ -norm of f .

Remark 5.1. *The wave function of a particle $\Psi(x, t) \in L^2(\mathbf{R})$.*

5.2 Hermitian operators

Definition 5.3 (Hermitian operator). *Consider an operator Q acts on a function of x , if $\langle f | Qg \rangle = \langle Qf | g \rangle$ for any functions f, g in the Hilbert space and vanish at infinity, then Q is called Hermitian operator.*

Remark 5.2. *Hamiltonian operator \hat{H} is a Hermitian operator.*

Proof. ...

□

Remark 5.3. *Linear momentum p is a Hermitian operator.*

Proof. ...

□

Remark 5.4. *Position x is a Hermitian operator. ($x \in \mathbf{R}$)*

Proof. ...

□

5.3 The Heisenberg uncertainty principle

Theorem 5.1 (The Schwarz inequality in $L^2(\mathbf{R})$). *For any two functions f, g in $L^2(\mathbf{R})$ the following inequality is satisfied:*

$$\|f\| \|g\| \geq |\langle f | g \rangle|$$

Proof. ...

□

Definition 5.4 (Commutator). *For two operators x and p , their commutator is*

$$[x, p] := xp - px.$$

Theorem 5.2 (The Heisenberg uncertainty principle).

$$\sigma_x \sigma_p \geq \frac{\hbar}{2}$$

Proof. ...

□