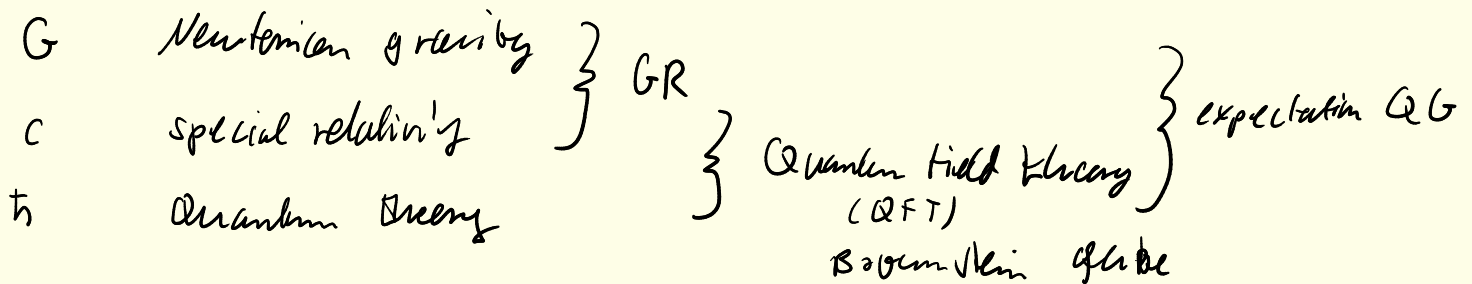


# Quantum Gravity

<https://pirsa.org/20070001> talk he has given  
Klaus Kiefer: Quantum Gravity (2004) book

Disclaimer: we do not have a theory of Quantum Gravity (QG)!  
No consensus about what it would be

Fundamental constants



Planck units (1899):

at which scale a theory comes relevant

$$L_P = \sqrt{\frac{\hbar G}{c^3}} \sim 10^{-35} \text{ m} \quad \text{Planck length}$$

$$M_P = \sqrt{\frac{\hbar c}{G}} \sim 10^{-8} \text{ kg} \sim 10^{19} \text{ GeV} = 10^{16} \text{ TeV} \quad \text{Planck mass}$$

$$t_P = \sqrt{\frac{\hbar G}{c^5}} \sim 10^{-43} \text{ s} \quad \text{Planck time}$$

LHC  $\sim 10^4 \text{ TeV}$  - energy can be measured  
 $\Rightarrow$  no direct tests of QG!

$$T_P = \frac{M_P c^2}{k_B} \sim 10^{32} \text{ K} \quad \text{Planck temperature}$$

What is the problem? Why should we do it?

both GR and QFT are  
incomplete (inconsistent?)

GR:

- singularity theorems:

spacetime singularities unavoidable?

QFT:

- UV divergences  $\rightarrow$  renormalization  
in perturbation

Standard model of particle physics

may not exist as mathematical theory!

$$R_{\mu\nu} - \frac{1}{2} g_{\mu\nu} R = \frac{8\pi G}{c^4} T_{\mu\nu} \quad \langle \Psi | T_{\mu\nu} | \Psi \rangle ?$$

space-time  
geometry  
classical

matter  
quantum

conceptional issue

collapse of wave  
function

Quantum mechanics solved  
this problem

pertubativly quantized GR is non-renormalizable

$$g_{\mu\nu} = \eta_{\mu\nu} + \kappa h_{\mu\nu} \leftarrow \text{spin-2 field}$$

$\ll 1$

$$\frac{1}{\kappa^2} \int \sqrt{g} R = S_{EH} = \int \left\{ \partial h \partial h + \underbrace{\kappa h \partial h \partial h + \kappa^2 h^2 \partial h \partial h + \dots}_{\text{complicated!}} \right\}$$

Einstein Hilbert action

One loop calculation very complicated

two loops: need counter term  $\Gamma_{div}^{(2)} = \frac{1}{\epsilon} \frac{209}{2880} \int R_{\mu\nu\rho\sigma} R^{\rho\sigma\lambda\tau} R_{\lambda\tau}{}^{\mu\nu}$

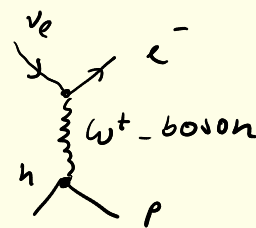
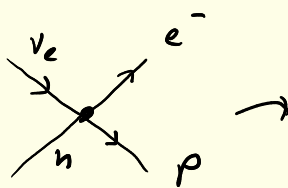
$\epsilon$ -dimensional regularisation  $\epsilon = D - 4$

not predictive! you have to fix an infinite amount of parameters

## (1) particle physics strategy

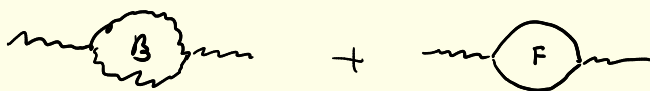
amend Einstein's theory

famous example:  $\beta$ -decay



renormalizable theory  
→ standard model of particle physics

eliminate divergences by adding fermions in calculations



⇒ supersymmetry!

⇒ supergravity = supersymmetric GR

there is a limit for bringing new symmetries

→ most "symmetric" extension of GR N=8 Supergravity

impossible to ask, if it is possible to eliminate all infinities

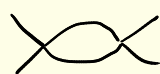
finite up to

$L \leq 5$  loops

point to

rotation of points

superstring theory



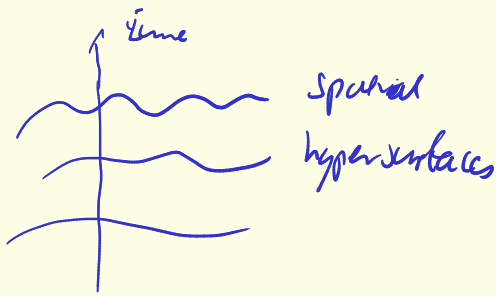
⇒



## (2) Canonical quantization strategy

keep Einstein and quantize non-perturbatively  
go to phase space

$$q, p \rightarrow \text{operators } q, p = i\hbar \partial_q$$



broke spacetime covariance

$$g_{\mu\nu} = g_{\mu\nu}(t, \vec{x}) \quad g_{00}, g_{0m} \quad m=1,2,3 \quad \text{no time dependency}$$

- Lagrange multiplier fields  $\rightarrow$  constraints

$$q'' = g_{mn}(\vec{x}) \quad \pi^{mn}(\vec{x}) = \frac{\delta S_{EH}}{\delta \dot{g}_{mn}}$$

Hamiltonian constraints from  $g_{00}$ ,

$$\mathcal{H}(\vec{x}) = G_{mn, pq} \pi^{mn} \pi^{pq} - \sqrt{g}^{(3)} R^{(3)} \quad \text{for each point a constraint}$$

$\Rightarrow$  infinitesimal constraints

$$\frac{1}{\sqrt{g}} (2g_{mp}g_{qn} - g_{mn}g_{pq})$$

$$\pi^{mn}(\vec{x}) = -i\hbar \frac{\delta}{\delta g_{mn}(\vec{x})} \quad \text{form that you could get Schrödinger equation}$$

$$\nabla^2_{\vec{x}} \mathcal{H}(\vec{x}) \Psi(g_{mn}) = 0 \quad \text{Wheeler-DeWitt equation (1963)}$$

(timeless equation) (like Schrödinger equation but without time dependency)

wavefunction of the universe

notion of time, problem of time in Quantum Gravity

loop quantum theory changing variables here, to tackle the problem

metric is fluctuates