

july 17th

exoplanet classification

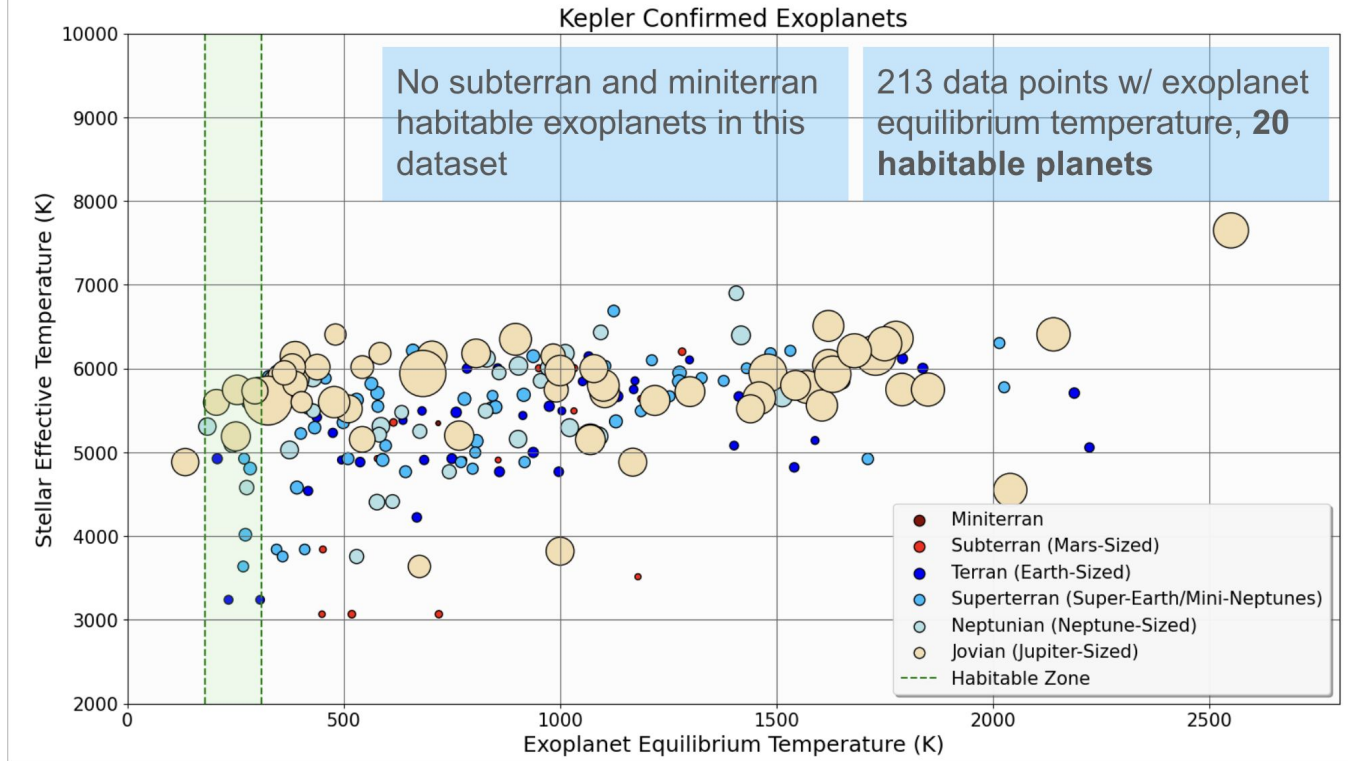
1. new data set received

- meeting on saturday, received a snapshot of the nasa data archive on **march 10, 2024** to better stay consistent with the rest of the HZ exoplanet classification paper

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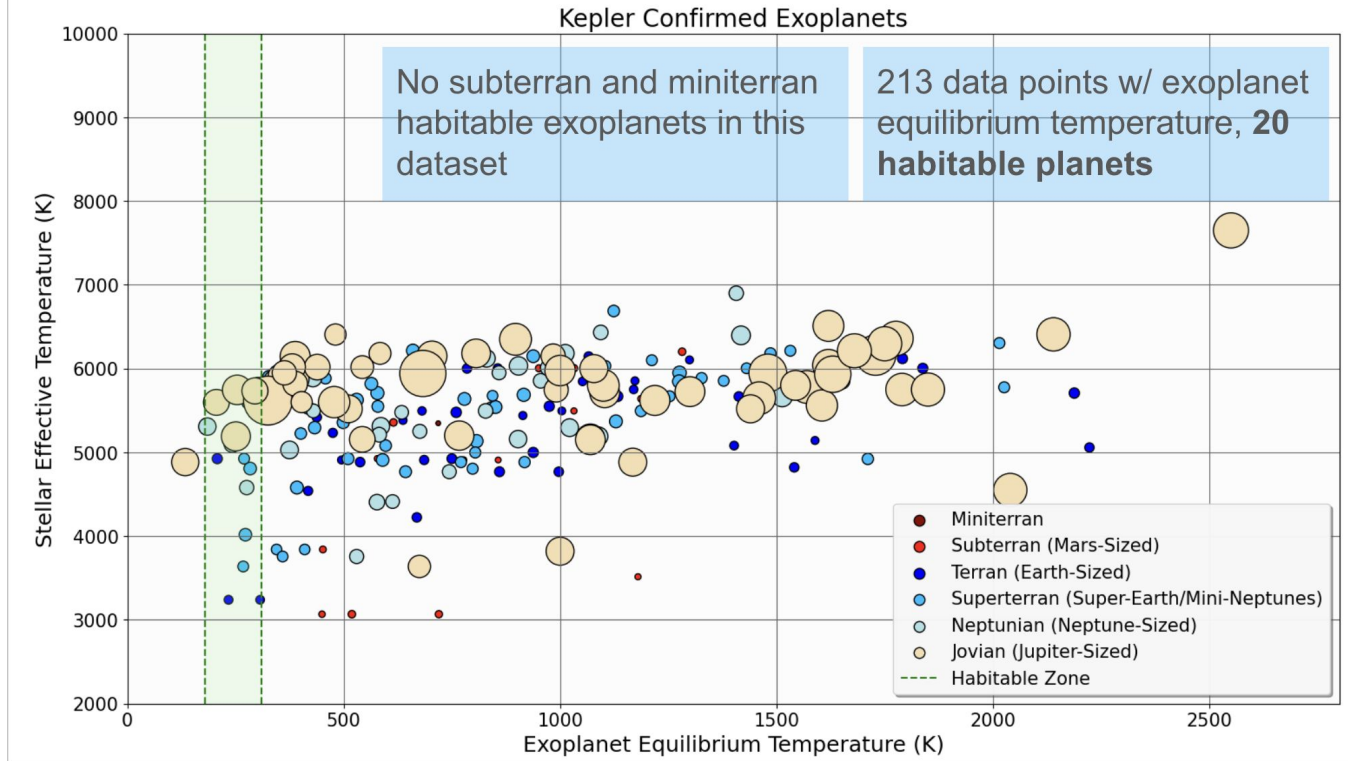
- meeting on saturday, received a snapshot of the nasa data archive on **march 10, 2024** to better stay consistent with the rest of the HZ exoplanet classification paper
- also received a draft of the paper, which i was able to use to better arrange how i calculated habitability zones (which was always a constant struggle in the past)

2. the original graph



< my **june 19th** presentation, with data from [Confirmed Planets Discovered by Kepler](#) (2774 entries)

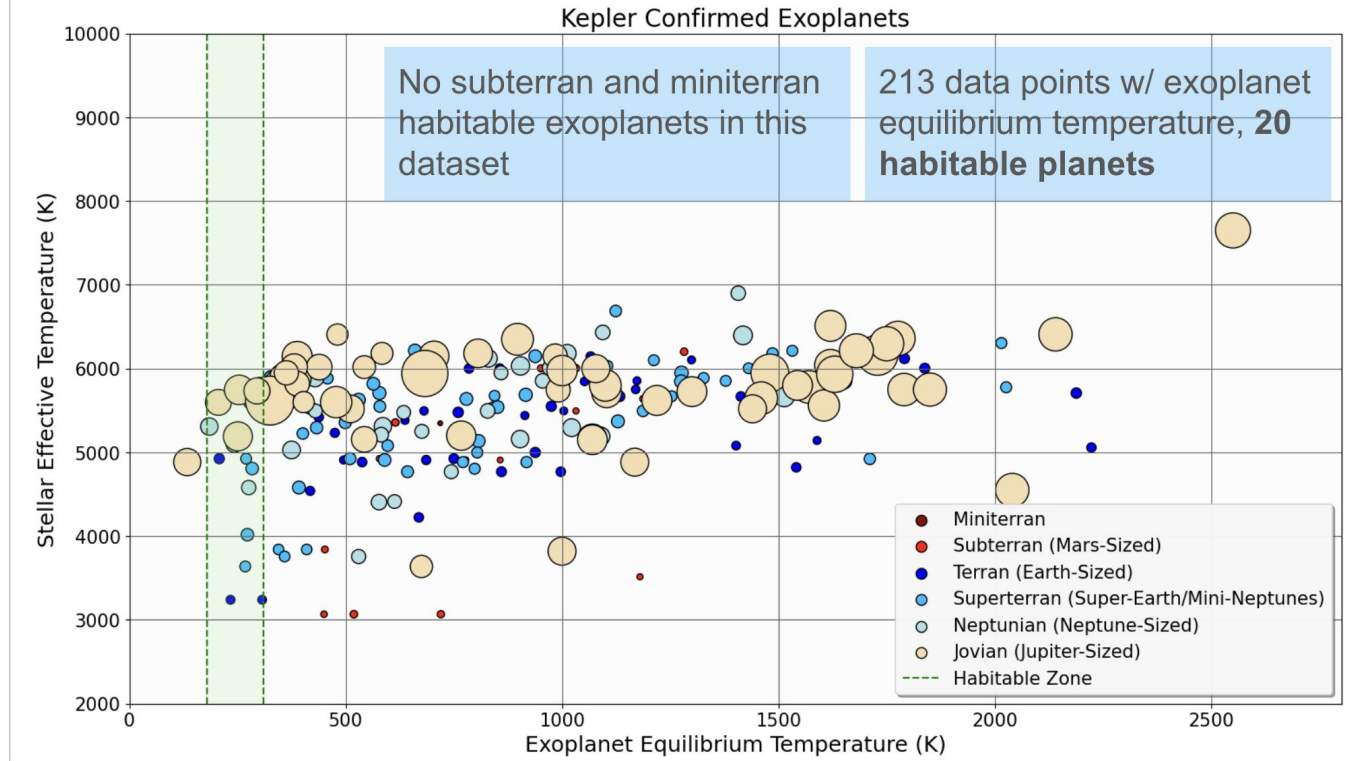
2. the original graph



< my **june 19th** presentation, with data from [Confirmed Planets Discovered by Kepler](#) (2774 entries)

ONLY 213 ENTRIES

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**ONLY 213 ENTRIES
ONLY TEMP. CONSIDERED**

3. defining the habitable zone

one of my biggest frustrations in past graph-making endeavors ... finally fixed!!

2.2 Habitability Zone Determination

The habitable zone of a given exoplanet was determined based on its calculated average surface temperature, denoted as $T_{\text{surf, ave}}$, a critical indicator of potential liquid water presence. We categorized exoplanets as "Too Hot" ($T_{\text{surf, ave}} > 100^{\circ}\text{C}$), "Too Cold" ($T_{\text{surf, ave}} < 0^{\circ}\text{C}$) or within the habitable zone ("In HZ") for $T_{\text{surf, ave}}$ between the benchmark temperature range of 0 to 100°C . This classification was vital to identifying exoplanets that could potentially support life under the assumed necessary precursor of liquid phase H_2O . Utilizing the basic Radiative Equilibrium equation as derived from first principles of radiative heat transfer, the exoplanet average surface temperature calculation considered several characterizing factors. These include the exoplanet's distance from its host star (d), the host's effective surface temperature (T_{\odot}) and radius (R_{\odot}), exoplanet albedo (A) and an additional scaler to account for bulk atmospheric greenhouse gas effect (k):

$$T_{\text{surf, ave}} = k T_{\odot} (1 - A)^{0.25} (R_{\odot} / (2d))^{0.5} \quad \text{Eq 2.2-1}$$

Applying equation 2.2-1 to each listing in the NASA Exoplanet Archive enabled selective sifting of the database to produce quantified results based on the aforementioned exoplanet HZ status categories. Additionally, where the Archive had no entry for the observable parameters T_{\odot} and/or R_{\odot} and/or d , a designation of "N/A" was made for the associated exoplanet to denote insufficient information for determining HZ status in those cases.

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$$T_{\text{surf, ave}} = k T_{\odot} (1 - A)^{0.25} (R_{\odot} / (2d))^{0.5}$$

^^ redefining the habitable zone based on this formula

```
✓ [11] exoplanets_data['pl_tsurf_k'] = 1.13 * exoplanets_data['st_teff'] * ((1 - 0.306) ** 0.25) * (((exoplanets_data['st_rad'] * 696000) / (2 * exoplanets_data['pl_orbsmax'] * 149598023)) ** 0.5)
exoplanets_data['pl_tsurf_c'] = exoplanets_data['pl_tsurf_k'] - 273.15
```

```
✓ [14] exoplanets_data.loc[(~np.isnan(exoplanets_data['pl_tsurf_c'])), 'pl_hz_status'] = 'N/A'
exoplanets_data.loc[(~np.isnan(exoplanets_data['pl_tsurf_c']) & (exoplanets_data['pl_tsurf_c'] >= 0) & (exoplanets_data['pl_rade'] <= 100)), 'pl_hz_status'] = 'In'
exoplanets_data.loc[(~np.isnan(exoplanets_data['pl_tsurf_c']) & (exoplanets_data['pl_tsurf_c'] > 100)), 'pl_hz_status'] = 'Too Hot'
exoplanets_data.loc[(~np.isnan(exoplanets_data['pl_tsurf_c']) & (exoplanets_data['pl_tsurf_c'] < 0)), 'pl_hz_status'] = 'Too Cold'
```

categorized exoplanets as "Too Hot" ($T_{\text{surf, ave}} > 100^{\circ}\text{C}$), "Too Cold" ($T_{\text{surf, ave}} < 0^{\circ}\text{C}$) or within the habitable zone ("In HZ") for $T_{\text{surf, ave}}$ between the benchmark temperature range of 0 to 100°C . This

4. exoplanet proposed classification

based on nasa's classification using **radius** located at

[https://science.nasa.gov/exoplanets/planet-types /](https://science.nasa.gov/exoplanets/planet-types/)

Types of exoplanets

Each planet type varies in interior and exterior appearance depending on composition.

Gas giants are planets the size of Saturn or Jupiter, the largest planet in our solar system, or much, much larger.

More variety is hidden within these broad categories. Hot Jupiters, for instance, were among the first planet types found – gas giants orbiting so closely to their stars that their temperatures soar into the thousands of degrees (Fahrenheit or Celsius).

Neptunian planets are similar in size to Neptune or Uranus in our solar system. They likely have a mixture of interior compositions, but all will have hydrogen and helium-dominated outer atmospheres and rocky cores. We're also discovering mini-Neptunes, planets smaller than Neptune and bigger than Earth. No planets of this size or type exist in our solar system.

Super-Earths are typically terrestrial planets that may or may not have atmospheres. They are more massive than Earth, but lighter than Neptune.

Terrestrial planets are Earth sized and smaller, composed of rock, silicate, water or carbon. Further investigation will determine whether some of them possess atmospheres, oceans or other signs of habitability.

4. exoplanet proposed classification

```
✓ [18] # Determine exoplanet types based on radius
# Type classification is based on https://science.nasa.gov/exoplanets/planet-types/
exoplanets_data.loc[(~np.isnan(exoplanets_data['pl_rade'])), 'pl_type'] = 'N/A'
exoplanets_data.loc[(~np.isnan(exoplanets_data['pl_rade'])) & (exoplanets_data['pl_rade'] <= 1), 'pl_type'] = 'Terrestrial-Panets'
exoplanets_data.loc[(~np.isnan(exoplanets_data['pl_rade'])) & (exoplanets_data['pl_rade'] > 1) & (exoplanets_data['pl_rade'] < 3.88), 'pl_type'] = 'Super-Earths'
exoplanets_data.loc[(~np.isnan(exoplanets_data['pl_rade'])) & (exoplanets_data['pl_rade'] >= 3.88) & (exoplanets_data['pl_rade'] < 9.5), 'pl_type'] = 'Neptunian-Planets'
exoplanets_data.loc[(~np.isnan(exoplanets_data['pl_rade'])) & (exoplanets_data['pl_rade'] > 9.5), 'pl_type'] = 'Gas-Giants'
```

terrestrial planets:

- ≤ 1 times the radius of the Earth

super-earths:

- $1 < r < 3.88$, where the radius of the exopl. is r times as big as that of Earth

neptunian planets:

- $3.88 \leq r < 9.5$

gas giants:

- ≥ 9.5 times the radius of the Earth

5. single-host systems

- cleaned data for the following graphs to just include **single host star systems** (to match with the rest of the paper)

6. general data + results

✓
0s [16] exoplanets_data['pl_hz_status'].value_counts()

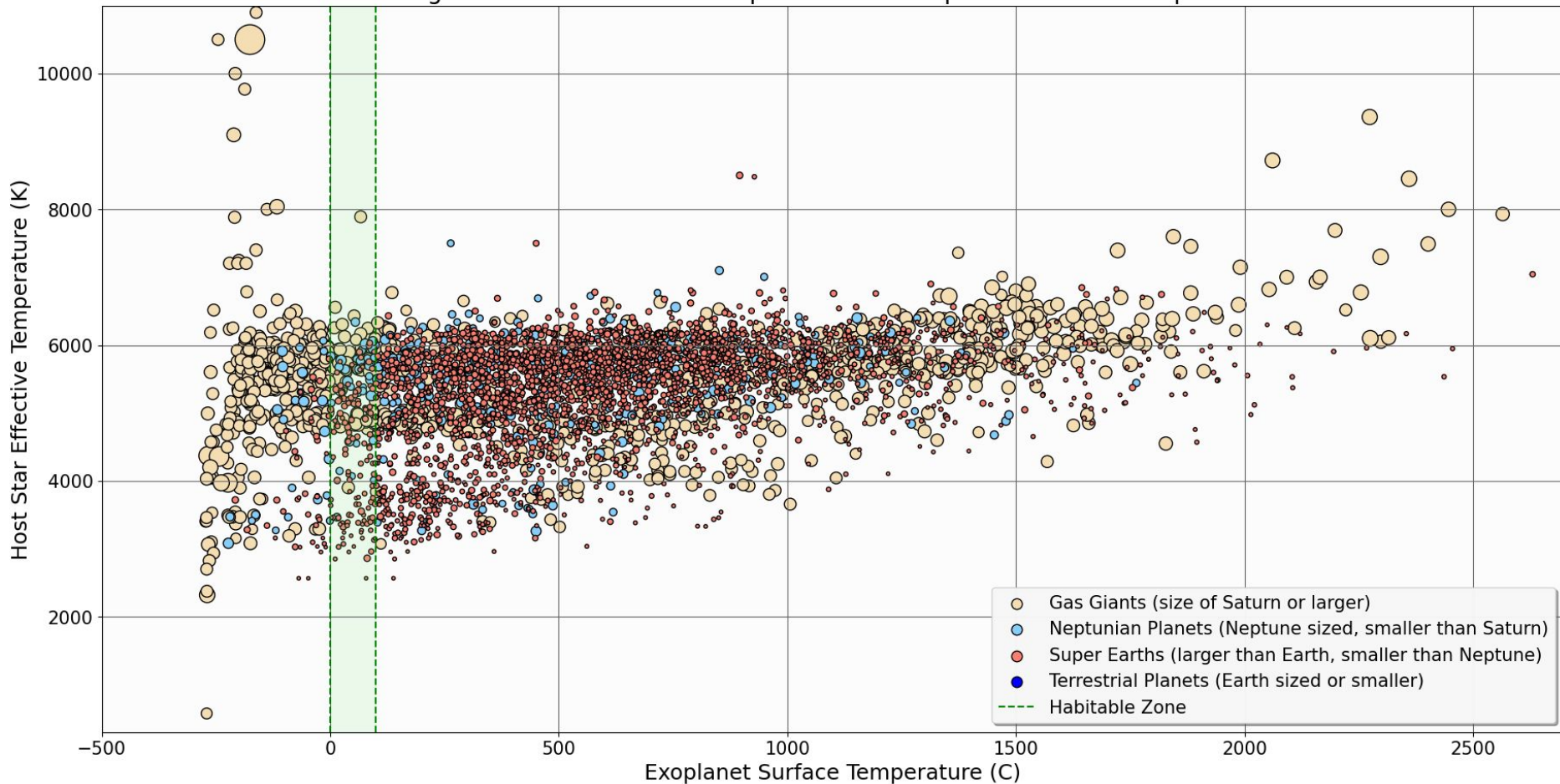
⇒ pl_hz_status
Too Hot 4310
N/A 536
Too Cold 490
In HZ 258
Name: count, dtype: int64

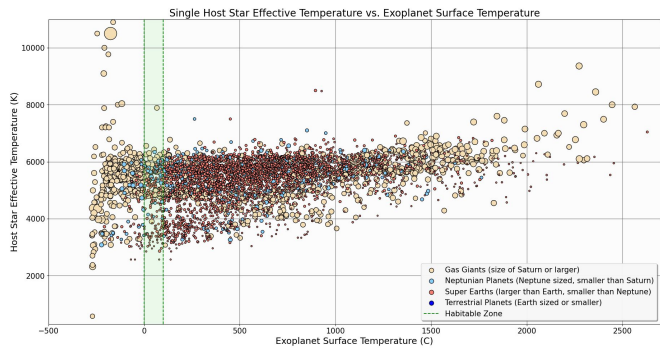
✓
0s [19] exoplanets_data['pl_type'].value_counts()

⇒ pl_type
Super-Earths 3255
Gas-Giants 1598
Neptunian-Planets 509
Terrestrial-Panets 214
N/A 19
Name: count, dtype: int64

6. the newer graphs (with the new 3.10 data)

Single Host Star Effective Temperature vs. Exoplanet Surface Temperature





4583 total planets

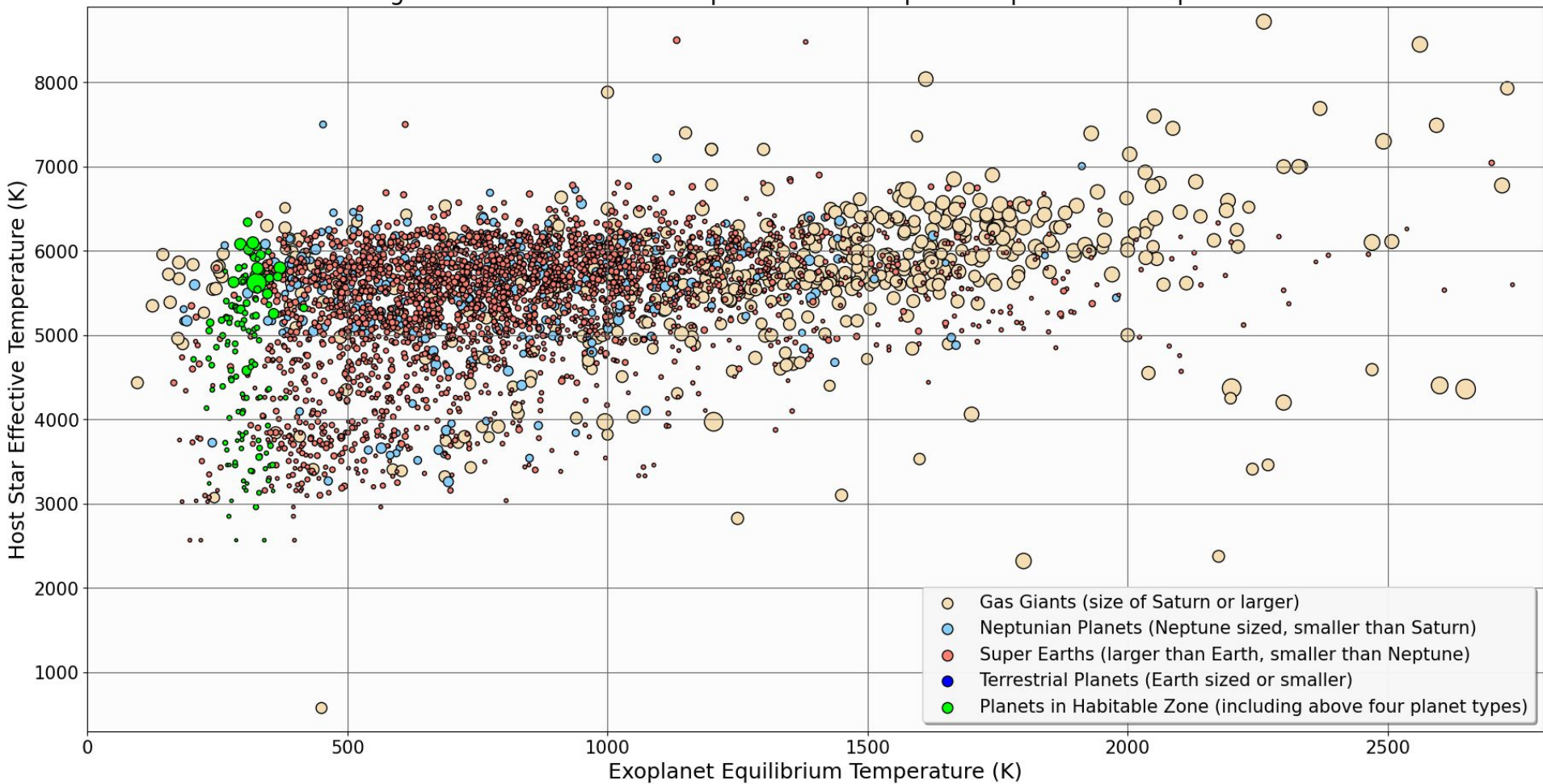
```
✓ [22] st_teff_vs_pl_tsurf_plot_data['pl_hz_status'].value_counts()
```

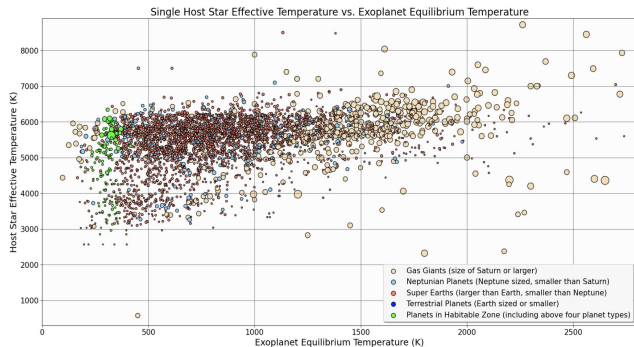
```
⇒ pl_hz_status
Too Hot      3953
Too Cold     403
In HZ        227
Name: count, dtype: int64
```

```
✓ 0s ▶ st_teff_vs_pl_tsurf_plot_data['pl_type'].value_counts()
```

```
⇒ pl_type
Super-Earths      2868
Gas-Giants        1146
Neptunian-Planets  387
Terrestrial-Panets 182
Name: count, dtype: int64
```


Single Host Star Effective Temperature vs. Exoplanet Equilibrium Temperature





3725 total planets

✓
0s

```
[28] st_teff_vs_pl_eqt_plot_data['pl_hz_status'].value_counts()
```



```
pl_hz_status
```

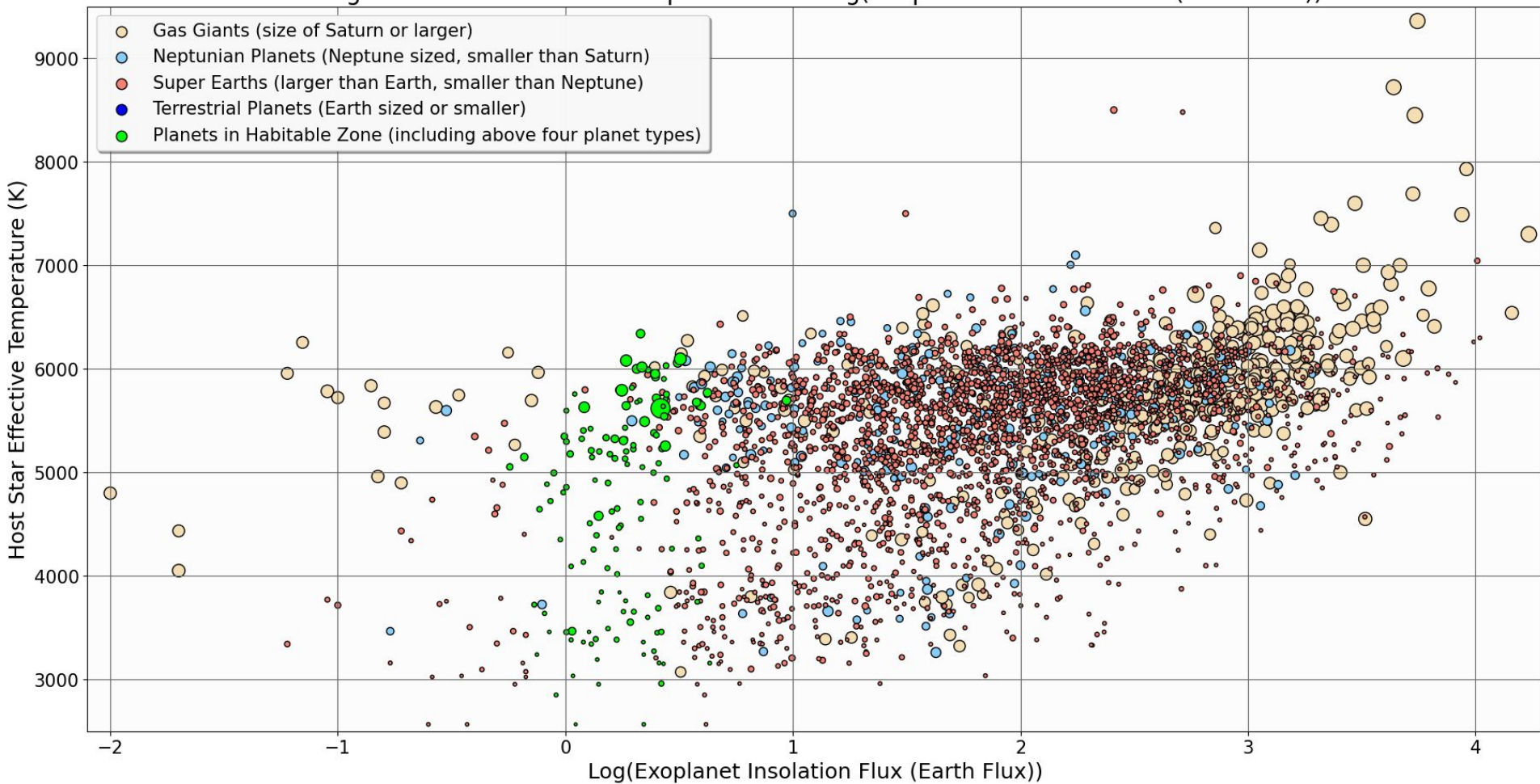
```
Too Hot      3503
```

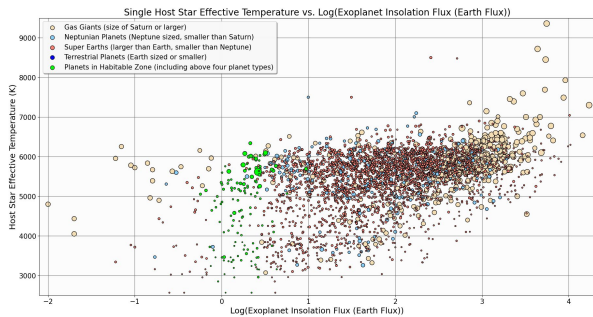
```
In HZ        132
```

```
Too Cold     90
```

```
Name: count, dtype: int64
```

Single Host Star Effective Temperature vs. Log(Exoplanet Insolation Flux (Earth Flux))





3563 total planets

✓
0s

```
[36] st_teff_vs_pl_insol_plot_data['pl_hz_status'].value_counts()
```



pl_hz_status

Too Hot 3373

In HZ 131

Too Cold 59

Name: count, dtype: int64

✓
0s



```
st_teff_vs_pl_insol_plot_data['pl_type'].value_counts()
```



pl_type

Super-Earths 2523

Gas-Giants 469

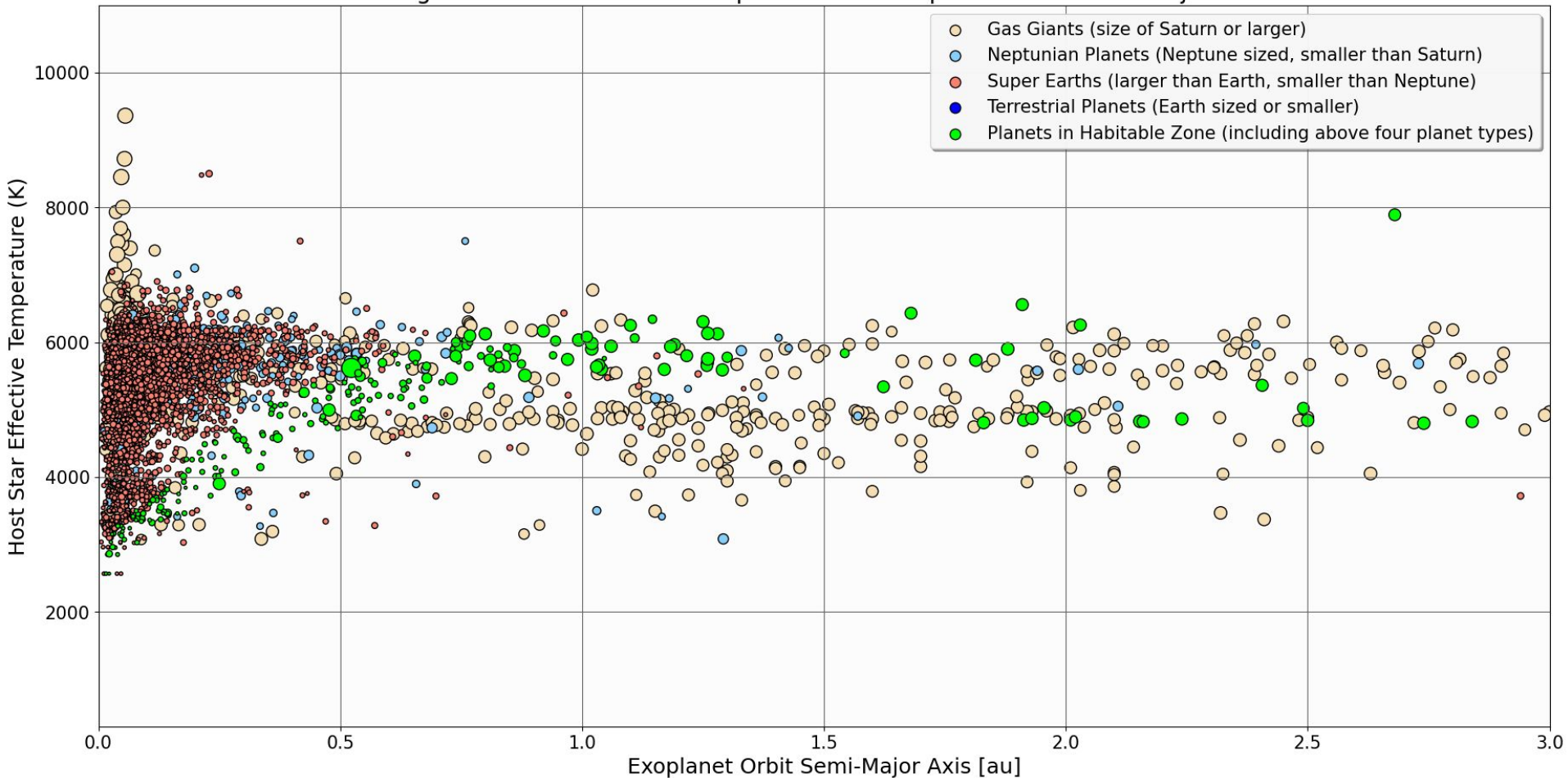
Neptunian-Planets 270

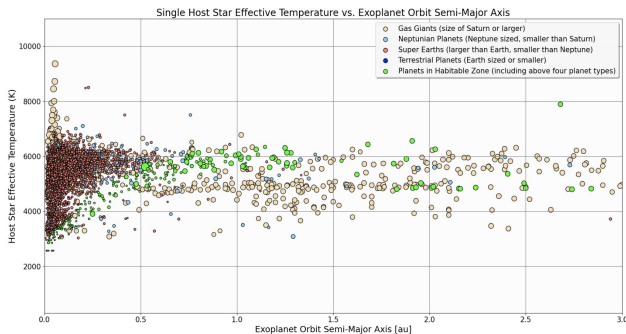
Terrestrial-Panets 170

HabitableZone-Planets 131

Name: count, dtype: int64

Single Host Star Effective Temperature vs. Exoplanet Orbit Semi-Major Axis





4583 total planets

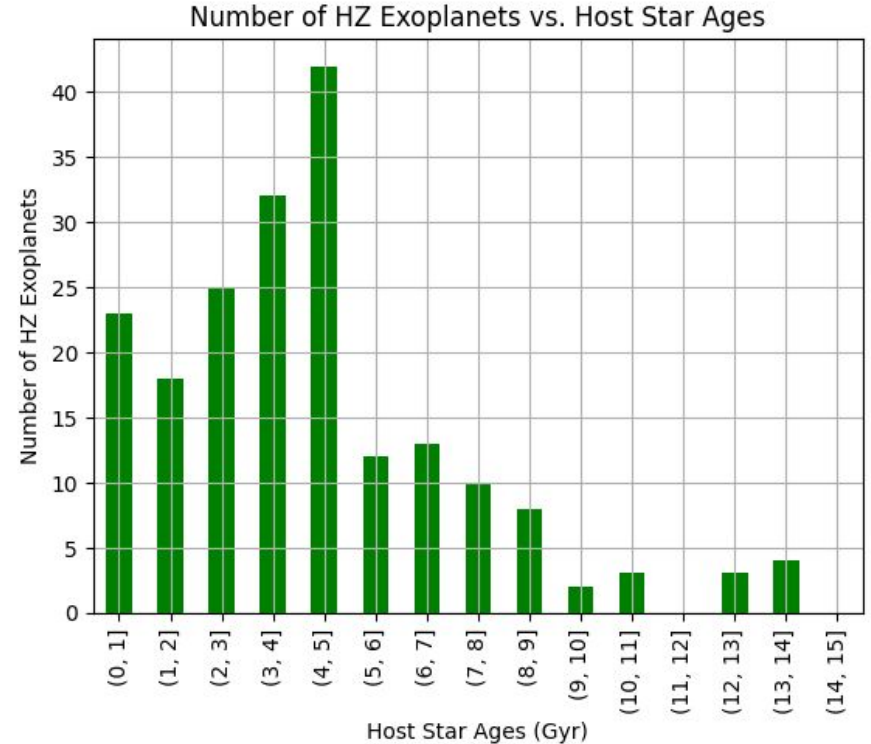
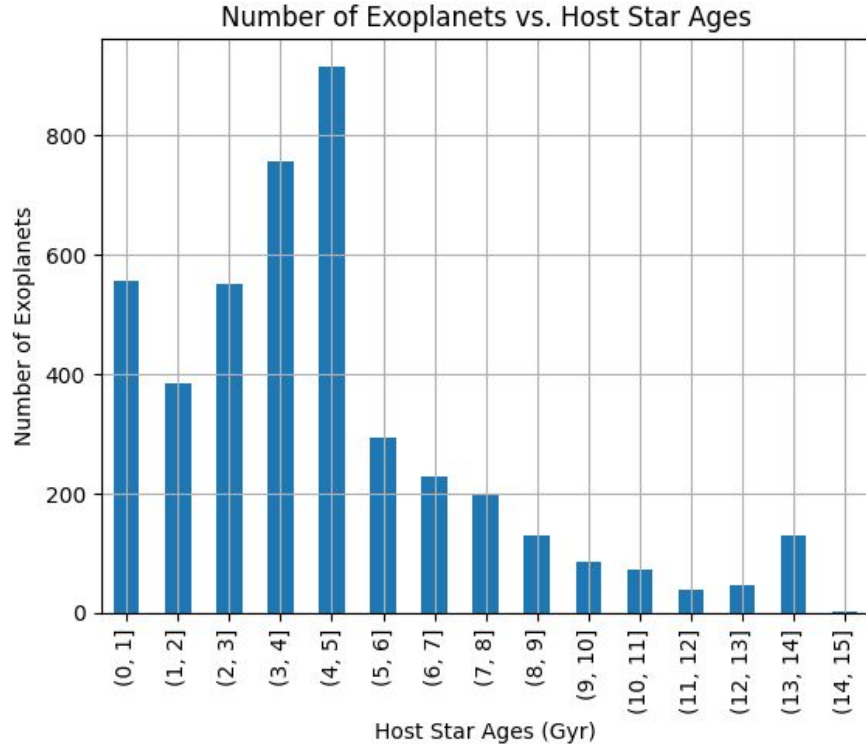
✓ [44] `st_teff_vs_pl_orb_plot_data['pl_hz_status'].value_counts()`

```
⇒ pl_hz_status
Too Hot      3953
Too Cold     403
In HZ        227
Name: count, dtype: int64
```

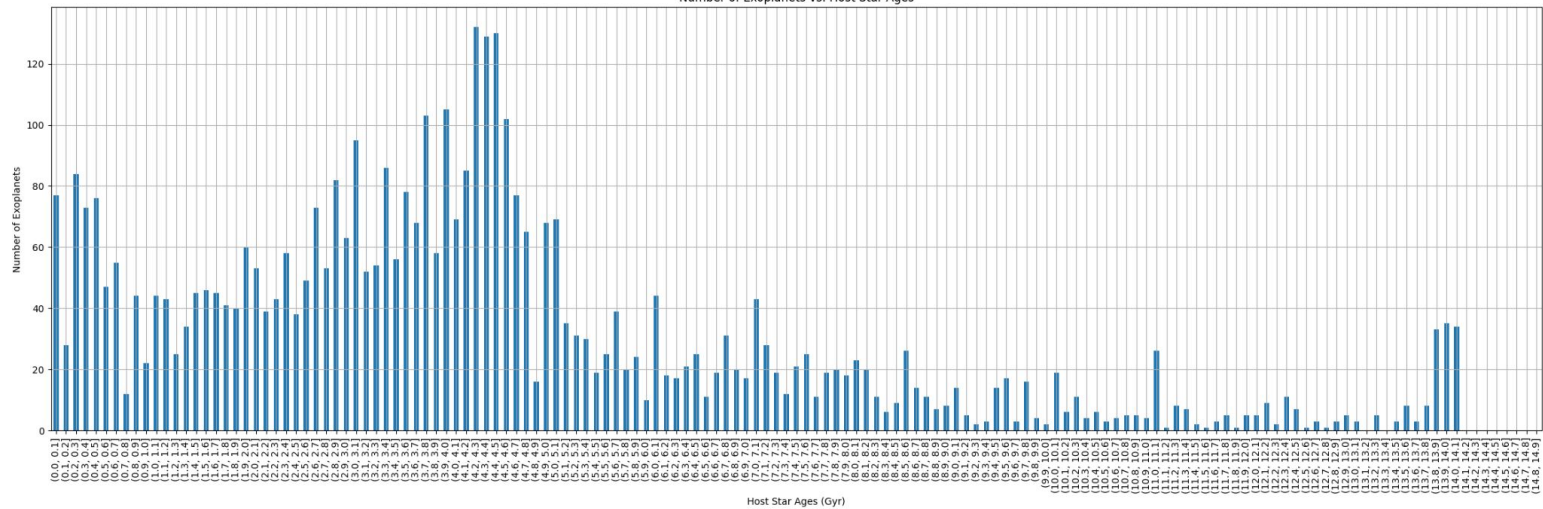
✓ [46] `st_teff_vs_pl_orb_plot_data['pl_type'].value_counts()`

```
⇒ pl_type
Super-Earths      2744
Gas-Giants        1080
Neptunian-Planets  353
HabitableZone-Planets  227
Terrestrial-Panets  179
Name: count, dtype: int64
```

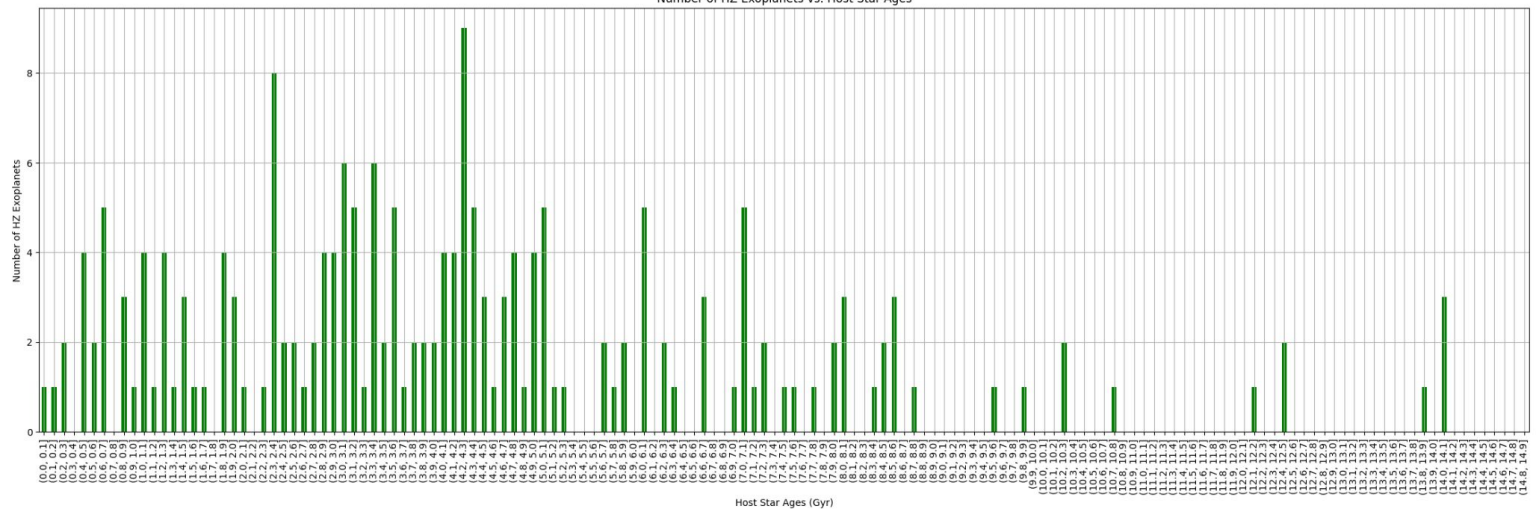
number of exoplanets or HZ exoplanets vs. host star ages



Number of Exoplanets vs. Host Star Ages



Number of HZ Exoplanets vs. Host Star Ages



other graphs

at this link:

https://github.com/christinaxliu/research/blob/main/Caltech-JPL-Intern/HZExoplanetsExploration/PlanetarySystemsCompositeData/HZExoplanetsExploration_PlanetarySystemsCompositeData.ipynb