Unit-level surprise in neural networks

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Abstract

To adapt to changes in real-world data distributions, neural networks must update their parameters. We argue that unit-level surprise *should* be useful for: (i) determining which few parameters should update to adapt quickly; and (ii) learning a modularization such that few modules need be adapted to transfer. We empirically validate (i) in simple settings and reflect on the challenges and opportunities of realizing both (i) and (ii) in more general settings.

1 Introduction

Neural networks have achieved remarkable successes on problems with *independent and identically-distributed* (IID) data [23]. However, real-world data is not IID—environmental conditions shift [8], new tasks are encountered [27], and agents alter their behaviour [6]. Ideally, we would like our networks to transfer or adapt *quickly* to such *out-of-distribution* (OOD) data [30].

Fast adaptation to OOD or *shifted* data is often achieved by updating only a few parameters—the final layer of an ImageNet-pretrained network [13, 42], the batch-normalization layers [24, 34, 38], or specific mechanisms or modules [15, 31]. For a given problem or (expected) *shift type*, a human usually decides: (i) which few parameters should be updated; and (ii) what architecture or *modularization* will allow the network to be adapted by updating only a few parameters (e.g. Rebuffi et al. [34] use residual adapters for domain adaptation). However, we often do not know what type of shift will occur and thus how to decide (i) and (ii).

In this work we argue that the calculation of *unit-level surprise* in neural networks can help identify *what* has changed, and this in turn can be used to: (i) infer which few parameters should be updated to adapt quickly by transferring past knowledge; and (ii) learn an appropriate modularization such that very few modules need be adapted to transfer. For (i), we propose the use of unit-level surprise—"*This looks the same as before, no need for me to update!*"— along with modulatory *update-in-progress* signals—"*This is being dealt with by someone else, hold tight!*". These two additional pieces of information at the unit-level allow the network to update only the appropriate parameters, facilitating fast adaptation by preserving learned structure. This idea is depicted in Figure 1 and further motivated in Appendix A. For (ii), we propose to use the number of surprised units (or groups of units) as a *proxy score* for the number of parameters and modules that need to be adapted to a given shift. Akin to Bengio et al.'s [5] use of adaptation speed, this score can then be optimized to learn a modularization that changes sparsely and thus can be adapted quickly (further discussed in Section 4).

As a first step, we empirically validate the usefulness of unit-level surprise by focusing on use case (i) above. Through experiments on shifted EMNIST [7] datasets we show that unit-level surprise can be used to create shift-dependent fine-tuning strategies that often align with our intuition about which few parameters should update. We then critique these results, identifying several challenges and opportunities of using unit-level surprise in more general settings where its full potential could be unlocked. Despite such challenges, the "beauty" of unit-level surprise—its alignment with our intuitions about how to handle OOD data, its seemingly strong signal for learning how to modularize knowledge, and its relations to neuroscience concepts like metaplasticity [1, 2, 41] and the free-energy principle [12]—convinces us that it *should be better* for realizing (i) and (ii) in more general settings.

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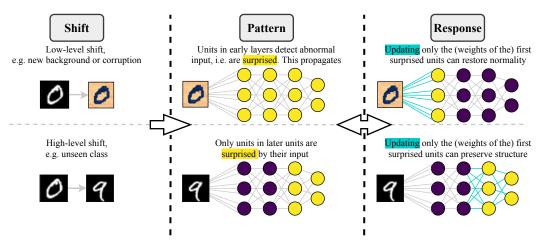


Figure 1: Unit-level surprise can help determine which few parameters should be updated. Purple units are unsurprised, yellow surprised. Blue indicates the weights to be updated. *Top row:* A low-level shift is noticed by units in the first layer (and all those that follow). By blocking those that follow, units in the first layer prevent unnecessary updates in later layers. *Bottom row:* A high-level shift is only noticed by units in later layers, so those in earlier layers don't need to update.

2 Initial validation

In this section we provide an initial empirical validation of the usefulness of unit-level surprise by showing how it can be used to determine which few parameters should be adapted to a given shift.

Experimental setup. We train a simple 5-layer network with 3 convolutional (conv) and 2 fully-connected (FC) layers on 37 of the 47 classes in EMNIST [7]. We then adapt this network to new data from 1 of 10 distribution shifts (see Appendix B.1) for which we expect it to be optimal to update different parts of the network. In particular, we use 7 low-level shifts where we expect only the early layers to need to adapt (e.g. new background) and 3 high-level shifts where we expect only the later layers to need to adapt (e.g. held-out/unseen classes). See Appendix B.2 for full experimental details.

Calculating unit-level surprise. For each unit in a network we store the 1D distribution of its activations under the training data, P(A), and compare this with the distribution of activations under a shifted data distribution, Q(A). More specifically, we parameterize P(A) and Q(A) as softly-binned histograms [40] and infer the parameters of these distributions from the training and shifted data respectively. We then calculate the (Bayesian) surprise of a unit as $s(A) = D_{KL}(Q(A)||P(A))$ [19]. We discuss this quantity in Appendix B.3 and how to calculate it from bin counts in Appendix B.4.

Surprise patterns. Together, a network's surprise values yield shift-dependent patterns. These patterns, shown in Figure 2, help us to understand the level of abstraction at which a shift occurs. For example, when the colour of the EMNIST characters is changed under the crystals shift (see Fig. 3 in Appendix B.1), we intuitively expect the low-level filters (e.g. black-and-white edge detectors) to be surprised—Figure 2a shows that this is indeed the case with units in the first layer exhibiting high surprise. Naturally, the abnormal activations which surprised these units propagate through the rest of the network causing the succeeding units to also be surprised. In contrast, when the network is presented with unseen classes, we intuitively expect the low-level features to be unsurprised (similar distribution of black-and-white edges) but the later layers to be surprised (new combinations of the lowlevel features)—Figure 2b shows that this is indeed the case with units in the later layers surprised while those in earlier layers are not.

Creating an update rule. Once units have been equipped with the ability to calculate their surprise, how can we best

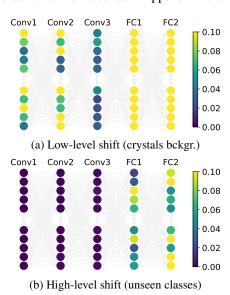


Figure 2: Surprise patterns.

Table 1: 5-shot accuracy: training single layers of a CNN vs. training all layers (R_0) vs. surprise-based update rule (R_2) . L: low-level shifts average, H: high-level shifts average, All: all shifts average.

	Conv1	Conv2	Conv3	FC1	FC2	R_0	R_1	R_2
_							82.7 ± 0.6 86.2 ± 0.3	
Al	157.9 ± 0.4	47.9 ± 0.8	48.6 ± 0.8	63.0 ± 1.6	64.3 ± 0.5	72.0 ± 1.2	$\textbf{83.8} \pm \textbf{0.4}$	81.4 ± 1.6

make use of this? We investigate whether or not surprise can give us some indication of the specific few units to update in order to adapt quickly to new data. In particular, we focus on few-shot learning as updating (the parameters of) fewer units should improve sample-efficiency. We consider three possible adaptation strategies or shift *responses*:

- R₀: SGD (ignore patterns). Ignore surprise patterns and use stochastic gradient descent (SGD).
- R₁: Human-selected layers based on patterns. Use the surprise patterns as a cue to help select specific layer(s) to update. Selecting the right layer can be a powerful fine-tuning technique [35].
- R₂: Update rule based on unit-level surprise. Use a rule that determines whether or not a unit should update based on surprise values. We design an update rule based on the insight that if a unit's parents (units in the preceding layer that are connected to it) are sufficiently surprised then it should not update as the shift that caused its surprise may be resolved by its parents (or even earlier units). The rule can be summarised as: "update only if you are surprised and your parents are not". We formally define this update rule in Appendix B.5.

Results. Table 1 gives the 5-shot results for training individual network layers and using each of our 3 responses. We see that low-level shifts are best dealt with by fine-tuning the first layer and high-level shifts the last. This aligns with our intuition about which units should update and also with the surprise patterns in Figure 2 (making R_1 a strong baseline). Without the need for human intervention (R_1) or a priori knowledge of the type of shift that will occur (choosing Conv1 for low-level shifts, FC2 for high-level), we achieve the best average performance over low- and high-level shifts with R_2 —where units only update if they are sufficiently surprised and their parents are not. These results show that it is not always optimal to fine-tune only the later layers of a network (as also found in [35]) and demonstrate that unit-level surprise can indeed help determine which few parameters should update to adapt quickly (i.e. with few samples) to new data. Results for individual shifts and different amounts of data are given in Appendix E.

Which units does SGD update? We calculate the "distance" that each unit's input parameters are moved when using SGD (R_0) and compare this with R_2 and the layers we find best for fine-tuning in Table 1. As shown in Figure 7 of Appendix E.2, whilst SGD (R_0) moves units in all layers of the network a similar amount, R_2 provides shift-specific behaviour that aligns well with the layers we find best for fine-tuning in Table 1—moving Conv1 for low-level shifts and FC1 for high-level shifts.

3 Critique

While we achieved some promising results in carefully-designed settings, it remains unclear how unit-level surprise should be used in more realistic settings to determine which few parameters to update. Below we discuss the main challenges discovered during the initial idea validation.

Networks are not sufficiently modular to show sparse surprise patterns. For the various shifts we considered, networks did not exhibit sparse or modular surprise patterns—low-level shifts surprised almost every unit in the network (Fig. 2a), while high-level shifts surprised all units in both FC layers (Fig. 2b). We posit that this occurs because, using standard training techniques, simple neural networks do not form modules or mechanisms that change independently (and thus can be surprised independently). Moreover, we find ourselves asking if its *ever* possible to learn such modules or mechanisms that change sparsely (i.e. align with the shifts in data distribution) without explicitly optimizing for it at training time over multiple shifts.

Surprise may not be sufficient. On inspection, we find that R_2 does not always select the "optimal" units to update, e.g. it updates FC1 rather than FC2 to adapt to high-level shifts (Table 1 and Figure 7 of Appendix E.2). This raises some questions—is surprise alone sufficient to determine which units should update? What other unit-level information may be required? Here we list some examples: (a)

gradient magnitude—may help us to update FC2 rather than FC1 to adapt to high-level shifts; (b) task importance [21, 43]—may help in continual learning settings to best preserve performance on past tasks when adapting to new ones; (c) parameter uncertainty—may help speed-up learning by only updating the parameters with high uncertainty [3], akin to Bayesian optimization. Finally, one can imagine shifts that are best resolved at a mid- or high-level but cause changes to the low-level feature distributions—e.g. if we occlude part of a digit, the low-level features (edges) should still be valid, we just need to update how they combine to form the class label. However, the changed low-level feature distributions (some edges are occluded) will surprise units in the early layers, causing them to update. One potential solution is to explore alternative measures that distinguish between being more or less surprised than expected (see Appendix C).

Lack of interpretability makes validation difficult. Finally, it is difficult to come up with experimental setups where we know what are the "optimal" units to update. This makes the design and validation of update rules quite challenging. While we may hope for interpretable edges-parts-wholes feature hierarchies (see Fig. 8 in Appendix E.3), this is often not the case. In fact, it can take as little as a change in random seed to change the semantic meaning of features, and with it, the optimal units to update. Work on unit-level interpretability [14, 29] may eventually help, but it is not currently at a level to decide which units should update for a given shift.

4 Future work

As discussed above, creating a general update rule using unit-level surprise is challenging as: (i) neural networks are not very modular by default; and (ii) surprise may not be sufficient to determine "who" should update. Below we discuss two potential avenues to overcome these two challenges.

Learning modular structure with unit-level surprise. Recent work in causal discovery assumes that the ground-truth data-generative process consists of independent mechanisms or modules, with many modules expected to behave similarly across different tasks and environments [5, 31, 32, 36, 37]. Thus, if an appropriate modular representation of the world has been learned, *very few modules should need to be adapted in order to transfer* [5, 15, 32, 37]. Bengio et al. [5] exploit this through a meta-learning objective that uses *adaptation speed* to optimize the way in which knowledge is represented or modularized. Here, adaptation speed is a proxy score for the number of modules and parameters that need be adapted and ultimately for how well the learned modularization fits the underlying causal dependencies. We believe surprise could provide an *alternative proxy score* (e.g. the number² of surprised units or groups of units) that is easier to meta-optimize (no inner-loop steps) and arguably a more direct measure of the number of modules and parameters that need be adapted.

Learning optimizers that combine unit-level surprise with other unit-level information. As discussed in the previous section, unit-level surprise may not be sufficient for some shifts—further information may be required to determine "who" should update. This makes it difficult to handcraft a general update rule as one must specify how surprise should be used with this other information. One solution is to instead *learn* an optimizer [4, 33], or the parameters of an update rule [25], across multiple shifts. However, learned optimizers are notoriously difficult to train [28] and it can be difficult to beat the heavily-tuned few-shot benchmarks against which they are often compared [16].

5 Conclusion

In this work we provided a preliminary validation of the usefulness of unit-level surprise in neural networks. In particular, we showed that it can be useful for analysing *where* dataset shifts are noticed in a network and for devising *shift-dependent* fine-tuning strategies. We also discussed some of the "beauty" that makes us believe that unit-level surprise *should* be useful in more general settings—it often aligns with our intuitions about how to handle OOD data, it seems a strong signal for learning how to modularize knowledge such that only a few modules are surprised (and thus need be adapted), and it has several neuroscience relations like metaplasticity and the free-energy principle. However, in formally specifying how to use unit-level surprise in simple settings, we encountered several challenges that made us wary of tackling the more complicated settings in which its full potential could be realized.

²A continuous proxy can be used for differentiability.

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A Example to further motivate surprise

Imagine you are given a network which has been trained to classify different Ford car models outdoors (in green pastures), and asked to classify pictures of these same models indoors (in a showroom). Intuitively, most of the learned function is still valid so only a small number of parameters need to be tweaked, requiring only a small number of indoor examples to do so. Perhaps tweaking parameters in the first layer would be sufficient, re-extracting the same features from e.g. the darker pixels of indoor images as were previously extracted from the outdoor images. Similarly, if you were asked to classify pictures of Fiat cars, you may expect that only a small number of parameters need to be adapted (likely in later layers this time), perhaps adjusting the (conceptual) wheel and mirror detectors for the smaller wheels and more rounded mirrors of Fiats. However, as we discuss in Section 2, minimizing a standard loss function (e.g. cross-entropy) with *stochastic gradient descent* (SGD) does not result in such intuitive updates—all of the network's parameters are updated and learnt structure is unnecessarily destroyed. These gradients tell us which parameters *can* reduce the error, but not which parameters *should* reduce the error in order to maximally-transfer knowledge and thus speed-up learning. Doing so requires additional information—such as unit-level surprise.

B Implementation details

B.1 Data



Figure 3: EMNIST-DA shifts. Figure adopted from [9].

Our networks are trained using the 47-class EMNIST dataset [7] ("identity" shift in Figure 3). We use 2000 samples per class from the training split with the remaining 400 forming a validation set, we report results on the separate test set also containing 400 examples per class.

We adapt to 10 new data distributions—7 low-level shifts from EMNIST-DA [9] and 3 high-level label shifts. From the 14 EMNIST-DA shifts depicted in Figure 3, we chose 7 shifts that adversely affect accuracy (without adaptation) and intuitively affect the early convolutional filters: crystals, fog, gaussian blur, grass, impulse noise, sky and stripe. To create the 3 label shifts, we train our networks on only the first 37 classes of EMNIST and then choose 5 of the 10 unseen classes three times from the range [38, 47] to arrive at: H1:[38, 39, 40, 41, 42], H2:[43, 44, 45, 46, 47], and H3:[38, 40, 42, 44, 46].

To evaluate sample efficiency we run experiments with varying amounts of data, we experiment with using 2, 5, 10, 20 and 50 samples per class as well as using all the data (2000 samples per class). The full results of these experiments are given in Appendix E.

B.2 Experimental setup

Table 2 provides the architectural details of the simple 5-layer convolutional neural network (CNN) that we use. During pretraining we use dropout between the layers, for adaptation we do not as it unnecessarily complicates the propagation of surprise and makes little difference to the final results.

During pretraining and adaptation we use a batch size of 256. We pretrain for 150 epochs with a learning rate of 0.01. During adaptation we train for 50 epochs with a learning rate of 0.1 for all experiments except for when fine-tuning all layers simultaneously which requires a learning rate of 0.01 to prevent divergence. We optimize using stochastic gradient descent with momentum set to 0.9.

For experiments using R_2 the thresholds α and β in Equation 3 are set to 0.01. Experiments are run over 3 seeds from which we report a mean and one standard deviation.

Table 2: Architecture of the CNN used. For conv. layers, the weights-shape is: num. input channels \times num. output channels \times filter height \times filter width.

Layer	Weights-Shape	Stride	Padding	Activation	Dropout Prob.
Conv	$3\times 64\times 5\times 5$	2	2	ReLU	0.1
Conv	$64 \times 128 \times 3 \times 3$	2	2	ReLU	0.3
Conv	$128 \times 256 \times 3 \times 3$	2	2	ReLU	0.5
FC	6400×128	N/A	N/A	ReLU	0.5
FC	128×47	N/A	N/A	Softmax	0

B.3 Measuring unit-level surprise

A single unit in a feed-forward neural network outputs an activation $a = g(\mathbf{w}^T\mathbf{h} + b)$, where \mathbf{h} is the hidden unit activations of the previous layer, \mathbf{w} the learned weight vector, b the learned bias and g some non-linearity. During training a unit can store a distribution P(A) which captures the distribution that its activation can take. A unit is surprised by new data if the activation distribution changes, i.e. $P(A) \neq Q(A)$, where Q(A) is the distribution of the unit's activations under the new data. We quantify surprise using the KL-divergence from P(A) to Q(A), i.e. $s(A) = D_{KL}(Q(A)||P(A))$ [19, 26].

The surprisal (or information content) of an event X=x, with $X\sim P(X)$, is given by $\log(1/P(x))$. Intuitively, this quantity represents how "surprised" we are to see X=x, with unlikely events having high surprisal. Surprise itself is a somewhat overloaded term but can be used to describe the entropy, $H(X)=-\sum P(x)\log P(x)$, that is the expected surprisal. If we now receive a sample of a different random variable $Y=y,Y\sim Q(Y)$, but we have assumed as a prior that we are receiving samples from P(X), the amount of additional surprisal we receive on account of our assumption is $\log(1/P(X=y))-\log(1/Q(Y=y))$. The expected value of this quantity over Q is the KL-Divergence $D_{KL}(Q||P)$ [22], that is the expected surprisal of receiving samples from Q when we have assumed the distribution to be P. We can also interpret $D_{KL}(Q||P)=H(Q,P)-H(Q)$ as the expected extra message-length per datum that must be communicated if a code that is optimal for P is used to communicate Q, compared to a code that is optimal for Q. We have seen this quantity referred to as Bayesian surprise, information gain, asymmetric surprise or simply surprise [10, 19]. Throughout this work we refer to this quantity as surprise for simplicity.

B.4 Calculating surprise from bin counts

After pre-training we parameterize activation distributions with softly binned histograms. To calculate P(A) we run one further forwards pass of the network over the training data and bin the activations using the same procedure as in [9], with 10 bins, which outputs 10 normalized bin counts, $\pi_1^p, \dots \pi_{10}^p$, for each unit. π_i^p represents the probability a falls into bin i and $\sum_{i=1}^{10} \pi_i^p = 1$. Example of such distributions for single units are shown in Figure 4 (blue curves), these distribution can be considered as representing the "normal" activation values the unit expects to take on.

We then receive some data from a new distribution, possibly the same as the pre-training data distribution, which is fed into the neural network. During adaptation we can parameterize Q(A) in the same way as P(A) using this new data (Figure 4–orange curves), to receive normalized bin counts $\pi_1^q, \ldots \pi_{10}^q$, which change as the network learns. The surprise for a unit can then be calculated as

$$s(A) = D_{KL} \left(Q(A; \{ \pi_i^q \}_{i=1}^{10}) \mid\mid P(A; \{ \pi_i^p \}_{i=1}^{10}) \right) = \sum_{i=1}^{10} \pi_i^q \log \frac{\pi_i^q}{\pi_i^p}$$
(1)

 $^{^3}$ For convolutions, when creating P(A) and Q(A) we take each spatial location of a feature map to be one sample of the activation, a.

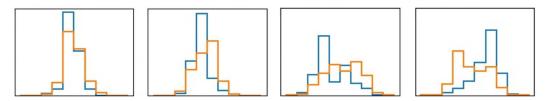


Figure 4: Examples of our histogram parameterizations of P(A), in blue, and Q(A), in orange. When new data is received, the activation distribution changes. From left to right, the surprise values s(A) are approximately 0.1, 0.2, 0.3, 0.4.

B.5 A surprise-based update rule

To formulate the update rule response, R_2 we consider a single unit in a neural network which has a surprise value s as defined in Equation 1. We also consider each of the K parents of this unit, that is the units in the preceding layer which are connected to the unit in question (the child unit). Each parent unit has a surprise value and a weight connecting it to the child unit, denoted $\{s_i\}_{i=1}^K$ and $\{w_i\}_{i=1}^K$ respectively. We aggregate the parent surprises into one value for the child unit using normalized weight values to get a comparable scale across units in a layer, specifically we calculate

$$p = \sum_{i} \frac{|w_i|}{\sum_{j} |w_j|} \cdot s_i \tag{2}$$

Each weight in $\{w_i\}_{i=1}^K$ updates if the aggregated parent surprise is lower than some threshold (that is the parents are unsurprised) and the child surprise is over some threshold (sufficiently surprised). So for a single weight w we have the update rule

$$w := w - \mathbb{I}[s > \alpha] \mathbb{I}[p < \beta] \cdot \eta \nabla \mathcal{L}, \tag{3}$$

where $\mathbb{I}[s>\alpha]$ is an indicator function that is 1 when $s>\alpha$ and 0 otherwise, $\mathbb{I}[p<\beta]$ is similarly defined, η the learning rate, and $\mathcal L$ the loss function (cross-entropy in our case). $\mathbb{I}[s>\alpha]$ ensures that only surprised units update, and can be compared with *metaplasticity*⁴ in the brain. $\mathbb{I}[p<\beta]$ prevents/blocks simultaneous changes in later units who may also surprised by their input, and can be compared with *neuromodulation*⁵ in the brain.

⁴The modification of a neuron's future capacity for learning as a function of recent synaptic history [1, 2]. Believed to regulate the plasticity mechanisms themselves in order to generate adaptive behaviour [11, 17, 39, 41].

⁵Neuromodulators are neurotransmitters which, instead of conveying excitation or inhibition, change the properties of other neurons or synapses [20].

C Alternative surprise measures

Imagine a unit or feature detector with a bi-modal activation distribution where the modes roughly represent on (detected) and off (not detected). Perhaps such a unit being off *more often* in the new data (as when part of an image is occluded, discussed in Section 3) is not a good signal to update. To differentiate this situation from e.g. the unit being *on* more often in the new data, we can define a new measure which we call the *surprise increase* (SI):

$$SI = H(Q, P) - H(P), \tag{4}$$

where H(P) is the entropy of P and H(Q,P) is the cross-entropy of P relative to Q. Unlike the KL-divergence, SI can be negative. SI is negative (i.e. a surprise decrease) when an event that is already quite likely under P becomes even more likely under Q—as shown in Figure 5a. This could be an interesting alternative surprise measurement as it can distinguish between surprise increases and decreases.



Figure 5: SI illustration. P is blue, Q is orange. For fixed P, SI depends only on H(Q, P).

D Does batch normalization solve the problem?

Batch normalization [18] standardizes activation distributions across batches, bringing P(A) and Q(A) closer together. This naturally raises the question as to whether or not batch norm solves the problem of differing unit-level distributions, thus removing the need for unit-level surprise. We investigate this below.

Do the same surprise patterns exist? Figure 6 shows the ideal situation for batch norm, where the batch norm statistics are calculated from the new data distribution before we calculate surprise. Compared to Figure 2, the magnitude of the surprise is indeed lessened as batch norm standardizes P and Q, but it does not make units unsurprised. We still see the same patterns of surprise in the early layers for low-level shifts and in the later layers for high-level shifts.

Is R_1 still the best adaptation strategy? It also still remains superior to train specific layers for specific shifts (akin to R_1 which chooses based on surprise) as shown in Table 3. That is, batch norm alone does not solve the problem of changes in activation distribution. In this table we also show that if we only use the batch norm statistics without any training (AdaBN [24]), or only update the batch norm parameters and statistics on the new data (BN params), performance is poor compared to the other strategies.

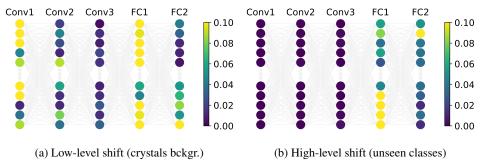


Figure 6: Network surprise patterns after updating the batch normalization statistics on the new data.

Table 3: 5-shot accuracy with batch norm. *L:* low-level shifts average, *H:* high-level shifts average. Zero-shot shows accuracy before adaptation. AdaBN [24] updates the batch-norm statistics on the new data. BN params updates only the batch-norm parameters and statistics. For all other methods/rows, only the layers listed are permitted to update.

	L	Н
Zero-Shot	22.1 ± 0.2	0.0 ± 0.0
AdaBN	50.7 ± 0.4	0.0 ± 0.0
BN params	66.0 ± 1.3	0.0 ± 0.0
Conv1, BN1	78.9 ± 0.8	0.0 ± 0.0
Conv2, BN2	67.8 ± 1.2	0.0 ± 0.0
Conv3, BN3	60.3 ± 0.8	0.2 ± 0.2
FC1, BN4	56.4 ± 0.5	61.4 ± 3.3
FC2	52.7 ± 0.6	95.4 ± 0.5
FC1, BN4, FC2	55.8 ± 0.8	94.4 ± 0.8
All	76.7 ± 1.2	90.5 ± 2.3

E Further Results

E.1 Varying the amount of data

After pre-training on the first 37 classes of EMNIST our network achieves a test accuracy of $92.0 \pm 0.1\%$. In Table 4 we also evaluate the results of applying each of the responses, described in Section 2, to new data distributions with different numbers of samples. We see that when using large amounts of data (2000 shot), all responses, including simply fine-tuning all layers of the network, perform well achieving similar accuracy as on the pre-training data. More interestingly as the number of samples reduces (5-50 shot) the choice of response becomes more important. With few samples, training all layers of the network with no additional considerations for structure preservation (R_0) is outperformed by both other strategies that use surprise patterns to select which units should update. R_1 , where a human selects which layers to train, in this case the first convolutional layer for the low-level shifts and both linear layers for the high-level shifts, consistently outperforms R_2 , although performance is similar except for in the very low data regime (2 shot).

Table 4: Accuracy for a varying number of shots (samples-per-class), averaged over different distribution shifts. Reported is the mean and 1 standard deviation over 3 random runs.

	2	5	10	20	50	2000
R_0	58.7 ± 2.2	72.0 ± 1.2	78.7 ± 0.4	82.3 ± 0.2	85.6 ± 0.2	91.6 ± 0.1
R_1	78.1 ± 0.6	83.8 ± 0.4	86.2 ± 1.0	88.0 ± 0.3	89.8 ± 0.0	92.3 ± 0.0
R_2	66.5 ± 2.9	81.4 ± 1.6	86.0 ± 0.3	87.9 ± 0.5	89.4 ± 0.1	92.3 ± 0.0

E.2 Which units does SGD update?

In order to evaluate which units are updated by different responses we calculate the euclidean distance between a parameter's initial value (after pre-training) and its value after adaptation (applying one of the responses). We do this for each parameter in the neural network and average the distance moved across all parameters in a layer (weights and biases) to get an overview of which layers are being changed by the responses. This is shown in Figure 7.

Figures 7a & 7b show that R_0 moves all layers of the network similarly and makes no distinction between different types of shift. This confirms that a simple gradient signal is not sufficient to optimally-update the network's structure. On the other hand, when selecting units to update with R_2 , Figures 7c & 7d show that we get shift-specific behaviour which aligns with the results from Table 1—training conv layers is better for low-level shifts and FC layers for high-level shifts.

E.3 Max-activating patches

Figure 8 shows the maximum-activating image patches on the EMNIST-DA [9] grass shift for selected units in each layer. Note that we do not have a perfect edges-parts-wholes hierarchy—the receptive field of Conv2 seems too large as it almost sees the entire image (can be a "whole" rather than a part).

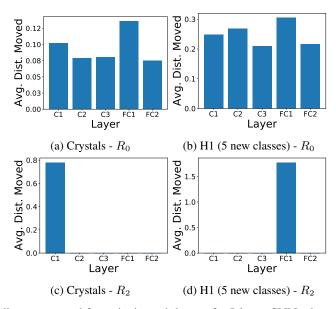


Figure 7: Mean distance moved for units in each layer of a 5-layer CNN when using R_0 and R_2 .

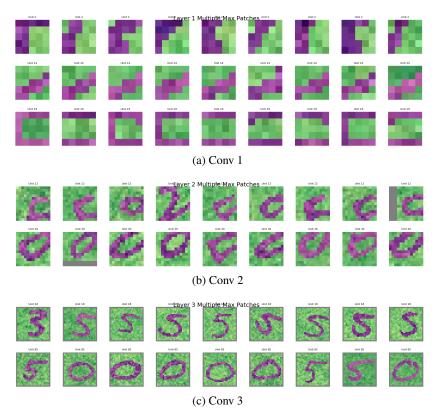


Figure 8: Max-activating patches on the EMNIST-DA grass shift.

E.4 Per-shift Results

What follows is tables of full results for each shift at each number of shots, these are provided for completeness with no further analysis.

Table 5: 2-shot accuracy across shifts: training different layers of a CNN

	Conv1	Conv2	Conv3	FC1	FC2	FC1 + FC2
H1	0.0 ± 0.0	0.0 ± 0.0	5.4 ± 5.6	65.8 ± 12.8	95.2 ± 1.4	77.5 ± 2.7
H2	0.0 ± 0.0	0.0 ± 0.0	4.5 ± 1.1	45.4 ± 8.5	80.5 ± 5.3	65.0 ± 6.3
Н3	0.0 ± 0.0	0.0 ± 0.0	7.1 ± 4.5	50.8 ± 6.2	93.0 ± 2.3	78.5 ± 8.8
Crystals	72.2 ± 4.0	51.5 ± 1.5	47.7 ± 1.1	45.0 ± 1.3	40.5 ± 1.3	43.6 ± 1.2
Fog	88.7 ± 0.7	83.9 ± 1.0	80.2 ± 0.8	77.2 ± 0.7	78.2 ± 0.3	76.9 ± 0.6
Gauss. Blur	82.3 ± 0.6	77.8 ± 0.1	75.5 ± 1.2	71.4 ± 0.6	70.6 ± 1.2	71.3 ± 0.8
Grass	81.2 ± 1.1	21.2 ± 8.3	8.7 ± 0.5	6.8 ± 0.5	7.1 ± 0.2	7.1 ± 0.9
Imp. Noise	87.9 ± 0.8	84.9 ± 0.3	82.6 ± 0.6	81.0 ± 0.8	78.8 ± 1.3	78.8 ± 0.6
Sky	74.9 ± 3.3	54.7 ± 2.2	33.5 ± 3.2	18.7 ± 0.9	14.0 ± 2.0	16.7 ± 0.3
Stripe	72.8 ± 1.8	59.9 ± 3.5	56.2 ± 2.6	42.8 ± 2.9	38.1 ± 0.5	45.2 ± 4.0
Avg High	0.0 ± 0.0	0.0 ± 0.0	5.7 ± 2.4	54.0 ± 3.5	89.6 ± 1.5	73.7 ± 1.7
Avg Low	80.0 ± 1.3	62.0 ± 2.0	54.9 ± 1.0	49.0 ± 0.5	46.8 ± 0.3	48.5 ± 0.6
Avg All	56.0 ± 0.9	43.4 ± 1.4	40.1 ± 0.5	50.5 ± 0.7	59.6 ± 0.4	56.1 ± 0.9

Table 6: 2-shot accuracy across shifts: comparison of different responses. Zero-shot is the accuracy before any updates are performed.

	Zero-shot	R_0	R_1	R_2
H1	0.0 ± 0.0	53.7 ± 6.1	77.5 ± 2.7	32.0 ± 21.2
H2	0.0 ± 0.0	45.3 ± 14.4	65.0 ± 6.3	35.0 ± 10.5
H3	0.0 ± 0.0	59.0 ± 1.9	78.5 ± 8.8	41.2 ± 9.6
Crystals	46.3 ± 0.4	54.2 ± 1.6	72.2 ± 4.0	71.0 ± 4.2
Fog	78.4 ± 0.4	82.5 ± 0.6	88.7 ± 0.7	88.7 ± 0.7
Gauss. Blur	60.4 ± 2.1	79.4 ± 0.8	82.3 ± 0.6	82.3 ± 0.8
Grass	5.8 ± 0.2	23.6 ± 9.9	81.2 ± 1.1	81.3 ± 1.2
Imp. Noise	76.9 ± 0.9	86.7 ± 0.8	87.9 ± 0.8	86.5 ± 1.1
Sky	4.1 ± 0.5	34.9 ± 4.7	74.9 ± 3.3	74.9 ± 2.9
Stripe	16.3 ± 1.1	68.3 ± 1.4	72.8 ± 1.8	72.0 ± 1.6
Avg High	0.0 ± 0.0	52.7 ± 2.4	73.7 ± 1.7	36.1 ± 6.8
Avg Low	41.2 ± 0.3	61.4 ± 2.3	80.0 ± 1.3	79.5 ± 1.4
Avg All	28.8 ± 0.2	58.7 ± 2.2	78.1 ± 0.6	66.5 ± 2.9

Table 7: 5-shot accuracy across shifts: training different layers of a CNN

	Conv1	Conv2	Conv3	FC1	FC2	FC1 + FC2
H1	0.0 ± 0.0	0.0 ± 0.1	24.3 ± 6.5	82.8 ± 8.2	96.4 ± 1	89.4 ± 2.7
H2	0.0 ± 0.0	0.0 ± 0.1	20.0 ± 2.9	77.7 ± 8.4	91.3 ± 1.5	82.6 ± 3.4
H3	0.0 ± 0.0	0.1 ± 0.1	14.5 ± 3.4	80.6 ± 6.1	96.0 ± 0.6	86.5 ± 3.1
Crystals	77.2 ± 1.2	57.4 ± 0.8	51.5 ± 0.1	49.2 ± 0.8	46.9 ± 1.3	48.9 ± 1.2
Fog	89.7 ± 0.1	86.7 ± 0.6	84.0 ± 0.2	81.4 ± 0.4	81.3 ± 0.4	81.6 ± 0.5
Gauss. Blur	84.4 ± 0.8	82.1 ± 0.9	79.7 ± 0.4	77.2 ± 0.5	75.7 ± 0.8	77.0 ± 0.2
Grass	82.8 ± 1.2	33.1 ± 8.0	12.0 ± 0.5	8.8 ± 0.4	7.6 ± 0.2	9.2 ± 0.5
Imp. Noise	88.8 ± 0.3	86.2 ± 0.2	84.4 ± 0.6	83.1 ± 0.3	80.9 ± 0.7	82.2 ± 0.3
Sky	80.5 ± 0.8	62.7 ± 1.3	44.9 ± 1.4	28.2 ± 1.0	19.3 ± 0.7	28.9 ± 0.3
Stripe	75.8 ± 0.6	70.6 ± 1.5	70.6 ± 0.5	60.6 ± 2.4	47.9 ± 1.0	60.8 ± 2.9
Avg High	0.0 ± 0.0	0.0 ± 0.1	19.6 ± 2.1	80.4 ± 5.6	94.5 ± 0.9	86.2 ± 0.3
Avg Low	82.7 ± 0.6	68.4 ± 1.1	61.0 ± 0.3	55.5 ± 0.5	51.4 ± 0.4	55.5 ± 0.4
Avg All	57.9 ± 0.4	47.9 ± 0.8	48.6 ± 0.8	63.0 ± 1.6	64.3 ± 0.5	64.7 ± 0.3

Table 8: 5-shot accuracy across shifts: comparison of different responses. Zero-shot is the accuracy before any updates are performed.

	Zero-shot	R_0	R_1	R_2
H1	0.0 ± 0.0	76.1 ± 2.7	89.4 ± 2.7	82.0 ± 8.3
H2	0.0 ± 0.0	75.7 ± 4.8	82.6 ± 3.4	76.8 ± 7.1
H3	0.0 ± 0.0	73.3 ± 6.0	86.5 ± 3.1	78.2 ± 9.5
Crystals	46.3 ± 0.4	59.1 ± 1.4	77.2 ± 1.2	75.8 ± 3.0
Fog	78.4 ± 0.4	86.1 ± 0.4	89.7 ± 0.1	89.7 ± 0.1
Gauss. Blur	60.4 ± 2.1	82.0 ± 1.0	84.4 ± 0.8	84.4 ± 0.7
Grass	5.8 ± 0.2	49.5 ± 8.3	82.8 ± 1.2	82.9 ± 1.3
Imp. Noise	76.9 ± 0.9	87.3 ± 0.5	88.8 ± 0.3	87.1 ± 0.6
Sky	4.1 ± 0.5	55.7 ± 4.2	80.5 ± 0.8	80.4 ± 1.0
Stripe	16.3 ± 1.1	75.1 ± 1.0	75.8 ± 0.6	76.7 ± 0.9
Avg High	0.0 ± 0.0	75.0 ± 2.7	86.2 ± 0.3	79.0 ± 4.9
Avg Low	41.2 ± 0.3	70.7 ± 2.0	82.7 ± 0.6	82.4 ± 0.7
Avg All	28.8 ± 0.2	72.0 ± 1.2	83.8 ± 0.4	81.4 ± 1.6

Table 9: 10-shot accuracy across shifts: training different layers of a CNN

	Conv1	Conv2	Conv3	FC1	FC2	FC1 + FC2
H1	0.0 ± 0.0	0.2 ± 0.3	47.5 ± 6.1	91.0 ± 1.9	96.1 ± 0.8	89.4 ± 5.5
H2	0.0 ± 0.0	0.0 ± 0.1	44.8 ± 6.4	88.3 ± 1.9	92.2 ± 1.3	86.3 ± 7.0
H3	0.0 ± 0.0	0.0 ± 0.1	42.0 ± 14.0	89.4 ± 5.4	94.9 ± 1.2	92.1 ± 1.9
Crystals	80.9 ± 0.3	60.7 ± 2.0	53.7 ± 0.6	51.8 ± 1.0	48.2 ± 1.1	51.8 ± 0.9
Fog	90.2 ± 0.2	87.9 ± 0.4	85.2 ± 0.2	83.4 ± 0.3	82.4 ± 0.8	83.3 ± 0.1
Gauss. Blur	85.9 ± 0.6	84.3 ± 0.1	81.9 ± 0.4	80.5 ± 0.7	78.7 ± 1.0	80.1 ± 0.5
Grass	84.2 ± 0.7	52.9 ± 4.2	15.4 ± 0.4	10.6 ± 0.6	7.7 ± 0.2	10.9 ± 0.7
Imp. Noise	89.0 ± 0.1	87.0 ± 0.5	84.8 ± 0.3	84.1 ± 0.2	82.3 ± 0.1	84.0 ± 0.3
Sky	83.2 ± 1.1	69.9 ± 1.1	52.7 ± 0.3	37.4 ± 0.8	26.3 ± 0.2	37.7 ± 0.6
Stripe	80.8 ± 0.5	76.3 ± 2.7	75.8 ± 0.9	69.0 ± 0.8	56.8 ± 0.9	69.2 ± 0.9
Avg High	0.0 ± 0.0	0.1 ± 0.1	44.7 ± 4.8	89.6 ± 1.2	94.4 ± 0.3	89.3 ± 3.0
Avg Low	84.9 ± 0.3	74.2 ± 1.1	64.2 ± 0.1	59.5 ± 0.5	54.6 ± 0.2	59.6 ± 0.1
Avg All	59.4 ± 0.2	51.9 ± 0.8	58.4 ± 1.5	68.6 ± 0.7	66.6 ± 0.1	68.5 ± 1.0

Table 10: 10-shot accuracy across shifts: comparison of different responses. Zero-shot is the accuracy before any updates are performed.

	Zero-shot	R_0	R_1	R_2
H1	0.0 ± 0.0	84.9 ± 1.7	89.4 ± 5.5	90.8 ± 2.3
H2	0.0 ± 0.0	79.4 ± 1.4	86.3 ± 7.0	88.1 ± 1.8
H3	0.0 ± 0.0	86.8 ± 2.3	92.1 ± 1.9	89.6 ± 5.1
Crystals	46.3 ± 0.4	61.8 ± 1.1	80.9 ± 0.3	79.2 ± 1.5
Fog	78.4 ± 0.4	87.0 ± 0.4	90.2 ± 0.2	90.2 ± 0.2
Gauss. Blur	60.4 ± 2.1	84.1 ± 0.5	85.9 ± 0.6	85.8 ± 0.7
Grass	5.8 ± 0.2	69.3 ± 2.1	84.2 ± 0.7	83.8 ± 0.2
Imp. Noise	76.9 ± 0.9	87.7 ± 0.3	89.0 ± 0.1	87.5 ± 0.3
Sky	4.1 ± 0.5	67.2 ± 1.6	83.2 ± 1.1	83.1 ± 1.2
Stripe	16.3 ± 1.1	78.6 ± 0.9	80.8 ± 0.5	82.0 ± 0.6
Avg High	0.0 ± 0.0	83.7 ± 0.1	89.3 ± 3.0	89.5 ± 1.0
Avg Low	41.2 ± 0.3	76.5 ± 0.5	84.9 ± 0.3	84.5 ± 0.1
Avg All	28.8 ± 0.2	78.7 ± 0.4	86.2 ± 1.0	86.0 ± 0.3

Table 11: 20-shot accuracy across shifts: training different layers of a CNN

	Conv1	Conv2	Conv3	FC1	FC2	FC1 + FC2
H1	0.0 ± 0.0	1.9 ± 3.3	63.8 ± 8.2	95.7 ± 0.8	97.2 ± 0.7	94.9 ± 0.9
H2	0.0 ± 0.0	2.9 ± 2.2	65.7 ± 5.3	89.9 ± 1.9	94.5 ± 0.4	90.0 ± 1.3
H3	0.0 ± 0.0	2.9 ± 1.9	68.9 ± 5.1	93.4 ± 2.3	95.8 ± 1.3	90.4 ± 1.2
Crystals	83.4 ± 0.1	65.9 ± 1.8	55.8 ± 0.3	53.8 ± 0.4	49.6 ± 0.9	54.1 ± 0.5
Fog	90.5 ± 0.2	88.6 ± 0.2	86.1 ± 0.4	84.5 ± 0.3	83.8 ± 0.3	84.5 ± 0.1
Gauss. Blur	86.8 ± 0.2	85.6 ± 0.3	83.2 ± 0.3	82.1 ± 0.4	80.6 ± 0.7	82.5 ± 0.6
Grass	85.6 ± 0.8	64.7 ± 2.9	21.1 ± 1.7	13.4 ± 0.9	8.2 ± 0.3	13.3 ± 1.0
Imp. Noise	89.2 ± 0.2	87.5 ± 0.2	85.3 ± 0.5	85.0 ± 0.3	82.8 ± 0.5	85.1 ± 0.3
Sky	84.6 ± 0.7	75.5 ± 0.4	59.5 ± 0.6	45.4 ± 0.9	31.9 ± 0.5	46.7 ± 0.5
Stripe	84.2 ± 0.2	80.0 ± 1.6	79.1 ± 0.9	74.4 ± 0.2	62.3 ± 1.0	75.0 ± 0.2
Avg High	0.0 ± 0.0	2.6 ± 1.6	66.1 ± 1.2	93.0 ± 1.3	95.8 ± 0.3	91.8 ± 0.8
Avg Low	86.3 ± 0.2	78.3 ± 0.8	67.2 ± 0.3	62.7 ± 0.2	57.0 ± 0.4	63.0 ± 0.2
Avg All	60.4 ± 0.1	55.5 ± 0.8	66.8 ± 0.3	71.8 ± 0.5	68.7 ± 0.4	71.7 ± 0.4

Table 12: 20-shot accuracy across shifts: comparison of different responses. Zero-shot is the accuracy before any updates are performed.

	Zero-shot	R_0	R_1	R_2
H1	0.0 ± 0.0	89.3 ± 1.7	94.9 ± 0.9	95.6 ± 1.0
H2	0.0 ± 0.0	85.4 ± 2.8	90.0 ± 1.3	90.1 ± 2.5
H3	0.0 ± 0.0	90.2 ± 0.5	90.4 ± 1.2	93.3 ± 2.4
Crystals	46.3 ± 0.4	66.6 ± 1.5	83.4 ± 0.1	81.1 ± 0.4
Fog	78.4 ± 0.4	87.5 ± 0.1	90.5 ± 0.2	90.5 ± 0.2
Gauss. Blur	60.4 ± 2.1	85.0 ± 0.7	86.8 ± 0.2	86.5 ± 0.8
Grass	5.8 ± 0.2	76.7 ± 0.9	85.6 ± 0.8	84.8 ± 0.4
Imp. Noise	76.9 ± 0.9	87.9 ± 0.4	89.2 ± 0.2	87.8 ± 0.2
Sky	4.1 ± 0.5	73.3 ± 0.7	84.6 ± 0.7	84.7 ± 0.6
Stripe	16.3 ± 1.1	81.2 ± 0.6	84.2 ± 0.2	84.6 ± 0.3
Avg High	0.0 ± 0.0	88.3 ± 1.2	91.8 ± 0.8	93.0 ± 1.6
Avg Low	41.2 ± 0.3	79.7 ± 0.2	86.3 ± 0.2	85.7 ± 0.1
Avg All	28.8 ± 0.2	82.3 ± 0.2	88.0 ± 0.3	87.9 ± 0.5

Table 13: 50-shot accuracy across shifts: training different layers of a CNN

	Conv1	Conv2	Conv3	FC1	FC2	FC1 + FC2
H1	0.0 ± 0.0	18.8 ± 16.9	83.2 ± 5.2	96.9 ± 0.1	97.2 ± 0.5	95.4 ± 0.4
H2	0.0 ± 0.0	15.7 ± 14.9	84.0 ± 1.9	94.5 ± 0.7	95.3 ± 0.8	93.0 ± 0.9
H3	0.0 ± 0.0	17.3 ± 18.6	84.6 ± 5.6	95.5 ± 1.9	97.3 ± 0.9	96.1 ± 0.7
Crystals	85.2 ± 0.2	72.5 ± 1.6	58.9 ± 0.1	57.3 ± 0.4	51.8 ± 0.4	57.6 ± 0.3
Fog	90.6 ± 0.2	89.3 ± 0.1	87.2 ± 0.2	85.9 ± 0.1	85.0 ± 0.3	85.9 ± 0.3
Gauss. Blur	87.6 ± 0.3	87.1 ± 0.4	85.0 ± 0.4	84.2 ± 0.3	82.9 ± 0.5	84.0 ± 0.2
Grass	86.9 ± 0.6	74.0 ± 2.2	29.9 ± 1.7	19.1 ± 0.6	9.7 ± 0.2	19.0 ± 1.0
Imp. Noise	89.3 ± 0.1	87.9 ± 0.1	85.8 ± 0.3	85.5 ± 0.1	83.8 ± 0.1	85.7 ± 0.1
Sky	86.6 ± 0.6	80.4 ± 0.2	66.4 ± 0.4	54.6 ± 1.1	38.2 ± 0.8	56.4 ± 0.4
Stripe	86.9 ± 0.3	84.5 ± 0.5	83.0 ± 0.6	79.6 ± 0.6	67.7 ± 0.7	79.1 ± 0.5
Avg High	0.0 ± 0.0	17.2 ± 15.3	83.9 ± 3.6	95.6 ± 0.6	96.6 ± 0.7	94.8 ± 0.2
Avg Low	87.6 ± 0.1	82.3 ± 0.5	70.9 ± 0.3	66.6 ± 0.2	59.9 ± 0.3	66.8 ± 0.1
Avg All	61.3 ± 0.1	62.8 ± 4.5	74.8 ± 0.9	75.3 ± 0.3	70.9 ± 0.3	75.2 ± 0.0

Table 14: 50-shot accuracy across shifts: comparison of different responses. Zero-shot is the accuracy before any updates are performed.

	Zero-shot	R_0	R_1	R_2
H1	0.0 ± 0.0	93.2 ± 0.7	95.4 ± 0.4	96.9 ± 0.2
H2	0.0 ± 0.0	89.9 ± 0.9	93.0 ± 0.9	94.5 ± 0.6
H3	0.0 ± 0.0	93.9 ± 1.5	96.1 ± 0.7	95.5 ± 2.0
Crystals	46.3 ± 0.4	73.0 ± 0.9	85.2 ± 0.2	82.9 ± 0.7
Fog	78.4 ± 0.4	88.0 ± 0.3	90.6 ± 0.2	90.5 ± 0.4
Gauss. Blur	60.4 ± 2.1	86.0 ± 0.6	87.6 ± 0.3	86.6 ± 0.7
Grass	5.8 ± 0.2	81.2 ± 0.2	86.9 ± 0.6	85.6 ± 0.3
Imp. Noise	76.9 ± 0.9	88.1 ± 0.1	89.3 ± 0.1	88.1 ± 0.2
Sky	4.1 ± 0.5	79.0 ± 0.4	86.6 ± 0.6	86.6 ± 0.5
Stripe	16.3 ± 1.1	84.1 ± 0.9	86.9 ± 0.3	86.8 ± 0.2
Avg High	0.0 ± 0.0	92.3 ± 0.5	94.8 ± 0.2	95.6 ± 0.5
Avg Low	41.2 ± 0.3	82.8 ± 0.2	87.6 ± 0.1	86.7 ± 0.1
Avg All	28.8 ± 0.2	85.6 ± 0.2	89.8 ± 0.0	89.4 ± 0.1

Table 15: 2000-shot (all data) accuracy across shifts: training different layers of a CNN

	Conv1	Conv2	Conv3	FC1	FC2	FC1 + FC2
H1	0.0 ± 0.0	63.7 ± 15.0	97.6 ± 0.1	98.7 ± 0.1	98.6 ± 0.2	98.8 ± 0.2
H2	0.3 ± 0.5	66.1 ± 13.4	96.9 ± 0.3	98.1 ± 0.2	97.8 ± 0.3	98.2 ± 0.2
H3	0.0 ± 0.0	67.6 ± 8.9	97.5 ± 0.5	98.6 ± 0.4	98.5 ± 0.0	98.8 ± 0.1
Crystals	88.0 ± 0.1	85.7 ± 0.5	71.8 ± 0.3	68.8 ± 0.4	57.5 ± 0.3	68.3 ± 0.3
Fog	91.2 ± 0.2	91.3 ± 0.1	90.4 ± 0.1	90.0 ± 0.1	88.0 ± 0.2	89.2 ± 0.1
Gauss. Blur	90.1 ± 0.1	90.8 ± 0.1	90.2 ± 0.1	90.1 ± 0.2	86.9 ± 0.2	89.4 ± 0.3
Grass	89.1 ± 0.2	85.9 ± 0.5	55.7 ± 1.1	40.7 ± 1.0	17.0 ± 0.3	36.9 ± 1.3
Imp. Noise	89.8 ± 0.1	89.6 ± 0.2	87.5 ± 0.1	87.8 ± 0.2	86.6 ± 0.1	87.7 ± 0.1
Sky	89.1 ± 0.1	88.1 ± 0.3	81.7 ± 0.2	75.0 ± 0.3	49.2 ± 0.3	73.6 ± 0.5
Stripe	90.1 ± 0.1	90.8 ± 0.3	90.0 ± 0.1	89.0 ± 0.1	75.8 ± 1.4	88.1 ± 0.2
Avg High	0.1 ± 0.2	65.8 ± 6.7	97.3 ± 0.2	98.5 ± 0.0	98.3 ± 0.1	98.6 ± 0.1
Avg Low	89.6 ± 0.1	88.9 ± 0.1	81.1 ± 0.1	77.3 ± 0.2	65.9 ± 0.1	76.2 ± 0.2
Avg All	62.8 ± 0.0	82.0 ± 2.0	85.9 ± 0.0	83.7 ± 0.1	75.6 ± 0.1	82.9 ± 0.1

Table 16: 2000-shot (all data) accuracy across shifts: comparison of different responses. Zero-shot is the accuracy before any updates are performed.

	Zero-shot	R_0	R_1	R_2
H1	0.0 ± 0.0	97.9 ± 0.2	98.8 ± 0.2	98.7 ± 0.1
H2	0.0 ± 0.0	97.4 ± 0.3	98.2 ± 0.2	98.2 ± 0.2
H3	0.0 ± 0.0	98.0 ± 0.2	98.8 ± 0.1	98.6 ± 0.3
Crystals	46.3 ± 0.4	87.0 ± 0.1	88.0 ± 0.1	88.3 ± 0.2
Fog	78.4 ± 0.4	90.2 ± 0.1	91.2 ± 0.2	90.6 ± 0.1
Gauss. Blur	60.4 ± 2.1	90.0 ± 0.3	90.1 ± 0.1	90.5 ± 0.2
Grass	5.8 ± 0.2	88.6 ± 0.2	89.1 ± 0.2	89.5 ± 0.3
Imp. Noise	76.9 ± 0.9	88.7 ± 0.2	89.8 ± 0.1	89.5 ± 0.2
Sky	4.1 ± 0.5	88.4 ± 0.1	89.1 ± 0.1	89.5 ± 0.2
Stripe	16.3 ± 1.1	89.9 ± 0.3	90.1 ± 0.1	90.0 ± 0.2
Avg High	0.0 ± 0.0	97.8 ± 0.2	98.6 ± 0.1	98.5 ± 0.0
Avg Low	41.2 ± 0.3	89.0 ± 0.1	89.6 ± 0.1	89.7 ± 0.0
Avg All	28.8 ± 0.2	91.6 ± 0.1	92.3 ± 0.0	92.3 ± 0.0