

Computational Vision

Radiometry and Reflectance

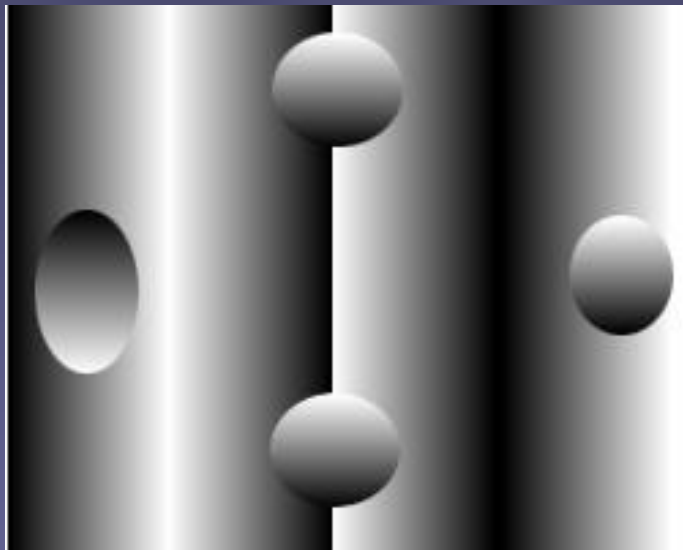
Szeliski: 2.2

Horn: Chapter 2: 22-24, Chapter 10: 202-215

From 2D to 3D

- DEPTH from TWO or MORE IMAGES
 - Stereo
 - Optical Flow -> Factorization Method.
- SHAPE from SINGLE IMAGE CUES
 - SHAPE from SHADING
 - SHAPE from TEXTURE
 - ...

Shading Encodes Shape



From Robot Vision (Horn)

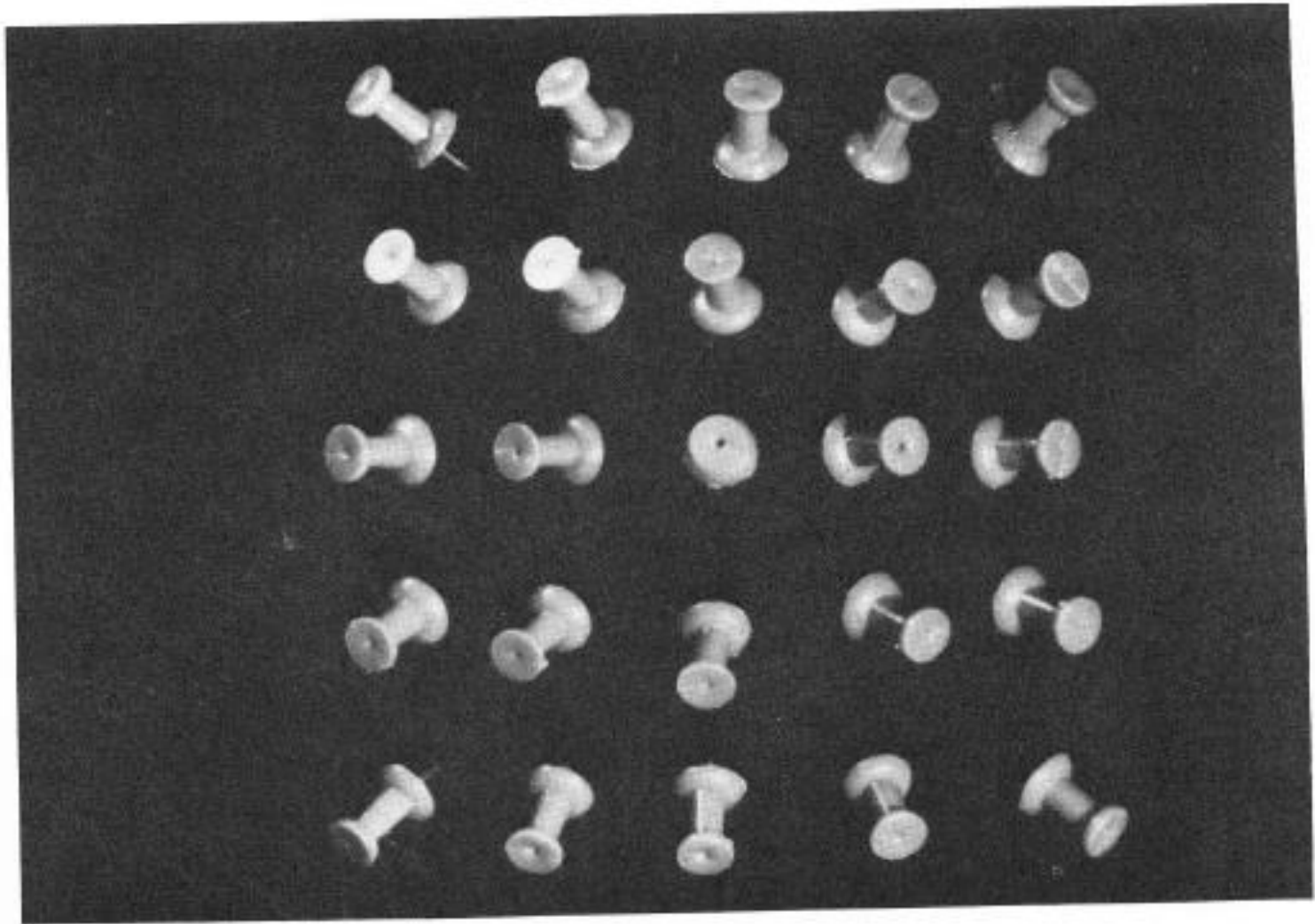
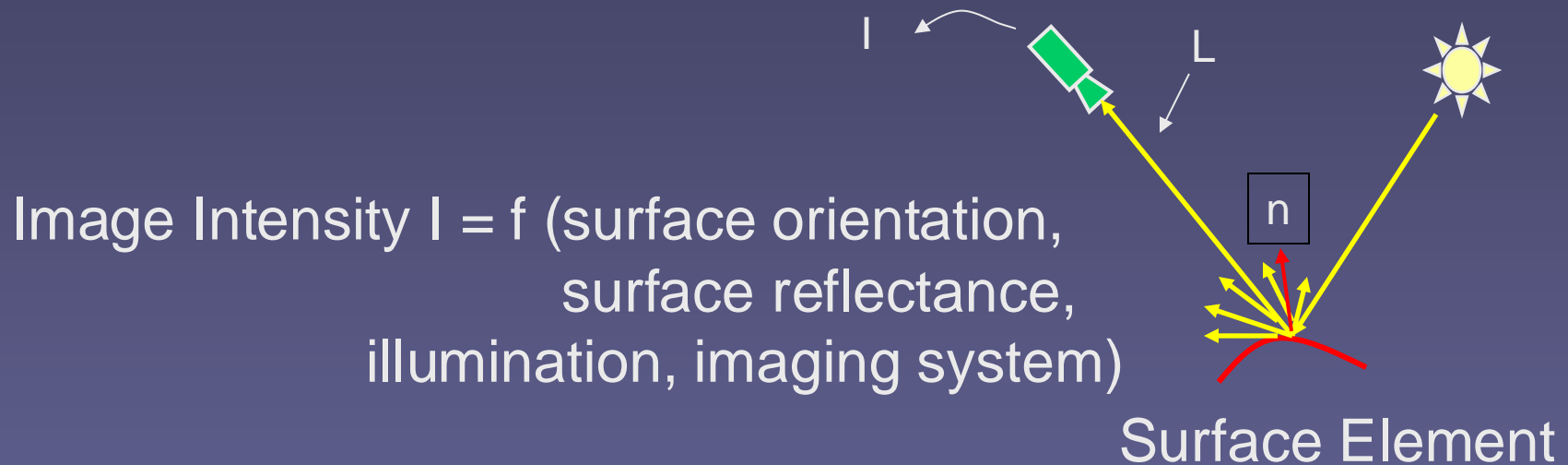


Figure 10-2. The appearance of an object depends greatly on its attitude in space relative to the viewer. Not only does the outline vary, but the brightness pattern within the silhouette changes.

Radiometry and Reflectance



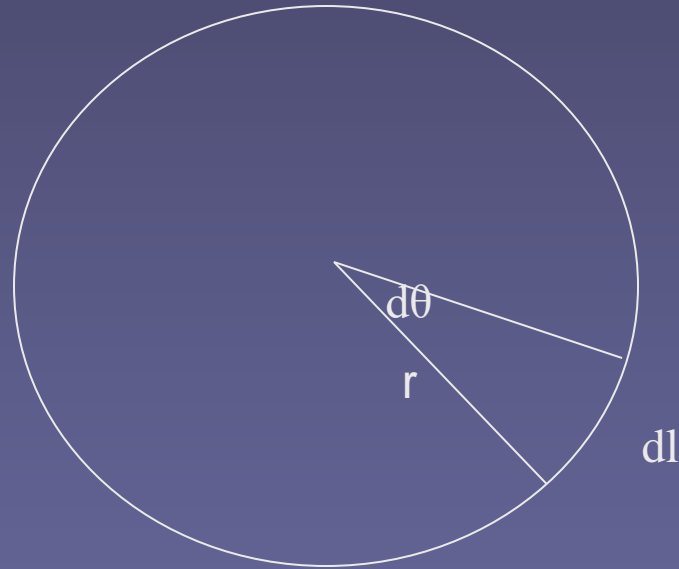
Note: Image Intensity Understanding is an **under-constrained** problem!

Angles in 2D & 3D

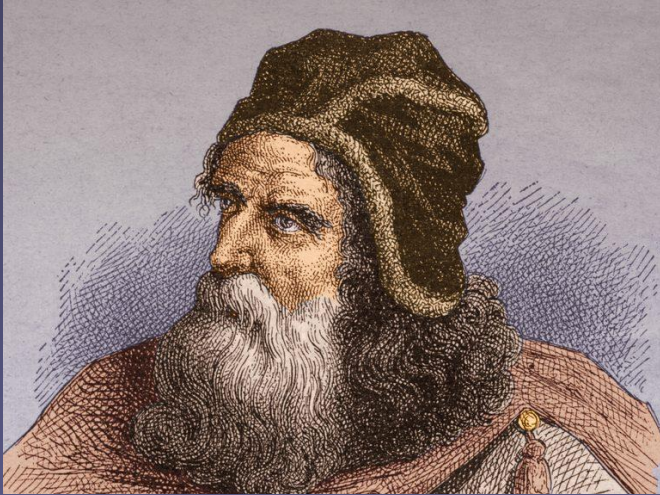
Angle definition on the plane

$$d\theta = dl / r \text{ (radian)}$$

$$d\theta = dl \text{ in unit circle } (r = 1)$$



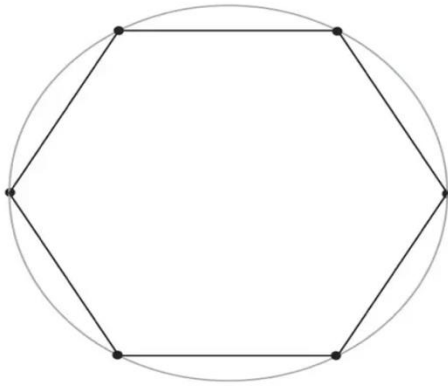
π day : 3/14



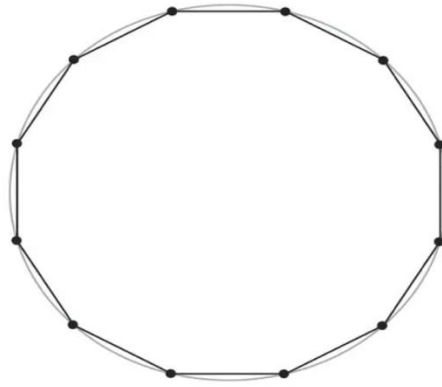
Archimedes (287-211BCE) computation of π

https://www.nytimes.com/article/pi-day-math-geometry-infinity.html?unlocked_article_code=1.ck0.WDnQ.5rBU-rR2oQHR&smid=url-share

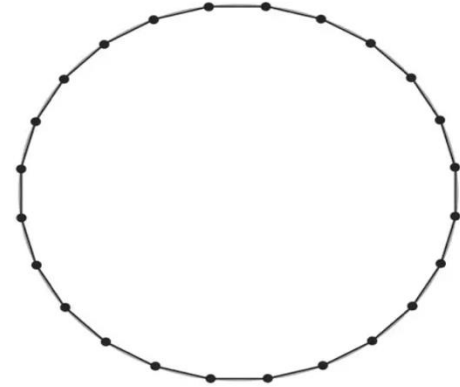
$$3 + \frac{10}{71} < \pi < 3 + \frac{10}{70}$$



6

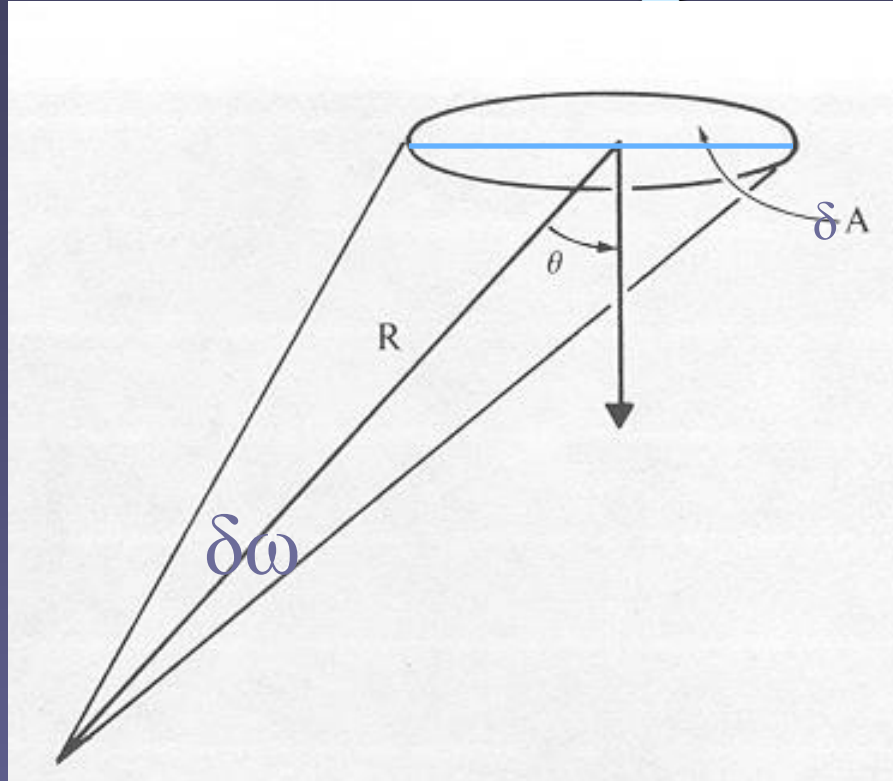


12



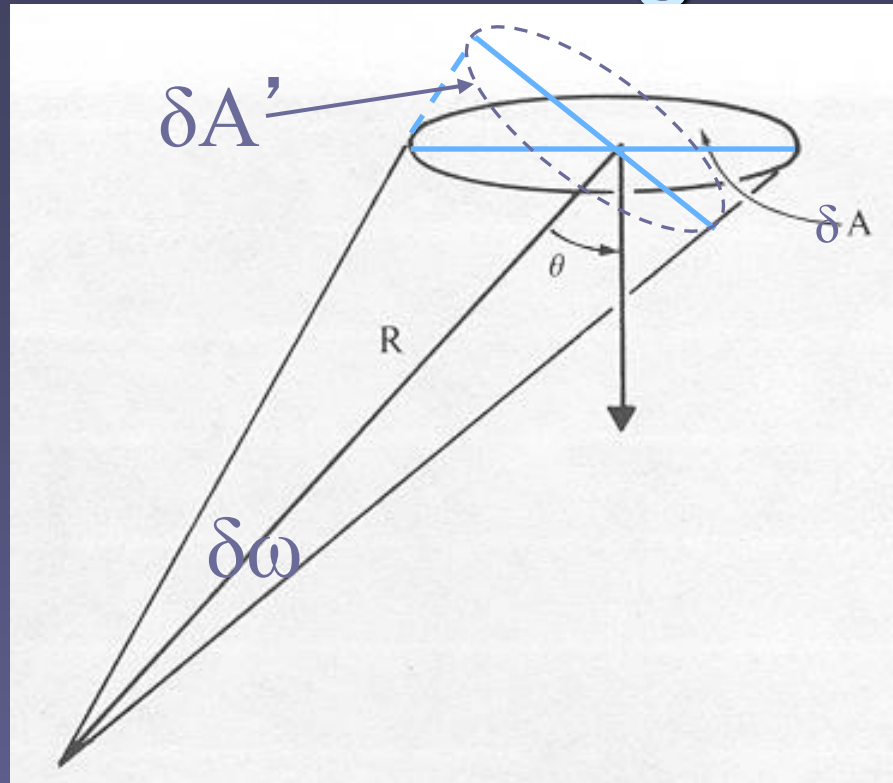
24

Solid Angle



$$\delta\omega = (\delta A \cos\theta) / R^2 \text{ (steradian)}$$

Solid Angle

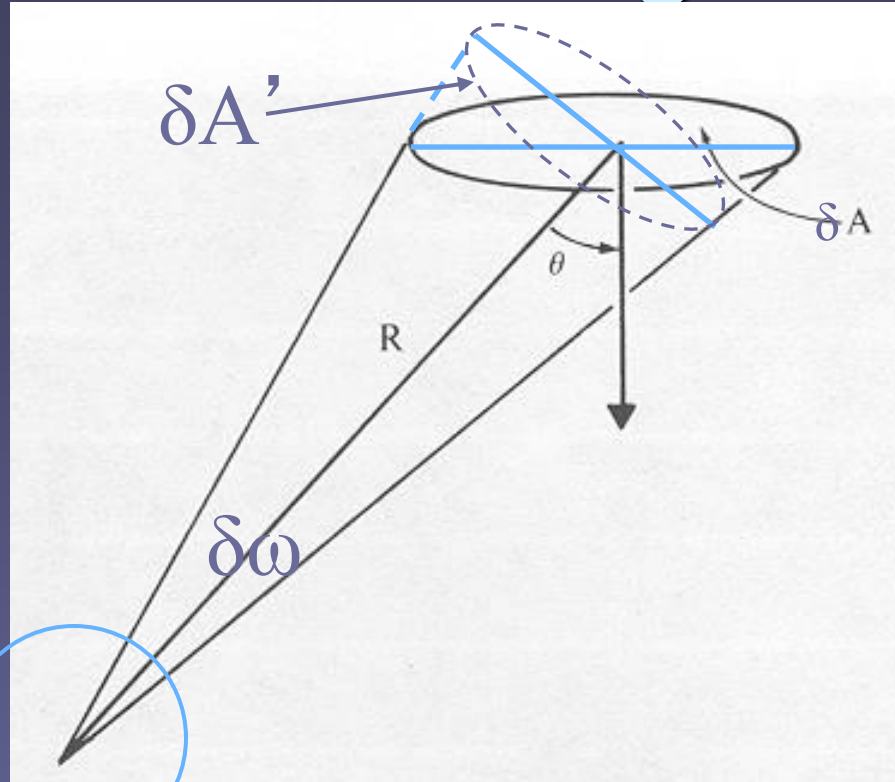


$$\delta \omega = \frac{(\delta A \cos \theta)}{R^2} \text{ (steradian)}$$

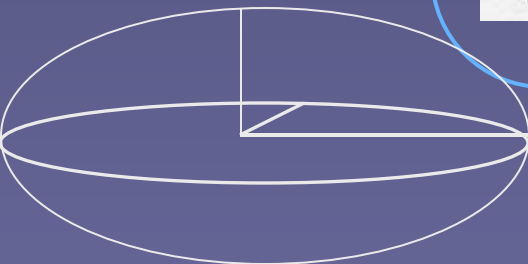
Foreshortened Area

$$\delta \omega = \delta A' / R^2 \text{ (steradian)}$$

Solid Angle



Unit
Sphere

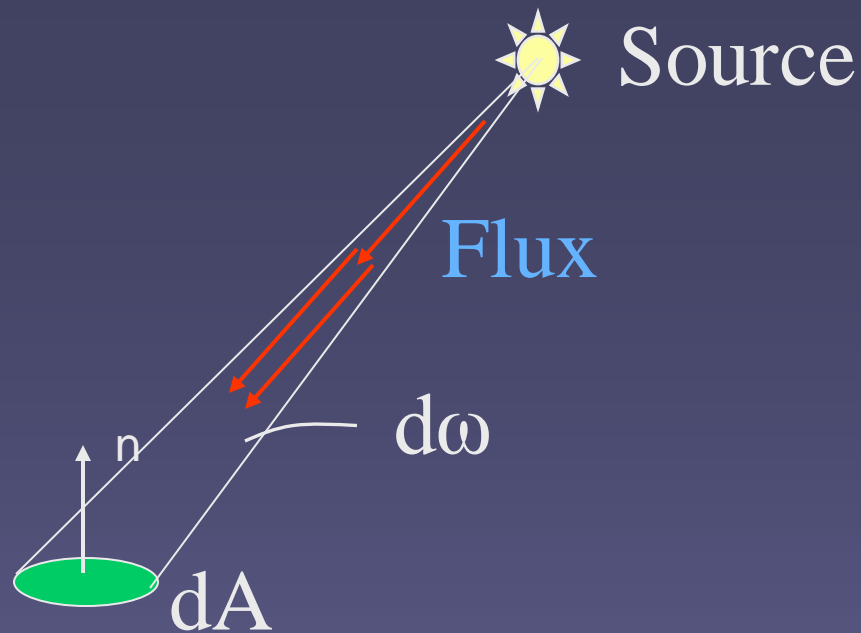


Solid Angle Sustained
by a hemisphere = 2π

$$\delta\omega = \frac{(\delta A \cos\theta)}{R^2} \text{ (steradian)}$$

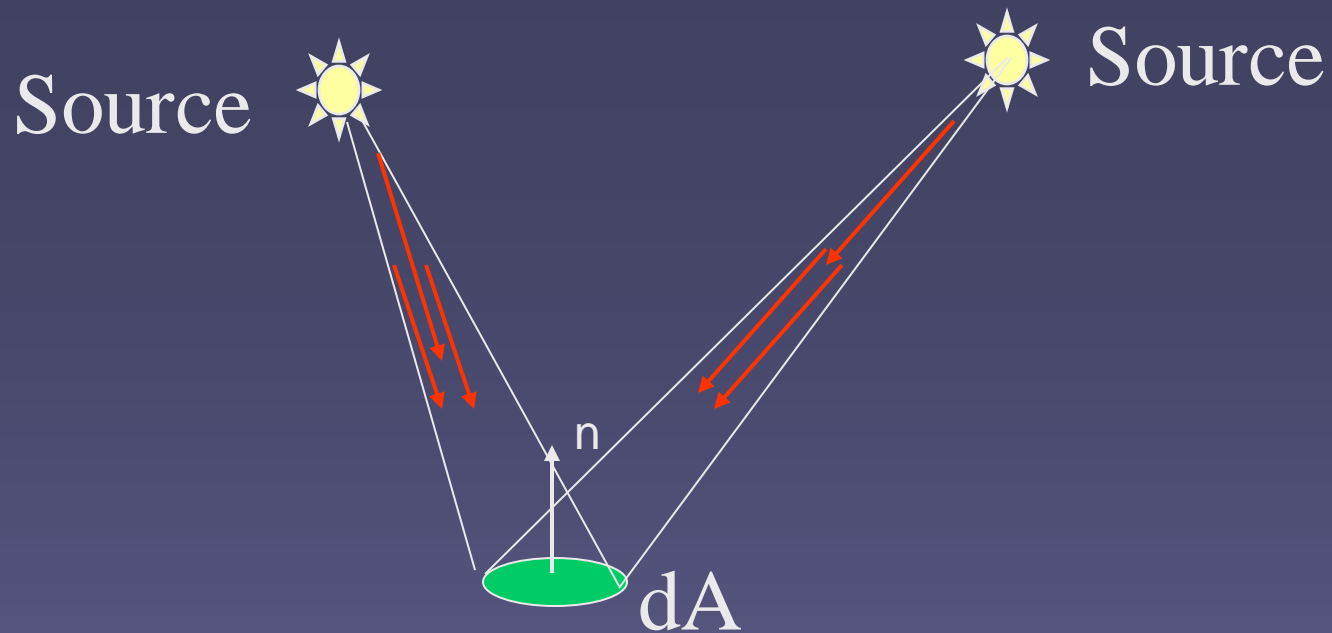
Foreshortened Area

$$\delta\omega = \frac{\delta A'}{R^2} \text{ (steradian)}$$



Radiant Intensity of **Source**: Light flux (power) emitted per unit solid angle:

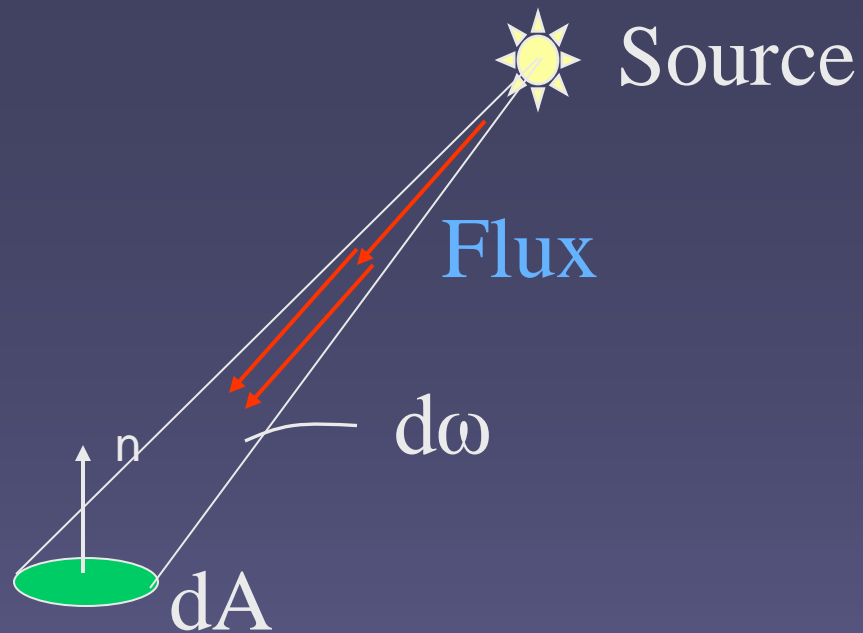
$$J = d\Phi / d\omega \text{ (watts/steradian)}$$



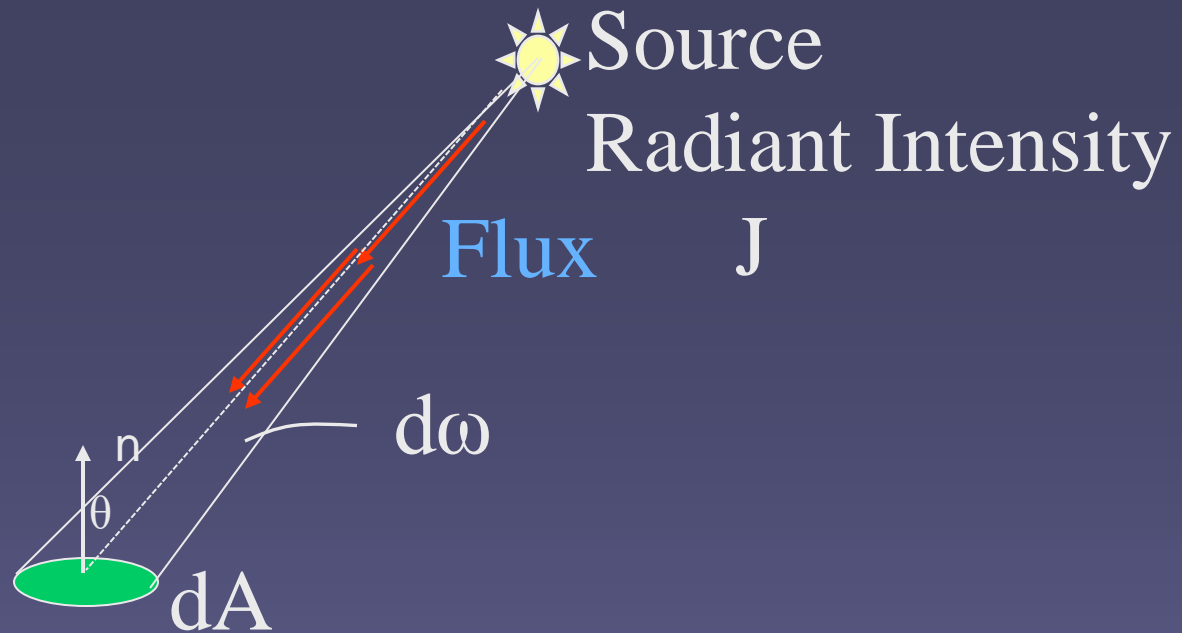
Surface Irradiance: Flux incident per unit surface area:

$$E = d\Phi / dA \text{ (watts/m}^2\text{)}$$

Does not depend on where the light is coming from!
One or more light sources.



Surface Irradiance: Flux incident per unit surface area:
 $E = d\Phi / dA$ (watts/m²)



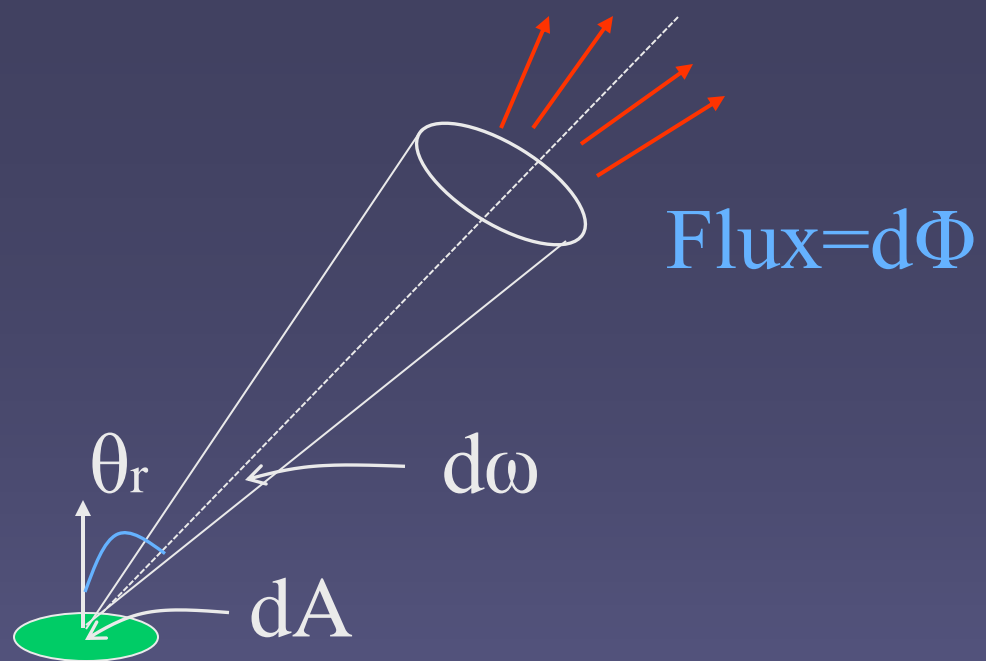
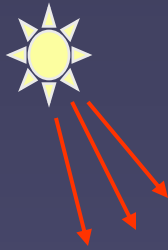
Surface Irradiance as a function of J (one source):

$$E = d\Phi / dA \Rightarrow E = (J d\omega) / dA \Rightarrow E = (J \cos\theta) / R^2$$

E inversely proportional to R

E proportional to $\cos\theta$ (when is E maximized?)

Source

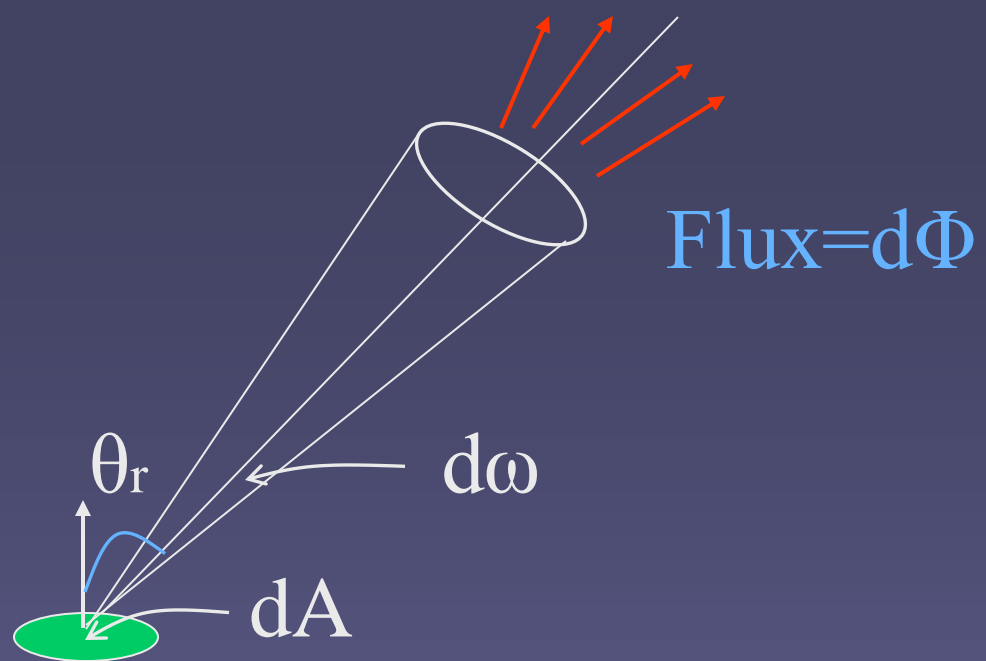
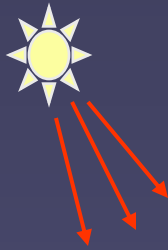


Surface Radiance(Brightness):

Flux emitted per unit foreshortened area, per unit solid angle:

$$L = d\Phi / (dA \cos\theta_r) d\omega \text{ (watts/m}^2 \text{ steradian)}$$

Source



Surface Radiance(Brightness):

Flux emitted per unit foreshortened area,
per unit solid angle:

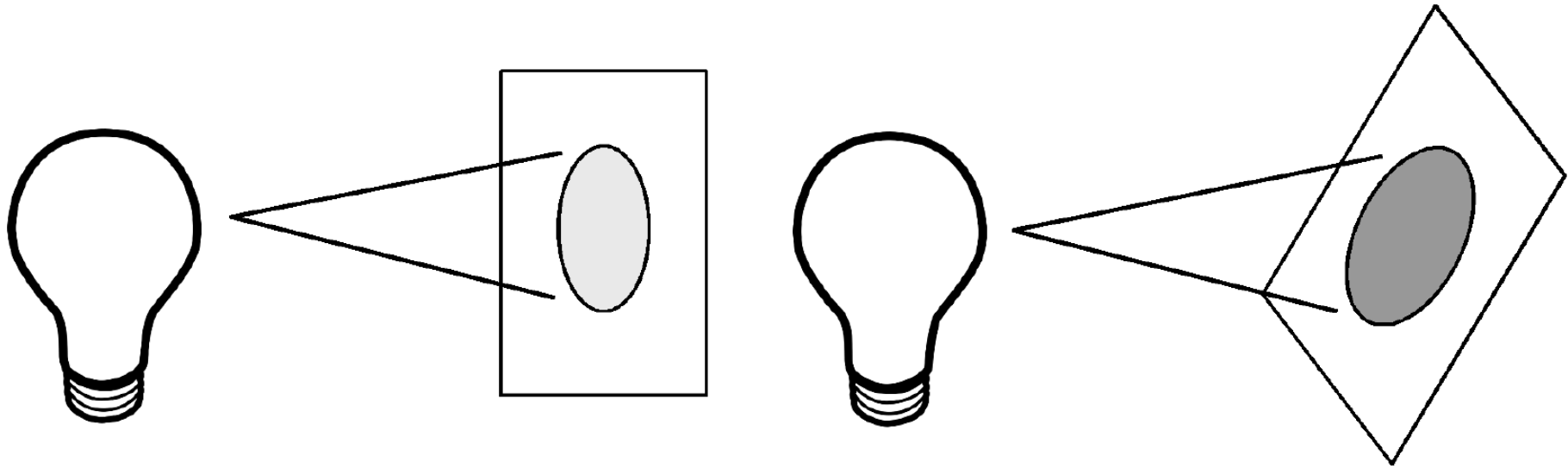
$$L = d\Phi / (dA \cos\theta_r) d\omega \text{ (watts/m}^2 \text{ steradian)}$$

L depends on direction θ_r

Surface can radiate into whole hemisphere

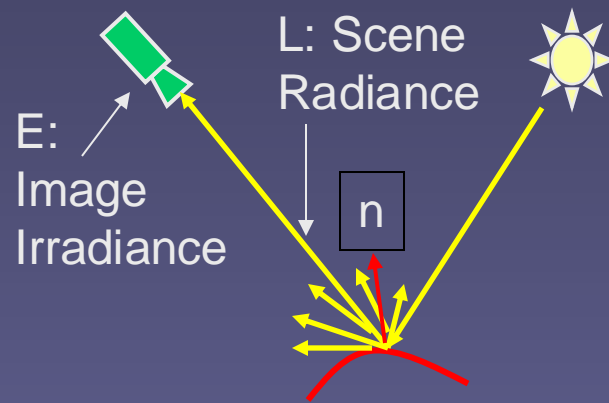
Depends on reflectance properties of surface

Foreshortening



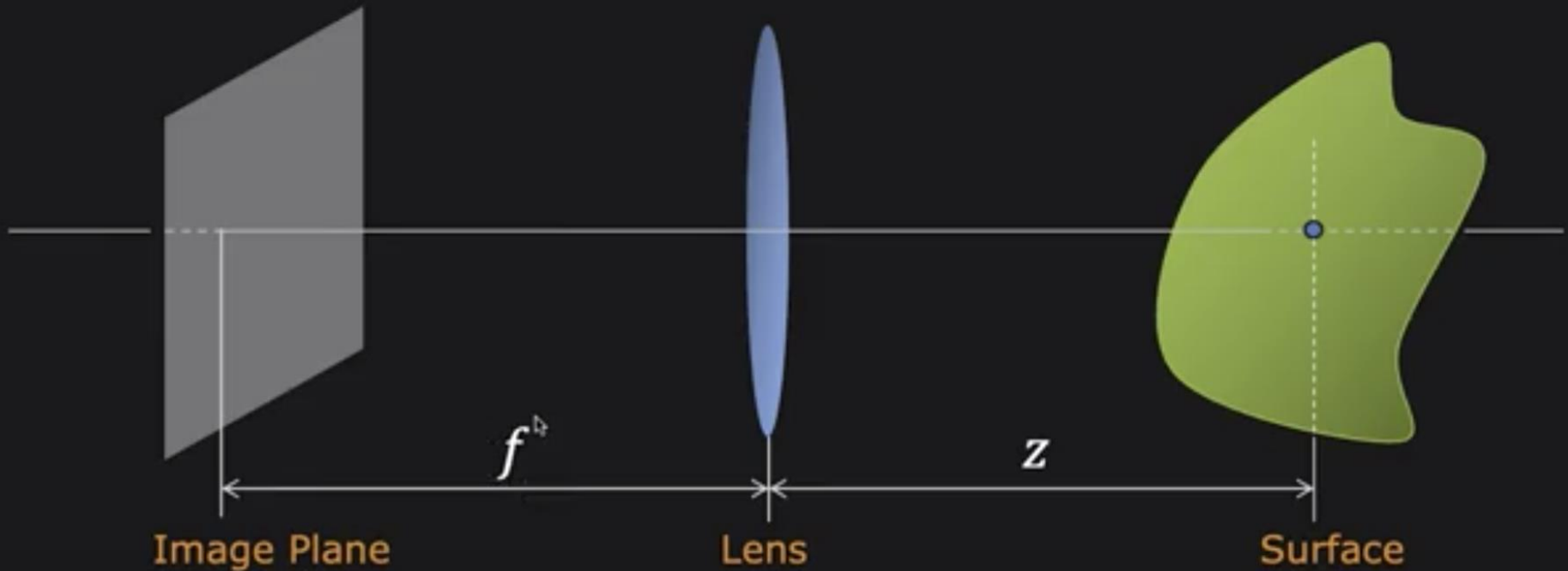
As the surface tilts away from the light source the same light energy is spread over a larger area, making the surface darker

Surface Radiance & Image Irradiance



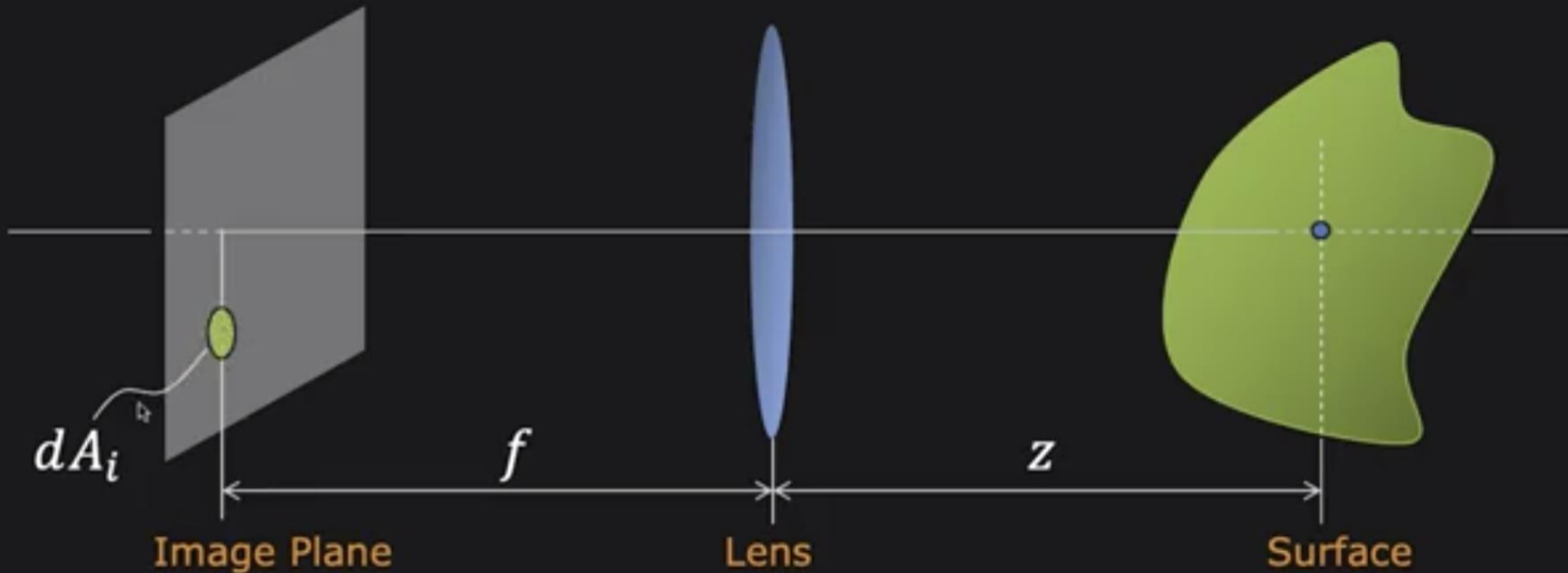
What is the relationship between L & E ?

Surface Radiance & Image Irradiance



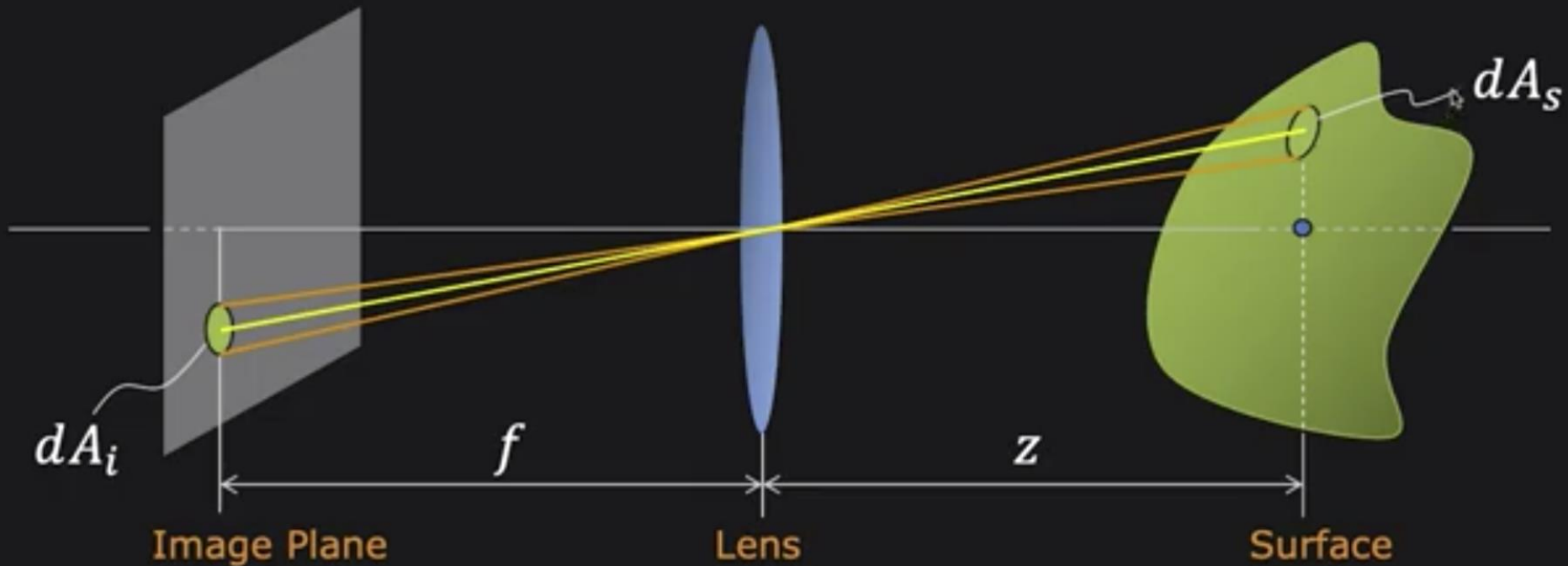
f : effective focal length

Surface Radiance & Image Irradiance



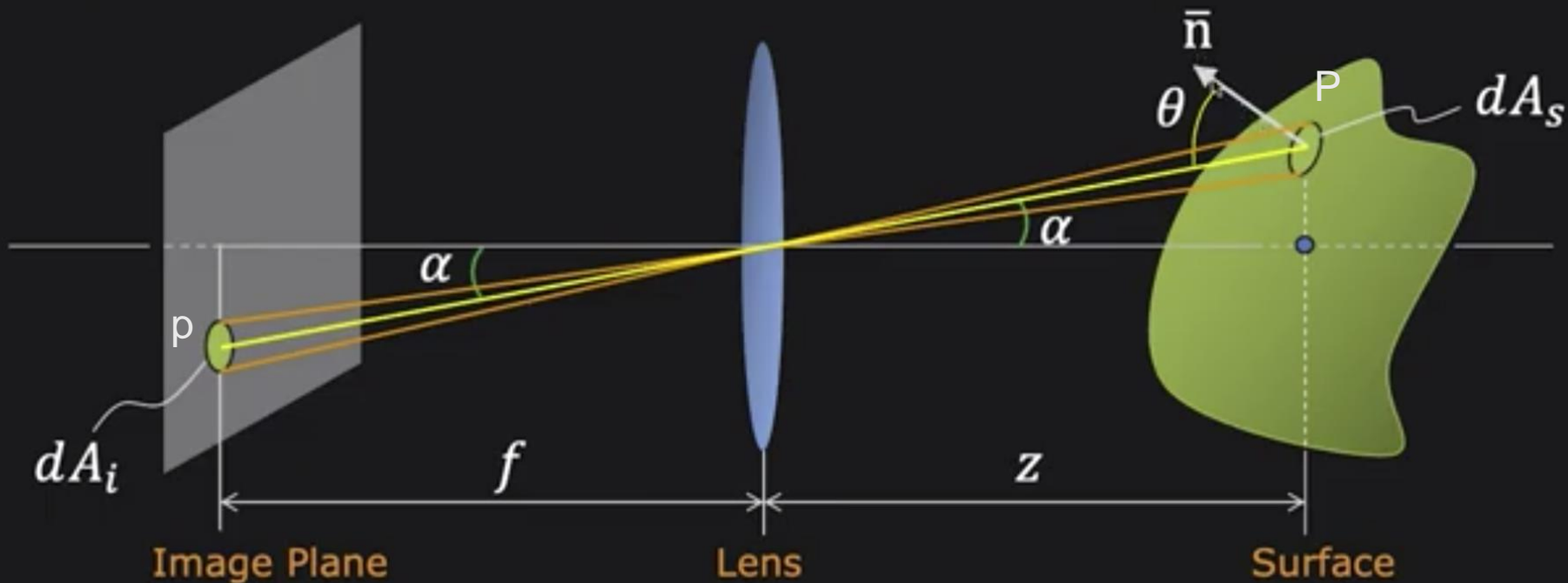
f : effective focal length

Surface Radiance & Image Irradiance



f : effective focal length

Surface Radiance & Image Irradiance



f : effective focal length
 p is perspective projection of scene point P
 dA_i and dA_s are very small surfaces

Surface Radiance & Image Irradiance

Trucco & Verri

L (Watts/m²* steradian)

Here

δI is dA_i

δO is dA_s

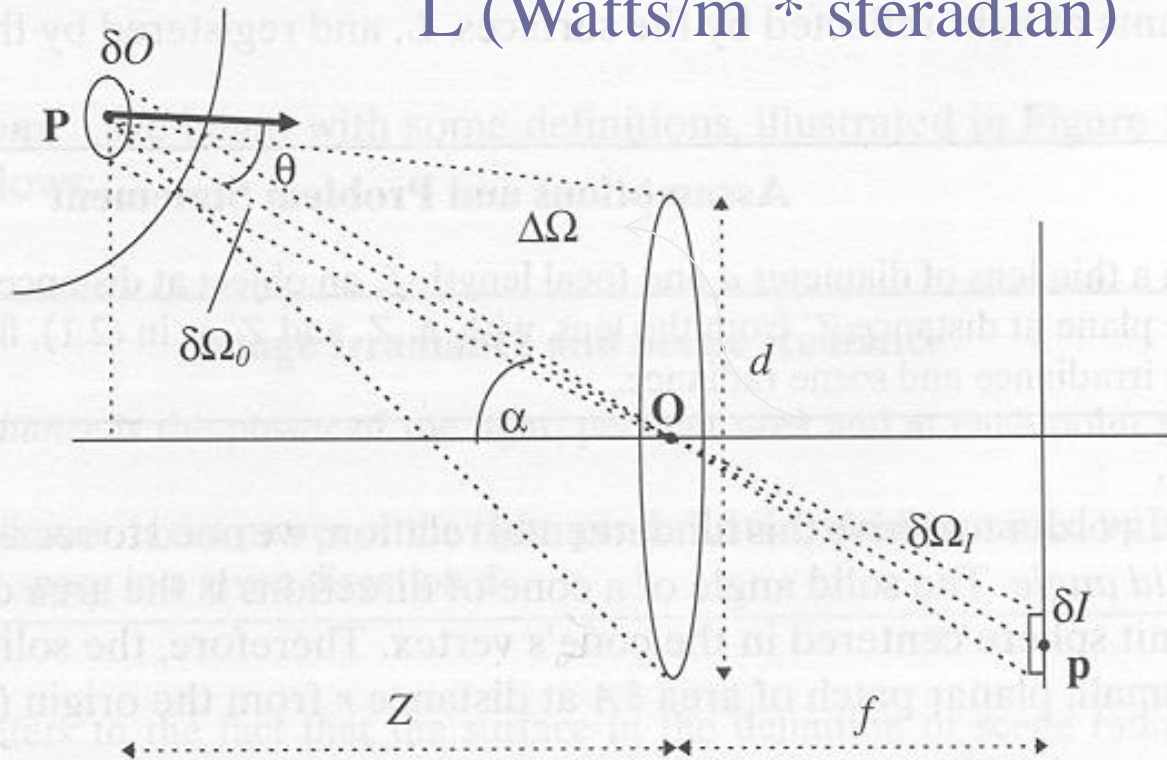


Figure 2.7 Radiometry of the image formation process.

Surface Radiance & Image Irradiance

Trucco & Verri

L (Watts/m²* steradian)

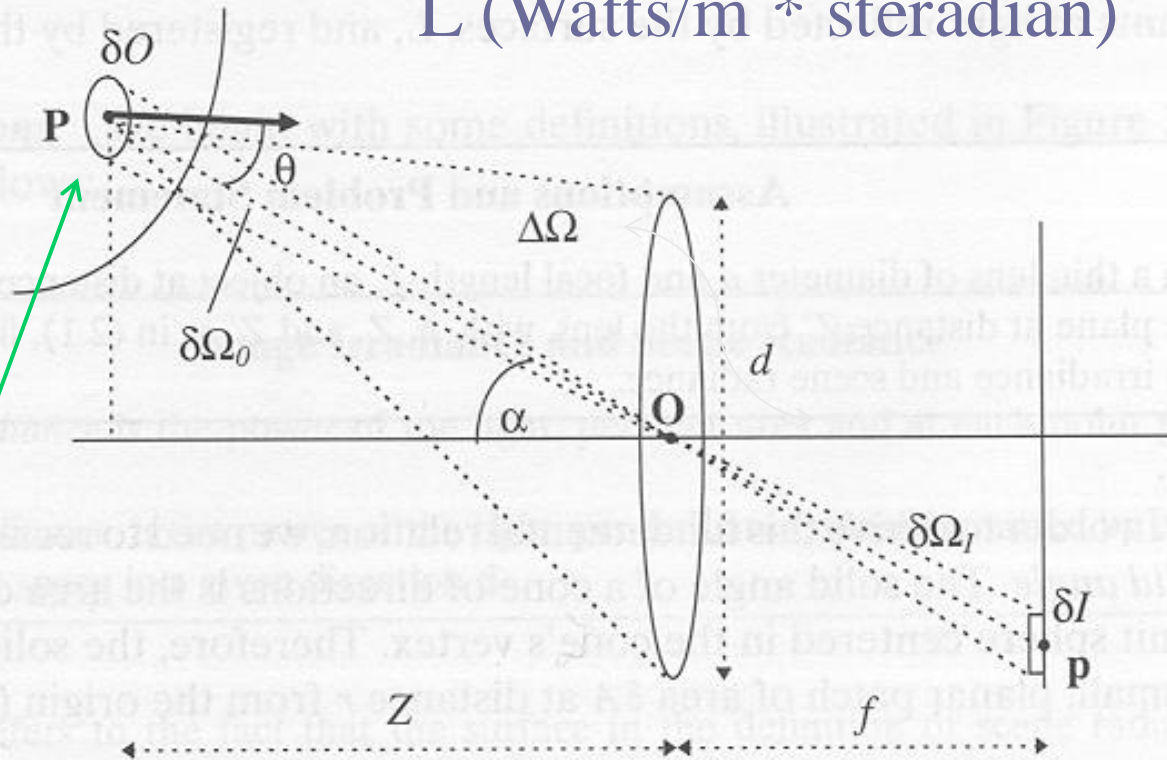


Figure 2.7 Radiometry of the image formation process.

$$L = \delta\Phi / (\delta O * \cos\theta * \Delta\Omega), L \text{ scene radiance at } P$$

Surface Radiance & Image Irradiance

Trucco & Verri

L (Watts/m²* steradian)

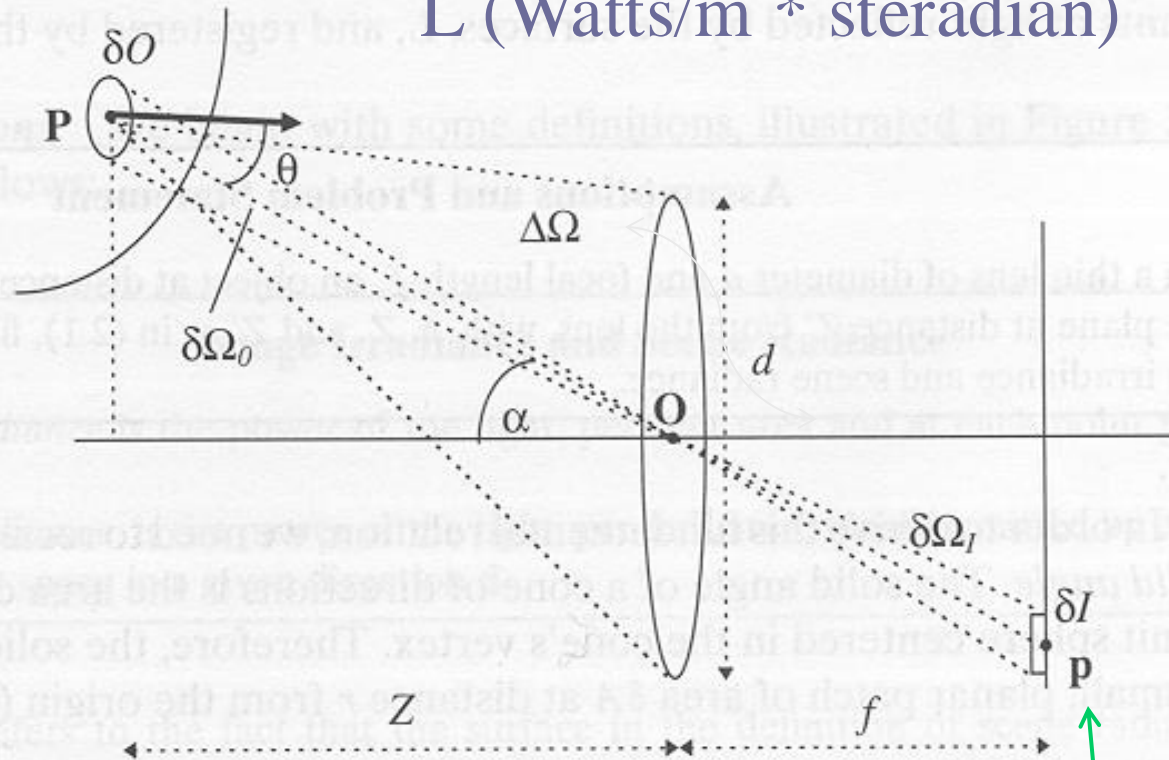


Figure 2.7 Radiometry of the image formation process.

$$E = \delta\Phi / \delta I, \text{ image irradiance at } p$$

Surface Radiance & Image Irradiance

Trucco & Verri

L (Watts/m²* steradian)

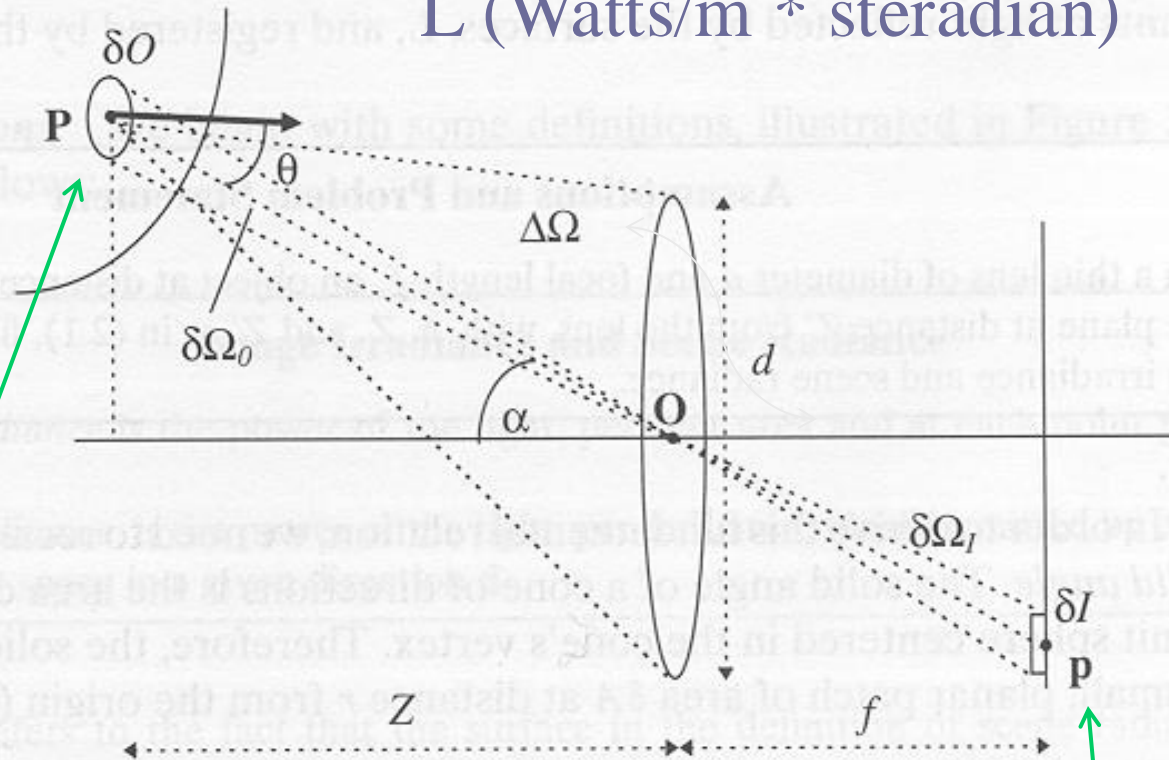


Figure 2.7 Radiometry of the image formation process.

$E = \delta\Phi / \delta I$, image irradiance at p.

$L = \delta\Phi / (\delta O * \cos\theta * \Delta\Omega)$, L scene radiance at P.

Fundamental Equation of Radiometric Image Formation

$$E = L * \Delta\Omega * \cos\theta * \delta O / \delta I$$

Solid angles:

$$\begin{aligned}\Delta\Omega &= \delta A * \cos\alpha / R^2 = (\pi d^2 / 4) * \cos\alpha / (Z / \cos\alpha)^2 \\ &= (\pi / 4) d^2 \cos^3\alpha / Z^2\end{aligned}$$

$$\delta\Omega_o = \delta O * \cos\theta / (Z / \cos\alpha)^2$$

$$\delta\Omega_I = \delta I * \cos\alpha / (f / \cos\alpha)^2$$

$$\delta\Omega_o = \delta\Omega_I \Rightarrow \delta O / \delta I = (\cos\alpha / \cos\theta) * (Z / f)^2$$

Fundamental Equation of Radiometric Image Formation

$$E = L * (\pi/4) d^2 (\cos^3 \alpha / Z^2) * \cos \theta * (\cos \alpha / \cos \theta) * (Z / f)^2$$

Finally:

$$E = L * (\pi / 4) * (d / f)^2 * \cos^4 \alpha$$

Surface Radiance & Image Irradiance

Image irradiance

$$E = L \frac{\pi}{4} \left(\frac{d}{f} \right)^2 \cos^4(\alpha)$$

Brightness falloff

1 / Effective F-number

Scene radiance

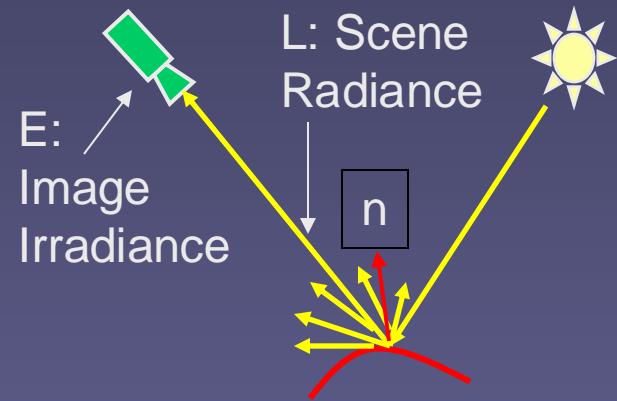


Image irradiance is proportional to scene radiance

$$E \propto L$$

Surface Radiance & Image Irradiance

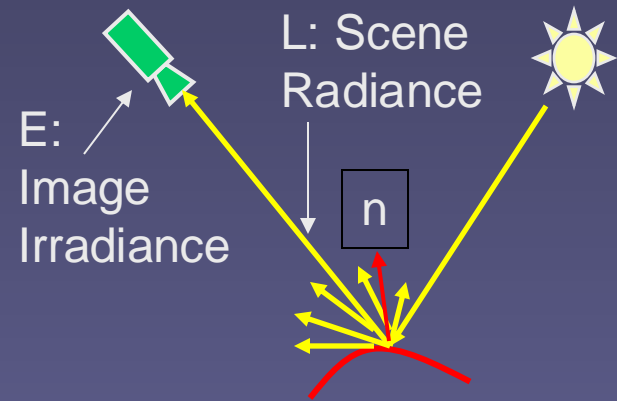
Image irradiance

$$E = L \frac{\pi}{4} \left(\frac{d}{f} \right)^2 \cos^4(\alpha)$$

Brightness falloff

1 / Effective F-number

Scene radiance



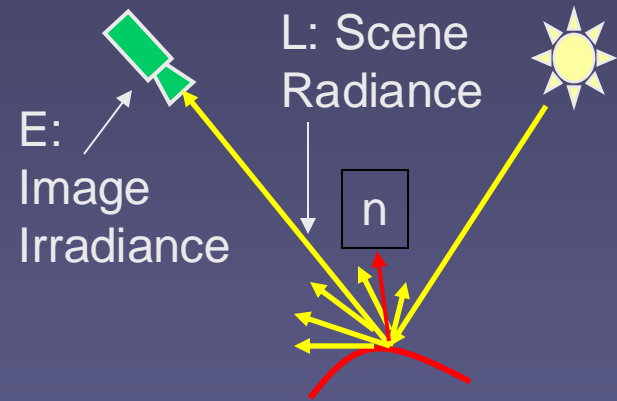
- Image irradiance is proportional to scene radiance $E \propto L$
- Image irradiance falls off as a function of $\cos^4(a)$
- Small field of view => small $\cos^4(a)$

Surface Radiance & Image Irradiance

Image irradiance

$$E = L \frac{\pi}{4} \left(\frac{d}{f} \right)^2 \cos^4(\alpha)$$

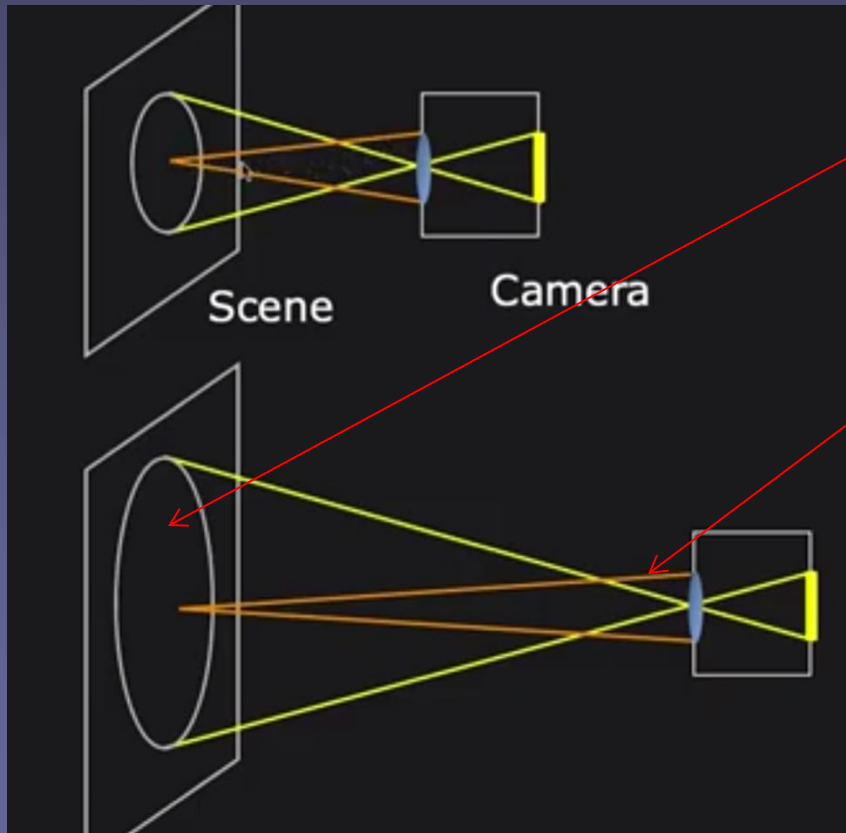
Scene radiance



DOES Image Irradiance VARY WITH SCENE DEPTH?

Image Irradiance Does Not Vary with Scene Depth

Varying Scene Depth



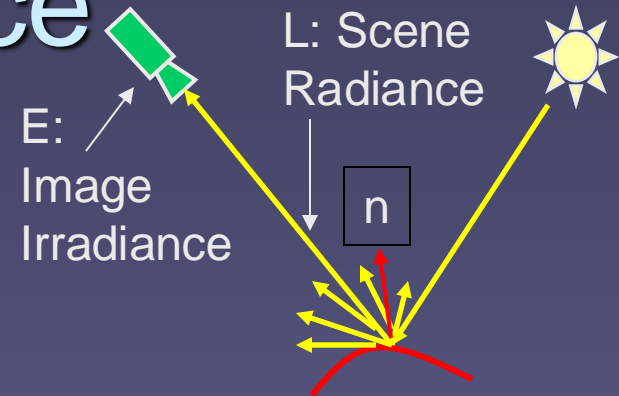
- Larger Scene Depth, Larger light accumulation
- Larger Scene Depth, Smaller Solid Angle subtended by each point

$$E = L \frac{\pi}{4} \left(\frac{d}{f} \right)^2 \cos^4(\alpha)$$

Surface Radiance & Image Irradiance

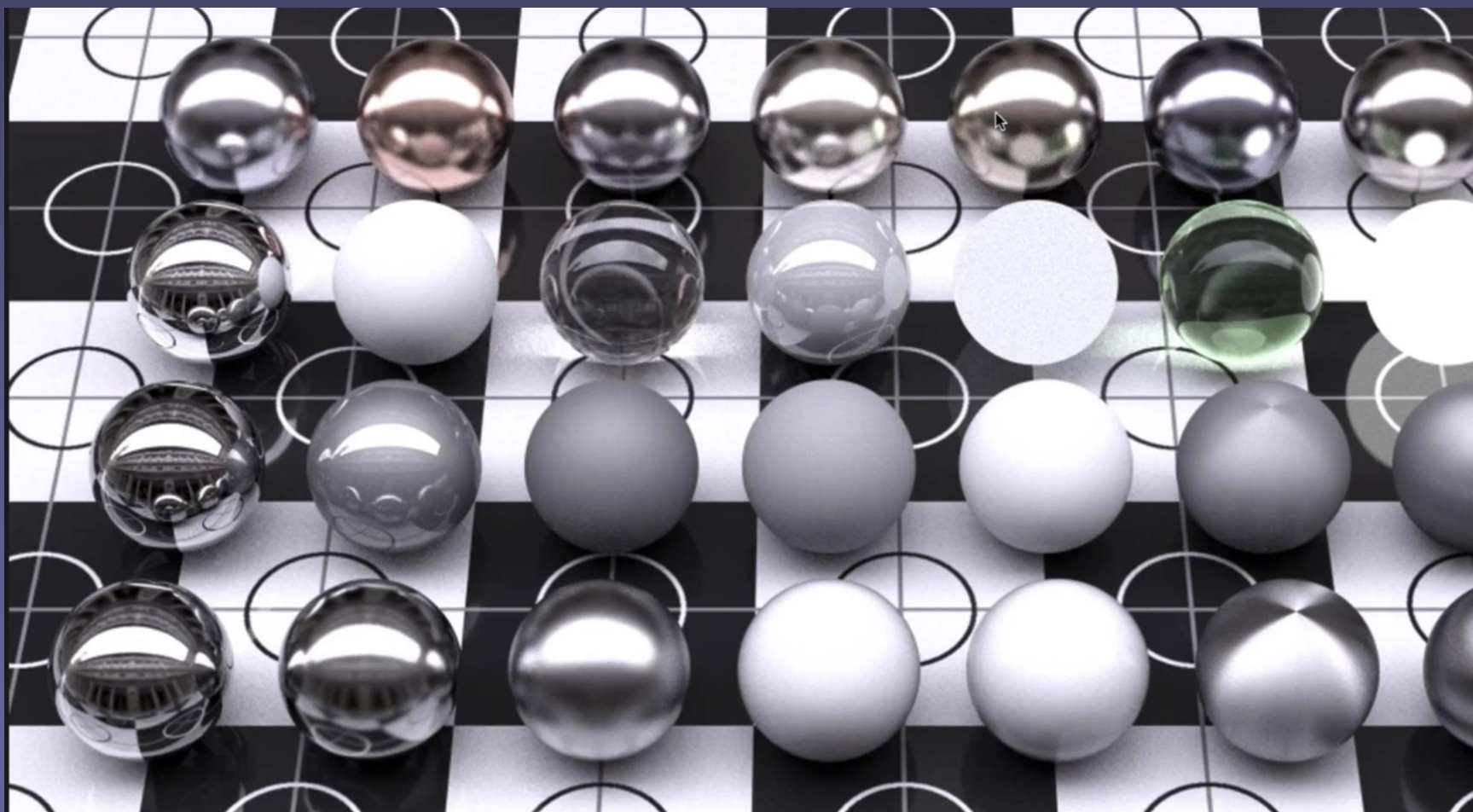
Image irradiance

$$E = L \frac{\pi}{4} \left(\frac{d}{f} \right)^2 \cos^4(\alpha)$$



DOES Image Irradiance VARY WITH SCENE DEPTH? NO

Surface Appearance



Measuring appearance

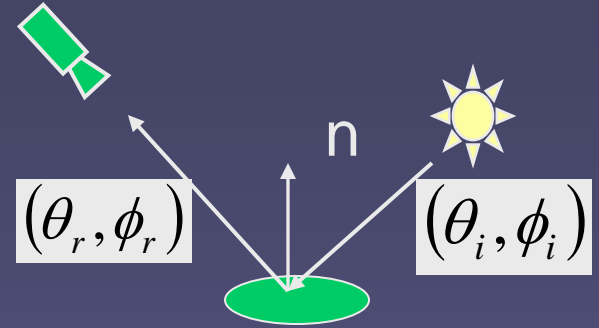
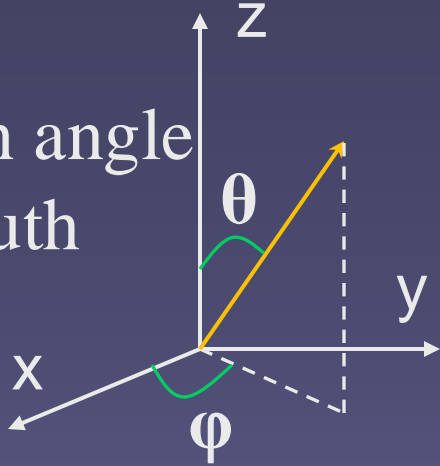


From “Fraunhofer IOSB”

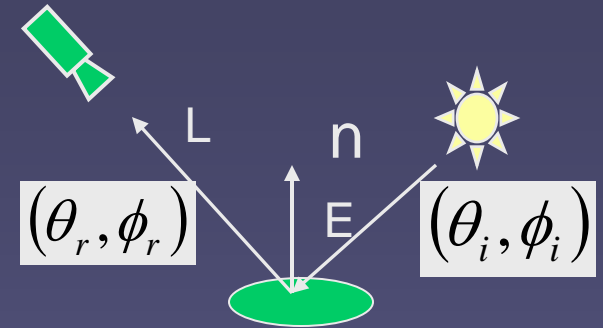
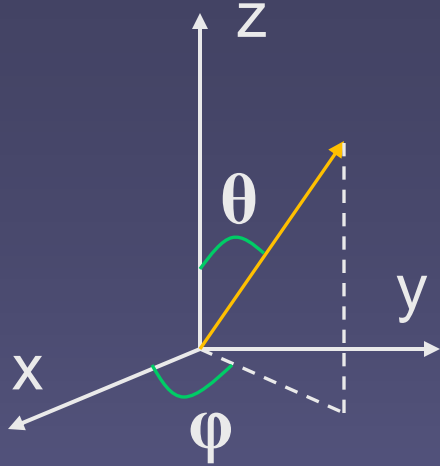
Bi-Directional Reflectance Distribution Function (BRDF)

θ : zenith angle

ϕ : azimuth
angle



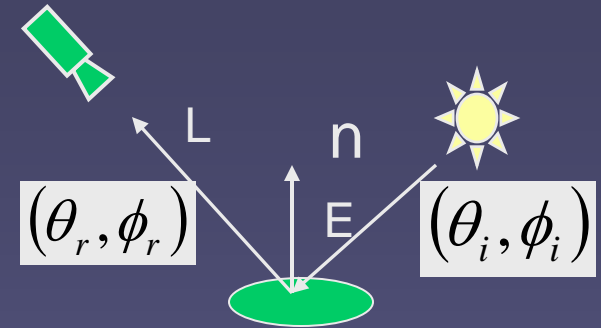
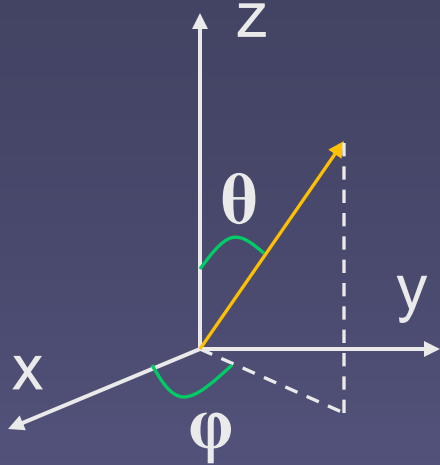
Bi-Directional Reflectance Distribution Function (BRDF)



$E(\theta_i, \phi_i)$: Irradiance due to source in direction (θ_i, ϕ_i)

$L(\theta_r, \phi_r)$: Radiance of surface in direction (θ_r, ϕ_r)

Bi-Directional Reflectance Distribution Function (BRDF)



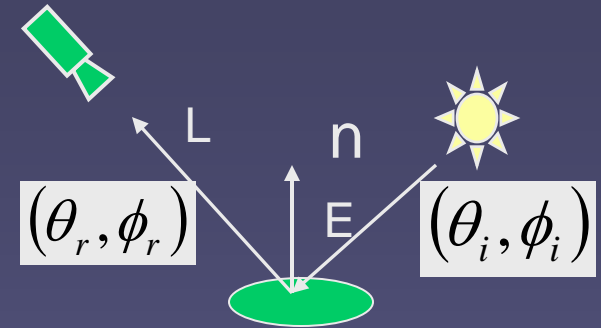
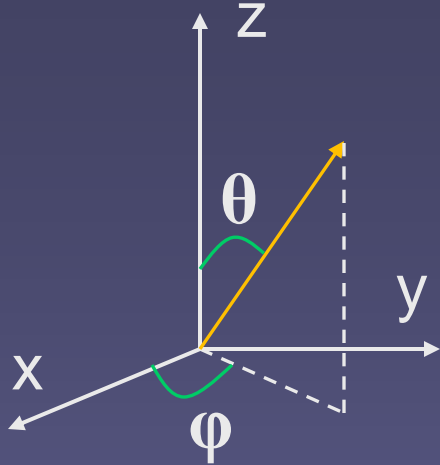
$E(\theta_i, \phi_i)$: Irradiance due to source in direction (θ_i, ϕ_i)

$L(\theta_r, \phi_r)$: Radiance of surface in direction (θ_r, ϕ_r)

BRDF:
$$f(\theta_i, \phi_i, \theta_r, \phi_r) = \frac{L(\theta_r, \phi_r)}{E(\theta_i, \phi_i)}$$

4-D function
Unit: 1 / steradian

Bi-Directional Reflectance Distribution Function (BRDF)



BRDF:

$$f(\theta_i, \phi_i, \theta_r, \phi_r) = \frac{L(\theta_r, \phi_r)}{E(\theta_i, \phi_i)}$$

4-D function

Unit: 1 / steradian

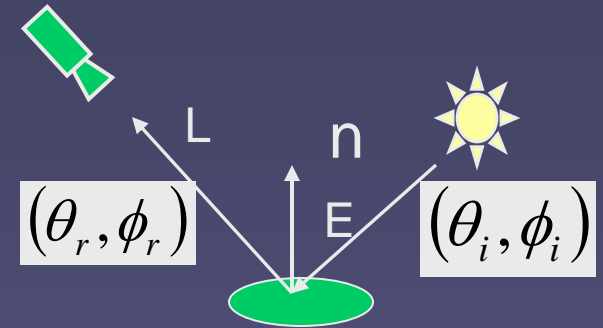
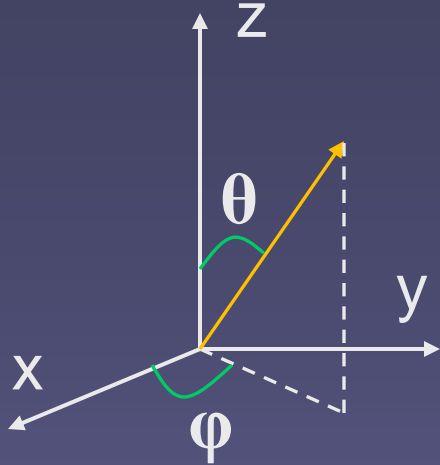
PROPERTIES:

1. BRDF f is always > 0

2. Helmholtz Reciprocity:

$$f(q_i, j_i, q_r, j_r) = f(q_r, j_r, q_i, j_i)$$

Bi-Directional Reflectance Distribution Function (BRDF)



BRDF:

$$f(\theta_i, \phi_i, \theta_r, \phi_r) = \frac{L(\theta_r, \phi_r)}{E(\theta_i, \phi_i)}$$

4-D function

Unit: 1 / steradian

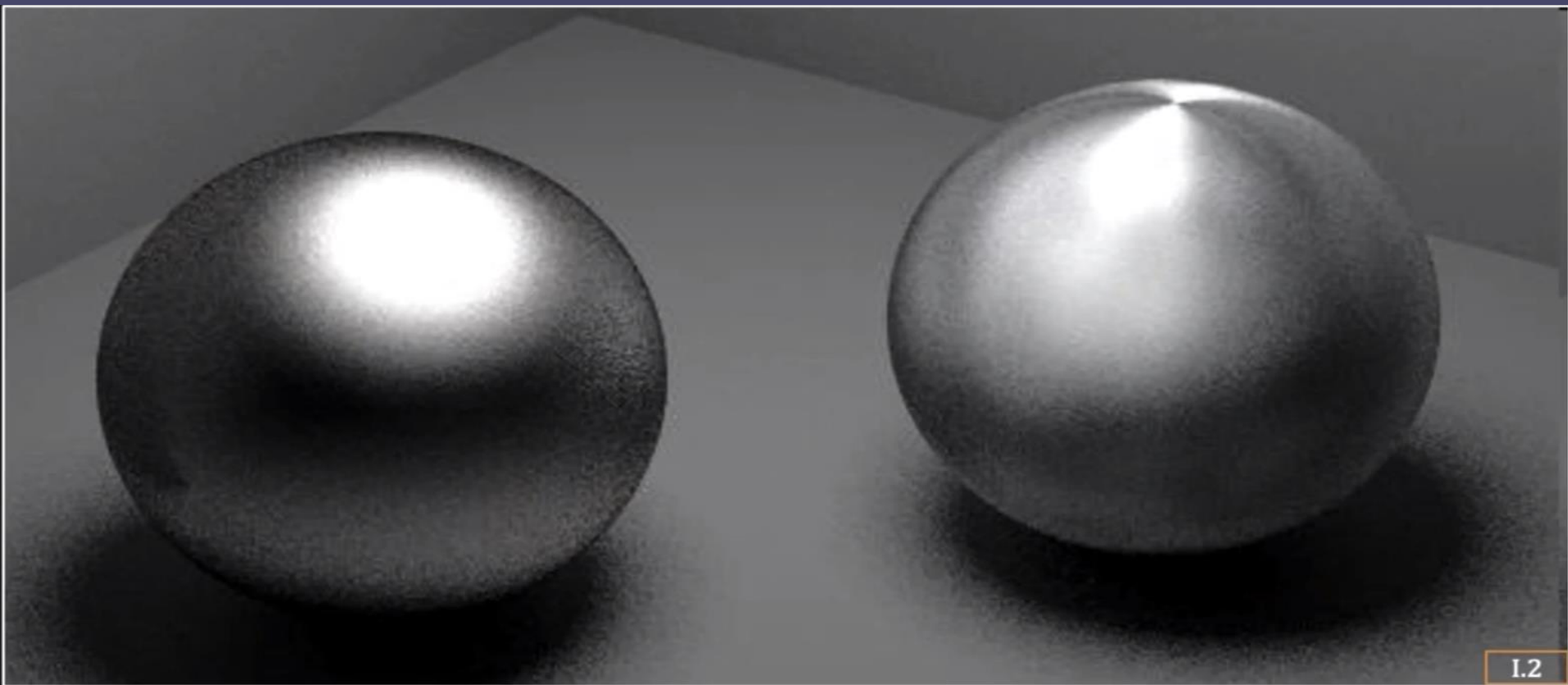
For Rotationally Symmetric Reflectance Properties :

BDRF:

$$f(\theta_i, \theta_r, (\phi_r - \phi_i))$$

(ISOTROPIC SURFACES: 3-D)

Isotropic vs. Anisotropic BRDF



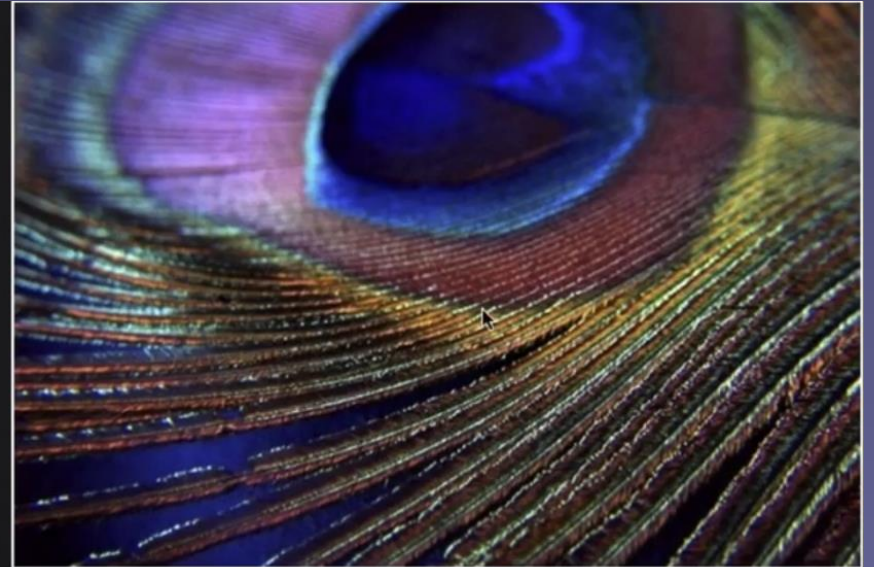
Isotropic BRDF

Anisotropic BRDF

Anisotropic BRDF in nature



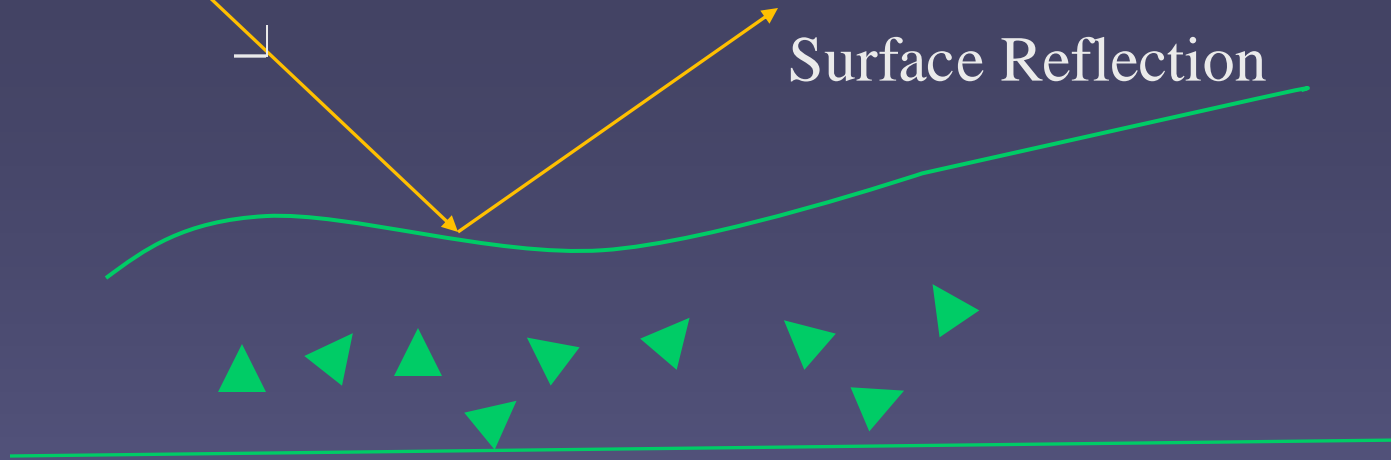
Butterfly wings



Peacock feathers



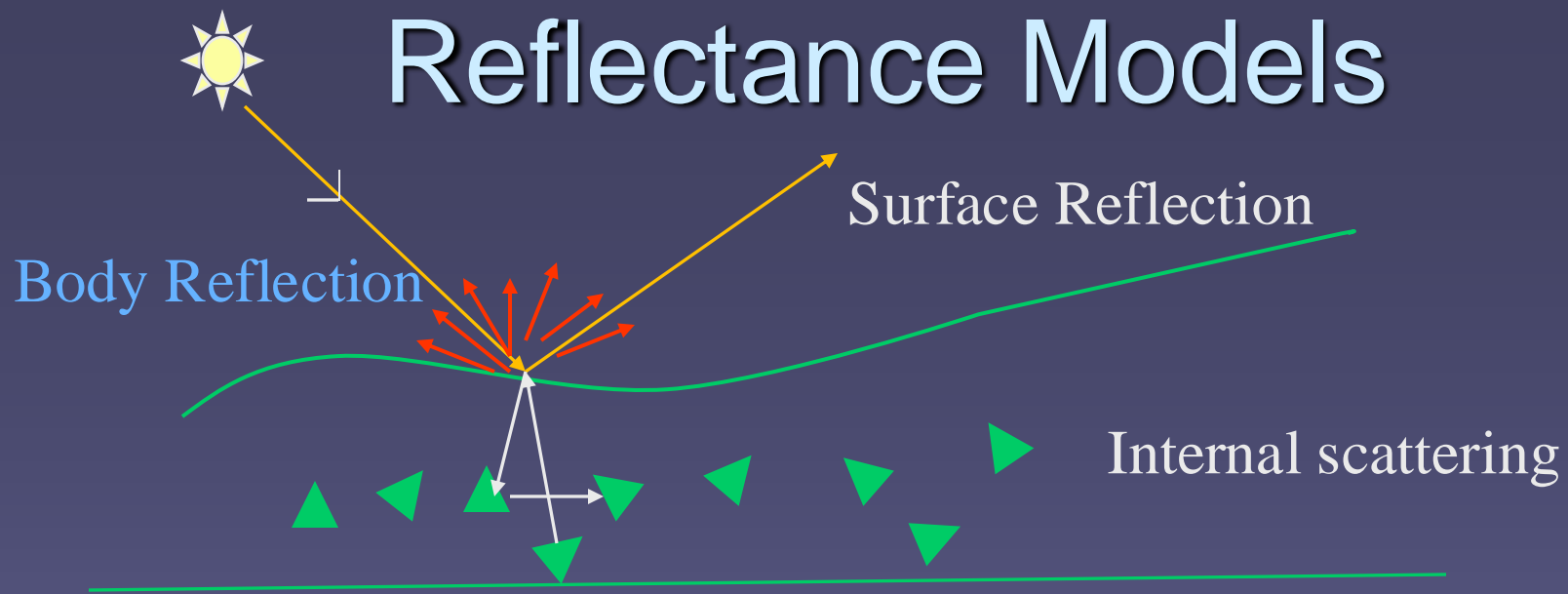
Reflectance Models



Surface Reflection:

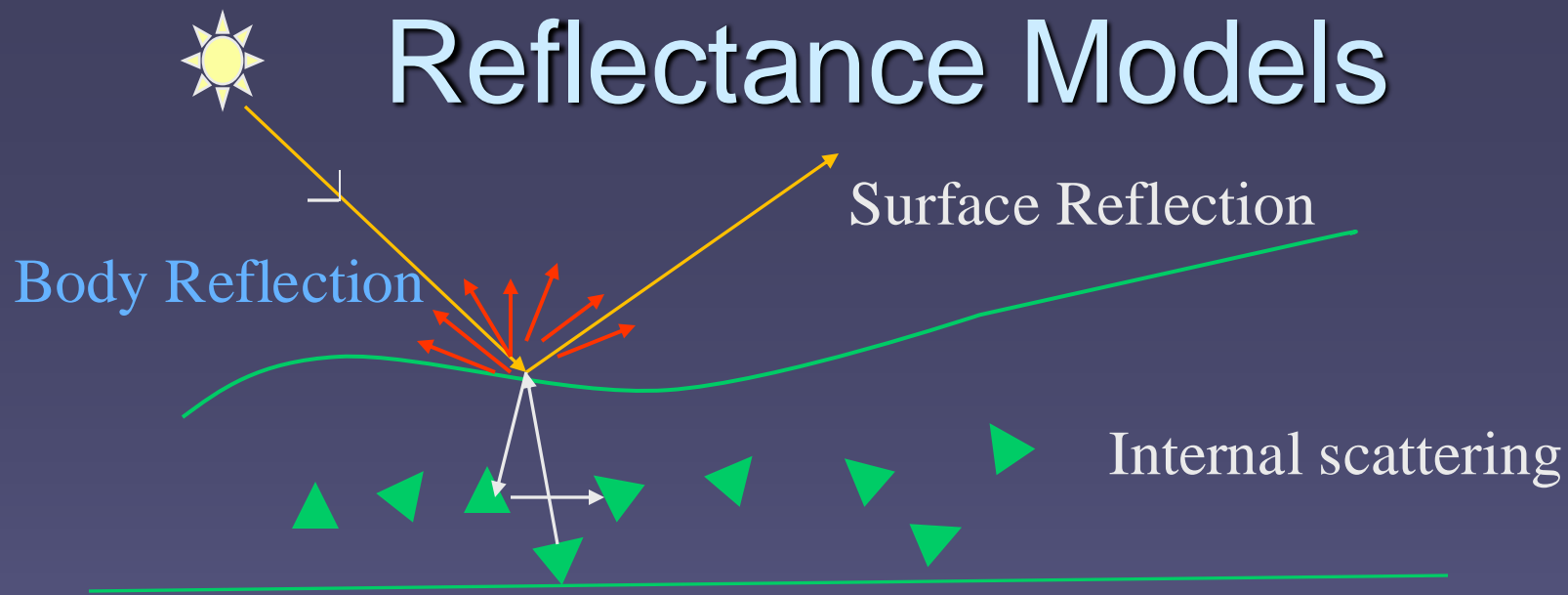
- *Specular Reflection
- *Glossy Appearance
- *Highlights.
- *Smooth surfaces (ex. metal mirror , glass)

Reflectance Models



Body Reflection: *Diffuse Reflection
*Matte Appearance
*Non-Homogeneous Medium (ex. clay, paper)

Reflectance Models

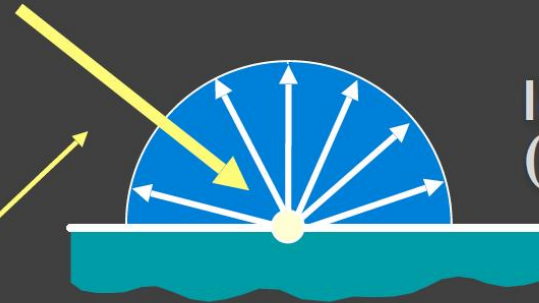


Body Reflection: *Diffuse Reflection
*Matte Appearance
*Non-Homogeneous Medium (ex. clay, paper)

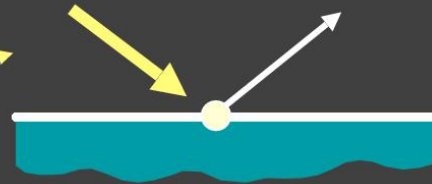
Surface Reflection: *Specular Reflection
*Glossy Appearance
*Highlights.
*Smooth surfaces (ex. metal mirror , glass)

Object appearance: Diffuse Reflection + Specular Reflection

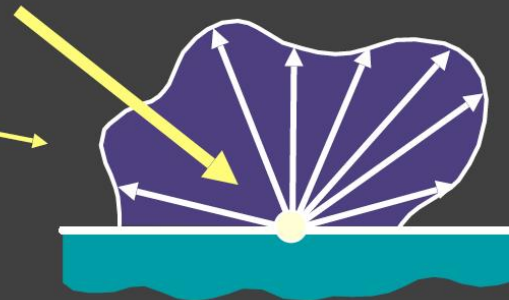
Materials - Three Forms



Ideal diffuse
(Lambertian)

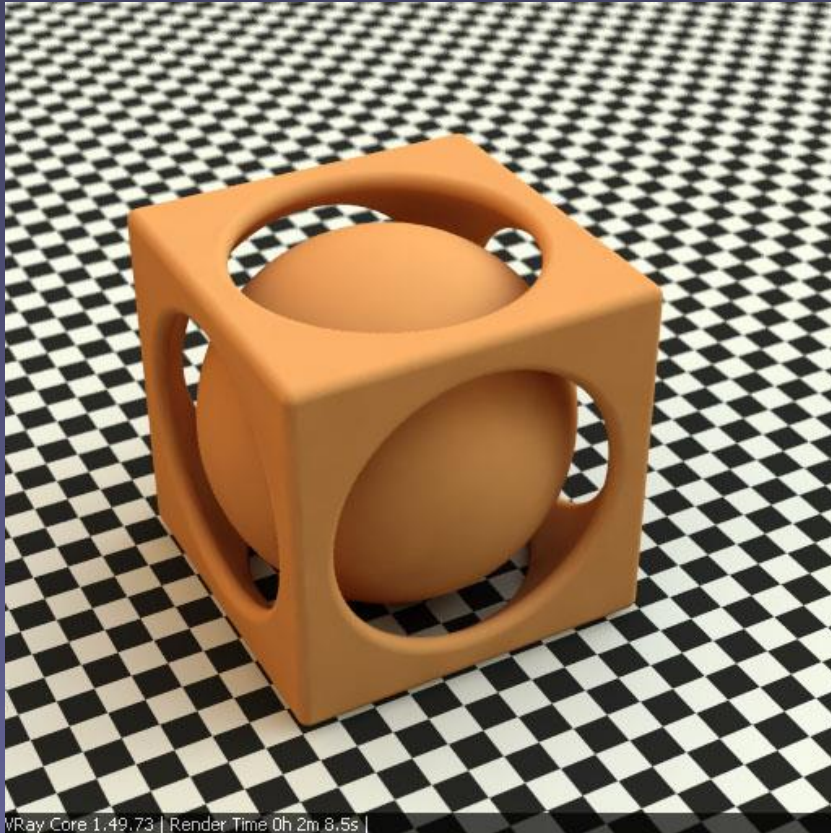


Ideal
specular



Directional
diffuse

Diffuse vs Specular Reflection



Mostly diffuse (no highlights)



Mostly specular (highlights)

Body Reflection:



Surface Reflection:

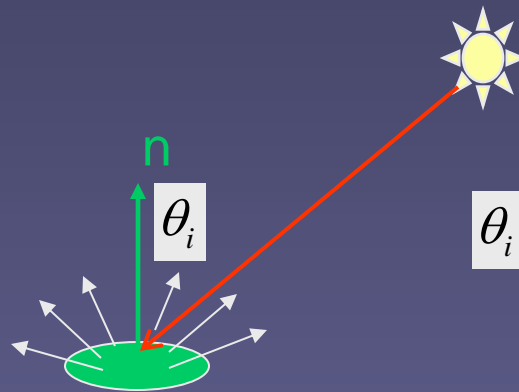


Hybrid Reflection
(Body + Surface):



[Nayar 1991]

Lambertian Reflectance Model (Body Reflection)



θ_i is the angle between light source direction and surface normal

A Lambertian (i.e. diffuse) surface scatters light equally in all directions!

Lambert (1760)

Very widely used in Vision & Graphics

Lambertian Reflectance Model

A Lambertian (diffuse) surface scatters light equally in all directions

Constant **BRDF** f :

$$f = \frac{r}{\rho}$$

← Albedo ρ :
intrinsic brightness of
surface

$$0 \leq \rho \leq 1$$

0: black surface

1: fully reflective

Surface appears equally bright from all viewing directions

Lambertian Reflectance Model

A Lambertian (diffuse) surface scatters light equally in all directions

$$f(q_i, j_i, q_r, j_r) = \frac{r}{\rho} \Rightarrow$$
$$L(q_r, j_r) = \frac{r}{\rho} E(q_i, j_i)$$
$$L = \frac{r}{\rho} E$$

Albedo ρ :
intrinsic brightness of
surface

$$0 \leq \rho \leq 1$$

Surface appears equally bright from all viewing directions

Lambertian Reflectance Model

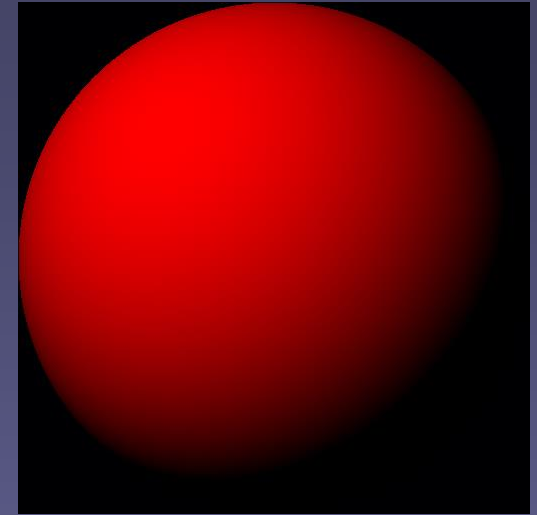
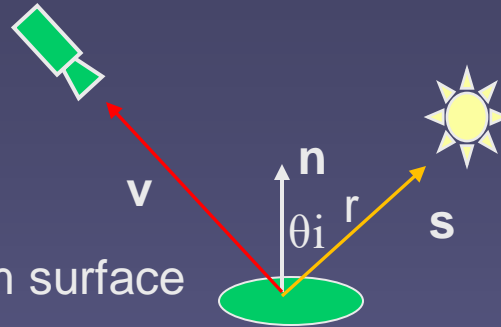
Unit Vectors

v: viewing direction

n: surface normal

s: light source direction

r: distance of light source from surface



A Lambertian sphere

Lambertian Reflectance Model

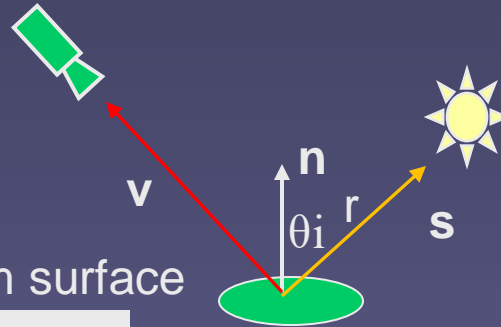
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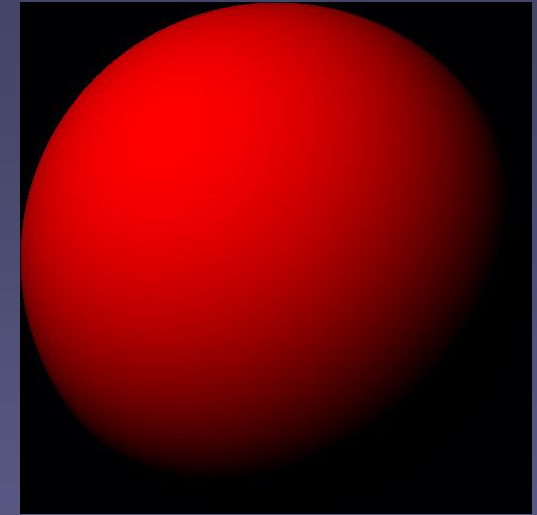
$$L = \frac{r}{\rho} E$$

$$E = (J \cos \theta_i / r^2) = \frac{J}{r^2} (\mathbf{n} \cdot \mathbf{s}) \Rightarrow$$

$$L = \frac{r}{\rho} \frac{J}{r^2} (\mathbf{n} \cdot \mathbf{s})$$

or

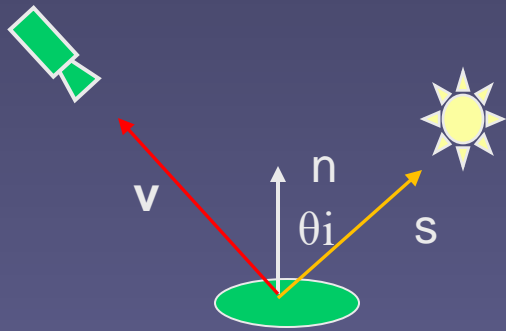
$$L = r k (\mathbf{n} \cdot \mathbf{s}), k = \frac{J}{\rho r^2}$$



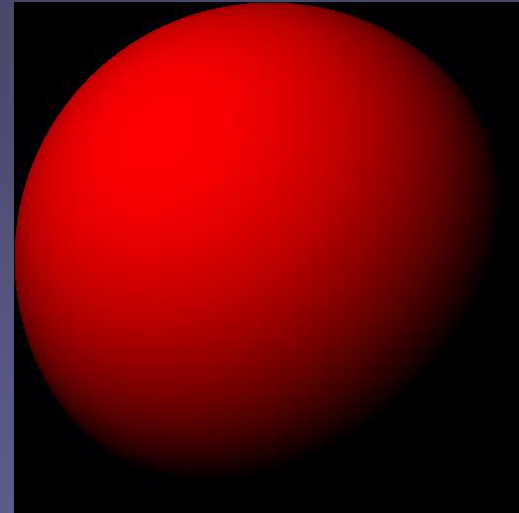
A Lambertian sphere

Dot product between unit vectors = cos

Lambertian Reflectance Model



Surface normal \mathbf{n}
Direction of illumination \mathbf{s}



A Lambertian sphere

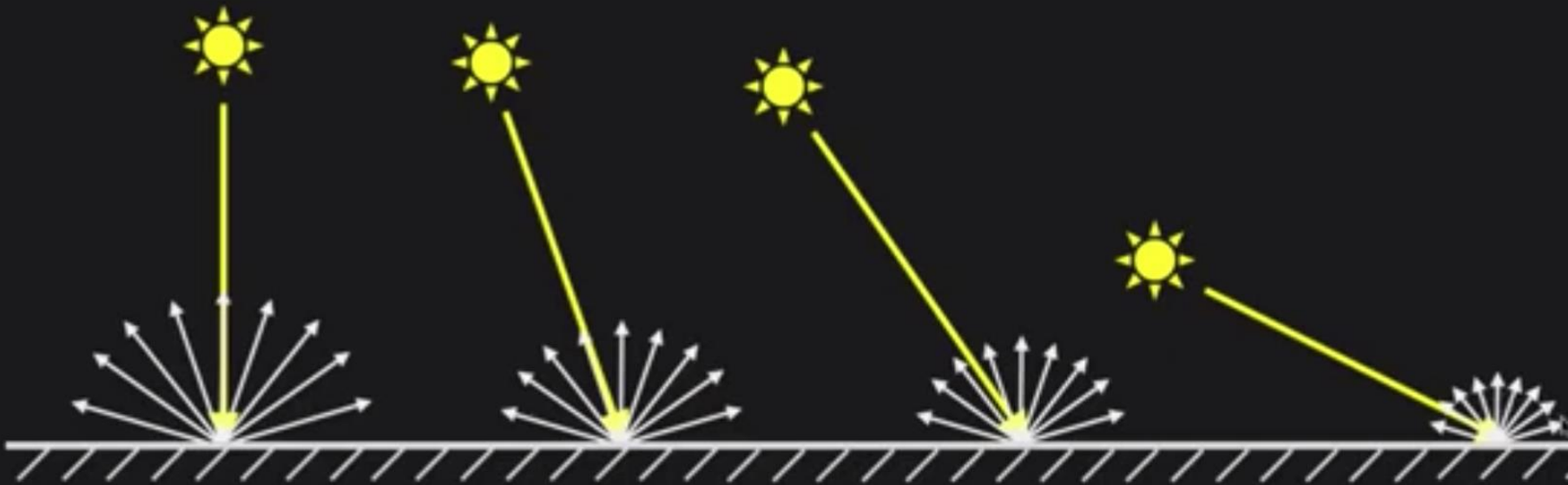
$$L = rk \cos q_i$$
$$L = r'(\mathbf{n} \times \mathbf{s})$$

Effective albedo

Commonly used in Computer Vision
and Graphics

Lambertian Reflectance Model

$$L = \frac{\rho_d}{\pi} \frac{J}{r^2} (\bar{n} \cdot \bar{s})$$



What Information Does Shading Encode

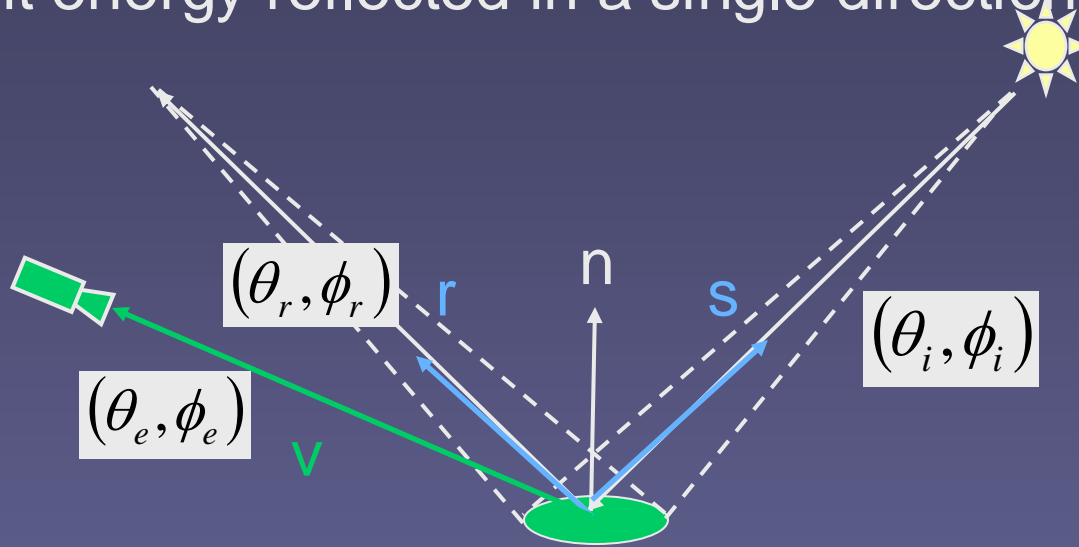
In regions of constant albedo, changes of intensity correspond to changes in the surface normal of the scene

$$I \propto L = r'(\mathbf{n} \cdot \mathbf{s})$$



Ideal Specular Model (Mirrors)

- *Very SMOOTH surface
- *All incident energy reflected in a single direction



Perfect reflector

$$f(q_i, j_i, q_e, j_e) = \frac{d(q_i - q_e) d((j_i + p) - j_e)}{\cos q_i \sin q_i}$$

Viewer received light only when $v = r$

Glossy Surfaces

- Delta function $\delta()$ too harsh a BRDF model
(valid only for highly polished mirrors and metals).
- Many glossy surfaces show broader highlights in addition to mirror reflection.



- Surfaces are not perfectly smooth – they show micro-surface geometry (roughness).
- Example Models : Phong model

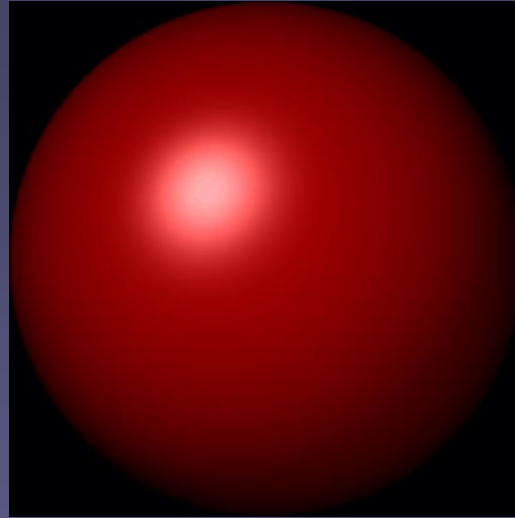
Torrance Sparrow model

Phong Reflectance Model

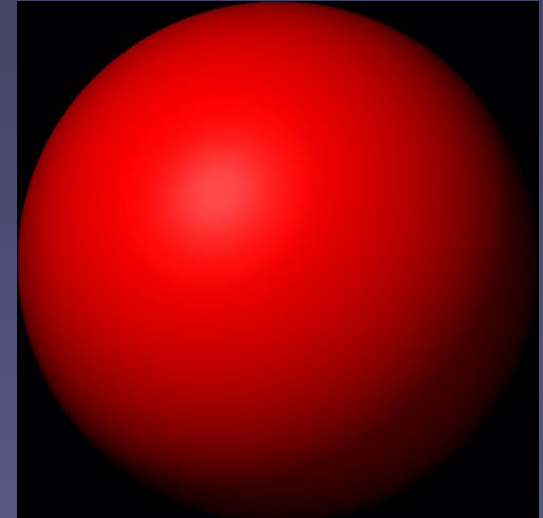
$$L = b \cos \theta_i + c \cos^n(\alpha)$$

diffuse

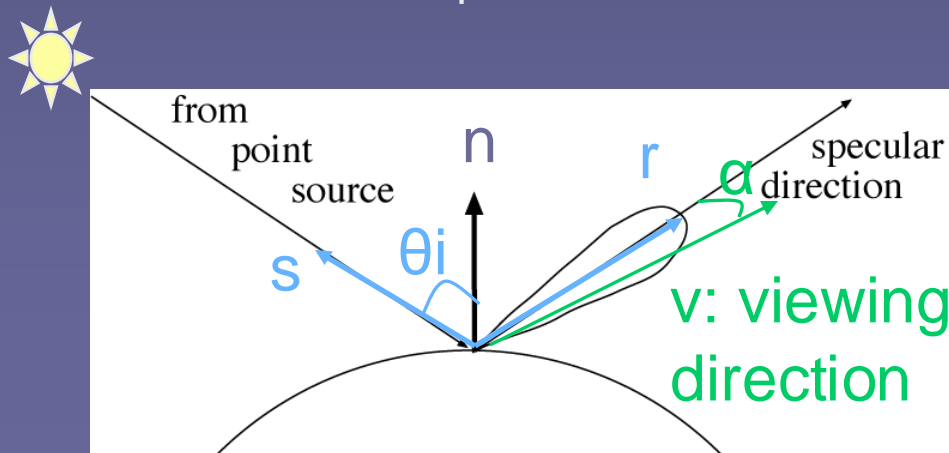
specular



$b=0.3, c=0.7, n=2$



$b=0.7, c=0.3, n=0.5$

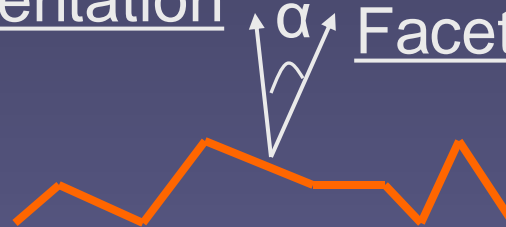


Torrance-Sparrow Model

Specular Reflection from Rough Surfaces.

Surface Micro-Structure Model – Each facet perfect mirror

Mean orientation



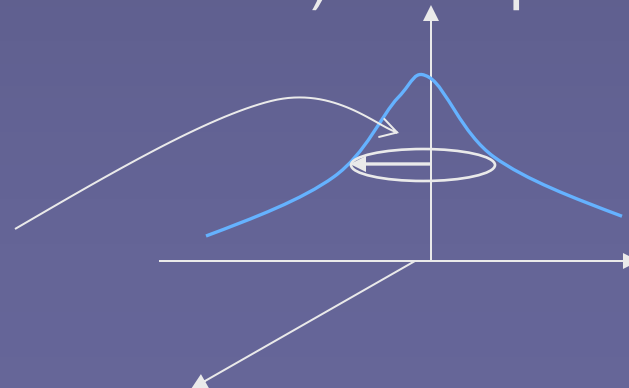
Facet orientation

Micro-facet Orientation Model: (example)

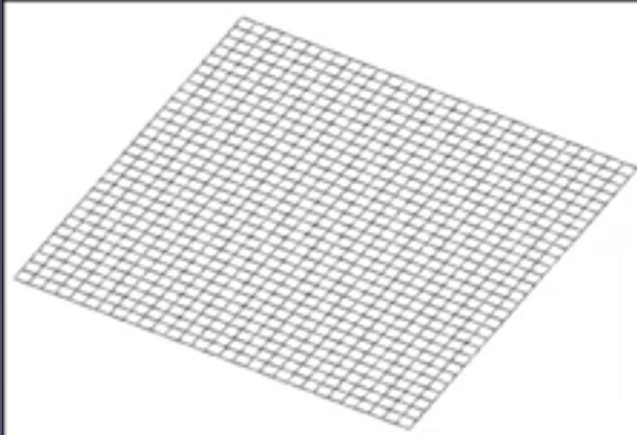
$$p(a) = \frac{1}{\sqrt{2\pi}\sigma} e^{-\frac{1}{2}\frac{a^2}{\sigma^2}}$$

(Gaussian Model) Isotropic

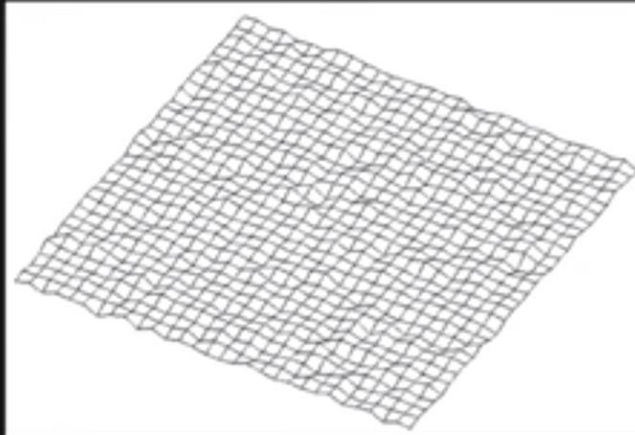
σ : roughness parameter



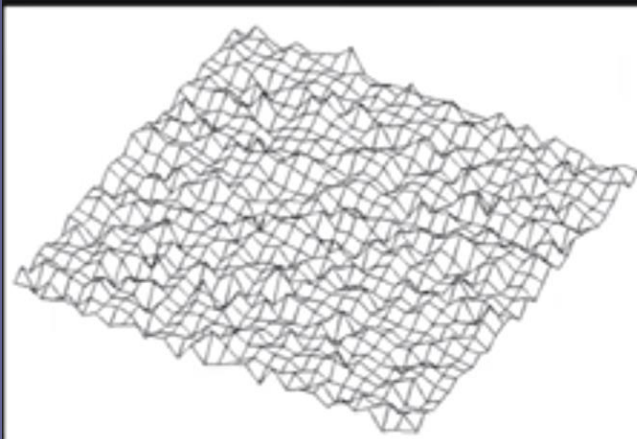
Micro-facet Orientation



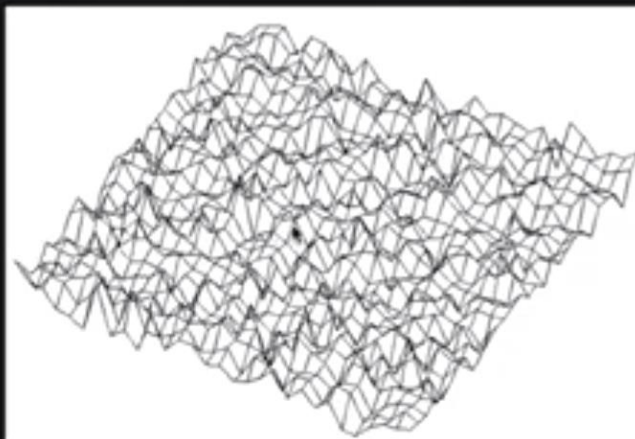
$\sigma = 0$



$\sigma = 0.1$



$\sigma = 0.3$

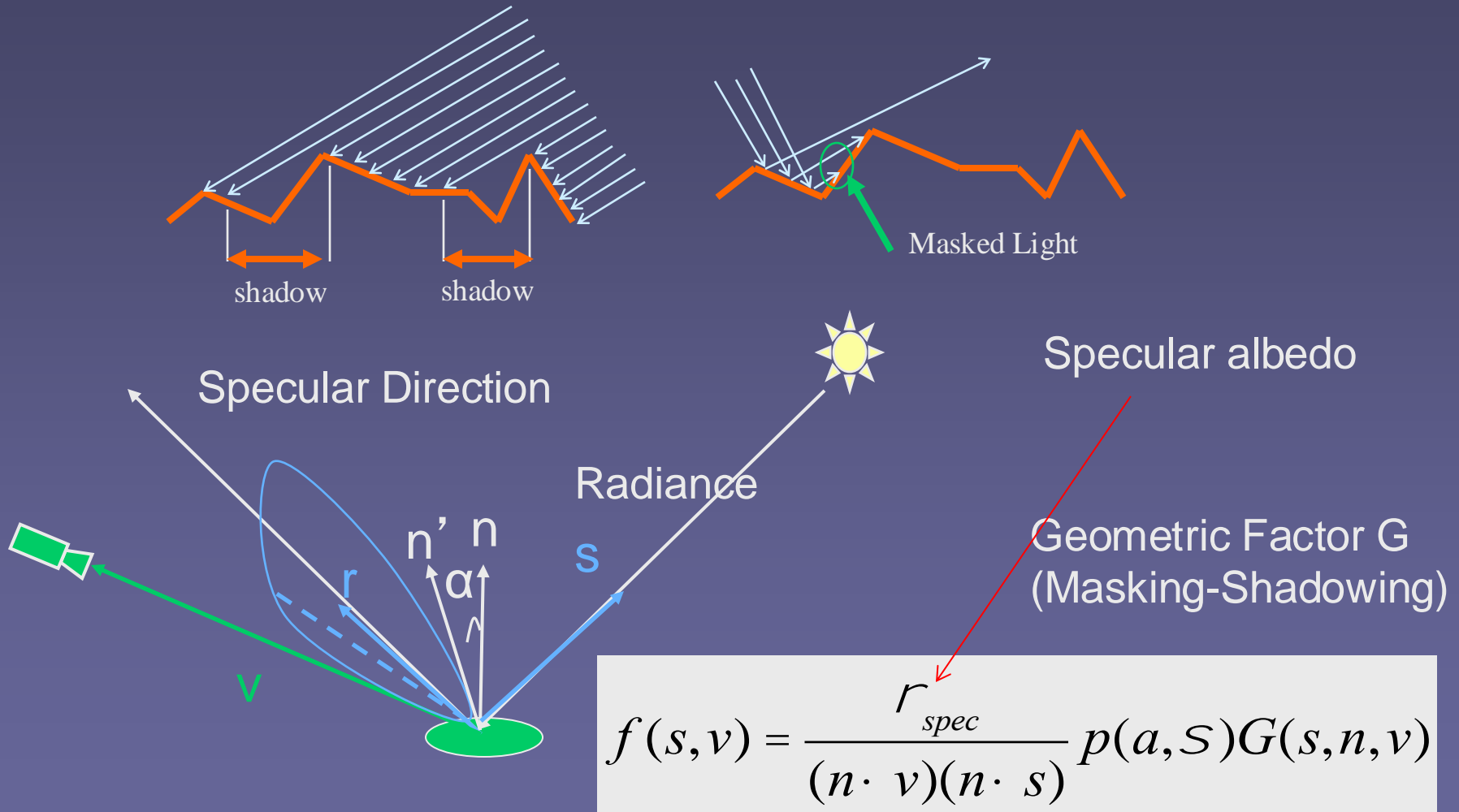


$\sigma = 0.6$

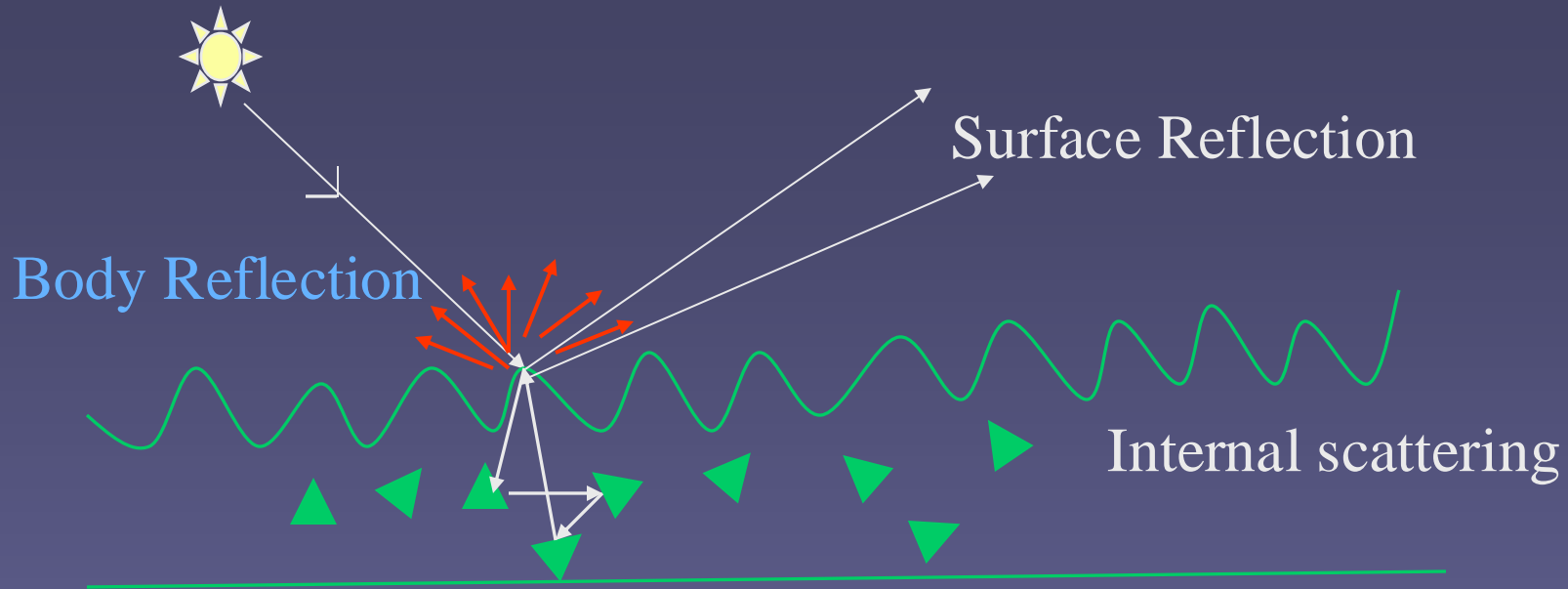
Shree Nayar

Torrance-Sparrow Model

Masking and Shadowing Effects



Dichromatic Model

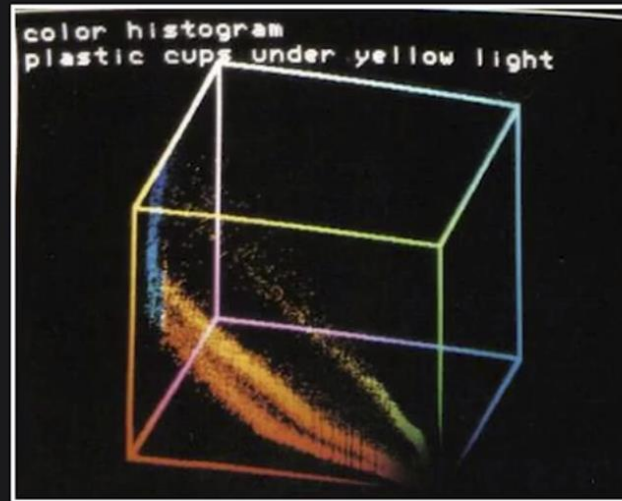


Color of body (diffuse) component = color of object x color of illumination

Color of surface (specular) component = color of illumination

We measure the sum of the diffuse and specular components.

Removing Specularities

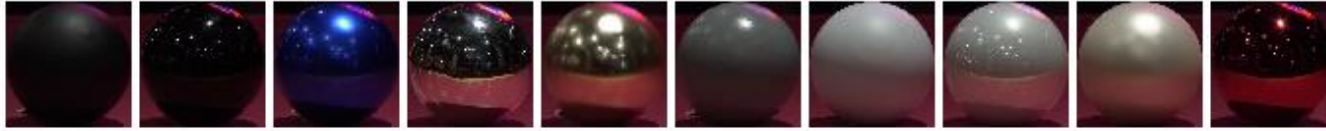


Klinker, Shafer,
Kanade,
1990,
International
Journal of
Computer
Vision

Illumination: Office scene



Illumination: Kendall Food Court



Illumination: Adelson Lab



Illumination: NE20 4th floor lobby



Illumination: Street scene



Illumination: By a window



Illumination: Under a desk lamp



Dror,
Adelson,
Wilsky