## 2018 Proposal Flux Plots

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### 1 Flux Estimates and Measurements

#### 1.1 Theoretical Predictions

Two very common theoretical predictions for the GZK cosmogenic neutrino flux are those by *Ahlers and Halzen* [1] and by *Kotera et. al.* [2].

### 1.2 IceCube Measurement

We will utilize three of the most recent IceCube flux measurements. The first is their measurement of the astrophysical muon neutrino spectrum using eight-years of through-going muons [3]:

$$\frac{d\Phi_{\nu+\bar{\nu}}}{dE} = 3 \times (1.01^{+0.26}_{-0.23}) \left(\frac{E}{100 \text{ TeV}}\right)^{-2.19 \pm 0.10} \cdot 10^{-18} \text{ GeV cm}^{-2} \text{ s}^{-1} \text{ sr}^{-1}$$
 (1)

The second is their measurement of the all-flavor astrophysical neutrino spectrum using four years of cascades [4]:

$$\frac{d\Phi_{\nu+\bar{\nu}}}{dE} = 3 \times (1.57^{+0.23}_{-0.22}) \left(\frac{E}{100 \text{ TeV}}\right)^{-2.48 \pm 0.08} \cdot 10^{-18} \text{ GeV cm}^{-2} \text{ s}^{-1} \text{ sr}^{-1}$$
 (2)

Their most recent *peer-reviewed* result is their combined-likelihood analysis, which utilizes both tracks and cascades [5]:

$$\frac{d\Phi_{\nu+\bar{\nu}}}{dE} = (6.7^{+1.1}_{-1.2}) \left(\frac{E}{100 \text{ TeV}}\right)^{-2.50 \pm 0.09} \cdot 10^{-18} \text{ GeV cm}^{-2} \text{ s}^{-1} \text{ sr}^{-1}$$
 (3)

It is interesting to note that the spectral index  $\gamma$  in the muon-based measurement is considerably harder than the explicitly all-flavor measurements, but the tension is only  $\sim 2\sigma$ .

## 2 Event Number Estimate

To estimate the number of events that would be detected by an experiment, we must complete the following integral:

$$N = \int \left(\frac{dN}{dE dA d\Omega dt}\right) [\Omega A_{eff}] dt dE$$
 (4)

where we are integrating over a flux model. This can be discretized into a sum over energy bins:

$$N = \Delta t \sum_{i} \left( \frac{dN}{dE dA d\Omega dt} \right)_{i} [\Omega A_{eff}]_{i} \Delta E_{i}$$
 (5)

# References

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- [2] K. Kotera, D. Allard, A. V. Olinto, Cosmogenic Neutrinos: parameter space and detectabilty from PeV to ZeV, JCAP 1010 (2010) 013. arXiv: 1009.1382, doi:10.1088/1475-7516/2010/10/013.
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