

Generalizing the isothermal compressibility for QCD

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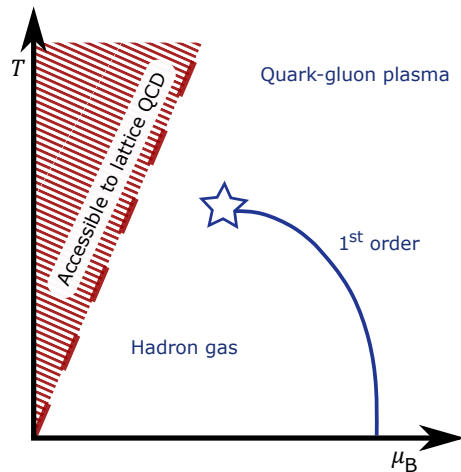
Broadly interested in **phase diagram** of strongly interacting systems.

At high enough temperatures and/or densities, hadrons dissociate to quark-gluon plasma.

Relevant to several systems:

- ▶ Early universe
- ▶ Neutron stars (NS)
- ▶ Heavy ion collisions (HIC)

Look for **critical end point (CEP)**!

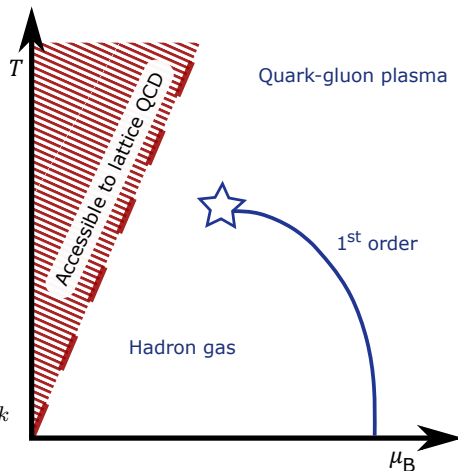


$$\begin{aligned}\langle O \rangle &\sim \int d\bar{\psi} d\psi e^{-\bar{\psi} D \psi} \int dU e^{-S(U)} O(U) \\ &= \int \det D \int dU e^{-S(U)} O(U)\end{aligned}$$

Complication (**sign problem**):

- ▶ $\det D \in \mathbb{R}$ when $\mu = 0$
- ▶ But if $\mu \neq 0$, it is complex...
- ▶ Can use tricks:

$$\frac{p}{T^4} \sim \sum_{i,j,k=0}^{\infty} \chi_{ijk}^{BQS} \left(\frac{\mu_B}{T}\right)^i \left(\frac{\mu_Q}{T}\right)^j \left(\frac{\mu_S}{T}\right)^k$$



Want to learn about composition and properties of strongly interacting systems. One way is through **material parameters** like **isothermal compressibility**

$$\kappa_T = -\frac{1}{V} \left(\frac{\partial V}{\partial P} \right)_{T,N}$$

Material parameters give physical intuition how strongly interacting matter deforms, expands, etc. For κ_T :

- ▶ Look for critical point (assuming it exists)

$$\kappa_T \sim |T - T^{\text{CEP}}|^{-\gamma}, \quad \gamma \approx 1.23$$

- ▶ location^{a,b}: $T^{\text{CEP}} < 100$, $\mu_B^{\text{CEP}} \gtrsim 420$ MeV
- ▶ κ_T also relates to n_B fluctuations

^aD. A. Clarke et al., Phys. Rev. D, 112.9, L091504 (2025).

^bD. A. Clarke et al., arXiv:2601.04782, (2026).

GOAL: compute material parameters at $\mu_B > 0$.

But only have *direct* access to $\mu_B = 0$ on the lattice. Commonly played game:

1. Write p/T^4 as **Taylor expansion** in μ_i/T
2. Derive material parameters from p/T^4 using **thermodynamics**
3. **Measure** Taylor coefficients on lattice
4. Compare against **HRG** for $T < T_{\text{pc}}$ (**crossover temp** ~ 156 MeV)

For convenience, $\hat{X} \equiv XT^{-k}$ with k s.t. \hat{X} dimensionless (e.g. $\hat{\mu} = \mu/T$)

Dealing with 3 chemical potentials $\hat{\mu}_B, \hat{\mu}_Q, \hat{\mu}_S$

To make contact with T - $\hat{\mu}_B$ plane, need to eliminate 2 independent variables

Can **impose external constraint**, e.g.

1. $\hat{\mu}_Q = \hat{\mu}_S = 0$
2. $n_S = 0, n_Q/n_B = 0.4$ (RHIC-like initial conditions, collide Au nuclei)
3. $n_S = 0, n_Q/n_B = 0.5$ (isospin-symmetric; yields $\hat{\mu}_Q = 0$)

and think of **expansions in $\hat{\mu}_B$ only**:

$$\hat{p} \sim \sum_{i,j,k=0}^{\infty} \chi_{ijk}^{BQS} \hat{\mu}_B^i \hat{\mu}_Q^j \hat{\mu}_S^k, \quad \chi_2^B \equiv \left. \frac{\partial^2 \hat{p}}{\partial \hat{\mu}_B^2} \right|_{\mu=0} \rightarrow \hat{p} = \sum_{k \text{ even}} P_k(T) \hat{\mu}_B^k$$

Non-interacting, quantum, relativistic gas eventually gives (single species)

$$\hat{p}^{\text{HRG}} \sim m^2 g \sum_{k=1}^{\infty} \frac{\eta^{k+1} z^k}{k^2} K_2 \left(\frac{mk}{T} \right), \quad z \equiv e^{\hat{\mu}_B B + \hat{\mu}_Q Q + \hat{\mu}_S S},$$

with K_2 modified Bessel function 2nd kind. HRG:

- ▶ Assume such gas where hadrons and resonances only d.o.f.
- ▶ Hence valid up to $\sim T_{\text{pc}}$
- ▶ Sum over all such states, each with g_i , m_i , etc.
- ▶ K_2 exponentially suppressed, so can keep few terms

With one chemical potential:

$$\kappa_T = -\frac{1}{V} \left(\frac{\partial V}{\partial P} \right)_{T,N} = \frac{1}{n^2} \frac{\partial^2 P}{\partial \mu^2}, \quad n = N/V$$

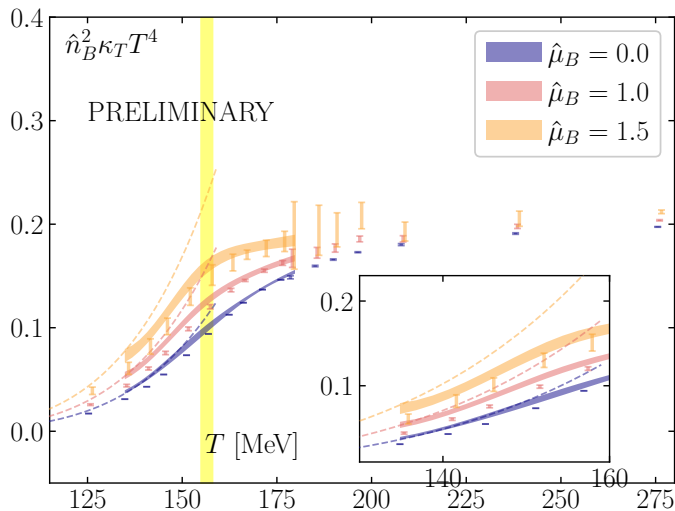
With B , Q , and S , $n_S = 0$, isothermal compressibility comes out to be

$$\kappa_T = \frac{F(X_{ijk}^{BQS})}{n_B^2 G(n_Q/n_B, X_{ijk}^{BQS})},$$

where e.g. $X_2^B \equiv \partial_{\mu_B}^2 p(\mu_B)$ depends on μ_B . Since $n_B \propto \mu_B$

κ_T diverges¹ as $\mu_B \rightarrow 0$

¹Also a problem for material parameters like C_p and α .



$$\kappa_T = -\frac{1}{V} \left(\frac{\partial V}{\partial P} \right)_{T,N} = \frac{1}{n^2} \frac{\partial^2 P}{\partial \mu^2}$$

Fix instead f s.t. $f \neq 0$ at $\mu = 0$:

$$\begin{aligned} \kappa_{T,f} T^4 &= -\frac{T^4}{V} \left(\frac{\partial V}{\partial P} \right)_{T,f,n_Q/n_B,n_S/n_B} \\ \kappa_{T,f} T^4 &= -\frac{1}{f \hat{n}_B} \left(\frac{\partial f}{\partial \hat{\mu}_B} k_B + \frac{\partial f}{\partial \hat{\mu}_Q} k_Q + \frac{\partial f}{\partial \hat{\mu}_S} k_S \right) \end{aligned}$$

at $n_S = 0$ where $k_i \left(n_Q/n_B, X_{ijk}^{BQS} \right)$.

$$\kappa_{T,f} T^4 = -\frac{1}{f \hat{n}_B} \left(\frac{\partial f}{\partial \hat{\mu}_B} k_B + \frac{\partial f}{\partial \hat{\mu}_Q} k_Q + \frac{\partial f}{\partial \hat{\mu}_S} k_S \right)$$

Take e.g. $f = V X_2^Q$. Advantages:

- ▶ **Meaningful** across T_{pc}
- ▶ Leads at $\mathcal{O}(\mu^2)$, so one derivative leads at $\mathcal{O}(\mu) \Rightarrow$ **finite**
- ▶ A good proxy for $N + \bar{N}$ in hadronic phase² \Rightarrow **probes similar physics**

Comes out to be

$$\kappa_{T,\sigma_Q^2} T^4 = \frac{\hat{X}_{12}^{BQ}}{\hat{n}_B \hat{X}_2^Q} \left(1 - \frac{\hat{X}_{21}^{QS} \hat{X}_{11}^{BS}}{\hat{X}_{12}^{BQ} \hat{X}_2^S} \right)$$

²P. Braun-Munzinger et al., Phys. Lett. B, 747, 292–298 (2015).

Related studies from the past, for instance $\hat{\mu}_B = 0$ ^{3,4} and $\hat{\mu}_B > 0$ ^{5,6}.

These studies compute c_s^2 and C_V . This study:

- ▶ Taylor series up to 6th order; converges well at least for $\hat{\mu}_B \lesssim 2$ ^{7,8}
- ▶ $N_f = 2 + 1$ with physical $m_s/m_l = 27$
- ▶ Focus⁹ on $n_S = 0$ and $n_Q/n_B = 0.5$
- ▶ First lattice determination of κ_T

³A. Bazavov et al., Phys. Rev. D, 90, 094503 (2014).

⁴S. Borsanyi et al., Phys. Lett. B, 730, 99–104 (2014).

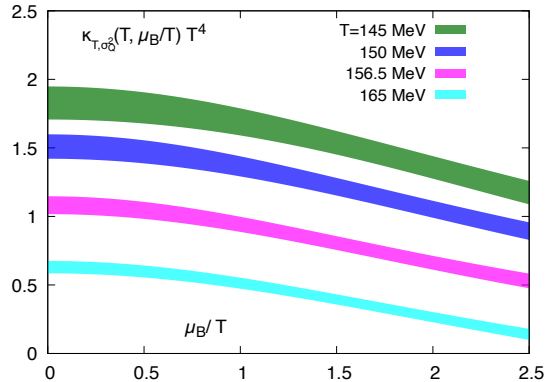
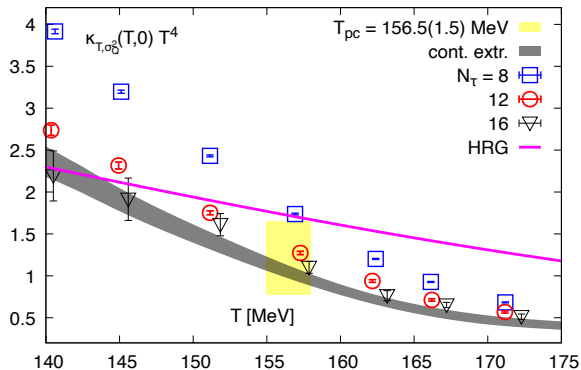
⁵A. Bazavov et al., Phys. Rev. D, 95.5, 054504 (2017).

⁶S. Borsanyi et al., JHEP, 10, 205 (2018).

⁷D. Bollweg et al., Phys. Rev. D, 105.7, 074511 (2022).

⁸Strictly speaking, convergence radius depends on temperature.

⁹Results at $n_Q/n_B = 0.4$ and $n_Q/n_B = 0.5$ are similar.



Decreasing function of μ_B and T
 Difference with HRG may be traceable back to Δ^{++} resonance

¹⁰D. A. Clarke et al., arXiv:2506.22816, (2025).

In HRG, κ_T can be expressed as¹¹

$$\kappa_{T,\vec{N}}^{-1} = -V \left(\frac{\partial P}{\partial V} \right)_{T,\vec{N}} = T^4 \left(\sum_{i \in \text{charged}} \frac{\hat{n}_i}{\omega_i} + \sum_{i \in \text{neutral}} \frac{\hat{n}_i}{\omega_i} \right), \quad \omega_i \equiv \sigma_i^2 / N_i$$

Variances ω_i similar, so to a good approximation,

$$\kappa_{T,\vec{N}}^{-1} \approx T^4 \left(\frac{\hat{n}_{\text{ch}}}{\omega_{\text{ch}}} + \frac{\hat{n}_0}{\omega_0} \right) \approx T \frac{n_{\text{tot}}}{\omega_{\text{tot}}},$$

where

$$\hat{n}_X = \sum_{i \in X} \hat{n}_i, \quad \omega_X = \frac{\sum_{i \in X} \sigma_i^2}{\sum_{i \in X} N_i}, \quad X = \text{ch}, 0, \text{tot}$$

¹¹M. Mukherjee et al., Phys. Lett. B, 784, 1–5 (2018).

$$\kappa_{T,\vec{N}}^{-1} \approx T^4 \left(\frac{\hat{n}_{\text{ch}}}{\omega_{\text{ch}}} + \frac{\hat{n}_0}{\omega_0} \right) \approx T \frac{n_{\text{tot}}}{\omega_{\text{tot}}},$$

- ▶ Neutral term often neglected in calculations from HIC
- ▶ Plugging in HRG pressure at T_{pc} yields $p\kappa_{T,\vec{N}} \approx 1.03$
- ▶ For an ideal gas, $p\kappa_{T,\vec{N}} = 1$

T_{pc} is a decreasing function of μ_B :

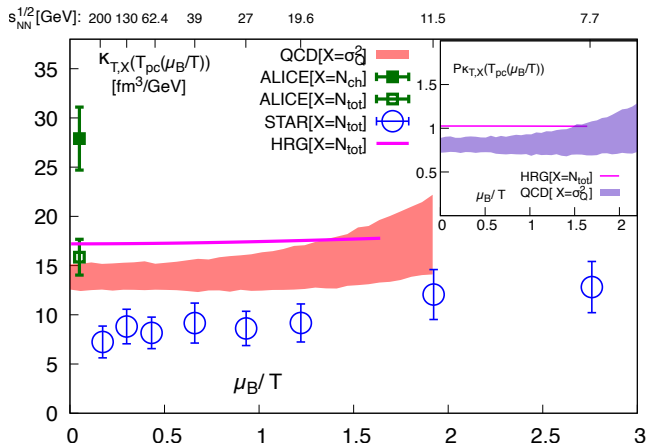
$$T_{\text{pc}}(\hat{\mu}_B) = T_{\text{pc},0} (1 - \kappa_2 \hat{\mu}_B^2 + \mathcal{O}(\hat{\mu}_B^4))$$

From lattice¹², $T_{\text{pc},0} = 156.5(1.5)$ MeV and $\kappa_2 = 0.012(4)$. Plug this expansion into expression for κ_{T,σ_Q^2} . Find

$$p\kappa_{T,\sigma_Q^2} \Big|_{T_{\text{pc}}(\hat{\mu}_B)} = 0.80(12) + 0.06(1)\hat{\mu}_B^2 + \mathcal{O}(\hat{\mu}_B^4)$$

Both definitions suggest similarity to ideal gas

¹²A. Bazavov et al., Phys. Lett. B, 795, 15–21 (2019).



- ▶ $p\kappa_{T,\sigma_Q^2}$ and $p\kappa_T$ from HRG are close within uncertainty
- ▶ Agreement with ALICE using both charged and neutral hadrons
- ▶ Compressibility computed from STAR using both charged and neutral hadrons slowly increases with μ_B/T
- ▶ Increase seems driven by N_{tot} ; no indication of characteristic divergence/CEP

¹³D. A. Clarke et al., arXiv:2506.22816, (2025).

- ▶ Introduced $\kappa_{T,f}$ finite at $\mu = 0$ in field theories!
- ▶ Consistent with ALICE
- ▶ κ_T from STAR shows no critical behavior
- ▶ κ_{T,σ_Q^2} on pseudocritical line stays close to that of an ideal gas

Thanks for your attention.

$$\kappa_{T,\sigma_Q^2} T^4 \propto \frac{\hat{X}_{12}^{BQ}}{\hat{n}_B \hat{X}_2^Q} = \frac{\chi_{22}^{BQ}}{\chi_2^B \chi_2^Q} \left(\frac{1 - \frac{\chi_{121}^{BQS} \chi_{11}^{BS}}{\chi_{22}^{BQ} \chi_2^S} + \mathcal{O}(\vec{\mu}^2)}{1 - \frac{(\chi_{11}^{BS})^2}{\chi_2^B \chi_2^S} + \mathcal{O}(\vec{\mu}^2)} \right)$$

