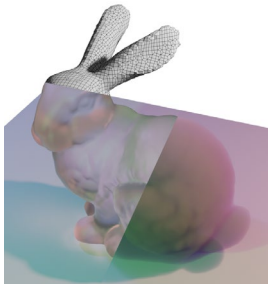


# 1. Radiometry and Light transport

---

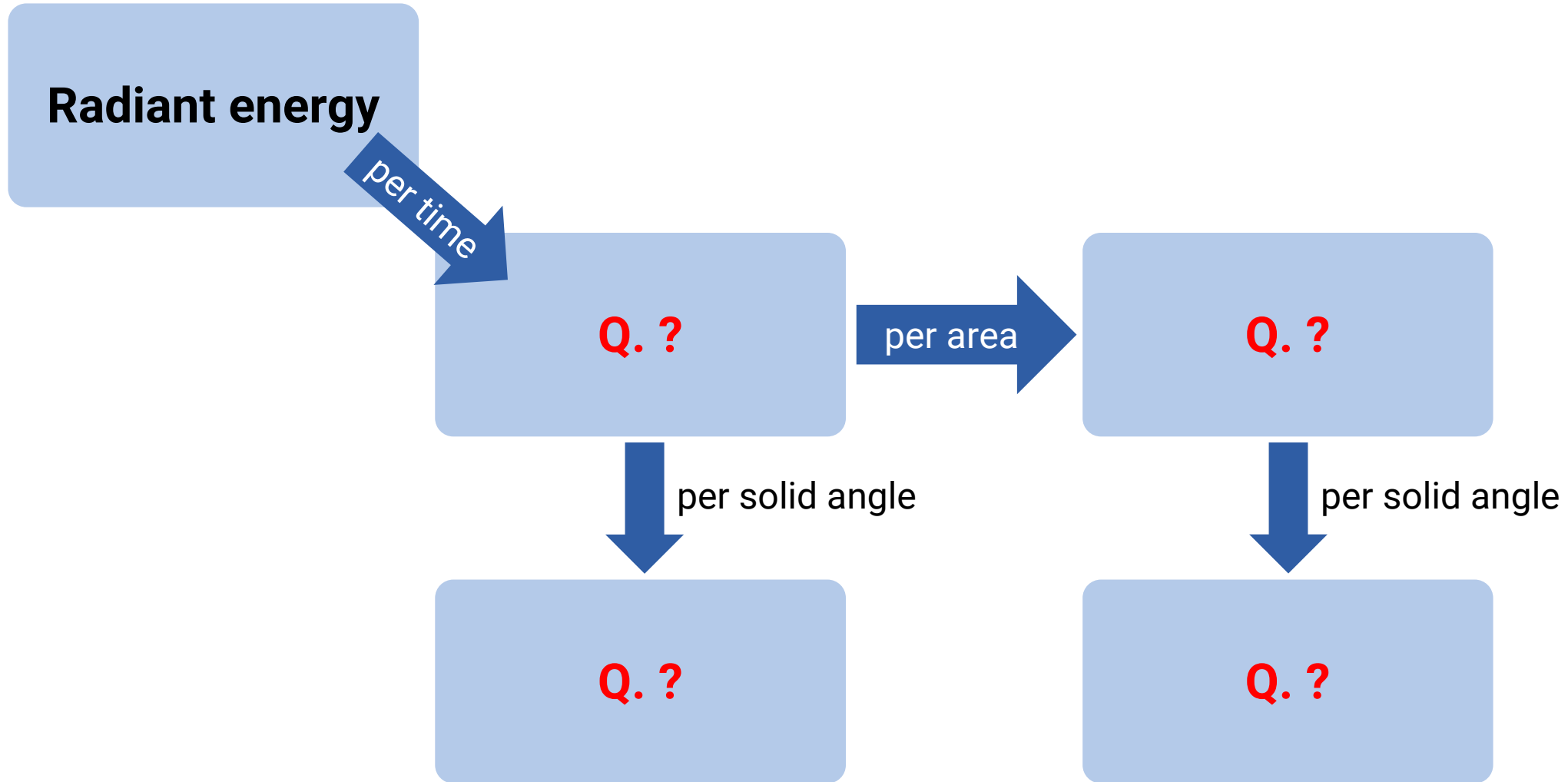
Physically Based Rendering

Shinyoung Yi (이신영)

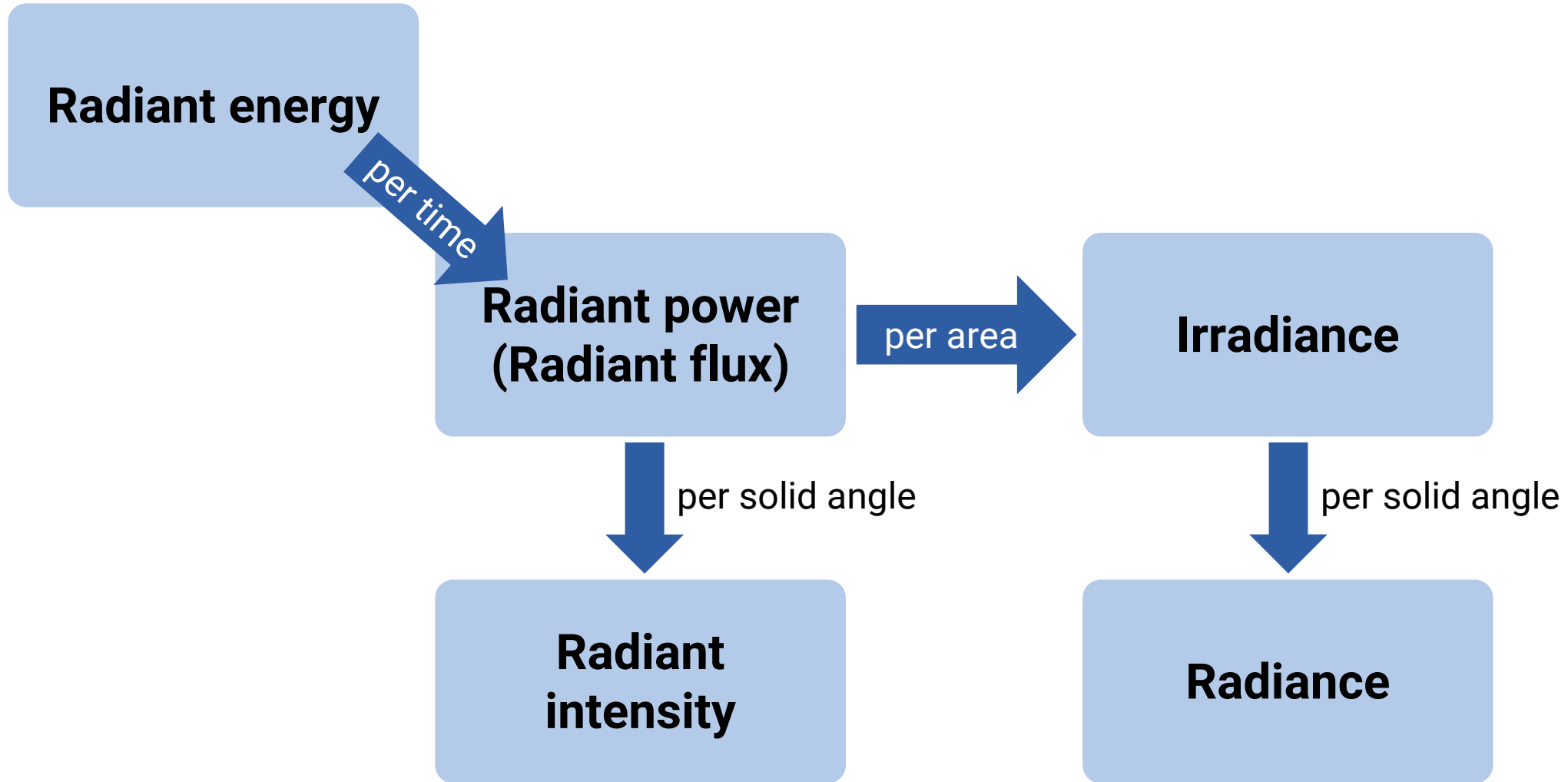


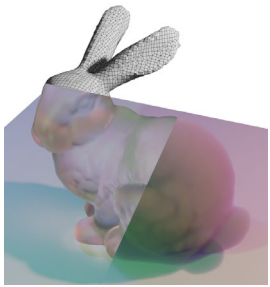
# Quiz & Preview

# Radiometric quantities



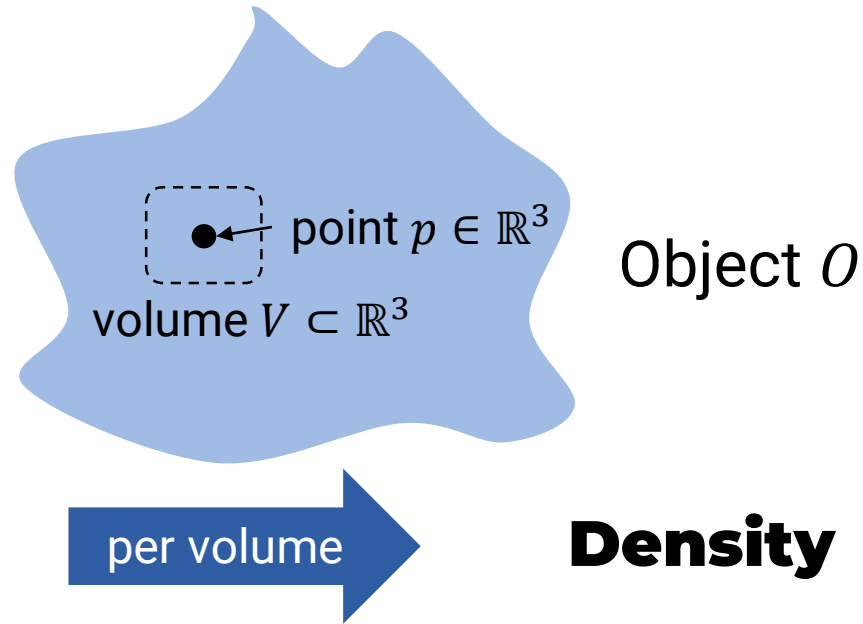
# Radiometric quantities





# Mass vs. Density

# Concepts of mass vs. density



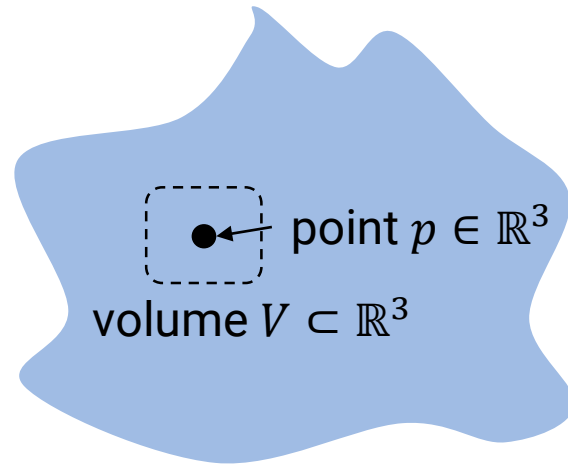
## Mass

- ✓ Mass of the object  $O$
- ✓ Mass of some region (volume)  $V$
- ✗ Mass at the point  $p$   
→ illegal or meaningless (always zero)

## Density

- ✗ Density of the object  $O$   
→ illegal or “average density” of the object  $O$
- ✗ Density of some region (volume)  $V$   
→ illegal or “average density” of the volume  $V$
- ✓ Density at the point  $p$

# Concepts of mass vs. density

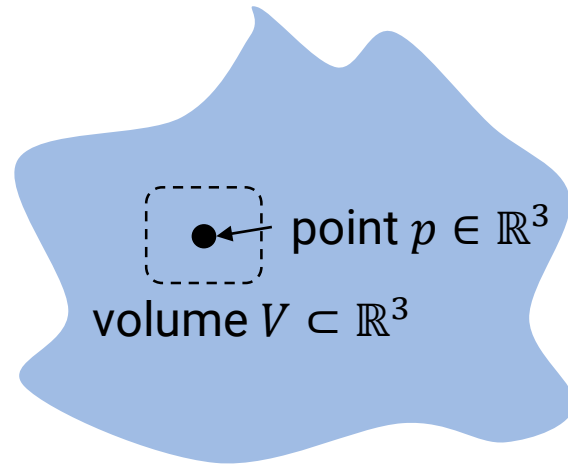


**Mass of  $V \subset \mathbb{R}^3$  *What?***

**Density of  $\mathbb{R}^3$  *what?***

- “mass of an object  $O$ ”  
= “mass of the volume of  $O$ ”

# Concepts of mass vs. density



**Mass of**  $V \subset \mathbb{R}^3$

**Density of**  $p \in \mathbb{R}^3$

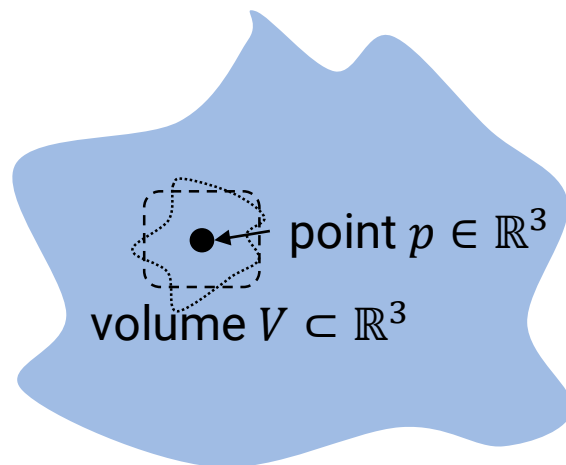
$\text{mass}(V)$

per volume →

$$\text{density}(p) = \lim_{\substack{\text{vol}(V) \rightarrow 0 \\ p \in V}} \frac{\text{mass}(V \text{ [blue box]})}{\text{vol}(V \text{ [blue box]})}$$



# Concepts of mass vs. density



**Mass of**  $V \subset \mathbb{R}^3$

**Density of**  $p \in \mathbb{R}^3$

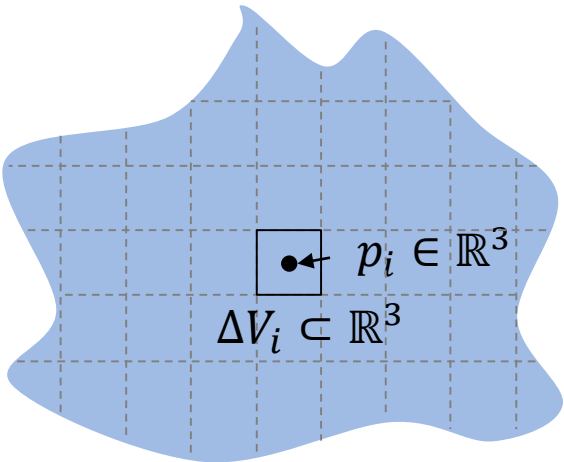
mass( $V$ )

per volume

$$\begin{aligned} \text{density}(p) &= \lim_{\substack{\text{vol}(V) \rightarrow 0 \\ p \in V}} \frac{\text{mass}(V \text{ [dashed box]})}{\text{vol}(V \text{ [dashed box]})} \\ &= \lim_{\substack{\text{vol}(V) \rightarrow 0 \\ p \in V}} \frac{\text{mass}(V \text{ [irregular shape]})}{\text{vol}(V \text{ [irregular shape]})} \end{aligned}$$

\* The limit converges to the same value whenever  $\text{vol}(V) \rightarrow 0$  and  $p \in V$ .

# Concepts of mass vs. density



**Mass of**  $V \subset \mathbb{R}^3$

**Density of**  $p \in \mathbb{R}^3$

→

$\text{vol}(\Delta V_i) \rightarrow 0$

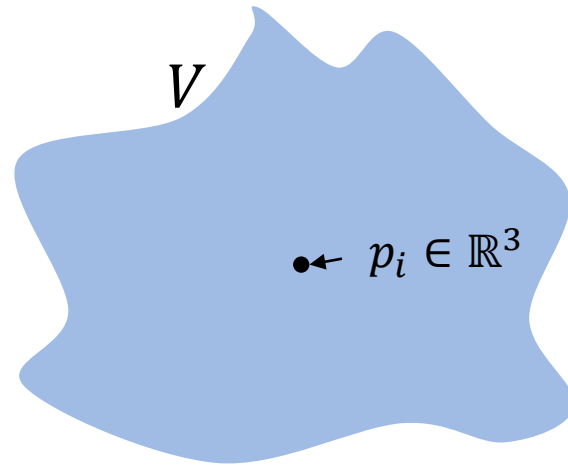
$$\text{mass}(V) = \sum_i \text{mass}(\Delta V_i)$$

$$\approx \sum_i \text{density}(p_i) \text{vol}(\Delta V_i)$$

← over volume

$$\text{density}(p) = \lim_{\substack{\text{vol}(V) \rightarrow 0 \\ p \in V}} \frac{\text{mass}(V)}{\text{vol}(V)}$$

# Concepts of mass vs. density



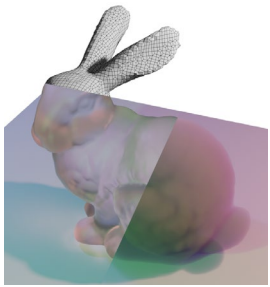
**Mass of**  $V \subset \mathbb{R}^3$  [ kg ]

**Density at**  $p \in \mathbb{R}^3$  [ kg/m<sup>3</sup> ]

## Proposition

Relations between mass  $m$  of a volume  $V \subset \mathbb{R}^3$  and density  $\rho$  at a point  $p \in \mathbb{R}^3$  is that:

$$m(V) = \int_V \rho(p) dp \quad \begin{array}{c} \xrightarrow{\text{per volume}} \\ \xleftarrow{\text{over volume}} \end{array} \quad \rho(p) = \lim_{\substack{|V| \rightarrow 0 \\ p \in V}} \frac{m(V)}{|V|}$$



# Solid Angles



We roughly say...

***“Solid angles” are  
3D versions of “angles”***

How strictly is this sentence correct?

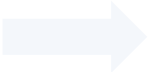
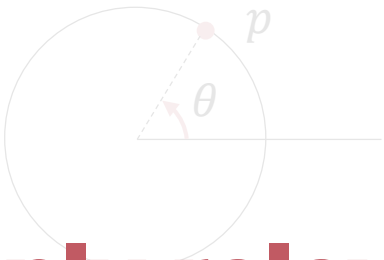
# Solid angles



Can following statements be converted to sphere ( $\mathbb{S}^2$ ) version using “solid angles”?

1. The position of a point  $p$  in the unit circle ( $\mathbb{S}^1$ ) can be represented as the angle  $\theta$

two coordinates



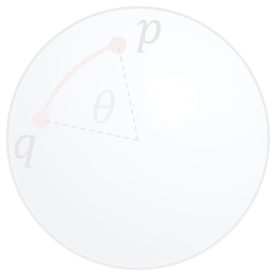
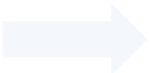
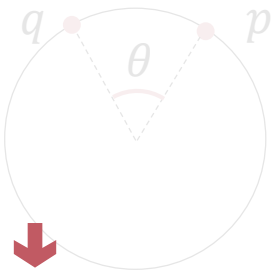
e.g.  $p = (\theta, \phi)$



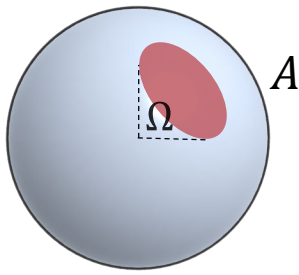
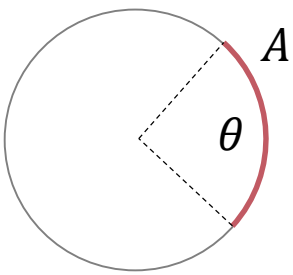
**“Solid angles” are only relevant to 3.**

2. How far apart two points  $p$  and  $q \in \mathbb{S}^1$  can be represented as the angle  $\theta$  between them.

angle  $\theta$



3. The size of a region  $A \subset \mathbb{S}^1$  can be represented as the angle  $\theta$  solid angle  $\Omega$

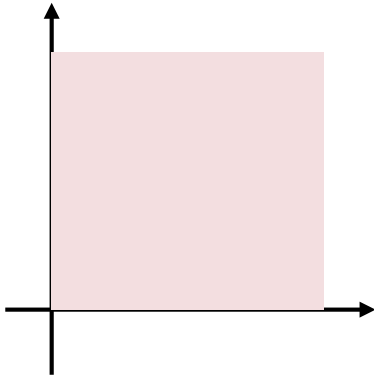




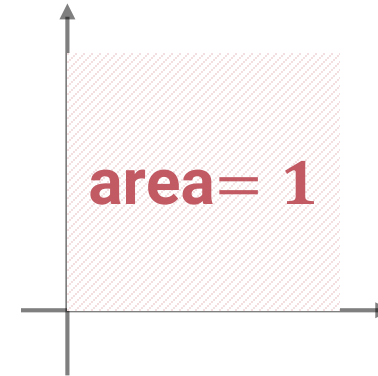
**In many times,  
several concepts can be treated  
as a single concept in lower dimensions,  
but they become different in higher dimensions**



**We call the both “area”.** (similarly for “volume”)



a 2-dimensional *subset*  
 $A = \{(x, y) \in \mathbb{R}^2 \mid 0 \leq x, y \leq 1\}$

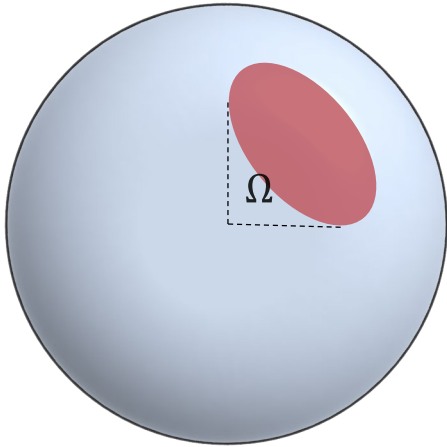


*measure* of a 2-dimensional subset  
 $|A| = 1$

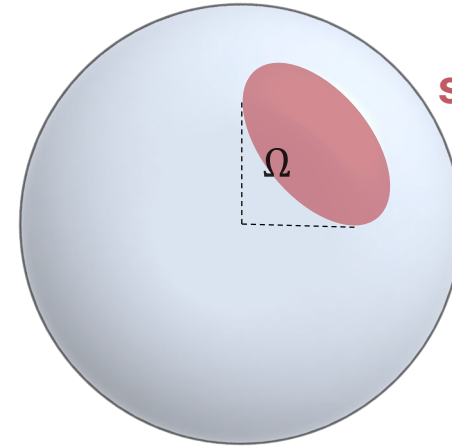




## Also for “solid angle”



a spherical *subset*  
 $\Omega = \{\hat{\omega} \in \mathbb{S}^2 | \hat{u} \cdot \hat{\omega} \geq 0.7\}$

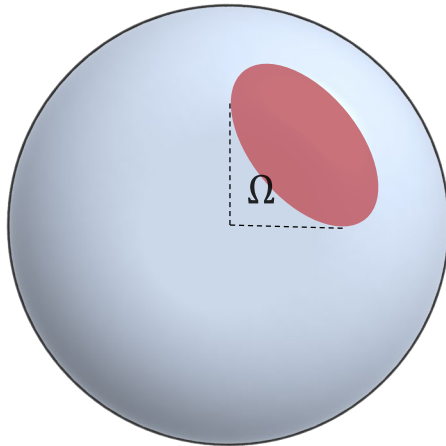


solid angle =  $0.2\pi$  sr

*measure* of a spherical subset  
 $|\Omega| = 1$



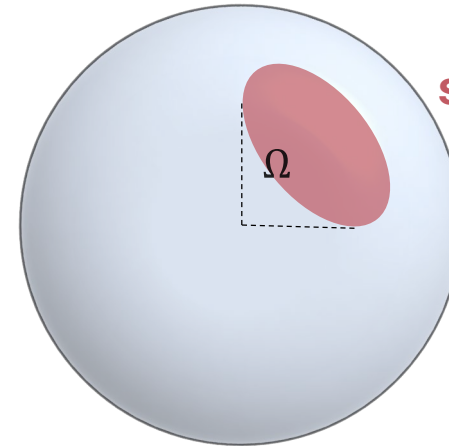
## (not common) terminology in this seminar



a spherical *subset*

$$\Omega = \{\hat{w} \in \mathbb{S}^2 | \hat{u} \cdot \hat{w} \geq 0.7\}$$

**solid angle *region***



**solid angle =  $0.2\pi$  sr**

*measure* of a spherical subset

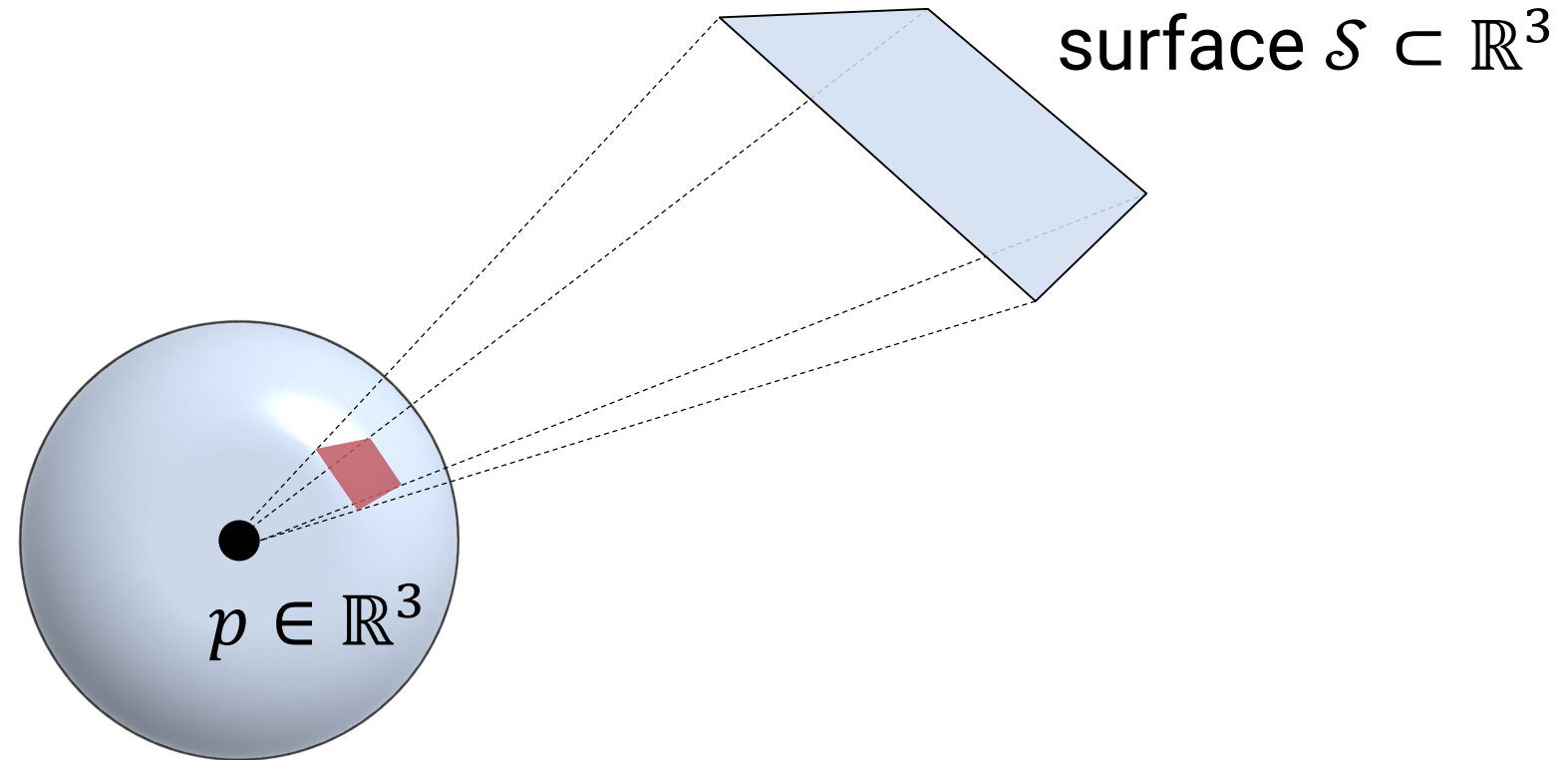
$$|\Omega| = 1$$

**solid angle *measure***

# Surface to solid angle



**How large  $\mathcal{S}$  appears to an observer at  $p$ ?**

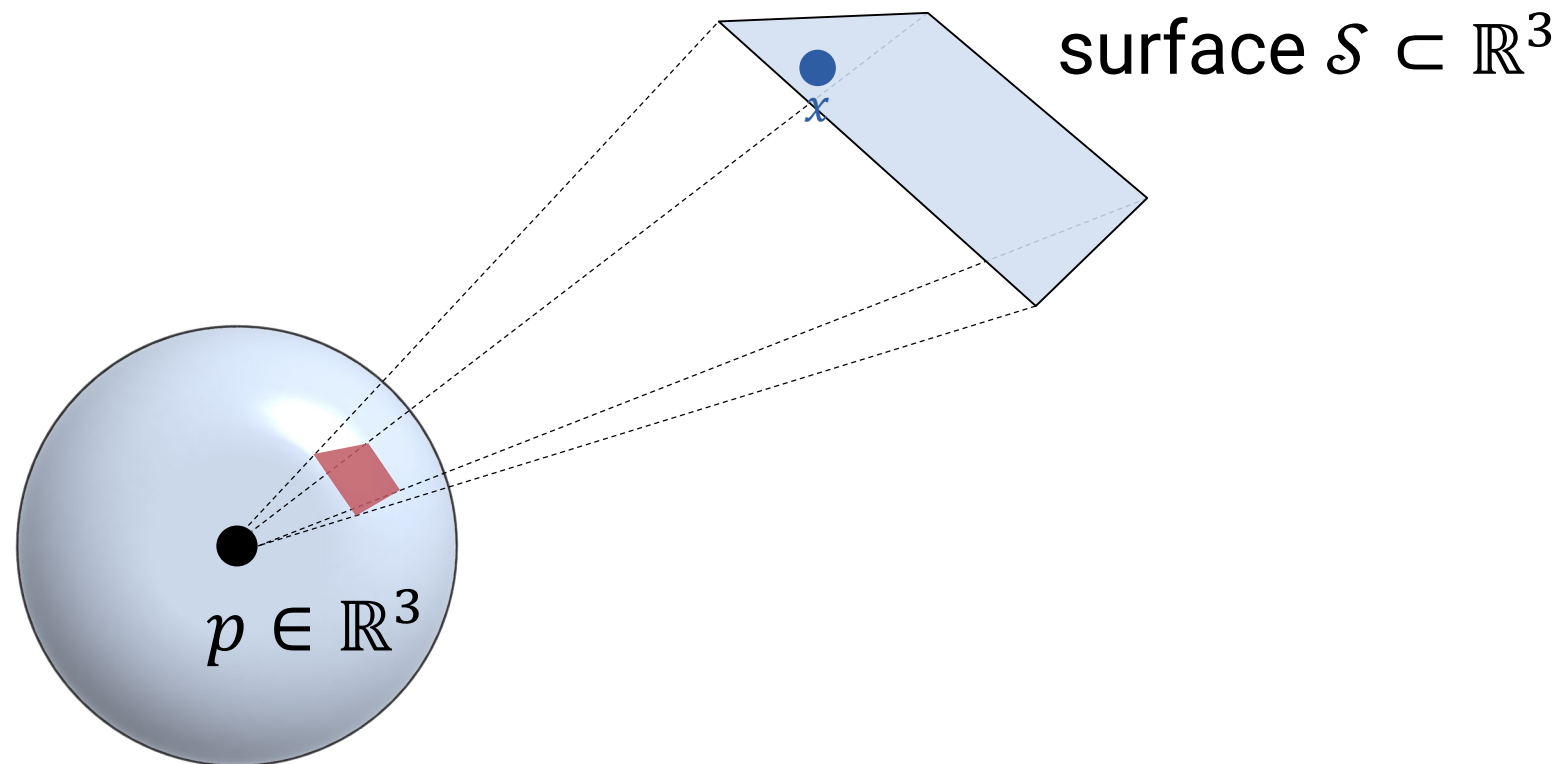


**Solid angle!**

# Surface to solid angle



**Try to write in an area integral on  $\mathcal{S}$**

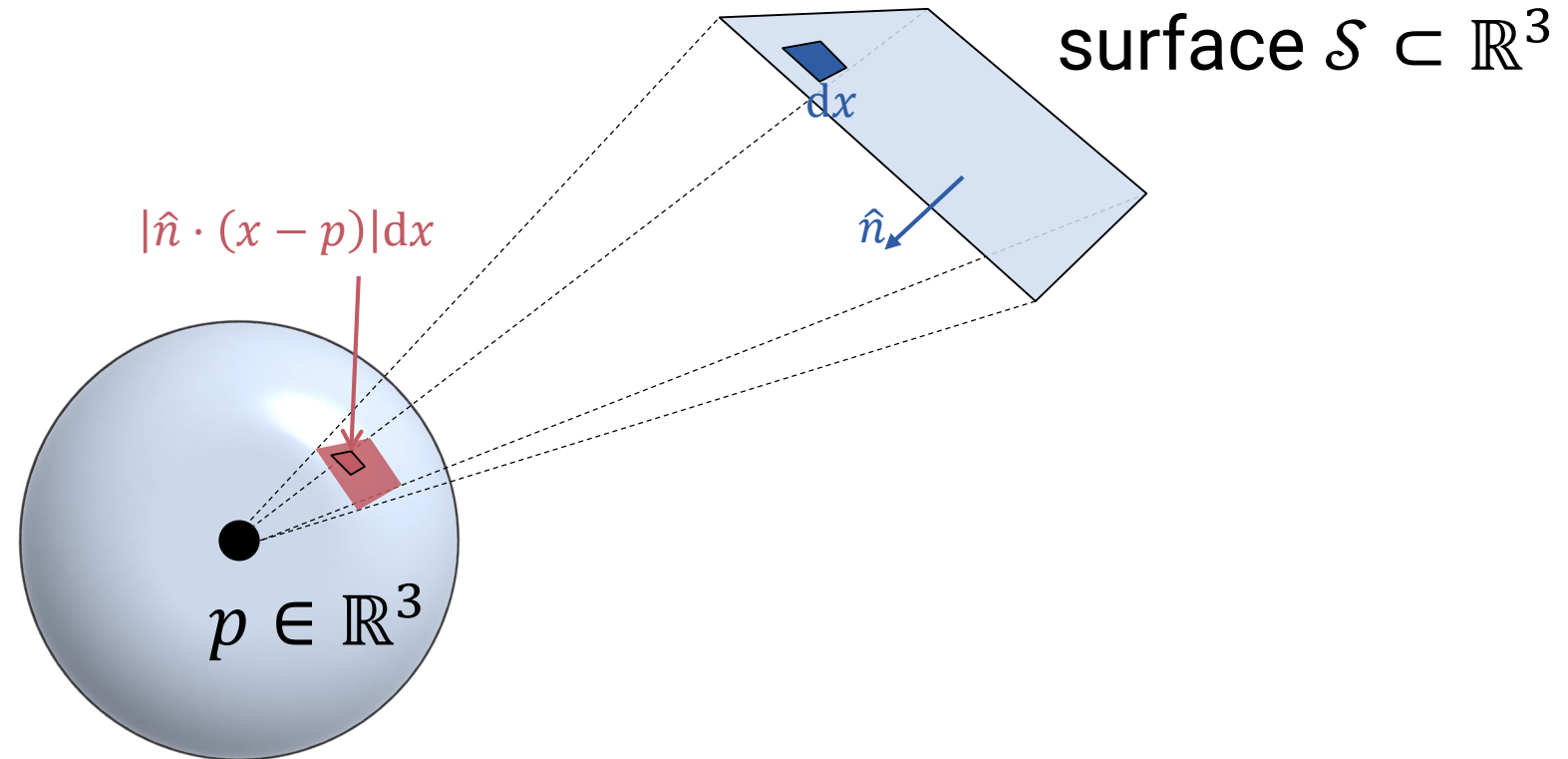


**Solid angle =**  $\int_{\mathcal{S}} ? \, dx$

# Surface to solid angle



**Try to write in an area integral on  $\mathcal{S}$**

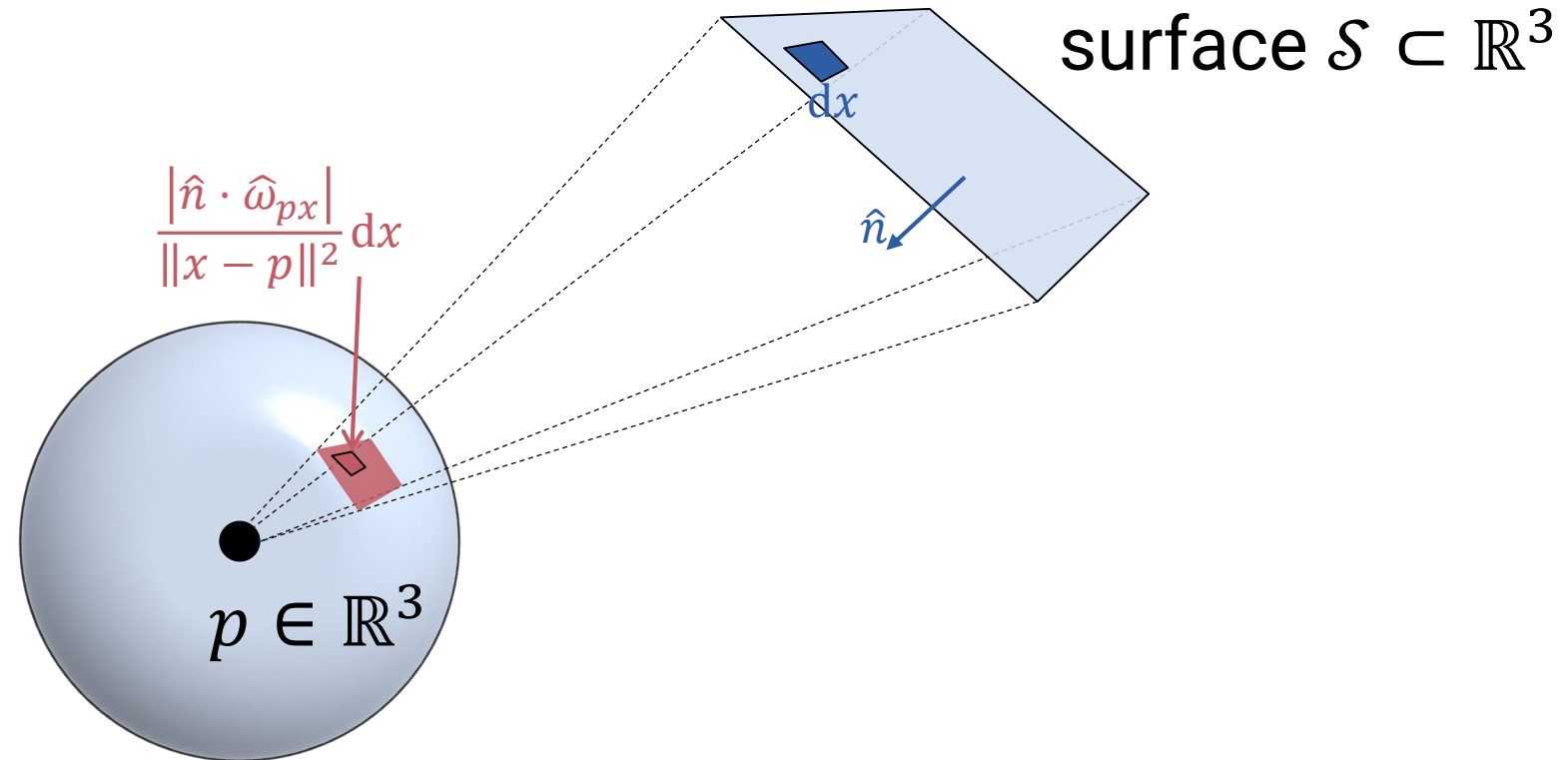


**Solid angle =**  $\int_{\mathcal{S}} ? \, dx$

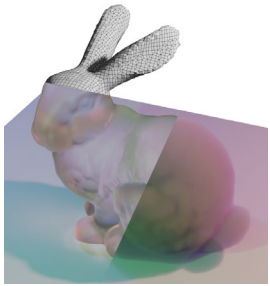
# Surface to solid angle



**Try to write in an area integral on  $\mathcal{S}$**

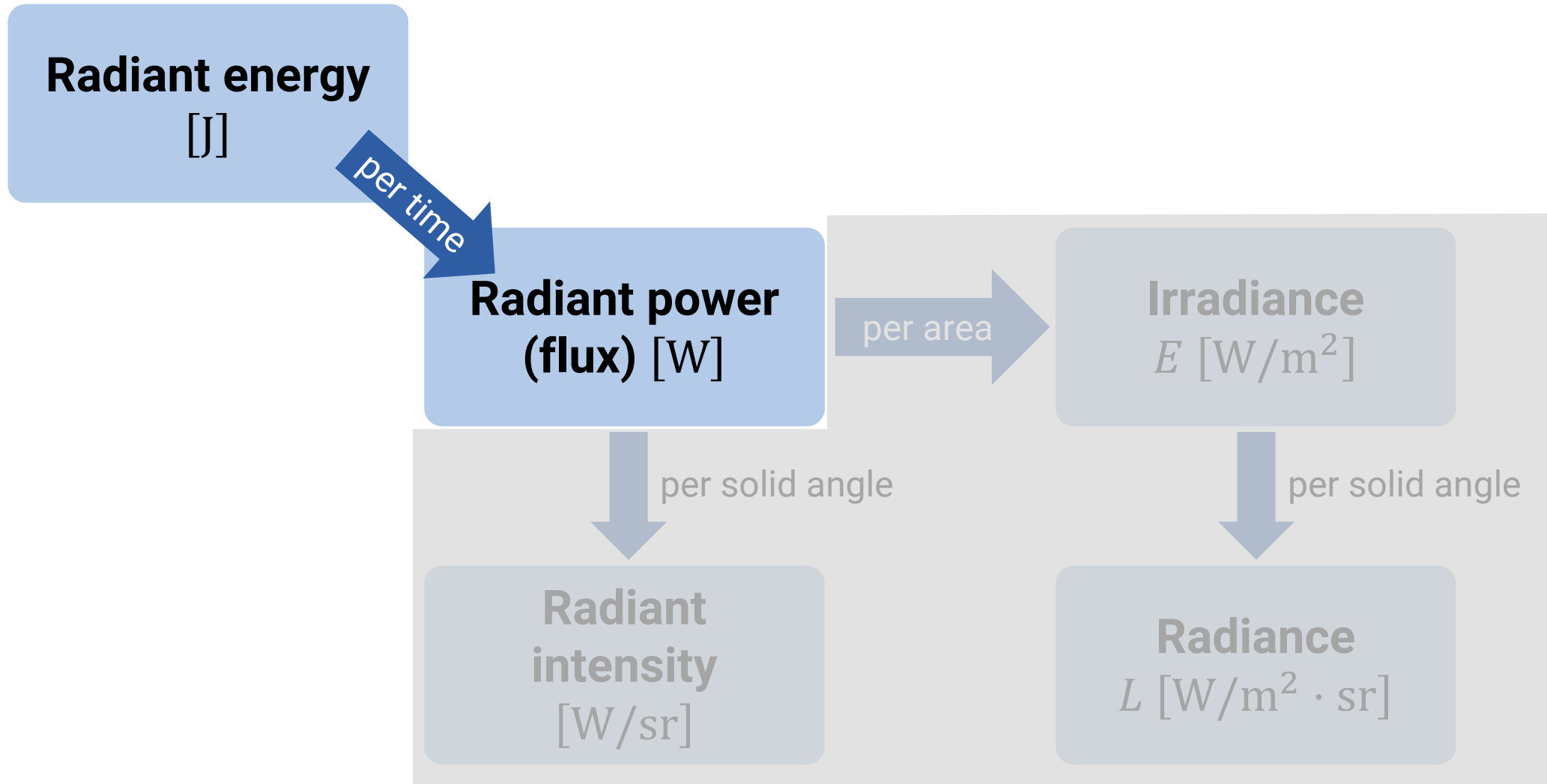


**Solid angle=** 
$$\int_{\mathcal{S}} \frac{|\hat{n} \cdot \hat{\omega}_{px}|}{\|x - p\|^2} dx$$

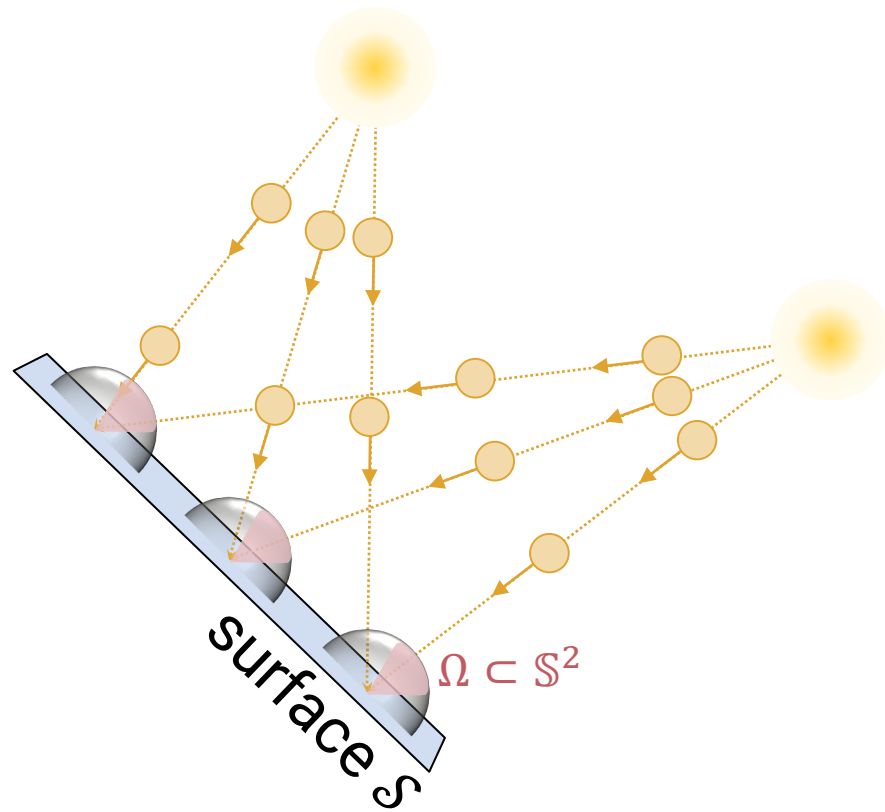


# Radiometric Quantities

# Radiometric quantities

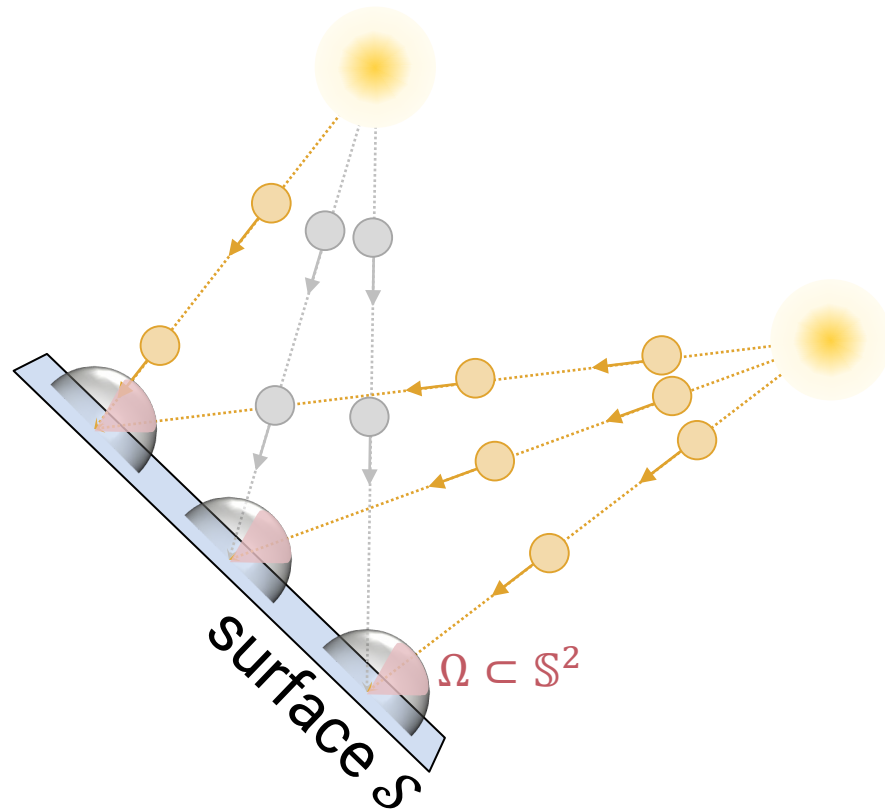






**Radiant energy of what?**  $\mathcal{S} \subset \mathbb{R}^3$ ,  
solid angle  $\Omega \subset \mathbb{S}^2$ ,

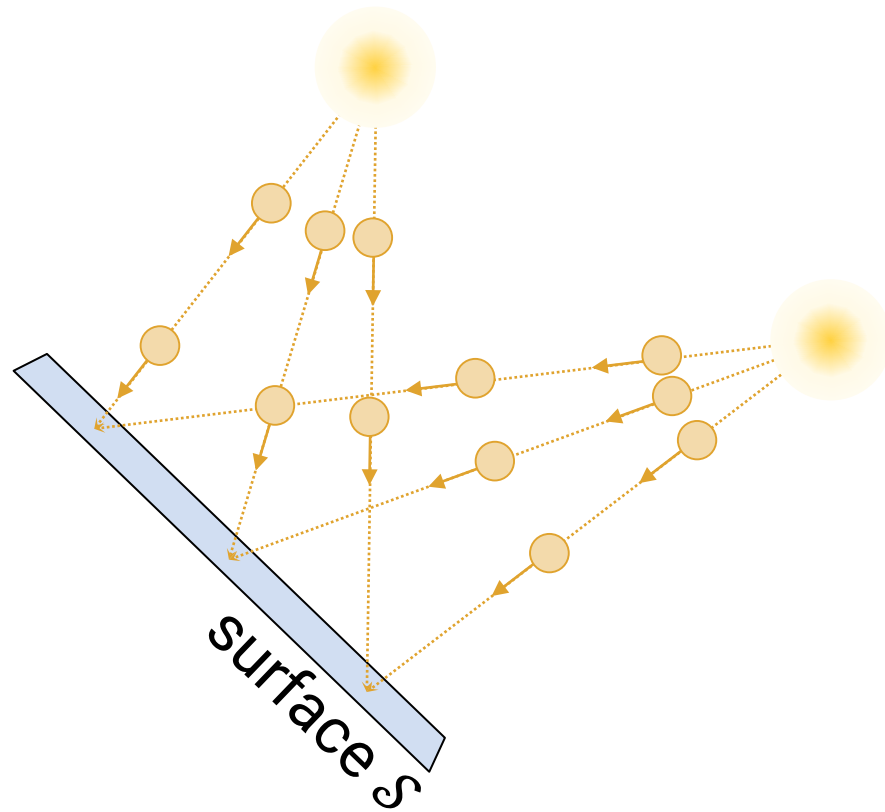
- “Energy” is “energy”!
- *Number of “hits” of photons on the surface*



**Radiant energy of** surface  $\mathcal{S} \subset \mathbb{R}^3$ ,  
solid angle  $\Omega \subset \mathbb{S}^2$ ,  
time interval  $[t_1, t_2] \subset \mathbb{R}$

- “Energy” is “energy”!
- *Number of “hits” of photons on the surface*

# Radiant energy

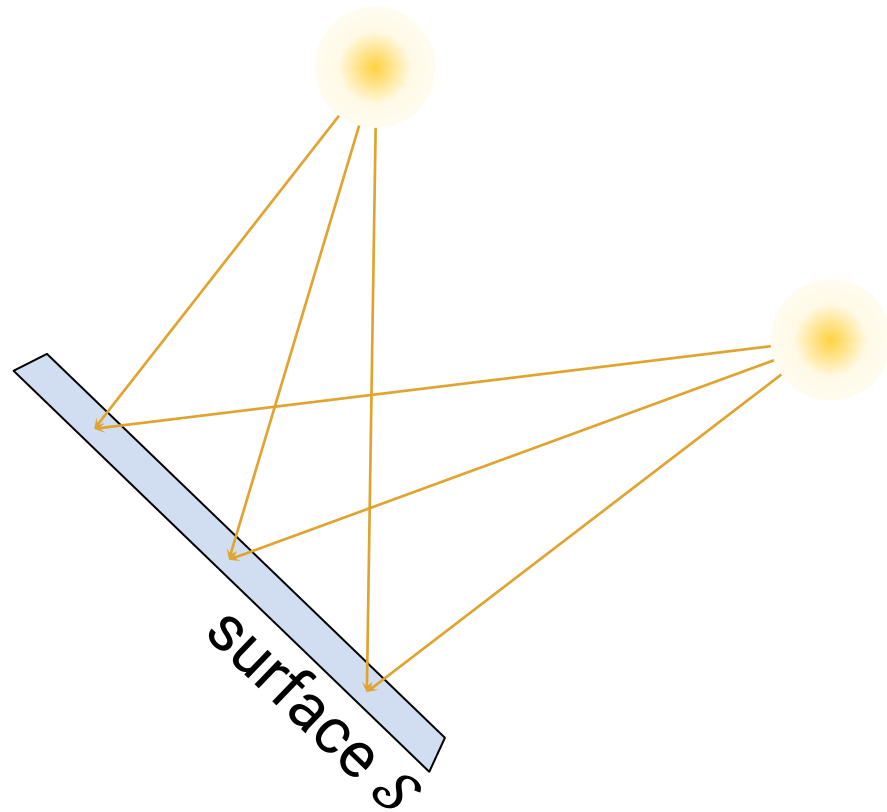


**Radiant energy of** surface  $\mathcal{S} \subset \mathbb{R}^3$ ,  
solid angle  $\Omega \overset{\text{default}}{\in} \mathbb{S}^2$ ,  
time interval  $[t_1, t_2] \subset \mathbb{R}$

$$Q(\mathcal{S}, \Omega, [t_1, t_2]) \quad \text{[1]}$$

- “Energy” is “energy”!
- *Number of “hits” of photons on the surface*

# Radiant flux (radiant power)

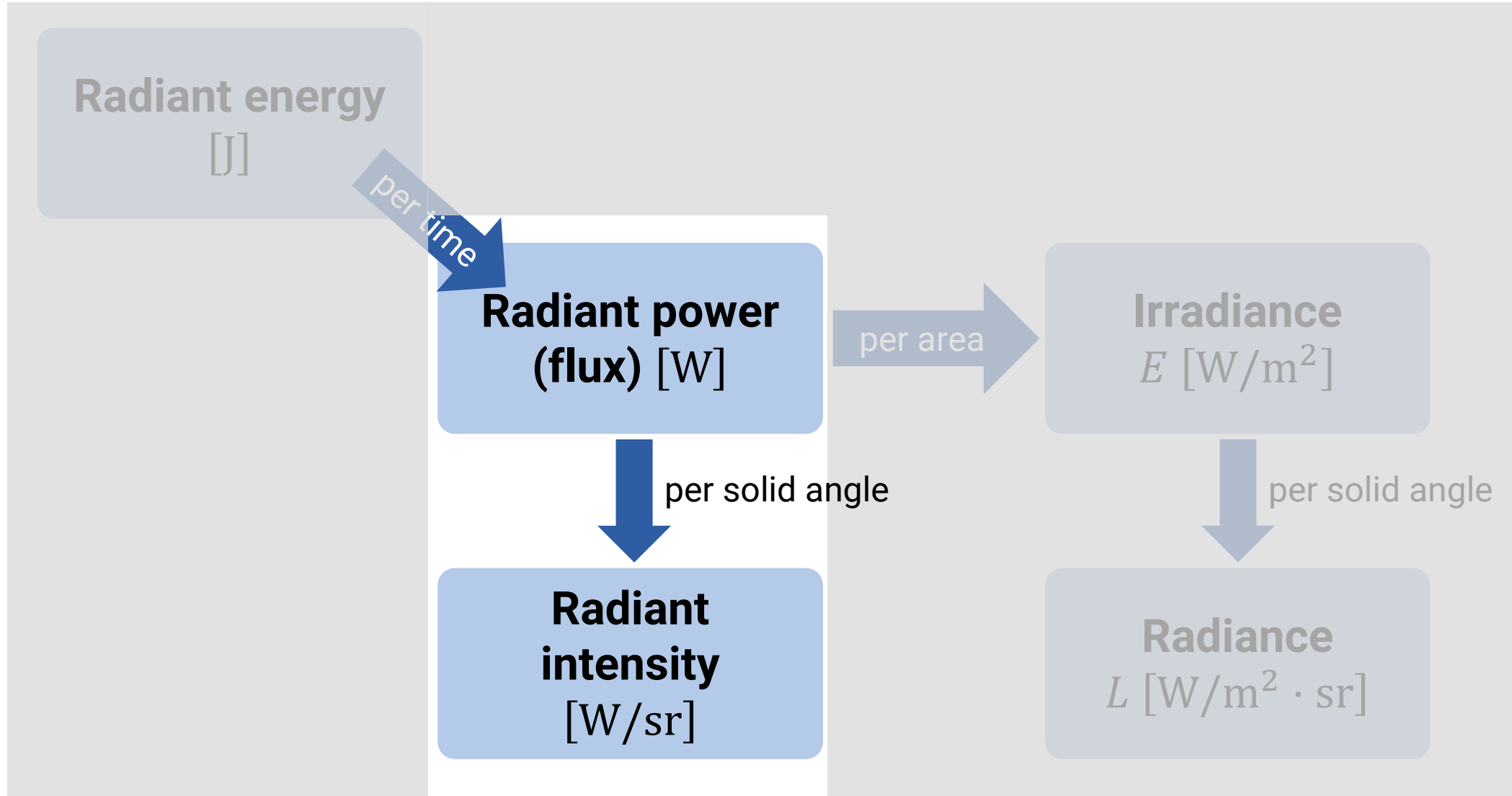


**Radiant flux of what?**  $\mathcal{S} \subset \mathbb{R}^3$ ,  
(Radiant power) solid angle  $\Omega \subset \mathbb{S}^2$   
time  $t \in \mathbb{R}$  (steady state)

$$\Phi(\mathcal{S}, \Omega, t) \quad [\text{J/s} = \text{W}]$$

- “Power” is “energy per time”!
- *Number of “intersecting rays” on the surface*

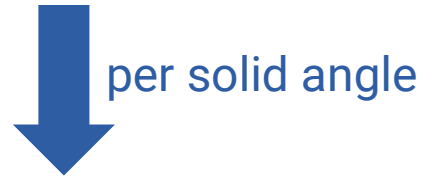
# Radiant intensity



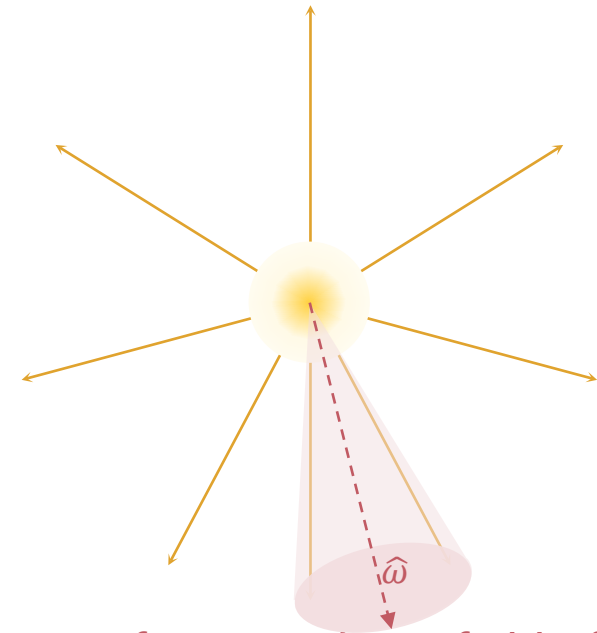
# Radiant intensity



**Radiant flux** of surface  $\mathcal{S} \subset \mathbb{R}^3$ ,  
solid angle  $\Omega \subset \mathbb{S}^2$



**Radiant intensity** of point source,  $\mathbb{R}^3$ , ?  
direction  $\hat{\omega}$



surface  $\mathcal{S}_\Omega$  whose field of view  
from the point source is  $\Omega \subset \mathbb{S}^2$

## Proposition

Relations between radiant flux  $\Phi$  of a surface  $\mathcal{S} \subset \mathbb{R}^3$  and  
radiant intensity of a point source at a position  $p \in \mathbb{R}^3$  is that:

$$\Phi(\mathcal{S}_\Omega) = \int_{\Omega} I(\hat{\omega}) d\hat{\omega}$$

per solid angle

over solid angle

$$I(\hat{\omega}) = \lim_{\substack{|\Omega| \rightarrow 0 \\ \hat{\omega} \in \Omega}} \frac{\Phi(\mathcal{S}_\Omega)}{|\Omega|},$$

solid angle measure

where  $\Omega$  is the solid angle region

# Radiant intensity

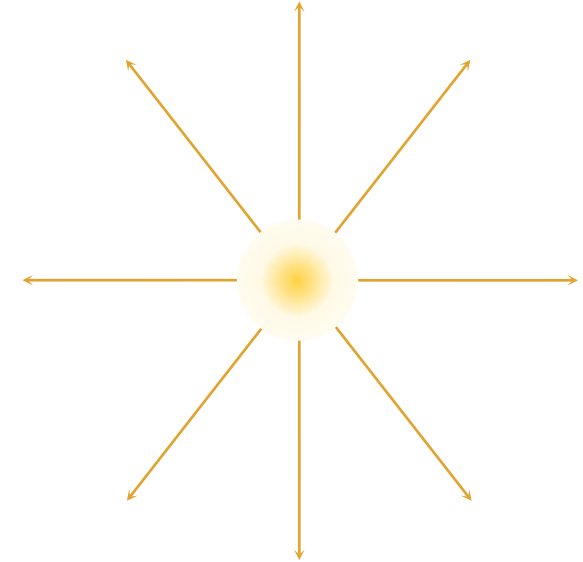


**Radiant flux** of surface  $\mathcal{S} \subset \mathbb{R}^3$ ,  
solid angle  $\Omega \subset \mathbb{S}^2$



per solid angle

**Radiant intensity** of point source,  
direction  $\hat{\omega}$



## Practice

There is an isotropic point light source with radiant flux  $\Phi$ .  
The radiant intensity of the source is?

$$I(\hat{\omega}) = \frac{\Phi}{4\pi}$$

# Radiant intensity



**Defined under any situation  
such as emission, incidence, reflection, refraction...**

**Radiant energy**  
[J]

per time

**Radiant power  
(flux) [W]**

per area

**Irradiance**  
 $E$  [W/m<sup>2</sup>]

per solid angle

**Radiant  
intensity**  
[W/sr]

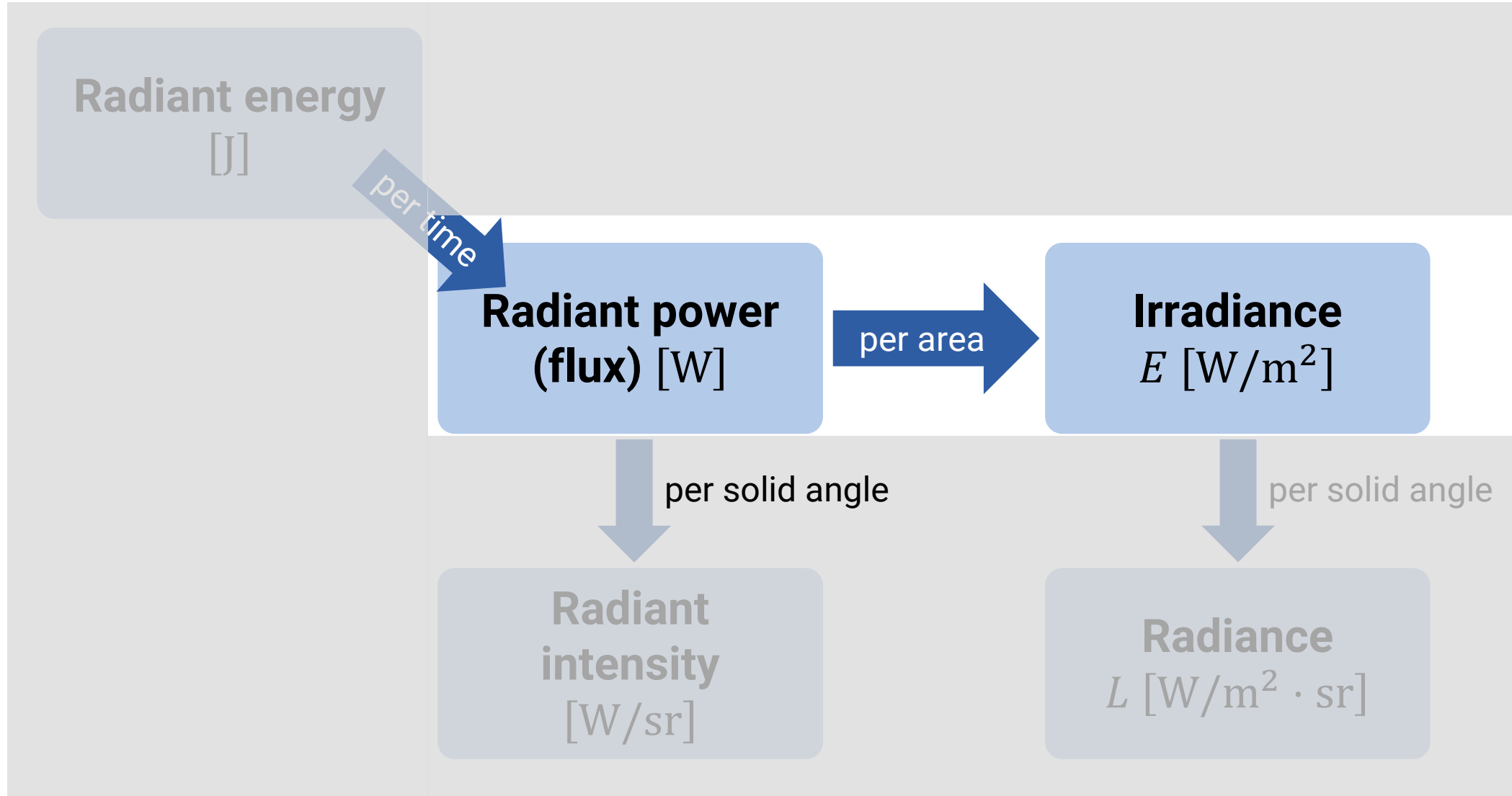
per solid angle

**Radiance**  
 $L$  [W/m<sup>2</sup> · sr]

**Defined on  
emission of point light source**



# Irradiance



# Irradiance

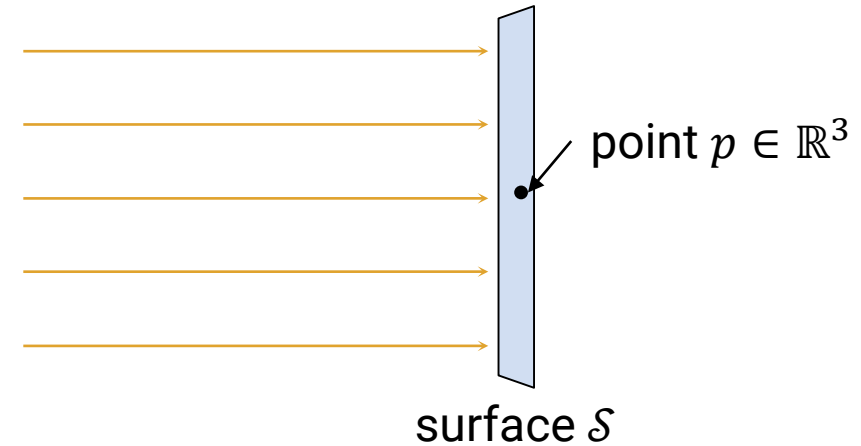


**Radiant flux** of surface  $\mathcal{S} \subset \mathbb{R}^3$ ,  
solid angle  $\Omega \subset \mathbb{S}^2$



per area

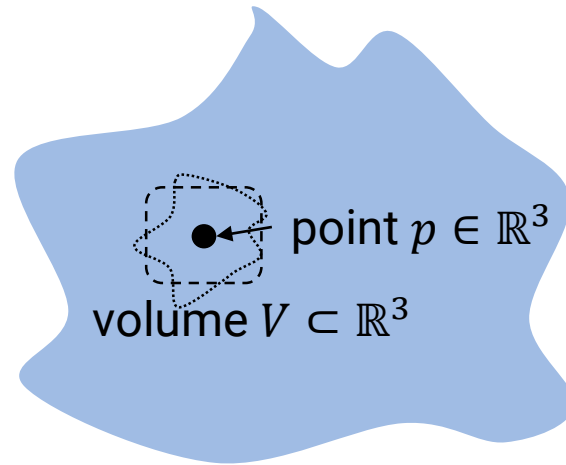
**irradiance** of point  $p \in \mathbb{R}^3$ , ?  
solid angle  $\Omega \subset \mathbb{S}^2$



$$? \quad E(p, \Omega) = \lim_{\substack{\text{area}(\mathcal{S}) \rightarrow 0 \\ p \in \mathcal{S}}} \frac{\Phi(\mathcal{S}, \Omega)}{\text{area}(\mathcal{S})}$$

**is it enough?**

# Review: Concepts of mass vs. density



**Mass of**  $V \subset \mathbb{R}^3$

**Density of**  $p \in \mathbb{R}^3$

mass( $V$ )

per volume

$$\begin{aligned} \text{density}(p) &= \lim_{\substack{\text{vol}(V) \rightarrow 0 \\ p \in V}} \frac{\text{mass}(V \text{ [dashed box]})}{\text{vol}(V \text{ [dashed box]})} \\ &= \lim_{\substack{\text{vol}(V) \rightarrow 0 \\ p \in V}} \frac{\text{mass}(V \text{ [irregular shape]})}{\text{vol}(V \text{ [irregular shape]})} \end{aligned}$$

\* The limit converges to the same value whenever  $\text{vol}(V) \rightarrow 0$  and  $p \in V$ .

# Irradiance

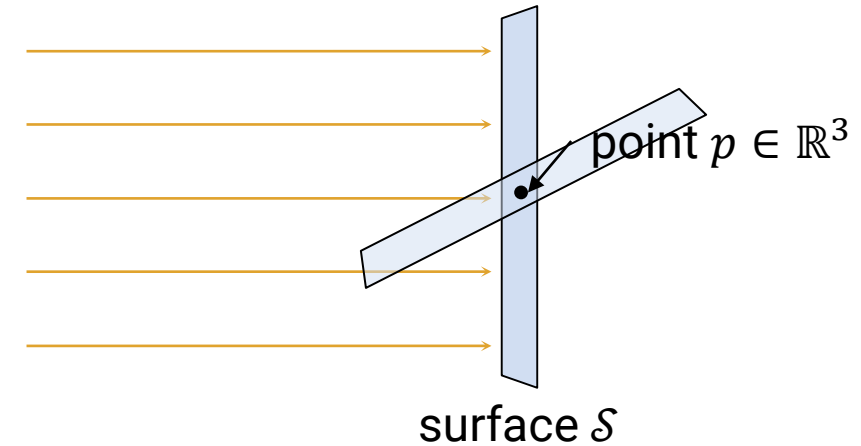


**Radiant flux** of surface  $\mathcal{S} \subset \mathbb{R}^3$ ,  
solid angle  $\Omega \subset \mathbb{S}^2$



per area

**irradiance** of point  $p \in \mathbb{R}^3$ , ?  
solid angle  $\Omega \subset \mathbb{S}^2$



? 
$$E(p, \Omega) = \lim_{\substack{\text{area}(\mathcal{S}) \rightarrow 0 \\ p \in \mathcal{S}}} \frac{\Phi(\mathcal{S}, \Omega)}{\text{area}(\mathcal{S})}$$

**is it enough?**

$\mathcal{S} = ( \text{vertical rectangle} )$  and  $\mathcal{S} = ( \text{tilted rectangle} )$  yield different limits

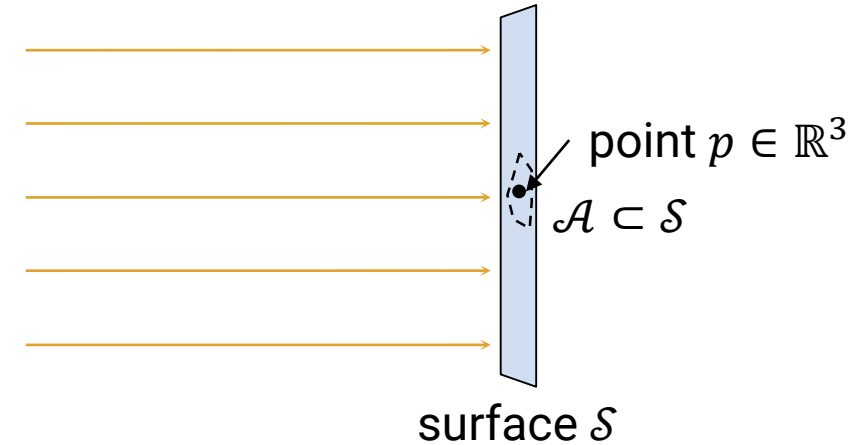
# Irradiance



**Radiant flux** of surface  $\mathcal{S} \subset \mathbb{R}^3$ ,  
solid angle  $\Omega \subset \mathbb{S}^2$



**irradiance** of point  $p \in \mathbb{R}^3$ , ?  
solid angle  $\Omega \subset \mathbb{S}^2$



$$E_{\mathcal{S}}(p, \Omega) = \lim_{\substack{\text{area}(\mathcal{A}) \rightarrow 0 \\ p \in \mathcal{A} \subset \mathcal{S}}} \frac{\Phi(\mathcal{S}, \mathcal{A}, \Omega)}{\text{area}(\mathcal{A})}$$

Irradiance defined as the limit about a subset of given fixed surface

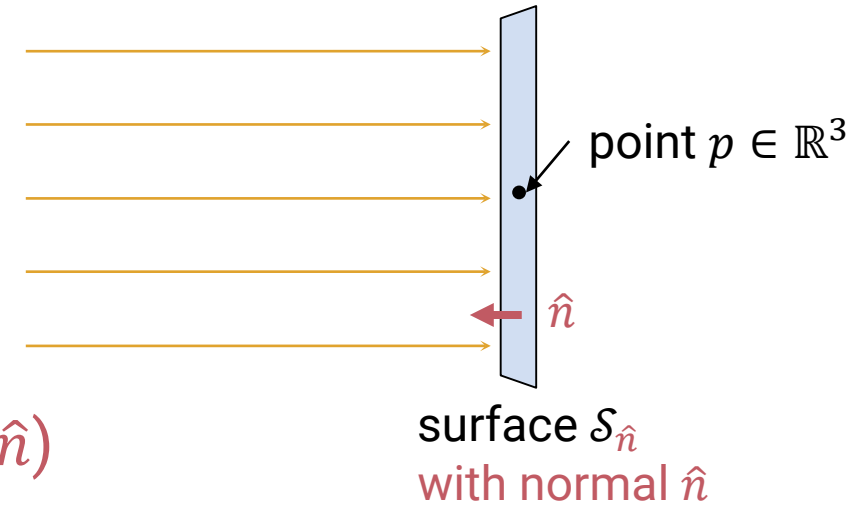
# Irradiance



**Radiant flux** of surface  $\mathcal{S} \subset \mathbb{R}^3$ ,  
solid angle  $\Omega \subset \mathbb{S}^2$



**irradiance** of point  $p \in \mathbb{R}^3$  (or  $p \in \mathbb{R}^3$  and  $\hat{n}$ )  
solid angle  $\Omega \subset \mathbb{S}^2$



$$E_{\hat{n}}(p, \Omega) = \lim_{\substack{\text{area}(\mathcal{S}_{\hat{n}}) \rightarrow 0 \\ p \in \mathcal{S}_{\hat{n}}}} \frac{\Phi(\mathcal{S}_{\hat{n}}, \Omega)}{\text{area}(\mathcal{S}_{\hat{n}})}$$

Irradiance defined as the limit about a subset of given fixed surface,  
or a surface with given fixed normal

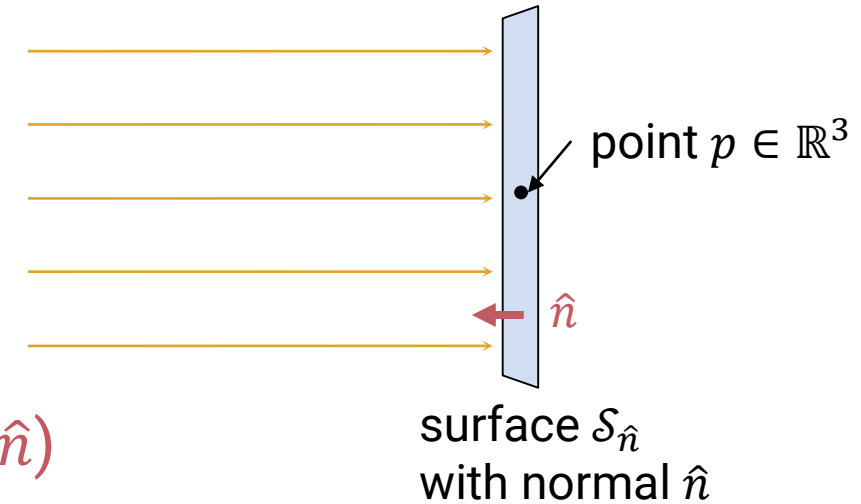
# Irradiance



**Radiant flux** of surface  $\mathcal{S} \subset \mathbb{R}^3$ ,  
solid angle  $\Omega \subset \mathbb{S}^2$



**irradiance** of point  $p \in \mathcal{S}$  (or  $p \in \mathbb{R}^3$  and  $\hat{n}$ )  
solid angle  $\Omega \subset \mathbb{S}^2$



$$E_{\hat{n}}(p, \Omega) = \lim_{\substack{\text{area}(\mathcal{S}_{\hat{n}}) \rightarrow 0 \\ p \in \mathcal{S}_{\hat{n}}}} \frac{\Phi(\mathcal{S}_{\hat{n}}, \Omega)}{\text{area}(\mathcal{S}_{\hat{n}})}$$

Irradiance defined as the limit about a subset of given fixed surface,  
or a surface with given fixed normal

We don't say just "irradiance at  $p$ " when  $p$  is not on any surface

# Irradiance

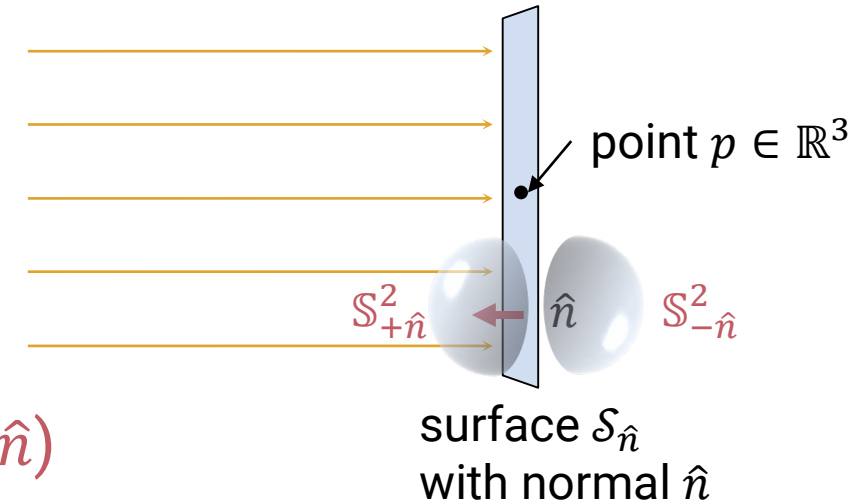


**Radiant flux** of surface  $\mathcal{S} \subset \mathbb{R}^3$ ,  
solid angle  $\Omega \subset \mathbb{S}^2$



per area

**irradiance** of point  $p \in \mathcal{S}$  (or  $p \in \mathbb{R}^3$  and  $\hat{n}$ )  
default: solid angle  $\Omega = \mathbb{S}^2, \mathbb{S}_{+\hat{n}}^2$ , or  $\mathbb{S}_{-\hat{n}}^2$



$$E_{\hat{n}}(p, \Omega) = \lim_{\substack{\text{area}(\mathcal{S}_{\hat{n}}) \rightarrow 0 \\ p \in \mathcal{S}_{\hat{n}}}} \frac{\Phi(\mathcal{S}_{\hat{n}}, \Omega)}{\text{area}(\mathcal{S}_{\hat{n}})}$$

The default solid angle changes depending on the context

***How?***



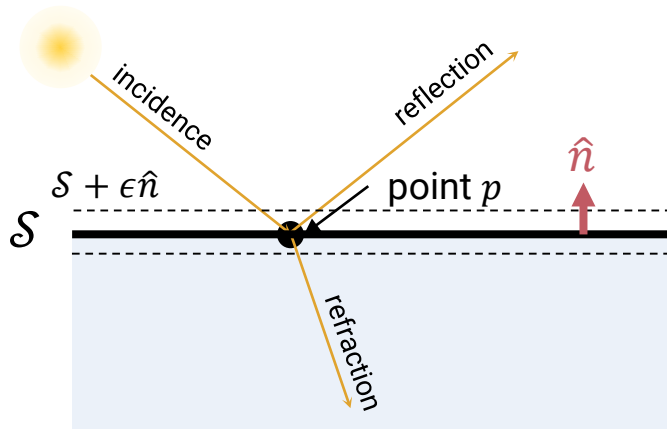


**Radiant flux** **of** surface  $\mathcal{S} \subset \mathbb{R}^3$ ,  
solid angle  $\Omega \subset \mathbb{S}^2$



**irradiance** **of** point  $p \in \mathcal{S}$  (or  $p \in \mathbb{R}^3$  and  $\hat{n}$ )  
default: solid angle  $\Omega = \mathbb{S}^2, \mathbb{S}_{+\hat{n}}^2$ , or  $\mathbb{S}_{-\hat{n}}^2$

Some details depending on context...



- “incoming” irradiance:  $E_{\mathcal{S}}^{(\text{in})}(p) = E_{\mathcal{S}+\epsilon\hat{n}}(p, \mathbb{S}_{-\hat{n}}^2)$
- “reflected” irradiance:  $E_{\mathcal{S}}^{(\text{refl})}(p) = E_{\mathcal{S}+\epsilon\hat{n}}(p, \mathbb{S}_{+\hat{n}}^2)$
- “refracted” irradiance:  $E_{\mathcal{S}}^{(\text{refr})}(p) = E_{\mathcal{S}-\epsilon\hat{n}}(p, \mathbb{S}_{-\hat{n}}^2)$
- “outgoing” irradiance:  $E_{\mathcal{S}}^{(\text{out})}(p) = E_{\mathcal{S}}^{(\text{refl})}(p) + E_{\mathcal{S}}^{(\text{refr})}(p)$

**Our intuition easily can do this!**

# Irradiance

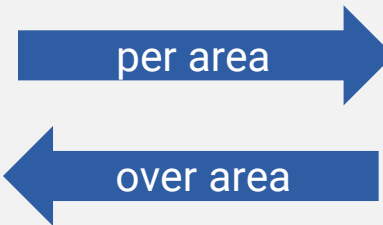


**Radiant flux** of surface  $\mathcal{S} \subset \mathbb{R}^3$ ,  
solid angle  $\Omega \subset \mathbb{S}^2$



**irradiance** of point  $p \in \mathcal{S}$  (or  $p \in \mathbb{R}^3$  and  $\hat{n}$ )  
solid angle  $\Omega \subset \mathbb{S}^2$

$$\Phi(\mathcal{S}, \Omega) = \int_{\mathcal{S}} E_{\mathcal{S}}(p, \Omega) dp$$



$$E_{\mathcal{S}}(p, \Omega) = \lim_{\substack{\text{area}(\mathcal{A}) \rightarrow 0 \\ p \in \mathcal{A} \subset \mathcal{S}}} \frac{\Phi(\mathcal{A}, \Omega)}{\text{area}(\mathcal{A})}$$

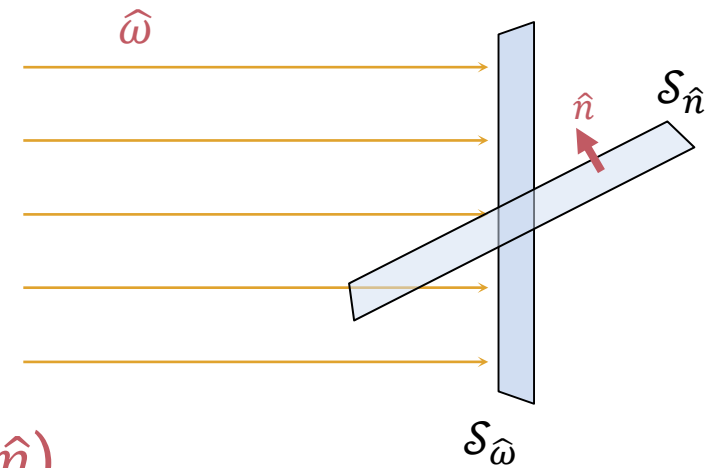
# Irradiance



**Radiant flux** of surface  $\mathcal{S} \subset \mathbb{R}^3$ ,  
solid angle  $\Omega \subset \mathbb{S}^2$



**irradiance** of point  $p \in \mathcal{S}$  (or  $p \in \mathbb{R}^3$  and  $\hat{n}$ )  
solid angle  $\Omega \subset \mathbb{S}^2$



## Practice

There is directional light with  $\hat{\omega}$ .

What is the relationship between  $E_{\hat{\omega}}(p)$  and  $E_{\hat{n}}(p)$ ?

$$E_{\hat{n}}(p) = E_{\hat{\omega}}(p) |\hat{n} \cdot \hat{\omega}|$$

# Irradiance

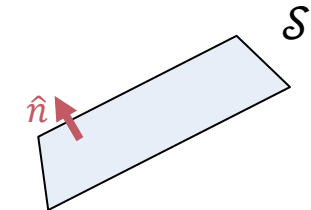
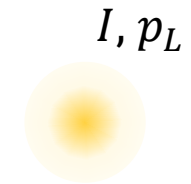


**Radiant flux** of surface  $\mathcal{S} \subset \mathbb{R}^3$ ,  
solid angle  $\Omega \subset \mathbb{S}^2$



per area

**irradiance** of point  $p \in \mathcal{S}$  (or  $p \in \mathbb{R}^3$  and  $\hat{n}$ )  
solid angle  $\Omega \subset \mathbb{S}^2$



## Practice

There is a point light source at  $p_L$  with the radiant intensity  $I(\hat{\omega})$ .  
What is the incident irradiance at  $p$  on a surface  $\mathcal{S}$ ,  $E_{\mathcal{S}}(p) = ?$

$$E_{\mathcal{S}}(p) = \lim_{\substack{\text{area}(\mathcal{A}) \rightarrow 0 \\ p \in \mathcal{A} \subset \mathcal{S}}} \frac{\Phi(\mathcal{A})}{\text{area}(\mathcal{A})} = \lim_{\substack{\text{area}(\mathcal{A}) \rightarrow 0 \\ p \in \mathcal{A} \subset \mathcal{S}}} \frac{I(\hat{\omega}_{p_L p}) \text{sol. ang.}(\mathcal{A}, p_L)}{\text{area}(\mathcal{A})} = \frac{I(\hat{\omega}_{p_L p}) |\hat{n} \cdot \hat{\omega}_{p_L p}|}{\|p - p_L\|^2}$$

Definition of irradiance

Definition of radiant intensity, small areal  $\mathcal{A}$

Relation between an area and a solid angle, small areal  $\mathcal{A}$

# Radiance



**Radiant energy**  
[J]

per time

**Radiant power  
(flux)** [W]

per area

**Irradiance**  
 $E$  [W/m<sup>2</sup>]

per solid angle

**Radiant  
intensity**  
[W/sr]

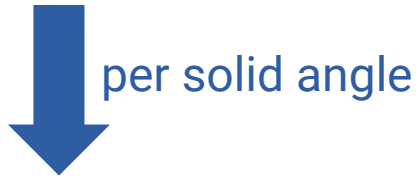
per solid angle

**Radiance**  
 $L$  [W/m<sup>2</sup> · sr]

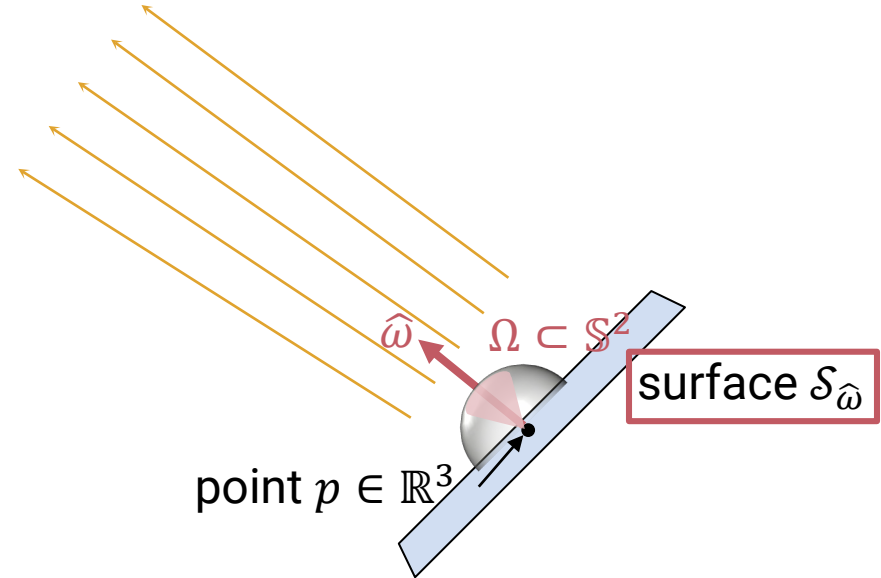
# Radiance



**Irradiance** of point  $p \in \mathcal{S}$ ,  
solid angle  $\Omega \subset \mathbb{S}^2$



**Radiance** of point  $p \in \mathbb{R}^3$ ,  
direction  $\hat{\omega} \in \mathbb{S}^2$



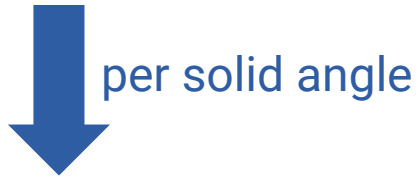
$$L(p, \hat{\omega}) = \lim_{\substack{\text{sol.ang.}(\Omega) \rightarrow 0 \\ \hat{\omega} \in \Omega}} \frac{E_{\mathcal{S}_{\hat{\omega}}}(p, \Omega)}{\text{sol.ang.}(\Omega)}$$

For a point  $p \in \mathbb{R}^3$  (on or not on a surface),  
the radiance  $L(p, \hat{\omega})$  is defined as the limit about a **virtual surface facing  $\hat{\omega}$**

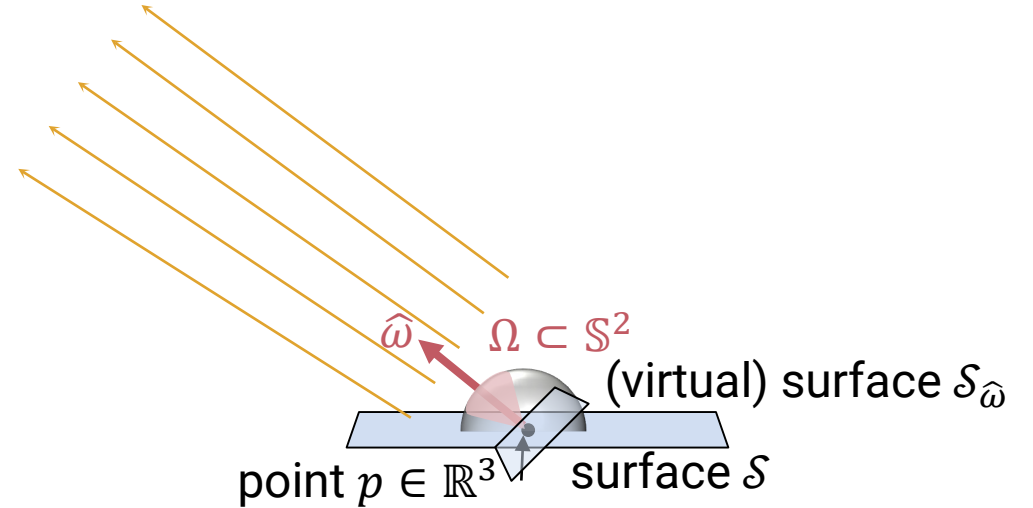
# Radiance



**Irradiance** of point  $p \in \mathcal{S}$ ,  
solid angle  $\Omega \subset \mathbb{S}^2$



**Radiance** of point  $p \in \mathbb{R}^3$ ,  
direction  $\hat{\omega} \in \mathbb{S}^2$



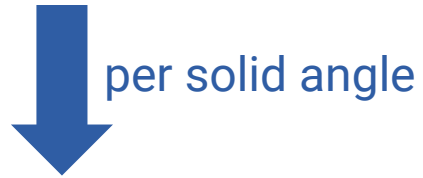
$$L(p, \hat{\omega}) = \lim_{\substack{\text{sol.ang.}(\Omega) \rightarrow 0 \\ \hat{\omega} \in \Omega}} \frac{\boxed{E_{\mathcal{S}_{\hat{\omega}}}(p, \Omega)}}{\text{sol. ang.}(\Omega)} \stackrel{\text{small solid angle around } \hat{\omega}!}{=} \lim_{\substack{\text{sol.ang.}(\Omega) \rightarrow 0 \\ \hat{\omega} \in \Omega}} \frac{\boxed{\frac{1}{|\hat{n} \cdot \hat{\omega}|}} \boxed{E_{\mathcal{S}}(p, \Omega)}}{\text{sol. ang.}(\Omega)}$$

For a point  $p$  on given surface  $\mathcal{S}$ ,  
the radiance  $L(p, \hat{\omega})$  can also be written as the limit about the irradiance on  $\mathcal{S}$

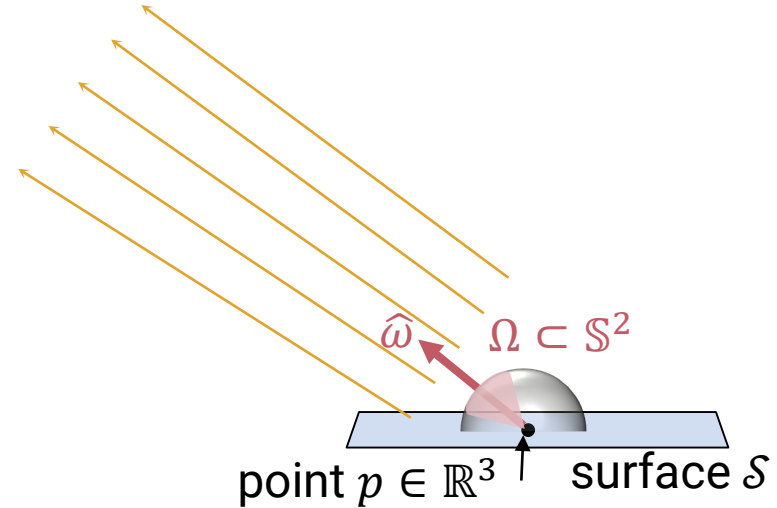
# Radiance



**Irradiance** of point  $p \in \mathcal{S}$ ,  
solid angle  $\Omega \subset \mathbb{S}^2$



**Radiance** of point  $p \in \mathbb{R}^3$ ,  
direction  $\hat{\omega} \in \mathbb{S}^2$



$$E_{\mathcal{S}}(p, \Omega) = \int_{\Omega} L(p, \hat{\omega}) |\hat{n} \cdot \hat{\omega}| d\hat{\omega}$$

per solid angle

over solid angle

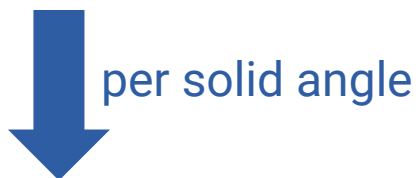
$$L(p, \hat{\omega}) = \lim_{\substack{\text{sol. ang.}(\Omega) \rightarrow 0 \\ \hat{\omega} \in \Omega}} \frac{1}{|\hat{n} \cdot \hat{\omega}|} \frac{E_{\mathcal{S}}(p, \Omega)}{\text{sol. ang.}(\Omega)}$$



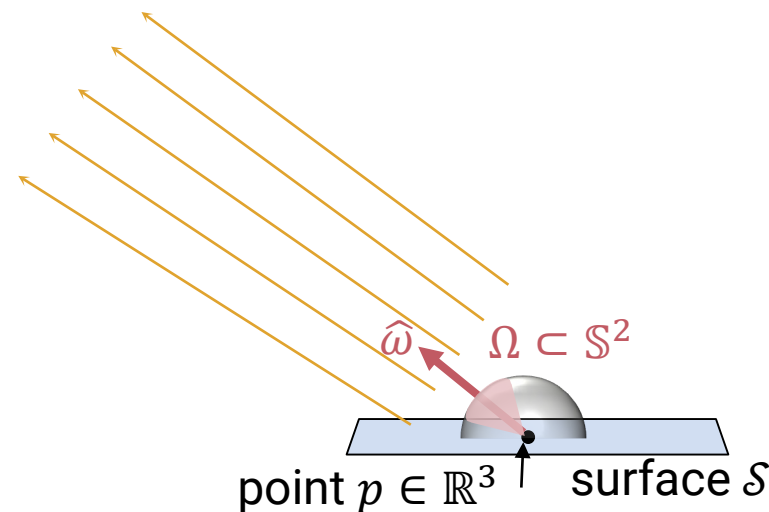
# Radiance



**Irradiance** of point  $p \in \mathcal{S}$ ,  
solid angle  $\Omega \subset \mathbb{S}^2$



**Radiance** of point  $p \in \mathbb{R}^3$ ,  
direction  $\hat{\omega} \in \mathbb{S}^2$



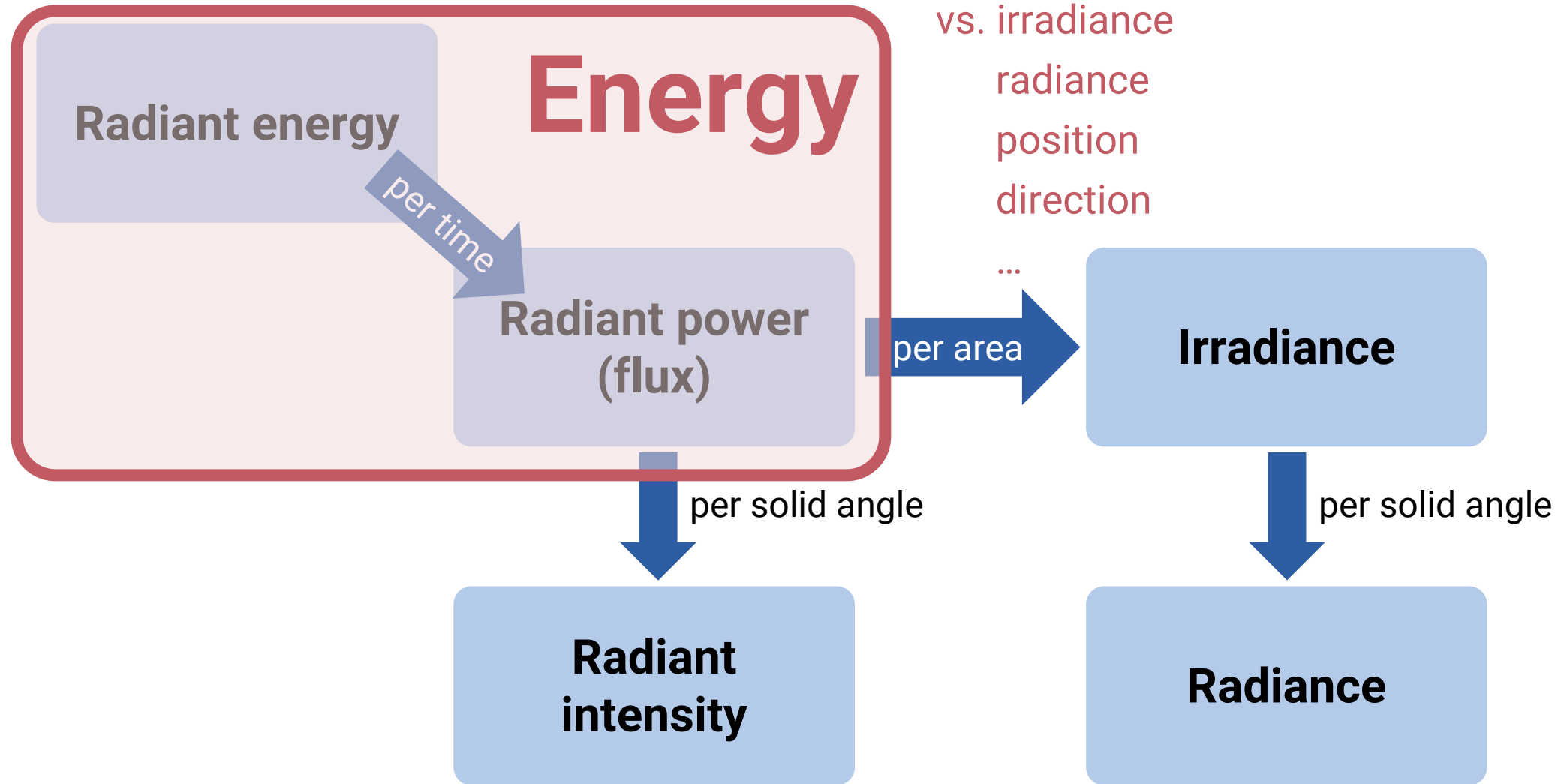
## Practice

Radiance is invariant along ray:

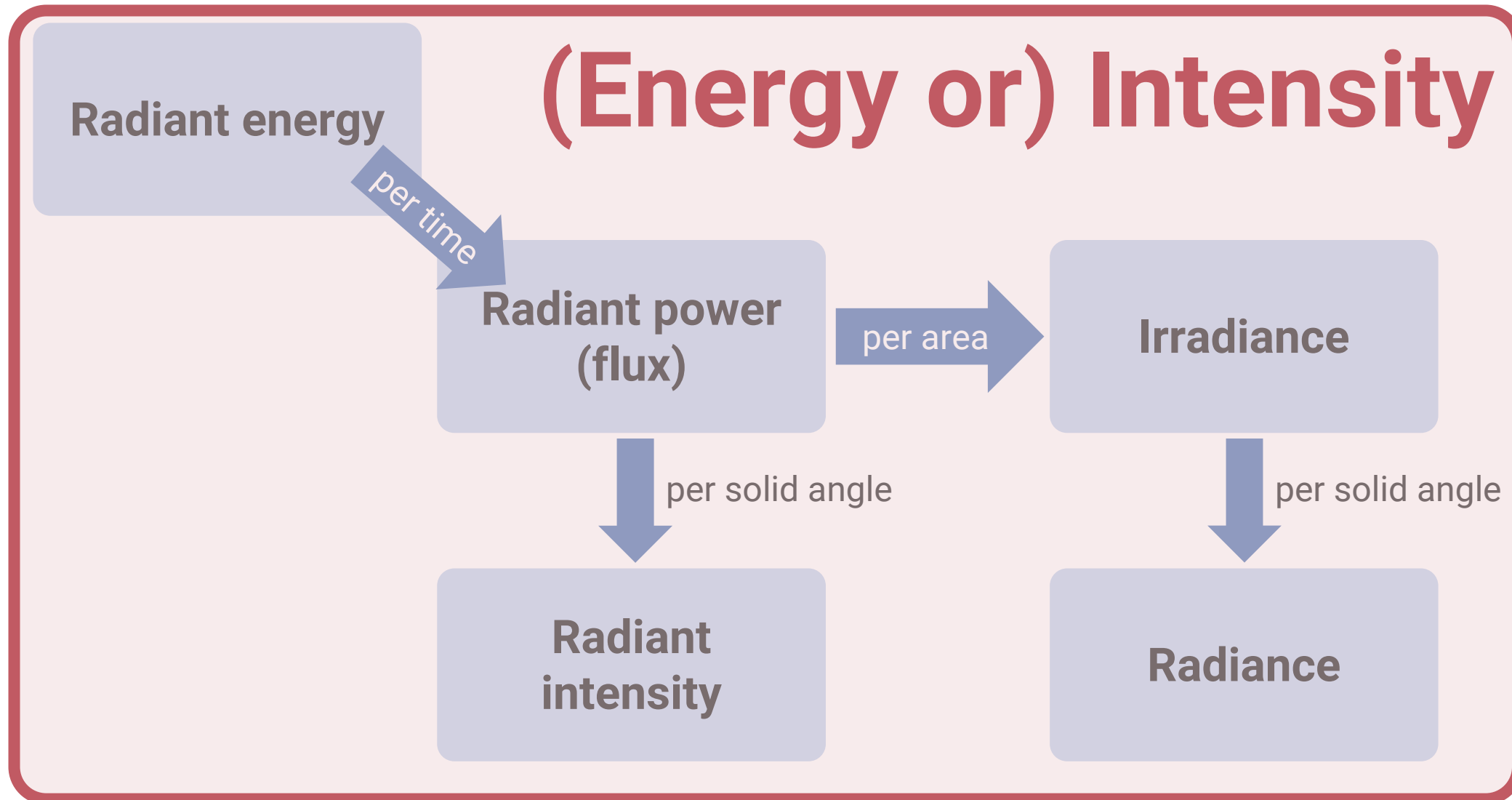
$$L(p, \hat{\omega}) = L(p + t\hat{\omega}, \hat{\omega}) \quad \forall t \in \mathbb{R}$$

whenever there is no material between  $p$  and  $p + t\hat{\omega}$

# Slight abuse of terminology



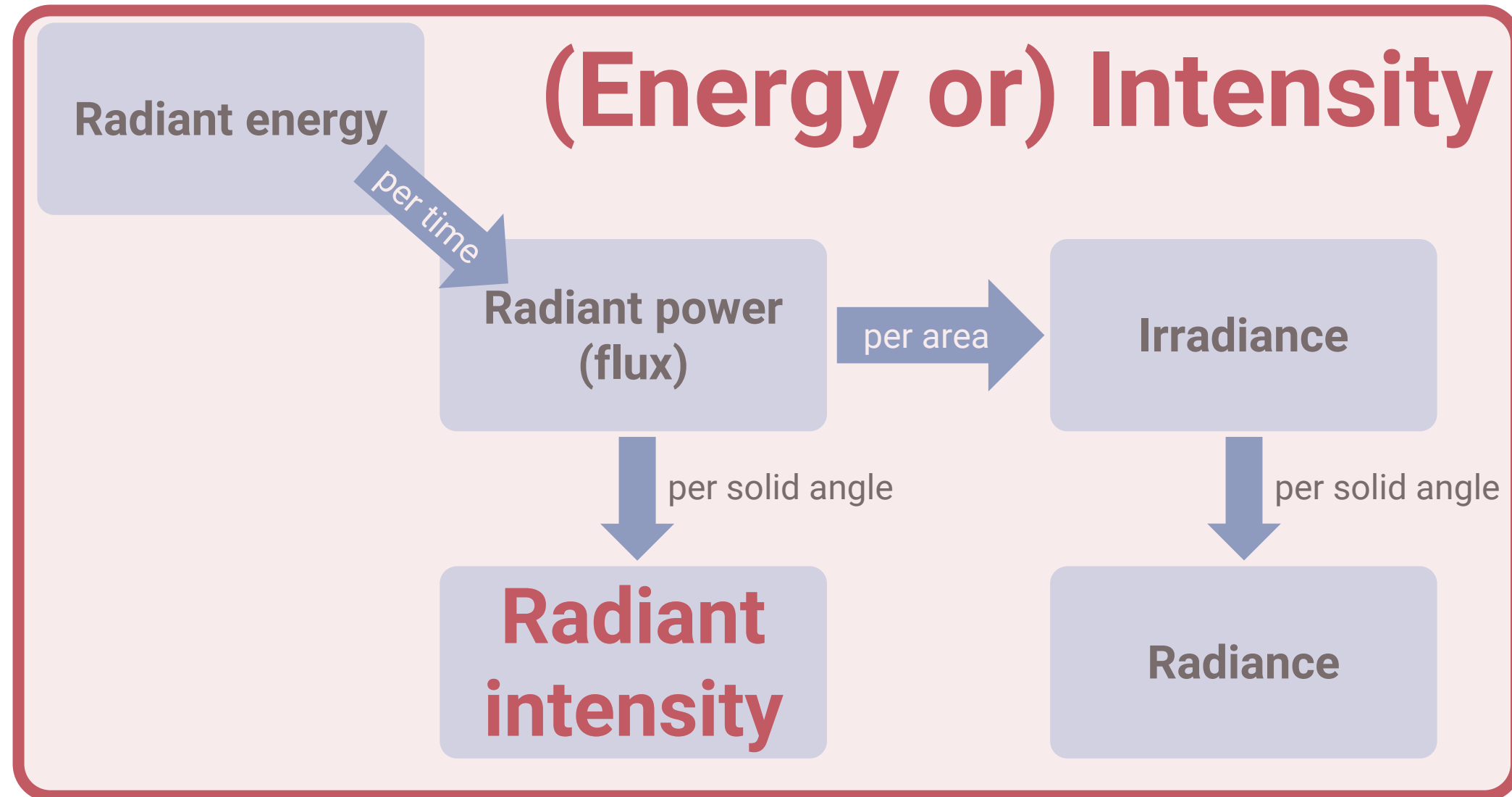
# Slight abuse of terminology



vs. position  
direction

...

# Unfortunate ambiguity



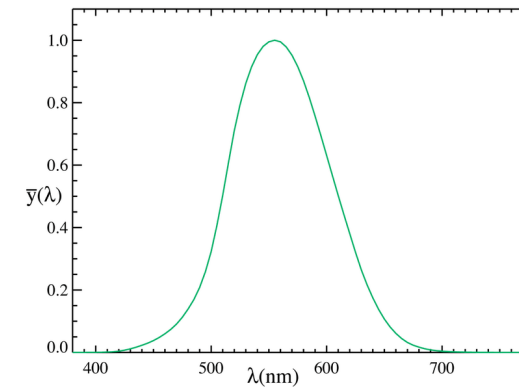
# Photometry?



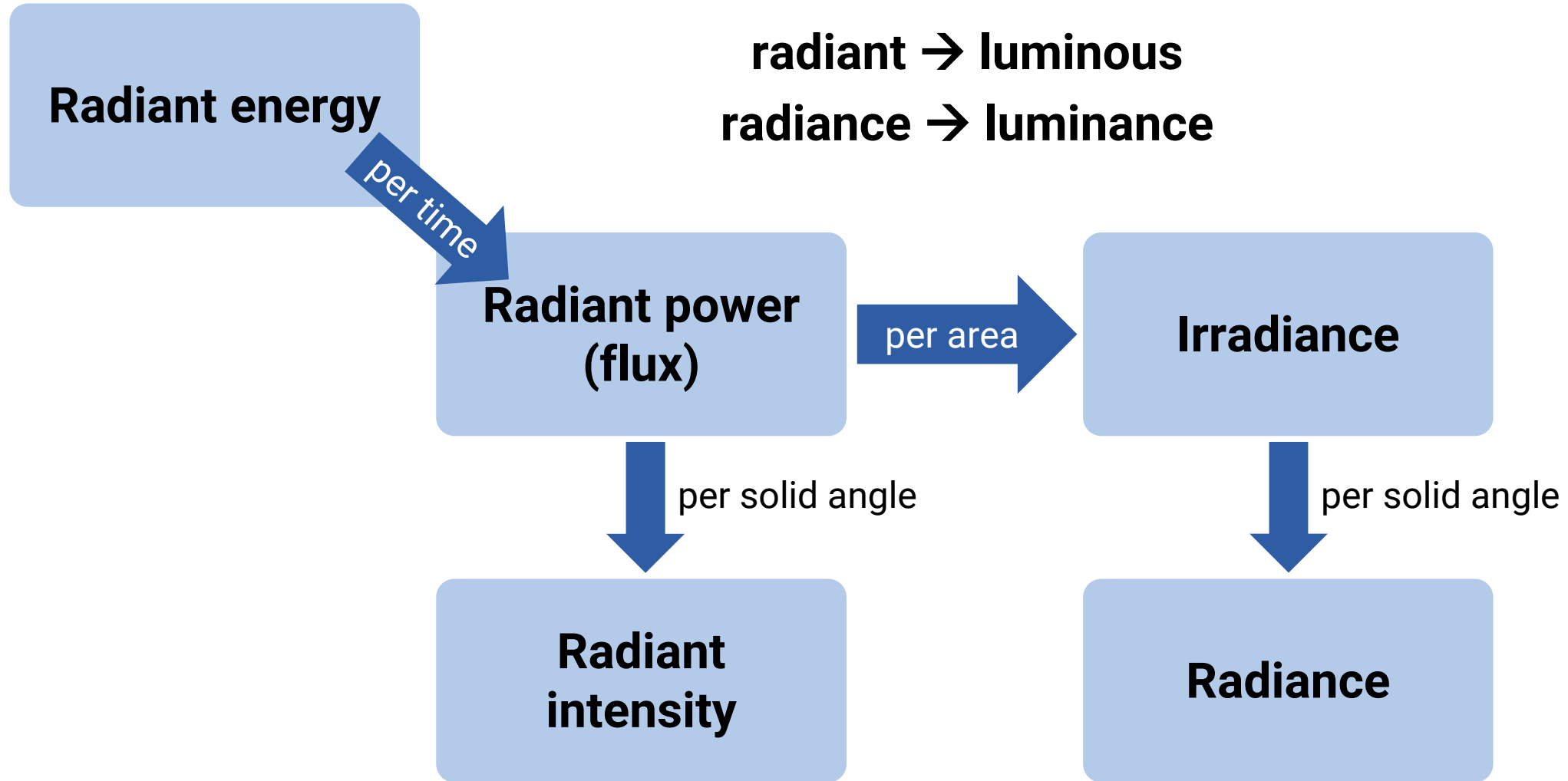
**Radiometry:** physical energy

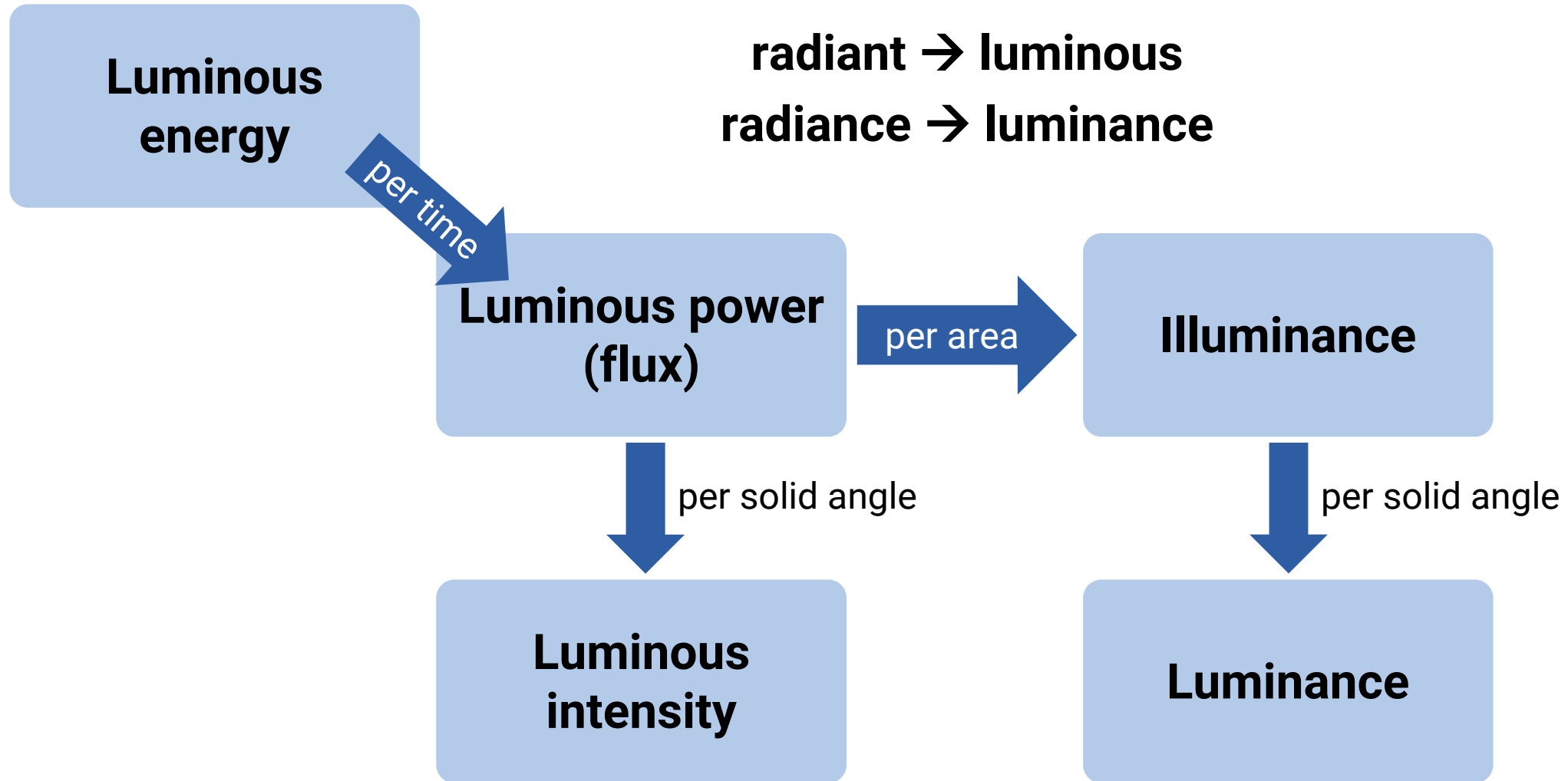
**Photometry:** how bright human perceive

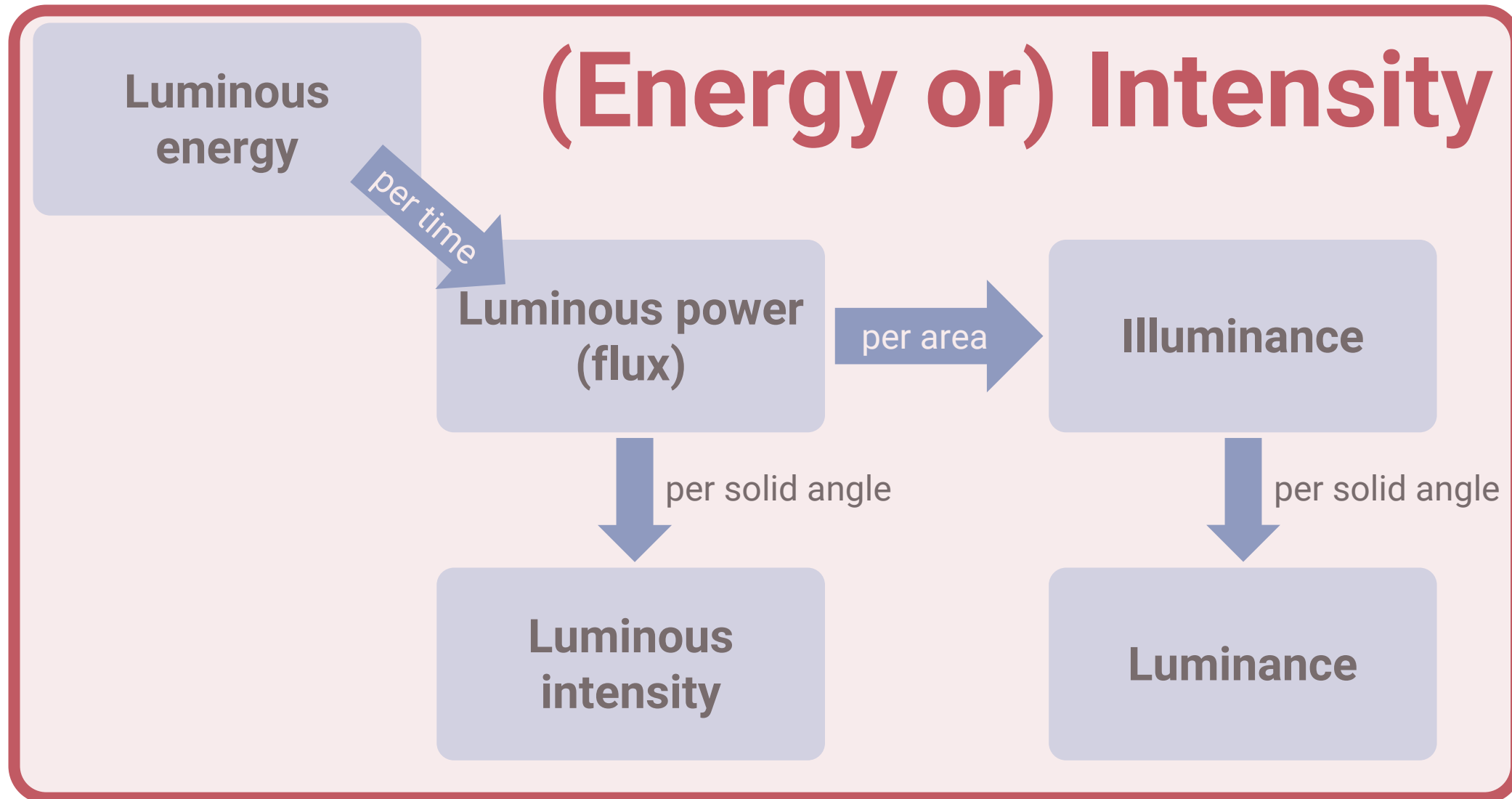
$$\int_{380\text{nm}}^{700\text{nm}} (\text{radiometric quantity per wavelength}) (\text{luminous efficiency function}) d\lambda$$



# Photometry?

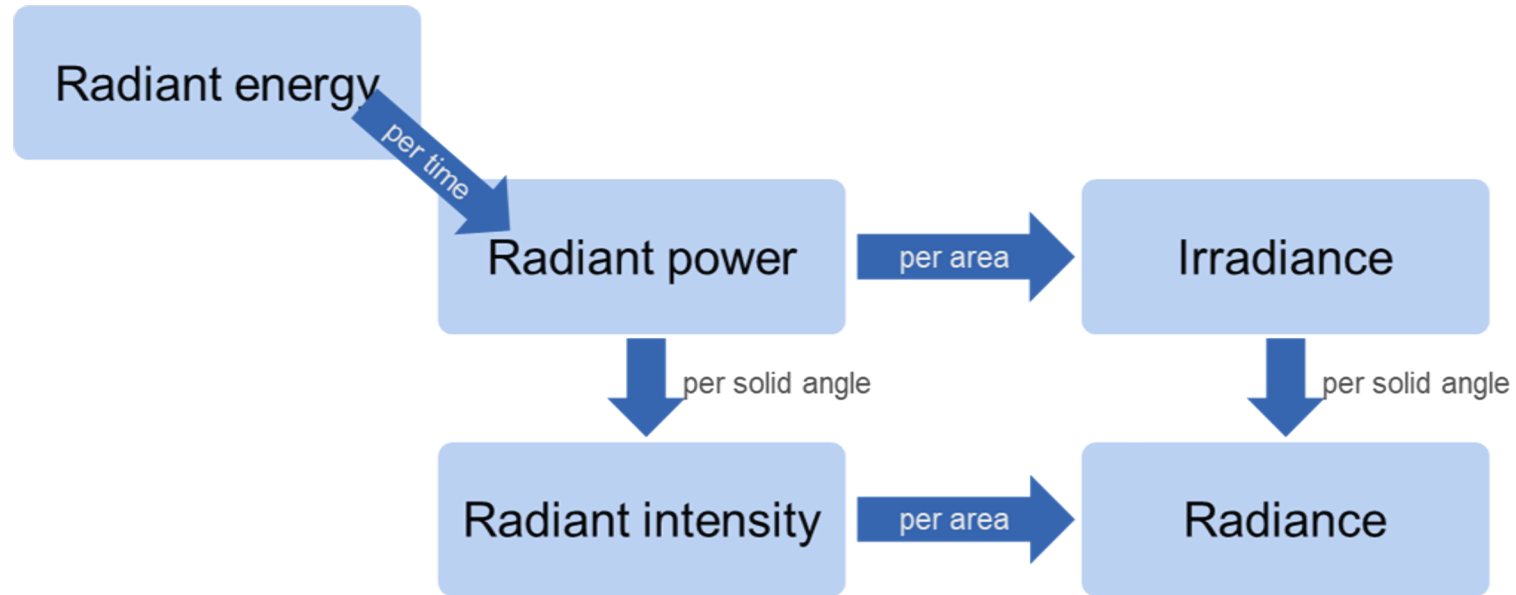








# Radiometric quantities

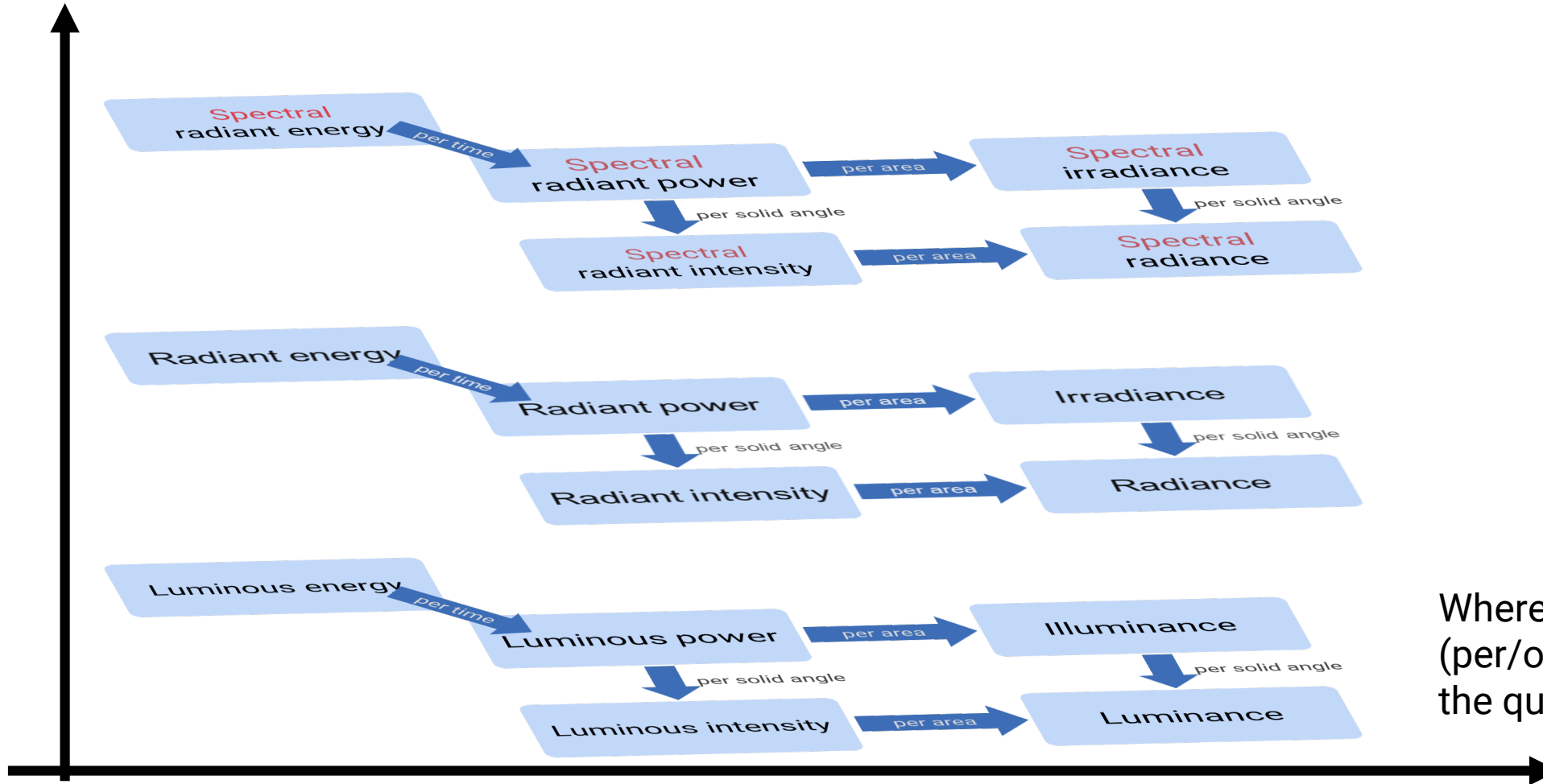


# Radiometric quantities



Physical energy? (per wavelength?)

Visible brightness for human?



Where  
(per/over position? direction?)  
the quantity is defined

# Bidirectional reflectance distribution function



***We roughly say...***

$$\text{BRDF } f_s: \frac{\text{outgoing radiance along } \omega_o}{\text{incident irradiance at } \omega_i}$$

Previous slide

**Irradiance**

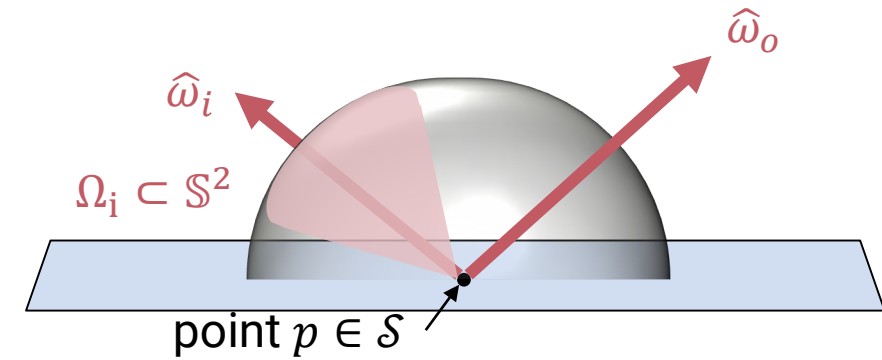
**at a point *on a surface***  
**(*dosen't depend on direction*)**

????

# Bidirectional reflectance distribution function

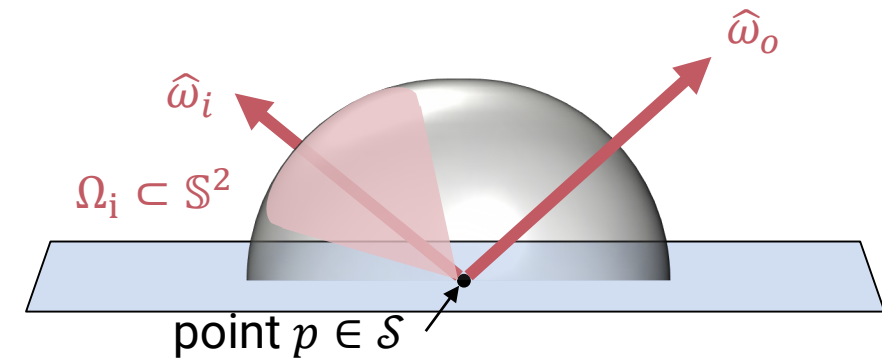


Concept of density	
<b>Mass</b>	<b>Density:</b> mass per volume
<ul style="list-style-type: none"> <li>✓ Mass of the object <math>O</math></li> <li>✓ Mass of some region (volume) <math>V</math></li> <li>✗ Mass at the point <math>p</math> → illegal or meaningless (always zero)</li> </ul>	<ul style="list-style-type: none"> <li>✗ Density of the object <math>O</math> → illegal or "average density" of the object <math>O</math></li> <li>✗ Density of some region (volume) <math>V</math> → illegal or "average density" of the volume <math>V</math></li> <li>✓ Density at the point <math>p</math></li> </ul>



$$\text{BRDF } f_s(p, \hat{\omega}_i, \hat{\omega}_o) = \lim_{\substack{\text{sol.ang.}(\Omega_i) \rightarrow 0 \\ \omega_i \in \Omega_i}} \frac{L(p, \hat{\omega}_o)}{E_s^{(\text{in})}(p, \Omega_i)}$$

# Bidirectional reflectance distribution function



$$\text{BRDF } f_s(p, \hat{\omega}_i, \hat{\omega}_o) = \lim_{\substack{\text{sol.ang.}(\Omega_i) \rightarrow 0 \\ \omega_i \in \Omega_i}} \frac{L(p, \hat{\omega}_o)}{E_s^{(\text{in})}(p, \Omega_i)}$$

$$= \lim_{\substack{\text{sol.ang.}(\Omega_i) \rightarrow 0 \\ \omega_i \in \Omega_i}} \frac{L(p, \hat{\omega}_o)}{\int_{\Omega_i} L^{(\text{in})}(p, \hat{\omega}_i) |\hat{n} \cdot \hat{\omega}_i| d\hat{\omega}_i} = \lim_{\substack{\text{sol.ang.}(\Omega_i) \rightarrow 0 \\ \omega_i \in \Omega_i}} \frac{L(p, \hat{\omega}_o)}{L^{(\text{in})}(p, \hat{\omega}_i) |\hat{n} \cdot \hat{\omega}_i| \text{sol.ang.}(\Omega_i)}$$

# Rendering equation



$+L_e(p, \hat{\omega}_o) \Rightarrow$ , then we get the rendering equation

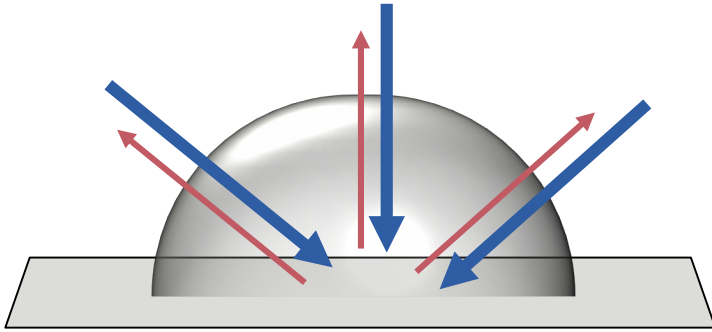
$$L^{(\text{out})}(p, \hat{\omega}_o) = \int_{\mathbb{S}^2} L^{(\text{in})}(p, \hat{\omega}_i) f_s(p, \hat{\omega}_i, \hat{\omega}_o) |\hat{n} \cdot \hat{\omega}_i| d\hat{\omega}_i$$

per solid angle

over solid angle

$$f_s(p, \hat{\omega}_i, \hat{\omega}_o) = \lim_{\substack{\text{sol.ang.}(\Omega_i) \rightarrow 0 \\ \omega_i \in \Omega_i}} \frac{L(p, \hat{\omega}_o)}{L^{(\text{in})}(p, \hat{\omega}_i) |\hat{n} \cdot \hat{\omega}_i| \text{sol.ang.}(\Omega_i)}$$

# Properties of BRDF: energy conservation



$\frac{\text{outgoing irradiance}}{\text{incident irradiance}} \leq 1$ , for any illumination condition

$$\frac{\int_{\mathbb{S}^2} L^{(\text{out})}(p, \hat{\omega}) |\hat{n} \cdot \hat{\omega}| d\hat{\omega}}{\int_{\mathbb{S}^2} L^{(\text{in})}(p, \hat{\omega}) |\hat{n} \cdot \hat{\omega}| d\hat{\omega}} =$$

rendering equation

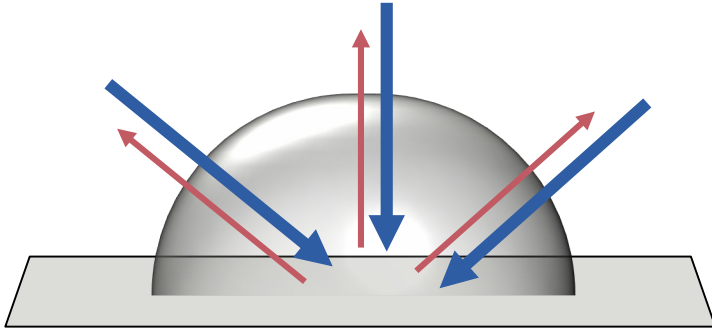
$$\frac{\int_{\mathbb{S}^2} \int_{\mathbb{S}^2} L^{(\text{in})}(p, \hat{\omega}_i) f_s(p, \hat{\omega}_i, \hat{\omega}_o) |\hat{n} \cdot \hat{\omega}_i| d\hat{\omega} |\hat{n} \cdot \hat{\omega}_o| d\hat{\omega}_o}{\int_{\mathbb{S}^2} L^{(\text{in})}(p, \hat{\omega}) |\hat{n} \cdot \hat{\omega}| d\hat{\omega}} \leq 1,$$

for **any positive function**  $L^{(\text{in})}(p, \cdot)$ .

Taking  $L^{(\text{in})}(p, \cdot)$  as a Dirac delta function centered at  $\hat{\omega}_i$ ,

$$\therefore \int_{\mathbb{S}^2} f_s(p, \hat{\omega}_i, \hat{\omega}_o) |\hat{n} \cdot \hat{\omega}_o| d\hat{\omega}_o \leq 1, \forall \hat{\omega}_i$$

# Properties of BRDF: energy conservation



$\frac{\text{outgoing irradiance}}{\text{incident irradiance}} \leq 1$ , for any illumination condition

$< 1$ : energy losses

$= 1$ : energy conserves

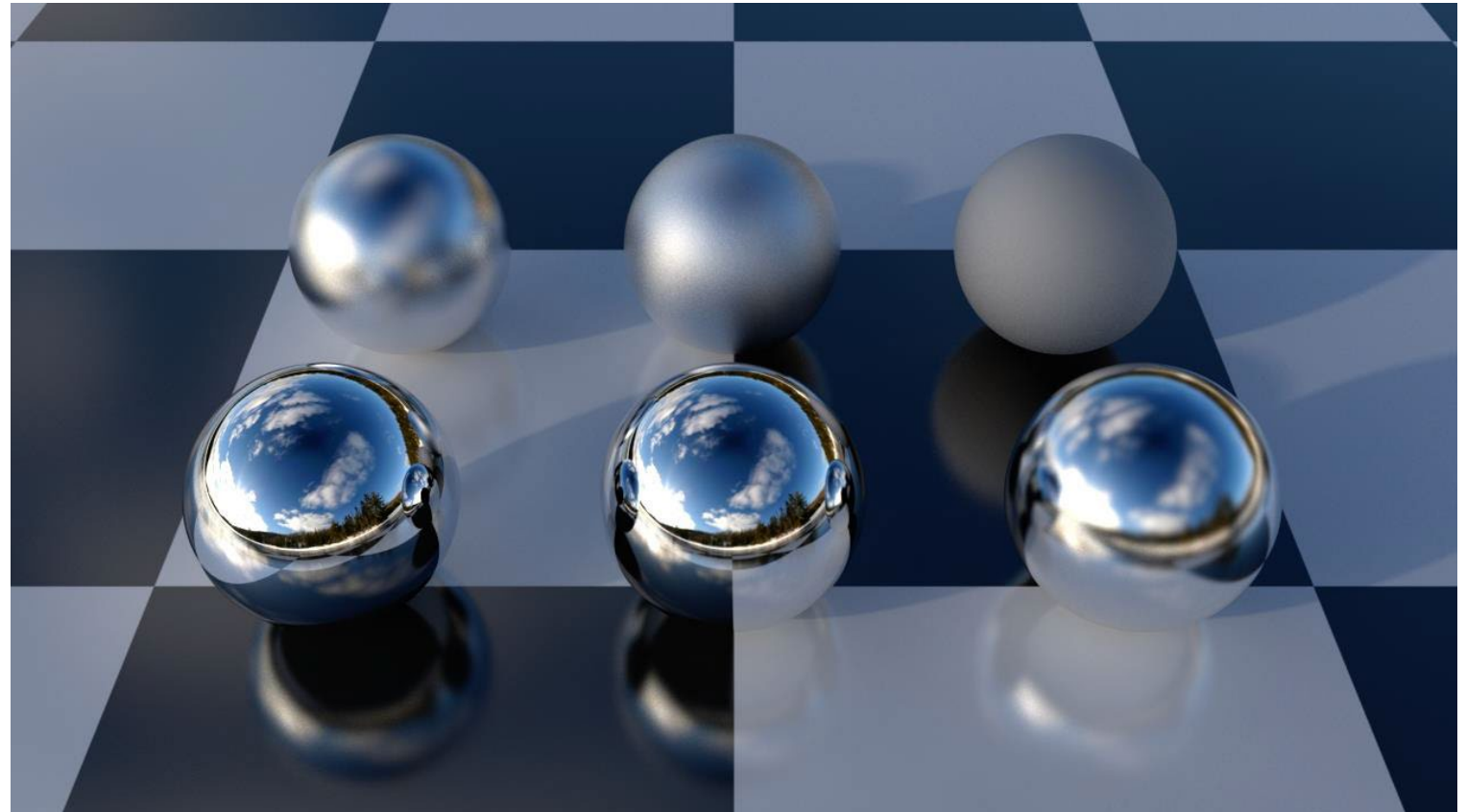
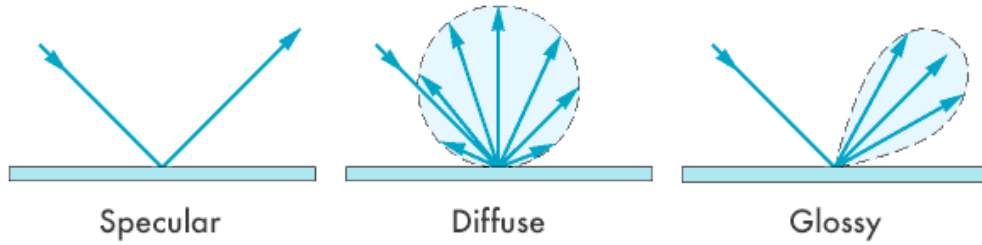
$> 1$ : impossible!

## Energy Conservation

$$\int_{\mathbb{S}^2} f_s(p, \hat{\omega}_i, \hat{\omega}_o) |\hat{n} \cdot \omega_o| d\hat{\omega}_o \leq 1, \forall \hat{\omega}_i$$



# Example BRDFs



source: Keenan Crane [\[1\]](#)

# Example BRDFs



- Pure diffuse (Lambertian reflection)
  - Albedo  $\rho_d$ : ratio of energy conservation
  - $f_s$  is a constant function on  $\mathbb{S}_{\hat{\mathbf{z}}}^2$

$$\int_{\mathbb{S}_{\hat{\mathbf{z}}}^2} f_s |\hat{\mathbf{n}} \cdot \hat{\omega}_o| d\omega_o = \pi f_s = \rho_d$$

$$\therefore f_s = \frac{\rho_d}{\pi}$$

# Example BRDFs



- Pure specular
  - A Dirac delta function centered at  $\text{refl}_{\hat{n}}(\hat{\omega}_i)$ ....
  - Be careful when you treat Dirac delta functions

$$f_s(\hat{\omega}_i, \hat{\omega}_o) = a \cdot \delta_{\mathbb{S}^2}(\hat{\omega}_o, \text{refl}_{\hat{n}}(\hat{\omega}_i)), \int_{\mathbb{S}_{\hat{n}}^2} f_s(\hat{\omega}_i, \hat{\omega}_o) |\hat{n} \cdot \hat{\omega}_o| d\hat{\omega}_o = 1$$

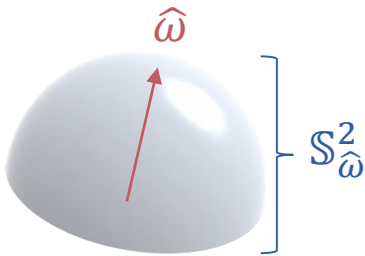
$$\therefore f_s(\hat{\omega}_i, \hat{\omega}_o) = \frac{\delta_{\mathbb{S}^2}(\hat{\omega}_o, \text{refl}_{\hat{n}}(\hat{\omega}_i))}{\hat{n} \cdot \hat{\omega}_o}$$

# Notation table



## Sets

$\mathbb{R}^n$	Euclidean space
$\mathbb{S}^2$	the unit sphere (the set of all unit vectors)
$\mathbb{S}^2_{\hat{\omega}}$	the hemisphere facing a direction $\hat{\omega} \in \mathbb{S}^2$



## Convention of variables

$p \in \mathbb{R}^3$	point in the space (or a surface)
$\hat{\omega} \in \mathbb{S}^2$	direction (unit vector)
$\bullet \hat{\omega}_{p_1 p_2} := \frac{p_2 - p_1}{\ p_2 - p_1\ }$	for any $p_1, p_2 \in \mathbb{R}^3$
$\hat{n} \in \mathbb{S}^2$	surface normal, where a point and a surface are given in context
$\mathcal{S} \subset \mathbb{R}^3$	surface
$\mathcal{V} \subset \mathbb{R}^3$	volume
$\Omega \subset \mathbb{S}^2$	solid angle (region on the unit sphere $\mathbb{S}^2$ )

## Radiometric quantities

\* time dependency is omitted for simplicity  
\* ( $\cdot$ , " $\Omega$ ") is usually omitted and assumed as an entire  $\mathbb{S}^2$  or hemisphere

$\Phi(\mathcal{S}, \Omega)$ [W]	radiant power (flux) at a surface $\mathcal{S} \subset \mathbb{R}^3$ and a solid angle $\Omega \subset \mathbb{S}^2$
$I(\hat{\omega})$ [W/sr]	radiant intensity at a direction $\hat{\omega} \in \mathbb{S}^2$ , where a point source is given in context
$E(p, \Omega)$ [W/m <sup>2</sup> ]	irradiance at $p \in \mathcal{S}$ and $\Omega \subset \mathbb{S}^2$ , where the surface $\mathcal{S} \subset \mathbb{R}^3$ is given in context
$L(p, \omega)$ [W/m <sup>2</sup> sr]	radiance at $p \in \mathbb{R}^3$ and $\hat{\omega} \in \mathbb{S}^2$
$f_s(p, \omega_i, \omega_o)$ [sr <sup>-1</sup> ]	BSDF at $p \in \mathcal{S}$ from $\hat{\omega}_i \in \mathbb{S}^2$ to $\hat{\omega}_o \in \mathbb{S}^2$ , where the surface $\mathcal{S}$ is given in context