**Help on module operational\_sep\_quantities:**

**NAME**

operational\_sep\_quantities

**FUNCTIONS**

**all\_program\_info()**

: This program will calculate various useful pieces of operational

information about SEP events from GOES-08, -10, -11, -12, -13, -14, -15

data and the SEPEM (RSDv2) dataset.

SEP event values are always calculated for threshold definitions:

>10 MeV exceeds 10 pfu

>100 MeV exceed 1 pfu

The user may add an additional threshold through the command line.

This program will check if data is already present in a 'data' directory. If

not, GOES data will be automatically downloaded from NOAA ftp site. SEPEM

(RSDv2) data must be downloaded by the user and unzipped inside the 'data'

directory. Because the SEPEM data set is so large (every 5 minutes from 1974

to 2015), the program will break up the data into yearly files for faster

reading.

The values calculated here are important for space radiation operations:

Onset time, i.e. time to cross thresholds

Peak intensity

Time of peak intensity

Rise time (onset to peak)

End time, i.e. fall below 0.85\*threshold for 3 points (15 mins for GOES)

Duration

Event-integrated fluences

User may choose differential proton fluxes (e.g. [MeV s sr cm^2]^-1) or

integral fluxes (e.g. [s sr cm^2]^-1). The program has no internal checks or

requirements on units - EXCEPT FOR THE THRESHOLD DEFINITIONS OF >10, 10

and >100, 1. If you change those thresholds in the main program accordingly,

you should be able to use other units. Also, all of the plots and messages

refer to MeV, pfu, and cm.

If a previous event is ongoing and the specified time period starts with a

threshold already crossed, you may try to set the --DetectPreviousEvent

flag. If the flux drops below threshold before the next event starts, the

program will identify the second event. This will only work if the

threshold is already crossed for the very first time in your specified

time period, and if the flux drops below threshold before the next event

starts.

If the event has an initial increase above threshold for a few points, falls

below threshold, then continues to increase above threshold again, you

may try to use the --TwoPeaks feature to capture the full duration of the

event. The initial increase above threshold must be less than a day. An

example of this scenario can be seen in >100 MeV for 2011-08-04.

**RUN CODE FROM COMMAND LINE** (put on one line), e.g.:

python3 operational\_sep\_quantities.py --StartDate 2012-05-17

--EndDate '2012-05-19 12:00:00' --Experiment GOES-13

--FluxType integral --showplot

**RUN CODE FROM COMMAND FOR USER DATA SET** (put on one line), e.g.:

python3 operational\_sep\_quantities.py --StartDate 2012-05-17

--EndDate '2012-05-19 12:00:00' --Experiment user --ModelName MyModel

--UserFile MyFluxes.txt --FluxType integral --showplot

**RUN CODE IMPORTED INTO ANOTHER PYTHON PROGRAM**, e.g.:

import operational\_sep\_quantities as sep

start\_date = '2012-05-17'

end\_date = '2012-05-19 12:00:00'

experiment = 'GOES-13'

flux\_type = 'integral'

model\_name = '' #if experiment is user, set model\_name to describe data set

user\_file = '' #if experiment is user, specify filename containing fluxes

showplot = True #Turn to False if don't want to see plots

detect\_prev\_event = True #Helps if previous event causes high intensities

two\_peaks = False #Helps if two increases above threshold in one event

threshold = '100,1' #default; modify to add a threshold to 10,10 and 100,1

sep.run\_all(start\_date, end\_date, experiment, flux\_type, model\_name,

user\_file, showplot, detect\_prev\_event, two\_peaks, threshold)

Set the desired directory locations for the data and output at the beginning

of the program in datapath and outpath. Defaults are 'data' and 'output'.

In order to calculate the fluence, the program determines time\_resolution

(seconds) from two (fairly random) data points at the start of the SEP

event. GOES and SEPEM data sets have a time resolution of 5 minutes. If the

user wishes to use a data set with measurements at irregular times, then the

subroutine calculate\_fluence should be modified.

**OUTPUT:** This program outputs 3 to 4 files, 1 per defined threshold plus

a summary file containing all of the values calculated for each threshold.

A file named as e.g. fluence\_GOES-13\_differential\_gt10\_2012\_3\_7.csv contains

the event-integrated fluence for each energy channel using the specified

threshold (gt10) to determine start and stop times.

A file named as e.g. sep\_values\_GOES-13\_differential\_2012\_3\_7.csv contains

start time, peak flux, etc, for each of the defined thresholds.

Added functionality to ouput the >10 MeV and >100 MeV time series for the

date range input by the user. If the original data were integral fluxes,

then the output files simply contain the >10 and >100 MeV time series from

the input files. If the original data were differential fluxes, then the

estimated >10 and >100 MeV fluxes are output as time series.

**USER INPUT DATA SETS:** Users may input their own data set. For example, if an

SEP modeler would like to feed their own intensity time series into this

code and calculate all values in exactly the same way they were calculated

for data, it is possible to do that. Fluxes should be in units of

1/[MeV cm^2 s sr] or 1/[cm^2 s sr] and energy channels in MeV for the plot

labels to be correct. You can use any units, as long as you are consistent

with energy units in energy channel/bin definition and in fluxes and you

MODIFY THE THRESHOLD VALUES TO REFLECT YOUR UNITS. You may then want to

modify plot labels accordingly if not using MeV and cm.

NOTE: The first column in your flux file is assumed to be time in format

YYYY-MM-DD HH:MM:SS. IMPORTANT FORMATTING!!

NOTE: The flux file may contain header lines that start with a hash #,

including blank lines.

NOTE: Any bad or missing fluxes must be indicated by a negative value.

NOTE: Put your flux file into the "datapath" directory.

NOTE: Please use only differential or integral channels. Please do not mix

them. You may have one integral channel in the last bin, as this is the way

HEPAD works and the code has been written to include that HEPAD >700 MeV

bin along with lower differential channels.

**USER VARIABLES:** The user must modify the following variables at the very

top of the code (around line 30):

user\_col - identify columns in your file containing fluxes to analyze;

even if your delimeter is white space, consider the date-time

column as one single column. SET AT TOP OF CODE.

user\_delim - delimeter between columns, e.g. " " or "," Use " " for

any amount of whitespace. SET AT TOP OF CODE.

user\_energy\_bins - define your energy bins at the top of the code in the

variable user\_energy\_bins. Follow the format in the subroutine

define\_energy\_bins. SET AT TOP OF CODE.

user\_fname - specify the name of the file containing the fluxes

through an argument in the command line. --UserFile The

user\_fname variable will be updated with that filename. ARGUMENT

time\_resolution - will be calculated using two time points in your file;

if you have irregular time measurements, calculate\_fluence()

must be modified/rewritten. AUTOMATICALLY DETERMINED.

**calculate\_event\_info(energy\_thresholds, flux\_thresholds, dates, integral\_fluxes, detect\_prev\_event, two\_peaks)**

: Uses the integral fluxes (either input or estimated from differential

channels) and all the energy and flux thresholds set in the main program

to calculate SEP event quantities.

Threshold crossing time (onset)

Peak Flux in date range specified by user

Time of Peak Flux

Rise Time (onset to peak flux)

Event End Time (below 0.85\*threshold for 3 data points, e.g 15 min)

Duration (onset to end)

If the detect\_prev\_event flag is set to true and the threshold is crossed

on the first time in the specified date range, this indicates that fluxes

were already high due to a previous event. The code will look for the

flux to drop below threshold and then increase above threshold again

during the specified time period. If this flag is not set, then the code

simply take the first threshold crossing as the start of the SEP event.

If the two\_peaks flag is set to true, the code will use the first

identified threshold crossing as the start time. A second threshold

crossing will be allowed. The event end will be determined as the drop

below threshold after the second threshold crossing. The duration will

be the end time - first threshold crossing time. The peak will be

determined as the largest flux value found in both threshold crossings.

The peak and rise times will be calculated from the first threshold

crossing time.

**calculate\_fluence(dates, flux)**

: This subroutine sums up all of the flux in the 1D array "flux". The

"dates" and "flux" arrays input here should reflect only the intensities

between the SEP start and stop times, determined by the subroutine

integral\_threshold\_crossing. "flux" should contain the intensity

time series for a single energy bin or single integral channel (1D

array). The subroutine does not differentiate between differential or

integral flux. dates contains the 1D array of datetimes that

correspond to the flux measurements.

The extract\_date\_range subroutine is used prior to calling this one to

make the dates and fluxes arrays covering only the SEP time period.

The flux will be multiplied by time\_resolution and summed for all of the

data between the start and end times. Data gaps are taken into account.

Fluence units will be 1/[MeV cm^2 sr] for GOES or SEPEM differential

fluxes or 1/[cm^2 sr] for integral fluxes.

For SRAG operational purposes, the event start and end is determined by

the >100 MeV fluxes, as they are most pertinent to astronaut health

inside of the space station.

**check\_data(startdate, enddate, experiment, flux\_type)**

: Check that the files containing the data are in the data directory. If

the files for the requested dates aren't present, they will be

downloaded from the NOAA website. For SEPEM (RSDv2) data, if missing,

the program prints the URL from which the data can be downloaded and

unzipped manually.

The RSDv2 data set is very large and takes a long time to read as a

single file. This program will generate files containing fluxes for

each year for faster reading.

**check\_for\_bad\_data(dates, fluxes, energy\_bins)**

: Search the data for bad values (flux < 0) and fill the missing data with

an estimate flux found by performing a linear interpolation with time,

using the good flux values immediately surrounding the data gap.

**check\_paths()**

: Check that the paths that hold the data and output exist. If not, create.

**define\_energy\_bins(experiment, flux\_type)**

: Define the energy bins for the selected spacecraft or data set.

If the user inputs their own file, they must set the user\_energy\_bins

variable at the top of the code.

**do\_interpolation(i, dates, flux)**

: If bad fluxes (flux < 0) are found in the data, find the first prior

data point and the first following data point that have good flux values.

Perform linear interpolation in time:

F(t) = F1 + (t - t1)\*(F2 - F1)/(t2 - t1)

This subroutine does the calculation for a single instance of bad data

that corresponds to array index i.

**extract\_date\_range(startdate, enddate, all\_dates, all\_fluxes)**

: Extract fluxes only for the dates in the range specified by the user.

**extract\_integral\_fluxes(fluxes, experiment, flux\_type, flux\_thresholds, energy\_thresholds, energy\_bins)**

: Select or create the integral fluxes that correspond to the desired

energy thresholds.

If the user selected differential fluxes, then the

differential fluxes will be converted to integral fluxes with the

minimum energy defined by the set energy thresholds.

If the user selected integral fluxes, then the channels corresponding to

the desired energy thresholds will be identified.

**find\_goes\_data\_dimensions(filename)**

: Input open csv file of GOES data. Identifies the start of the data by

searching for the string 'data:', then returns the number of header

rows and data rows present in the file.

**from\_differential\_to\_integral\_flux(experiment, min\_energy, energy\_bins, fluxes)**

: If user selected differential fluxes, convert to integral fluxes to

caluculate operational threshold crossings (>10 MeV protons exceed 10

pfu, >100 MeV protons exceed 1 pfu).

Assume that the measured fluxes correspond to the center of the energy

bin and use power law interpolation to extrapolate integral fluxes

above user input min\_energy.

The intent is to calculate >10 MeV and >100 MeV fluxes, but leaving

flexibility for user to define the minimum energy for the integral flux.

An integral flux will be provided for each timestamp (e.g. every 5 mins).

**get\_fluence\_spectrum(experiment, flux\_type, model\_name, energy\_threshold, flux\_threshold, sep\_dates, sep\_fluxes, energy\_bins, save\_file)**

: Calculate the fluence spectrum for each of the energy channels in the

user selected data set. If the user selected differential fluxes, then

the fluence values correspond to each energy bin. If the user selected

integral fluxes, then the fluence values correspond to each integral bin.

Writes fluence values to file according to boolean save\_file.

**get\_west\_detector(filename, dates)**

: For GOES-13+, identify which detector is facing west from the

orientation flag files. Get an orientation for each data point.

EPEAD orientation flag. 0: A/W faces East and B/E faces West.

1: A/W faces West and B/E faces East. 2: yaw-flip in progress.

**integral\_threshold\_crossing(energy\_threshold, flux\_threshold, dates, fluxes)**

: Calculate the time that a threshold is crossed.

Operational thresholds used by the NASA JSC Space Radiation Analysis

Group to determine actions that should be taken during an SEP event are:

>10 MeV proton flux exceeds 10 pfu (1/[cm^2 s sr])

>100 MeV proton flux exceeds 1 pfu (1/[cm^2 s sr])

An SEP event is considered to start if 3 consecutive points are

above threshold. The start time is set to the first point that crossed

threshold.

If the user input differential flux, then the program has converted it

to integral fluxes to calculate the integral threshold crossings.

If the user selected fluxes that were already integral fluxes, then

the correct channel was identified to determine these crossings.

GOES and SEPEM data have a 5 minute time resolution, so initial

crossings are accurate to 5 minutes.

This program also calculates peak flux and time of peak flux for time

period specified by user. It then calculates rise time by comparing:

rise time = time of peak flux - threshold crossing time

If no thresholds are crossed during specified time period, peak time

and rise time will be 0.

The event end time is much more subjective. For this program, I follow

the code that officially calls the end of an event when SRAG supports

ISS operations. When the flux has three consecutive points (15 minutes

of GOES data) below 0.85\*threshold, then the event is ended. The end time

is corrected back to the first of the three data points. This end

criteria has no physics basis. It simply ensures that a new event will

not erroneously begin within the SRAG SEP alarm code due to flux

fluctuations around threshold as the SEP flux decays away slowly.

If the time period specified by the user ends before the event falls

below 0.85\*threshold, then the event\_end\_time is set to the last time

in the specified time range. The program will give a message indicating

that the user may want to extend the requested time frame in order to

better-capture the end of the event.

**make\_yearly\_files(filename)**

: Convert a large data set into yearly files.

**print\_values\_to\_file(experiment, flux\_type, model\_name, energy\_thresholds, flux\_thresholds, crossing\_time, peak\_flux, peak\_time, rise\_time, event\_end\_time, duration, integral\_fluences)**

: Write all calculated values to file for all thresholds. Event-integrated

fluences for >10, >100 MeV (and user-defined threshold) will also be

included. Writes out file with name e.g.

output/sep\_values\_experiment\_fluxtype\_YYYY\_M\_D.csv

**read\_in\_files(experiment, flux\_type, filenames1, filenames2, filenames\_orien)**

: Read in the appropriate data files with the correct format. Return an

array with dates and fluxes. Bad flux values (any negative flux) are set

to -1. Format is defined to work with the files downloaded directly from

NOAA or the RSDv2 (SEPEM) website as is.

The fluxes output for the GOES-13+ satellites are always from the

westward-facing detector (A or B) by referring to the orientation flags

provided in the associated orientation file. Data taken during a yaw

flip (orientation flag = 2) are excluded and fluxes are set to -1.

Note that the EPS detectors on GOES-08 and -12 face westward. The

EPS detector on GOES-10 faces eastward. GOES-11 is a spinning satellite.

**read\_in\_user\_files(filenames1)**

: Read in file containing flux time profile information that was

specified by the user.

The first column MUST contain the date in YYYY-MM-DD HH:MM:SS

format. The remaining flux columns to be read in are specified by the

user in the variable user\_col at the very beginning of this program.

The date column should always be considered column 0, even if you used

whitespace as your delimeter. The code will consider the date format

YYYY-MM-DD HH:MM:SS as one column even though it contains whitespace.

Any number of header lines are allowed, but they must be indicated by #

at the very beginning, including empty lines.

Be sure to add the energy bins associated with your flux columns in the

subroutine define\_energy\_bins under the "user" is statement.

**report\_threshold\_fluences(experiment, flux\_type, model\_name, energy\_thresholds, energy\_bins, sep\_dates, sep\_fluxes)**

: Report fluences for specified thresholds, typically >10, >100 MeV.

These values are interesting to use for comparison with literature and

for quantifying event severity.

If the user has input integral channels, then this information has also

been saved in the output fluence file. If the user selected differential

channels, then these values come from the estimated integral fluxes.

**run\_all(str\_startdate, str\_enddate, experiment, flux\_type, model\_name, user\_file, showplot, detect\_prev\_event, two\_peaks, str\_thresh)**

: "Runs all subroutines and gets all needed values. Takes the command line

areguments as input. Written here to allow code to be imported into

other python scripts.

str\_startdate, str\_enddate, experiment, flux\_type are strings.

model\_name is a string. If model is "user", set model\_name to describe

your model (e.g. MyModel), otherwise set to ''.

user\_file is a string. Defaul is ''. If user is selected for experiment,

then name of flux file is specified in user\_file.

showplot, detect\_prev\_event, and two\_peaks are booleans.

Set str\_thresh to be '100,1' for default value or modify to add your own

threshold.

**save\_integral\_fluxes\_to\_file(experiment, flux\_type, model\_name, energy\_thresholds, crossing\_time, dates, integral\_fluxes)**

**DATA**

\_\_email\_\_ = 'kathryn.whitman@nasa.gov'

\_\_maintainer\_\_ = 'Katie Whitman'

badval = -1

datapath = 'data'

exp = <ufunc 'exp'>

outpath = 'output'

user\_col = array('i', [1, 2, 3, 4, 5, ...])

user\_delim = ','

user\_energy\_bins = [[300, -1], [100, -1], [60, -1], [50, -1], [30, -1]...

user\_fname = []

**VERSION**

0.6

**AUTHOR**

Katie Whitman

**FILE**

Operational-sep/operational\_sep\_quantities.py