

# STO 229 HW#2

## Problem 2.1

$$\vec{g} = -\nabla V$$

$$= -\nabla G \left( \frac{M_E}{r} - \frac{a^2}{r^3} J_2 \frac{1}{2} (3 \cos^2 \theta - 1) + \dots \right)$$

$$\approx -G \left( \frac{\partial}{\partial r} \hat{r} + \frac{1}{r} \frac{\partial}{\partial \theta} \hat{\theta} + \frac{1}{r \sin \theta} \frac{\partial}{\partial \phi} \hat{\phi} \right) \left( \frac{M_E}{r} - \frac{a^2}{r^3} J_2 \frac{1}{2} (3 \cos^2 \theta - 1) \right)$$

$$= -\frac{\partial}{\partial r} G \left( \frac{M_E}{r} - \frac{a^2}{r^3} J_2 \frac{1}{2} (3 \cos^2 \theta - 1) \right) \hat{r}$$

$$- \frac{\partial}{\partial \theta} \frac{G}{r} \left( \frac{M_E}{r} - \frac{a^2}{r^3} J_2 \frac{1}{2} (3 \cos^2 \theta - 1) \right) \hat{\theta}$$

$$- \frac{G}{r \sin \theta} \frac{\partial}{\partial \phi} \left( \frac{M_E}{r} - \frac{a^2}{r^3} J_2 \frac{1}{2} (3 \cos^2 \theta - 1) \right) \hat{\phi}$$

$$\vec{g} \approx G \left[ \left( \frac{M_E}{r^2} - \frac{3a^2 J_2 (3 \cos^2(\theta) - 1)}{2r^4} \right) \hat{r} - \left( \frac{3a^2 J_2 \sin(\theta) \cos(\theta)}{r^4} \right) \hat{\theta} \right]$$

$$r_{\text{pole}} = 6357 \text{ km}$$

$$\theta_{\text{pole}} = 0^\circ$$

$$\vec{g}_{\text{pole}} = (9.8633, 0, 0) \text{ m/s}^2$$

$$r_{\text{equator}} = 6378 \text{ km}$$

$$\theta_{\text{equator}} = 90^\circ$$

$$\vec{g}_{\text{equator}} = (9.7984, 0) \text{ m/s}^2$$

$$|\vec{g}|_{\text{pole}} = 9.863 \text{ m/s}^2 \quad \dots >$$

$$|\vec{g}|_{\text{equator}} = 9.798 \text{ m/s}^2$$

## Problem 2.2

$$U_0 = -\frac{G M_E}{r_0(\theta)} + \frac{G M_E a^2}{r_0(\theta)^3} J_2 P_2(\cos\theta) - \frac{1}{2} \omega^2 r_0(\theta)^2 \sin^2\theta$$

where  $r_0(\theta) = a(1 - f \cos^2\theta)$

$$f = \frac{3}{2} J_2 + \frac{1}{2} \frac{a^3 \omega^2}{G M_E}$$

$$J_2 \approx 0.001$$

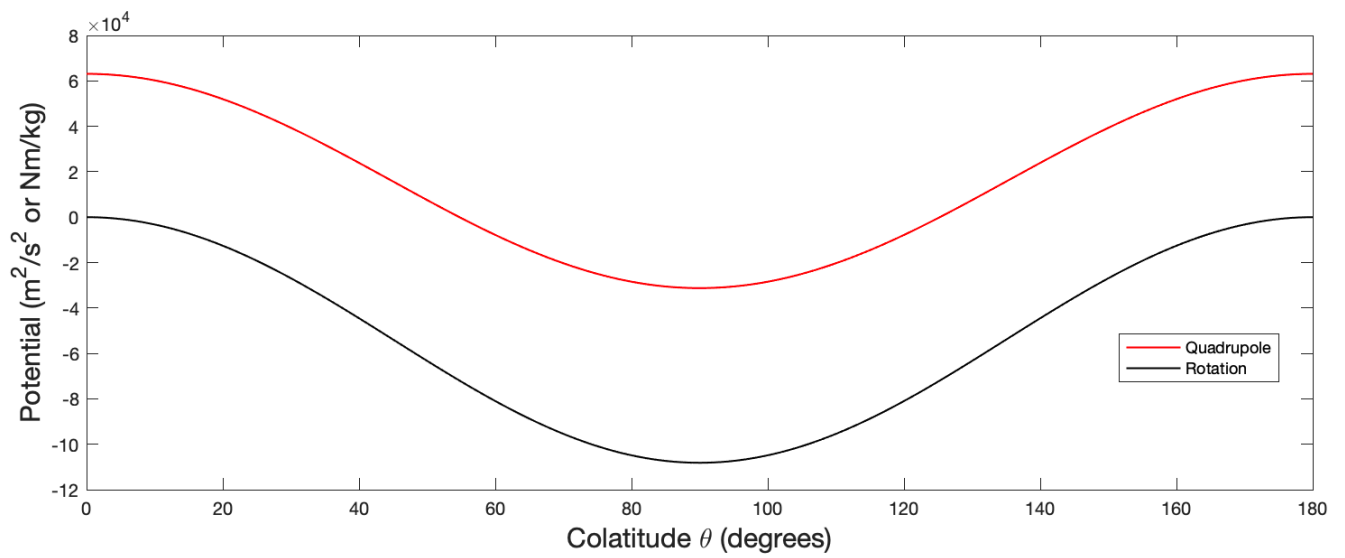
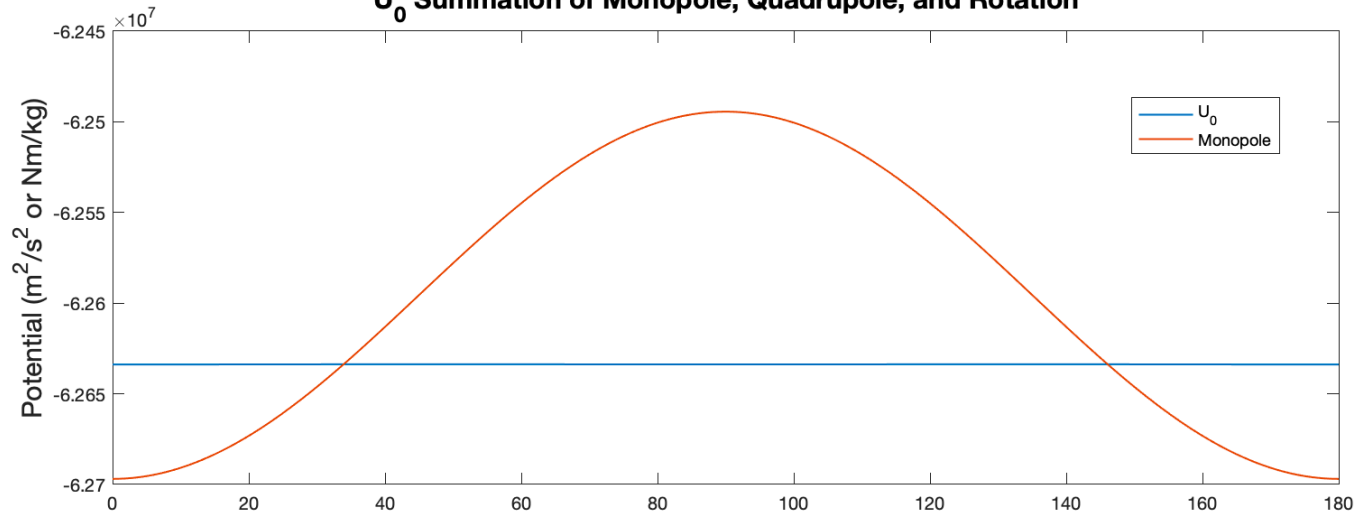
$$P_2(\cos\theta) = \frac{1}{2} (3 \cos^2\theta - 1)$$

$$\omega = 7.291150 \times 10^{-5} \frac{\text{rad}}{\text{s}}$$

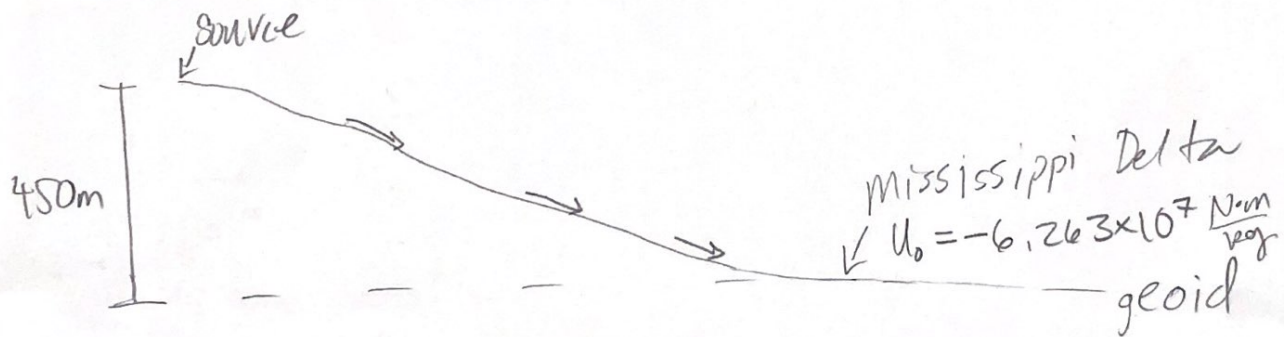
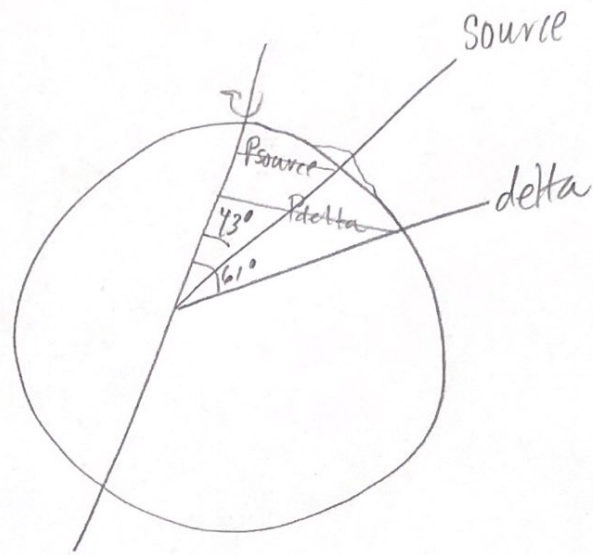
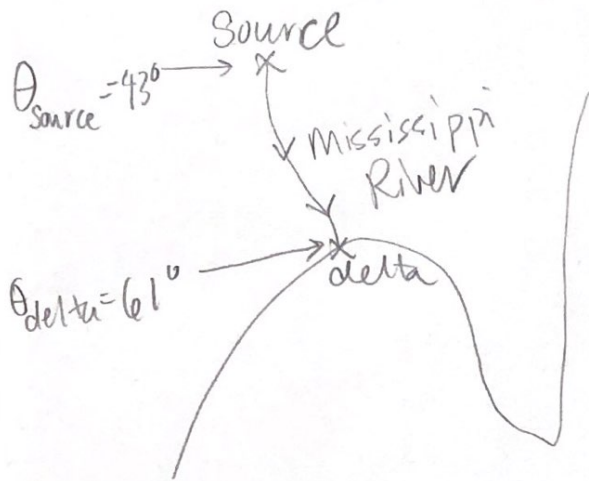
The geopotential  $U_0$  on the geoid is approximately  $-6.263 \times 10^7 \frac{\text{N} \cdot \text{m}}{\text{kg}}$  or  $\frac{\text{m}^2}{\text{s}^2}$ .

The calculation was done on Matlab under problem 2.2. The graph shows  $U_0$  being constant around the Earth. The graph also separates out the monopole, quadrupole, and spin rotation.

**$U_0$  Summation of Monopole, Quadrupole, and Rotation**



## Problem 2.3



The Mississippi delta is located on the geoid so the potential difference is the 450 m elevation and the  $18^\circ$  change in colatitude. The Mississippi river will stop flowing when  $U_{\text{source}} = U_{\text{delta}}$ .

$$U_{\text{source}} = \underbrace{\frac{-Gm_E}{r_0(43^\circ) + 450\text{m}} - \frac{Gm_E a^2 J_2 P_2(\cos(43^\circ))}{(r_0(43^\circ) + 450\text{m})^3}}_{V_{\text{source gravitational potential}}} - \underbrace{\frac{1}{2} \omega^2 (r_0(43^\circ) + 450\text{m}) \sin^2(43^\circ)}_{\text{spin potential } S_{\text{source}}}$$



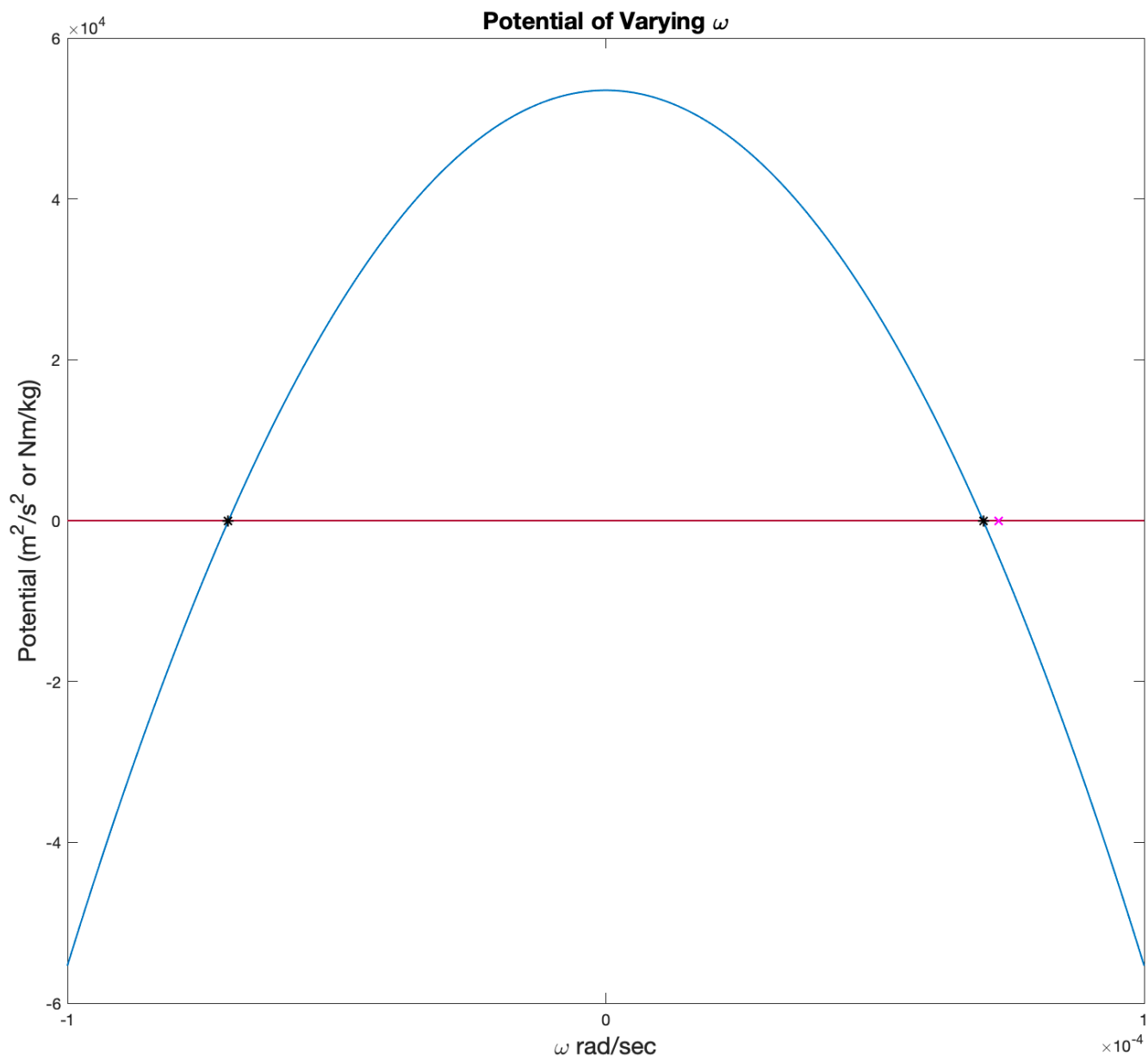
$$U_{\text{delta}} = U_0 = \underbrace{\frac{-Gm_E}{r_0(61^\circ)} + \frac{Gm_E a^2 J_2 P_2(\cos(61^\circ))}{(r_0(61^\circ))^3}}_{V_{\text{delta}}} - \underbrace{\frac{1}{2} \omega^2 (r_0(61^\circ))^2 \sin^2(61^\circ)}_{S_{\text{delta}}}$$

$$V_{\text{source}} + S_{\text{source}} = V_{\text{delta}} + S_{\text{delta}}$$

$J_2$  and  $P_2(\cos \theta)$  does not depend on  $\omega$ , however,  $r_0$  does. I varied  $\omega$  from  $-1 \times 10^{-4} \frac{\text{rad}}{\text{sec}}$  to  $1 \times 10^{-4} \frac{\text{rad}}{\text{sec}}$  and recalculated  $r_0$  for the changing spin. The graph shows where the potential is zero which is the point where there wouldn't be flow.

$$0 = -V_{\text{source}} - S_{\text{source}} + V_{\text{delta}} + S_{\text{delta}}$$

The red x is the current spin rate and the black \* are the spin rate when flow stops. The current spin rate would need to slow down to ~24.89 hour days.



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```
% SIO 229 Homework 2
clear all; close all; clc
```

## Problem 2.1

Use MacCullagh's Formula to compare the difference in gravity at the pole and at the equator without Earth's spin and assume present shape and interior structure.

```
% g = -gradient of V
% V is MacCullagh's Formula
mEarth = 5.972e24;           % Mass of Earth (kg)
J = 0.001;                   % J_2 value for flattening roughly,
    unitless
a = 6378000;                  % Radius of Earth at equator m
G = 6.6743e-11;               % Gravitational constant N*m^2/kg^2

% radius and theta
rPole = 6357000;              % radius at the pole meters
thetaPole = 0;                % theta at the pole degrees

rEquator = a;                 % radius at the equator meters
thetaEquator = 90;            % theta at the equator degrees

% component for g at the poles without phi
g_rPole = G*(mEarth/rPole^2-(3*a^2*J*(3*cosd(thetaPole)^2-1))/(
    (2*rPole^4)));
g_thetaPole = G*(-3*a^2*J*sind(thetaPole)*cosd(thetaPole))/rPole^4;

% component for g at the equator without phi
g_rEquator = G*(mEarth/rEquator^2-
    (3*a^2*J*(3*cosd(thetaEquator)^2-1))/(2*rEquator^4));
g_thetaEquator = G*(-3*a^2*J*sind(thetaEquator)*cosd(thetaEquator))/
    rEquator^4;
```

## Problem 2.2

```
% Calculate the value of the geopotential U on the geoid

w = 7.2921150e-5;            % Earth's spin rate (rad/s)
theta = 0:180;                % degrees of latitude
```

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```

f = (3/2)*J+(a^3*w^2)/(2*G*mEarth);      % flattening, unitless
r0 = a*(1-f*cosd(theta).^2);              % changing radius value in
    meters
P = (1/2)*(3*cosd(theta).^2-1);           % Legendre Polynomial,
    unitless

mono = -G*mEarth./r0;                     % Monopole N*m/kg
quad = (G*mEarth*a^2)*(J*P)./(r0.^3);     % Quadrupole N*m/kg
spin = -(1/2)*w^2*r0.^2.*sind(theta).^2;  % Rotation m^2/s^2

U0= mono+quad+spin;                       % Geopotential U on the geoid

U0mean = mean(U0)

figure(1)
subplot(2,1,1)
plot(theta, U0, 'LineWidth', 1)
hold on
plot(theta, mono, 'LineWidth', 1)
ylabel('Potential (m^2/s^2 or Nm/kg)', 'FontSize', 15)
legend('U_0', 'Monopole')
hold off
title('U_0 Summation of Monopole, Quadrupole, and
    Rotation', 'FontSize', 15)
subplot(2,1,2)
plot(theta, quad, 'r', 'LineWidth', 1)
hold on
plot(theta, spin, 'k', 'LineWidth', 1)
ylabel('Potential (m^2/s^2 or Nm/kg)', 'FontSize', 15)
legend('Quadrupole', 'Rotation')
xlabel('Colatitude \theta (degrees)', 'FontSize', 15)
set(gcf, 'color', 'w');
hold off

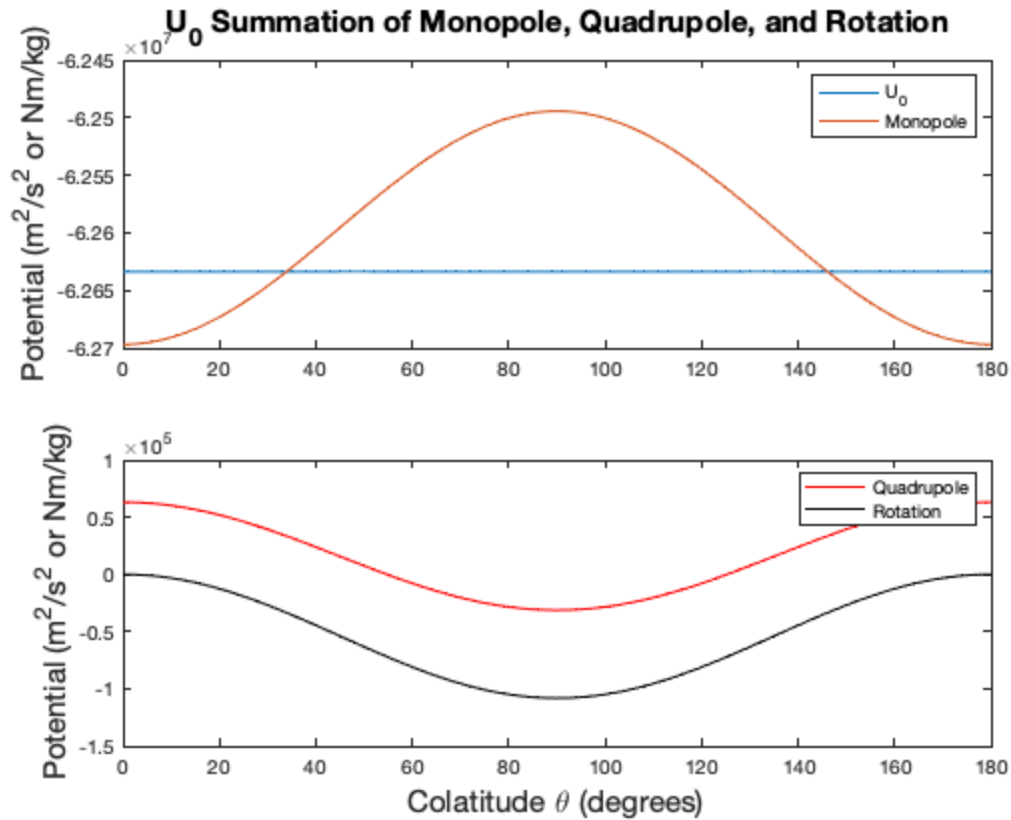
U0mean =

    -6.2634e+07

```

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## Problem 2.3

```
% Find the rotation rate, omega, when the Mississippi starts to flow
% backward from South to North. Assuming the Earth does not change
% shape.
% The source of the Mississippi is 47 degrees North and the delta is
% 29
% degrees North of the Equator.

MissLatS = 43;           % Mississippi source colatitude from the pole
degrees
MissLatD = 61;           % Mississippi delta colatitude from the pole
degrees

radiusSource = r0(44)+450; % radius of source in meters
Psource = (1/2)*(3*cosd(MissLatS).^2-1); % Legendre Polynomial
for Source

% Vary omega for when potential for the source and delta are equal
omega = (-1e-4:1e-7:1e-4); % Spin radian/sec

% Calculate the source gravitational potential, V source in N*m/kg or
% m^2/s^2
Vsource = -G*mEarth/radiusSource + G*mEarth*a^2*J*Psource/
(radiusSource^3);
```

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```

% Calculate the source apparent centrifugal force, depends on omega
m^2/s^2
SpinSource = -(1/2)*omega.^2*radiusSource^2*sind(MissLatS)^2;

% Calculate the delta gravitational potential, geoid potential, V in
N*m/kg
% depends on omega
f = (3/2)*J+(a^3*omega.^2)/(2*G*mEarth); % flattening depends on
rot. rate
r0Delta = a*(1-f.*cosd(MissLatD).^2); % radius meters
depends on rotation rate
Pdelta = (1/2)*(3*cosd(MissLatD)^2-1); % Legendre Polynomial,
unitless
Vdelta = -G*mEarth./r0Delta + G*mEarth*a^2*J*Pdelta./(r0Delta.^3);

% Calculate the delta apparent centrifugal force m^2/s^2, depends on
omega
SpinDelta = -(1/2)*omega.^2.*r0Delta.^2*sind(MissLatD)^2;

% Source and Delta potential equalling zero
SandDequal0 = -Vsource - SpinSource + Vdelta + SpinDelta;

T = (2*pi./omega)/(60*60); % Period of rotation of Earth in hours
Topp = T(300) % Period in the opposite direction
Tsame = T(1702) % Period for the same direction today

figure(2)
plot(omega, SandDequal0, 'LineWidth',1)
hold on
plot(omega, zeros(length(omega)), 'LineWidth',1)
plot([omega(1702) omega(300)], [SandDequal0(1702)
SandDequal0(300)], 'k*', 'LineWidth',1)
% plot(omega(300), SandDequal0(300), 'k*', 'LineWidth',1)
plot(w, 0, 'mx', 'LineWidth',1)
ylabel('Potential (m^2/s^2 or Nm/kg)', 'FontSize',15)
xlabel('\omega rad/sec', 'FontSize',15)
title('Potential of Varying \omega', 'FontSize',15)
set(gcf, 'color', 'w')
hold off

Topp =

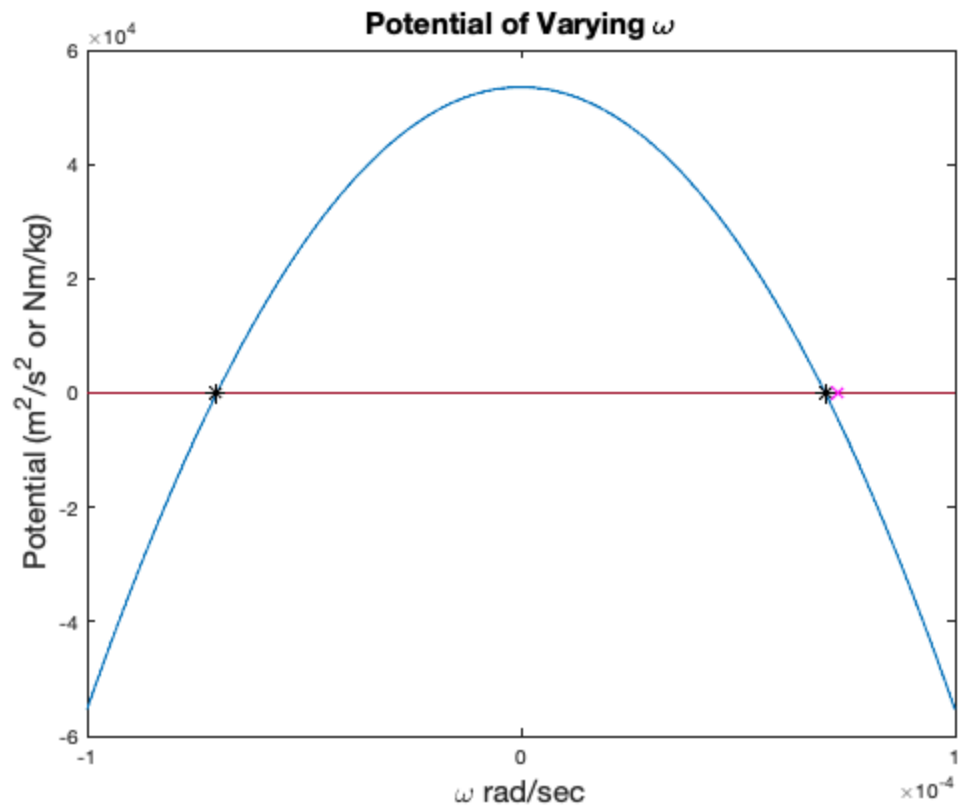
-24.8977

Tsame =

24.8977

```

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