# Computational Fluid Dynamics: Lecture 2 (ME EN 6720)

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# Mathematical representations

#### The first step:

- After identifying our physical process of interest we need to conceptualize it mathematically
- Most problems in fluid mechanics are described by a combination of:

- Conservation of Mass: 
$$\left(\frac{dm}{dt}\right)_{sys} = 0$$

- Conservation of Momentum: 
$$\sum \vec{F} = \frac{d(m\vec{V})}{dt} \Big|_{sys}$$

- Conservation of Energy: 
$$\dot{Q} - \dot{W} = \frac{dE}{dt} \Big|_{sys}$$

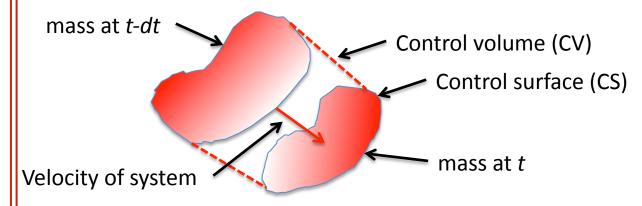
- No mass can enter or leave a system, it is a Lagrangian reference frame
- Typically in experiments and in CFD solutions we want a fixed **Eulerian** reference frame

# Reynolds Transport Theorem

How do we go from a system (Lagrangian) to a control volume (Eulerian) approach?

#### **Reynold's Transport Theorem** (RTT):

RTT converts a Lagrangian reference frame to an Eulerian one



We can derive RTT by noting that the mass of the fluid at *t* and *t-dt* is the same and examining the geometric changes

See fluids text for full derivation (e.g., Fox et al., 2009 pg 95-99)

$$\left(\frac{dN}{dt}\right)_{sys} = \frac{\partial}{\partial t} \int_{CV} \rho \eta d \nabla + \int_{CS} \rho \eta \vec{V} \cdot d\vec{A}$$

*N* = extensive property (mass, momentum, etc.)

 $\eta$  = intensive property (N per unit mass)

### Conseravtion of mass

• We can apply RTT to the system formulation of conservation of mass

$$\frac{dm}{dt}\Big|_{sys} = 0 \qquad \text{set } N = m \Rightarrow \eta = \frac{\partial N}{\partial m} = \frac{\partial m}{\partial m} = 1$$

$$\Rightarrow \frac{dm}{dt}\Big|_{sys} = \frac{\partial}{\partial t} \int_{CV} \rho dV + \int_{CS} \rho \vec{V} \cdot d\vec{A} = 0$$

**Finite Volumes** 

Use Gauss's theorem on the convective term to put both terms in volume form ->

$$\frac{\partial}{\partial t} \int_{CV} \rho d\mathbf{\nabla} + \int_{CV} \nabla \cdot (\rho \vec{V}) d\mathbf{\nabla} = 0$$

We can move the time rate of change term inside the integral and then write

$$\int_{CV} \frac{\partial \rho}{\partial t} dV + \int_{CV} \nabla \cdot (\rho \vec{V}) dV = 0 \Rightarrow \int_{CV} \left( \frac{\partial \rho}{\partial t} + \nabla \cdot (\rho \vec{V}) \right) dV$$

**Finite Differences** 

If we shrink the CV to an infinitesimal size we get:

$$\frac{\partial \rho}{\partial t} + \nabla \cdot (\rho \vec{V}) = 0 \quad \text{or in index notation} \quad \frac{\partial \rho}{\partial t} + \frac{\partial}{\partial x_i} (\rho u_i) = 0$$

sum repeated indexes: for  $i = 1 \rightarrow 3$ ,  $u_i u_i = u_1 u_1 + u_2 u_2 + u_3 u_3$ 

### Conservation of Momentum

Conservation of Momentum (Newton's 2<sup>nd</sup> law):

-For a system: 
$$\sum \vec{F} = \frac{d(m\vec{V})}{dt} \bigg|_{sys}$$

-Or using RTT (same as above for mass):

$$\sum \vec{F} = \frac{\partial}{\partial t} \int_{CV} \rho \vec{V} dV + \int_{CS} \vec{V} \rho \vec{V} d\vec{A}$$

What are the forces for flow in a fluid volume?

- body forces: gravity, magnetism (for magnetohydrodynamics problems)
- surface forces: **stresses**, surface tension, etc.

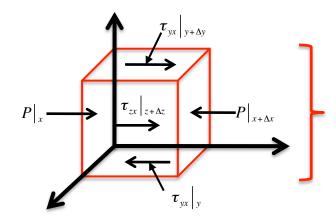
We are mostly concerned with the pressure (normal stress) and shear stresses

If we define our surface and body forces in integral forms we can write:

$$\frac{\partial}{\partial t} \int_{CV} \rho \vec{V} dV + \int_{CS} \vec{V} \rho \vec{V} d\vec{A} = \int_{CS} \mathbf{T} \cdot \hat{n} d\vec{A} + \int_{CV} \rho \vec{b} dV$$

### Conservation of Momentum

- The stresses **T** depend on molecular processes
- To use conservation of momentum we have to <u>choose a form for the stress tensor</u> **T**Recall "stress cube" for the x-direction:



**T** is a tensor (matrix) composed of all of the stress components in the x,y,z directions

The most common assumption is that the fluid is **Newtonian** (i.e. has a linear relation between stress and rate of strain)→

$$\mathbf{T} = -\left(P + \frac{2}{3}\mu\nabla\cdot\vec{V}\right)\mathbf{I} + 2\mu\mathbf{D}$$

where  $\mathbf{D} = \frac{1}{2} \left( \nabla \vec{V} + \nabla \vec{V}^T \right)$  is the deformation (rate of strain) tensor and  $\mathbf{I}$  is the unit tensor (or identity matrix)

### Conservation of Momentum

• The equivalent index-notation (differential) form of the momentum equation is:

$$\frac{\partial(\rho u_i)}{\partial t} + \frac{\partial(\rho u_i u_j)}{\partial x_j} = \frac{\partial(\tau_{ij})}{\partial x_j} - \frac{\partial(P)}{\partial x_i} + \rho g_i$$

where the stress has been split into shear (viscous) and normal (pressure) components using:

This using:
$$\tau_{ij} = 2\mu D_{ij} - \frac{2}{3}\mu \delta_{ij} \frac{\partial u_k}{\partial x_k} \quad \text{where } D_{ij} = \frac{1}{2} \left( \frac{\partial u_i}{\partial x_j} + \frac{\partial u_j}{\partial x_i} \right) \quad \text{and} \quad \delta_{ij} = \begin{cases} 1 \text{ for } i = j \\ 0 \text{ for } i \neq j \end{cases}$$

- Forms of the Momentum Equation
  - Oifferential vs. Integral
  - **@ Eulerian vs. Lagrangian**
  - © Conservative vs. Non conservative
- For the integral equations, conservative means all terms are written in the form of a divergence (e.g.,  $\nabla \cdot \vec{V}$ )
- For Finite Volume formulations, using a conservative form instead of the non conservative form guarantees the velocity field will be divergent free (a requirement for realizeability)

# Conservation of Energy

#### **Conservation of Energy**

Several forms of the conservation of energy exist (all mathematically equivalent)

- internal energy per volume
- internal energy per mass
- total energy  $(e + \frac{1}{2}u_iu_i)$
- enthalpy

For a system conservation of energy is:  $\dot{Q} - \dot{W} = \frac{dE}{dt}\Big|_{sys}$  or: (in-out) + produced = stored

- (in-out) is the convective flux of energy
- -Production is the heat conducted in + the work done on the volume (e.g., thermal flux and shear stress)
- if we use  $e = c_V T$  (specific internal energy)
- and define  $q_i = -k \frac{\partial T}{\partial x_i}$  as the thermal conductive flux where  $c_V$  is the specific heat and

*T* is temperature. We can derive the following differential form for energy (see text for integral)

$$-\frac{\partial}{\partial x_{i}} \left[ \rho u_{i} \left( e + \frac{1}{2} u_{i} u_{i} \right) \right] - \frac{\partial q_{i}}{\partial x_{i}} - \frac{\partial \left( P u_{i} \right)}{\partial x_{i}} + \frac{\partial \left( \tau_{ij} u_{i} \right)}{\partial x_{i}} = \frac{\partial}{\partial t} \left[ \rho \left( e + \frac{1}{2} u_{i} u_{i} \right) \right]$$

### Conservation of a scalar

#### Conservation of Mass (revisited):

- we can interpert density as  $\rho = \frac{\text{mass}}{\text{volume}}$
- we can use the same basic definition for any scalar quantity  $\varphi$  if we interpert  $\varphi$  as a mass of the scalar per volume (concentration)
- we than have:  $\frac{d\phi}{dt} = 0$  or using R.T.T.

$$\frac{\partial}{\partial t} \int_{CV} \phi dV + \int_{CS} \phi \vec{V} d\vec{A} = 0$$
accumulation flux

accumulation = (inflow - outflow) + generation - degradation

### Conservation of a scalar

accumulation = (inflow - outflow) + generation - degradation

$$\frac{\partial}{\partial t} \int_{CV} \phi dV = -\int_{CS} \phi \vec{V} \cdot d\vec{A} + \int_{CV} r_g dV - \int_{CV} r_d dV$$
net accumulation flux generation degradation

- $r_g$  = rate of generation of mass in the CV
- r<sub>d</sub> = rate of degradation of mass in the CV
- note if  $r_g$  =  $r_d$  (or  $r_g$  = 0 =  $r_d$  etc.) we get our standard form if  $\varphi$  =  $\rho$

# **Conservation Equations**

#### summing up our equations:

- conservation of mass
- conservation of momentum
- conservation of energy
- conservation of scalar concentration

#### **Other equations of interest**

-Inviscid flow equations in conservative form (Euler equations):

Continuity - 
$$\frac{\partial \rho}{\partial t} + \frac{\partial}{\partial x_i} (\rho u_i) = 0$$

$$\frac{\partial (\rho u_i)}{\partial t} + \frac{\partial}{\partial x_j} (u_i u_j) = -\frac{\partial P}{\partial x_i} + \rho F_i$$

$$\frac{\partial (E)}{\partial t} + \frac{\partial}{\partial x_i} (E u_i) = \rho \dot{q} - \frac{\partial u_i P}{\partial x_i} + \rho F_i u_i \quad \text{where } E = \rho \left( e + \frac{1}{2} u_i u_i \right)$$

# **CFD Equations**

#### A few Notes:

Each of these sets of equations are coupled nonlinear PDEs:

If we look at the N-S (or Euler) equations we have:

5 equations → 6 unknowns therefore we need:

• Equations of state: (for air we can use the ideal gas law):

$$P = \rho RT$$

- we still need one more equation to close the set. We need a thermodynamic relation relating the different state variables.
- usually of the form:  $e = f(T, \rho)$
- ullet for a calorically perfect gas (constant specific heat) we have:  $\ensuremath{\it e} = c_{_V} T$
- The momentum, continuity, and energy equations combined are usually referred to as the Navier-Stokes (N-S) equations

# **Boundary and Initial Conditions**

#### **Boundary and Initial Conditions:**

- in CFD we start with a physical process and then formulate a P.D.E
- we need to make sure we have a **well-posed** problem. What does this mean?
- A solution must exist and it must be unique!

We need proper Initial Conditions and Boundary Conditions

- 3 main types of B.C.s
  - u = a at the boundaries - Direchlet:
  - $\frac{\partial u}{\partial n} = a$  at the boundaries - Neumann:
  - Mixed (Robin):  $a_1 u + a_2 \frac{\partial u}{\partial n} = a_3$  at the boundaries

## More equations

- Other equations that we will discuss of interest to fluid mechanics and heat transfer:
  - 1<sup>st</sup> order linear (and nonlinear) wave equation:

$$\frac{\partial u}{\partial t} + c \frac{\partial u}{\partial x} = 0$$

• Poisson equations (and Laplace equations):

$$\nabla^2 \theta = \frac{\partial^2 \theta}{\partial x^2} + \frac{\partial^2 \theta}{\partial y^2} = f(x, y)$$

• Burger's equation:

$$\frac{\partial u}{\partial t} + u \frac{\partial u}{\partial x} = v \frac{\partial^2 u}{\partial x^2}$$
 (inviscid form is nonlinear wave)

• unsteady heat equation (Fourier's law)

$$\frac{\partial T}{\partial t} = D \frac{\partial^2 T}{\partial x^2}$$

# Classification of equations

• In general we can write the PDEs that we deal with in CFD (and many other fields) in the following form:

$$a\frac{\partial^2 \phi}{\partial x^2} + b\frac{\partial^2 \phi}{\partial x \partial y} + c\frac{\partial^2 \phi}{\partial y^2} + d\frac{\partial \phi}{\partial x} + e\frac{\partial \phi}{\partial y} = g$$

• we can classify PDEs based on how they fit into this general form

<u>Classification</u>	<u>discriminant</u>	type of problems	<u>example PDE</u>
Elliptical	$b^2 - 4ac < 0$	boundary value	Laplace equation
Parabolic	$b^2 - 4ac = 0$	marching (IVP)	Heat equation
Hyperbolic	$b^2 - 4ac > 0$	marching (IVP)	Wave equation

• for information on equation classification see Anderson chapter 3.