Cosmology HW7

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Problem 1: 4/4

part 1. 
$$O = P^{\kappa} \nabla_{\alpha} P^{\mu} = P^{\kappa} \left( \partial_{\alpha} P^{\mu} + \Gamma^{\mu}_{\alpha \beta} P^{\beta} \right) \rightarrow P^{\kappa} \partial_{\alpha} P^{\mu} = -\Gamma^{\mu}_{\alpha \beta} P^{\alpha} P^{\beta}$$

2/2

Note that  $\frac{1}{d\lambda} P^{\mu} = \frac{1}{d\lambda^{\kappa}} \frac{\partial P^{\mu}}{\partial x^{\kappa}} = P^{\kappa} \partial_{\alpha} P^{\mu} = \infty$ 
 $\lambda$  is a function if  $\eta$  above so  $\frac{1}{d\lambda} P^{\mu} = \frac{1}{d\lambda} \frac{1}{d\eta} P^{\mu} = P^{\alpha} \frac{1}{d\eta} P^{\kappa} = P^{\kappa} \partial_{\alpha} P^{\mu} = -\Gamma^{\mu}_{\alpha \beta} P^{\kappa} P^{\beta}$ .

Thus  $P^{\kappa} \frac{dP^{\mu}}{d\eta} = -\Gamma^{\mu}_{\alpha \beta} P^{\kappa} P^{\beta}$ .  $\rightarrow \mu = 0$ :  $P^{\kappa} \frac{dP^{\mu}}{d\eta} = -\Gamma^{\kappa}_{\alpha \beta} P^{\kappa} P^{\beta}$ .

 $P^{\kappa} \partial_{\alpha} P^{\mu} = P^{\kappa} P^{\mu}_{\alpha}$ , we have

 $Q = g_{\mu\nu} P^{\nu} P^{\nu} = g_{\mu\nu} (P^{\nu})^{\kappa} + 2g_{\nu\nu} P^{\nu} P^{\nu} + g_{\nu\nu} P^{\nu} P^{\nu} = -a^{\kappa} (1+2\pi) (P^{\nu})^{\kappa} + a^{\kappa} (1-24) \delta_{ij} P^{\nu} P^{i}$ 
 $S^{\kappa} \delta_{ij} P^{\nu} P^{i} = \frac{1+2\pi}{1-2+} (P^{\nu})^{\kappa} \approx [1+2(4+\pi)](P^{\nu})^{\kappa}$ 
 $S^{\kappa} \delta_{ij} P^{\nu} P^{i} = P^{\kappa} [1+2(4+\pi)](P^{\nu})^{\kappa} \approx P^{\kappa} (1+4+\pi) \rightarrow P^{\kappa} \approx \sqrt{\delta_{ij} P^{\nu} P^{i}} (1-4-\pi)$ 
 $\Gamma^{\alpha}_{\alpha\beta} P^{\mu} P^{\beta} = \Gamma^{\alpha}_{\alpha \alpha} P^{\mu} P^{\kappa} + \Gamma^{\alpha}_{\alpha \beta} P^{\kappa} P^{i} + \Gamma^{\alpha}_{\nu \beta} P^{\nu} P^{i} + \Gamma^{\alpha}_{\nu \beta} P^{\nu} P^{i} + \Gamma^{\alpha}_{\nu \beta} P^{\nu} P^{i}$ 
 $S^{\kappa} P^{\mu} P^{\mu} P^{\mu} = \Gamma^{\alpha}_{\alpha \alpha} P^{\mu} P^{\nu} P^{\mu} P^{\nu} P^{\mu} P^{\nu} P^$ 

 $\frac{dP^{\circ}}{d\eta} = \left(-2\partial \ell + \Phi' - \Psi' - 2\partial_{i}\Psi \hat{\rho}^{i}\right) P^{\circ}$ 

part 2. compute the components of the 4velocity of comoving observers:

2/2 
$$U_{com}^{\mu} = \left( U_{com}^{\circ}, \overrightarrow{O} \right)$$
 since  $dx^{i} = 0$  for comoving
$$-1 = U^{\mu}U_{\mu} = g_{\mu\nu} U^{\mu}U^{\nu} = g_{\sigma\sigma} \left( U^{\circ} \right)^{i} = -a^{2} \left( 1 + 2\Psi \right) \left( U^{\circ} \right)^{i}$$

$$\Rightarrow \left( U^{\sigma} \right)^{i} = a^{-1} \left( 1 - 2\Psi \right) \Rightarrow U^{\circ} = \frac{1}{a} \left( 1 - \Psi \right)$$

$$\rightarrow U_{com}^{\mu} = \left( \frac{1 - \Psi}{a}, \overrightarrow{O} \right).$$

photon energy measured by comoving observes is

Problem 2: code on Gi+Hub

which I think

is causily the

in our Do s.

downstream differences

can use last week's code to compute  $F_6(k) = \frac{1}{4} \delta_8(k) + \phi(k)$ ,  $V_5(k)$  at  $\eta_*$ past 1. want Sachs-Wolfe and Doppler contributions to  $D_{\ell} = T_0^2 \frac{\ell^2}{2\pi} C_{\ell}$ 3/3

 $\underline{SW} = \frac{D_{k}^{SW}}{T_{k}^{2}} \simeq |T_{F_{k}}(\eta, k = \frac{1}{2}\chi_{k})|^{2} \Delta_{R}^{L}(k = \frac{1}{2}\chi_{k})$ I think my code for this poster is  $\underline{Dapplur} \quad \frac{D_R^0}{T^2} \simeq \frac{1}{3} \left| T_{V_b}(\eta, k = \frac{9}{2} \chi_u) \right|^2 \Delta_R^2(k = \frac{9}{2} \chi_u)$ all good; hover, it lodes like my

 $F_{o}$  is a bit differen,  $F_{o}(\eta, \vec{k}) = T_{F_{o}}(\eta, k) R_{i}(\vec{k})$   $V_{b}(\eta, \vec{k}) = T_{V_{b}}(\eta, k) R_{i}(\vec{k})$ in code, we picked  $\phi_1 = 1 \text{ VK}$ , so  $R_1 = \frac{3}{2}$ 

 $\rightarrow T_{F_0}(\eta, k) = \frac{2}{3} F_0(k) \quad \text{sim for } V_b.$ 

feel free to deduce if you think it's  $\Delta_R^2 = A_s \left(\frac{k}{k_p}\right)^{n_s-1}$   $A_s = 2 \times 10^{-9}$   $n_s = 0.96$   $k_p = 0.05 \text{ Mpc}^{-1}$ nussey.

> part 2. 2/2  $\langle \Theta^2 \rangle = \sum_{\ell} \frac{2\ell+1}{U_{TT}} C_{\ell} \simeq \int_{\ell-1}^{\infty} d\ell_{0j} \ell \frac{\ell^2}{2\pi} C_{\ell}$ So I numerically integrated De/To our log l

 $\rightarrow$   $\left(\theta^{2}\right)^{1/2} \simeq 4.9 \times 10^{-5}$  which is roughly the "I part in 25,000" anisotropy

Stated on Wikipedia

$$\begin{array}{lll} & \text{Problem 3: } 5.5 \ / 7 \\ & \frac{d}{d\eta} \stackrel{?}{\eta} & \approx 2 \left( \delta^{ij} - \hat{\rho}^{i} \stackrel{?}{\varphi}^{i} \right) \partial_{j} \varphi \\ & \Delta \hat{\rho}^{i} \left( \eta_{n} \right) = 2 \left( \delta^{ij} - \hat{\rho}^{i} \stackrel{?}{\varphi}^{i} \right) \left( \int_{\eta_{n}}^{\eta_{n}} d\eta - \partial_{j} \varphi \right)_{\eta_{n}, \pi_{n} = -\kappa_{p}}^{\eta_{n}} & \mathcal{F} = \eta_{n} - \eta \\ & \text{part } 1. & \left( |\Delta \hat{\rho}^{i}|^{\lambda} \right) = \left( \Delta \hat{\rho}^{i} \Delta \hat{\rho}_{i} \right)_{\eta_{n}}^{\eta_{n}} d\eta - \partial_{j} \varphi \right)_{\eta_{n}, \pi_{n} = -\kappa_{p}}^{\eta_{n}} & \mathcal{F} = \eta_{n}^{\eta_{n}} d\eta - \partial_{j} \varphi \right)_{\eta_{n}}^{\eta_{n}} d\eta - \int_{\eta_{n}}^{\eta_{n}} d\eta - \partial_{j} \varphi \right)_{\eta_{n}}^{\eta_{n}} d\eta - \int_{\eta_{n}}^{\eta_{n}} d\eta - \partial_{j} \varphi \right)_{\eta_{n}}^{\eta_{n}} d\eta - \left( \delta^{ij} - \hat{\rho}^{i} \stackrel{?}{\varphi}^{i} \right) \partial_{j} \varphi \right)_{\eta_{n}}^{\eta_{n}} d\eta - \partial_{j} \varphi \right)_{\eta_{n}}^{\eta_{n}} d\eta - \left( \delta^{ij} - \hat{\rho}^{i} \stackrel{?}{\varphi}^{i} \right) \partial_{j} \varphi \right)_{\eta_{n}}^{\eta_{n}} \partial_{j} \varphi \right)_{\eta_{n}}^{\eta_{n}} d\eta - \left( \delta^{ij} - \hat{\rho}^{i} \stackrel{?}{\varphi}^{i} \right) \partial_{j} \varphi \right)_{\eta_{n}}^{\eta_{n}} \partial_{j} \varphi \right)_{\eta_{n}}^{\eta_{n}} d\eta - \left( \delta^{ij} - \hat{\rho}^{i} \stackrel{?}{\varphi}^{i} \right) \partial_{j} \varphi \right)_{\eta_{n}}^{\eta_{n}} \partial_{j} \varphi \right)_{\eta_{n}}^{\eta_{n}} d\eta - \left( \delta^{ij} - \hat{\rho}^{i} \stackrel{?}{\varphi}^{i} \right) \partial_{j} \varphi \right)_{\eta_{n}}^{\eta_{n}} \partial_{j} \varphi \right)_{\eta_{n}}^{\eta_{n}} d\eta - \left( \delta^{ij} - \hat{\rho}^{i} \stackrel{?}{\varphi}^{i} \right) \partial_{j} \varphi \right)_{\eta_{n}}^{\eta_{n}} \partial_{j} \varphi \right)_{\eta_{n}}^{\eta_{n}} d\eta - \left( \delta^{ij} - \hat{\rho}^{i} \stackrel{?}{\varphi}^{i} \right) \partial_{j} \varphi \right)_{\eta_{n}}^{\eta_{n}} \partial_{j} \varphi \right)_{\eta_{n}}^{\eta_{n}} d\eta - \left( \delta^{ij} - \hat{\rho}^{i} \stackrel{?}{\varphi}^{i} \right) \partial_{j} \varphi \right)_{\eta_{n}}^{\eta_{n}} \partial_{j} \varphi \right)_{\eta_{n}}^{\eta_{n}} d\eta - \left( \delta^{ij} - \hat{\rho}^{i} \stackrel{?}{\varphi}^{i} \right) \partial_{j} \varphi \right)_{\eta_{n}}^{\eta_{n}} \partial_{j} \varphi \right)_{\eta_{n}}^{\eta_{n}} d\eta - \left( \delta^{ij} - \hat{\rho}^{i} - \hat{$$

assume between y , and yo the Universe is matter dominated.

thun 
$$\Phi \sim const$$
, so  $T_{\Phi}(k, \eta) \approx T_{\Phi}(k, \eta_{\star})$ .

So  $\left\langle \left| \Delta \hat{\rho} \right|^2 \right\rangle = 4 \int d^3k \left[ k^2 - (\hat{\rho} \cdot \vec{k})^2 \right] P_R(k) \left| T_{\Phi}(k, \eta_{\star}) \right|^2 \int_{\eta_{\star}}^{\eta_0} d\eta \int_{\eta_{\star}}^{\eta_0} d\eta' \ e^{4i\vec{k} \cdot \hat{\rho}(x - x')}$ 
 $= 4 \int d^3k \left[ k^2 - (\hat{\rho} \cdot \vec{k})^2 \right] P_R(k) \left| T_{\Phi}(k, \eta_{\star}) \right|^2 \left[ (\vec{k} \cdot \hat{\rho}) \right]$ 

introduce  $\Delta_R^2(k) = \frac{k^3}{2\pi^2} P_R(k)$ .  $\rightarrow \left\langle \left| \Delta \hat{\rho} \right|^2 \right\rangle = 8\pi^2 \int d^3k \left[ k^2 - (\hat{\rho} \cdot \vec{k})^2 \right] \frac{\Delta_R^2(k)}{L^2} \left| T_{\Phi}(k, \eta_{\star}) \right|^2 \left[ (\vec{k} \cdot \hat{\rho}) \right]$ 

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part 2. \mathbb{I}(\vec{k}\cdot\hat{p}) = \int_{-\infty}^{\pi} dx \int_{-\infty}^{\pi} dx' e^{i\vec{k}\cdot\hat{p}(x'-x')}
    Let u = x' - x, v = x' + x. Then \left| \frac{\partial(u,v)}{\partial(x,x')} \right| = \left| \frac{\partial y_{xx}}{\partial y_{xx}} \frac{\partial y_{xx}}{\partial y_{xx}} \right| = \left| \frac{-1}{+1} + 1 \right| = \left| -1 - 1 \right| = 2.
                So dx dx' = \left| \frac{\partial(u,v)}{\partial(x,x')} \right|^{-1} du dv = \frac{1}{2} du dv.
     if k \gg \eta_{n}^{-1} then I(\vec{k} \cdot \hat{p}) \approx \int_{-\infty}^{+\infty} du (\chi_{\alpha} - |u|) e^{i\vec{k} \cdot \hat{p} \cdot u}
                                                       = F[ x, - |4](k)
                                                       = 2\pi \chi_{k} \delta(\vec{k} \cdot \hat{p}) + (\frac{2}{\vec{k} \cdot \hat{b}})^{2}
                                                     \approx 2\pi \gamma_0 \delta(\vec{l} \cdot \hat{p}) using x_n \approx \gamma_0, large k.
    \approx \  \, \gamma_o \  \, \left[ \  \, dk_x \  \, dk_y \  \, \frac{\Delta_R^2(L)}{L} \  \, \left| \  \, \gamma_\varphi(k,\gamma_u) \right|^2 \quad \, \text{picking coords st} \  \, \hat{\rho} = \hat{\epsilon} \, , \  \, \delta(\vec{k}\cdot\hat{\rho}) \, \, \text{enform} \, \, \vec{k} \perp \hat{\epsilon} \, \, \right]
                                        = 2\pi \eta_o \left[ dk \Delta_g^2(k) \left| \Upsilon_{\phi}(k, \eta_*) \right|^2 \right]
                                        = 2\pi \eta_o \int dk \Delta_R^{(k)} | 1_{\phi}^{(k)} \eta_e |
= 2\pi \int dk \eta_o k \Delta_R^2(k) | \gamma_{\phi}^{(k)} \eta_e |^2 correct up to the missing factor of 2 (1 used the given report)
                        for k \gtrsim k_{eq}, mode enters horizon during rad. dow. and decays as \sim \frac{1}{72} \sim \frac{1}{4}.

50, by \eta_*, \varphi is highly suppressed
0.5/2
                                               taled somewhere one pre
                                                                                                                        → To( k ≥ kay, y*) ~ 1/4 ~ 1/4
                             for k < key, mude does not evolve during sad dom. and o remains = of
                                                 \phi_1 = \frac{2}{3}R: so T_{\phi} \simeq \frac{2}{3}
                             thum \left\langle |\Delta p|^2 \right\rangle \simeq \frac{8\pi}{9} \, \eta_0 \, \int_{a}^{k_{eq}} \, d \ln k \, \, k \, \, \Delta_R^2(k)
                       =\frac{8\pi}{q}\eta_{o}\int_{0}^{k_{eq}}dk\ A_{s}\left(\frac{k}{k_{p}}\right)^{n_{s}-1}
=\frac{8\pi}{q}\eta_{o}A_{s}\frac{k_{p}}{n_{s}}\left(\frac{k_{eq}}{k_{p}}\right)^{n_{s}}
Costeen \eta to missing face of 2.

=\frac{8\pi}{q}\eta_{o}A_{s}\frac{k_{p}}{n_{s}}\left(\frac{k_{eq}}{k_{p}}\right)^{n_{s}}
Costeen \eta to missing face of 2.

A_{s}=2\times 10^{-3}, n_{s}=0.96, k_{p}=0.05 \text{ Mpc}^{-1}, \eta_{o}\approx 16 \text{ Gpc}, k_{eq}\approx 0.01 \text{ Mpc}^{-1}
k_{q}\sim 2k_{q}=\frac{\mu(q=3400)}{3400}
                                          \rightarrow \langle |\Delta p|^2 \rangle \approx 9.9 \times 10^{-7} rad \approx 0.003 aremin not order 1!
                                                                              forgot this was squared -1/2
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Problem 4: 
$$\P/\P$$

Adding  $\frac{d^3k}{(4\pi)^3} = \frac{d^3k}{(4\pi)^3}$ .

Problem 4:  $R(\vec{x}) = G(\vec{x}) + f(G(\vec{x})^3 - G^4)$ 
 $Q(\vec{x}) = F[R(\vec{x})]$ 
 $= F[G(\vec{x})] + f(F[G(\vec{x})^2] - F[1](G^4)$ 
 $= G(\vec{x}) + f(\int d^3k' G(\vec{x}) G(\vec{x})^2 - (\pi^0)^3 G(\vec{x}) G^2)$ 
 $= G(\vec{x}) + f(\int d^3k' G(\vec{x}) G(\vec{x})^2 - (\pi^0)^3 G(\vec{x}) G^2)$ 

Part 2:  $3/3$ 
 $R(\vec{k}) R(\vec{k}_1) R(\vec{k}_2) = ((G(\vec{k}_1) + fG_{M_1}G_{K_1}))(G(\vec{k}_1) + fG_{M_2}G_{K_2}) + f(G_{G_{M_1}G_{M_2}}G_{K_2})$ 
 $= (G(\vec{k}) + f(G_{M_1}G_{K_1}G_{K_2}))(G(\vec{k}_1) + fG_{M_1}G_{K_2}) + f(G_{G_{M_1}G_{M_2}}G_{K_2})$ 
 $+ f^3 (G_{M_1}G_{M_2}G_{M_2}G_{K_2}) + f(G_{G_{M_1}G_{M_2}}G_{K_2}) + f(G_{G_{M_1}G_{M_2}}G_{K_2})$ 
 $+ f^3 (G_{M_1}G_{M_2}G_{M_2}G_{M_2}) + f(G_{G_{M_1}G_{M_2}}G_{K_2}) + f(G_{G_{M_1}G_{M_2}}G_{M_2}) + f(G_{G_{M_1}G_{M_2}}G_{M_2}G_{M_2}) + f(G_{G_{M_1}G_{M_2}}G_{M_2}G_{M_2}G_{M_2}) + f(G_{G_{M_1}G_{M_2}}G_{M_2}G_{M_2}G_{M_2}G_{M_2}) + f(G_{G_{M_1}G_{M_2}}G_{M_2}G_{M_2}G_{M_2}G_{M_2}G_{M_2}G_{M_2}G_{M_2}) + f(G_{G_{M_1}G_{M_2}}G_{M_2}G_$ 

1/1

$$\begin{array}{lll} & \text{part 3.} & \mathcal{R}(\vec{k}) = G(\vec{k}) + g \cdot \mathcal{F}[G] + \mathcal{F}[G] & \text{by constitute, association of } d \cdot \pi \\ & = (\ln^{3} \left[ \int_{0}^{d} d^{3}k, \, G(k) \, G(k^{-1}k) \, G(k^{-1}k) \, - \mathcal{F}[G] & \text{by constitute, association of } d \cdot \pi \\ & = (\ln^{3} \left[ \int_{0}^{d} d^{3}k, \, G(k) \, G(k^{-1}k) \, G(k^{-1}k) \, - \mathcal{F}[G] & \text{by constitute, association of } d \cdot \pi \\ & = (\ln^{3} \left[ \int_{0}^{d} d^{3}k, \, G(k) \, G(k^{-1}k) \, G(k^{-1}k) \, - \mathcal{F}[G] & \text{by constitute, association of } d \cdot \pi \\ & = (\ln^{3} \left[ \int_{0}^{d} d^{3}k, \, G(k) \, G(k^{-1}k) \, G(k^{-1}k) \, - \mathcal{F}[G] & \text{constitute, association of } d \cdot \pi \\ & = (\ln^{3} \left[ \int_{0}^{d} d^{3}k, \, G(k) \, G(k^{-1}k) \, G(k^{-1}k) \, - \mathcal{F}[G] & \text{constitute, association of } d \cdot \pi \\ & = (\ln^{3} \left[ \int_{0}^{d} d^{3}k, \, G(k) \, G(k^{-1}k) \, - \mathcal{F}[G] \, G(k^{-1}k) \, - \mathcal{F}[G] \, - \mathcal{F}[G] \, G(k^{-1}k) \, - \mathcal{F}[G] \, - \mathcal{F}$$