

26/30

Cosmology HW6

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Problem 1: 10/14

part 1. $dU = -P dV + \delta Q$, $dU = \rho dV + V d\rho \rightarrow \delta Q = V d\rho + (\rho + P) dV$

2/2 $\frac{\delta Q}{dt} = V \dot{\rho} + (\rho + P) \dot{V}$. $V \sim a^3$ so $\frac{dV}{V} = 3 \frac{da}{a} \Rightarrow \dot{V} = 3H V$

$\Rightarrow \frac{\delta Q}{dt} = [\dot{\rho} + 3H(\rho + P)] V \Rightarrow \dot{q} = \frac{\partial}{\partial V} \frac{\delta Q}{dt} = \dot{\rho} + 3H(\rho + P)$.

$\rho = mn + \frac{3}{2} nT$, $P = nT$. $\dot{\rho} = (m + \frac{3}{2} T) \dot{n} + \frac{3}{2} n \dot{T}$. assuming $a^3 n$ is constant, $\dot{n} = -3Hn$.

thus $\dot{\rho} = -3H\rho + \frac{3}{2} n \dot{T}$. $\Rightarrow \dot{q} = \frac{3}{2} n \dot{T} + 3HnT$

rearranging, $\dot{T} + 2HT = \frac{2}{3} \frac{\dot{q}}{n}$

part 2. in e^- rest frame: initially, e^- has $\vec{v}_i = 0$, γ has \vec{p}_i
1/1 finally, e^- has \vec{v}_f , γ has \vec{p}_f

conservation of momentum: $\vec{p}_f + m_e \vec{v}_f = \vec{p}_i \Rightarrow \vec{v}_f = \frac{\vec{p}_i - \vec{p}_f}{m_e}$

assuming $|\vec{p}_i| = |\vec{p}_f|$, let $\vec{p}_i = p \hat{p}_i$, $\vec{p}_f = p \hat{p}_f$. then $\vec{v}_f = \frac{p}{m_e} (\hat{p}_i - \hat{p}_f)$

final electron energy is $E_f = \frac{1}{2} m_e \vec{v}_f^2 = \frac{1}{2} m_e \frac{p^2}{m_e^2} (\hat{p}_i - \hat{p}_f)^2 = \frac{p^2}{m_e} (1 - \cos \theta)$

since initial e^- energy is 0 and photon energy = p (up to c^2), then $\Delta E_e = \frac{E_f^2}{m_e} (1 - \cos \theta)$

part 3. a scatter with photon of energy E_γ at angle θ imparts ΔE_e on the electron

3/3 the scattering rate is $\frac{dN_s}{dt dV dE_\gamma d\cos \theta}$. the heating rate is $\dot{q} = \frac{dQ}{dt dV} = \frac{dN_s \Delta E_e}{dt dV}$

$\rightarrow \dot{q} = \int dE_\gamma d\cos \theta d\varphi \frac{dN_s}{dt dV dE_\gamma d\cos \theta} \Delta E$
 $= \int_0^\infty dE_\gamma \int_{-1}^{+1} d\cos \theta \int_0^{2\pi} d\varphi n_e \frac{d\sigma_T}{d\cos \theta} \frac{dn_\gamma}{dE_\gamma} (1 + f_\gamma(E_\gamma)) \frac{E_\gamma^2}{m_e} (1 - \cos \theta)$

components: $\frac{d\sigma_T}{d\cos \theta} = \frac{3}{16\pi} \sigma_T (1 + \cos^2 \theta)$, $f_\gamma(E_\gamma) = [\exp E_\gamma/T_\gamma - 1]^{-1}$

$n_\gamma = \frac{2}{h^3} \int d^3p f_\gamma(E(p)) = \frac{2}{h^3} \int dp 4\pi p^2 f_\gamma(E(p)) = \frac{8\pi}{h^3} \int dE E^2 f(E) \Rightarrow \frac{dn}{dE} = \frac{8\pi}{h^3} E^2 f(E)$

so $\dot{q} = \int_0^\infty dE_\gamma \int_{-1}^{+1} dc \int_0^{2\pi} d\varphi n_e \frac{3\sigma_T}{16\pi} (1 + c^2) \frac{8\pi}{h^3} E_\gamma^2 f_\gamma (1 + f_\gamma) \frac{E_\gamma^2}{m_e} (1 - c)$

$= \frac{n_e}{m_e} 2\pi \int_{-1}^{+1} dc \frac{3\sigma_T}{16\pi} (1 + c^2)(1 - c) \int_0^\infty dE \frac{8\pi}{h^3} E^4 f(1 + f)$

$= \frac{n_e \sigma_T}{m_e} \frac{8\pi}{h^3} T_\gamma^5 \int_0^\infty du u^4 e^u (e^u - 1)^{-2}$

$= \frac{n_e \sigma_T}{m_e} \frac{32\pi^5}{15 h^3} T_\gamma^5 \frac{4\pi^4/15}{4\pi^4/15}$

$= \frac{4 n_e \sigma_T}{m_e} \rho_\gamma T_\gamma$ since $\rho_\gamma = \frac{8\pi^5}{15 h^3} T_\gamma^4$

$dE_\gamma \frac{dn_\gamma}{dE_\gamma} = dE \frac{8\pi}{h^3} E^2 f(E)$

$= dp 4\pi p^2 \frac{2}{h^3} f(p)$

$= d^3p \frac{2}{h^3} f(p)$

part 4. $n_b = n_e + \underbrace{n_{H^0} + n_{H^+}}_{\equiv n_H} + \underbrace{n_{He^0} + n_{He^+} + n_{He^{++}}}_{\equiv n_{He}}$ recall $x_e \equiv \frac{n_e}{n_H}$ and $f_{He} \equiv \frac{n_{He}}{n_H}$.

thus $n_b = x_e n_H + n_H + f_{He} n_H = n_H (1 + x_e + f_{He})$

part 5. see code on GitHub. I found $z_T \approx 105$.

3/3 this means for $z > z_T$, Thomson scattering remains effective enough to keep

the baryon temperature equal to the photon temperature so $T_b \sim T_\gamma \sim a^{-1}$.

afterward, $\dot{T}_b \approx -2HT_b$ so the baryon temperature evolves

independent of the photon temperature with $T_b \sim a^{-2}$.

part 6. continuity equations: $\delta'_\gamma = -\frac{4}{3} \nabla \cdot \vec{V}_\gamma$, $\delta'_b = -\nabla \cdot \vec{V}_b$

0/4 momentum eqs: $\vec{V}'_\gamma = -\frac{1}{4} \nabla \delta_\gamma + a n_e \sigma_T (\vec{V}_b - \vec{V}_\gamma)$

$$\vec{V}'_b = -\partial \vec{V}_b - \frac{\nabla P_b}{\bar{\rho}_b} - \frac{4}{3} \frac{\bar{P}_\gamma}{\bar{\rho}_b} a n_e \sigma_T (\vec{V}_b - \vec{V}_\gamma)$$

in Fourier space: $\delta'_\gamma = -\frac{4}{3} (i\vec{k}) \cdot (i\vec{k} V_\gamma) = +\frac{4}{3} k V_\gamma$

$$\delta'_b = +k V_b$$

$$V'_\gamma = -\frac{1}{4} k \delta_\gamma + a n_e \sigma_T (V_b - V_\gamma)$$

$$V'_b = -\partial V_b - \frac{k P_b}{\bar{\rho}_b} - \frac{4}{3} \frac{\bar{P}_\gamma}{\bar{\rho}_b} a n_e \sigma_T (V_b - V_\gamma)$$

since $\Theta_0 \sim \delta$ and $\Theta_1 \sim V$, "neglecting baryon temperature fluctuations"

means $\delta_b, V_b \rightarrow 0$? using this to solve the photon equations, then

then: $\delta'_\gamma = \frac{4}{3} k V_\gamma$, $V'_\gamma = -\frac{1}{4} k \delta_\gamma - a n_e \sigma_T V_\gamma$

taking $a n_e \sigma_T \propto a^{-2}$, $\frac{d}{d\eta} (a n_e \sigma_T) = -2\partial a n_e \sigma_T$ so

$$V''_\gamma + a n_e \sigma_T V'_\gamma + \left(\frac{k^2}{3} - 2\partial a n_e \sigma_T\right) V_\gamma = 0.$$

if $\partial a n_e \sigma_T$ are slow compared to V_γ , then

$$V_\gamma = \left(\exp^{-\frac{a n_e \sigma_T}{2} \eta} \right) \left[C_1 \exp \sqrt{(a n_e \sigma_T)^2 + 8\partial a n_e \sigma_T - \frac{4}{3} k^2} \frac{\eta}{2} + C_2 \exp -\sqrt{(a n_e \sigma_T)^2 + 8\partial a n_e \sigma_T - \frac{4}{3} k^2} \frac{\eta}{2} \right]$$

$$\sim \exp \left(2\partial - \frac{k^2/3}{a n_e \sigma_T} \right) \eta$$

then $0 = -\partial V_b - \frac{\nabla P_b}{\bar{\rho}_b} + \frac{4}{3} \frac{\bar{P}_\gamma}{\bar{\rho}_b} a n_e \sigma_T V_\gamma$

$$\rightarrow \frac{\nabla P_b / \bar{\rho}_b}{\partial V_b} \approx \frac{-\partial V_b + \frac{4}{3} \frac{\bar{P}_\gamma}{\bar{\rho}_b} a n_e \sigma_T V_\gamma}{\partial V_b}$$

$$= \frac{4}{3} \frac{\bar{P}_\gamma}{\bar{\rho}_b} a n_e \sigma_T - 1$$

I didn't understand what we should assume, and it wasn't clear to me that we were computing this in the decoupled regime

(-4)

Problem 2: 4/4

$$(1) \quad \delta'_y - \frac{4}{3} k V_y = 4 \phi'$$

$$(2) \quad V'_y = -\frac{1}{4} k \delta_y + k \Pi_y - k \psi$$

$$(3) \quad \Pi'_y = -\frac{4}{15} k V_y - \frac{9}{10} a n_e \sigma_T \Pi_y$$

part 1. taking the quasistationary approximation for Π_y ($\Pi'_y \rightarrow 0$)

1/1 we get $\Pi_y \approx \left(-\frac{4}{15} k V_y\right) \left(\frac{10}{9} \frac{1}{a n_e \sigma_T}\right) = -\frac{8}{27} \frac{k}{a n_e \sigma_T} V_y$ ✓

plug into eq 2: $V'_y = -\frac{1}{4} k \delta_y + k \Pi_y - k \psi$
 $= -\frac{1}{4} k \delta_y - \frac{8}{27} \frac{k^2}{a n_e \sigma_T} V_y - k \psi$

eq 1: $\frac{4}{3} k V_y = \delta'_y - 4 \phi' \Rightarrow k V_y = \frac{3}{4} \delta'_y - 3 \phi'$

$$\Rightarrow V'_y = -\frac{1}{4} k \delta_y - \frac{8}{27} \frac{k}{a n_e \sigma_T} \left(\frac{3}{4} \delta'_y - 3 \phi'\right) - k \psi$$

$\frac{d}{d\eta}$ (eq 1): $\delta''_y - \frac{4}{3} k V'_y = 4 \phi''$

$$\Rightarrow \delta''_y + \frac{4}{3} k \left[\frac{1}{4} k \delta_y + \frac{8}{27} \frac{k}{a n_e \sigma_T} \left(\frac{3}{4} \delta'_y - 3 \phi'\right) + k \psi \right] = 4 \phi''$$

$$\delta''_y + \frac{8}{27} \frac{k^2}{a n_e \sigma_T} \delta'_y + \frac{1}{3} k^2 \delta_y = 4 \phi'' + \frac{32}{27} \frac{k^2}{a n_e \sigma_T} \phi' - \frac{4}{3} k^2 \psi$$
 ✓

part 2. seek WKB soln $\delta_y = A(\eta) e^{\pm i k \eta / \sqrt{3}} \equiv A e^{\pm i \varphi}$, assuming $\frac{A'}{A} \ll k$

3/3 $\delta = A e^{\pm i \varphi}$

$$\varphi' = \frac{k}{\sqrt{3}}, \quad \varphi'' = 0.$$

$$\delta' = \pm i \varphi' A e^{\pm i \varphi} + A' e^{\pm i \varphi}$$
 ✓

$$\delta'' = \pm i \varphi'' A e^{\pm i \varphi} \pm 2i \varphi' A' e^{\pm i \varphi} + (\pm i \varphi')^2 A e^{\pm i \varphi} + A'' e^{\pm i \varphi}$$
 ✓

plugging in: $0 = A'' \pm 2i \varphi' A' - \varphi'^2 A$

$$+ \frac{8}{27} \frac{k^2}{a n_e \sigma_T} A' \pm \frac{8}{27} \frac{k^2}{a n_e \sigma_T} i \varphi' A + \frac{1}{3} k^2 A$$

$$= A'' \pm 2i \varphi' A' + \frac{8}{27} \frac{k^2}{a n_e \sigma_T} (A' \pm i \varphi' A) + \left(\frac{k^2}{3} - \varphi'^2\right) A$$

$$= A'' \pm \frac{2ik}{\sqrt{3}} A' + \frac{8}{27} \frac{k^2}{a n_e \sigma_T} (A' \pm \frac{ik}{\sqrt{3}} A)$$

$$\sim 2\ell^2 A \quad 2\ell k A \quad \frac{k^2}{a n_e \sigma_T} (2\ell + k) A$$

to leading order, $0 = 2A' + \frac{8}{27} \frac{k^2}{a n_e \sigma_T} A$ ✓

$$\rightarrow \frac{A'}{A} = -\frac{4}{27} \frac{k^2}{a n_e \sigma_T}$$

$$\log A = -\frac{4k^2}{27} \int \frac{d\eta}{a n_e \sigma_T}$$

$$\rightarrow A(\eta) \propto \exp \left[-\frac{4}{27} k^2 \int^\eta \frac{d\eta'}{a n_e \sigma_T(\eta')} \right]$$
 ✓

Problem 3: 12/12

here I'll collect all the equations:

I think my plots
all look good!

neutrinos decoupled at all (relevant) times,

relativistic at all times (so \sim massless $\rightarrow P = \rho/3$)

and no anisotropic stress ($\rightarrow \Pi = 0, c_s^2 = 1/3$)

then: $\delta'_v = \frac{4}{3} k V_v + 4\phi'$ ✓

$V'_v = -\frac{1}{4} k \delta_v - k\phi$ ✓

cdm pressureless perfect fluid, decoupled at all times

$\delta'_c = k V_c + 3\phi'$ ✓

$V'_c = -\mathcal{H} V_c - k\phi$ ✓

photons & baryons for $z > z_d = 10^3$, tight coupling

$\delta'_\gamma = \frac{4}{3} k V_\gamma + 4\phi'$ ✓

$V'_\gamma = -\frac{k}{4(1+R)} \delta_\gamma - k\phi - \frac{R}{1+R} \mathcal{H} V_\gamma$ ✓ $R \equiv \frac{3}{4} \frac{\bar{\rho}_b}{\bar{\rho}_\gamma}$ ✓

$\delta_b = \frac{3}{4} \delta_\gamma, V_b = V_\gamma$ ✓

for $z < z_d$, total decoupling

photons assume $\Pi_\gamma = 0, c_s^2 = 1/3$

$\delta'_\gamma = \frac{4}{3} k V_\gamma + 4\phi'$ ✓

$V'_\gamma = -\frac{1}{4} k \delta_\gamma - k\phi$ ✓

baryons assume $P_b = 0$. note $\Pi_b = 0$ always.

$\delta'_b = k V_b + 3\phi'$ ✓

$V'_b = -\mathcal{H} V_b - k\phi$ ✓

potential evolves under $\phi'' + 3(1+w)\mathcal{H}\phi' + w k^2 \phi = 4\pi G a^2 (\delta\rho - w \delta p)$ $w \equiv \bar{P}/\bar{\rho}$.

want to split into two first-order ODEs for dimensionless variables.

let $\varphi_1 = \phi, \varphi_2 = \frac{d\phi}{da}$. then

$\phi' = \frac{d\phi}{d\eta} = \frac{da}{d\eta} \frac{d\phi}{da} = \mathcal{H} a \varphi_2$.

$\phi'' = \frac{d}{d\eta} (\mathcal{H} a \varphi_2) = \mathcal{H}' a \varphi_2 + \mathcal{H}^2 a \varphi_2 + \mathcal{H} a \varphi_2' = \frac{1-3w}{2} \mathcal{H}^2 a \varphi_2 + \mathcal{H} a \varphi_2'$

then our system of first-order ODEs is

$$\varphi_1' = 2\ell a \varphi_2$$

$$\varphi_2' = -\frac{7+3w}{2} 2\ell \varphi_2 - w \frac{k^2}{2\ell a} \varphi_1 + 4\pi G \frac{a}{2\ell} (\delta\rho - w\rho)$$

it's more convenient to consider $\log a$ the integration variable

$$\begin{aligned} \text{note } \frac{d}{d\eta} &= \frac{da}{d\eta} \frac{d}{da} \\ &= \frac{da}{d\eta} \frac{d \log a}{da} \frac{d}{d \log a} \\ &= 2\ell a \frac{1}{a} \frac{d}{d \log a} \\ &= 2\ell \frac{d}{d \log a} \end{aligned}$$