Problem 1:

part 2. compute the components of the 4velocity of comoving observers:

$$\begin{array}{lll} U_{com}^{\mu} = \left(\begin{array}{ccc} U_{com}^{\circ} & \overrightarrow{\circ} \end{array} \right) & \text{since } dx^{i} = 0 & \text{for comoving} \\ -1 & = & U^{\mu} U_{\mu} & = g_{\mu\nu} & U^{\mu} U^{\nu} & = g_{oo} \left(U^{\circ} \right)^{i} & = -a^{2} \left(1 + 2\Psi \right) \left(U^{\circ} \right)^{i} \\ & \Rightarrow \left(U^{\sigma} \right)^{i} = a^{-1} \left(1 - 2\Psi \right) & \Rightarrow U^{\circ} = \frac{1}{a} \left(1 - \Psi \right) \\ & \rightarrow & U_{com}^{\mu} = \left(\frac{1 - \Psi}{a} \right) & \overrightarrow{\circ} \end{array} \right). \end{array}$$

photon energy measured by comoving observes is

≈ (¢' - ĵi à; 4) p

Problem 2: code on GitHub

part 1. can use last week's code to compute
$$F_0(k) = \frac{1}{4} \delta_K(k) + \phi(k)$$
, $V_0(k)$ at γ_* want Sachs-Wolfe and Doppler contributions to $D_R = T_0^2 \frac{\ell^2}{2\pi} C_R$

$$\frac{SW}{T_0^2} \simeq \left| T_{F_0}(\gamma, k = \ell/\gamma_*) \right|^2 \Delta_R^2(k = \ell/\gamma_*)$$

$$\frac{Doppler}{T_0^2} \simeq \frac{1}{3} \left| T_{V_0}(\gamma, k = \ell/\gamma_*) \right|^2 \Delta_R^2(k = \ell/\gamma_*)$$

$$\begin{split} F_o\left(\eta,\vec{k}\right) &= T_{F_o}\left(\eta,k\right) \; R_i(\vec{k}) & V_b\left(\eta,\vec{k}\right) = T_{V_b}\left(\eta,k\right) \; R_i\left(\vec{k}\right) \\ \text{in code, we picked } \varphi_i &= 1 \; \forall \vec{k}, \; \text{so } \; R_i = \frac{3}{2} \\ &\rightarrow T_{F_o}\left(\eta,k\right) = \frac{2}{3} \; F_o\left(k\right) \; , \; \; \text{sim for V_b} \; . \end{split}$$

$$\Delta_R^2 = A_s \left(\frac{k}{k_p}\right)^{n_s-1}$$
 $A_s = 2*10^{-9}, \quad n_s = 0.96, \quad k_p = 0.05 \text{ Mpc}^{-1}$

past 2.

$$\left\langle \Theta^{2} \right\rangle = \sum_{l} \frac{2l+1}{4\pi} C_{l} \approx \int_{lmin}^{\infty} d\log l \frac{l^{2}}{2\pi} C_{l}$$

So I numerically integrated D_{l} / T_{0}^{2} over log l

$$= \frac{\left\langle \Theta^{2} \right\rangle^{1/2}}{\left\langle \Theta^{2} \right\rangle^{1/2}} \approx \frac{4.9 \times 10^{-5}}{10^{-5}} \quad \text{which is coughly the "1 part in 25,000" suisotropy}$$
Stated on Wikipedia

Problem 3:

$$\begin{array}{lll} \frac{1}{d\eta} \stackrel{\circ}{\eta^{i}} \approx & 2 \left(\delta^{ij} - \stackrel{\circ}{\eta^{i}} \stackrel{\circ}{\eta^{j}} \right) \partial_{j} \varphi \\ & \Delta_{j}^{\lambda}(\eta_{o}) = 2 \left(\delta^{ij} - \stackrel{\circ}{\eta^{i}} \stackrel{\circ}{\eta^{j}} \right) \int_{\eta_{o}}^{\eta_{o}} d\eta \quad \partial_{j} \varphi \Big|_{\eta_{o}, \pi^{i} = -\chi_{0}^{i}} \\ & \gamma = \eta_{o} - \eta \end{array}$$

$$\begin{array}{lll} \text{part } I. & \left(|\Delta_{j}^{\lambda}|^{2} \right) = \left(\Delta_{j}^{\lambda} |\Delta_{j}^{\lambda}| \right) = I \left(\delta^{ij} - \stackrel{\circ}{\eta^{i}} \stackrel{\circ}{\eta^{j}} \right) \left(\delta_{ik} - \stackrel{\circ}{\eta^{i}} \right) \left(\delta_{ik} - \stackrel{\circ}{\eta^{i}} \stackrel{\circ}{\eta^{j}} \right) \left(\delta_{ik} - \stackrel{\circ}{\eta^{i}} \right) \left(\delta_{ik} - \stackrel{\circ}$$

assume between y , and yo the Universe is matter dominated.

then $\phi \sim const$, so $T_{\phi}(k, \gamma) \approx T_{\phi}(k, \gamma_{*})$.

so
$$\langle |\Delta \hat{P}|^2 \rangle = 4 \int d^3k \left[k^2 - (\hat{p} \cdot \vec{k})^2 \right] P_R(k) \left| \Upsilon_{\phi}(k, \eta_{\alpha}) \right|^2 \int_{\eta_{\alpha}}^{\eta_{\alpha}} d\eta \int_{\eta_{\alpha}}^{\eta_{\alpha}} d\eta' e^{4i\vec{k} \cdot \hat{p}(\gamma - \gamma')}$$

$$= 4 \int d^3k \left[k^2 - (\hat{p} \cdot \vec{k})^2 \right] P_R(k) \left| \Upsilon_{\phi}(k, \eta_{\alpha}) \right|^2 I(\vec{k} \cdot \hat{p})$$

introduce
$$\Delta_{R}^{2}(k) = \frac{k^{3}}{2\pi^{3}}P_{R}(k)$$
. $\rightarrow \left\langle \left| \Delta \hat{p} \right|^{2} \right\rangle = 8\pi^{2} \int d^{3}k \left[k^{2} - (\hat{p} \cdot \vec{k})^{2} \right] \frac{\Delta_{R}^{2}(k)}{k^{3}} \left| \Upsilon_{\Phi}(k, \eta_{\pi}) \right|^{2} I(\vec{k} \cdot \hat{p})$

part 2.
$$I(\vec{k},\hat{p}) = \int_{x_{0}}^{x_{0}} dy \int_{x_{0}}^{x_{0}} dy' = \int_{x_{0}}^{x_{0}} dy' = \int_{x_{0}}^{x_{0}} (x'-y) dy' =$$

 $\rightarrow \langle |\Delta p|^2 \rangle \approx 9.9 \times 10^{-7}$ and ≈ 0.003 aremin not order 1!

$$\begin{array}{lll} & \text{part 3.} & \mathcal{R}(\overline{k}^{2}) = G(\overline{k}) + g \cdot \mathcal{F}[G^{2}] \\ & \mathcal{F}[G^{3}] = (\pi h^{3} + \mathbb{F}[G) + \mathbb{F}[G] + \mathbb{F}[G] + \mathbb{F}[G] & \text{by convisions}, \text{ detection in the } g + n \\ & = (\pi h^{3})^{2} \left[\int g^{3}k_{1} \cdot G(\overline{k}_{1}) G(\overline{k}_{1} - \overline{k}_{1}) G(\overline{k}_{1} - \overline{k}_{1}) \\ & = \int g^{3}k_{1} \cdot g^{3}k_{2} \cdot G(f_{1}) G(\overline{k}_{1} - \overline{k}_{1}) G(\overline{k}_{1} - \overline{k}_{1}) \\ & \to & \mathcal{R}(\overline{k}^{2}) = G(\overline{k}) + g \cdot \int g^{3}k_{1} \cdot g^{3}k_{2} \cdot G(\overline{k}) G(\overline{k}_{1} - \overline{k}_{1}) G(\overline{k}_{1} - \overline{k}_{1}) \\ & \to & \mathcal{R}(\overline{k}^{2}) = G(\overline{k}) + g \cdot \int g^{3}k_{1} \cdot g^{3}k_{2} \cdot G(\overline{k}) G(\overline{k}_{1} - \overline{k}_{1}) G(\overline{k}_{1} - \overline{k}_{1}) G(\overline{k}_{1} - \overline{k}_{1}) \\ & \to & \mathcal{R}(\overline{k}^{2}) = G(\overline{k}) + g \cdot \int g^{3}k_{1} \cdot g^{3}k_{2} \cdot G(\overline{k}) G(\overline{k}_{1} - \overline{k}_{1}) G(\overline{$$