## Extragalactic Astrophysics HW7

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## Dynamics, Analytic Exercise #5.

Using the spherical symmetric first-order Jeans Equation,

$$\partial_r(n\sigma_{rr}^2) + \frac{n}{r} \left[ 2\sigma_{rr}^2 - (\sigma_{\theta\theta}^2 + \sigma_{\phi\phi}^2) \right] = -n\partial_r \Phi, \tag{1}$$

show that

$$M(\langle r) = -\frac{r\sigma_{rr}^2}{G} \left[ \frac{\mathrm{d}\ln n}{\mathrm{d}\ln r} + \frac{\mathrm{d}\ln \sigma_{rr}^2}{\mathrm{d}\ln r} + 2\beta \right]. \tag{2}$$

Note that, due to spherical symmetry, n and  $\sigma_{rr}^2$  and  $\Phi$  are functions only of r, and thus we can switch between partial and total derivatives w.r.t. r without an issue.

Let's begin from the Jeans Equation. Under spherical symmetry, we have  $\sigma_{\theta\theta}^2 = \sigma_{\phi\phi}^2$ , so

$$\partial_r(n\sigma_{rr}^2) + \frac{2n}{r} \left(\sigma_{rr}^2 - \sigma_{\theta\theta}^2\right) = -n\partial_r \Phi. \tag{3}$$

We can also introduce the anisotropy parameter  $\beta \equiv 1 - \sigma_{\theta\theta}^2/\sigma_{rr}^2$ . Plugging this in simplifies the Jeans Equation to

$$\partial_r(n\sigma_{rr}^2) + \frac{2n}{r}\sigma_{rr}^2\beta = -n\partial_r\Phi. \tag{4}$$

Now we want to derive an expression for the total enclosed mass, which requires us consider the relation between the gravitational potential  $\Phi$  and the mass density  $\rho$ . This relation is

$$\nabla^2 \Phi = 4\pi G \rho. \tag{5}$$

The enclosed mass can then be written in terms of  $\Phi$  as

$$M(< r) = \int_0^r dr' \ 4\pi r'^2 \rho \tag{6}$$

$$= \int_0^r \mathrm{d}r' \, \frac{r'^2}{G} \nabla'^2 \Phi \tag{7}$$

$$= \frac{1}{G} \int_0^r dr' \, \partial_{r'} r'^2 \partial_{r'} \Phi \tag{8}$$

$$=\frac{r^2}{G}\partial_r\Phi. (9)$$

Now  $\partial_r \Phi$  is a term in the Jeans Equation, which we can solve for and plug in. This gives

$$M(\langle r) = -\frac{1}{G} \left[ 2r\sigma_{rr}^2 \beta + \frac{r^2}{n} \partial_r (n\sigma_{rr}^2) \right]$$
 (10)

$$= -\frac{r\sigma_{rr}^2}{G} \left[ 2\beta + \frac{r}{n\sigma_{rr}^2} \partial_r (n\sigma_{rr}^2) \right]. \tag{11}$$

Finally, we can identify the logarithmic derivative:

$$\frac{\mathrm{d}\ln f}{\mathrm{d}\ln x} = \frac{\mathrm{d}x}{\mathrm{d}\ln x} \frac{\mathrm{d}\ln f}{\mathrm{d}x} = \frac{x}{f} \frac{\mathrm{d}f}{\mathrm{d}x}.$$
 (12)

Making this replacement yields

$$M(\langle r) = -\frac{r\sigma_{rr}^2}{G} \left[ 2\beta + \frac{\mathrm{d}\ln(n\sigma_{rr}^2)}{\mathrm{d}\ln r} \right]$$
 (13)

or, expanding using log rules,

$$M(\langle r) = -\frac{r\sigma_{rr}^2}{G} \left[ 2\beta + \frac{\mathrm{d}\ln n}{\mathrm{d}\ln r} + \frac{\mathrm{d}\ln\sigma_{rr}^2}{\mathrm{d}\ln r} \right],\tag{14}$$

as desired.