

# Cosmological Perturbations

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Most of the conventions and notations in [6, ch. 5] will be followed.  
Suppose the metric can be expanded up to the linear order as

$$g = g^{(0)} + \epsilon g^{(1)} + O(\epsilon^2). \quad (1)$$

The background metric  $g^{(0)}$  takes the Robertson–Walker form

$$g_{\mu\nu}^{(0)} dx^\mu dx^\nu = -N^2(t) dt^2 + a^2(t) d\Omega_{3F}^2, \quad (2)$$

in which  $d\Omega_{3F}^2 = d\chi^2 + \chi^2(d\theta^2 + \sin^2\theta d\phi^2)$  is the dimensionless flat spatial metric. The linear perturbation can be decomposed into scalar, vector and tensor parts

$$g_{00}^{(1)} = -E, \quad (3)$$

$$g_{i0}^{(1)} = g_{0i}^{(1)} = F_{,i} + G_i, \quad (4)$$

$$g_{ij}^{(1)} = A\delta_{ij} + B_{,i,j} + C_{i,j} + C_{j,i} + D_{ij}. \quad (5)$$

Here one has some weird condition, where the contractions do not follow the one-up-one-down tradition

$$C_{i,i} = G_{i,i} = 0, \quad D_{ij,i} = 0, \quad D_{ii} = 0. \quad (6)$$

## 1 Metric perturbation under diffeomorphism

Consider a diffeomorphism generated by  $\xi^\mu$

$$x^\mu \rightarrow \bar{x}^\mu = x^\mu - \epsilon \xi^\mu. \quad (7)$$

The generator  $\xi^\mu$  can in turn be decomposed into  $\xi_0 = \zeta$ ,  $\xi_i = \xi_{,i}^S + \xi_i^V$ .

One has here again some weird condition, where the contractions do not follow the one-up-one-down tradition

$$\xi_{i,i}^V = 0. \quad (8)$$

The Lie derivative of the metric  $\mathbb{L}_\xi g$  is

$$(\mathbb{L}_\xi g)_{\mu\nu} = \xi^\lambda g_{\mu\nu,\lambda} + \xi^\lambda_{,\mu} g_{\lambda\nu} + \xi^\lambda_{,\nu} g_{\mu\lambda}. \quad (9)$$

In components and expansion, these are

$$(\mathbb{L}_\xi g)_{00} = 2\dot{\zeta} - 2\zeta \frac{\dot{N}}{N} + O(\epsilon), \quad (10)$$

$$(\mathbb{L}_\xi g)_{i0} = (\mathbb{L}_\xi g)_{0i} = \left( \zeta - 2\frac{\dot{a}}{a} \xi^S + \dot{\xi}^S \right)_{,i} + \left( -2\frac{\dot{a}}{a} \xi_i^V + \dot{\xi}_i^V \right) + O(\epsilon), \quad (11)$$

$$(\mathbb{L}_\xi g)_{ji} = (\mathbb{L}_\xi g)_{ij} = -\frac{2a\dot{a}}{N^2} \zeta \delta_{ij} + 2\xi_{,i,j}^S + \xi_{i,j}^V + \xi_{j,i}^V + O(\epsilon). \quad (12)$$

## 2 Scalar perturbations

$$-N^2 - \epsilon E + O(\epsilon^2) \rightarrow -N^2 - \epsilon E + \epsilon \left( 2\dot{\zeta} - 2\zeta \frac{\dot{N}}{N} \right) + O(\epsilon^2), \quad (13)$$

so one can write

$$\mathbb{L}_\xi E = -2\dot{\zeta} + 2\zeta \frac{\dot{N}}{N}. \quad (14)$$

Similarly one can read-off

$$\mathbb{L}_\xi F = \zeta - 2\frac{\dot{a}}{a} \xi^S + \dot{\xi}^S, \quad (15)$$

$$\mathbb{L}_\xi A = -\frac{2a\dot{a}}{N^2} \zeta, \quad (16)$$

$$\mathbb{L}_\xi B = 2\xi^S. \quad (17)$$

The four scalar perturbations are generated by  $\zeta$  and  $\xi^S$ , so that only two independent perturbations exists. It is clear that

$$\mathbb{L}_\xi \left( \frac{F}{a} - \frac{\mathbb{D}}{\mathbb{D}t} \frac{B}{2a} \right) = \frac{\zeta}{a}. \quad (18)$$

One can verify that

$$\mathbb{L}_\xi \left\{ \frac{E}{2N} + \frac{\mathbb{D}}{\mathbb{D}t} \left[ \frac{a}{N} \left( \frac{F}{a} - \frac{\mathbb{D}}{\mathbb{D}t} \frac{B}{2a} \right) \right] \right\} = 0, \quad (19)$$

$$\mathbb{L}_\xi \left\{ \frac{A}{2} + \frac{a^2 \dot{a}}{N^2} \left( \frac{F}{a} - \frac{\mathbb{D}}{\mathbb{D}t} \frac{B}{2a} \right) \right\} = 0. \quad (20)$$

### 3 Vector perturbations

### 4 Tensor perturbations

### 5 Scalar field perturbation under diffeomorphism

### 6 Perturbation in Arnowitt–Deser–Misner Hamiltonian formalism

Up to boundary terms, the Hamiltonian action for General Relativity in terms of Arnowitt–Deser–Misner variables is [3, ch.4.2.2]

$$S = \int dt dx^3 \left\{ \mathfrak{p}^{ij} \dot{h}_{ij} + \mathfrak{P} \dot{N} + \mathfrak{P}^i \dot{N}_i - N \mathfrak{H}^\perp - N_i \mathfrak{H}^i - \mathfrak{P} V - \mathfrak{P}^i V_i \right\}, \quad (21)$$

$$\mathfrak{H}^\perp = 2\mathscr{K} \mathfrak{G}_{ijkl} \mathfrak{p}^{ij} \mathfrak{p}^{kl} - \frac{\sqrt{\mathfrak{h}}}{2\mathscr{K}} R[h] = 2\mathscr{K} \mathfrak{F}^{ijkl} h_{ij} h_{kl} - \frac{\sqrt{\mathfrak{h}}}{2\mathscr{K}} R[h], \quad (22)$$

$$\mathfrak{H}^i = -2\mathfrak{p}^{ij}{}_{|j}, \quad (23)$$

$$\mathfrak{G}_{ijkl} := \frac{1}{2\sqrt{\mathfrak{h}}} (h_{ik} h_{jl} + h_{il} h_{kj} - h_{ij} h_{kl}), \quad (24)$$

$$\mathfrak{F}^{ijkl} := \frac{1}{2\sqrt{\mathfrak{h}}} (\mathfrak{p}^{ik} \mathfrak{p}^{jl} + \mathfrak{p}^{il} \mathfrak{p}^{kj} - \mathfrak{p}^{ij} \mathfrak{p}^{kl}), \quad (25)$$

where  $V$  and  $V_i$  are velocities of  $N$  and  $N_i$  and play the role of Lagrange multipliers. Technical details about the boundary terms can be found in [5, ch. 4.2] and the references therein. Note that  $\{N, N_i, h_{ij}, \mathfrak{P}, \mathfrak{P}^i, \mathfrak{p}^{ij}\}$  are not the unique choice of canonical variables for General Relativity in Hamiltonian formalism; instead, they are a special parametrisation of the phase space. One can also choose the components of the original four-metric and their conjugate momenta  $\{g_{\mu\nu}, \mathfrak{p}^{\mu\nu}\}$  as canonical variables, as Dirac has done in [2]. The two approaches are different in some subtle aspects; see [4] for a comparison.

Gauge transformations in the Arnowitt–Deser–Misner canonical variables are generated by [1]

$$G = - \int d^3x \left\{ \left[ \xi_\perp \left( \mathfrak{H}^\perp + N_{|i} \mathfrak{P}^i + (N \mathfrak{P}^i)_{|i} + (N_i \mathfrak{P})^{|i} \right) + \dot{\xi}_\perp \mathfrak{P} \right] \right. \\ \left. + \left[ \xi_i \left( \mathfrak{H}^i + N_j^{|i} \mathfrak{P}^j + (N_j \mathfrak{P}^i)^{|j} + N^{|i} \mathfrak{P} \right) + \dot{\xi}_i \mathfrak{P}^i \right] \right\}. \quad (26)$$

Possible boundary terms have not been discussed so far. The infinitesimal gauge transformation of  $\{N, N_i\}$  is

$$\delta N = [N, G]_{\text{P}} = \xi_\perp^{|i} N_{z|i} - \dot{\xi}_\perp - \xi_i N^{|i}, \quad (27)$$

$$\delta N_i = -\xi_\perp N_{|i} + \xi_\perp^{|i} N - \xi_j N_i^{|j} + \xi_i^{|j} N_j - \dot{\xi}_i, \quad (28)$$

which can be found in [4]. Transformations for  $g_{ij}$  and the momenta have to be worked out as

$$\delta \mathfrak{P} = -(\xi_{\perp} \mathfrak{P}^i)_{|i} - \xi_{\perp|i} \mathfrak{P}^i - (\xi_i \mathfrak{P})^{|i}, \quad (29)$$

$$\delta \mathfrak{P}^i = -\xi_{\perp|i} \mathfrak{P} - (\xi_j \mathfrak{P}^i)^{|j} - \xi_j^{|i} \mathfrak{P}^j, \quad (30)$$

where only the primary constraints are involved;

$$\begin{aligned} \delta g_{ij} &= -\frac{\partial}{\partial \mathfrak{p}^{ij}} (\xi_{\perp} \mathfrak{H}^{\perp} + \xi_i \mathfrak{H}^i) \\ &= -\xi^{\perp} \frac{2\mathfrak{H}}{\sqrt{\mathfrak{h}}} (h_{ik} h_{jl} + h_{il} h_{kj} - h_{ij} h_{kl}) \mathfrak{p}^{kl} - \xi_{i|j} - \xi_{j|i}, \quad (31) \\ \delta \mathfrak{p}^{ij} &= \frac{\partial}{\partial g_{ij}} (\xi_{\perp} \mathfrak{H}^{\perp} + \xi_i \mathfrak{H}^i) \\ &= \xi_{\perp} \left\{ \frac{\mathfrak{H}}{\sqrt{\mathfrak{h}}} \left[ (4\delta^i_k \delta^j_m - h^{ij} h_{km}) h_{ln} \right. \right. \\ &\quad \left. \left. - \frac{1}{2} (4\delta^i_k \delta^j_l - h^{ij} h_{kl}) h_{mn} \right] \mathfrak{p}^{kl} \mathfrak{p}^{mn} \right. \\ &\quad \left. + \frac{\sqrt{\mathfrak{h}}}{2\mathfrak{H}} G^{ij}[h] \right\} + \frac{\sqrt{\mathfrak{h}}}{2\mathfrak{H}} (\xi_{\perp|k}^{|k} h^{ij} - \xi_{\perp}^{|(i|j)}) \\ &\quad - (\xi_k \mathfrak{p}^{ij})^{|k} + 2\xi_{k|i} h^{k(i} \mathfrak{p}^{j)l}, \quad (32) \end{aligned}$$

where  $G^{ij}[h] = R^{ij}[h] - h^{ij} R[h]/2$ , and only the secondary constraints are involved. In eq. (32), the first two lines come from the variation of the ‘kinetic’ term in  $\mathfrak{H}^{\perp}$ , the third comes from the ‘potential’ term in  $\mathfrak{H}^{\perp}$ , and the last line from  $\mathfrak{H}^i$ . The results can be checked with [5, p. 4.2.7].

## 6.1 Expansion of the action with fluctuations

### Some useful results

#### First variations

First variation of  $h^{ij}$

$$\delta h^{ij} = -h^{ik} h^{jl} \delta h_{kl} = -h^{i(k} h^{l)j} \delta h_{kl}. \quad (33)$$

First variation of  $\mathfrak{h} = \det h_{ij}$

$$\delta \mathfrak{h} = \mathfrak{h} h^{ij} \delta h_{ij}. \quad (34)$$

First variation of  $\Gamma^i_{jk}$

$$\delta \Gamma^i_{jk} = \frac{1}{2} h^{il} \left\{ -(\delta h_{jk})_{|l} + (\delta h_{kl})_{|j} + (\delta h_{lj})_{|k} \right\} \quad (35)$$

$$= \frac{1}{2} \left\{ -h^{il} \delta^m_j \delta^n_k + h^{in} \delta^l_j \delta^m_k + h^{im} \delta^n_j \delta^l_k \right\} (\delta h_{mn})_{|l}. \quad (36)$$

First variation of  $R_{ij}[h]$  and  $R^{ij}[h]$

$$\delta R_{ij}[h] = (\delta \Gamma^k_{ij})_{|k} - (\delta \Gamma^k_{ik})_{|j}, \quad (37)$$

$$\delta R^{ij}[h] = -2R^{ki}h^{jl}\delta h_{kl} + h^{ik}\bar{\delta}u^{jl}_{k|l}, \quad (38)$$

where

$$\bar{\delta}u^{ij}_k := h^{il}\delta\Gamma^j_{kl} - h^{ij}\delta\Gamma^l_{kl}. \quad (39)$$

Equation (37) can be obtained by using normal coordinates.

First variation of  $\sqrt{h}R[h]$

$$\delta(\sqrt{h}R[h]) = \sqrt{h}\{-G^{ij}[h]\delta h_{ij} + \bar{\delta}u^{ji}_{j|i}\}. \quad (40)$$

First variation of  $\mathfrak{G}_{ijkl}\mathfrak{p}^{ij}\mathfrak{p}^{kl} \equiv \mathfrak{F}^{ijkl}h_{ij}h_{kl}$

$$\begin{aligned} \delta(\mathfrak{G}_{ijkl}\mathfrak{p}^{ij}\mathfrak{p}^{kl}) &\equiv \delta(\mathfrak{F}^{ijkl}h_{ij}h_{kl}) \\ &= \delta h_{ij}\left(-\frac{1}{2}h^{ij}\mathfrak{F}^{klmn}h_{mn} + 2\mathfrak{F}^{ijkl}\right)h_{kl} + \delta\mathfrak{p}^{ij}2\mathfrak{G}_{ijkl}\mathfrak{p}^{kl}. \end{aligned} \quad (41)$$

First variation of  $\mathfrak{H}^\perp$

$$\begin{aligned} \delta\mathfrak{H}^\perp &= \delta h_{ij}\left\{2\mathcal{N}\left(-\frac{1}{2}h^{ij}\mathfrak{F}^{klmn}h_{mn} + 2\mathfrak{F}^{ijkl}\right)h_{kl}\right. \\ &\quad \left.+ \frac{\sqrt{h}}{2\mathcal{N}}\left(G^{ij}[h] - \bar{\delta}u^{ji}_{j|i}\right)\right\} + \delta\mathfrak{p}^{ij}4\mathcal{N}\mathfrak{G}_{ijkl}\mathfrak{p}^{kl}, \end{aligned} \quad (42)$$

where the term with a  $\sqrt{h}/2\mathcal{N}$  factor comes from variation of the three Ricci scalar.

First variation of  $\mathfrak{H}^i$

$$\delta\mathfrak{H}^i = (\delta p^{ij})_{|j} + \delta\Gamma^i_{jk}p^{jk}. \quad (43)$$

First variation of  $\mathfrak{F}^{ijkl}h_{kl}$

$$\begin{aligned} \delta(\mathfrak{F}^{ijkl}h_{kl}) &= \delta h_{kl}\left(-\frac{1}{2}\mathfrak{F}^{ijmn}h^{kl}h_{mn} + \mathfrak{F}^{ijkl}\right) \\ &\quad + \delta\mathfrak{p}^{kl}(\delta^i_k\mathfrak{G}^j_{lmn} + \delta^i_m\mathfrak{G}^j_{nkl})\mathfrak{p}^{mn}. \end{aligned} \quad (44)$$

First variation of  $\mathfrak{G}_{ijkl}\mathfrak{p}^{kl}$

$$\begin{aligned} \delta(\mathfrak{G}_{ijkl}\mathfrak{p}^{kl}) &= \delta h_{kl}\left(-\frac{1}{2}\mathfrak{G}_{ijmn}h^{kl} + \delta^k_i\mathfrak{G}^l_{jmn} + \delta^k_m\mathfrak{G}^l_{nij}\right)\mathfrak{p}^{mn} \\ &\quad + \delta p^{kl}\mathfrak{G}_{ijkl}. \end{aligned} \quad (45)$$

First variation of  $\sqrt{\mathfrak{h}} G^{ij}[h]$

$$\begin{aligned} \delta(\sqrt{\mathfrak{h}} G^{ij}[h]) &= \sqrt{\mathfrak{h}} \left\{ \delta h_{kl} \right. \\ &\quad \left( -\frac{1}{2} \right) (R^{ik} h^{lj} + R^{il} h^{kj} - R^{ij} h^{kl} + h^{ik} G^{lj} + h^{il} G^{kj} - h^{ij} G^{kl}) \\ &\quad \left. + \left( h^{il} \bar{\delta} u^{jk}{}_{|l} - \frac{1}{2} h^{ij} \bar{\delta} u^{lk}{}_{|k} \right) \right\}. \end{aligned} \quad (46)$$

## Second variations

Second variation of  $\mathfrak{H}^\perp$

$$\begin{aligned} &\delta^2(\mathfrak{H}^\perp) \\ &= \delta h_{ij} \delta h_{kl} \left\{ 2\mathscr{N} \left[ \frac{1}{4} (h^{ik} h^{lj} + h^{il} h^{kj} + h^{ij} h^{kl}) \mathfrak{F}^{mnrs} h_{mn} h_{rs} \right. \right. \\ &\quad \left. \left. - (h^{ij} \mathfrak{F}^{klmn} + \mathfrak{F}^{ijmn} h^{kl}) h_{mn} + \mathfrak{F}^{ijkl} \right] \right. \\ &\quad \left. - \frac{\sqrt{\mathfrak{h}}}{4\mathscr{N}} (R^{ik} h^{lj} + R^{il} h^{kj} - R^{ij} h^{kl} + h^{ik} G^{lj} + h^{il} G^{kj} - h^{ij} G^{kl}) \right\} \\ &\quad + \delta h_{ij} \frac{\sqrt{\mathfrak{h}}}{2\mathscr{N}} \left( h^{il} \bar{\delta} u^{jk}{}_{|l} - \frac{1}{2} h^{ij} \bar{\delta} u^{lk}{}_{|l} \right)_{|k} \\ &\quad + \delta h_{ij} \delta \mathfrak{p}^{kl} 4\mathscr{N} \{ -h^{ij} \mathfrak{G}_{klmn} + 2(\delta^i{}_k \mathfrak{G}^j{}_{lmn} + \delta^i{}_m \mathfrak{G}^j{}_{nkl}) \} \mathfrak{p}^{mn} \\ &\quad + \delta p^{ij} \delta p^{kl} 4\mathscr{N} \mathfrak{G}_{ijkl} \\ &\quad - \delta h_{ij} \delta \left( \frac{\sqrt{\mathfrak{h}}}{2\mathscr{N}} \bar{\delta} u^{lk}{}_{|l|k} \right). \end{aligned} \quad (47)$$

Second variation of  $\Gamma^i{}_{jk}$

$$\delta^2 \Gamma^i{}_{jk} = -h^{im} \delta \Gamma^l{}_{jk} \delta h_{lm}. \quad (48)$$

Second variation of  $\mathfrak{H}^i$

$$\delta^2(\mathfrak{H}^i) = -h^{im} p^{jk} \delta \Gamma^l{}_{jk} \delta h_{lm} + 2 \delta \Gamma^i{}_{jk} \delta p^{jk}. \quad (49)$$

## Other second variations

Second variation of  $h^{ij}$

$$\delta^2 h^{ij} = (h^{im} h^{jl} h^{kn} + h^{ik} h^{jm} h^{ln}) \delta h_{kl} \delta h_{mn} \quad (50)$$

Second variation of  $\mathfrak{h} = \det h_{ij}$

$$\delta^2 \mathfrak{h} = -\frac{1}{4} \mathfrak{h} (h^{ik} h^{jl} + h^{il} h^{kj} - h^{ij} h^{kl}) \delta h_{ij} \delta h_{kl}. \quad (51)$$

First variation of  $(\delta h_{ij})|_k$

$$\delta \left\{ (\delta h_{ij})|_k \right\} = -2\delta \Gamma^l{}_{k(i} \delta h_{j)l}. \quad (52)$$

In a general background, the second variations of the quantities are much more tedious.

In Robertson–Walker background,

## References

- [1] Leonardo Castellani. “Symmetries in constrained Hamiltonian systems”. In: *Annals of Physics* 143.2 (Oct. 1982), pp. 357–371. DOI: 10.1016/0003-4916(82)90031-8.
- [2] Paul A. M. Dirac. “The Theory of Gravitation in Hamiltonian Form”. In: *Proceedings of the Royal Society A: Mathematical, Physical and Engineering Sciences* 246.1246 (Aug. 1958), pp. 333–343. DOI: 10.1098/rspa.1958.0142.
- [3] Claus Kiefer. *Quantum Gravity*. 3rd ed. Oxford University Press, Apr. 2012. DOI: 10.1093/acprof:oso/9780199585205.001.0001.
- [4] Natalia Kiriushcheva and Sergei Kuzmin. “The Hamiltonian formulation of General Relativity: myths and reality”. In: *Central Eur.J.Phys.* 9:576-615,2011 9.3 (Aug. 31, 2008). DOI: 10.2478/s11534-010-0072-2. arXiv: 0809.0097 [gr-qc].
- [5] Eric Poisson. *A Relativist’s Toolkit*. Cambridge University Press, 2004. DOI: 10.1017/cbo9780511606601.
- [6] Steven Weinberg. *Cosmology*. Oxford University Press, 2008. ISBN: 9780198526827. URL: <https://global.oup.com/academic/product/cosmology-9780198526827>.