

# Integrable Cosmological Models with Liouville Fields

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# Outline

1. Introduction
2. Classical model and the implicitised trajectories
3. Dirac quantisation and the wave functions
4. Conclusions



# Introduction

## Introduction



- Flat Robertson–Walker metric  $ds^2 = -N^2(t) dt^2 + \varkappa^{-1} e^{2\alpha(t)} d\Omega_3^2$ , where  $\varkappa := 8\pi G$ ,  $d\Omega_3^2$  dimensionless spacial metric.
- Homogeneous real Klein–Gordon field with potential  $V e^{\lambda\phi}$  (Liouville), where  $\lambda, V \in \mathbb{R}$ , and kinetic term with sign  $\ell = \pm 1$  (quintessence / phantom model).
- Total action  $\mathcal{S} := S_{\text{EH}} + S_{\text{GHY}} + S_{\text{L}} = \int d\Omega_3^2 \int dt L$ , in which the effective Lagrangian reads

$$L := \varkappa^{3/2} N e^{3\alpha} \left( -\frac{3}{\varkappa} \frac{\dot{\alpha}^2}{N^2} + \ell \frac{\dot{\phi}^2}{2N^2} - V e^{\lambda\phi} \right), \quad (1)$$

in which dot means  $d/dt$  and  $\ell = \pm 1$ .



# Decoupling the variables

Via *rescaled* special orthogonal transformation

- Setting  $\overline{N} := N e^{-3\alpha}$ , eq. (1) transforms to

$$L = \kappa^{3/2} \overline{N} \left( -\frac{3}{\kappa} \frac{\dot{\alpha}^2}{\overline{N}^2} + \ell \frac{\dot{\phi}^2}{2\overline{N}^2} - V e^{\lambda\phi+6\alpha} \right) \quad (2)$$

- Defining  $\Delta := \lambda^2 - 6\ell\kappa$ ,  $\jmath := \text{sgn } \Delta$  and  $g := \jmath\sqrt{|\Delta|} \equiv \jmath\sqrt{\jmath\Delta}$ , the *rescaled* special orthogonal transformation

$$\begin{pmatrix} \alpha \\ \phi \end{pmatrix} = \frac{\jmath}{g} \begin{pmatrix} \lambda & -\ell\kappa \\ -6 & \lambda \end{pmatrix} \begin{pmatrix} \jmath_\beta \beta \\ \jmath_\chi \chi \end{pmatrix} \quad \text{where } \jmath_\beta, \jmath_\chi = \pm 1 \quad (3)$$

gives the decoupled Lagrangian

$$L = \kappa^{3/2} \overline{N} \left( -\jmath \frac{3}{\kappa} \frac{\dot{\beta}^2}{\overline{N}^2} + \ell \jmath \frac{\dot{\chi}^2}{2\overline{N}^2} - V e^{\jmath_\chi g \chi} \right). \quad (4)$$

- The Euler–Lagrange equations w.r.t.  $\overline{N}$ ,  $\beta$  and  $\chi$  will be called the trsfed. 1st, 2nd Friedmann eqs. and the Klein–Gordon eq., respectively.



- Beyond isotropy
  - Bianchi Type-I
- Beyond one Liouville field
  - Two exponential potentials:  $V_1 = V_2$  and special  $\lambda_i$
  - Multiple Liouville fields: mixing kinetic terms<sup>1</sup>
- Beyond classic matter
  - Non-Hermitian,  $PT$ -symmetric Liouville fields<sup>2</sup>

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<sup>1</sup> Alexander A. Andrianov, Oleg O. Novikov, and Chen Lan. In: *Theoretical and Mathematical Physics* 184.3 (2015), pp. 1224–1233.

<sup>2</sup> Alexander A. Andrianov et al. In: *International Journal of Modern Physics D* 19.01 (2010), pp. 97–111, Alexander A. Andrianov, Chen Lan, and Oleg O. Novikov. In: *Springer Proceedings in Physics*. Springer International Publishing, 2016, pp. 29–44.



# Implicitised integration

$$p_\beta \neq 0$$

- Since  $\beta$  is cyclic in eq. (4), the trsfed. 2nd Friedmann eq. can be integrated<sup>3</sup>

$$\text{const.} \equiv p_\beta := \frac{\partial L}{\partial \dot{\beta}} = -6\mathfrak{s}\kappa^{1/2} \frac{\dot{\beta}}{\bar{N}} \equiv -6\mathfrak{s}\mathfrak{s}_\beta \frac{\kappa^{1/2}}{g} \frac{\lambda \dot{\alpha} + \ell \kappa \dot{\phi}}{\bar{N}}. \quad (5)$$

- For  $p_\beta \neq 0$ , fixing the *implicitising gauge*  $\bar{N} = -6\mathfrak{s}\sqrt{\kappa}\dot{\beta}/p_\beta$ , the trsfed. 1st Friedmann equation can be integrated

$$\mathfrak{e}^{\mathfrak{s}_\times g\chi} = \frac{p_\beta^2}{12\kappa^2|V|} S^2 \left( \mathfrak{s}_\beta \sqrt{\frac{3}{2\kappa}} g\beta + C \right), \quad (6)$$

in which  $\mathfrak{v} := \text{sgn } V$ , (sgn,sgn) means  $(\ell, \mathfrak{s}\mathfrak{v})$ , and

$$\begin{aligned} (+,+)S(\gamma) &:= \text{sech}(\gamma), & (+,-)S(\gamma) &:= \text{csch}(\gamma), \\ (-,+)S(\gamma) &:= \sec(\gamma), & (-,-)S(\gamma) &:= \mathfrak{i} \csc(\gamma). \end{aligned} \quad (7)$$

<sup>3</sup> Chen Lan. PhD thesis. Saint Petersburg State University, 2016, Alexander A. Andrianov, Oleg O. Novikov, and Chen Lan. In: *Theoretical and Mathematical Physics* 184.3 (2015), pp. 1224–1233.



# Integration

## Discussions

- The integrals are consistent with the trsfed. Klein–Gordon equation.
- The integral for  $(+, +)$ 
  - has two asymptotes
- The implicitised integral for  $(+, -)$ 
  - contains two distinct solutions
  - has three asymptotes
- The implicitised integral for  $(-, +)$ 
  - is  $\beta$ -even for  $C = 0$
  - contains infinite distinct solutions
  - has infinite asymptotes, which are pairwise parallel
- The integral for  $(-, -)$ 
  - is not real
- The implicitised integral enables one to compare trajectories with wave functions, see below.





# Implicitised integration

$$p_\beta = 0$$

- For  $p_\beta = 0$ , one has  $\beta \equiv \beta_0$  or  $\phi - \phi_0 = -\ell\lambda\alpha/\kappa$ , which is the familiar power-law special solution<sup>4</sup>.
- Further integrating the first Friedmann equation demands  $(+, -)$  or  $(-, +)$  to guarantee  $\bar{N} > 0$ , and the result is automatically consistent with the trsfed. Klein–Gordon equation.
- Fixing  $\bar{N} = (2\kappa^2|V|)^{-1/2}$  yields

$$e^{g\beta\chi} = \left( \frac{2\kappa}{g(t-t_0)} \right)^2. \quad (8)$$

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<sup>4</sup> Mariusz P. Dąbrowski, Claus Kiefer, and Barbara Sandhöfer. In: *Physical Review D* 74.4 (2006).



- The primary Hamiltonian and the Hamiltonian constraint reads

$$H^p = \bar{N}H_\perp + p_{\bar{N}}v^{\bar{N}}, \quad (9)$$

$$H_\perp = -\imath \frac{p_\beta^2}{12\kappa^{1/2}} + \ell \imath \frac{p_\chi^2}{2\kappa^{3/2}} + \kappa^{3/2} V e^{g_\imath \chi}. \quad (10)$$

- Applying the Dirac quantisation rules with Laplace–Beltrami operator, one gets the mss. Wheeler–DeWitt eq. with  $(\beta, \chi)$

$$0 = \widehat{H}_\perp \Psi(\beta, \chi) := \left( \imath \frac{\hbar^2}{12\kappa^{1/2}} \partial_\beta^2 - \ell \imath \frac{\hbar^2}{2\kappa^{3/2}} \partial_\chi^2 + \kappa^{3/2} V e^{g_\imath \chi} \right) \Psi. \quad (11)$$

- Equation (11) is KG-like, hyperbolic for  $\ell = +1$  and *elliptic* for  $\ell = -1$ .

# Separation of the variables and mode functions

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- Writing  $\Psi(\beta, \chi) = \varphi(\beta)\psi(\chi)$ , eq. (11) can be separated into

$$\partial_\beta^2 \varphi(\beta) = k_\beta^2 \varphi(\beta); \quad (12)$$

$$\mathbb{D}\psi(\chi) = k_\beta^2 \psi(\chi), \quad \mathbb{D} := -\ell \frac{6}{\chi} \partial_\chi^2 + \imath v \frac{12\chi^2 |V|}{\hbar^2}. \quad (13)$$

- Equation (13) turns out to be Besselian, and the mode functions are

$$\Psi_\nu(\beta, \chi) := \sum_{i=1}^2 c_i \varphi_\nu^{(i)}(\gamma) \sum_{j=1}^2 a_j B_\nu^{(j)}(\sigma), \quad \nu \geq 0; \nu := \sqrt{\frac{2\chi}{3}} \frac{1}{g} k_\beta, \quad \gamma := \sqrt{\frac{3}{2\chi}} \frac{g}{1} \beta, \quad (14)$$

$$_{(+,+)}B_\nu^{(i)}(\sigma) := K \text{ and } I_{i\nu}(\sigma), \quad (+,-)B$$

$$_{(-,+)}B_\nu^{(i)}(\sigma) := J \text{ and } Y_\nu(\sigma), \quad (-,-)B$$

- Adapted to imaginary order,  $F_\nu(\sigma)$  and  $G_\nu(\sigma)$  are defined in<sup>5</sup>.

<sup>5</sup> T. Mark Dunster. In: *SIAM Journal on Mathematical Analysis* 21.4 (1990), pp. 995–1018.



- Physical mode functions are expected to be regular on the boundary.
- $(+, +)$ :  $|I_{i\nu}(\sigma)| \rightarrow +\infty$  as  $\alpha \rightarrow +\infty$
- $(-, +)$ :
  - $\forall n \in \mathbb{N}$ ,  $|Y_n(\sigma)| \rightarrow +\infty$  as  $\alpha \rightarrow -\infty$ .
  - $\forall \nu \in \mathbb{R}^+ \setminus \mathbb{N}$ , choose  $J_{-\nu}$  instead of  $Y_\nu$ , since  $J_{\pm\nu}$  are also linearly independent.
  - $\forall \nu \in \mathbb{R}^+ \setminus \mathbb{N}$ ,  $|J_{-\nu}(\sigma)| \rightarrow +\infty$  as  $\alpha \rightarrow -\infty$ .
- $(-, -)$ :  $|K_\nu(\sigma)| \rightarrow +\infty$  as  $\alpha \rightarrow -\infty$ ;  $|I_\nu(\sigma)| \rightarrow +\infty$  as  $\alpha \rightarrow +\infty$
- These are not to be included in the space of physical wave functions.  
 $\forall \nu \geq 0$ ,
  - $(+, +)$ :  $K_{i\nu}(\sigma)$  survives
  - $(+, -)$ :  $F$  and  $G_{i\nu}(\sigma)$  survives
  - $(-, +)$ :  $J_\nu(\sigma)$  survives
  - $(-, -)$ : drops out



# Matching quantum number with classical first integral

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- Baustelle
- In order to match the quantum number  $k_\beta$  (or linearly,  $\nu$ ) with the classical first integral  $p_\beta$ , one may apply the *principle of constructive interference*.
- Writing the mode function in the WKB form

$$\Psi_{k_\beta}(\beta, \chi) \sim \sqrt{\rho} e^{iS/\hbar}, \quad S/\hbar \gg 1 \text{ and } k_\beta \gg 1, \quad (15)$$

the principle demands that  $\partial S / \partial k_\beta = 0$  be equivalent to the classical trajectory.



# Matching quantum number with classical first integral

(+, +) as exemplar

- Fixing  $\nu/\sigma > 1$ , the asymptotic expansion reads

$$K_{i\nu}(\sigma) \sim \frac{\sqrt{2\pi} \cos\left(\sqrt{\nu^2 - \sigma^2} - \nu \arccos \frac{\nu}{\sigma} - \frac{\pi}{4}\right)}{(\nu^2 - \sigma^2)^{1/4} e^{\pi\nu/2}} + O\left(\frac{1}{\sigma}\right). \quad (16)$$

- There are two phases with opposite signs. Assuming  $c, a$  are real and applying *the principle* to  $\Psi_\nu(\sigma)$ , one has  $\sigma/\nu = \operatorname{sech}(\beta\gamma)$ , which matches the trajectory with  $C = 0$  if

$$\hbar k_\beta \equiv \hbar \sqrt{\frac{3}{2\kappa}} g \hbar \nu = p_\beta, \quad (17)$$

- Non-vanishing  $C$  can be compensated by the phase of  $c$ 's and  $a$ 's.
- Fixing  $\nu/\sigma < 1$ , the asymptotic expansion is not oscillatory, but exponential; it is not within the WKB regime.
- The conclusions also hold for  $F_{i\nu}(\sigma)$ ,  $G_{i\nu}(\sigma)$  for  $(+, -)$ , and  $J_\nu(\sigma)$  for  $(-, +)$ .



# Inner product for wave functions

## Schrödinger product

- To make sense of amplitude and wave packet, an inner product is necessary
- A common starting point of most approaches is the *Schrödinger product*

$$(\Psi_1, \Psi_2)_S := \int d\chi \Psi_1^*(\beta, \chi) \Psi_2(\beta, \chi); \quad (18)$$

- In terms of a norm,  $(\Psi, \Psi)_S \equiv \int d\chi \rho_S(\beta, \chi)$ , in which  $\rho_S := \Psi^* \Psi$ . Manifestly  $\rho_S \geq 0$ ; one has  $(\Psi, \Psi)_S > 0$ .
- The corresponding Schrödinger current does not satisfy continuity equation  $\dot{\rho}_S + \nabla \cdot \vec{j}_S = 0$ , because eq. (11) is KG-like.
- $K_{i\nu}$ <sup>6</sup> for  $(+, +)$ ,  $F_{i\nu}$  and  $G_{i\nu}$  for  $(+, -)$  can be proved to be orthogonal and complete individually, as well as can be normalised.
  - $J_{i\nu}$ 's for  $(+, -)$  are not orthogonal

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<sup>6</sup> Semyon B. Yakubovich. In: *Opuscula Math.* 26.1 (2006), pp. 161–172, A. Passian et al. In: *Journal of Mathematical Analysis and Applications* 360.2 (2009), pp. 380–390, Radosław Szmytkowski and Sebastian Bielski. In: *Journal of Mathematical Analysis and Applications* 365.1 (2010), pp. 195–197.



# Peculiarity of the phantom model $(-, +)$

Orthogonality for mode functions; Hermiticity for operators

- $J_\nu(\sigma)$ 's are not orthogonal under the Schrödinger product

$$(J_\nu, J_{\tilde{\nu}})_S \propto \int_{-\infty}^{+\infty} dx J_\nu^*(e^x) J_{\tilde{\nu}}(e^x) = \frac{2 \sin(\pi(\nu - \tilde{\nu})/2)}{\pi(\nu^2 - \tilde{\nu}^2)}, \quad (19)$$

therefore  $\mathbb{D}$  in eq. (13) is not Hermitian (though we do not need it so far)

- $\hat{p}_\chi^2$  is not Hermitian for  $\{J_\nu(\sigma)\}$  under the Schrödinger product

$$\int_{-\infty}^{+\infty} dx J_\nu^*(-\partial_x^2 J_{\tilde{\nu}}) - \int_{-\infty}^{+\infty} dx (-\partial_x^2 J_\nu)^* J_{\tilde{\nu}} = \frac{2}{\pi} \sin \frac{\pi(\nu - \tilde{\nu})}{2}. \quad (20)$$

- In order to save Hermiticity for  $p_\chi^2$  and  $\mathbb{D}$  and orthogonality of the modes under Schrödinger product, one can restrict

$$\nu = 2n + \nu_0, \quad n \in \mathbb{N}, \quad \nu_0 \in [0, 2). \quad (21)$$

- The classical trajectory is  $\beta$ -even; imposing the same condition fixes  $\nu_0 = 1$ .





## Levels of the phantom model are discretised

...if one imposes Hermiticity of squared momenta under the Schrödinger product

- This kind of subtlety on Hermiticity is named as *self-adjoint extension*, which arises already for infinite square well<sup>7</sup>.
- It also applies to  $x^{-2}$  potentials<sup>8</sup>, which is of cosmological relevance<sup>9</sup>.

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<sup>7</sup> Guy Bonneau, Jacques Faraut, and Galliano Valent. In: *American Journal of Physics* 69.3 (2001), pp. 322–331.

<sup>8</sup> Andrew M. Essin and David J. Griffiths. In: *American Journal of Physics* 74.2 (2006), pp. 109–117, Vanilse S. Araujo, F. A. B. Coutinho, and J. Fernando Perez. In: *American Journal of Physics* 72.2 (2004), pp. 203–213.

<sup>9</sup> Mariam Bouhmadi-López et al. In: *Physical Review D* 79.12 (2009).



# Further inner products for wave functions

## Klein–Gordon and Mostafazadeh product

- Since eq. (11) is KG-like, another popular choice is the KG product

$$(\Psi_1, \Psi_2)_{\text{KG}}^g := \text{i}g \left\{ (\Psi_1, \dot{\Psi}_2)_S - (\dot{\Psi}_1, \Psi_2)_S \right\}, \quad g > 0. \quad (22)$$

- $\rho_{\text{KG}}$  is real but may go negative
- Unique  $\vec{J}_{\text{KG}}$  exists such that  $\dot{\rho}_{\text{KG}} + \nabla \cdot \vec{J}_{\text{KG}} = 0$ .
- Mostafazadeh<sup>10</sup> found an inner product *for Hermitian  $\mathbb{D}$  with positive spectrum*:  $\forall \kappa > 0$ ,

$$(\Psi_1, \Psi_2)_{\text{M}}^{\kappa} := \kappa \left\{ (\Psi_1, \mathbb{D}^{+1/2} \Psi_2)_S + (\dot{\Psi}_1, \mathbb{D}^{-1/2} \dot{\Psi}_2)_S \right\}, \quad (23)$$

and  $-1 < a < 1$ .

- $\rho_{\text{M}}^{\kappa}$  may go complex
- Unique  $\vec{J}_{\text{M}}^{\kappa}$  exists such that  $\dot{\rho}_{\text{M}}^{\kappa} + \nabla \cdot \vec{J}_{\text{M}}^{\kappa} = 0$ .

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<sup>10</sup> Ali Mostafazadeh. In: *Classical and Quantum Gravity* 20.1 (2002), pp. 155–171.



- It can be shown<sup>11</sup> that

$$\varrho_M^\kappa := \kappa \left\{ |\mathbb{D}^{+1/4}\Psi|^2 + |\mathbb{D}^{-1/4}\dot{\Psi}|^2 \right\} \quad (24)$$

satisfies

- $\varrho_M^\kappa$  is non-negative
- $\int d\chi \varrho_M^\kappa = \int d\chi \rho_M^\kappa \equiv (\Psi_1, \Psi_2)_M^\kappa$ : it may be understood as a prob. density; but
- The corresponding conserved  $\tilde{\mathcal{J}}_M^\kappa$  does not exist.

<sup>11</sup> Ali Mostafazadeh and F. Zamani. In: *Annals of Physics* 321.9 (2006), pp. 2183–2209.



- A normalising  $\kappa$  for  $(\cdot, \cdot)_M^\kappa$  has not yet been able to be evaluated, hence a quantitative comparison of  $(\cdot, \cdot)_S$  and  $(\cdot, \cdot)_M^\kappa$  is not yet possible.



- $PT$ -symmetric instead of phantom field
- Cosmological perturbation



# Integration of the transformed first Friedmann equation

$$p_\beta \neq 0$$

In order to integrate the equation under the implicitising gauge

$$\mathcal{J} \frac{p_\beta^2}{12} \left( -\ell \frac{\kappa^{1/2}}{6} \left( \frac{d\chi}{d\beta} \right)^2 + \kappa^{-1/2} \right) - \kappa^{3/2} V e^{g_\beta \chi} = 0, \quad (25)$$

one can substitute

$$\gamma := \sqrt{\frac{3}{2\kappa}} g_\beta, \quad \tilde{\sigma}^2 := \frac{p_\beta^2}{12\kappa^2 |V|} e^{-g_\beta \chi}, \quad (26)$$

to get

$$\left( \frac{d\tilde{\sigma}}{d\gamma} \right)^2 + \ell (\mathcal{J}v - \tilde{\sigma}^2) = 0 \quad \Rightarrow \quad \frac{d\gamma}{d\tilde{\sigma}} = \pm \frac{1}{\sqrt{\ell(-\mathcal{J}v + \tilde{\sigma}^2)}}, \quad (27)$$

which is of the standard inverse hyperbolic / trigonometric form for  $(+, +)$ ,  $(+, -)$  and  $(-, +)$ .



# Integration of the separated mss. Wheeler–DeWitt equation

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In order to integrate the separated minisuperspace Wheeler–DeWitt equation

$$\mathbb{D}\psi(\chi) = k_\beta^2 \psi(\chi), \quad \mathbb{D} := -\ell^2 \frac{6}{\varkappa} \partial_\chi^2 + \imath \nu \frac{12\varkappa^2 |V|}{\hbar^2}, \quad (13 \text{ rev.})$$

one can transform

$$\nu := \sqrt{\frac{2\varkappa}{3}} \frac{k_\beta}{g}, \quad \sigma^2 := \frac{8\varkappa^3 |V| e^{g\imath_\chi \chi}}{\hbar^2 g^2}, \quad (14 \text{ rev.})$$

to get

$$\sigma^2 \psi''(\sigma) + \sigma \psi'(\sigma) + \ell(\nu^2 - \imath \nu \sigma^2) \psi(\sigma) = 0, \quad (28)$$

which is of the standard Besselian form.







- Mit diesem *beamer theme* ist es möglich, Präsentationen in  $\text{\LaTeX}$  mit der Beamer-Klasse zu erstellen, die dem Corporate Design der Universität zu Köln entsprechen
- Auf die Beamer-Klasse wird in diesem Dokument nicht näher eingegangen, nähere Informationen finden Sie unter <http://latex-beamer.sourceforge.net/>



Das Theme kann mit den folgenden Optionen geladen werden

```
\usepackage[%  
% uk,          %% Farben aller Fakultäten  
wiso,          %% Wiso-Fakultät  
% jura,        %% Rechtswissenschaftliche Fakultät  
% medizin,     %% Medizinische Fakultät  
% philo,       %% Philosophische Fakultät  
% matnat,      %% Mathematisch-Naturwissenschaftliche Fakultät  
% human,       %% Humanwissenschaftliche Fakultät  
% verw,        %% Universitätsverwaltung  
{UzK}
```



- Es stehen verschiedene Fußzeilen zur Auswahl, die als Option beim Laden des *themes* übergeben werden:
  - Balken mit allen Fakultätsfarben (Option `uk`)
  - Balken in jeweils einer Fakultätsfarbe (Optionen `wiso`, `jura`, `medizin`, `philo`, `matnat`, `human`, `verw`)<sup>12</sup>
- "'Universität zu Köln"' sowie der Name der Fakultät sind im Theme definiert, das Institut oder Seminar kann mit dem Befehl `\institute{}` festgelegt werden
- Die Optionen sind im Quellcode dieser Präsentation dokumentiert

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<sup>12</sup>Es werden die offiziellen RGB-Werte aus dem 2-D Handbuch Corporate Design verwendet.



- Der Universitäts- sowie die Fakultätsnamen werden standardmäßig auf Deutsch angezeigt.
- Übergeben Sie dem Paket babel die Option english, so werden diese Namen entsprechen angepasst.
- Die Übersetzungen können in der Theme-Datei beamerthemeUzK.sty geändert werden



# block-Umgebungen

## Standard (block)

Verwendet die Farbe "Blaugrau Mittel" als Blocktitel-Hintergrund

## exampleblock

Bei Verwendung der Fußzeile mit allen Fakultätsfarben Titelhintergrund in Wiso-Grün, sonst in der jeweiligen Fakultätsfarbe

## alertblock

Verwendet das Rot der Folientitel

# Installation

- Das Theme besteht aus den Dateien `beamerthemeUzK.sty` und `beamercolorthemeUzK.sty` sowie den Grafikdateien `logo.pdf` und `logo-small.pdf`.
- Das Theme kann auf zwei Arten verwendet werden:
  1. Die vier Dateien werden in den selben Ordner wie die zu erstellende Präsentation gelegt
  2. Die vier Dateien werden im lokalen *texmf*-Baum abgelegt
- Die zweite Variante ist der ersten vorzuziehen, da das Theme so an einem zentralen Ort vorliegt



## Was noch zu tun ist...

- Erstellen einer eigenen Titelseite
- ...