Notes on Canonical Singular Dynamics

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1 Classical formalism

Lagrangian with velocity

$$L^{\mathrm{v}} := \left. L \right|_{\dot{q}=v} \tag{1.1}$$

Equations of motion

$$\sum_j M_{ij} \dot{v}_j = K_i^{\mathrm{v}}, \quad \dot{q}_i = v_i. \tag{1.2}$$

where

$$M_{ij}(q,v) := \frac{\partial^2 L^{\mathrm{v}}}{\partial v_i \, \partial v_j}. \tag{1.3}$$

Adding

$$p_i := \frac{\partial L^{\mathrm{v}}}{\partial v_i}. \tag{1.4}$$

Variation of

$$S[q,p;v] := \int \mathrm{d}t \left[L^{\mathrm{v}} + \sum_i p_i(\dot{q}_i - v_i) \right]. \tag{1.5}$$

gives the *Euler–Lagrange equations with velocities*

$$\dot{q}_i = v_i, \quad \dot{p}_i = \frac{\partial L^{\mathrm{v}}}{\partial q_i}, \quad p_i = \frac{\partial L^{\mathrm{v}}}{\partial v_i}. \tag{1.6}$$

Hamiltonian with velocity

$$H^{\mathrm{v}}(q,p;v) := \sum_i p_i v_i - L^{\mathrm{v}}. \tag{1.7}$$

Identities
$$\frac{\partial H^{\mathrm{v}}}{\partial q_i} \equiv -\frac{\partial L^{\mathrm{v}}}{\partial q_i}, \quad \frac{\partial H^{\mathrm{v}}}{\partial p_i} \equiv v_i, \quad \frac{\partial H^{\mathrm{v}}}{\partial v_i} \equiv p_i - \frac{\partial L^{\mathrm{v}}}{\partial v_i}. \tag{1.8}$$

Variation of

$$S[q,p;v] := \int \mathrm{d}t \left[\sum_i p_i \dot{q}_i - H^{\mathrm{v}} \right] \tag{1.9}$$

gives the *canonical equations with velocities*

$$\dot{q}_i = [q_i, H^{\mathrm{v}}]_{\mathrm{P}}, \quad \dot{p}_i = [p_i, H^{\mathrm{v}}]_{\mathrm{P}}, \quad \frac{\partial H^{\mathrm{v}}}{\partial v_i} = 0, \tag{1.10}$$

where the *Poisson bracket* is defined as

$$[f^{\mathrm{v}}, g^{\mathrm{v}}]_{\mathrm{P}} := \sum_i \left(\frac{\partial f^{\mathrm{v}}}{\partial q_i} \frac{\partial g^{\mathrm{v}}}{\partial p_i} - \frac{\partial f^{\mathrm{v}}}{\partial p_i} \frac{\partial g^{\mathrm{v}}}{\partial q_i} \right). \tag{1.11}$$

$v_a = \bar{v}_a(q, p; \{v_\alpha\})$ can be solved, $a = 1, 2, \dots, r_M$; v_α cannot be solved, $\alpha = r_M + 1, \dots, n$, where $r_M = \mathrm{rank}\, M$.

(need to show $v_a = \bar{v}_a(q, p_a)$)

Primary constraints in the standard form

$$\Phi_\alpha(q, p) := \left. \frac{\partial H^{\mathrm{v}}}{\partial v_\alpha} \right|_{\{v_\alpha = \bar{v}_\alpha\}} \equiv p_\alpha - \bar{p}_\alpha(q, \{p_a\}), \tag{1.12}$$

where

$$\bar{p}_\alpha(q, \{p_a\}) := \left. \frac{\partial L^{\mathrm{v}}}{\partial v_\alpha} \right|_{\{v_a = \bar{v}_a\}}. \tag{1.13}$$

Hamiltonian with primary constraint

$$H^{\mathrm{p}} := \left. H^{\mathrm{v}} \right|_{\{v_a = \bar{v}_a\}} \equiv H^{\mathrm{v}}(q, p; \{\bar{v}^a(q, p_a; \{v_\alpha\}), v_\alpha\}). \tag{1.14}$$

Subspace of primary constraints

$$\Gamma_{\mathrm{P}} = \{(q, p) \,|\, \Phi_\alpha(q, p) = 0, \forall \alpha\} \tag{1.15}$$

Since

$$\frac{\partial H^{\mathrm{p}}}{\partial v_\alpha} = \left. \frac{\partial H^{\mathrm{v}}}{\partial v_\alpha} \right|_{\{v_a = \bar{v}_a\}} = \Phi_\alpha \equiv p_\alpha - \bar{p}_\alpha(q, \{p_a\}), \tag{1.16}$$

H^{p} is linear in v_α. One writes

$$H^{\mathrm{p}}(q, \{p_a\}; \{p_\alpha\}, \{v_\alpha\}) = H(q, \{p_a\}) + \sum_\alpha v_\alpha \Phi_\alpha, \tag{1.17}$$

where H^{c} is the *canonical Hamiltonian* or simply *Hamiltonian*.

Proposition H^{c} is independent of $\{p_\alpha\}$.

Proposition   Canonical equations with primary constraints

$$\dot{q}_i = [q_i, H]_{\mathrm{P}} + \sum_\beta v_\beta [q_i, \phi_\beta]_{\mathrm{P}}, \tag{1.18}$$

$$\dot{p}_i = [p_i, H]_{\mathrm{P}} + \sum_\beta v_\beta [p_i, \phi_\beta]_{\mathrm{P}}, \tag{1.19}$$

$$\Phi_\alpha(q, p) = 0, \tag{1.20}$$

where v_β's are undetermined. Note that eq. (1.18) for $i = \alpha$ holds identically: $\dot{q}_\alpha = \dot{q}_\alpha$.

Weak equality: $f_1 \approx f_2$ iff $\left. f_1 \right|_{\Gamma_{\mathrm{P}}} = \left. f_2 \right|_{\Gamma_{\mathrm{P}}}$.

Proposition   if f and g are two functions over the phase space Γ, and $f \approx h$, then

$$\frac{\partial}{\partial q_i}\left(f - \sum_\beta \phi_\beta \frac{\partial f}{\partial p_\beta}\right) \approx \frac{\partial}{\partial q_i}\left(h - \sum_\beta \phi_\beta \frac{\partial h}{\partial p_\beta}\right), \qquad (1.21)$$

$$\frac{\partial}{\partial p_i}\left(f - \sum_\beta \phi_\beta \frac{\partial f}{\partial p_\beta}\right) \approx \frac{\partial}{\partial p_i}\left(h - \sum_\beta \phi_\beta \frac{\partial h}{\partial p_\beta}\right). \qquad (1.22)$$

Corollary $\forall H_1 \approx H$,

$$\dot{q}_i \approx [q_i, H]_{\mathrm{p}}, \qquad \dot{p}_i \approx [p_i, H]_{\mathrm{p}}. \qquad (1.23)$$

Primary and second constraints $\phi_\mu^{(1,)}$, $\phi_\omega^{(2,)}$; first and second class constraints $\phi_u^{(,1)}$, $\phi_w^{(,2)}$.

2   Examples

2.1   Toy examples

Example 0

Gitman and Tyutin 1990, sec. 1.2

$$L = \frac{1}{2}(\dot{x} - y)^2 \qquad (2.1)$$

Example 1

$$L = \frac{1}{2}\dot{x}^2 + \dot{x}y - \frac{1}{2}(x - y)^2. \qquad (2.2)$$

One has

$$L^{\mathrm{v}} = \frac{1}{2}v_x^2 + v_x y - \frac{1}{2}(x - y)^2, \qquad (2.3)$$

so that

$$p_x = \frac{\partial L^{\mathrm{v}}}{\partial v_x} = v_x + y, \qquad p_y = 0, \qquad (2.4)$$

thus

$$\bar{v}_x = p_x - y. \qquad (2.5)$$

So that v_y is the primary inexpressible velocity.

The Hamiltonian with velocity reads

$$H^{\mathrm{v}}(q, p; v) = v_x p_x + v_y p_y - \frac{1}{2}v_x^2 - v_x y + \frac{1}{2}(x - y)^2, \qquad (2.6)$$

whilst the Hamiltonians are

$$H^{\mathrm{p}}\big(q, p; \overline{v}_x, v_y\big) = H^{\mathrm{c}} + v_y \Phi_1, \tag{2.7}$$

$$H^{\mathrm{c}}(q, p) = \frac{1}{2}(p_x - y)^2 + \frac{1}{2}(x - y)^2, \tag{2.8}$$

where the only primary constraint $\Phi_1 = p_y$.

Persistence condition of Φ_1 leads to

$$0 \approx [\Phi_1, H^{\mathrm{p}}]_{\mathrm{P}} = p_x + x - 2y =: \Phi_2. \tag{2.9}$$

Note that $[\Phi_1, \Phi_2] = p_x - x$ does not vanish on Γ, thus $\Phi_{1,2}$ are second-class constraints, and no more constraint can be generated. To solve for v_y one evaluates

$$0 =: [\Phi_1, H^{\mathrm{p}}]_{\mathrm{P}} = p_x - x - 2v_y, \tag{2.10}$$

so that $v_y := (p_x - x)/2$ solves the constraint.

One also has

$$Q_{\alpha\beta} = [\Phi_\alpha, \Phi_\beta]_{\mathrm{P}} = \begin{pmatrix} 0 & +2 \\ -2 & 0 \end{pmatrix}, \qquad (Q^{-1})_{\alpha\beta} = \begin{pmatrix} 0 & -1/2 \\ +1/2 & 0 \end{pmatrix}, \tag{2.11}$$

so that the Dirac brackets are defined as

$$[f, g]_{\mathrm{D}} := [f, g]_{\mathrm{P}} + \frac{1}{2}\Big([f, p_y]_{\mathrm{P}}[p_x + x - 2y, g]_{\mathrm{P}} - [f, p_x + x - 2y]_{\mathrm{P}}[p_y, g]_{\mathrm{P}}\Big). \tag{2.12}$$

The fundamental ones different from Possion brackets are

$$[x, y]_{\mathrm{D}} = [y, p_x]_{\mathrm{D}} = \frac{1}{2}, \qquad [y, p_y]_{\mathrm{D}} = 0. \tag{2.13}$$

Last one different from book?

Example 2

$$L = \frac{1}{2}\dot{x}^2 + \dot{x}y + \frac{1}{2}(x - y)^2 \tag{2.14}$$

Primary constraint

$$p_y = 0; \tag{2.15}$$

Hamiltonian with primary constraint

$$H^{\mathrm{p}} = \frac{1}{2}p_x^2 - p_x y - \frac{1}{2}x^2 + xy + v_y p_y. \tag{2.16}$$

Example 3

$$L = \frac{1}{2}(\dot{q}_2 - \mathbb{e}^{q_1})^2 + \frac{1}{2}(\dot{q}_3 - q_2)^2. \tag{2.17}$$

2.2 Parametrised systems

Non-relativistic point particle

Kiefer 2012, sec. 3.1.1

$$S[q(t)] := \int_{t_1}^{t_2} \mathrm{d}t\, L\left(q, \frac{\mathrm{d}q}{\mathrm{d}t}\right) \tag{2.18}$$

Relativistic charged point particle

Landau and Lifshitz 1975, sec. 16, Kiefer 2012, sec. 3.1.2

$$S := \int -m\,\mathrm{d}s + eA_\mu(x)\,\mathrm{d}x^\mu =: \int \mathrm{d}\tau\, L, \tag{2.19}$$

where the Lagrangian reads

$$L = -m\sqrt{-\eta_{\mu\nu}\dot{x}^\mu \dot{x}^\nu} + q\dot{x}^\mu A_\mu(x). \tag{2.20}$$

$$M_{\mu\nu} := \frac{\partial^2 L^{\mathrm{v}}}{\partial v^\mu \, \partial v^\nu} = m\frac{-\eta_{\mu\nu}\eta_{\alpha\beta} + \eta_{\mu\alpha}\eta_{\nu\beta}}{\left(-\eta_{\rho\sigma}v^\rho v^\sigma\right)^{3/2}}v^\alpha v^\beta, \tag{2.21}$$

which has one and only one eigenvector with null eigenvalue

$$v^\mu M_{\mu\nu} = 0. \tag{2.22}$$

Momenta

$$p_\mu = \frac{\partial L^{\mathrm{v}}}{\partial v^\mu} = \frac{m\eta_{\mu\nu}v^\nu}{\sqrt{-\eta_{\rho\sigma}v^\rho v^\sigma}} + qA_\mu. \tag{2.23}$$

If one chooses v^0 to be the primary inexpressible velocity, then eliminating p_0 in eq. (2.23) yields

$$v^i = \frac{\xi\eta^{ij}\left(p_j - qA_j\right)v^0}{\sqrt{m^2 + \eta^{kl}(p_k - qA_k)(p_l - qA_l)}}, \tag{2.24}$$

where $\xi = \mathrm{sgn}\, v^0$. In the following $\xi = +1$ will be chosen.

Inserting eq. (2.24) into the Hamiltonian with velocity

$$H^{\mathrm{v}} = v^\mu p_\mu - L^{\mathrm{v}} = m\sqrt{-\eta_{\mu\nu}v^\mu v^\nu} + v^\mu\left(p_\mu - qA_\mu(x)\right), \tag{2.25}$$

one obtains the Hamiltonian with primary constraint

$$H^{\mathrm{p}} = v^0\left(p_0 - qA_0 + \sqrt{m^2 + \eta^{kl}(p_k - qA_k)(p_l - qA_l)}\right), \tag{2.26}$$

where only a primary constraint survives, which is obviously a first-class constraint

$$\phi^{(1,1)} = p_0 - qA_0 + \sqrt{m^2 + \eta^{kl}(p_k - qA_k)(p_l - qA_l)}, \tag{2.27}$$

and the canonical Hamiltonian vanishes

$$H^c = 0. \tag{2.28}$$

To compare, note in the non-covariant formalism (Landau and Lifshitz 1975, sec. 8)

$$S = \int \mathrm{d}t\, L, \qquad L = -m\sqrt{1 - \dot{\vec{x}}^2} - q\phi + q\dot{\vec{x}} \cdot \vec{A}, \tag{2.29}$$

the system is regular, and the canonical Hamiltonian reads

$$H^c = \sqrt{m^2 + \left(\vec{p} - q\vec{A}\right)^2} + q\phi, \tag{2.30}$$

which corresponds to setting $\phi^{(1,1)} = 0$, $p_0 \to -H^c$ ($p_\mu = (-E, \vec{p})$), and noting $A_\mu = \left(-\phi, \vec{A}\right)$.

Relativistic point particle with einbein

Blumenhagen, Lüst, and Theisen 2013, sec. 2.1

$$L := \frac{1}{2}\left(e^{-1}\eta_{\mu\nu}\dot{x}^\mu \dot{x}^\nu - m^2 e\right) \tag{2.31}$$

$$p_\mu = \frac{\partial L^{\mathrm{v}}}{\partial v^\mu} = e^{-1}\eta_{\mu\nu}v^\nu, \qquad p_e = 0. \tag{2.32}$$

Choosing v^e to be the primary inexpressible velocity, one has

$$v^\mu = e\eta^{\mu\nu}p_\nu. \tag{2.33}$$

Hamiltonian with velocity

$$H^{\mathrm{v}} = v^\mu p_\mu + v^e p_e + \frac{1}{2}\left(-e^{-1}\eta_{\mu\nu}v^\mu v^\nu + m^2 e\right); \tag{2.34}$$

Hamiltonian with primary constraint

$$H^{\mathrm{p}} = \frac{e}{2}\left(\eta^{\mu\nu}p_\mu p_\nu + m^2\right) + v^e p_e; \tag{2.35}$$

canonical Hamiltonian

$$H^c = \frac{e}{2}\left(\eta^{\mu\nu}p_\mu p_\nu + m^2\right). \tag{2.36}$$

The only primary constraint

$$\Phi^{(1,)} = p_e; \tag{2.37}$$

its time evolution

$$\left[\Phi^{(1,)}, H^{\mathrm{p}}\right]_{\mathrm{P}} = [p_e, e]_{\mathrm{P}} \frac{1}{2}\left(\eta^{\mu\nu}p_\mu p_\nu + m^2\right)$$
$$= -\frac{1}{2}\left(\eta^{\mu\nu}p_\mu p_\nu + m^2\right). \tag{2.38}$$

Choose

$$\Phi^{(2,)} = \eta^{\mu\nu}p_\mu p_\nu + m^2, \tag{2.39}$$

whose Possion bracket with H^{p} vanishes; furthermore,

$$\left[\Phi^{(1,)}, \Phi^{(2,)}\right]_{\mathrm{P}} \equiv 0. \tag{2.40}$$

Thus one ends up with two first-class constraints.

2.2.1 Neutral scalar field

Kiefer 2012, sec. 3.3

2.3 Maxwell–Proca theory

$$\mathcal{L} = -\frac{1}{4}F_{\mu\nu}F^{\mu\nu} - \frac{1}{2}m^2 A_\mu A^\mu + A_\mu J^\mu, \tag{2.41}$$

where $m > 0$ corresponds to the Proca theory Gitman and Tyutin 1990, sec. 2.3, and $m = 0$ the Maxwell theory H. J. Rothe and K. D. Rothe 2010, sec. 3.3.3, Gitman and Tyutin 1990, sec. 2.4.

Lagrangian density with velocity

$$\mathcal{L}^{\mathrm{v}} = \frac{1}{2}(V_i - \partial_i A_0)^2 - \frac{1}{4}F_{ij}^2 + \frac{m^2}{2}(A_0^2 - A_i^2) + A_0 J^0 + A_i J^i; \tag{2.42}$$

momenta density

$$B^0 := \frac{\partial \mathcal{L}^{\mathrm{v}}}{\partial V_0} = 0, \qquad B^i := \frac{\partial \mathcal{L}^{\mathrm{v}}}{\partial V_i} = V^i - \partial^i A_0; \tag{2.43}$$

Hamiltonians

$$\mathcal{H}^{\mathrm{p}} = \mathcal{H}^{\mathrm{c}} + V_0 \Phi_1, \tag{2.44}$$

$$\mathcal{H}^{\mathrm{c}} = \frac{1}{2}(B^i)^2 + B^i \partial_i A_0 + \frac{1}{4}F_{ij}^2 + \frac{m^2}{2}(-A_0^2 + A_i^2) - A_0 J^0 - A_i J^i, \tag{2.45}$$

$$\tag{2.46}$$

where

$$\Phi_1 = B^0 \tag{2.47}$$

is the only primary constraint.

$$[\Phi_1(\vec{x}_1), \mathcal{H}^{\mathrm{p}}(\vec{x}_2)]_{\mathrm{P}} = \left[(B^0)_1, \left(B^i \partial_i A_0 - \frac{m^2}{2}A_0^2 - A_0 J^0\right)_2\right]_{\mathrm{P}}$$
$$= (-B^i \partial_i + m^2 A_0 - J_0)_2 \,\delta(\vec{x}_1 - \vec{x}_2), \qquad (2.48)$$

where $J_0 = -J^0$. Integration with $\mathbb{d}^d x_2$ yields the secondary constraint

$$[\Phi_1, H^{\mathrm{p}}]_{\mathrm{P}} = \partial_i B^i + m^2 A_0 - J_0 =: \Phi_2, \qquad (2.49)$$

so that

$$[\Phi_1(\vec{x}_1), \Phi_2(\vec{x}_2)]_{\mathrm{P}} = -m^2 \delta(\vec{x}_1 - \vec{x}_2). \qquad (2.50)$$

One may further compute

$$[\Phi_2(\vec{x}_1), \mathcal{H}^{\mathrm{c}}(\vec{x}_2)]_{\mathrm{P}} = \left[(\partial_i B^i)_1, \left(\frac{1}{4}F_{jk}^2 + \frac{m^2}{2}A_j^2 - A_j J^j\right)_2\right]_{\mathrm{P}}, \qquad (2.51)$$

in which

$$\left[(\partial_i B^i)_1, \left(\frac{1}{4}F_{jk}^2\right)_2\right]_{\mathrm{P}} = (\partial_j A_k - \partial_k A_j)_2 (\partial_i)_1 \left[(B^i)_1, (\partial^j A^k)_2\right]_{\mathrm{P}}$$
$$= -(F^{ij}\partial_j)_2 (\partial_i)_1 \delta(\vec{x}_1 - \vec{x}_2). \qquad (2.52)$$

The Poisson bracket can be evaluated to be

$$[\Phi_2(\vec{x}_1), \mathcal{H}^{\mathrm{c}}(\vec{x}_2)]_{\mathrm{P}} = -\left(F^{ij}\partial_j + m^2 A^i + J^i\right)_2 (\partial_i)_1 \delta(\vec{x}_1 - \vec{x}_2). \qquad (2.53)$$

Integration with $\mathbb{d}^3 x_2$ yields

$$[\Phi_2, H^{\mathrm{c}}]_{\mathrm{P}} = -\partial_i(m^2 A^i + J^i). \qquad (2.54)$$

For Proca theory $m > 0$, then the algorithm terminates, and one obtains a pure second-class system.

Introducing the standard coordinates

$$\alpha_i = A_i + m^{-2}(\partial_i B^0 - J_i), \qquad \beta^i = B^i; \qquad (2.55)$$
$$\alpha_0 = A_0 + m^{-2}(\partial_i B^i - J_0), \qquad \beta^0 = B^0. \qquad (2.56)$$

Note that $J^i = J_i$. One has

$$\mathcal{H}^{\mathrm{p}} = \mathcal{H}^{\mathrm{phy}} + \mathcal{H}^{\mathrm{con}} + \mathcal{H}^{\mathrm{irr}}, \qquad (2.57)$$

where

$$\mathcal{H}^{\mathrm{phy}} = \frac{1}{2}(\beta^i)^2 + \frac{m^2}{2}\alpha_i^2 + \frac{1}{4}(\partial_i\alpha_j - \partial_j\alpha_i)^2 + \frac{1}{2m^2}(\partial_i\beta^i)^2$$
$$+ \frac{1}{m^2}J^0\partial_i\beta^i, \tag{2.58}$$

$$\mathcal{H}^{\mathrm{con}} = -\frac{m^2}{2}\alpha_0^2 - \frac{1}{2m^2}(\partial_i\beta^0)^2, \tag{2.59}$$

$$\mathcal{H}^{\mathrm{irr}} = \partial_i\left(\alpha_0\beta^i - \beta^0\alpha_i + \frac{1}{m^2}(\beta^0\partial_i\beta^0 - \beta^i\partial_j\beta^j - J^0\beta^i)\right)$$
$$+ \frac{1}{2m^2}\left((J^0)^2 - (J^i)^2\right). \tag{2.60}$$

Further more,

$$\Phi_1 = beta^0, \qquad \Phi_2 = m^2\alpha_0 \propto \alpha_0. \tag{2.61}$$

For Maxwell theory $m = 0$.

2.4 String theories

Nambu–Gotō action

Generalising the kinetic part of (2.19), one has

$$S_{\mathrm{NG}} := -T\int_\Sigma \mathrm{d}A =: -T\int_\Sigma \mathrm{d}^2\sigma\mathcal{L}, \tag{2.62}$$

where the Lagrangian density

$$\mathcal{L} = \sqrt{-\Gamma}, \quad \Gamma := \det\Gamma_{\alpha\beta}, \quad \Gamma_{\alpha\beta} := \frac{\partial X^\nu}{\partial\sigma^\alpha}\frac{\partial X_\nu}{\partial\sigma^\alpha}. \tag{2.63}$$

Historically Gotō 1971; Nambu 1970; Reference e.g. Blumenhagen, Lüst, and Theisen 2013 Kiefer 2012, sec. 3.2

Polyakov action

Generalising (2.31)

$$S_{\mathrm{P}}\big[X^\mu, h_{\alpha\beta}\big] = -\frac{T}{2}\int_\Sigma \mathcal{L}, \tag{2.64}$$

where

$$\mathcal{L} := \sqrt{-h}h^{\alpha\beta}\Gamma_{\alpha\beta}. \tag{2.65}$$

Historically Brink, Di Vecchia, and Howe 1976; Deser and Zumino 1976; Polyakov 1981; Reference Kiefer 2012, sec. 3.2

2.5 Gravitation theories

Closed Friedmann universe

This part adapts ibid., sec. 8.1.2.

The total action reads

$$S := S_{\text{EG}} + S_\phi, \tag{2.66}$$

where S_{EG} follows (2.83), and

$$S_\phi := \int_{\mathcal{M}} \mathrm{d}^4 x \sqrt{-g} \left(-\frac{1}{2} g^{\mu\nu} (\nabla_\mu \phi)(\nabla_\nu \phi) - m^2 \phi^2 \right). \tag{2.67}$$

Adapting

$$\mathrm{d}s^2 = -N^2(t)\,\mathrm{d}t^2 + a^2(t)\,\mathrm{d}\Omega_3^2, \tag{2.68}$$

where

$$\Omega_3^2 = \mathrm{d}\chi^2 + \sin^2\chi \left(\mathrm{d}\theta^2 + \sin^2\theta\,\mathrm{d}\phi^2 \right). \tag{2.69}$$

One has

$$\sqrt{-g} = Na^3 \sin^2\chi \sin\theta, \qquad \sqrt{h} = a^3 \sin^2\chi \sin\theta; \tag{2.70}$$

whereas

$$R = \frac{6}{N^2} \left(-\frac{\dot{N}\dot{a}}{Na} + \frac{\ddot{a}}{a} + \left(\frac{\dot{a}}{a}\right)^2 \right) + \frac{6}{a^2}, \qquad K = \frac{3\dot{a}}{Na}. \tag{2.71}$$

$$S_{\text{EG}} = \frac{A_3}{16\pi G} \left(\int_{t_1}^{t_2} \mathrm{d}t\, Na^3 (R - 2\Lambda) - \left[\frac{6\dot{a}a^2}{N} \right]_{t_1}^{t_2} \right), \tag{2.72}$$

where

$$A_3 = \int \sin^2\chi \sin\theta\,\mathrm{d}\chi\,\mathrm{d}\theta\,\mathrm{d}\phi = 2\pi^2. \tag{2.73}$$

The term proportional to $\ddot{a}/a$ in the integrand can be integrated by parts

$$\int_{t_1}^{t_2} \mathrm{d}t\, Na^3 \frac{6}{N^2} \frac{\ddot{a}}{a} = 6 \left(\left[\frac{\dot{a}a^2}{N} \right]_{t_1}^{t_2} - \int_{t_1}^{t_2} \mathrm{d}t\, \dot{a} \frac{\mathrm{d}}{\mathrm{d}t} \frac{a^2}{N^2} \right), \tag{2.74}$$

in which the first term cancels the Gibbons–Hawking–York term. One has

$$S_{\text{EG}} = \frac{3\pi}{4G} \int_{t_1}^{t_2} \mathrm{d}t \left(-\frac{a}{N}\dot{a}^2 + Na - \frac{\Lambda}{3} Na^3 \right). \tag{2.75}$$

The matter part of the action reads

$$S_\phi = \pi^2 \int_{t_1}^{t_2} \mathrm{d}t\, a^3 \left(\frac{1}{N}\dot{\phi}^2 - m^2 N\phi^2 \right). \tag{2.76}$$

Lagrangian with velocity

$$L^{\mathrm{v}} = \frac{3\mathbb{m}}{4G}\left(-\frac{a}{N}v^{a\,2} + Na - \frac{\Lambda}{3}Na^3\right) + \mathbb{m}^2 a^3\left(\frac{1}{N}v^{\phi\,2} - m^2 N\phi^2\right). \qquad (2.77)$$

Canonical momenta

$$p_N := \frac{\partial L^{\mathrm{v}}}{\partial v^N} = 0, \quad p_a := \frac{\partial L^{\mathrm{v}}}{\partial v^a} = -\frac{3\mathbb{m}}{2G}\frac{a}{N}v^a, \quad p_\phi := \frac{\partial L^{\mathrm{v}}}{\partial v^\phi} = 2\mathbb{m}^2\frac{a^3}{N}v^\phi. \tag{2.78}$$

Choosing v^N to be the primary inexpressible velocity, one obtains

$$H^{\mathrm{p}} = -NH_\perp + v^N\Phi, \qquad (2.79)$$
$$H^{\mathrm{c}} = -NH_\perp, \qquad (2.80)$$

where

$$H_\perp := \frac{G}{3\mathbb{m}}\frac{p_a^2}{a} - \frac{1}{4\mathbb{m}^2}\frac{p_\phi^2}{a^3} - \frac{3\mathbb{m}}{4G}\left(\frac{\Lambda}{3}a^2 - 1\right)a - \mathbb{m}^2 m^2 a^3\phi^2, \qquad (2.81)$$
$$\Phi = p_N \qquad (2.82)$$

are the *Hamiltonian constraint* and the primary constraint, respectively.

Evaluating the time evolution of Φ yields

$$[\Phi, H^t]_{\mathrm{p}} = H_\perp, \qquad (2.83)$$

so that the Hamiltonian constraint is indeed a constraint. There is no further constraint, and $[\Phi, H_\perp]_{\mathrm{p}}$ vanishes identically. Therefore there are two and only two first-class constraints.

2.5.1 Einstein–Hilbert action

$$S_{\mathrm{EG}} = S_{\mathrm{EH}} + S_{\mathrm{GHY}}, \qquad (2.84)$$
$$S_{\mathrm{EH}} = \frac{1}{16\mathbb{m}G}\int_{\mathcal{M}} \mathrm{d}^4x\,\sqrt{-g}(R - 2\Lambda), \qquad (2.85)$$

and

$$S_{\mathrm{GHY}} = -\frac{1}{8\mathbb{m}G}\int_{\partial\mathcal{M}} \mathrm{d}^3x\,\sqrt{h}K, \qquad (2.86)$$

which is named after Gibbons and Hawking 1977; York 1972 but actually already mentioned in Einstein 1916. See Dyer and Hinterbichler 2009 for a brief review.

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